1	Magnesium isotope analysis of olivine and pyroxene by SIMS:
2	Evaluation of matrix effects
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28 Abstract

29 The performance of multi-collector secondary ion mass spectrometry (MC-SIMS) for Mg 30 isotope ratio analysis was evaluated using 17 olivine and 5 pyroxene reference materials (RMs). 31 The Mg isotope composition of these RMs was accurately and precisely determined by multi-32 collector inductively coupled plasma mass spectrometry (MC-ICP-MS), and these measured 33 isotope ratios were used to evaluate SIMS instrumental mass bias as a function of the forsterite 34 (Fo) content of olivine. The magnitude of the Mg isotope matrix effects were $\sim 3\%$ in δ_{25} Mg, 35 and are a complex function of olivine Fo content, that ranged from Fo59.3 to F0100. In addition to 36 these Mg isotope matrix effects, Si+ ion yields and Mg+/Si+ ion ratios varied as a complex 37 function of the Fo content of the olivine RMs. For example, Si_{+} ion yields varied by ~33%. 38 Based on the observations, we propose instrumental bias correction procedures for SIMS Mg 39 isotope analysis of olivine using a combination of Mg+/Si+ ratios and Fo content of olivine. 40 Using this correction method, the accuracy of δ_{25} Mg analyses is 0.3‰, except for analysis of 41 olivine with Fo86-88 where instrumental biases and Mg+/Si+ ratios change dramatically with Fo 42 content, making it more difficult to assess the accuracy of Mg isotope ratio measurements by 43 SIMS over this narrow range of Fo content.

44 Five pyroxene RMs (3 orthopyroxenes and 2 clinopyroxenes) show smaller ranges of 45 instrumental bias (~1.4‰ in δ_{25} Mg) as compared to the olivine RMs. The instrumental bias for 46 the 3 orthopyroxene RMs do not define a linear relationship with respect to enstatite (En) 47 content, that ranged from En85.5-96.3. The clinopyroxene RMs have similar En and wollastonite 48 (Wo) contents but have δ_{25} Mg values that differ by 0.5‰ relative to their δ_{25} Mg values 49 determined by MC-ICP-MS. These results indicate that additional factors (e.g., minor element 50 abundances) likely contribute to SIMS instrumental mass fractionation. In order to better correct 51 for these SIMS matrix effects, additional pyroxene RMs with various chemical compositions and 52 known Mg isotope ratios are needed.

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- 55 **1. Introduction**
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57 Magnesium is a major rock-forming element and seventh most abundant element in the 58 Solar System. Magnesium has three naturally occurring stable isotopes (24Mg, 25Mg, and 26Mg) 59 allowing us to study mass-dependent and mass-independent isotope fractionation induced by 60 natural processes. Three analytical approaches are commonly taken to obtain high precision Mg 61 isotope data: solution nebulization multi-collector inductively coupled plasma mass spectrometry 62 (SN-MC-ICP-MS), laser ablation MC-ICP-MS (LA-MC-ICP-MS), and multi-collector secondary ion mass spectrometry (MC-SIMS). Among them, SN-MC-ICP-MS analyses of Mg 63 64 isotopes are the most reliable. This method entails dissolution of the sample followed by 65 quantitative separation of Mg from all other cations by ion exchange chromatography and Mg 66 isotope ratio analysis using a standard bracketing technique (e.g., Galy et al., 2001; Teng et al., 67 2010). SN-MC-ICP-MS has been applied to solve important geochemical and cosmochemical 68 questions, such as mantle metasomatism (e.g., Xiao et al., 2013; Hu et al., 2016a; Teng, 2017 69 and references therein), continental weathering (e.g., Tipper et al., 2006; Pogge von Strandmann 70 et al., 2008), distribution of Mg isotopes in the Earth and early solar system (e.g., Handler et al., 71 2009; Teng et al., 2010; Schiller et al., 2010; Bourdon et al., 2010; Sedaghatpour and Teng, 72 2016; Van Kooten et al., 2016; Olsen et al., 2016), radiogenic 26Mg excesses produced by the 73 decay of short-lived 26Al (e.g., Jacobsen et al., 2008; Larsen et al., 2011; Wasserburg et al., 74 2012), and condensation and evaporation processes in the solar nebula (e.g., Richter et al., 2007; 75 Davis et al., 2015). Meteoritic samples, such as Ca, Al-rich inclusions (CAIs), show micro-scale 76 variations in radiogenic 26Mg and/or the degree of mass-dependent fractionation, and these Mg 77 isotope composition have been determined by both LA-MC-ICP-MS and MC-SIMS at the 5-100 78 µm spatial resolutions (e.g., Young et al., 2002; Kita et al., 2012; MacPherson et al., 2012, 2017; 79 Bullock et al., 2013; Kawasaki et al., 2015, 2017, 2018, 2019; Ushikubo et al., 2017; Mendybaev 80 et al., 2017; Williams et al., 2017).

81 Isotope ratios measured by the above three mass spectrometric techniques are inherently 82 different from the absolute Mg isotope ratios of samples due to instrumental mass fractionation 83 (IMF) (e.g., Eiler et al., 1997; Albarède et al., 2004; Horn and von Blanckenburg, 2007). 84 Typically, IMF is corrected by normalization to a standard run under the same conditions as the 85 sample. A bias in this IMF correction can be produced if the sample and standard are not 86 identical in composition, and this bias is often referred to as a matrix or non-spectral interference 87 (e.g., Albarede and Beard, 2004). For SN-MC-ICP-MS, matrix effects are eliminated by 88 purifying samples by ion exchange chromatography and matching the Mg concentration of 89 samples and standards when doing isotope ratio measurements. In contrast, for in-situ 90 techniques, this purification step is not possible and thus LA-ICP-MS and SIMS analysis can be 91 strongly affected by sample matrix. Chaussidon et al. (2017) argued that matrix correction is 92 typically much larger for SIMS than LA-ICP-MS. Thus, matrix-matched standards are always 93 required to obtain accurate Mg isotope data using SIMS techniques. For LA-ICP-MS matrix 94 effects can be minimized or eliminated by use of lasers with ultra-short pulse widths and by 95 loading the plasma with small amounts of water that is co-aspirated into the instrument with the 96 ablated material (Oeser et al., 2014).

97 In spite of matrix effects that compromise the accuracy of in-situ Mg isotope analyses, 98 the SIMS technique has been successfully used to identify large mass-dependent isotope 99 fractionations (>10%/amu) in Mg isotopes that have been observed in CAIs with Fractionation 100 and Unidentified Nuclear (FUN) effects (e.g., Krot et al., 2014 and reference therein; Park et al., 101 2017 and reference therein). Magnesium isotope ratios of FUN CAIs are positively fractionated 102 up to 45‰/amu (e.g., Park et al., 2017), suggesting that FUN CAIs experienced intense 103 evaporation events under near vacuum conditions (Mendybaev et al., 2013; 2017). Technical 104 improvements on the MC-SIMS Mg isotope analysis over the last decade (e.g., Kita et al., 2012; 105 Luu et al., 2013) allow us to investigate mass-dependent fractionation effects smaller than 106 10‰/amu in CAIs. For example, Mg isotope zoning of melilite in coarse grained CAIs was 107 successfully determined at sub‰ level by MC-SIMS by correcting the SIMS matrix effects of

δ25Mg in melilite solid-solution that changes linearly with Åkermanite contents (Kita et al.,
2012; Bullock et al., 2013).

110 Ushikubo et al. (2013) conducted high precision SIMS Mg isotope analysis of multiple 111 phases in chondrules and calibrated instrumental bias on olivine based on a linear correlation 112 between instrumental bias and Fo content (Fo = [Mg]/[Mg + Fe] molar %) observed in 4 olivine 113 standards (Fo59 to Fo100). Olivine in chondrules tends to show positively fractionated Mg isotope 114 ratios (up to 2.3‰/amu) relative to coexisting phases (pyroxene and plagioclase). Based on these 115 observations, Ushikubo et al. (2013) suggested that chondrule melting occurred in an open 116 system. Recently, however, Chaussidon et al. (2017) evaluated matrix effects on Mg isotope 117 analyses of olivine and silicate glasses by SIMS and found complex IMF in olivine as a function 118 of Fo content. Consequently, Mg isotope ratios of olivine in chondrules reported in Ushikubo et 119 al. (2013) might not be accurate due to unrecognized matrix effects. Furthermore, Ushikubo et al. 120 (2013) assumed that the Mg isotope composition of these four olivine standards (including one 121 synthetic olivine) is the same as the Earth's mantle (-0.13‰ relative to the DSM-3 standard; 122 Teng et al., 2010). This assumption, at least for the synthetic olivine standard, may not be valid 123 because Mg isotope compositions of starting materials do not need to be the same as the Earth's 124 mantle and synthetic processes may induce mass-dependent fractionation (e.g., Kita et al., 2012). 125 In order to evaluate mass-dependent fractionation effects in olivine and pyroxene from 126 extraterrestrial materials, it is important to develop suitable reference materials so that the matrix 127 effects associated with SIMS analysis can be fully evaluated.

In this paper, we evaluate matrix effects on SIMS Mg isotope analysis of olivine and pyroxene. We prepared 17 olivine and 5 pyroxene reference materials (RMs) with various chemical compositions. The Mg isotope ratios of individual RMs were determined relative to the DSM-3 scale by either SN-MC-ICP-MS or by LA-MC-ICP-MS techniques, depending on the quantity of available RM and the degree to which high purity mineral separates could be produced. RMs with limited amounts of material, or which were difficult to separate, were analyzed by LA-MC-ICP-MS. These RMs were then subjected to MC-SIMS three Mg isotope

135 ratio measurements. Moreover, major- and minor-element analyses of these olivine RMs were 136 also performed by SIMS and electron microprobe in order to evaluate the relationship between 137 instrumental biases and ionization efficiencies of each element as compared to the chemical 138 composition of the olivine RMs that were determined by electron microprobe analysis. Our goal 139 is to evaluate the applicability of high precision and high spatial resolution SIMS Mg isotope 140 analysis. This work is critical to allow one to assess the accuracy of Mg isotope ratio analysis of 141 extraterrestrial olivine and pyroxene samples, such as those in chondrule and amoeboid olivine 142 aggregate from primitive meteorites, as well as small and precious particles obtained by sample 143 return missions (e.g., Stardust, Hayabusa2, and OSIRIS-REx).

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146 **2. Sample preparation**

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148 Seventeen olivine (11 terrestrial, 5 meteoritic, and one synthetic) and 5 terrestrial 149 pyroxene RMs were used in this study. Four olivine (SC-Ol, HN-Ol, IG-Ol, and OR-Ol) and all 150 pyroxene RMs in the suite of standards have been used as SIMS calibration standards for oxygen 151 isotope analyses (Kita et al. 2010). Mg isotope ratios of 3 olivine RMs (CL09-08, 19, and 33-Ol) 152 were previously reported in the literature (Xiao et al., 2013). Importantly, the RMs considered in 153 this study also include 5 meteoritic olivine samples. Meteoritic olivine has distinct chemical 154 compositions (in terms of both Fo contents and minor-element abundances), as compared to 155 terrestrial olivine samples. A list of the RMs is reported in Table 1. Eleven olivine and all 156 pyroxene RMs were either taken from mineral separates or handpicked from sieved grain-size 157 fractions that had a range of grain sizes from 100 µm-1000 µm. Aliquots of 5 olivine RMs (KN-158 Ol, CL09-19-Ol, CL09-08-Ol, CL09-33-Ol, and WN-Ol) were separated from host rocks and 159 meteorites in this study. Olivine grains in Winona meteorite (WN-Ol) were small (~100 µm) and 160 intermixed with metals, so several ~1 mm sized chips of Winona meteorites were mounted in 161 epoxy resin and polished; chips containing olivine-rich areas were then extracted from the epoxy

162 resin and used for this study. The RM, N7325-Ol, is from olivine grains from an ungrouped 163 achondrite (NWA 7325), that has Mg-rich olivine (F097.4). This RM was obtained from a 164 polished thick section studied by Goodrich et al. (2017). Detail descriptions of each RM, and 165 procedures used for separating olivine grains are summarized in Appendix EA1. Except the WN-166 Ol and N7325-Ol, 1 to 42 grains (typically 20 grains) of each RM were handpicked under a 167 binocular microscope and mounted in 25 mm diameter epoxy disks for electron microprobe, LA-168 MC-ICP-MS, and SIMS analysis. Among them, limited numbers of grains (one or two grains) 169 were handpicked for SC-Ol and HN-Ol because they have larger grain sizes (~1 mm) and are 170 known to be homogeneous in their Mg isotope ratios (e.g., Ushikubo et al., 2013). In addition, 171 only two grains of A77257-Ol were handpicked and used for this study because we only have 3 172 olivine grains of this RM. All grains and Winona chips were placed within a radius of 7 mm 173 from the geometrical center of the mount to minimize instrumental mass fractionation effects due 174 to sample geometry and topography (Kita et al., 2009; Peres et al., 2013). San Carlos olivine 175 (hereafter, SC-OI) grains were placed near the center of each mount and these grains were used 176 as a running standard during LA-MC-ICP-MS and SIMS analysis (the NWA 7325 thick section 177 was originally mounted with a grain of SC-Ol). All grain mounts of olivine and pyroxene RMs and the NWA 7325 thick section were coated with carbon (20 nm thickness) for electron 178 179 microprobe and SIMS analyses. 180 181

182 **3. Experimental methods**

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184 **3.1. Electron microprobe analysis**

Olivine RMs were measured by electron probe micro analyzer (EPMA) in two sessions.
The first session (August 2018) considered the three major and two minor element oxide (MgO,

187 SiO₂, CaO, MnO, and FeO) concentrations of 17 olivine RMs using a Cameca SXFive FE

188 electron microprobe at the University of Wisconsin-Madison. Analyses were performed with an

accelerating voltage of 15 kV and a beam current of 20 nA, with a 3 micron diameter beam.
Counting time for the peak was 10 seconds. Backgrounds were determined using the mean
atomic number procedure (Donovan and Tingle, 1996). The following standards were used for
olivine analyses: synthetic forsterite (Mg, Si), NMNH 122142 Kakanui augite (Ca), synthetic
FeO (Fe), and synthetic Mn₂SiO₄ (Mn). The probe for EPMA_{TM} (PFE) software was used for
data reduction. Calculated detection limits (99% confidence) for the measured oxides listed
above were 0.01, 0.02, 0.02, 0.06, and 0.04 wt%, respectively.

A second session (January 2019) for olivines added Cr₂O₃ and NiO to the list of elements, with the same column conditions. Here off peak backgrounds were acquired, with 10 second counts on the background and 10 seconds on the peaks. The same standards were used as above, with the addition of synthetic Cr₂O₃ for Cr and synthetic Ni₂SiO₄ for Ni. Detection limits (wt%) were: MgO-0.02, SiO₂-0.03, CaO-0.02, Cr₂O₃-0.07, MnO-0.07, FeO-0.06 and NiO-0.08.

201 Major and minor element oxide (Na2O, MgO, Al2O3, SiO2, CaO, TiO2, Cr2O3, MnO, and 202 FeO) concentrations of 5 pyroxene RMs were obtained with the SXFive FE electron microprobe 203 under the same column condition for olivine analyses. Off peak backgrounds were acquired, with 204 10 seconds background and 10 seconds peak counting time. The following standards were used 205 for analyses: Burma jadeite (Na), NMNH 122142 Kakanui augite (Mg, Ca), Grass Valley 206 anorthite (Al), synthetic enstatite (Mg, Si), synthetic TiO₂ (Ti), synthetic Cr₂O₃ (Cr), synthetic 207 Mn₂SiO₄ (Mn) and synthetic FeO (Fe). Calculated detection limits for the measured oxides in 208 wt% were Na2O-0.02, MgO-0.02, Al2O3-0.03, SiO2-0.03, CaO-0.02, TiO2-0.05, Cr2O3-0.06, 209 MnO-0.05, and FeO-0.06 wt%. In order to evaluate grain-scale homogeneity of each of olivine 210 and pyroxene RMs, 10-20 or 5-8 grains were typically analyzed for each of the olivine or 211 pyroxene RMs, respectively, and individual grains were analyzed by 5 times. 212 In the following, Mg contents in olivine and pyroxene RMs are expressed as a Fo (=

213 $Mg/[Mg + Fe] \mod \%$) and an En (= $Mg/[Mg + Fe + Ca] \mod \%$) contents, respectively.

Likewise, Ca content in pyroxene RMs is expressed as a Wo (= Ca/[Mg + Fe + Ca] molar %)

215 content. The Fo, En, and Wo contents were calculated by considering total iron as ferrous iron.

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217 3.2. Chemical procedures and Mg isotope analysis by SN-MC-ICP-MS Aliquots of 4 olivine RMs (SC-Ol, HN-Ol, IG-Ol, and OR-Ol) and all pyroxene RMs, 218 219 weighing between 4 mg and 19 mg, were used for SN-MC-ICP-MS analysis. 220 All sample processing was performed under clean laboratory conditions in the Isotope 221 Cosmochemistry and Geochronology Laboratory (ICGL) at Arizona State University (ASU). 222 Each aliquot was cleaned by ultrasonication in Milli-Q® H₂O and then crushed using an agate 223 mortar and pestle. These aliquots were dissolved using in-house distilled, high-purity acids. All 224 olivine fractions were treated with 6 N HCl on a hotplate at 120 °C for 48 hours followed by 225 evaporation to dryness and further treated with a 3:1 mixture of HNO₃:HF at 120 °C for 24 226 hours. After evaporating to dryness, a final treatment in 12N HCl was used to ensure complete 227 dissolution was achieved. The pyroxene fractions were digested in a 3:1 mixture of HNO3:HF on 228 a hotplate at 120°C for 48 hours followed by evaporation to dryness; this process was repeated 229 until the sample was fully converted to fluorides. The fluorides were subsequently dissolved 230 using 6N HCl on the hotplate at 120 °C for 48 hr. Following complete dissolution, a 1% fraction 231 was taken from each aliquot for bulk chemical analysis on the iCAP-Q quadrupole ICP-MS at 232 ASU.

