

An abrupt decrease in Southern Hemisphere terrestrial temperature during the Eocene–Oligocene transition



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ABSTRACT

Earth's climate transitioned from the greenhouse conditions of the Eocene to the icehouse of the Oligocene ~34 million years ago. Global cooling was triggered by a decrease in atmospheric $p\text{CO}_2$ and is associated with the first appearance of large, permanent ice-sheets on Antarctica. The Eocene–Oligocene transition (EOT) is marked by a large two-step increase in the $\delta^{18}\text{O}$ of benthic foraminifera, reflecting a combination of cooler global temperatures and continental ice growth. We measured the hydrogen isotope composition of hydrated volcanic glasses deposited in Patagonia during the EOT to reconstruct stable isotope hydrology and associated changes in temperature during this critical climate transition. These data show changes in stable hydrogen isotope ratios that are consistent with Patagonian terrestrial mean annual temperature decreasing by ~5 °C during the EOT before returning to 2 °C below pre-EOT values by the Early Oligocene. This is approximately synchronous with Northern Hemisphere cooling during the EOT, and indicates close coupling of temperature and atmospheric $p\text{CO}_2$.

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1. Introduction

The abrupt shift of global climate from the ice-free greenhouse climate of the Late Eocene (Pearson et al., 2007) to the icehouse climate of the Early Oligocene (Galeotti et al., 2016) is perhaps the most dramatic climate transition of the Cenozoic. The transition between climate states occurs at the end of a cooling trend in the Late Eocene, following which temperature fluctuated before settling into a stable cool temperature regime by ~32 Ma (e.g., Lear et al., 2008). This change in climate state is marked by a rise in the $\delta^{18}\text{O}$ of benthic foraminifera (Zachos et al., 2001). While the onset of Antarctic glaciation (Galeotti et al., 2016) has been attributed to the opening of ocean gateways and thermal isolation of Antarctica (Kennett, 1977), coupled climate and ice sheet models show that it was far more likely a result of long-term decreases in atmospheric $p\text{CO}_2$ (DeConto and Pollard, 2003; Goldner et al., 2014). These models predict that the transition from ice-free to ice-present conditions should have been rapid and occurred in two steps. This two-step transition is documented in the

$\delta^{18}\text{O}$ of well-preserved benthic foraminifera (Coxall et al., 2005; Lear et al., 2008) and ε_{Nd} estimates of Antarctic weathering (Scher et al., 2011), and is contemporaneous with the timing of falling atmospheric $p\text{CO}_2$ (Pagani et al., 2005; Pearson et al., 2009; Zhang et al., 2013).

However, the terrestrial record of the EOT is less clear. This is due to the fact that in the terrestrial realm, high-quality age control is often lacking and sedimentation may be more intermittent in nature than in deep water marine settings. Records of the EOT in the Northern Hemisphere are marked by regionally variable changes in hydrology and temperature (Hren et al., 2013; Sheldon, 2009; Sheldon et al., 2012, 2016; Terry, 2001; Zanazzi et al., 2007), although most show cooling in the range of 4–8 °C (Hren et al., 2013; Zanazzi et al., 2007) from the Late Eocene to Early Oligocene. In contrast, the only terrestrial records of the EOT in the Southern Hemisphere come from Australia (Korasidis et al., 2018) and Patagonia, where phytoliths (Dunn et al., 2015; Strömberg et al., 2013) and $\delta^{18}\text{O}$ of tooth enamel and soil carbonate (Kohn et al., 2004, 2015) have been interpreted to show no change in climate across the EOT. The only large change observed in Patagonia during this period is a transient shift in the delivery of dust and suppression of wildfire associated with the earliest Oligocene glacial maximum (EOGM) (Selkin et al., 2015). A lack of observed cooling in Patagonia across the EOT is particularly

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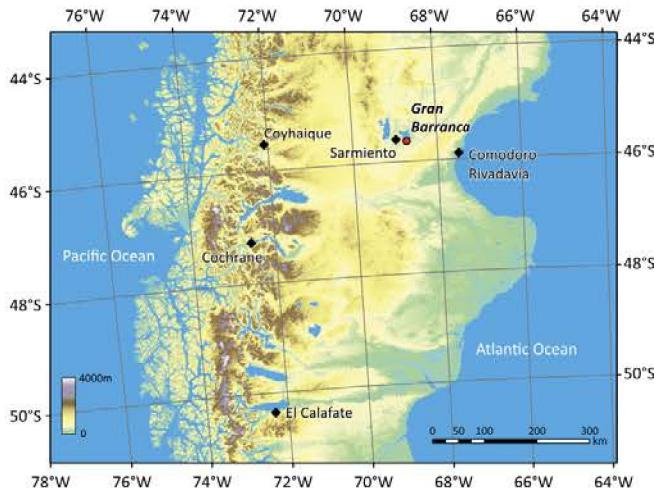


Fig. 1. Topography of central Patagonia at 1:4,000,000 scale. The location of the Gran Barranca section is indicated by the red circle; black diamonds identify regional cities of note. (For interpretation of the colors in the figure(s), the reader is referred to the web version of this article.)

