

$\bar{B} \rightarrow D^* \ell \bar{\nu}$ at non-zero recoil

Alejandro Vaquero Avilés-Casco^{1,*}, Carleton DeTar¹, Daping Du², Aida El-Khadra^{3,4}, Andreas Kronfeld^{4,5}, John Laiho², and Ruth S. Van de Water⁴

¹Department of Physics and Astronomy, University of Utah, Salt Lake City, Utah, 84112, USA

²Department of Physics, Syracuse University, Syracuse, New York, 13244, USA

³Department of Physics, University of Illinois, Urbana, Illinois, 61801, USA

⁴Fermi National Accelerator Laboratory, Batavia, Illinois, 60510, USA

⁵Institute for Advanced Study, Technische Universität München, 85748 Garching, Germany

Abstract. We present preliminary results from our analysis of the form factors for the $\bar{B} \rightarrow D^* \ell \bar{\nu}$ decay at non-zero recoil. Our analysis includes 15 MILC asqtad ensembles with $N_f = 2 + 1$ flavors of sea quarks and lattice spacings ranging from $a \approx 0.15$ fm down to 0.045 fm. The valence light quarks employ the asqtad action, whereas the heavy quarks are treated using the Fermilab action. We conclude with a discussion of future plans and phenomenological implications. When combined with experimental measurements of the decay rate, our calculation will enable a determination of the CKM matrix element $|V_{cb}|$.

1 Introduction

Although the Standard Model (SM) is widely regarded as a very successful theory, its description of nature is not fully satisfactory, and it is believed to be incomplete. This fact has set a large part of the scientific community in the search for physics Beyond the Standard Model (BSM). One of the most promising places to look for new physics is the flavor sector of the SM. In particular, the unitarity triangle and the CKM matrix elements [1, 2] have been object of extensive research [3], as newer methods and algorithms improved the precision of the measurements. The latest experimental averages [4, 5] reveal a tension of about 3σ between the inclusive and exclusive determinations of $|V_{cb}|$, and it is of foremost importance to investigate the origins of this tension.

One way to obtain $|V_{cb}|$ is from the decay rate of $\bar{B} \rightarrow D^* \ell \bar{\nu}$. Experimental measurements of this decay rate are precise for large momenta, but noisy at low recoil, due to a suppression of the phase space [4]. The lattice approach, on the other hand, excels at low recoil, where it becomes more precise. Forthcoming experiments at LHCb [6] and Belle-II [7], are expected to reduce the experimental uncertainties, so improved lattice calculations are needed to maximize their impact on $|V_{cb}|$ determinations. In this work we present first, preliminary results of our lattice QCD calculation of the form factors for the $\bar{B} \rightarrow D^* \ell \bar{\nu}$ process at non-zero recoil.

*Speaker, e-mail: alexvaq@physics.utah.edu

2 Form factor definitions and ratios

The $\bar{B} \rightarrow D^* \ell \bar{\nu}$ decay is mediated by a $V - A$ weak current. The matrix elements can be decomposed in the following way:

$$\frac{\langle D^*(p_{D^*}, \epsilon^\nu) | \mathcal{V}^\mu | \bar{B}(p_B) \rangle}{2 \sqrt{M_B M_{D^*}}} = \frac{1}{2} \epsilon^{\nu*} \epsilon^{\mu\nu} v_B^\rho v_{D^*}^\sigma h_V(w), \quad (1)$$

$$\frac{\langle D^*(p_{D^*}, \epsilon^\nu) | \mathcal{A}^\mu | \bar{B}(p_B) \rangle}{2 \sqrt{M_B M_{D^*}}} = \frac{i}{2} \epsilon^{\nu*} \left[g^{\mu\nu} (1 + w) h_{A_1}(w) - v_B^\nu \left(v_B^\mu h_{A_2}(w) + v_{D^*}^\mu h_{A_3}(w) \right) \right], \quad (2)$$

where \mathcal{V}^μ and \mathcal{A}^μ are the vector and axial, continuum, electroweak currents, respectively. The 4-vectors v_X^μ ($X = \bar{B}, D^*$) are the 4-velocities, defined as $v_X^\mu = p_X^\mu / M_X$, $w = v_B \cdot v_{D^*}$ is the recoil parameter, the ϵ^ν is the polarization of the D^* final stat eand the $h_k(w)$ ($k = V, A_{1,2,3}$) are the different form factors that contribute to the decay rate.