A ~10 μ g Mg equivalent fraction of each of the dissolved RM aliquots was loaded onto a quartz column packed with AG® 50W-X8 200-400 mesh cation resin. The Mg was purified using procedures similar to those described previously in Spivak-Birndorf et al. (2009). Following this chemical separation procedure, the Mg yields were verified using the iCAP-Q ICP-MS and were consistently >99%. The purified Mg cuts were dried and then brought into 250 ppb Mg solutions in 3% HNO3 for analysis of Mg isotopes via SN-MC-ICP-MS.

Magnesium isotope analyses were performed on the ThermoFinnigan Neptune MC-ICP-MS in medium-resolution mode following procedures described in Spivak-Birndorf et al. (2009) and Bouvier et al. (2011). Instrumental mass fractionation was corrected by sample-standard bracketing using the DSM-3 Mg standard (Galy et al., 2003), and the mass-dependent Mg

243 isotope composition of each RM is reported as $\delta_{25,26}$ Mg values relative to this standard. To verify 244 the accuracy and precision of our measured Mg isotope compositions, the USGS basaltic rock 245 standard BCR-2 was processed through the entire chemical procedure alongside the olivine and 246 pyroxene RMs and analyzed using the same SN-MC-ICP-MS protocols as these RMs during two 247 analytical session. The reproducibility of our Mg isotope measurements was evaluated using 248 repeat analyses of the DSM-3 Mg standard and the BCR-2 rock standard over the course of these 249 two analytical sessions and was $\pm 0.06\%$ (2SD) and $\pm 0.11\%$ (2SD) for δ_{25} Mg and δ_{26} Mg, 250 respectively.

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252 **3.3.** Mg isotope analysis by LA-MC-ICP-MS

253 Magnesium three-isotope analyses of 14 olivine RMs were made using a Nu plasma II 254 MC-ICP-MS with a Nd:YLF-pumped Ti:sapphire femtosecond laser ablation (fs-LA) system at 255 the University of Wisconsin-Madison. Typical operating conditions of fs-LA and MC-ICP-MS 256 are summarized in Table 2. The laser spot size at the sample surface was $\sim 13 \mu m$ in diameter as 257 determined by the measured width of a line scan conducted using the laser and raster conditions 258 reported in Table 2 on a grain of SC-Ol. For this laser ablation system, the laser spot size is 259 controlled by adjusting the distance between sample and laser objective lens and by an adjustable 260 iris placed before the objective lens. For each analysis, a raster ablation was used that ranged in 261 size from $\sim 1200-6000 \,\mu\text{m}^2$ with various scan speeds ($\sim 2-10 \,\mu\text{m/sec}$) and laser repetition rates 262 such that the dosage of laser shots delivered to a spot was similar using different conditions. 263 These changes in laser repetition rate and raster settings were done so that the ion intensity of the 264 RM's matched the ion intensity of the SC-Ol used as a bracketing standard. The ICP-MS was 265 operated in low-mass-resolution mode using a 0.3 mm defining slit. Magnesium isotope ratios 266 were measured using a standard-sample-bracketing method; SC-Ol grains (Foss.8) were used as 267 the bracketing standard. A 60-second on-peak gas blank (laser not firing) was analyzed before 268 each analysis, and this gas blank was subtracted from the signal analyzed for 60 seconds for both 269 unknown samples and the bracketing standard. In general, fs-LA-MC-ICP-MS analysis have

270 been adopted to overcome matrix effects compared with nanosecond laser ablation MC-ICP-MS 271 analysis (e.g., Horn and von Blanckenburg 2007; Steinhoefel et al., 2009; Oeser et al., 2014; 272 Zheng et al., 2018). In this study, to further minimize possible matrix effects on fs-LA-MC-ICP-273 MS, Mg isotope analyses were performed under wet plasma conditions by co-aspirating ultra-274 pure water along with the ablated material into the plasma (Oeser et al., 2014). Two analysis 275 sessions were conducted. During one session, ultra-pure water was introduced at a rate of 12 276 μ L/min to a peltier cooled (7 °C) cyclonic spray chamber (condition LA1) and in the second, 277 water was aspirated at a rate of 40 μ L/min (condition LA2). Typical 24Mg+ intensities of SC-Ol 278 under the condition LA1 and LA2 were 4.8 V and 2.9 V, respectively. The differences in ion 279 intensities reflects the fact that instrument sensitivity is reduced as more water is added to the 280 system (Zheng et al., 2018). The δ_{25} Mg external reproducibilities of the bracketing standard (SC-281 Ol) were $\pm 0.08\%$ (2SD) and $\pm 0.17\%$ (2SD) over the course of 12 hours for condition LA1 and 282 LA2, respectively. Mg isotope ratios of each RM were determined as $\delta_{25,26}$ Mg values relative to 283 the SC-Ol bracket standard and these $\delta_{25,26}$ Mg values were normalized to DSM-3 scale by using 284 δ25,26MgDSM-3 values of SC-Ol, as determined using SN-MC-ICP-MS analysis at IGCL, ASU.

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3.4. Mg isotope analysis by MC-SIMS

287 Magnesium three-isotope analyses of olivine and pyroxene were performed in 4 sessions 288 utilizing the WiscSIMS Cameca IMS 1280 at the University of Wisconsin-Madison equipped 289 with a radio-frequency (RF) plasma ion source. We used two primary ion species (O- and O₂-), 290 which were accelerated by 23 kV (-13 kV at the ion source and +10 kV at the sample surface) 291 and were focused to ~9 or ~7.5 µm in diameter at 2.6 nA or 1 nA, respectively. Beam sizes were 292 determined by SEM observations after SIMS analyses (Fig. 1). Secondary ion optics were 293 adjusted to $200 \times$ magnification from the sample surface to the field aperture (4000 µm square) 294 with a 50 eV energy window. Secondary Mg ions (24Mg+, 25Mg+, 26Mg+) were detected on multi-295 collector Faraday Cups (FCs) using one 1010 ohm and two 1011 ohm resistors for 24Mg+, and 25, 296 $_{26}Mg_{+}$, respectively. The mass resolving power at 10% peak height was set to ~ 2500 (entrance

297	slit; 90 μ m and exit slit 500 μ m) and contributions of 48Ca2+ and 24Mg1H+ to 24Mg and 25Mg+		
298	peaks were negligibly small. A typical mass spectrum is provided in Appendix EA2. The		
299	secondary ion optics are similar to that described in Kita et al. (2012) and Ushikubo et al. (2013)		
300	SC-Ol (Foss.s) grains were used as a bracketing standard for both olivine and pyroxene		
301	measurements. Typical 24Mg+ count rates of SC-Ol by using O- and O2- primary beams were		
302	\sim 2.4 \times 108 and \sim 2.3 \times 108 cps, respectively. Note that secondary ionization efficiency of Mg with		
303	O2- primary beam is higher than that with O- primary beam (Kita et al., 2000). A single analysis		
304	takes 8 min, including 100 s of presputtering, ~80 s for automated centering of secondary beam		
305	(DTFA-X, DTFA-Y), and 300 s of integration (10 s \times 30 cycle) of the secondary ion signals. The		
306	baseline of the FC detectors was monitored during each presputtering and averaged over eight		
307	analyses. External reproducibilities (2SD) of $\delta_{25}Mg_m$ and $\delta_{26}Mg_m$ were 0.10‰ and 0.17‰ for O-		
308	analysis and 0.09‰ and 0.16‰ for O ₂₋ analysis, respectively.		
309	We follow a data reduction scheme described in Ushikubo et al. (2017). Mass-dependent		
310	instrumental bias (<i>f</i> *25 in the unit of ‰) of a RM is expressed as:		
311			
312	$f^{*}_{25 (RM)} = \left[(1 + \delta_{25}Mg_{m}/1000) / (1 + \delta_{25}Mg_{DSM-3}/1000) - 1 \right] \times 1000 $ (1)		
313			
314	$\delta_{25}Mg_m$ represents a raw-measured, background-corrected $\delta_{25}Mg$ value of a RM measured by		
315	SIMS, which is expressed in δ -notation by normalizing to the absolute Mg isotope ratio		
316	$(25Mg/24Mg = 0.12663$, Catanzaro et al., 1966). $\delta_{25}Mg_{DSM-3}$ represents a $\delta_{25}Mg$ value relative to		
317	DSM-3 that is determined by SN- or fs-LA-ICP-MS analyses. In order to correct instrumental		
318	drift during a SIMS session, the instrumental bias (f^{*25}) of each RM was normalized to the bias		
319	of the running standard (SC-Ol). The relative bias* for a RM is defined as:		
320			
321	bias*25 (RM-SCOI) = $[(1 + f^{*}_{25} (RM)/1000)/(1 + f^{*}_{25} (SCOI)/1000) - 1] \times 1000$ (2)		
322			

323 The f^{*25} (scol) represents an average f^{*25} value that is calculated from eight bracket 324 analyses of the running standard (SC-OI) for each bracket. In general, instrumental mass bias is 325 better-evaluated using δ_{26} Mg values because the mass difference between ${}_{24}$ Mg and ${}_{26}$ Mg 326 isotopes is larger as compared to the difference between 24Mg and 25Mg isotopes. However, the 327 fs-LA-MC-ICP-MS analyses were done at a mass resolving power of 400, which is insufficient 328 to resolve $26Mg_{+}$ from possible $12C14N_{+}$ and $52Cr_{2+}$ isobars, making it challenging to assess the 329 accuracy of δ_{26} MgdsM-3 values determined by fs-LA. Thus we prefer to use δ_{25} Mg values to 330 evaluate instrumental mass bias. Moreover, meteoritic olivine RMs may have excess radiogenic 331 $_{26}$ Mg, making δ_{25} Mg values the best choice for evaluating IMF.

In addition to the above analysis conditions, a limited numbers of olivine RMs were analyzed for 300 cycles for each spot (~50 min) in order to examine the drift of $f^{*_{25}}$ with the depth of the analysis. To obtain complete depth profile from the surface, only a short presputtering time (10s) was applied and no secondary deflector adjustments (DTFA-X, DTFA-Y) were performed for these analyses. Secondary deflector adjustments for these analyses were performed on a spot adjacent to and just prior to the depth profile analysis.

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339 **3.5.** Major and minor element analyses by SIMS

340 Major and minor element analyses (Na, Mg, Al, Si, Ca, Cr, Mn, Fe, and Ni) of 17 olivine 341 RMs were performed with the WiscSIMS Cameca IMS 1280. As with the case for Mg isotope 342 analysis, two different primary ion species (O_{-} and O_{2-}) were used, which were focused to ~ 2 or 343 \sim 1.5 µm diameter at 16 pA or 6 pA, respectively. Secondary ion optics was operated under the 344 same conditions for Mg isotope analysis described in 3.4, except for the mass resolving power 345 that was set to ~3000. A secondary 24Mg+ ion was detected by an axial FC (FC2) with a 1011 346 ohm resistor and the other secondary ions (23Na+, 27Al+, 28Si+, 40Ca+, 52Cr+, 55Mn+, 56Fe+, and 347 60Ni+) and the mass 22.7 for stabilizing magnetic field for 23Na+ were detected by an axial 348 electron multiplier (EM) with magnetic peak jumping mode. SC-Ol grains were used as a 349 running standard. Typical $_{24}Mg_{+}$ count rates of SC-Ol for O₋ and O₂₋ analyses were $\sim 8.5 \times 10^{5}$

350	and ~ 8.6×105 cps, respectively. Per cycle, the count duration for 24Mg+, 28Si+, and 56Fe+ ions		
351	were 2 seconds, and for the mass 22.7 and 23Na+, 27Al+, 40Ca+, 52Cr+, 55Mn+, and 60Ni+ ions were		
352	1 second. The waiting duration for 23Na+, 27Al+, 28Si+, 55Mn+, 56Fe+, and 60Ni+ ions were 1.6		
353	seconds, for 24Mg+ and 52Cr+ ions were 2.4 seconds, for the 40Ca+ ion was 3.0 second, and for the		
354	mass 22.7 was 2.0 second. A single analysis takes 6 min, including 100 s of presputtering, ~80 s		
355	for automated centering of secondary beam (DTFA-X, DTFA-Y), and 160 s of integration (32 s		
356	\times 5 cycle) of the secondary ion signals.		
357	For SIMS analyses, an ion yield of Mg (or Si) is defined as count rate (per second) of Mg		
358	(or Si) divided by primary ion beam intensity (nA). A relative sensitivity factor (RSF) of the		
359	24Mg+/28Si+ ratio (hereafter referred as Mg/Si RSF) is expressed as		
360			
361	Mg/Si RSF = (24Mg+/28Si+)/(Mg cpfu/Si cpfu) (3)		
362			
363	where $(24Mg+/28Si+)$ is a ratio of count rates of secondary ions $24Mg+$ and $28Si+$, and Mg (or		
364	Si) cpfu represents a number of cations of Mg (or Si) per oxygen in the oxide formula,		
365	which is determined by EPMA analyses.		
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368	4. Results		
369			
370	Definitions of measured parameters used in the following are summarized in Table 3.		
371			
372	4.1. EPMA analyses		
373	Representative chemical compositions of olivine and pyroxene RMs are summarized in		
374	Table 4. Complete electron microprobe data of olivine and pyroxene RMs are given in Appendix		
375	EA3. Electron microprobe data were collected during three analytical sessions. The first session		
376	(August 2018) was carried out to evaluate the Fo homogeneity of 8 olivine RMs (Table EA3-1).		

377 The second session (January 2019) was carried out to evaluate Fo homogeneity of additional 9 378 olivine RMs (Table EA3-1) and to determine minor element abundances including Cr and Ni in 379 all of olivine RMs (Table EA3-2). The third session (September 2019) was carried out to 380 determine the major and minor element compositions and homogeneity of all pyroxene RMs 381 (Table EA3-1). Homogeneities of Fo and En contents are summarized in Appendix EA4. The Fo 382 contents of 17 olivine RMs and the En contents of 5 pyroxene RMs range from 59.3 to 100 and 383 48.6 to 96.3, respectively. Variations in Fo or En contents for each olivine or pyroxene RM are 384 within ± 0.7 unit (1SD) and most of them are within ± 0.3 unit (1SD).

385

386 4.2. SN- and LA-MC-ICP-MS analyses

387 The Mg isotope ratios of the USGS basaltic rock standard (BCR-2), 4 olivine (HN-Ol,

IG-Ol, SC-Ol, and OR-Ol), and 5 pyroxene (Sp79-11, IG-Opx, IG-Cpx, JE En, and 95AK-6)

389 RMs analyzed by SN-MC-ICP-MS are reported as $\delta_{25,26}$ Mg_{DSM-3} values (Table 5). The

390 δ25,26MgDSM-3 values of 3 olivine RMs (CL09-08-Ol, CL09-19-Ol, and CL09-33-Ol) which have

been determined by Xiao et al. (2013) are also shown in Table 5. The BCR-2 standard resulted in

 $\delta_{25}Mg_{DSM-3} = -0.05 \pm 0.08\%$ (2SD), which is in agreement with the recommended value

393 ($\delta_{25}Mg_{DSM-3} = -0.12 \pm 0.02\%$, 2SD; Teng, 2017) within uncertainties. Four olivine and 5

394 pyroxene RMs have δ_{25} Mg_{DSM-3} values ranging from $-0.37 \pm 0.09\%$ to $-0.01 \pm 0.06\%$ and -0.74

 $\pm 0.05\%$ to $-0.04 \pm 0.08\%$ (2SD), respectively. The δ_{25} Mgdsm-3 value of the SC-Ol is

determined to be $-0.07 \pm 0.09\%$ (2SD), which is also in agreement with the average value of San

397 Carlos olivine (δ_{25} Mg_{DSM-3} = -0.13 ± 0.03 ‰, 2SD) reported in Hu et al. (2016b).

We analyzed 14 olivine RMs by fs-LA-MC-ICP-MS under the two different analytical

399 conditions (LA1 and LA2, see the section 3.3). Under the condition LA1, δ_{25} Mg_{DSM-3} values of

400 OR-Ol (Fos9.3) and IG-Ol (Fos9.6) are determined to be $-0.06 \pm 0.08\%$ (2SD) and $-0.04 \pm 0.05\%$

401 (2SD), respectively, which are consistent with those obtained by SN-MC-ICP-MS ($-0.01 \pm$

- 402 0.06‰ (2SD) and -0.03 ± 0.09 ‰ (2SD), respectively). δ_{25} Mg_{DSM-3} values of SW-Ol (Fo_{81.8})
- 403 from the Springwater meteorite and N7325-Ol (Fo97.4) from the NWA 7325 meteorite are