boring Antarctica through this interval (Galeotti et al., 2016), as well as marine U^{37} and TEX_{86} surface temperature data from DSDP Site 511 on the Falkland Plateau adjacent to eastern Patagonia (Liu et al., 2009) that show a large decrease in sea surface temperature (SST) during the EOT. The interpretation of a static climate in Patagonia during the EOT therefore challenges our understanding of ocean-atmosphere connectivity. Here we present a new high-resolution record of stable hydrogen isotopes (δD) from volcanic glass deposited in South America ($\sim 46^{\circ}$ S) during this critical transition. Stable isotopes of precipitation are, to first order, controlled by temperature, and thus provide way to quantify regional climatic

change during the EOT. These new data address the apparent dichotomy between hemispheres, and indicate a significant isotopic and climatic response to falling atmospheric pCO_2 similar to that observed in the Northern Hemisphere.

2. Geologic setting

We collected 23 volcanic ash samples from the Sarmiento Formation (Ameghino, 1889, 1906; Bellosi, 2010b; Madden et al., 2010; Simpson, 1930) at Gran Barranca in Chubut, Argentina (Fig. 1) to develop a new hydrogen isotope record of hydration water in ancient volcanic glass. The stratigraphic section at Gran Barranca is remarkable in its exceptional degree of stratigraphic completeness and highly resolved geochronological constraints (Dunn et al., 2013; Ré et al., 2010a, 2010b). The Eocene–Miocene Sarmiento Formation contains six members, including the valley-filling Eocene–Oligocene Vera Member. The Vera Member represents relatively uninterrupted deposition, showing poor soil development and no evidence of internal discontinuities (Bellosi, 2010a). This unit is composed primarily of felsic wind-transported volcanic material that was deposited following eruptions of the Andean volcanic arc to the west (Bellosi, 2010a) (Fig. 2). We sampled a continuous stratigraphic profile at Profile K, one of the thickest exposures of the Vera Member (Bellosi, 2010b). The high sedimentation rates in the Vera Member enabled sampling across the EOT with ca. 25 kyr resolution. All sample heights were referenced to the measured section and dated volcanic tuffs (Dunn et al., 2013; Ré et al., 2010a) in the Vera Member and the Lower and Upper Puesto Almendra Members. While the Vera Member lies stratigraphically between the Lower and Upper Puesto Almendra Members, it is not physically between them in outcrop due to the valley-filling nature of the Vera. Age control is provided by zircon U/Pb geochronology on volcanic tuffs (Dunn et al., 2013) and magnetostratigraphy (Ré et al., 2010b) (Fig. 2).

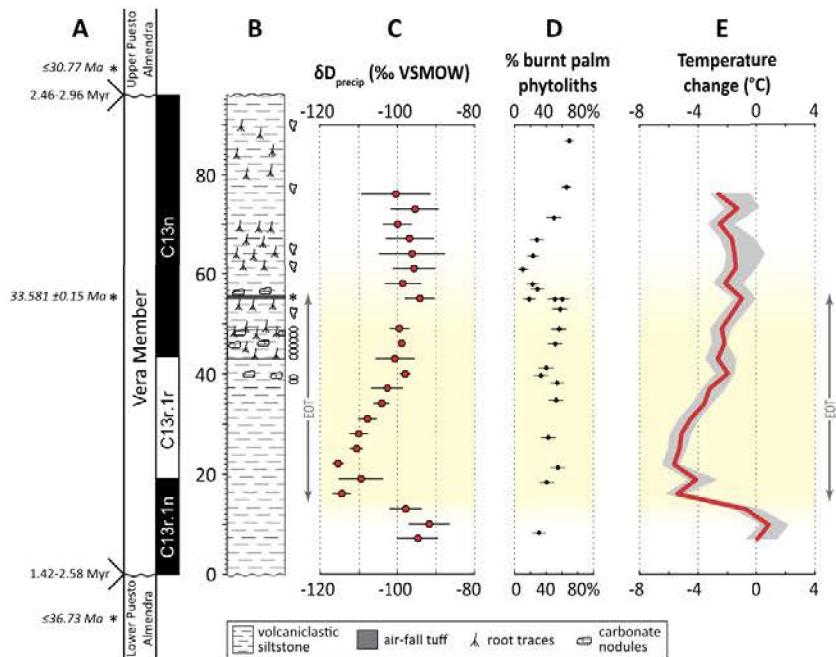


Fig. 2. Stratigraphy and geochemistry of studied sedimentary section of the Vera Member of the Sarmiento Formation (Bellosi, 2010b). (A) Radiometric (Dunn et al., 2013), lithostratigraphic (Bellosi, 2010b), and magnetostratigraphic (Dunn et al., 2013; Ré et al., 2010b) framework for the Vera Member. Radiometric dates indicated by asterisks