The differential decay rate can be expressed as

$$\frac{d\Gamma}{dw} = \frac{G_F^2 M_B^5}{4\pi^3} r^3 (1 - r)^2 (w^2 - 1)^{\frac{1}{2}} |\eta_{EW}|^2 |V_{cb}|^2 \chi(w) |\mathcal{F}(w)|^2, \quad (3)$$

with $r = M_{D^*}/M_B$, the ratio of the meson rest masses and η_{EW} , the electroweak correction factor for NLO box diagrams. The kinematic factor $\sqrt{w^2 - 1}$ is what makes experimental measurements so difficult at low recoil. The functions $\chi(w)$ and $\mathcal{F}(w)$ are standard, motivated by the heavy-quark limit:

$$\chi(w) = (1 + w)^2 \lambda(w) \quad \text{with} \quad \lambda(w) = \frac{1 + \frac{4w}{w+1} t^2(w)}{12}, \quad t^2(w) = \frac{1 - 2wr + r^2}{(1 - r)^2}, \quad (4)$$

$$\mathcal{F}(w) = h_{A_1}(w) \sqrt{\frac{H_0^2(w) + H_+^2(w) + H_-^2(w)}{\lambda(w)}}, \quad (5)$$

where

$$X_V(w) = \sqrt{\frac{w-1}{w+1}} \frac{h_V(w)}{h_{A_1}(w)}, \quad (6) \quad X_2(w) = (w-1) \frac{h_{A_2}(w)}{h_{A_1}(w)}, \quad (7)$$

$$X_3(w) = (w-1) \frac{h_{A_3}(w)}{h_{A_1}(w)}, \quad (8) \quad H_0(w) = \frac{w - r - X_3(w) - r X_2(w)}{1 - r}, \quad (9)$$

$$H_\pm(w) = t(w) (1 \mp X_V(w)), \quad (10)$$

The calculation of $\mathcal{F}(w)$ will allow us to determine $|V_{cb}|$ from experimental measurements of the differential decay rate.

3 Simulation details

For this analysis we use the same set of ensembles as in the zero recoil publication [9]. These are 15 different MILC asqtad ensembles with $N_f = 2 + 1$ flavors of sea quarks [10]. The valence light quarks are simulated with the asqtad action [11], whereas the heavy quarks use the clover action with the Fermilab interpretation [12]. The calculations on these ensembles are performed at $\mathbf{p}^2 = 0, \left(\frac{2\pi}{L}\right)^2, \left(\frac{4\pi}{L}\right)^2$ in lattice units, with L the spatial size of the lattice. We distinguish two different cases: momentum perpendicular to the D^* polarization and an average over parallel and perpendicular directions. Fig. 1 shows the size of the ensemble as a function of the lattice spacing and the quark masses. For further details we refer to tables I and II of [9].

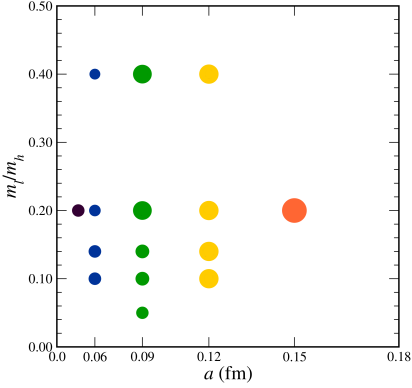


Figure 1. List of available ensembles. The area of the circle is proportional to the statistics.