404 determined to be $-0.09 \pm 0.06\%$ and $-0.32 \pm 0.07\%$ (2SD), respectively. The δ_{25} Mg_{DSM-3} value 405 of SW-Ol is also agreement with a literature value ($-0.11 \pm 0.03\%$, 2SE; Handler et al., 2009). 406 However, the δ_{25} Mg_{DSM-3} value of N7325-Ol with high Fo97.4 is ~0.16‰ lower than a literature 407 value ($-0.16 \pm 0.02\%$, 2SE; Koefoed et al., 2016). Interestingly, the δ_{25} Mgdsm-3 value of WN-Ol 408 with high F095.5 from the Winona meteorite, is $-0.29 \pm 0.07\%$ (2SD), which is also $\sim 0.14\%$ 409 lower than that of a bulk Winona meteorite (-0.15 ± 0.04 %, 2SD; Sedaghatpour and Teng, 410 2016). Although it is possible that Mg isotope ratios of WN-Ol and the bulk Winona meteorite 411 are different, the low δ_{25} MgdsM-3 values observed in high Fo samples relative to the literature 412 values suggest that small extent of matrix effects (~0.16‰/amu) exist on fs-LA-MC-ICP-MS 413 analyses of olivine with high Mg contents under the condition LA1. 414 In order to overcome the problem, we introduced more water to the ICP (40 μ L/min, 415 condition LA2) compare to the condition LA1 (12 μ L/min). The Mg isotope ratios of 13 olivine 416 RMs analyzed under the condition LA2 are reported as $\delta_{25,26}$ MgDsM-3 values in Table 5. Thirteen 417 olivine RMs have δ_{25} Mgdsm-3 values ranging from $-0.89 \pm 0.20\%$ to $0.04 \pm 0.17\%$ (2SD). For 418 comparison, $\delta_{25,26}$ MgdsM-3 values of olivine grains from the Springwater and NWA 7325 419 meteorites, which have been determined by Handler et al. (2009) and Koefoed et al. (2016), 420 respectively, are also shown in Table 5. Under the condition LA2, high Fo content samples (WN-421 Ol and N7325-Ol) have greater δ_{25} Mgdsm-3 values (-0.04 ± 0.08‰ (2SD) and -0.07 ± 0.14‰ 422 (2SD), respectively) relative to those obtained using LA1 conditions. The δ_{25} Mg_{DSM-3} values for 423 these forsteritic olivines measured using LA2 conditions agree with the literature values of the 424 bulk Winona meteorite (Sedaghatpour and Teng, 2016) and olivine from NWA 7325 meteorite 425 (Koefoed et al., 2016), respectively. Likewise, δ_{25} Mg_{DSM-3} values determined using LA2 426 conditions for OR-OI (Fo59.3), IG-OI (Fo89.5), and HN-OI (Fo100) are consistent with those 427 obtained by SN-MC-ICP-MS within 2SD uncertainties (see Table. 5). The lower precision 428 obtained for LA2 is likely driven by the decrease in sensitivity associated with the large amount 429 of water co-aspirated with the ablated material as compared to LA1 conditions. The consistency 430 between fs-LA- and SN-MC-ICP-MS results verifies no significant matrix effects on fs-LA-MC-

- 431 ICP-MS measurements within analytical uncertainties ($\leq 0.20\%$ in δ_{25} Mg, 2SD) under LA2 432 conditions.
- 433

434 **4.3. SIMS analyses**

435

436 **4.3.1. Instrumental bias of olivine RMs**

437 Multiple sessions of SIMS Mg isotope analysis were conducted (Appendix EA5). Olivine 438 RMs were evaluated for Mg isotope heterogeneities by conducting multiple grain analyses 439 (Appendix EA4). The measured δ_{25} Mgm values of each olivine RM had limited variability (\leq 440 0.23‰, 2SD). The *f**₂₅ and bias*₂₅ (RM-SCOI) values of all RMs are shown in Table 6 and complete 441 SIMS Mg isotope data for each session are provided in Appendix EA6.

442 Correction for IMF of isotope ratio analysis by SIMS is typically done by applying an 443 empirical correction. This empirical correction is established by determining a calibration line or 444 curve of IMF values for standards that have a range of major element compositions (e.g., Eiler et 445 al., 1997; Valley and Kita, 2009; Kita et al., 2012; Śliwiński et al., 2016a; 2016b; 2018; Isa et al., 446 2017; Chaussidon et al., 2017; Scicchitano et al., 2018). The IMF for the 17 olivine RMs that 447 range from Fo59 to Fo100 defined a $\sim 3\%$ range of $f^{*}25$ values. The $f^{*}25$ value for olivine analyses 448 using a O- primary beam ranged from -3.5 to -0.3%, and from -3.5 to -0.7% using an O₂₋ 449 primary beam (Table 6). The bias*25 (RM-SCOI) values of 17 olivine RMs defines a complex 450 function versus Fo content. This function is slightly different if an O- or O₂- primary ion beam is 451 used (Fig. 2a and 2b, respectively). In general, bias*25 (RM-SCOI) values are maximum at around 452 Foso and are lowest for near-pure forsterite RMs (SK-Ol and HN-Ol). The change in IMF for 453 olivine from Fo59 to Fo80 is approximately 1 %. There is a 3 % change in IMF from Fo80 to 454 F0100, but this change is not a smooth function of Fo content. In the case of O- analysis, A77257-455 Ol (F085.7) shows the bias*25 (RM-SCOI) value ~1.5‰ higher those of RMs with similar Fo contents 456 (Fig. 2a). The same RM shows a slightly enhanced bias*25 (RM-SCOI) value if an O2- primary beam 457 was used (Fig. 2b). In both primary beam conditions, the bias*25 (RM-SCOI) values zigzag between

Fos9 and Fo97, and decrease by ~1‰ at Fo100. We conducted four sessions of SIMS Mg isotope
analyses (FI1 and FI3 with O2- and FI2 and FI4 with O-) and confirmed that the complex
relationship between IMFs and Fo contents was reproduced. The comparisons of the
relationships between IMF and Fo content obtained from each session are shown in Appendix
EA6. Overall, the complex instrumental biases against Fo content observed in 17 olivine RMs
suggest that the instrumental bias on olivine is not entirely controlled by Mg and Fe contents.

464 Eleven olivine RMs were analyzed for a longer analysis time (total 300 cycles, ~50 min) 465 to determine the changes of IMF as a function of the depth of sputtering. The results are shown 466 in Figs. 3a-d as f^{*}_{25} values versus analysis cycle numbers. During the first 100 cycles (1000 sec) 467 there are large changes in f^{*25} values, especially for O- analysis (Figs. 3a-b). For the rest of the 468 cycles, f^{*}_{25} values is nearly constant or monotonically decreasing. Figure 4a and 4b show the 469 average bias*25 (RM-SCOI) values for each depth (per 50 cycles) as a function of Fo content. In both 470 O- and O₂- analyses, the bias*₂₅ (RM-scol) values for first ~100 cycles are not a smooth function of 471 Fo content and those for after 100 cycles tend to be smoother than those for first ~100 cycles.

472

473 4.3.2. Ion yields of Mg and Si among olivine RMs

474 Results of SIMS major and minor element analyses of olivine RMs are listed in Appendix 475 EA7. We examined the variations in ion yields of Mg and Si among 17 olivine RMs (Figs. 5a-d). 476 Ion yields of Mg are not positively correlated with Fo content and show a complex behavior as a 477 function of Fo content (Figs. 5a-b). Ion yields of Si vary by ~33% and ~26% for O- and O2-478 analyses, respectively (Figs. 5c-d). In the case of O- and O₂- analysis, Si ion yield of olivine RMs 479 with F059.3 to F078.0 decreases with F0 content, whereas Si ion yield increases in RMs from F078.0 480 to F091.9, except for A77257-Ol (F085.7). The Si ion yield for RMs with Fo contents from F091.9 to 481 Fo100 defines a complex behavior. For analyses done using a O- primary beam the Si ion yield 482 largely decreases but for analysis done with a O₂- primary beam the Si ion yield defines a "zig-483 zag" pattern showing an overall decrease in Si ion yield (Figs. 5c-d). We do not find a complex 484 behavior for Fe+ yields against Fo contents (see Appendix EA7).

485 The ion yields of Mg+ and Si+ may also change depending on the sputter-rate of each RM 486 that would differ by Fo contents (e.g., Isa et al., 2017). Thus, we calculate RSFs of secondary ion 487 yields 24Mg+/28Si+ ratios in Figs. 5e-f. The Mg/Si RSF is a complex function of Fo content and it 488 is very similar to the trend in bias*25 (RM-SCOI) as a function of Fo content (Figs. 2a-b). The Mg/Si 489 RSFs are nearly constant for RMs with Fo59.3 to Fo78.0 or from Fo59.3 to Fo85.7 for O- or O2-490 analyses, respectively. The Mg/Si RSFs are negatively correlated with olivine RMs of Fo 78.0 to 491 F091.9 (at the exception of A77257-Ol) or from F085.7 to F091.9 for O- or O2- analyses, 492 respectively. The exact same systematic is observed in bias*25 (RM-SCOI). Moreover, irregularity of 493 Mg/Si RSFs against Fo content follows that of bias*25 (RM-SCOI), such as A77257-Ol (F085.7), WK-494 Ol (F094.3), WN-Ol (F095.5), and N7325-Ol (F097.4) for both O- or O2- analyses. 495

496 **4.3.3. Instrumental bias of pyroxene RMs**

497 Five pyroxene RMs that range from En48.6 to En96.3 were analyzed for their Mg isotope 498 ratios using an O₂- primary beam and the IMF, (f^{*}_{25}) , ranged from 0 to 1.4‰ (Table 6). The 499 range of the pyroxene IMF (1.4‰/amu for En48.6 to En96.3) is two times smaller than the IMF 500 determined for the olivine RMs (3.3%/amu for Fo59.3 to Fo100). However, the absolute magnitude 501 of the IMF for the pyroxene RMs is more positive than the IMF for the olivine RMs (Table 6). 502 The bias*25 (RM-SCOI) values of three orthopyroxene RMs (JE En, IG-Opx, and Sp79-11 En) are 503 not a smooth function of En content (Fig. 6). The two clinopyroxene RMs, IG-Cpx and 95AK-6 504 Di, have identical En content (En48.6) but the bias*25 (RM-SCOI) value of IG-Cpx is ~0.5‰ higher 505 than that bias value of 95AK-6 Di. 506

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508	5. DIS	scussion

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510 5.1. Matrix effects on SIMS Mg isotope analysis of olivine and pyroxene

511

512 5.1.1. Instrumental bias of olivine and calibration against Fo contents

Longer SIMS Mg isotope analyses (total 300 cycles, \sim 50 min) of 11 olivine RMs show significant IMF changes (f^{*}_{25} , up to \sim 2.9‰) as a function of sputtering depth (Figs. 3a-d). The present results suggest that instrumental biases with different depth should be characterized for depth profiling Mg isotope analysis of olivine even if samples have homogeneous chemical compositions. The depth-dependent IMF changes also indicate that the IMF behaves differently depending on different analytical conditions such as primary beam current, beam size, integration time, and cycle numbers.

520 For olivine RMs with F059.3 to F078.0, the instrumental biases on both O- and O2- analyses 521 varies by $\leq 0.3\%$ /amu. Thus, both accurate and precise Mg isotope analysis of olivine with low 522 Fo can be performed. Moreover, Mg/Si RSFs of these low Fo content olivine RMs are near-523 constant (Figs. 5e-f), suggesting no irregular behavior of secondary ion generation of Mg and Si. 524 For olivine with Fo > 78.0, instrumental biases range over $\sim 3\%$ /amu and this variability is a 525 complex function of the olivine Fo content. Fitting the variation of IMF as a function of Fo 526 content would result in a poor fit to these data indicating that the IMF is not solely a function of 527 Fo content.

528 As noted previously, the variation in Mg/Si RSFs as a function of Fo content and the 529 variation in IMF as a function of Fo content are similar, and we consider this to imply that the 530 variations in IMF are in part a result in changes in the ionization efficiency of Mg and Si. The 531 complex relationship between Mg and Si ion yields and Mg and Si contents has been observed in 532 other studies (Steele et al., 1981; Chaussidon et al., 2017, Villeneuve et al., 2019). For example, Villeneuve et al. (2019) found large matrix effects on Si ion yield and instrumental mass 533 534 fractionation in SIMS Si isotope analysis of olivine and low-Ca pyroxene using two different 535 primary ion species (Cs+ and O-). In both cases, relationships between Si ion yield of olivine as a 536 function of Fo content changed at ~ Foso, although the trends for Cs+ and O- analyses were in 537 opposite directions (see Fig. 2 in Villeneuve et al., 2019). The reasons for such complex 538 instrumental bias and ionization efficiency are not well understood (Villeneuve et al., 2019).

539 The observed relationships between ion yields of Mg+ and Si+ and Fo content in this study are

540 broadly consistent with the results for O₋ analyses by Chaussidon et al. (2017) and Villeneuve et

al. (2019). However, previous studies did not include olivine RMs similar to A77257-Ol (Fo85.7)

and Mg-rich samples (Fo > 95) that resulted in a significant irregularities in bias^{*}25 (RM-SCOI) -

- against Fo content as observed in this work.
- 544

545 **5.1.2.** Effect of minor element concentrations in olivine

546 Olivine RMs with higher Cr concentrations ($Cr_2O_3 > 0.2$ wt %) tend to have higher 547 bias*25 (RM-SCOI) values relative to those of olivine RMs without Cr (Figs. 2a-b). Many meteoritic 548 olivine samples, in which Cr occurs as divalent cation (Cr_{2+}) under relatively reducing 549 conditions, tend to have higher Cr contents as compared to terrestrial olivine samples. As shown 550 in Figs. 2a-b, Cr-bearing olivine RMs ($Cr_2O_3 > 0.2$ wt%) tend to show larger bias*25 (RM-SCOI) 551 values relative to those of olivine RMs with similar Fo contents, although this is not the case for 552 all of them. Figure 7 shows the difference between bias*25 (RM-SCOI) values of Cr-bearing olivine 553 RMs compared to mean bias^{*25} (RM-scol) values of other olivine RMs with similar Fo contents 554 (hereafter referred as residual bias). The detail calculation of the residual biases are shown in 555 Appendix EA8. In the case of O- analysis, the residual biases of Cr-bearing olivine RMs with Fo 556 contents \geq 86 are positively correlated with Cr₂O₃ contents (Fig. 7), and those with Fo \geq 94 for 557 O₂- analysis follow a similar correlation (Fig. 7). The other Cr-bearing olivine RMs with lower 558 Fo contents (Fo78 for O- and Fo86 and greater for O2-) do not follow these correlations. Note that 559 both the IMF and Mg/Si RSFs as a function of Fo content (Figs. 5e-f) are nearly constant for 560 olivine with a Fo content below Fo78 for O- and Fo86 for O2-, suggesting that secondary 561 ionization processes are not sensitive to sample matrix below these Fo contents. In contrast, 562 small changes in Cr contents (< 1 wt% as Cr₂O₃) in Mg-rich olivine (Fo > 78 for O₋ and > 86 for 563 O₂–) might drastically change the secondary ionization yields and bias^{*}₂₅ (RM-SCOI) values. These

observations suggest that the secondary ionization process is sensitive to sample matrix in Mg-rich olivine.

566

567 5.1.3. Alternative bias correction scheme for olivine

568 Here, we explore an alternate calibration scheme using the ion yield 24Mg+/28Si+ ratios for 569 the bias^{*}25 (RM-SCOI) values. Figure 8a and 8b show the relationships between bias^{*}25 (RM-SCOI) and 570 24Mg+/28Si+ ratio divided by Fo content. Since olivine should have near-constant Si atomic 571 abundances, differences in the 24Mg+/28Si+/Fo content values represent differences in Mg/Si 572 RSFs among olivine RMs. Although it may be logical to use molar Mg/Si ratio instead of Fo 573 content for the normalization, calibrations using Fo content and molar Mg/Si ratio are very 574 similar to each other. In addition, the Fo content of unknown olivine samples can be easily 575 determined by electron microscopic techniques (e.g., SEM-EDS). Therefore, we use 576 (24Mg+/28Si+/Fo) values instead of Mg/Si RSFs for the sake of convenience. In the case of O-577 analysis (Fig. 8a), the bias*25 (RM-SCOI) values of 15 olivine RMs with various Fo contents ranging 578 from 59.3 to 97.4 are positively correlated with (24Mg+/28Si+/Fo) values and can be fitted with a 579 quadratic function. However, two olivine RMs that are near-pure forsterite (SK-Ol and HN-Ol) 580 do not follow the regression curve from other 15 olivine RMs. In the case of O₂- analysis (Fig. 581 8b), the bias^{*25} (RM-SCOI) values can be fitted into two separate quadratic curves for Fo contents 582 ranging from 59.3 to 88.8 and 88.8 to 100, respectively. These calibration curves are much 583 smoother than those against Fo contents, and this would improve the accuracy of SIMS Mg 584 isotope analyses of unknown olivine samples. Figure 8c and 8d show residual bias^{*25} (RM-SCOI) 585 values calculated from these calibration curves, which are plotted as a function of Fo content. In 586 the case of O- analysis (Fig. 8c), the residual bias^{*25} (RM-SCOI) values are within $\pm 0.4\%$, except 587 for two forsterite (SK-Ol and HN-Ol). In the case of O₂- analysis (Fig. 8d), the residual bias*25 588 (RM-SCOI) values of all olivine RMs are within $\pm 0.2\%$ that is smaller than analytical uncertainties 589 $(\leq \pm 0.3\%, 2\sigma)$. According the calibration scheme, precision and accuracy of bias^{*}25 (RM-SCOI) 590 correction for O₂- analysis would be 0.3‰ for olivine samples with Fo₅₉₋₁₀₀.