3. Hydration of volcanic glass

Reconstructing ancient stable isotopes of precipitation requires a material that samples the stable isotope chemistry of precipitation during formation and then remains stable over geologic timescales. We measured δD of volcanic glass preserved in the Vera Member to reconstruct the isotopic composition of precipitation through time. During volcanic eruptions, magma degasses and forms glass fragments, which typically retain 0.1–0.3 wt. % of magmatic water (Friedman and Smith, 1958; Martin et al., 2017; Ross and Smith, 1955). These glasses are then hydrated by meteoric water, which saturates at up to 6 wt. % water (Friedman et al., 1993a; Cassel and Breecker, 2017). Glass hydration occurs over 1–10 kyr (Cassel et al., 2009; Friedman et al., 1993b; Seligman et al., 2016), and is rapid compared to timescales of climate change during the Late Eocene to Early Oligocene. During the saturation process, alkali cations (e.g., Na^+ , K^+) are exchanged for H^+ from the meteoric water (Cerling et al., 1985; Cassel and Breecker, 2017). As this exchange occurs, the rim of the glass particle becomes impermeable to further diffusion of water (Friedman and Long, 1976) and thus glass particles do not undergo significant further isotopic exchange with the environment (Friedman and Long, 1984). Preservation of glass in and of itself attests to the fact that ancient sedimentary environments in which glass is found did not allow sufficient fluid availability to allow complete conversion of glass to secondary clays. Following hydration, the resulting glass particles effectively preserve a record of the time-averaged δD of precipitation during hydration (Dettinger and Quade, 2015; Friedman et al., 1993b; Seligman et al., 2016).

4. Materials and methods

Fresh rock samples were collected from felsic volcaniclastic sediments of aeolian origin (Bellosi, 2010a) in the measured sedimentary section (Fig. 2). We separated volcanic glasses from clays and other hydrous minerals using established mineral separation procedures. Rock samples were crushed with a mortar and pestle and then sieved to a 63–180 μm fraction. The sieved fraction was agitated in deionized water and left to settle until particles $>63 \mu\text{m}$ had settled; the remaining clay-rich liquid was decanted. This was repeated 3–5 times, depending on the concentration of clay particles. Samples were then placed in a sonicating bath with sodium pyrophosphate deflocculant for 10 min in order to remove any clay particles adhering to grain surfaces. Following sonication, the samples were agitated in a 10% HCl solution to remove any carbonate. Samples were then rinsed repeatedly and dried at 40 °C. Magnetic grains were separated using repeated runs through a Frantz magnetic separator at progressively increasing magnet strengths. Resulting non-magnetic separates were density-separated using lithium polytungstate heavy liquid and then inspected using a microscope to determine the purity of glass separates. A subset of samples was inspected using SEM to confirm their preservation and the lack of significant secondary hydrous minerals (Fig. 3).

High-purity samples were weighed into silver capsules, stored in a desiccating jar, and then dried in a vacuum oven at 80 °C for 24 h prior to analysis. Drying at higher temperature and longer time produces essentially identical δD and water contents (Fig. S1). The δD of glass separates was measured using a thermal conversion elemental analyzer (TC/EA) attached to a Thermo MAT 253 IRMS at the University of Connecticut. Isotopic compositions of

relative to VSMOW and errors reflect standard deviation of multiple analyses. The δD of the most repeated standards was PEF-1 = $-102 \pm 2\text{‰}$, NBS-22 = $-118 \pm 2\text{‰}$, and KGa-1b = $-65 \pm 2\text{‰}$. The known water content of NBS-22 was used to calculate water contents of glasses. Measured δD of glass was used to calculate the δD of hydration water (precipitation) using the fractionation factor of Friedman et al. (1993a), which has been replicated in recent studies (Seligman et al., 2016; Porter et al., 2016). During interpretation, we assumed that any remnant magmatic water had a negligible effect on the measured δD because (1) the δD of nearby degassed rhyolite (Castro et al., 2014) is similar to the range of values we observe and (2) magmatic water is likely to have undergone complete isotopic exchange with hydration water in glasses of Eocene–Oligocene age.

5. Results and age models

The Vera Member is known to span a time interval of ~ 0.7 to ~ 2.0 Myr that includes the EOT (Dunn et al., 2013). During this time, the δD of volcanic glasses in the Vera Member abruptly decreases from $-125 \pm 3\text{‰}$ to $-142 \pm 3\text{‰}$ and then recovers more gradually between 10 and 40 m in the section to nearly pre-excursion values of $-128 \pm 2\text{‰}$ (Fig. 2, Table 1). To place this geochemical excursion into a time framework, we use the depositional age model of Dunn et al. (2013), which integrates the available radiometric ages and magnetostratigraphy. Their age model for the Vera Member includes the 33.581 ± 0.15 Ma La Cancha tuff (Fig. 2), as well as radiometric constraints from the Lower Puesto Amenda Member and Upper Puesto Almendra Member, which, respectively, lie stratigraphically below and above the Vera Member (Dunn et al., 2013; see Fig. 2). The preferred age model of Dunn et al. (2013) correlates the reversed magnetic polarity interval in the lowest part of the Vera Member with chron C13r.1n. The existence of 'cryptochron' C13r.1 is well established (Channell et al., 2003), and we follow this interpretation for the age model of the Vera Member. This indicates that the abrupt δD decrease occurs at ~ 34.2 Ma, followed by the slower recovery between ~ 34.0 and ~ 33.5 Ma (Table 1). This chronology is generally consistent with other records of the timing of change associated with the EOT elsewhere on land and in the ocean (e.g., Coxall et al., 2005; Hren et al., 2013; Lear et al., 2008).