4 Calculation of correlation functions

4.1 Two-point functions and calculation of rest masses and excited states

The 2-point functions give us information about the energy levels of the states, as well as the overlap factors Z . In this work, the definitions and notation of the 2-point functions follow ref. [13]. We fit 2-point functions for the \bar{B} and the D^* meson with smeared ($1S$) and point (d) sources at source and sink, giving rise to four possible combinations per ensemble and momentum: $(1S, 1S)$, (d, d) , $(1S, d)$ and $(d, 1S)$, where we combined the latter two into a symmetric average.

For our analysis, we perform the joint fit of the three 2-point functions with different smearings at the same time on each ensemble. At non-zero momentum (D^*), we also distinguished between momentum parallel or perpendicular with respect to the meson polarization. Therefore at non-zero momentum our simultaneous fits involve six 2-point functions. We use Bayesian constraints with loosely constrained priors. The main function of the priors here was to prevent the fitter from arriving at absurd values. The prior errors were large enough so small variations in their values didn't change the final results. The correlation matrix for the fitter was obtained using the jackknife method.

We vary the number of states in our fit function: $1 + 1$, $2 + 2$ and $3 + 3$, where the first figure is the number of standard states, and the second gives the number of oscillating states. We choose the $2 + 2$ fits, which balance reasonable p values results with low errors, and are consistent with the results coming from the $3 + 3$ fits. The initial time of the fit is chosen after reaching a plateau in the value of the meson mass, and it is adjusted to be similar in physical units for all the ensembles. The final time is taken to ensure the resulting p value is reasonable, i.e., we check that the p values follow a uniform distribution.

As a consistency check, we test the continuum dispersion relation $E^2 = \mathbf{p}^2 + m^2$ by plotting $E^2/(\mathbf{p}^2 + m^2)$ vs \mathbf{p}^2 (fig. 2). Discretization effects with large masses (in lattice units) could cause this quantity to deviate from ~ 1 . Nonetheless, we clearly see an improvement as the lattice spacing is reduced.

4.2 Three-point function ratios and general considerations

We extract the matrix elements from ratios of 3-point correlation functions. To this end, we use the same setup and notation as in [9, 13]. We construct the non-zero momentum correlation functions as

$$C_J^{\bar{B} \rightarrow D^*}(\mathbf{p}, t, T) = \langle O_{D^*}(-\mathbf{p}, 0) J(\mathbf{p}, t) O_B(\mathbf{0}, T) \rangle. \quad (11)$$

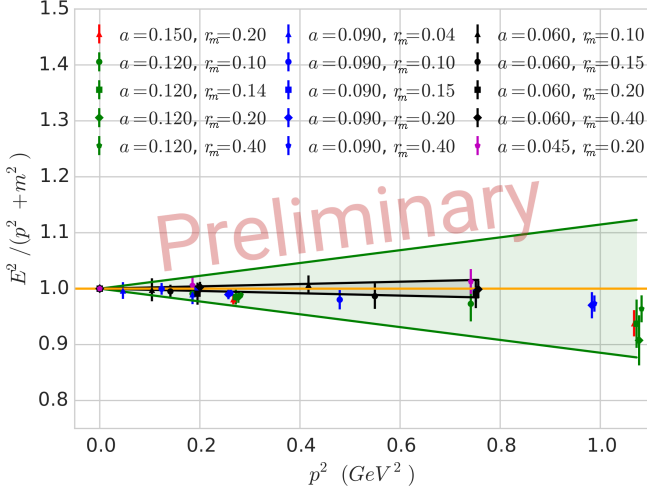


Figure 2. Dispersion relation for the D^* meson, all ensembles. The legend gives the lattice spacing a and the ratio between the light and the strange quark masses, called here r_m . The cones represent the expected discretization effects for the green and the black points. Deviations from $E^2/(p^2 + m^2)$ are observed only for the coarsest ensembles at high momenta, and even in the worst case they are well within the expected discretization errors, of order $O(\alpha_s(ap)^2)$.