591 Note that the current suite of olivine RMs does not cover the range of Fo contents at 592 around 87 and between 98 and 100, where both the bias*25 (RM-SCOI) and (24Mg+/28Si+/Fo) change 593 significantly within a small range of Fo content. Unknown analyses of olivine with these Fo 594 contents may require additional RMs that match the minor element content of the unknown 595 sample. Confirmation of consistent ionization yields of Mg and Si between RM and unknown 596 olivine with similar Fo content could be used to evaluate reliability of SIMS Mg isotope 597 analyses. It should be noted that we combine the IMF and 24Mg+/28Si+ ratios obtained with very 598 different primary ion intensities (e.g., 1 nA and 6 pA for analyses with O₂-), suggesting that these 599 data are acquired from different depths. As mentioned above, the functionality of IMF against Fo 600 content changes significantly with depth (Fig. 4), while relative 24Mg+/28Si+ ratios do not (see 601 Appendix EA7). As a result, the proposed calibration scheme would not work very well for Mg 602 isotope analyses obtained from greater depths. This also means that the complex matrix effects in 603 unknown olivine samples can be properly corrected only if a series of RMs are analyzed under 604 the same analytical conditions, such as primary beam setting, integration time and cycle 605 numbers, differences in which may change the depths of the analyses. The new RF plasma ion 606 source maintains constant primary beam intensity (typically $\pm 2\%$, 1SD) without changes in 607 beam diameters in a week-long analysis session, in contrast to significant changes observed for a 608 Duoplasmatron ion source (>30% in intensity and >50% in beam diameters; e.g., Tenner et al. 609 2019). Thus, the use of stable ion source, like RF plasma ion source, would be critical in 610 obtaining more reliable SIMS Mg isotope analyses.

611

612 **5.1.4. Instrumental bias of pyroxene**

Five pyroxene RMs define a complex variation in IMF as a function of En content for O₂₋ analysis (Fig. 6). Ushikubo et al. (2013, 2017) evaluated IMF on the same set of pyroxene RMs using an O- primary ion beam, and assuming that the $\delta_{25,26}$ MgDSM-3 values of these pyroxenes matched that of pyroxene from the Earth's mantle (-0.13 ± 0.04‰, 2SD; Teng et al., 2010). Here the bias*₂₅ (RM-SCOI) values reported in Ushikubo et al. (2013, 2017) are recalculated by using

 δ_{25} MgDSM-3 values of each pyroxene RM obtained by SN-MC-ICP-MS analyses and these, along with the values determined using a O₂₋ primary beam from this study are shown in Table 6 and Fig. 6. Ushikubo et al. (2013, 2017) also conducted multiple grain analyses of 5 pyroxene RMs so that we used the results as δ_{25} Mg homogeneities of each pyroxene RM to evaluate errors for bias*₂₅ (RM-SCOI) values. The original SIMS raw data obtained by Ushikubo et al. (2013, 2017) are shown in Appendix EA6. Homogeneities on δ_{25} Mgm values of each pyroxene RM are summarized in Appendix EA4 and are within ± 0.23‰ (2SD).

625 For the three orthopyroxene RMs (JE En, IG-Opx, and Sp79-11 En), bias^{*25} (RM-SCOI) 626 values obtained by O- and O2- analyses are similar to each other (Fig. 6). However, the bias*25 627 (RM-SCOI) values of two clinopyroxene RMs (IG-Cpx and 95AK-6 Di) differ by ~2‰ between O-628 and O2- analyses. In both cases of O- and O2- analyses, the bias*25 (RM-SCOI) values of IG-Cpx are 629 ~0.5‰ higher than that of 95AK-6 Di even though these clinopyroxenes have identical En 630 content (En48.6). Note that IG-Cpx contains Cr ($Cr_2O_3 = 0.9 \text{ wt\%}$), but 95AK-6 Di does not 631 (Cr₂O₃ < 0.07 wt%). Moreover, the two clinopyroxene RMs differ in their Al₂O₃ contents (4.5 632 wt% in IG-Cpx and 0.9 wt% in 96AK-6 Di), implying that the instrumental bias on pyroxene 633 may be sensitive to minor element abundances. However, because only 5 pyroxene RMs have 634 been studied, further studies on a range of pyroxenes that differ in their En and Wo contents is 635 needed to fully evaluate matrix effects on SIMS Mg isotope analysis of pyroxene.

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637 5.2. Selection and preparation of olivine and pyroxene reference materials for SIMS Mg 638 isotope analysis

In order to conduct SIMS analyses on olivine and pyroxene that are both accurate and precise it is a best practice to characterize the Mg isotope composition of the RM using either SN-MC-ICP-MS or fs-LA-MC-ICP-MS with water addition. However, for some materials such as olivine and pyroxene from terrestrial mantle-derived peridotites (e.g., mantle xenoliths) or igneous rocks one can assume the material has a mantle-like Mg isotope composition (δ_{25} MgDsM- $3 = -0.13 \pm 0.04\%$, 2SD; Teng et al., 2010). For example, four terrestrial olivine (SC-OI, UWOI-

645 1, IG-Ol, and HaK-Ol) and 2 pyroxene (IG-Opx and IG-Cpx) RMs were obtained from mantle-646 derived peridotites and these RMs have $\delta_{25}Mg_{DSM-3}$ that ranged from -0.14 ± 0.10 to $-0.03 \pm$ 647 0.09‰ (2SD). Moreover, a terrestrial olivine RM from a gabbro (FJ-Ol), a terrestrial olivine RM 648 from a komatiite (WK-Ol), and a terrestrial orthopyroxene (Sp79-11 En) also have mantle-like 649 δ_{25} Mgdsm-3 values ranging from $-0.09 \pm 0.13\%$ to $-0.06 \pm 0.05\%$ (2SD). Meteoritic samples 650 also tend to have a limited range in Mg isotope compositions. Where, for example, chondrites 651 have an average δ_{25} Mg_{DSM-3} = -0.15 ± 0.04‰ (2SD; Teng et al., 2010), and bulk achondrites 652 overlap with the terrestrial mantle and chondritic Mg isotope ratios (Sedaghatpour and Teng, 653 2016). In our study, meteorite olivine RMs include samples from an ureilites (Kenna and 654 ALH77257), a pallasite (Springwater), a winonaite (Winona), and an ungrouped achondrite 655 (NWA 7325). These meteoritic oliving RMs have a limited variation in δ_{25} Mgds_{M-3} values 656 ranging from -0.07 ± 0.14 to $0.04 \pm 0.17\%$ (2SD) that are indistinguishable from the average 657 δ_{25} MgDSM-3 value of the Earth's mantle and chondrites.

658 Olivine and pyroxene RMs that were either synthetically produced or came from a 659 metamorphosed carbonate have Mg isotope compositions that are significantly different as 660 compared to terrestrial mantle or chondritic Mg isotope compositions. For these types of 661 samples, one must determine their Mg isotope composition by either SN-MC-ICP-MS or fs-LA-662 MC-ICP-MS if they are to be used to conduct accurate and precise Mg isotope analysis by SIMS. 663 The synthetic pure forsterite RM (HN-Ol), that is endmember Fo, has a δ_{25} MgdsM-3 value -0.37 664 $\pm 0.09 \$ (2SD) as determined by SN-MC-ICP-MS. This Mg isotope composition reflects the 665 Mg isotope composition of the reagents used to synthesize the olivine, as well as any mass-666 dependent fractionation that occurred during its synthesis (e.g., Kita et al., 2012). Additionally, 667 olivine and pyroxene RMs that were obtained from metamorphosed carbonate rocks have Mg 668 isotope compositions that are significantly different than mantle or chondritic compositions. 669 These differences arise because Mg-bearing carbonates have lower δ_{25} Mg_{DSM-3} values as 670 compared to silicates (e.g., Young and Galy, 2004). These low $\delta_{25}Mg_{DSM-3}$ values reflect the fact 671 that at equilibrium Mg-bearing carbonates have low δ_{25} Mg_{DSM-3} values by 2-3‰ relative to

672 aqueous Mg at room temperature (Li et al., 2015). Moreover, Mg-bearing dolomite in 673 hydrothermally altered rocks show δ_{25} Mgdsm-3 values ranging from -1.15 to -0.23% (e.g., Azmy 674 et al., 2013; Lavoie et al., 2014; Geske et al., 2015). Two of the olivine RMs are from 675 metamorphosed carbonates (SK-Ol and 95AK-6 Di) and these samples have distinct Mg isotope 676 ratios compared to igneous rocks. SK-Ol is nearly pure forsterite (Fo99.8) that came from a 677 marble from the Beinn an Dubhaich aureole, Isle of Skye, Scotland, where the olivine was 678 produced by the contact metamorphism of dolomite (e.g., Ferry et al., 2011). This olivine has 679 the lowest δ_{25} Mgdsm-3 value (-0.89 ± 0.20‰, 2SD) of all the RMs considered in this study. 680 Additionally, 95AK-6 Di is a clinopyroxene RM, that is from a marble, and this sample has the 681 second lowest δ_{25} MgDSM-3 value (-0.74 ± 0.05‰, 2SD).

682

5.3. Implications for Mg isotope fractionations of olivine in meteoritic components

684 In the WiscSIMS laboratory, 4 out of 17 olivine RMs (SC-Ol, HN-Ol, IG-Ol, and OR-Ol) 685 have been used for instrumental bias corrections for Mg isotope analyses of olivine in Ca, Al-686 rich inclusions (CAIs), amoeboid olivine aggregates (AOAs), and chondrules (MacPherson et al., 687 2012, 2017; Ushikubo et al., 2013, 2017; Hertwig et al., 2019; Tenner et al., 2019). Although these previous analyses were conducted to determine the excess $\delta_{26}Mg^*$ after mass fractionation 688 689 corrections, MacPherson et al. (2012, 2017) and Ushikubo et al. (2017) reported δ_{25} Mg values of 690 olivine grains in CAIs and AOAs, which show large variations ranging from -2.8 to 11.3%. The 691 observed variations are much larger than the observed range of the instrumental bias on 17 692 olivine RMs in this study so that some extent of variations may exist in Mg isotope ratios of 693 olivine among CAIs and AOAs. Furthermore, Ushikubo et al. (2013) reported δ_{25} Mg values of 694 olivine grains in chondrules, which are systematically higher (up to 2.3‰) than those of 695 coexisting phases (pyroxene and plagioclase). Based on the observations, Ushikubo et al. (2013) 696 suggested that chondrule melting likely occurred in an open system process. However, these data 697 were corrected for IMF using a calibration scheme based on the relationship between 698 instrumental bias and Fo content (Fig. 2a), which fortuitously showed linear relationship if only

699 these 4 RMs (SC-Ol, HN-Ol, IG-Ol, and OR-Ol) were utilized. The IMF of these 4 RMs that 700 were obtained by using the Duoplasmatron source (Ushikubo et al., 2013) are consistent with our 701 results that are obtained by using the RF plasma ion source (Fig. 2a), suggesting that the complex 702 IMF observed in this study is not due to the difference in the primary ion sources (i.e., 703 Duoplasmatron versus RF plasma). Moreover, as noted in section 5.1.2, Cr-bearing olivine RMs 704 (up to 0.8 wt% in Cr2O₃) tend to show larger bias*25 (RM-SCOI) values relative to those of olivine 705 RMs with similar Fo contents (Fig. 7), which were not included for the calibration during the 706 analysis session (Fig. 2a). Olivine grains in chondrules measured by Ushikubo et al. (2013) 707 typically have substantial amount of Cr (up to 1.1 wt% in Cr2O3) so that at least a part of the 708 observed Mg isotope variability identified in chondrule olivine may be an analytical artifacts 709 caused by matrix effects. Note that the simple linear regression between IMF and Fo content as 710 done by Ushikubo et al. (2013) may induce additional uncertainties on the IMF correction even if 711 Cr-bearing RMs are not considered for the regression. Further studies based on a suitable set of 712 standards and new calibration scheme based on the combination of Mg+/Si+ ratios and Fo 713 contents is required to more accurately determine Mg isotope compositions in chondrule olivine 714 grains. Accurate Mg isotope data that is free of matrix effects will allow one to better address if 715 chondrules formed in an open system in the protoplanetary disk. 716 717 718 6. Conclusion 719 720 We performed Mg isotope analyses of 17 olivine and 5 pyroxene RMs using MC-ICP-721 MS and MC-SIMS and evaluated SIMS matrix effects on Mg isotope analysis. 722 723 (1) No significant matrix effects on fs-LA-MC-ICP-MS analysis of olivine (Fo59.3 to Fo100) 724 within analytical uncertainties ($\leq 0.2\%$ in δ_{25} Mg, 2SD) under the wet plasma condition (a 725 rate of introducing water = $40 \mu L/min$).

- 726(2) The IMF for SIMS Mg isotope analysis of olivine (Fos9.3 to Fo100) ranges over $\sim 3.3\%$ in727 δ_{25} Mg and is a complex function of Fo content. Moreover, the inferred IMF changes by728 $\sim 2.9\%$ in δ_{25} Mg during a long duration (50 min) spot analysis, suggesting that the IMF of729an olivine analysis may change as a function of the depth of the analysis.
- (3) For olivine analyses, the relationship between variations in the Mg/Si RSF and IMF, as a
 function of olivine Fo content, are similar suggesting that the instrumental mass bias
 changes in part because of differences in the ionization efficiency of Mg and Si. Minor
 element abundances (e.g., Cr) may also influence secondary ionization processes,
 especially for Mg-rich olivine. This minor element variability may be the cause of the
 complex behavior of the IMF as a function of Fo content for the high Fo olivine RMs.
- (4) The instrumental bias among most olivine RMs can be fitted using a quadratic function
 relative to the sensitivity of Mg and Si ions, expressed as (24Mg+/28Si+/Fo). According to
 the regression curve, precision and accuracy of bias*25 (RM-SCOI) correction would be
 0.3‰ for O2- analysis.
- 740(5) The magnitudes of IMF on SIMS Mg isotope analysis of pyroxene (En48.6 to En96.3) range741over $\sim 1.4\%$ in δ_{25} Mg, which is approximately one half the range as measured in olivine742over a similar range in Fe/Mg variation. The Mg isotope IMF on pyroxene is not a743smooth function of En content, indicating that additional factors (e.g., minor element744abundances) may influence matrix effects. Further studies are required for accurate SIMS745Mg isotope analysis of pyroxene.
- 746(6) Most olivine and pyroxene RMs from igneous rocks do not show significant variations in747 $\delta_{25}Mg_{DSM-3}$ values ($\leq 0.3\%$). Therefore, $\delta_{25}Mg_{DSM-3}$ values of these samples may be748assumed to have the same value of the Earth's mantle (-0.13‰: Teng et al., 2010) and749such samples can be used as RMs for SIMS Mg isotope analysis. In contrast, $\delta_{25}Mg_{DSM-3}$ 750values of RMs from olivine and pyroxene derived from metamorphosed carbonate rocks751have $\delta_{25}Mg_{DSM-3}$ values that are ~1‰ lower than that of the Earth's mantle. Synthetic752olivine and pyroxene are also likely to have $\delta_{25}Mg_{DSM-3}$ values that do not match the

terrestrial mantle and for these samples the δ_{25} Mgdsm-3 values of these RMs must be determined by independent methods such as SN-MC-ICP-MS or fs-LA-MC-ICP-MS. Acknowledgements We are grateful to Kouki Kitajima for his valuable comments and developments of SIMS operation and data reduction procedures. We thank Timothy McCoy and Julie Hoskin (Smithsonian Institution) for allocating Winona, Kenna, and Springwater meteorites for the study. We thank Katsuyuki Yamashita, Cyrena Goodrich, Alexander Sobolev, Shichun Huang, and Fang-Zhen Teng for generously providing meteorite and rock samples (ALH77257, NWA 7325, Weltevreden komatiite, and CL09 series). Several WiscSIMS oxygen standards used in this work were provided by John Valley. We thank Guillaume Siron for assistance with electron microprobe analyses. We are grateful to Steve Romaniello and Rebekah Hines for their invaluable assistance in the IGCL at ASU. We also thank Kazuhide Nagashima and an anonymous reviewer for constructive comments that improved the quality of the paper and Balz Kamber for prompt editorial handling of this paper. This work is supported by the NASA program (NNX16AG80G to NK and NNX15AH41G to MW). WiscSIMS is partly supported by NSF (EAR 1658823). References

- Albarède, F., Beard, B., 2004. Analytical Methods for Non-Traditional Isotopes. Rev. Mineral.
 Geochemistry 55, 113–152.
- Albarède, F., Telouk, P., Blichert-Toft, J., Boyet, M., Agranier, A., Nelson, B., 2004. Precise and
 accurate isotopic measurements using multiple-collector ICPMS. Geochim. Cosmochim.
- 784 Acta 68, 2725–2744.
- Azmy, K., Lavoie, D., Wang, Z., Brand, U., Al-Aasm, I., Jackson, S., Girard, I., 2013.
- 786 Magnesium-isotope and REE compositions of Lower Ordovician carbonates from eastern
- 787 Laurentia: Implications for the origin of dolomites and limestones. Chem. Geol. 356, 64–75.
- Bourdon, B., Tipper, E.T., Fitoussi, C., Stracke, A., 2010. Chondritic Mg isotope composition of
 the Earth. Geochim. Cosmochim. Acta 74, 5069–5083.
- Bouvier, A., Spivak-Birndorf, L.J., Brennecka, G.A., Wadhwa, M., 2011. New constraints on
 early Solar System chronology from Al-Mg and U-Pb isotope systematics in the unique
 basaltic achondrite Northwest Africa 2976. Geochim. Cosmochim. Acta 75, 5310–5323.
- 793 Bullock, E.S., Knight, K.B., Richter, F.M., Kita, N.T., Ushikubo, T., MacPherson, G.J., Davis,
- A.M., Mendybaev, R.A., 2013. Mg and Si isotopic fractionation patterns in types B1 and B2
- 795 CAIs: Implications for formation under different nebular conditions. Meteorit. Planet. Sci.
 796 48, 1440–1458.
- Catanzaro, E.J., Murphy, T.J., Garner, E.L., Shields, W.R., 1966. Absolute isotopic abundance
 ratios and atomic weight of magnesium. J. Res. Natl. Bur. Stand. Sect. A Phys. Chem. 70A,
 453–458.
- 800 Chaussidon, M., Deng, Z., Watson, B., Richter, F., 2017. In Situ Analysis of Non-Traditional
- Isotopes by SIMS and LA–MC–ICP–MS: Key Aspects and the Example of Mg Isotopes in
 Olivines and Silicate Glasses. Rev. Mineral. Geochemistry 82, 127–163.
- 803 Davis, A.M., Richter, F.M., Mendybaev, R.A., Janney, P.E., Wadhwa, M., McKeegan, K.D.,
- 804 2015. Isotopic mass fractionation laws for magnesium and their effects on 26Al–26Mg
- systematics in solar system materials. Geochim. Cosmochim. Acta 158, 245–261.
- 806 Donovan, J.J., Tingle, T.N., 1996. An improved mean atomic number background correction for