6. Precipitation stable isotopes and temperature

The hydrogen isotopic composition of volcanic glass within the Vera Member records a decrease of $\sim 20\text{‰}$ during the time period associated with the EOT. This indicates a significant change in the isotopic composition of ambient water in the region during this transition period. There are a number of factors that influence the isotopic composition of water in South America, including changes in atmospheric circulation (e.g., Hoke et al., 2013; Liu et al., 2010) and topography (e.g., Poage and Chamberlain, 2001). However, globally and locally, stable isotopes of precipitation have a clear observed relationship with temperature at mid and high latitudes (Bowen and Wilkinson, 2002; Dansgaard, 1964; Rozanski et al., 1993) due to temperature-dependent fractionation that occurs during (1) evaporation from the source and (2) partitioning between vapor and precipitation during rainout (e.g., Bowen and Revenaugh, 2003; Gat, 1996). More specifically, the net fractionation of stable isotopes reflects the difference between the temperature at which this partitioning occurs (Pierrehumbert,

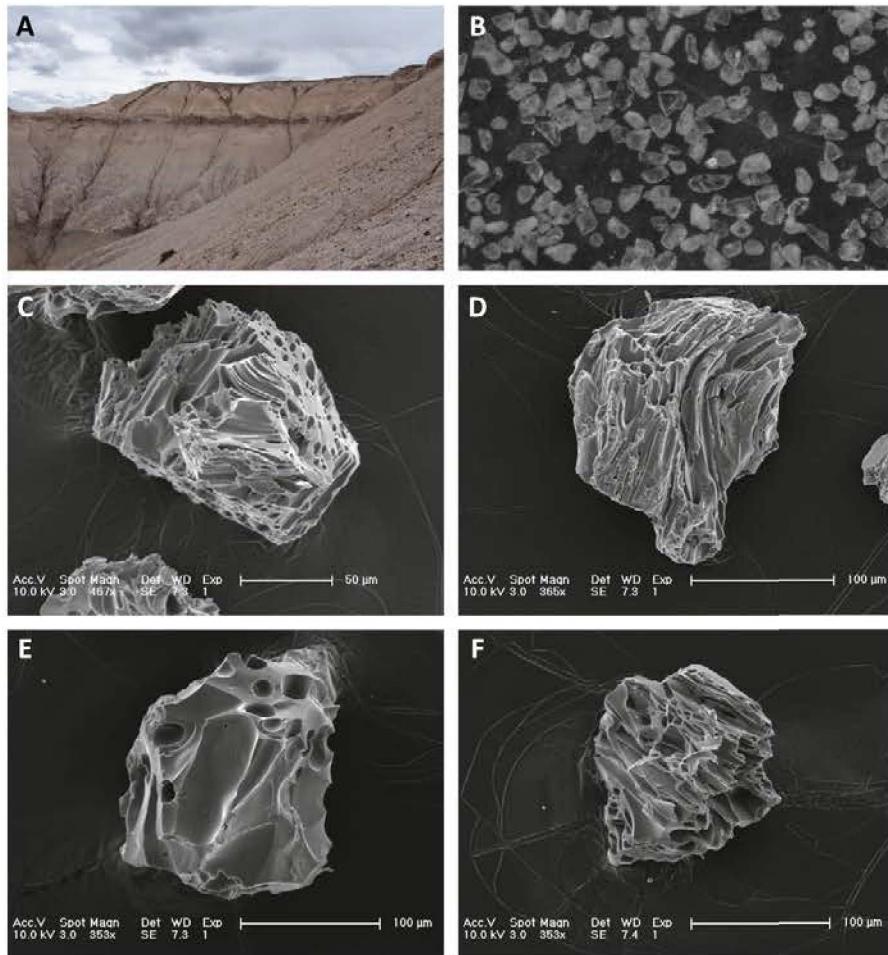


Fig. 3. (A) Field photo of the Vera Member at Gran Barranca, near Profile K. (B) Light microscope view of typical volcanic glass separate. 2.5X magnification, field of view is 3 mm across. (C–F) Scanning electron microscope images of Vera Member glass separates from samples 13GB13, 13GB17, 13GB23, and 13GB25, respectively. Images show primary volcanic textures and sharp edges, as well as minimal secondary clay particles on grain surfaces.

Table 1
Measured δD of volcanic glasses in the Vera Member, Sarmiento Formation, and reconstructed meteoric water values.