where $O_{D^*}(\mathbf{p}, 0)$ and $O_B(\mathbf{0}, T)$ are interpolating operators with the quantum numbers of the D^* and the \bar{B} mesons, respectively, with 3-momentum \mathbf{p} and $\mathbf{0}$, and J is a vector or axial current. Then we can relate the correlation functions to the matrix elements using

$$C_J^{\bar{B} \rightarrow D^*}(\mathbf{p}, t, T) = \sqrt{Z_{D^*}(\mathbf{p})Z_B(\mathbf{0})} e^{-E_{D^*}t} e^{-M_B(T-t)} \langle D^*(\mathbf{p}) | J | B(\mathbf{0}) \rangle + \dots \quad (12)$$

The Z_X factors are the overlap factors obtained from the 2-point functions of the meson X , and the exponentials decay with the energy of the constructed states. The missing extra-terms refer to higher excited states, which are suppressed at large values of t and $T - t$, as well as extra, oscillatory terms, whose contribution is proportional to $(-1)^t$ and $(-1)^{T-t}$. Regarding renormalization and matching, we follow references [9, 13], but at the time of the conference the perturbative ρ factors were not available yet.

In our calculations, the parent meson is always at rest and \bar{B} propagates from time T . The daughter meson propagates from time 0, and both of them are tied together at the current insertion point t . We fix the sink and move the insertion in the range $[0, T]$, expecting the excited states in (12) to die as we move far from the extremes. The different source-sink separations used in this analysis are listed in table IV of [9].

From (12) we construct ratios of 3-point functions that remove as many Z factors and exponentials as possible from the rhs of (12). In order to suppress the effect of the oscillatory terms coming from the 3-point functions, we use the following weighted average [8]

$$\bar{R}(t, T) = \frac{1}{2}R(t, T) + \frac{1}{4}R(t, T+1) + \frac{1}{4}R(t+1, T+1), \quad (13)$$

where R is the ratio to be analyzed, T is the sink time and \bar{R} is the weighted average we use for the analysis. This average ensures that the errors introduced by oscillating states are negligible compared with other errors. As a general ansatz to fit the ratios, we use the following function:

$$r(t) = R \left(1 + A e^{-\Delta E_{M_0}t} + B e^{-\Delta E_{M_T}(T-t)} \right), \quad (14)$$

where R is the ratio of matrix elements we want to extract, T is the sink time, ΔE_{M_0} and ΔE_{M_T} are the differences between the energies of the first excited and the ground state for the mesons living at $t = 0$ and $t = T$ respectively, taken from the 2-point fits, and A and B are fit parameters.

In our analysis we use constrained fits with priors. Whenever the parameters are largely unknown or known with poor precision, a loose prior is set just to guide the fitter. In contrast, whenever the parameters involve quantities obtained in previous fits, the priors are set to the fit outcome and are counted as data points. To remove autocorrelations in our data, we compute the covariance matrix given to the fitter after processing the data with jackknife.

Finally, the parent meson is always smeared using Richardson smearing, but the daughter meson can be smeared or not. This fact gives rise to several versions of each ratio, and following the same technique as in the 2-point function analysis, we perform a joint fit of all the available data to obtain the final results. As an example, we show in fig. 3 one of the fits we use to estimate the recoil parameter.

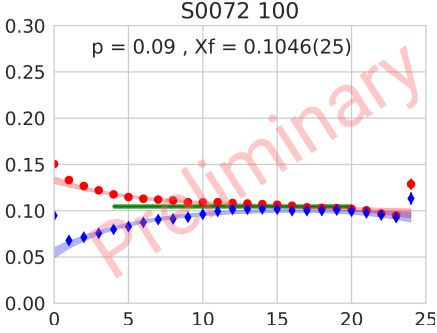


Figure 3. Sample ratio fit for a 3-point function ratio, x_f in this case. The blue points show the data with a smeared interpolating operator, whereas the red points represent a point interpolating operator. The green horizontal bar represents the result for R in (14). In this particular case, we force $B = 0$ because the \bar{B} smearing suppresses all the excited states coming from the sink.

5 Form factors from lattice matrix elements

In order to extract the matrix elements from the lattice, we consider several ratios that reduce the errors in the determination of the form factors.