- 807 quantitative microanalysis. Microsc. Microanal. 2, 1–7.
- Eiler, J.M., Graham, C., Valley, J.W., 1997. SIMS analysis of oxygen isotopes: Matrix effects in
 complex minerals and glasses. Chem. Geol. 138, 221–244.
- Ferry, J.M., Ushikubo, T., Valley, A.W., 2011. Formation of forsterite by silicification of
 dolomite during contact metamorphism. J. Petrol. 52, 1619–1640.
- Galy, A., Belshaw, N.S., Halicz, L., O'Nions, R.K., 2001. High-precision measurement of
- 813 magnesium isotopes by multiple-collector inductively coupled plasma mass spectrometry.
 814 Int. J. Mass Spectrom. 208, 89–98.
- 815 Galy, A., Yoffe, O., Janney, P.E., Williams, R.W., Cloquet, C., Alard, O., Halicz, L., Wadhwa,
- 816 M., Hutcheon, I.D., Ramon, E., Carignan, J., 2003. Magnesium isotope heterogeneity of the
- 817 isotopic standard SRM980 and new reference materials for magnesium-isotope-ratio
 818 measurements. J. Anal. At. Spectrom. 18, 1352.
- 819 Geske, A., Goldstein, R.H., Mavromatis, V., Richter, D.K., Buhl, D., Kluge, T., John, C.M.,
- 820 Immenhauser, A., 2015. The magnesium isotope (δ_{26} Mg) signature of dolomites. Geochim.
- 821 Cosmochim. Acta 149, 131–151.
- 822 Goodrich, C.A., Kita, N.T., Yin, Q.Z., Sanborn, M.E., Williams, C.D., Nakashima, D., Lane,
- 823 M.D., Boyle, S., 2017. Petrogenesis and provenance of ungrouped achondrite Northwest
- 824 Africa 7325 from petrology, trace elements, oxygen, chromium and titanium isotopes, and
- mid-IR spectroscopy. Geochim. Cosmochim. Acta 203, 381–403.
- Handler, M.R., Baker, J.A., Schiller, M., Bennett, V.C., Yaxley, G.M., 2009. Magnesium stable
 isotope composition of Earth's upper mantle. Earth Planet. Sci. Lett. 282, 306–313.
- 828 Hertwig, A.T., Kimura, M., Ushikubo, T., Defouilloy, C., Kita, N.T., 2019. The 26Al-26Mg
- systematics of FeO-rich chondrules from Acfer 094: Two chondrule generations distinct in
 age and oxygen isotope ratios. Geochim. Cosmochim. Acta 253, 111–126.
- 831 Horn, I., von Blanckenburg, F., 2007. Investigation on elemental and isotopic fractionation
- during 196 nm femtosecond laser ablation multiple collector inductively coupled plasma
- mass spectrometry. Spectrochim. Acta Part B At. Spectrosc. 62, 410–422.

- Hu, Y., Teng, F.Z., Zhang, H.F., Xiao, Y., Su, B.X., 2016a. Metasomatism-induced mantle
 magnesium isotopic heterogeneity: Evidence from pyroxenites. Geochim. Cosmochim. Acta
 185, 88–111.
- Hu, Y., Harrington, M.D., Sun, Y., Yang, Z., Konter, J., Teng, F.Z., 2016b. Magnesium isotopic
- 838 homogeneity of San Carlos olivine: a potential standard for Mg isotopic analysis by multi-
- collector inductively coupled plasma mass spectrometry. Rapid Commun. Mass Spectrom.
- 840 2123–2132.
- 841 Isa, J., Kohl, I.E., Liu, M.C., Wasson, J.T., Young, E.D., McKeegan, K.D., 2017. Quantification
- of oxygen isotope SIMS matrix effects in olivine samples: Correlation with sputter rate.
- 843 Chem. Geol. 458, 14–21.
- Jacobsen, B., Yin, Q. zhu, Moynier, F., Amelin, Y., Krot, A.N., Nagashima, K., Hutcheon, I.D.,
 Palme, H., 2008. 26Al–26Mg and 207Pb–206Pb systematics of Allende CAIs: Canonical solar
 initial 26Al/27Al ratio reinstated. Earth Planet. Sci. Lett. 272, 353–364.
- Kawasaki, N., Kato, C., Itoh, S., Wakaki, S., Ito, M., Yurimoto, H., 2015. 26Al-26Mg chronology
- and oxygen isotope distributions of multiple melting for a Type C CAI from Allende.
- 849 Geochim. Cosmochim. Acta 169, 99–114.
- 850 Kawasaki, N., Itoh, S., Sakamoto, N., Yurimoto, H., 2017. Chronological study of oxygen
- isotope composition for the solar protoplanetary disk recorded in a fluffy Type A CAI from
 Vigarano. Geochim. Cosmochim. Acta 201, 83–102.
- 853 Kawasaki, N., Simon, S.B., Grossman, L., Sakamoto, N., Yurimoto, H., 2018. Crystal growth
- and disequilibrium distribution of oxygen isotopes in an igneous Ca-Al-rich inclusion from
- the Allende carbonaceous chondrite. Geochim. Cosmochim. Acta 221, 318–341.
- 856 Kawasaki, N., Park, C., Sakamoto, N., Park, S.Y., Kim, H.N., Kuroda, M., Yurimoto, H., 2019.
- 857 Variations in initial 26Al/ 27Al ratios among fluffy Type A Ca–Al-rich inclusions from
- reduced CV chondrites. Earth Planet. Sci. Lett. 511, 25–35.
- Kita, N.T., Nagahara, H., Togashi, S., Morishita, Y., 2000. A short duration of chondrule
- 860 formation in the solar nebula: Evidence from 26Al in Semarkona ferromagnesian

- 861 chondrules. Geochim. Cosmochim. Acta 64, 3913–3922.
- Kita, N.T., Ushikubo, T., Fu, B., Valley, J.W., 2009. High precision SIMS oxygen isotope
 analysis and the effect of sample topography. Chem. Geol. 264, 43–57.
- Kita, N.T., Nagahara, H., Tachibana, S., Tomomura, S., Spicuzza, M.J., Fournelle, J.H., Valley,
- J.W., 2010. High precision SIMS oxygen three isotope study of chondrules in LL3
- chondrites: Role of ambient gas during chondrule formation. Geochim. Cosmochim. Acta
 74, 6610–6635.
- Kita, N.T., Ushikubo, T., Knight, K.B., Mendybaev, R.A., Davis, A.M., Richter, F.M., Fournelle,
- J.H., 2012. Internal 26Al–26Mg isotope systematics of a Type B CAI: Remelting of
 refractory precursor solids. Geochim. Cosmochim. Acta 86, 37–51.
- Koefoed, P., Amelin, Y., Yin, Q.Z., Wimpenny, J., Sanborn, M.E., Iizuka, T., Irving, A.J., 2016.
- 872 U–Pb and Al–Mg systematics of the ungrouped achondrite Northwest Africa 7325.
- 873 Geochim. Cosmochim. Acta 183, 31–45.
- Krot, A.N., Nagashima, K., Wasserburg, G.J., Huss, G.R., Papanastassiou, D., Davis, A.M.,
- 875 Hutcheon, I.D., Bizzarro, M., 2014. Calcium-aluminum-rich inclusions with fractionation
- and unknown nuclear effects (FUN CAIs): I. Mineralogy, petrology, and oxygen isotopic
 compositions. Geochim. Cosmochim. Acta 145, 206–247.
- 878 Larsen, K.K., Trinquier, A., Paton, C., Schiller, M., Wielandt, D., Ivanova, M.A., Connelly, J.N.,
- 879 Nordlund, Å., Krot, A.N., Bizzarro, M., 2011. Evidence for magnesium isotope
- heterogeneity in the solar protoplanetary disk. Astrophys. J. Lett. 735.
- 881 Lavoie, D., Jackson, S., Girard, I., 2014. Magnesium isotopes in high-temperature saddle
- dolomite cements in the lower Paleozoic of Canada. Sediment. Geol. 305, 58–68.
- Li, W., Beard, B.L., Li, C., Xu, H., Johnson, C.M., 2015. Experimental calibration of Mg isotope
- fractionation between dolomite and aqueous solution and its geological implications.
- 885 Geochim. Cosmochim. Acta 157, 164–181.
- Luu, T.-H., Chaussidon, M., Mishra, R.K., Rollion-Bard, C., Villeneuve, J., Srinivasan, G.,
- 887 Birck, J.-L., 2013. High precision Mg isotope measurements of meteoritic samples by

- secondary ion mass spectrometry. J. Anal. At. Spectrom. 28, 67–76.
- MacPherson, G.J., Kita, N.T., Ushikubo, T., Bullock, E.S., Davis, A.M., 2012. Well-resolved
- 890 variations in the formation ages for Ca-Al-rich inclusions in the early Solar System. Earth
- 891 Planet. Sci. Lett. 331–332, 43–54.
- 892 MacPherson, G.J., Bullock, E.S., Tenner, T.J., Nakashima, D., Kita, N.T., Ivanova, M.A., Krot,
- A.N., Petaev, M.I., Jacobsen, S.B., 2017. High precision Al–Mg systematics of forsterite-

bearing Type B CAIs from CV3 chondrites. Geochim. Cosmochim. Acta 201, 65–82.

895 Mendybaev, R.A., Richter, F.M., Georg, R.B., Janney, P.E., Spicuzza, M.J., Davis, A.M.,

Valley, J.W., 2013. Experimental evaporation of Mg- and Si-rich melts: Implications for the
origin and evolution of FUN CAIs. Geochim. Cosmochim. Acta 123, 368–384.

898 Mendybaev, R.A., Williams, C.D., Spicuzza, M.J., Richter, F.M., Valley, J.W., Fedkin, A. V.,

899 Wadhwa, M., 2017. Thermal and chemical evolution in the early Solar System as recorded

- 900 by FUN CAIs: Part II Laboratory evaporation of potential CMS-1 precursor material.
- 901 Geochim. Cosmochim. Acta 201, 49–64.

902 Oeser, M., Weyer, S., Horn, I., Schuth, S., 2014. High-precision fe and mg isotope ratios of

silicate reference glasses determined in situ by femtosecond LA-MC-ICP-MS and by
solution nebulisation MC-ICP-MS. Geostand. Geoanalytical Res. 38, 311–328.

905 Olsen, M.B., Wielandt, D., Schiller, M., Van Kooten, E.M.M.E., Bizzarro, M., 2016. Magnesium

and 54Cr isotope compositions of carbonaceous chondrite chondrules – Insights into early
disk processes. Geochim. Cosmochim. Acta 191, 118–138.

908 Park, C., Nagashima, K., Krot, A.N., Huss, G.R., Davis, A.M., Bizzarro, M., 2017. Calcium-

aluminum-rich inclusions with fractionation and unidentified nuclear effects (FUN CAIs):

- 910 II. Heterogeneities of magnesium isotopes and 26Al in the early Solar System inferred from
- 911 in situ high-precision magnesium-isotope measurements. Geochim. Cosmochim. Acta 201,
- 912 6–24.

913 Peres, P., Kita, N.T., Valley, J.W., Fernandes, F., Schuhmacher, M., 2013. New sample holder

geometry for high precision isotope analyses. Surf. Interface Anal. 45, 553–556.

- 915 Pogge von Strandmann, P.A.E., Burton, K.W., James, R.H., van Calsteren, P., Gislason, S.R.,
- 916 Sigfússon, B., 2008. The influence of weathering processes on riverine magnesium isotopes
 917 in a basaltic terrain. Earth Planet. Sci. Lett. 276, 187–197.
- 918 Richter, F.M., Janney, P.E., Mendybaev, R.A., Davis, A.M., Wadhwa, M., 2007. Elemental and
- 919 isotopic fractionation of Type B CAI-like liquids by evaporation. Geochim. Cosmochim.
 920 Acta 71, 5544–5564.
- Schiller, M., Handler, M.R., Baker, J.A., 2010. High-precision Mg isotopic systematics of bulk
 chondrites. Earth Planet. Sci. Lett. 297, 165–173.
- 923 Scicchitano, M.R., Rubatto, D., Hermann, J., Majumdar, A.S., Putnis, A., 2018. Oxygen isotope
- analysis of olivine by ion microprobe: Matrix effects and applications to a serpentiniseddunite. Chem. Geol. 499, 126–137.
- Sedaghatpour, F., Teng, F.Z., 2016. Magnesium isotopic composition of achondrites. Geochim.
 Cosmochim. Acta 174, 167–179.
- 928 Śliwiński, M.G., Kitajima, K., Kozdon, R., Spicuzza, M.J., Fournelle, J.H., Denny, A., Valley,
- 929 J.W., 2016a. Secondary Ion Mass Spectrometry Bias on Isotope Ratios in Dolomite-
- 930 Ankerite, Part I: δ18O Matrix Effects. Geostand. Geoanalytical Res. 40, 157–172.
- 931 Śliwiński, M.G., Kitajima, K., Kozdon, R., Spicuzza, M.J., Fournelle, J.H., Denny, A., Valley,
- 932 J.W., 2016b. Secondary Ion Mass Spectrometry Bias on Isotope Ratios in Dolomite-
- 933 Ankerite, Part II: δ13C Matrix Effects. Geostand. Geoanalytical Res. 40, 173–184.
- 934 Śliwiński, M.G., Kitajima, K., Spicuzza, M.J., Orland, I.J., Ishida, A., Fournelle, J.H., Valley,
- 935 J.W., 2018. SIMS Bias on Isotope Ratios in Ca-Mg-Fe Carbonates (Part III): δ18O and δ13C
- 936 Matrix Effects Along the Magnesite–Siderite Solid-Solution Series. Geostand.
- 937 Geoanalytical Res. 42, 49–76.
- 938 Spivak-Birndorf, L., Wadhwa, M., Janney, P., 2009. 26Al–26Mg systematics in D'Orbigny and
- 939 Sahara 99555 angrites: Implications for high-resolution chronology using extinct
- 940 chronometers. Geochim. Cosmochim. Acta 73, 5202–5211.
- 941 Steele, I.M., Hervig, R.L., Hutcheon, I.D., Smith, J. V., 1981. Ion microprobe techniques and

- analyses of olivine and low-Ca pyroxene. Am. Mineral. 66, 526–546.
- Steinhoefel, G., Horn, I., von Blanckenburg, F., 2009. Matrix-independent Fe isotope ratio
 determination in silicates using UV femtosecond laser ablation. Chem. Geol. 268, 67–73.
- 945 Teng, F.Z., Li, W.Y., Ke, S., Marty, B., Dauphas, N., Huang, S., Wu, F.Y., Pourmand, A., 2010.
- 946 Magnesium isotopic composition of the Earth and chondrites. Geochim. Cosmochim. Acta
 947 74, 4150–4166.
- 948 Teng, F.Z., 2017. Magnesium Isotope Geochemistry. Rev. Mineral. Geochemistry 82, 219–287.
- 949 Tenner, T.J., Nakashima, D., Ushikubo, T., Tomioka, N., Kimura, M., Weisberg, M.K., Kita,
- 950 N.T., 2019. Extended chondrule formation intervals in distinct physicochemical
- 951 environments: Evidence from Al–Mg isotope systematics of CR chondrite chondrules with
- unaltered plagioclase. Geochim. Cosmochim. Acta 260, 133–160.
- Tipper, E.T., Galy, A., Bickle, M.J., 2006. Riverine evidence for a fractionated reservoir of Ca
 and Mg on the continents: Implications for the oceanic Ca cycle. Earth Planet. Sci. Lett.
 247, 267–279.
- 956 Ushikubo, T., Nakashima, D., Kimura, M., Tenner, T.J., Kita, N.T., 2013. Contemporaneous
- 957 formation of chondrules in distinct oxygen isotope reservoirs. Geochim. Cosmochim. Acta
 958 109, 280–295.
- Ushikubo, T., Tenner, T.J., Hiyagon, H., Kita, N.T., 2017. A long duration of the ¹⁶O-rich
 reservoir in the solar nebula, as recorded in fine-grained refractory inclusions from the least
 metamorphosed carbonaceous chondrites. Geochim. Cosmochim. Acta 201, 103–122.
- Valley, J.W., Kita, N.T., 2009. In situ oxygen isotope geochemistry by ion microprobe. Mineral.
 Assoc. Canada Short Course 19–63.
- Van Kooten, E.M.M.E., Wielandt, D., Schiller, M., Nagashima, K., Thomen, A., Larsen, K.K.,
- 965 Olsen, M.B., Nordlund, Å., Krot, A.N., Bizzarro, M., 2016. Isotopic evidence for primordial
- 966 molecular cloud material in metal-rich carbonaceous chondrites. Proc. Natl. Acad. Sci. 113,
- 967 2011–2016.
- 968 Villeneuve, J., Chaussidon, M., Marrocchi, Y., Deng, Z., Watson, E.B., 2019. High-precision In
- situ silicon isotopic analyses by MC-SIMS in olivine and low-Ca pyroxene. Rapid
- 970 Commun. Mass Spectrom. 33, 1589–1597.
- 971 Wasserburg, G.J., Wimpenny, J., Yin, Q.Z., 2012. Mg isotopic heterogeneity, Al–Mg isochrons,

And canonical 26Al/27Al in the early solar system. Meteorit. Planet. Sci. 47, 1980–1997.