Sample	Stratigraphic height (m)	Model age (Ma)	$\delta D_{\text{glass}}^a$ (‰ VSMOW)	N	2 σ error	H_2O (wt. %)	$\delta D_{\text{water}}^b$ (‰ VSMOW)
13GB12	7	34.331	-125	3	5.4	0.7%	-95
13GB13	10	34.284	-122	2	1.7	4.3%	-92
13GB14	13	34.237	-128	3	4.1	0.6%	-98
13GB15	16	34.190	-144	3	1.9	1.0%	-114
13GB16	19	34.085	-139	2	5.8	4.1%	-110
13GB17	22	34.038	-145	3	1.5	0.9%	-115
13GB18	25	33.990	-140	3	1.7	0.8%	-111
13GB19	28	33.943	-139	3	2.5	0.6%	-110
13GB20	31	33.895	-137	3	2.4	0.7%	-108
13GB21	34	33.848	-134	3	2.1	0.8%	-104
13GB22	37	33.800	-132	2	4.2	6.1%	-103
13GB23	40	33.753	-128	3	1.4	0.7%	-98
13GB25	43	33.705	-130	3	4.6	0.6%	-101
13GB26	46	33.674	-129	3	0.6	0.7%	-99
13GB28	49	33.643	-129	3	2.6	0.7%	-99
13GB32	55	33.581	-124	2	3.7	4.9%	-94
13GB33	58	33.571	-128	3	4.8	0.8%	-99
13GB34	61	33.560	-126	3	5.6	0.7%	-96
13GB35	64	33.550	-126	2	5.1	2.6%	-96
13GB36	67	33.539	-127	3	6.3	0.5%	-97
13GB37	70	33.529	-130	3	3.8	0.5%	-100
13GB38	73	33.518	-125	3	6.3	0.5%	-95

records requires an assessment of the potential effects of temperature change on isotopic composition as well as factors such as changes in water source or mixing of air masses.

Gran Barranca is situated approximately 150 km from the Atlantic Ocean to the east and less than 500 km from the Pacific to the west at $\sim 45^{\circ}\text{S}$ latitude. Over the relatively short time interval represented in the Vera Member sediments, it is not geologically plausible for the latitudinal position of Patagonia, the distance of Gran Barranca from vapor sources in the Pacific Ocean, or the height of the Andes to have changed significantly. Precipitation in Patagonia is fed by water vapor carried by the Southern Hemisphere westerly wind belt (SWWB) (Garreaud et al., 2009), although minor mixing with vapor from Atlantic can occur near the east coast (Garreaud et al., 2013). The position and direction of the SWWB results from the speed and direction of Earth's rotation (Schneider, 2006; Williams, 1988), which has changed little over the Cenozoic (Parrish et al., 1982). Thus we consider the SWWB a persistent feature of global and Patagonian climate. While the SWWB has been observed to shift north and south by several degrees of latitude in response to changes in global climate (Koffman et al., 2014; Lamy et al., 2001; Moy et al., 2008), volcanic ash dispersal patterns confirm that Patagonia has remained within the band of westerly winds throughout the Cenozoic (Bellosi, 2010a; Rapela et al., 1988). Wind strength might have changed over the Cenozoic, although the pattern of westerly circulation did not. Consequently, changes in moisture source are not a significant factor in determining precipitation δD over geologic time scales in Patagonia. We also infer that evaporation had little to no effect on the δD of meteoric water during both precipitation and storage, based on modern observations of precipitation δD (Smith and Evans, 2007; Stern and Blisniuk, 2002) and the sedimentary facies present at Gran Barranca (Bellosi, 2010a, 2010b; Bellosi and Gonzalez, 2010), respectively.

Over short time scales (hours to years), δD of precipitation can be highly variable (e.g., Coplen et al., 2008), but values collapse around a mean at decadal or longer timescales (Liu et al., 2010). Thus, records that integrate precipitation δD over long time scales preserve such a time-averaged value. Because volcanic glass hydrates over 1–10 kyr (Friedman and Long, 1976), it thus records averages of precipitation over those millennial timescales. Conversely, it also means that variation on shorter-than-kyr timescales—synoptic events and shorter-term variation such as ENSO and the sunspot cycle—does not appear in the record.

Because the mid-latitude climate of Patagonia minimizes non-temperature effects on hydrogen isotopes of precipitation, we use modern hydrogen isotope and temperature data from IAEA-GNIP stations in Patagonia to quantify the relationship between regional temperature and precipitation δD . This approach that has been previously utilized to reconstruct Cenozoic temperatures in this region (Kohn et al., 2004) and specifically includes seasonal data that is known to reflect the range of potential variability in moisture source. We selected four low-elevation stations across Patagonia with relatively long precipitation stable isotope records: Bariloche (-41.15°S , -71.33°W , records span 1996–2002), Puerto Montt (-41.47°S , -72.93°W , 1964–2014), Coyhaique (-45.35°S , -72.07°W , 1988–1999), and Punta Arenas (-53.0°S , -70.51°W , 1990–2014). These sites span a range of modern climatic conditions, including wind speed, MAT, and precipitation amount. Data from each station was averaged monthly, which is reported in Table S1.