5.1 Recoil parameter

Considering the \bar{B} meson at rest, one can compute the recoil parameter as

$$w(\mathbf{p}) = \frac{1 + \mathbf{x}_f^2(\mathbf{p})}{1 - \mathbf{x}_f^2(\mathbf{p})}, \quad (15) \quad \mathbf{x}_f(\mathbf{p}) = \frac{\langle D^*(\mathbf{p}) | \mathbf{V} | D^*(\mathbf{0}) \rangle}{\langle D^*(\mathbf{p}) | V^4 | D^*(\mathbf{0}) \rangle}, \quad (16)$$

where the ratio for \mathbf{x}_f comes from the flavor-diagonal transition $D^*(p, s) \rightarrow D^*(p', s')$.

5.2 Axial form factors

5.2.1 Axial double ratio R_{A_1}

Although R_{A_1} is directly obtained from $\langle D^*(\mathbf{p}_\perp) | A_1 | B(0) \rangle$, it is advantageous to use a ratio,

$$R_{A_1}(\mathbf{p}) = \frac{\langle D^*(\mathbf{p}_\perp) | A_1 | \bar{B}(0) \rangle}{\langle D^*(0) | A_1 | \bar{B}(0) \rangle}. \quad (17)$$

Eq. (17) has an important drawback: when expressed as a ratio of 3-point functions, exponentials depending on time are not cancelled for the daughter meson. Hence, we need to add an energy-dependent correction factor before we can fit the ratio on the lattice. In the zero momentum case a

double ratio was used to get rid of the exponential and some overlap factors, reducing the errors and improving the quality of the fit [8, 9]. Here we can do the same,

$$|R_{A_1}(\mathbf{p})|^2 = \frac{\langle D^*(\mathbf{p}_\perp) | A_1 | B(0) \rangle \langle B(0) | A_1 | D^*(\mathbf{p}_\perp) \rangle}{\langle D^*(0) | V_4 | B(0) \rangle \langle B(0) | V_4 | D^*(0) \rangle}. \quad (18)$$

From this quantity we can directly extract $h_{A_1}(w) = 2/(1+w)R_{A_1}$. As we don't have any measurements for the matrix element $\langle B(0) | A_1 | D^*(\mathbf{p}_\perp) \rangle$, we use time reversal \mathcal{T} to reconstruct it from known data,

$$C^{D^* \rightarrow \bar{B}}(\mathbf{p}_\perp, t, T) \xrightarrow{\mathcal{T}} C^{\bar{B} \rightarrow D^*}(\mathbf{p}_\perp, T - t, T). \quad (19)$$

Our preliminary results for $h_{A_1}(w)$, computed from eq. (18) is shown in the right pane of fig. 4.

5.2.2 Ratios R_0 and R_1

The quantities R_0 and R_1

$$R_0(\mathbf{p}) = \frac{\langle D^*(\mathbf{p}_\parallel) | A_4 | B(0) \rangle}{\langle D^*(\mathbf{p}_\perp) | A_1 | B(0) \rangle}, \quad (20) \quad R_1(\mathbf{p}) = \frac{\langle D^*(\mathbf{p}_\parallel) | A_1 | B(0) \rangle}{\langle D^*(\mathbf{p}_\perp) | A_1 | B(0) \rangle}, \quad (21)$$

encode the behavior of $h_{A_2}(w)$ and $h_{A_3}(w)$ as

$$R_0 = \frac{\sqrt{w^2 - 1}(1 - h_{A_2} + wh_{A_3})}{(1 + w)h_{A_1}}, \quad (22) \quad R_1 = w - \frac{(w^2 - 1)h_{A_3}}{(1 + w)h_{A_1}}. \quad (23)$$

Preliminary results for $h_{A_2}(w)$ and $h_{A_3}(w)$ are reported in fig. 5.