- 973 Williams, C.D., Ushikubo, T., Bullock, E.S., Janney, P.E., Hines, R.R., Kita, N.T., Hervig, R.L.,
- 974 MacPherson, G.J., Mendybaev, R.A., Richter, F.M., Wadhwa, M., 2017. Thermal and
- 975 chemical evolution in the early solar system as recorded by FUN CAIs: Part I Petrology,
- 976 mineral chemistry, and isotopic composition of Allende FUN CAI CMS-1. Geochim.
- 977 Cosmochim. Acta 201, 25–48.
- Xiao, Y., Teng, F.Z., Zhang, H.F., Yang, W., 2013. Large magnesium isotope fractionation in
 peridotite xenoliths from eastern North China craton: Product of melt-rock interaction.
 Geochim. Cosmochim. Acta 115, 241–261.
- Young, E.D., Ash, R.D., Galy, A., Belshaw, N.S., 2002. Mg isotope heterogeneity in the Allende
 meteorite measured by UV laser ablation-MC-ICPMS and comparisons with O isotopes.
- 983 Geochim. Cosmochim. Acta 66, 683–698.
- Young, E.D., Galy, A., 2004. The isotope geochemistry and cosmochemistry of magnesium.
 Rev. Mineral. Geochemistry 55, 197–230.
- Zheng, X.Y., Beard, B.L., Johnson, C.M., 2018. Assessment of matrix effects associated with Fe
 isotope analysis using 266 nm femtosecond and 193 nm nanosecond laser ablation multicollector inductively coupled plasma mass spectrometry. J. Anal. At. Spectrom. 33, 68–83.
- 989
- 990

991 Figure captions

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Fig. 1. A secondary electron image of an example of SIMS pit. This image was taken after Mg
isotope analyses by using O₂- primary ion beam (1 nA, session FI3). One analysis takes 8 min.

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Fig. 2. The bias*25 (RM-SCOI) values of 17 olivine RMs obtained by (a) O- and (b) O2- analyses as a function of Fo content. The dashed line in Fig. 2a represents the best-fit line for data obtained by using the Duoplasmatron source reported in Ushikubo et al. (2013). Cr-bearing (~0.8 wt% in Cr2O3) RMs tend to show higher bias*25 (RM-SCOI) values relative to those of RMs with similar Fo contents, except for KN-Ol (Fo78.0). Errors are 2σ .

1001

1002Fig. 3. The f^{*}_{25} values of 11 olivine RMs for (a and b) O- and (c and d) O2- analyses against1003cycle number. One cycle takes 10 seconds so that one analysis (300 cycles) takes ~50 min. Gray1004bands represent cycle numbers correspond to depths that Mg isotope data (listed in Table 6) were1005obtained.

1006

Fig. 4. The bias*25 (RM-SCOI) values of 11 olivine RMs with different depths obtained by (a) Oand (b) O₂- analyses as a function of Fo content. The different symbols and lines represent the average of 50 cycles taken over the total 300 cycle analysis. Error bars are not shown for clarity, but would typically be \pm 0.5‰ (2SD), except for the first 100 cycles of O- analysis that show large changes in *f**25 values (Figs. 3a-b) so that errors are larger (~ 1‰) than those for the rest of the cycles.

1013

1014 Fig. 5. Compositional dependences of secondary ion yields and relative sensitivities of Mg+/Si+ 1015 as a function of olivine Fo content. (a, b) Secondary 24Mg+ ion yields vs. Fo contents for O- and 1016 O₂- analyses, respectively. (c, d) Secondary 28Si+ ion yields vs. Fo contents for O₋ and O₂-1017 analyses, respectively. (e, f) Relative sensitivity factors of Mg+/Si+ (Mg/Si RSFs) vs. Fo contents 1018 for O- and O2- analyses, respectively. Errors of ion yields and Mg/Si RSFs are 2SE. For 1019 comparison, the bias*25 (RM-SCOI) values of 17 olivine RMs as shown in Fig. 2a and 2b are also 1020 plotted as gray square symbols. Errors of the bias^{*}25 (RM-SCOI) values are 2σ . Yield = count rate 1021 (per second) / primary ion beam intensity (nA). Mg/Si RSF = (24Mg+/28Si+)/(Mg cpfu/Si cpfu). 1022

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Fig. 6. The bias*25 (RM-SCOI) values of 5 pyroxene RMs obtained by O- and O2- analyses as a function of En content. Data for O- analysis (open symbols) are from Ushikubo et al. (2013, 2017), which are re-normalized to the DSM-3 scale using the δ_{25} MgDSM-3 values reported in this study. Errors are 2σ .

1027

1028 Fig. 7. Differences between bias*25 (RM-SCOI) and modelled bias*25 (RM-SCOI) cale values of Cr-1029 bearing olivine RMs as a function of Cr2O3 content obtained by EPMA analyses. The modelled 1030 value (bias*25 (RM-SCOI) calc) is calculated by assuming linear correlation between bias*25 (RM-SCOI) 1031 value and Fo content among olivine RMs with similar Fo contents (see Appendix EA8 for 1032 details). A solid and dashed lines represent least-squares regression lines for Cr-bearing olivine 1033 RMs with Fo \geq 86 (O₋ analysis) and that with Fo \geq 94 (O₂₋ analysis), respectively. Assuming 1034 intercepts with zero, the slopes are determined to be 1.7 ± 0.3 (2 σ , MSDW = 0.6) for O₋ analysis 1035 and 1.4 ± 0.4 (2 σ , MSDW = 0.4) for O₂- analysis, respectively. Errors of Cr₂O₃ (wt%) and 1036 residual bias are 2SD and 2σ , respectively.

1037

1038 Fig. 8. The relationships between bias*25 (RM-SCOI) values and 24Mg+/28Si+ ratios divided by Fo 1039 contents of 17 olivine RMs. (a) In the case of O- analysis, the bias^{*25} (RM-SCOI) values of 15 1040 olivine RMs with F059.3 to F097.4 are positively correlated with (24Mg+/28Si+/F0) values. The 1041 dashed curve represents a quadratic function regression curve (bias*25 (RM-SCOI) = $-0.08 \times$ 1042 $(24Mg+/28Si+/Fo)^2 + 1.76 \times (24Mg+/28Si+/Fo) - 7.48; R_2 = 0.98)$. Note that two nearly pure 1043 forsterite RMs (SK-Ol and HN-Ol) do not follow the regression curve. (b) In the case of O₂-1044 analysis, the bias^{*25} (RM-SCOI) values can be fitted into two separate quadratic curves for Fo 1045 contents ranging from 59.3 to 88.8 and 88.8 to 100, respectively. The dashed curve represents 1046 the quadratic function regression curve for Fo content ≤ 88.8 (bias*25 (RM-SCOI) = $-0.31 \times$ 1047 $(24Mg+/28Si+/Fo)^2 + 6.77 \times (24Mg+/28Si+/Fo) - 36.14; R_2 = 0.96)$ and the solid curve represents the 1048 quadratic function regression curve for Fo content \geq 88.8 (bias*25 (RM-SCOI) = $-0.87 \times$ $(24Mg+/28Si+/Fo)^2 + 15.43 \times (24Mg+/28Si+/Fo) - 68.67; R_2 = 0.97)$. (c, d) Plots of the calibration 1049

- 1050 residuals as a function of Fo content for O- and O2- analyses, respectively. For O2- analysis,
- 1051 bias* $_{25}$ (RM-SCOI) values of all RMs differs by < 0.3‰ (depicted by dashed lines) from the value
- 1052 predicted by the calibration (solid line). Errors are 2σ .
- 1053
- 1054 Appendix
- 1055
- 1056 Appendix EA1: Sample descriptions and laser fluorination oxygen isotope data
- 1057 Appendix EA2: Typical mass spectrum for SIMS Mg isotope analysis
- 1058 Appendix EA3: EPMA data
- 1059 Appendix EA4: Summary of Fo and En contents, and $\delta_{25}Mg_m$ homogeneities
- 1060 Appendix EA5: List of SIMS analytical sessions
- 1061 Appendix EA6: SIMS Mg isotope data
- 1062 Appendix EA7: SIMS major and minor element data
- 1063 Appendix EA8: Calculation of residual biases for Cr-bearing olivine reference materials

Sample name	Sample name Fo or En content _a Provenance / Meteorite nam		Rock type
Olivine			
OR-Ol	59.3	Orikabe plutonic complex, Japan	gabbro
FJ-Ol	73.4	Mount Fuji, Japan	gabbro
KN-Ol	78.0	Kenna (Ureilite)	meteorite
CL09-19-Ol	79.4	Beiyan, China	peridotite
CL09-08-O1	80.5	Beiyan, China	peridotite
SW-Ol	81.8	Springwater (pallasite)	meteorite
A77257-Ol	85.7	ALHA77257 (Ureilite)	meteorite
CL09-33-Ol	86.6	Beiyan, China	peridotite
SC-Ol	88.8	San Carlos, USA	peridotite
UWOL-1	89.2	Kilbourne Hole, USA	peridotite
IG-Ol	89.6	Ichinome-gata, Japan	peridotite
HaK-Ol	91.9	Harrat al Kishb, Saudi Arabia	peridotite
WK-Ol	94.3	Barberton, South Africa	komatiite
WN-Ol	95.5	Winona (Winonaite)	meteorite
N7325-Ol	97.4	NWA 7325 (ungrouped achondrite)	meteorite
SK-Ol	99.8	Isle of Skye, Scotland	marble
HN-Ol	100.0	Synthetic forsterite	
Pyroxene			
95AK-6 Di	48.6	Adirondack Mt., USA	marble
IG-Cpx	48.6	Ichinome-gata, Japan	peridotite
JE En	85.5	Unknown	unknown
IG-Opx	88.9	Ichinome-gata, Japan	peridotite
Sp79-11 En	96.3	Adirondack Mt., USA	enstatite

Table 1 List of olivine and pyroxene RMs used in this study.

a Molar % of forsterite or enstatite end-members based on EPMA analyses performed in this study

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Table 2Summary of operating conditions of LA-MC-ICP-MS measurements.

^a Total He gas flow. Sub-equal amounts were delivered to the arm and the chamber

b Total Ar flow. Sub-equal amounts were delivered from the nebulizer used for water addition and Ar delivered directly to the aerosol stream

Table 3 Measured parameters used in this study

-		•
Parameter	Method	Definition
Foa Ena Woa $\delta x M g D S M - 3b$ $\delta 2 s M g m c$	EPMA EPMA EPMA ICP-MS SIMS	Fo = Mg/[Mg + Fe] molar % En = Mg/[Mg + Fe + Ca] molar % Wo = Ca/[Mg + Fe + Ca] molar % $\delta xMg_{DSM-3} = [(xMg/24Mg)_{RM}/(xMg/24Mg)_{DSM-3} - 1] \times 1000 (\%)$ $\delta zSMg_m = [(2SMg/24Mg)_{RM}/0.12663 - 1] \times 1000 (\%)$
<i>f</i> *25 (RM)	SIMS	$f^{*}_{25 (\text{RM})} = [(1 + \delta_{25} \text{Mgm}/1000)/(1 + \delta_{25} \text{MgDSM}-3/1000) - 1] \times 1000 (\%)$
bias*25 (RM-SCOI)	SIMS	bias*25 (RM-SCOI) = $[(1 + f^{*25} (RM)/1000)/(1 + f^{*25} (SCOI)/1000) - 1] \times 1000 (\%)$
Mg/S1 RSFd	SIMS	Mg/S1 RSF = (24Mg+/28S1+)/(Mg cpfu/S1 cpfu)

 $_a$ Fo, En and Wo contents are calculated by considering total iron as ferrous iron. $_b$ X = 25 or 26

c Absolute 25Mg/24Mg ratio (=0.12663) is from Catanzaro et al. (1966) d cpfu = cations per formula unit

Sample name	EPMA session	Fo or En content	Na ₂ O	MgO	Al ₂ O ₃	SiO ₂	CaO	TiO ₂	Cr ₂ O ₃	FeO	NiO	MnO	Total
Olivine													
OR-Ol	2018 Aug & 2019 Jan.a	59.3	-	28.42	-	35.72	0.02	-	b.d.	34.81	b.d.	0.67	99.64
FJ-Ol	2018 Aug & 2019 Jan.a	73.4	-	37.24	-	37.63	0.06	-	b.d.	24.11	0.14	0.42	99.60
KN-Ol	2018 Aug & 2019 Jan.a	78.0	-	39.42	-	38.58	0.38	-	0.80	19.85	b.d.	0.42	99.45
CL09-19-Ol	2019 Jan.	79.4	-	41.41	-	38.26	0.08	-	b.d.	18.98	0.18	0.22	99.13
CL09-08-O1	2019 Jan.	80.5	-	42.08	-	38.54	0.05	-	b.d.	18.13	0.29	0.24	99.33
SW-Ol	2018 Aug & 2019 Jan.a	81.8	-	42.87	-	38.91	b.d.	-	b.d.	16.99	b.d.	0.31	99.08
A77257-Ol	2019 Jan.	85.7	-	44.95	-	39.07	0.32	-	0.80	13.37	b.d.	0.46	98.97
CL09-33-O1	2019 Jan.	86.6	-	46.19	-	39.17	0.06	-	b.d.	12.70	0.40	0.17	98.69
SC-Ol	2019 Jan.	88.8	-	48.22	-	40.01	0.07	-	b.d.	10.82	0.42	0.16	99.69
UWOL-1	2018 Aug & 2019 Jan.a	89.2	-	47.77	-	39.85	0.06	-	b.d.	10.27	0.41	0.14	98.50
IG-Ol	2019 Jan.	89.6	-	48.63	-	40.12	0.05	-	b.d.	10.09	0.42	0.16	99.48
HaK-Ol	2018 Aug & 2019 Jan.a	91.9	-	50.07	-	40.74	0.02	-	b.d.	7.92	0.51	0.10	99.36
WK-Ol	2018 Aug & 2019 Jan.a	94.3	-	51.44	-	41.17	0.12	-	0.23	5.54	0.47	0.09	99.05
WN-Ol	2018 Aug & 2019 Jan.a	95.5	-	53.63	-	41.80	b.d.	-	b.d.	4.49	b.d.	0.33	100.26
N7325-Ol	2019 Jan.	97.4	-	54.69	-	41.61	0.32	-	0.45	2.58	b.d.	0.12	99.77
SK-Ol	2019 Jan.	99.8	-	56.73	-	41.48	0.04	-	b.d.	0.19	b.d.	0.08	98.52
HN-Ol	2019 Jan.	100.0	-	57.31	-	42.04	b.d.	-	b.d.	b.d.	b.d.	b.d.	99.35
Pvroxene													
95AK-6 Di	2019 Sep.	48.6	0.10	17.92	0.92	54.91	25.41	b.d.	b.d.	1.26	-	0.07	100.59
IG-Cpx	2019 Sep.	48.6	0.50	16.70	4.52	51.70	22.51	0.45	0.91	2.65	-	0.09	100.05
JE Ên	2019 Sep.	85.5	b.d.	33.16	0.11	56.21	0.22	b.d.	b.d.	9.72	-	0.07	99.49
IG-Opx	2019 Sep.	88.9	b.d.	33.38	3.56	54.38	0.57	0.13	0.44	6.70	-	0.16	99.31
Sp79-11 En	2019 Sep.	96.3	b.d.	37.89	1.42	57.20	0.44	0.09	b.d.	2.03	-	0.41	99.49

Table 4 Major and minor element compositions (wt%) of 17 olivine and 5 pyroxene RMs obtained by EPMA

b.d. (below detection limit)

- (not measured)

^a Combined data obtained by two different sessions are summarized in this table. Fo content, MgO, SiO₂, CaO, FeO, and MnO are from session_2018Aug (Table EA3-1). Cr₂O₃ and NiO are from session_2019Jan. (Table EA3-2).

	Fo or En				SN-MC-ICP	-MS				fs	LA-M	C-ICP-MS	
Sample name	contenta	δ25Mgdsm-3	2SD	2SE	δ26Mgdsm-3	2SD	2SE	Nb	Ref.	δ25Mgdsm-3	2σc	δ26Mgdsm-3	$2\sigma_c$
Standards													
DSM-3 250 ppb		0.00	0.06		0.00	0.11		147	This study				
Spex Mg		-1.08	0.04		-2.09	0.07		8	This study				
BCR-2		-0.05	0.08		-0.08	0.13		10	This study				
Olivine									-				
OR-Ol	59.3	-0.01	0.06		-0.02	0.13		27	This study	0.02	0.09	-0.05	0.24
FJ-Ol	73.4								2	-0.07	0.20	-0.10	0.30
KN-Ol	78.0									0.04	0.18	-0.01	0.30
SW-Ol	81.8									-0.02	0.14	-0.03	0.16
A77257-Ol	85.7									-0.07	0.14	-0.29	0.13
SC-Ol	88.8	-0.07	0.09		-0.15	0.16		10	This study				
UWOL-1	89.2									-0.14	0.10	-0.24	0.20
IG-Ol	89.6	-0.03	0.09		-0.05	0.15		10	This study	-0.08	0.18	-0.19	0.20
HaK-Ol	91.9									-0.06	0.13	-0.16	0.19
WK-Ol	94.3									-0.09	0.13	-0.17	0.20
WN-Ol	95.5									-0.04	0.09	-0.14	0.16
N7325-Ol	97.4									-0.07	0.14	-0.06	0.10
SK-Ol	99.8									-0.89	0.20	-1.72	0.54
HN-Ol	100.0	-0.37	0.09		-0.70	0.14		10	This study	-0.49	0.14	-0.95	0.17
Pyroxene													
95AK-6 Di	48.6	-0.74	0.05		-1.43	0.11		8	This study				
IG-Cpx	48.6	-0.14	0.10		-0.28	0.19		8	This study				
JE En	85.5	-0.26	0.06		-0.50	0.11		8	This study				
IG-Opx	88.9	-0.04	0.08		-0.07	0.13		8	This study				
Sp79-11 En	96.3	-0.06	0.05		-0.11	0.11		18	This study				
Olivine (Literature data)													
CL09-19-Ol	79.4	-0.21	0.06		-0.37	0.06		2	[1]				
CL09-08-O1	80.5	-0.21	0.05		-0.41	0.07		4	[1]				
CL09-33-Ol	86.6	-0.25	0.05		-0.47	0.06		2	[1]				
Springwater olivine		-0.11		0.03	-0.22		0.04	7	[2]				
NWA 7325 olivine		-0.16		0.02	-0.23		0.04	6	[3]				

Table 5 Mg isotope data of 17 olivine and 5 pyroxene RMs obtained by SN- and/or LA-MC-ICP-MS

^a Molar % of forsterite or enstatite end-members based on EPMA analyses performed in this study

ь Number of analyses

c 2σ errors (‰) are propagated both 2SD of fs-LA-MC-ICP-MS analyses and 2SE of δ25MgDSM-3 (SC-Ol) obtained by SN-MC-ICP-MS analyses.