All sites show a strong linear relationship (r^2 ranges from 0.70

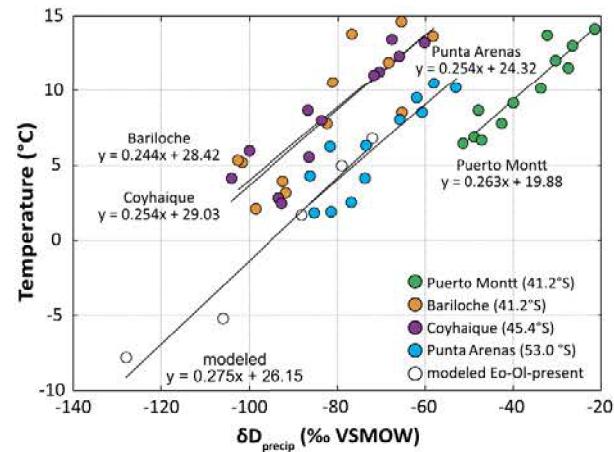


Fig. 4. δD –temperature relationships for IAEA-GNIP stations across Patagonia and isotope-enabled GCM data for the geologic past, including the time around the EOT (Feehins et al., 2014). Note the consistent slope of modern and ancient relationships in Patagonia.

temperature across space and short timescales and demonstrates that, independent of season, all southern Patagonia sites reflect the same dependence of stable hydrogen isotopes on temperature. A regional mean relationship was calculated from the entire data set, yielding a single equation that relates regional temperature to stable hydrogen isotope ratio: T ($^{\circ}\text{C}$) = $0.254 (\delta\text{D}) + 25.41$. As a conservative approach, we use the stations with the steepest δD – T slopes (Puerto Montt, $y = 0.263x + 19.88$) and shallowest (Bariloche, $y = 0.244x + 28.42$) as end members that describe the entire range of potential climate variation due to latitudinal shifts in climate belts and ocean temperature change. These regressions were then applied to the reconstructed precipitation δD from the Vera Member to estimate the mean and possible range of temperature change during the EOT (Fig. 2). Importantly, we base our interpretations on the *relative* change in δD , which can be effectively reconstructed in the geologic past using the slope of the δD – T relationship.

While GNIP observations support the use of this δD – T relationship across space and short time scales, some model predictions indicate that the slope of the δ – T relationship for northern hemisphere sites during warm periods in the geologic past may be shallower than modern (Fricke and O'Neil, 1999; Jouzel et al., 1997; Kohn et al., 2004; Secord et al., 2010). However, a comparison of modern GNIP data against modeled MAT and δD data for the southern hemisphere during the late Eocene to early Oligocene shows a similar δ – T relationship to the modern. Feehins et al. (2014) used an isotope-enabled GCM to simulate climate on the nearby Antarctic Peninsula for the late Eocene, Eocene–Oligocene transition, early and late Oligocene, and present. A regression through these data yields nearly the same slope as the modern GNIP data (Fig. 4). The decision to use a modern δ – T slope to reconstruct the past temperatures in Patagonia is supported by the agreement of modern observations with reconstructed slopes for the mid-latitudes in the geologic past (Kohn et al., 2004; Secord et al., 2010) and the observation that Patagonia has remained at approximately its modern latitude through the Cenozoic (Seton et al., 2012). In addition, model predictions for the precipitation stable isotope response for Patagonia from the late Eocene to early Oligocene (Feehins et al., 2014) are comparable to the magnitude of change that we observe. Agreement between modeled ancient

ancient, including vapor source, atmospheric circulation patterns, topography, convection, and continentality (Charles et al., 1994; Schmidt et al., 2007). While the use of the modern observed relationship in the geologic past of Patagonia appears justified, care should be used when applying this approach to other regions in deep time.

Using the observed $\delta D-T$ relationship, we reconstruct regional terrestrial temperature from the glass δD , taking into account the effect of stepwise changes in ice volume during the EOT (Bohaty et al., 2012) on source composition. To do this, we assume instantaneous ice growth and apply a correction of 0.6‰ in $\delta^{18}\text{O}$, or 4.8‰ δD . Following existing work on the effect of ice growth on marine $\delta^{18}\text{O}$ (Bohaty et al., 2012; Lear et al., 2008), we apply a decrease in $\delta^{18}\text{O}$ of 0.2‰ at the first step in the benthic foraminiferal record (16 m in our measured section) and a decrease of 0.4‰ at the second step (25 m in our section). While we acknowledge this does not likely reflect the true complexity of the pacing of ice volume growth, this approach provides a conservative treatment of the data. Modern precipitation stable isotope data and GCM simulations of Eocene to Oligocene precipitation stable isotopes for the nearby Antarctic Peninsula (Feehins et al., 2014) predict that for every 1 °C change in temperature, the δD of precipitation should change on the order of ~4‰ or more. The rapid decrease in glass and meteoric δD observed in the Vera Member at Gran Barranca indicates a rapid cooling of $\sim 5 \pm 1$ °C (range reflects uncertainty in the $\delta D-T$ relationship) during the onset of the EOT (Fig. 5). By ~33.5 Ma (above the 33.581 Ma La Cancha Tuff), temperatures return to approximately -2 °C relative to pre-EOT values (Fig. 5). The initial fall in temperature is coincident with the onset of the $\delta^{18}\text{O}$ shift in the marine benthic foraminifera record (Fig. 5), and the recovery occurs over ~ 600 kyr. Unlike marine records, terrestrial records did not suffer from carbonate dissolution (e.g., Coxall et al., 2005), thus the Vera Member may capture an interval of time typically missing from marine records of the EOT (albeit at much lower resolution).