5.3 Vector form factor

The previously defined quantity $X_V(w)$ can be measured as the following ratio of matrix elements:

$$h_V = \frac{R_{A_1}}{\sqrt{w - 1}} X_V \quad (24) \quad X_V(\mathbf{p}) = \frac{\langle D^*(\mathbf{p}_\perp) | V_1 | B(0) \rangle}{\langle D^*(\mathbf{p}_\perp) | A_1 | B(0) \rangle}. \quad (25)$$

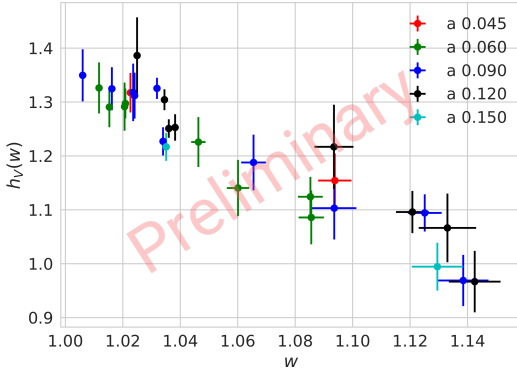
When this ratio is expressed in terms of lattice 3-point correlation functions, all the exponentials and overlap factors are cancelled, and the result yields directly the quotient of form factors we are looking for. Our preliminary result for $h_V(w)$ is shown in the left pane of fig. 4.

5.4 Results for the form factors as a function of the recoil parameter

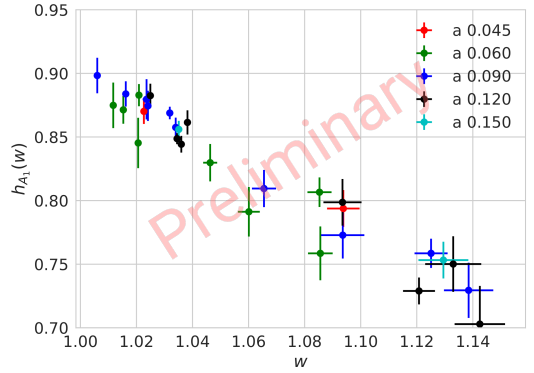
In figs. 4,5 we show the preliminary results for the axial and vector form factors, with statistical errors only, without rho factors, and before taking the chiral-continuum extrapolation. It is important to notice that near zero recoil (for small $w - 1$), $\mathcal{F}(w)$ is dominated by h_{A_1} , because the contributions from h_V , h_{A_2} , and h_{A_3} are suppressed by kinematic factors (see eqs. (7), (8) and (6)).

6 Summary and future work

In this paper we have presented first preliminary results from our lattice QCD calculations of the form factors for $\bar{B} \rightarrow D^* \ell \bar{\nu}$ at non-zero recoil. Still to be completed are further improvements in the fits to the correlation functions, after which we plan to study the chiral-continuum extrapolations, use the z -expansion to parametrize the shape, and construct a complete, systematic error budget.

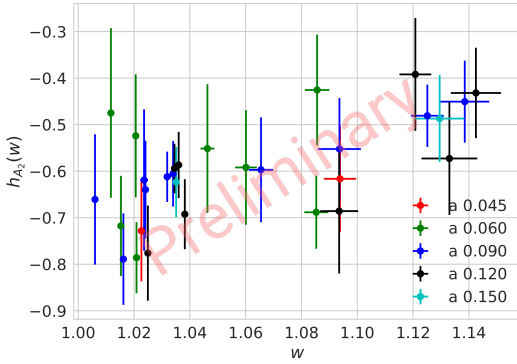


(a) $h_V(w)$ form factor.

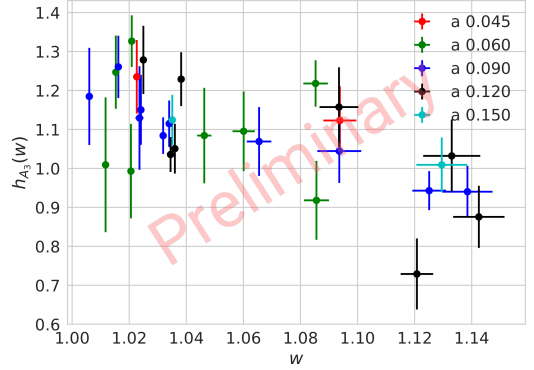


(b) $h_{A_1}(w)$ form factor.