References; [1] Xiao et al. (2013); [2] Handler et al. (2009); [3] Koefoed et al. (2016).

	EE.	Sessio	on UI1 and UI2	2 (Duo O-	-)	Sess	ion FI4 (RF O-	-)	Sessi	ion FI3 (RF O	2-)	Session FE2 (RF	⁷ O-)	Session FE1 (RF	O2-)
Sample name	content	f*25 (RM) (‰)	bias*25 (RM- SCOI) (‰)	$2 \sigma_a$	Ref.	f*25 (RM) (‰)	bias*25 (RM- SCOI) (‰)	2 σa	f*25 (RM) (‰)	bias*25 (RM- SCOI) (%)	2 σ _a	(24Mg+/28Si+)/Fo (× 100)	2 σь	(24Mg+/28Si+)/Fo (× 100)	2 σь
Olivine															
OR-O	59.3	-0.54	2.31	0.18	[1]	-0.49	1.93	0.23	-1.44	0.42	0.15	8.74	0.78	12.45	0.57
FJ-O	1 73.4					-0.27	2.15	0.30	-1.23	0.62	0.24	8.85	0.29	12.25	0.48
KN-O	1 78.0					-0.27	2.14	0.28	-1.18	0.70	0.23	8.86	0.39	11.92	0.51
CL09-19-O	1 79.4					-1.26	1.35	0.23	-0.94	0.86	0.17	7.59	0.36	11.67	0.46
CL09-08-O	80.5					-1.53	1.07	0.23	-0.86	0.94	0.20	7.10	0.35	11.73	0.48
SW-O	l 81.8					-1.60	0.77	0.26	-0.97	0.90	0.20	6.77	0.18	11.78	0.45
A77257-O	1 85.7					-0.63	1.87	0.27	-0.72	1.10	0.27	7.70	0.24	11.45	0.47
CL09-33-O	l 86.6					-2.16	0.44	0.23	-1.31	0.53	0.15	5.89	0.16	9.71	0.43
SC-O	88.8	-2.84	0.00	0.17	[1]	-2.41	0.00	0.24	-1.83	0.00	0.17	5.59	0.15	9.11	0.29
UWOL-	89.2					-2.39	-0.03	0.24	-2.01	-0.12	0.18	5.55	0.28	9.04	0.29
IG-O	l 89.6	-3.09	-0.25	0.15	[1]	-2.67	-0.07	0.24	-1.96	-0.12	0.17	5.52	0.16	8.75	0.32
HaK-O	l 91.9					-2.61	-0.23	0.26	-2.22	-0.34	0.19	5.38	0.15	8.28	0.26
WK-O	1 94.3					-2.38	0.03	0.25	-2.00	-0.19	0.19	5.63	0.15	8.57	0.31
WN-O	1 95.5					-2.62	-0.22	0.24	-2.31	-0.51	0.19	5.55	0.19	8.04	0.26
N7325-O	1 97.4					-2.07	0.25	0.26	-1.98	-0.21	0.20	5.87	0.27	8.57	0.27
SK-O	99.8					-3.35	-1.01	0.30	-3.14	-1.30	0.25	5.46	0.24	7.75	0.25
HN-O	1 100.0	-3.80	-0.97	0.15	[1]	-3.54	-1.17	0.24	-3.49	-1.62	0.16	5.40	0.23	7.55	0.24
Pyroxene															
95AK-6 D	i 48.6	0.83	3.72	0.18	[2]				0.00	1.81	0.18				
IG-Cp	48.6	1.33	4.23	0.19	[2]				0.55	2.36	0.17				
JE Ei	n 85.5	0.43	3.42	0.15	[1]				1.28	3.09	0.15				
IG-Op	88.9	0.39	3.38	0.24	[1]				1.41	3.22	0.24				
Sp79-11 Er	96.3	-0.27	2.72	0.15	[1]				0.65	2.46	0.15				

Table 6 SIMS δ25Mg instrumental biases of 17 olivine and 5 pyroxene RMs and (24Mg+/28Si+)/Fo values of 17 olivine RMs

 $_{a}$ 2 σ error (‰) = $\sqrt{(2SD_{SIMS})^2 + (2SD_{ICPMS})^2}$, where 2SD_{SIMS} represents either δ_{2S} Mg homogeneity of each RM or long-term reproducibility of the running standard (SC-OI), whichever is larger. b 2 σ error represents either 2 standard error (2SE) of each measurement or long-term reproducibility (2SD) of the running standard (SC-OI), whichever is larger. References; [1] Ushikubo et al. (2013); [2] Ushikubo et al. (2017). All data from Ushikubo et al. (2013, 2017) were renormalized to DSM-3 scale using SN-MC-ICP-MS results.

















Appendix EA1: Sample descriptions and laser fluorination oxygen isotope data

Olivine Reference Materials

- Orikabe Plutonic Complex, Kitakami Mountains, northeast Japan (OR-OI)
 Olivine grains were sampled from a hornblende-biotite-olivine-clinopyroxene gabbro
 from Tokusenjo type, Orikabe Plutonic Complex (sample No. 841022B-9; Mikoshiba et
 al., 2004) as hand-picked 250-500 µm sized fraction by Kita et al. (1998). Fe and Ni
 concentrations were measured by atomic absorption analysis (Kita et al., 1998). Fe, Co,
 and Ni concentrations were also obtained by INAA (Kita et al., 1998). Oxygen isotope
 ratio was obtained by laser fluorination analyses (Kita et al. 2010). An aliquot of the
 same fraction was used for the solution MC-ICP-MS analyses of Mg isotope ratios.
 Additional olivine grains were newly handpicked from 250-500 µm separate of the
 remaining sample in order to mount them in epoxy disk as SIMS standard.
- 2. Hoei Gabbro, Mount Fuji, Japan (FJ-Ol)

These olivine grains were hand-picked from the 500-1000 µm fraction of a gabbro from 1707 Hoei eruption, Mount Fuji (Sample No. FH NTK-3) by Kita et al. (1998). A host rock was among those collected from 1707 Hoei eruption gabbros by Dr. Shigeko Togashi (GSJ-AIST). A specific specimen (FH NTK-3) was selected with the purpose of extracting olivine grains for SIMS standards according to the occurrence of coarse (>1mm) and clean olivine. Fe and Ni concentrations were measured by atomic absorption analysis (Kita et al., 1998). Fe, Co, and Ni concentrations were also obtained by INAA (Kita et al., 1998).

Kenna, Ureilite meteorite (KN-Ol), USNM5825
 Kenna is an ureilite meteorite that was found in 1972, New Mexico, USA. Olivine from

Kenna shows a xenoblastic texture, which is commonly rimmed by forsterite (Berkley et al., 1976). A chip of the meteorite was provided by National Museum of Natural History (Smithsonian Institute). The chip (425 mg) was crushed and sieved to a grain size of 112 - 425 µm. The sieved grains were separated into three fractions (1. highly magnetic fraction, 2. less magnetic fraction, and 3. non-magnetic fraction) by hand magnet and then several grains from fraction 1 and 2 were handpicked under a binocular microscope (labeled as KN-Ol1 and KN-Ol2, respectively). Backscattered electron images of representative grains from KN-Ol1 and KN-Ol2 fractions are shown in Fig. EA1-1. Some of KN-Ol2 grains show chemical zoning and darker areas with tiny grains of metal (Fig. EA1-1d) tend to be MgO-rich compositions, suggesting that these olivine grains were reduced by carbonaceous matrix (Berkley et al., 1976; Goodrich, 1992). The apparent chemically zoned grains like KN-Ol2-7 (Fig. EA1-1d) were not used for electron probe, LA-MC-ICP-MS, and SIMS analyses.



Fig. EA1-1. Backscattered electron images of olivine grains from the Kenna meteorite. (a, b) from the non-magnetic fraction KN-Ol1. (c, d) from the less magnetic fraction KN-Ol2.

- 4. Peridotite xenoliths from eastern North China craton, China (CL09-08, 19, and 33-Ol) Three different peridotite chips were provided by Dr. Shichun Huang (University of Nevada Las Vegas). The Mg isotope ratios of olivine from three peridotite samples were obtained by solution-MC-ICP-MS measurements (Xiao et al., 2013).
- Springwater, pallasite meteorite (SW-Ol)
 Springwater is a pallasite meteorite that was found in 1931, Saskatchewan, Canada.
 Olivine grains (< 15 mg, sample No. USNM2566) provided by National Museum of
 Natural History (Smithsonian Institute), were previously separated from the host rock.
- 6. ALHA77257, ureilite meteorite (A77257-Ol)

ALHA77257 is an ureilite meteorite that was collected by the U.S.-Japan joint team in the 1977 and 1978 sessions in Antarctica. Two olivine grains provided by Dr. Katsuyuki Yamashita (Department of Earth Science, Okayama University), were previously separated from the host rock. Backscattered electron images of the two grains are shown in Fig. EA1-2a and 2b. One of the grain (A77257-OI5) is chemically zoned (see Fig. EA1-2b), which is similar to the feature found in olivine grains from the Kenna meteorite (see Fig. EA1-1d). The other grain (A77257-OI4) shows chemically homogeneous compositions in terms of MgO contents (Fo = 85.63 ± 0.13 , 1SD). Both two grains were used for electron probe, LA-MC-ICP-MS, and SIMS analyses. Regarding the A77257-OI5 grain, center part without chemically zoned MgO-rich parts was only used for these analyses. The Fo content of the center part of the A77257-OI5 grain is determined to be 85.77 ± 0.19 (1SD), which is identical to that of the A77257-Ol4 grain within uncertainties.



Fig. EA1-2a, b. Backscattered electron images of olivine grains from the ALHA77257 meteorite.

7. San Carlos, USA (SC-Ol)

Sample No. M-81-43-4, Gila Co. Arizona, USA, purchased from a private mineral collector. A large (1-2 cm) single olivine crystal was broken to fragments for SIMS standard (Kita et al., 1988; 2010). A small aliquot of fragments was used for solution ICP-MS analyses of Mg isotope ratios.

8. Mantle xenolith, Kilbourne Hole, USA (UWOL-1)

A peridotite from the Kilbourne Hole was provided by Dr. John Valley (University of Wisconsin-Madison). The peridotite was crushed and then olivine grains were separated from a host rock and hand-picked from the 300-500 µm fraction (sample name UWOL-1). All processes were done by Dr. Ilya Bindeman (University of Oregon).

Ichinome-gata spinel lherzolite nodule, Japan (IG-Ol)
 These olivine grains were hand-picked from the 500-1000 μm fraction processed by Kita et al. (1998), using a specimen GSJ R57877 that was originally collected by Dr. Shigeko

Togashi (GSJ-AIST). Fe and Ni concentrations were measured by atomic absorption analysis (Kita et al., 1998). Fe, Co, and Ni concentrations were also obtained by INAA (Kita et al., 1998). A small aliquot of the separate was used for solution MC-ICP-MS analyses of Mg isotope ratios.

10. Harrat al Kishb mantle xenolith, Saudi Arabia (HaK-Ol)

Petrology of the Harrat al Kishb mantle xenolith was studied by McGuire 1987. Olivine grains from the mantle xenolith (Sample name H30-82-1) were provided by Dr. Brian Beard (University of Wisconsin-Madison). Three olivine grains (30-47 mg per each) were crushed separately and then 27 grains were hand-picked from each crushed fraction (9 grains per each fraction).

11. Weltevreden komatiite, Barberton, South Africa (WK-Ol)

Olivine grains from the Weltevreden komatiite (sample No. 1521) were provided by Dr. Alexander Sobolev (ISTerre Institute of Earth Science). Backscattered electron images of representative grains are shown in Fig. EA1-3. Olivine grains are slightly chemically zoned, especially a rim of these grains. Center parts of the grains were only used for electron probe, LA-MC-ICP-MS, and SIMS analyses.



Fig. EA1-3a-d. Backscattered electron images of olivine grains from the Weltevreden komatiite.

12. Winona, winonaite meteorite (WN-Ol)

Winona is a winonaite meteorite that was found in 1928, Arizona, USA. A chip (1.4 g, sample No. USNM854) of the meteorite was provided by National Museum of Natural History (Smithsonian Institute). Since sizes of olivine grains are relatively small, several chips of the Winona meteorite were mounted in an epoxy round and then polished to obtain flat surfaces. After polishing, several chips were extracted from the epoxy round and were mounted again with the other olivine samples.

13. NWA7325 ungrouped achondrite (N7325-Ol)

NWA7325 is an ungrouped achondrite that was found in 2012, Morocco. Comprehensive study of this meteorite (e.g., Petrology, trace elements, oxygen, chromium and titanium isotopes, and mid-IR spectroscopy) was carried out by Goodrich et al. (2017). In this study, we used an olivine grain (N7325-O17) on a thick section that was studied by Goodrich et al. (2017). The oxygen isotope ratios of the grain are determined as $\delta^{18}O_{VSMOW} = 7.45$, $\delta^{17}O_{VSMOW} = 3.07$, and $\Delta^{17}O = -0.8\%$ by Goodrich et al. (2017; see supporting information).

14. Marble rock from Isle of Skye, Scotland (SK-Ol)

We do not have information about the exact sampling location of the marble rock. However, it is certain that the rock from the Beinn an Dubhaich aureole, Isle of Skye, Scotland.

15. Synthetic forsterite (HN-Ol)

The synthetic pure forsterite (sample name Fo16) was produced by Dr. Hiroko Nagahara (The University of Tokyo), which was previously used as oxygen isotope standard (Kita et al. 2010). An aliquot containing small fragments of the forsterite was used for the solution MC-ICP-MS analyses of Mg isotope ratios.

Pyroxene Reference Materials

 Enstatite, Mg-rich gneisses, Adirondack Mountains, New York, USA (Sp79-11 En) Petrology and chemistry of the host rock (Sample No. SP79-11) are described in Lamb and Valley (1988). The specific specimen was selected due to the low FeO contents in enstatite reported in the paper. Enstatite was hand-picked from the host rock for SIMS oxygen isotope standard by Kita et al. (2010). An aliquot of the same separate was used for the solution MC-ICP-MS analyses of Mg isotope ratios.

2. JE enstatite (JE En).

Enstatite separate from John Eiler (Eiler et al. 1997).

 Orthopyroxene and clinopyroxene, Ichinome-gata spinel lherzolite nodule, Japan (IG-Opx and IG-Cpx)

Orthopyroxene and clinopyroxene separates were originally prepared by Kita et al. (1998) from the same 500-1000 µm fraction of Ichinome-gata nodule (GSJ R57877, spinel lherzolite) that was used to extract olivine standard IG-OI. Orthopyroxene separate was analyzed by INAA and AAA for Fr, Ni, and Co concentrations (Kita et al. 1998). Both pyroxene standards were analyzed for oxygen isotopes to be used as SIMS standard (Kita et al. 2010). Aliquot from each separate was used for the solution MC-ICP-MS analyses of Mg isotope ratios.

4. Diopside from diopside marble, Adirondack mountains, New York, USA (95AK-6 Di). The diopside grains are among seven samples studied by Edwards and Valley (1998). Sample number 6 was selected as SIMS oxygen isotope standard (Kita et al. 2010) because of their internally homogeneous oxygen isotope ratios. Aliquot of the diopside separate was used for the solution MC-ICP-MS analyses of Mg isotope ratios.

Laser fluorination oxygen isotope analyses

In order to use 4 olivine RMs (FJ-Ol, SW-Ol, UWOL-1, and HaK-Ol) as SIMS calibration standards for future oxygen isotope analyses, $\delta^{18}O_{VSMOW}$ values of 4 olivine RMs

were obtained with a laser fluorination and a gas-source mass spectrometer at University of Wisconsin-Madison. Analytical procedures are described in Valley et al. (1995). The UWG-2 garnet was used as a running standard. External reproducibility of δ^{18} O is $\pm 0.11\%$ (2SD).

The results of laser fluorination analyses of 4 olivine RMs (FJ-Ol, SW-Ol, UWOL-1 and HaK-Ol) are shown in Table EA1. $\delta^{18}O_{VSMOW}$ values of 4 olivine RMs (FJ-Ol, SW-Ol, UWOL-1 and HaK-Ol) are determined to be 4.89‰, 3.51‰, 5.48‰, and 5.25‰, respectively. $\delta^{18}O_{VSMOW}$ values of 4 olivine RMs (SC-Ol, HN-Ol, IG-Ol, and OR-Ol) and all pyroxene RMs obtained by previous laser fluorination studies (Eiler et al., 1997; Edwards and Valley, 1998; Kita et al., 2010) are also shown in Table EA1.

Table EA1 Laser fluorination oxygen isotope ratios of olivine and pyroxene RMs used in this study.