A cooling of ~ 5 °C on land in Patagonia matches well with the temperature change observed during the EOT in Northern Hemisphere records (Hren et al., 2013; Zanazzi et al., 2007). In addition, this magnitude of change is identical to the modeled and observed magnitude of change in SST at high latitude marine sites, including DSDP Site 511 located on the nearby Falklands Plateau (Liu et al., 2009). While changes in hydrology appear to be heavily dependent on the regional setting, differing even in sign (Hren et al., 2013; Sheldon, 2009), the terrestrial cooling signal appears to be a common characteristic of Northern Hemisphere EOT records. Our observation that such cooling also occurred in the Southern Hemisphere emphasizes the global and at least broadly synchronous nature of the terrestrial climate response to the EOT.

7. Discussion

High latitude terrestrial cooling in South America was broadly synchronous with the initiation of cooling in the low-latitude ocean (Coxall et al., 2005; Lear et al., 2008) and the onset of the benthic $\delta^{18}\text{O}$ increase (Fig. 5). However, the temperature change at Gran Barranca reflects an apparently abrupt cooling and partial rebound within the longer-term trend (Galeotti et al., 2016; Lear et al., 2008) of Late Eocene to Early Oligocene cooling. We suggest the general synchrony of marine and terrestrial records indicates that terrestrial climate responded directly to falling global $p\text{CO}_2$. Broadly, this pattern of temperature change strengthens the idea that the first step in the benthic $\delta^{18}\text{O}$ record was primarily

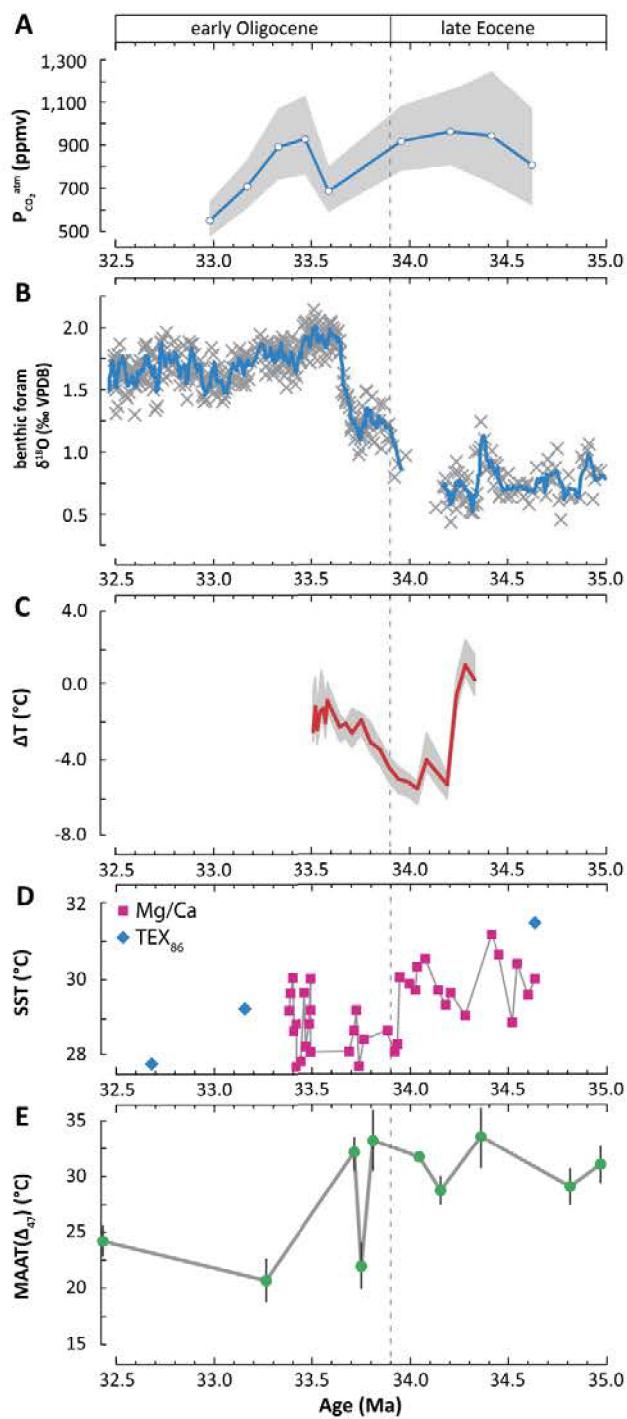


Fig. 5. Comparison of the terrestrial Patagonian EOT record to regional and global marine records of the EOT. (A) Marine $\delta^{11}\text{B}$ -derived record of atmospheric $p\text{CO}_2$ (Pearson et al., 2009), replotted for changes in the $\delta^{11}\text{B}$ of seawater (Anagnostou et al., 2016). (B) Marine benthic foraminiferal calcite $\delta^{18}\text{O}$ record from ODP Site 1218 (Coxall et al., 2005). (C) Terrestrial temperature change from pre-EOT mean based on our δD record from Patagonia, accounting for stepwise changes in ice volume (Bohaty et al., 2012). Error envelope reflects analytical and $\delta D-T$ relationship uncertainty. (D) Tropical SST record based on the Mg/Ca (Lear et al., 2008) and TEX_{86} (Pearson et al., 2007) of well-preserved planktonic foraminifera from Tanzania. (E) Mid-latitude mean annual air temperature record based on carbon-47 clumped isotopes of freshwater gastropods from Hampshire Basin, UK (Hren et al., 2013).

smaller than temperature shifts observed at sites with substantial continentality (Fan et al., 2018; Zanazzi et al., 2007), which underscores the coherent response of the global climate system during the EOT.