Figure 4: Preliminary results for $h_V(w)$ and $h_{A_1}(w)$ from X_V and the double ratio R_{A_1} . The contribution of $h_V(w)$ to $\mathcal{F}(w)$ is suppressed due to kinematic factors, and the final result is clearly dominated by $h_{A_1}(w)$.



(a) $h_{A_2}(w)$ form factor.



(b) $h_{A_3}(w)$ form factor.

Figure 5: Preliminary results for $h_{A_2}(w)$ and $h_{A_3}(w)$ from R_0 and R_1 . We expect to reduce the errors in these form factors when we refine the analysis. However, the contribution of $h_{A_2}(w)$ and $h_{A_3}(w)$ to $\mathcal{F}(w)$ is highly suppressed by kinematic factors at low recoil. Therefore the final total error is dominated by the errors coming from $h_{A_1}(w)$.

7 Acknowledgements

Computations for this work were carried out with resources provided by the USQCD Collaboration, the National Energy Research Scientific Computing Center and the Argonne Leadership Computing Facility, which are funded by the Office of Science of the U.S. Department of Energy; and with resources provided by the National Institute for Computational Science and the Texas Advanced Computing Center, which are funded through the National Science Foundation's Teragrid/XSEDE Program. This work was supported in part by the U.S. Department of Energy under grants No. DE-FC02-06ER41446 (C.D.) and No. DE-SC0015655 (A.X.K.), by the U.S. National Science Foundation under grants PHY10-67881 and PHY14-17805 (J.L.), PHY14-14614 (C.D., A.V.); by the

Fermilab Distinguished Scholars program (A.X.K.); by the German Excellence Initiative and the European Union Seventh Framework Program under grant agreement No. 291763 as well as the European Union's Marie Curie COFUND program (A.S.K.).

Fermilab is operated by Fermi Research Alliance, LLC, under Contract No. DE-AC02-07CH11359 with the United States Department of Energy, Office of Science, Office of High Energy Physics. The publisher, by accepting the article for publication, acknowledges that the United States Government retains a non-exclusive, paid-up, irrevocable, world-wide license to publish or reproduce the published form of this manuscript, or allow others to do so, for United States Government purposes.

References

- [1] N. Cabibbo, Phys. Rev. Lett. **10**, 531 (1963), [,648(1963)]
- [2] M. Kobayashi, T. Maskawa, Prog. Theor. Phys. **49**, 652 (1973)
- [3] M. Antonelli et al., Phys. Rept. **494**, 197 (2010), 0907.5386
- [4] C. Patrignani et al. (Particle Data Group), Chin. Phys. **C40**, 100001 (2016)
- [5] Y. Amhis et al. (2016), 1612.07233
- [6] R. Aaij, B. Adeva, M. Adinolfi, Z. Ajaltouni, S. Akar, J. Albrecht, F. Alessio, M. Alexander, S. Ali, G. Alkhazov et al. (LHCb Collaboration), Tech. Rep. CERN-LHCC-2017-003, CERN, Geneva (2017), <http://cds.cern.ch/record/2244311>
- [7] T. Abe et al. (Belle-II), Tech. rep., KEK (2010), 1011.0352, <http://arxiv.org/pdf/1011.0352v1>
- [8] C. Bernard et al., Phys. Rev. **D79**, 014506 (2009), 0808.2519
- [9] J.A. Bailey et al. (Fermilab Lattice, MILC), Phys. Rev. **D89**, 114504 (2014), 1403.0635
- [10] A. Bazavov et al. (MILC), Rev. Mod. Phys. **82**, 1349 (2010), 0903.3598
- [11] G.P. Lepage, Phys. Rev. **D59**, 074502 (1999), hep-lat/9809157
- [12] A.X. El-Khadra, A.S. Kronfeld, P.B. Mackenzie, Phys. Rev. **D55**, 3933 (1997), hep-lat/9604004
- [13] J.A. Bailey et al. (MILC), Phys. Rev. **D92**, 034506 (2015), 1503.07237