The terrestrial cooling we observe is followed by a partial recovery. This rebound is truncated in the Vera Member by erosion or cessation of deposition, possibly in response to the newly established icehouse climate state. A similar return to warmer temperatures is also observed in foraminiferal Mg/Ca temperatures (Lear et al., 2008). The parallel nature of this terrestrial record with marine records suggests that the Gran Barranca record is an expression of a global event rather than a localized phenomenon. The simplest reason for this brief warming is the rebound in atmospheric $p\text{CO}_2$ (Pearson et al., 2009) (Fig. 5A), which is consistent with a transient pulse of warming prior to the establishment of cooler icehouse conditions and a stable ice sheet on Antarctica (Galeotti et al., 2016). This rebound in temperature explains why previous climate data from Gran Barranca (Kohn et al., 2004, 2015, 2010) had suggested negligible change across the EOT: we observe that the levels at which tooth enamel and soil carbonate nodules are available is located in a largely invariant part of our record that post-dates the recovery (Fig. 2).

The reconstructed temperature change during the EOT in Patagonia does not parallel metrics of regional biotic change—phytolith assemblages (Strömberg et al., 2013) and grazing mammal fauna (Reguero et al., 2010) change minimally during the EOT. In contrast to the significant extinctions documented in North American lineages of mammals, amphibians, reptiles, and both marine and terrestrial molluscs (Evanoff et al., 1992; Hutchison, 1992; Ivany et al., 2000) at the EOT, Patagonian flora and fauna appear unchanged. This generally indicates that factors other than temperature were important in driving biotic change in Patagonia. The abundance of burnt palm phytoliths varies through the time interval represented by the Vera Member (Selkin et al., 2015; see Fig. 2D), although it does not parallel reconstructed temperature change.

Phytoliths show little or no apparent ecosystem change across the EOT at Gran Barranca showing dominance by palms (*Arecaceae*) (Strömberg et al., 2013) prior to and after the EOT. Despite the persistence of this ecosystem indicator, extant palms tolerate a wide range of temperatures above MATs of 10°C (Greenwood and Wing, 1995), and extend to $\sim 40^\circ\text{S}$ in South America in the present interglacial climate (Eiserhardt et al., 2011; Reichgelt et al., 2018). Reconstructed SST data from nearby DSDP Site 511 show minimum EOT MATs greater than $10\text{--}14^\circ\text{C}$ (Liu et al., 2009); thus palms likely remained viable in Patagonia during and following the EOT. The lack of apparent change in the floral community at the EOT could reflect the cosmopolitan nature of *Arecaceae* rather than a static climate. In short, precipitation stable isotopes and phytoliths record different environmental factors (temperature and vegetation change, respectively), but both records are consistent with changes in central Patagonia's climate during the EOT.

The current age model for the Vera Member (Dunn et al., 2013) indicates that our isotope data record cooling and an environmental response that is generally synchronous with records of the EOT from elsewhere on Earth. Alternative interpretation of paleomagnetic data (see discussion in Dunn et al., 2013) could place observed changes to stable isotope hydrology and cooling up to 1.3 Myr earlier, however a comparison of onshore (Gran Barranca) and South Atlantic offshore (ODP Site 511) data suggests a preference for the current depth-age model. These results demon-

in atmospheric $p\text{CO}_2$ and regional ice buildup on Antarctica during this climate transition.

8. Conclusions

Terrestrial Patagonia cooled at the same time as the ocean at the onset of the EOT. Observed changes in stable isotope hydrology in Patagonia during this interval are consistent with GCM predictions of high latitude and Antarctic cooling and associated change in precipitation δD across the EOT (Feakins et al., 2014). The rebound in Patagonian terrestrial temperature following initial cooling likely coincides with the transient resurgence of atmospheric $p\text{CO}_2$ at the EOGM (Pearson et al., 2009) and the initial expansion of ice on Antarctica (Galeotti et al., 2016). Though mechanisms for the transient rebound in atmospheric $p\text{CO}_2$ are debated (Zachos and Kump, 2005; Zachos et al., 1999) this increase in $p\text{CO}_2$ likely resulted in a near-return to pre-EOT conditions that is observed in both terrestrial and marine records (Lear et al., 2008). The correspondence between temperature change and atmospheric $p\text{CO}_2$ highlights the coupling of the ocean–atmosphere system in southern Patagonia and the role of $p\text{CO}_2$ in driving terrestrial climate change, particularly if changes in ocean circulation were responsible for the initial $p\text{CO}_2$ fall (Abelson and Erez, 2017; Elsworth et al., 2017).

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Appendix A. Supplementary material

Supplementary material related to this article can be found online at <https://doi.org/10.1016/j.epsl.2019.01.052>.

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