Sample name	Fo or En content ^a	$\delta^{18}O_{VSMOW}$ (‰)	References ^b
Olivine			
OR-Ol	59.3	5.88	[3]
FJ-Ol	73.4	4.89	This study
KN-Ol	78.0		
CL09-19-Ol	79.4		
CL09-08-O1	80.5		
SW-Ol	81.8	3.51	This study
A77257-Ol	85.7		
CL09-33-Ol	86.6		
SC-Ol	88.8	5.32	[3]
UWOL-1	89.2	5.48	This study
IG-Ol	89.6	5.23	[3]
HaK-Ol	91.9	5.25	This study
WK-Ol	94.3		
WN-Ol	95.5		
N7325-Ol	97.4		
SK-Ol	99.8		
HN-Ol	100.0	8.90	[3]
Pyroxene			
95AK-6 Di	48.6	24.14	[2]
IG-Cpx	48.6	5.15	[3]
JE Ēn	85.5	9.20	[1]
IG-Opx	88.9	5.89	[3]
Sp79-11 En	96.3	13.31	[3]

^a Molar % of forsterite or enstatite end-members based on EPMA analyses performed in this study

^b Data sources [1] Eiler et al. (1997); [2] Edwards and Valley (1998); [3] Kita et al. (2010)

References

- Berkley, J.L., Brown IV, H.G., Keil, K., Carter, N.L., Mercier, J.C.C., Huss, G., 1976. The Kenna ureilite: an ultramafic rock with evidence for igneous, metamorphic, and shock origin. Geochim. Cosmochim. Acta 40.
- Edwards, K.J., Valley, J.W., 1998. Oxygen isotope diffusion and zoning in diopside: The importance of water fugacity during cooling. Geochim. Cosmochim. Acta 62, 2265–2277.
- Eiler, J.M., Graham, C., Valley, J.W., 1997. SIMS analysis of oxygen isotopes: Matrix effects in complex minerals and glasses. Chem. Geol. 138, 221–244.
- Goodrich, C.A., 1992. Invited Review Ureilites : A critical review. Meteoritics 27, 327-352.
- Goodrich, C.A., Kita, N.T., Yin, Q.Z., Sanborn, M.E., Williams, C.D., Nakashima, D., Lane,
 M.D., Boyle, S., 2017. Petrogenesis and provenance of ungrouped achondrite Northwest
 Africa 7325 from petrology, trace elements, oxygen, chromium and titanium isotopes, and
 mid-IR spectroscopy. Geochim. Cosmochim. Acta 203, 381–403.
- Kita, N.T., Nagahara, H., Tachibana, S., Tomomura, S., Spicuzza, M.J., Fournelle, J.H., Valley, J.W., 2010. High precision SIMS oxygen three isotope study of chondrules in LL3 chondrites: Role of ambient gas during chondrule formation. Geochim. Cosmochim. Acta 74, 6610–6635.
- Kita, N.T., Togashi, S., Morishita, Y., Terashima, S., Yurimoto, H., 1998. Search for ⁶⁰Ni excesses in MET-78008 Ureilite: An ion microprobe study. Antarct. Meteor. Res. 11, 103– 121.
- Lamb, W.M., Valley, J.W., 1988. Granulite facies amphibole and biotite equilibria, and calculated peak-metamorphic water activities. Contrib. to Mineral. Petrol. 100, 349–360.
- Mcguire, A.V., 1988. Petrology of mantle xenoliths from harrat al kishb: The mantle beneath Western Saudi Arabia. J. Petrol. 29, 73–92.
- Mikoshiba, M.U., Kanisawa, S., Matsuhisa, Y., Togashi, S., 2004. Geochemical and isotopic characteristics of the Cretaceous Orikabe Plutonic Complex, Kitakami Mountains, Japan:

Magmatic evolution in a zoned pluton and significance of a subduction-related mafic parental magma. Contrib. to Mineral. Petrol. 146, 433–449.

- Valley, J.W., Kitchen, N., Kohn, M.J., Niendorf, C.R., Spicuzza, M.J., 1995. UWG-2, a garnet standard for oxygen isotope ratios: Strategies for high precision and accuracy with laser heating. Geochim. Cosmochim. Acta 59, 5223–5231.
- Xiao, Y., Teng, F.Z., Zhang, H.F., Yang, W., 2013. Large magnesium isotope fractionation in peridotite xenoliths from eastern North China craton: Product of melt-rock interaction. Geochim. Cosmochim. Acta 115, 241–261.

Appendix EA2: typical mass spectrum



(Fig. EA2-1)

A typical mass spectrum of a running standard SC-OI (Fo₈₈) collected during a session FI3 showing peaks of ²⁴Mg⁺, ²⁵Mg⁺, ²⁶Mg⁺, and ²⁷Al⁺. Expected interference peaks (⁴⁸Ca⁺⁺ and ²⁴MgH⁺) are also shown as dotted lines. The possible contribution of both interferences to the signal intensities (²⁴Mg⁺ and ²⁵Mg⁺) is less than 10⁻⁵ at the center of these peaks.

Sample name	EPMA session	No. of analyzed grains	No. of EPMA analyses	Fo ^a or En ^b content
Olivine				
OR-Ol	2018. Aug.	20	103	59.27
FJ-Ol	2018. Aug.	20	99	73.36
KN-Ol	2018. Aug.	22	102	77.97
CL09-19-O1	2019 Jan.	18	18	79.44
CL09-08-O1	2019 Jan.	19	19	80.53
SW-Ol	2018. Aug.	22	109	81.81
A77257-Ol	2019 Jan.	2	10	85.70
CL09-33-O1	2019 Jan.	19	19	86.63
SC-Ol	2019 Jan.	1	10	88.82
UWOL-1	2018. Aug.	42	213	89.23
IG-Ol	2019 Jan.	8	12	89.58
HaK-Ol	2018. Aug.	27	134	91.85
WK-Ol	2018. Aug.	22	101	94.30
WN-Ol	2018. Aug.	10	67	95.52
N7325-Ol	2019 Jan.	1	5	97.42
SK-Ol	2019 Jan.	5	21	99.81
HN-O1	2019 Jan.	1	5	100.00
Pyroxene				
95AK-6	2019 Sep.	8	40	48.55
IG-Cpx	2019 Sep.	6	30	48.60
JE En	2019 Sep.	6	30	85.53
IG-Opx	2019 Sep.	5	25	88.89
Sp79-11 En	2019 Sep.	7	35	96.31

Table EA4-1 Summary of homogeneities of Fo, En, and δ^{25} Mg values in 17 olivine and 5 pyro:

^a Fo = Mg/[Mg + Fe] × 100 (molar %) ^b En = Mg/[Mg + Fe + Ca] × 100 (molar %)

^{a, b} Fo and En contents are calculated by considering total iron as ferrous iron.

^c Standard deviations (Fo or En units) were calclated from averaged Fo or En cont Regarding three olivine RMs (SC-Ol, N7325-Ol, and HN-Ol), each one grain w

^b FeO wt% of all 5 analyses are below detection limit.

^c Four individual grains on 4 individual SIMS mounts were measured.

References; [1] Ushikubo et al. (2013).GCA, [2] Ushikubo et al. (2017).GCA.

xene R	RMs
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1SD ^c	SIMS session i.d.	No. of grains	No. of SIMS analyses	$\frac{\text{SIMS 2SD}}{(\delta^{25}\text{Mg}_m)}$	Ref.
0.74	FI1	10	10	0.12	This study
0.29	FI1	8	8	0.08	This study
0.14	FI1	13	15	0.13	This study
0.24	FI3	8	8	0.16	This study
0.62	FI3	8	8	0.19	This study
0.10	FI1	8	8	0.06	This study
0.10	FI3	2	6	0.23	This study
0.33	FI3	8	8	0.12	This study
0.11	FI3	4 ^c	62	0.14	This study
0.08	FI3	11	15	0.15	This study
0.09	FI3	8	8	0.05	This study
0.09	FI1	8	8	0.10	This study
0.52	FI1	8	8	0.12	This study
0.47	FI3	4	8	0.17	This study
0.08	FI3	1	4	0.14	This study
0.02	FI3	3	6	0.07	This study
b	FI3	2	6	0.08	This study
0.19	UI2	4	4	0.18	[2]
0.58	UI2	4	4	0.09	[2]
0.15	UI1	4	4	0.06	[1]
0.12	UI1	4	4	0.23	[1]
0.56	UI1	4	11	0.14	[1]

tents of each grain. as measured so that 1SD for three RMs represent intra-grain homogeneity.

Data	sossion i d	analyzaia		Primary
Date	SCSSIOII I.U.	anarysis	source	species
2018. Sept.	FI1	Mg isotopes	Radio frequency	$^{16}O_2^{-}$
2018. Sept.	FI2	Mg isotopes	Radio frequency	$^{16}O^{-}$
2019. Jan.	FI3	Mg isotopes	Radio frequency	$^{16}O_2^{-}$
2019. Apr.	FI4	Mg isotopes	Radio frequency	$^{16}O^{-}$
2019. Jan.	FE1	Major & minor elements	Radio frequency	$^{16}O_2^{-}$
2019. Apr.	FE2	Major & minor elements	Radio frequency	$^{16}O^{-}$
2009. Sep.	UI1	Mg isotopes	Duoplasmatron	$^{16}O^{-}$
2013. Apr.	UI2	Mg isotopes	Duoplasmatron	¹⁶ O ⁻

Table EA5-1 List of analytical sessions for SIMS data used in this study

^a in diameter
^b FC = faraday cup, EM = electron multiplier
^c O1 = olivine, Px = pyroxene
References; [1] Ushikubo et al. (2013); [2] Ushikubo et al. (2017).

ion		1-4b	Measured	Dof
intensity	beam size ^a	detection	mineral ^c	Kel.
1 nA	7.5 μm	Multi-FC	01	This study
2.4 nA	9 μm	Multi-FC	Ol	This study
1 nA	7.5 μm	Multi-FC	Ol & Px	This study
2.6 nA	9 µm	Multi-FC	Ol	This study
6 pA	1.5 µm	Mono-EM or FC	Ol	This study
16 pA	2 µm	Mono-EM or FC	Ol	This study
4.5 nA	15 μm	Multi-FC	Ol & Px	[1]
2.2 nA	10 µm	Multi-FC	Px	[2]

Appendix EA6-1: SIMS Mg isotope data (session FI1, 1nA, O_2^- prima

^a 2 σ error (A) = $\sqrt{(2SE_{SIMS})^2 + (2SD_{ICPMS})^2}$ ^b 2 σ error = $\sqrt{(A)^2 + (2SD_{SIMS_bracket})^2}$ Erros of averaged values are 2SD.

Filo	Analysis anot	Cycle#	Cycle#
FIIE	Analysis spot	(measured)	(used)
20180904@18.asc	WI-STD-91 SCOL	30	30
20180904@19.asc	WI-STD-91 SCOL	30	30
20180904@20.asc	WI-STD-91 SCOL	30	30
20180904@21.asc	WI-STD-91 SCOL	30	30
20180904@30.asc	WI-STD-91 SCOL	30	30
20180904@31.asc	WI-STD-91 SCOL	30	30
20180904@32.asc	WI-STD-91 SCOL	30	30
20180904@33.asc	WI-STD-91 SCOL	30	30
	Average (SC-OI)		
20180904@8.asc	WI-STD-91 KNOL1-2_02_1	30	30
20180904@9.asc	WI-STD-91 KNOI1-4	30	30
20180904@10.asc	WI-STD-91 KNOI1-8	30	30
20180904@11.asc	WI-STD-91 KNOI1-5	30	30
20180904@12.asc	WI-STD-91 KNOI1-9	30	30
20180904@13.asc	WI-STD-91 KNOL1-10_02_1	30	30
20180904@14.asc	WI-STD-91 KNOL1-12	30	30
20180904@15.asc	WI-STD-91 KNOI1-14	30	30
20180904@97.asc	WI-STD-91 KNOL2-6	30	30
20180904@98.asc	WI-STD-91 KNOL2-7	30	30
20180904@99.asc	WI-STD-91 KNOL2-8	30	30
20180904@100.asc	WI-STD-91 KNOL2-9_02_1	30	30
20180904@102.asc	WI-STD-91 KNOL2-5_O2_1	30	30
20180904@104.asc	WI-STD-91 KNOL1-10_02_2	30	30
20180904@105.asc	WI-STD-91 KNOL1-10_02_3	30	30
	Average (KN-OI)		
20180904@22.asc	WI-STD-91 WKOL-1_1	30	30
20180904@23.asc	WI-STD-91 WKOL-3	30	30
20180904@24.asc	WI-STD-91 WKOL-5_02_1	30	30
20180904@25.asc	WI-STD-91 WKOL-8	30	30
20180904@26.asc	WI-STD-91 WKOL-10_02_1	30	30
20180904@27.asc	WI-STD-91 WKOL-11	30	30
20180904@28.asc	WI-STD-91 WKOL-12	30	30
20180904@29.asc	WI-STD-91 WKOL-17	30	30
	Average (WK-OI)		
20180904@37.asc	WI-STD-91 WNOL-1_02_2	30	30

20180904@38.asc	WI-STD-91 WNOL-1_02_3	30	30
20180904@39.asc	WI-STD-91 WNOL-1_02_4	30	30
20180904@40.asc	WI-STD-91 WNOL-1 02 5	30	30
20180904@41.asc	WI-STD-91 WNOL-1_02_6	30	30
20180904@42.asc	WI-STD-91 WNOL-1_02_7	30	30
20180904@43.asc	WI-STD-91 WNOL-1_02_8	30	30
20180904@123.asc	WI-STD-91 WNOL-4_02_1	30	30
20180904@124.asc	WI-STD-91 WNOL-4_02_2	30	30
	Average (WN-OI)		
20180904@54.asc	WI-STD-91 FJOL-2_02_1	30	30
20180904@55.asc	WI-STD-91 FJOL-3_02_1	30	30
20180904@56.asc	WI-STD-91 FJOL-4	30	30
20180904@57.asc	WI-STD-91 FJOL-5	30	30
20180904@58.asc	WI-STD-91 FJOL-6	30	30
20180904@59.asc	WI-STD-91 FJOL-10	30	30
20180904@60.asc	WI-STD-91 FJOL-9	30	30
20180904@61.asc	WI-STD-91 FJOL-11	30	30
	Average (FJ-OI)		
20180904@66.asc	WI-STD-91 HaK-g1-2_O2_1	30	30
20180904@67.asc	WI-STD-91 HaK-g1-3_O2_1	30	30
20180904@68.asc	WI-STD-91 HaK-g1_5	30	30
20180904@69.asc	WI-STD-91 HaK-g1_8	30	30
20180904@70.asc	WI-STD-91 HaK-g2_1	30	30
20180904@71.asc	WI-STD-91 HaK-g2_2	30	30
20180904@72.asc	WI-STD-91 HaK-g3_1	30	30
20180904@73.asc	WI-STD-91 HaK-g3_3	30	30
	Average (HaK-OI)		
20180904@16.asc	WI-STD-91 HNOL	30	30
20180904@34.asc	WI-STD-91 HNOL	30	30
20180904@74.asc	WI-STD-91 HNOL	30	30
20180904@75.asc	WI-STD-91 HNOL	30	30
20180904@76.asc	WI-STD-91 HNOL	30	30
20180904@77.asc	WI-STD-91 HNOL	30	30
20180904@130.asc	WI-STD-91 HNOL	30	30
	Average (HN-OI)		
20180904@81.asc	WI-STD-91 SWOL-1_02_1	30	30
20180904@82.asc	WI-STD-91 SWOL-3_02_1	30	30
20180904@83.asc	WI-STD-91 SWOL-4	30	30
20180904@84.asc	WI-STD-91 SWOL-5	30	30
20180904@85.asc	WI-STD-91 SWOL-8	30	30
20180904@86.asc	WI-STD-91 SWOL-12	30	30
20180904@87.asc	WI-STD-91 SWOL-15_02_1	30	30
20180904@88.asc	WI-STD-91 SWOL-16	30	30
	Average (SW-OI)		
20180904@89.asc	WI-STD-91 A77257-5_O2_1	30	30
20180904@90.asc	WI-STD-91 A77257-5_02_2	30	30
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20180904@92.asc	WI-STD-91 A77257-5_O2_4	30	30
	Average (A77257-OI)		
20180904@110.asc	WI-STD-91 UWOL1-42_O2_1	30	30
20180904@111.asc	WI-STD-91 UWOL1-41_O2_1	30	30
20180904@112.asc	WI-STD-91 UWOL1-5	30	30
20180904@113.asc	WI-STD-91 UWOL1-6	30	30
	Average (UW-OI)		
20180904@17.asc	WI-STD-91 OROL-1	30	30
20180904@114.asc	WI-STD-91 OROL-2_02_1	30	30
20180904@115.asc	WI-STD-91 OROL-3_02_1	30	30
20180904@116.asc	WI-STD-91 OROL-4_02_1	30	30
20180904@117.asc	WI-STD-91 OROL-6_02_1	30	30
20180904@118.asc	WI-STD-91 OROL-12	30	30
20180904@119.asc	WI-STD-91 OROL-17_02_1	30	30
20180904@120.asc	WI-STD-91 OROL-18	30	30
20180904@121.asc	WI-STD-91 OROL-19	30	30
20180904@122.asc	WI-STD-91 OROL-20	30	30
	Average (OR-OI)		

ry ion beam)

		Measure	d			
24Mg	Yield (10 ⁸	d25Mg-m		d26Mg-m		d26Mg*-m
(100M cps)	cps/nA)	‰	25E ‰	‰	25E %	‰
1.80	1.93	-2.61	0.06	-4.78	0.11	0.31
1.80	1.92	-2.58	0.07	-4.68	0.12	0.34
1.81	1.94	-2.51	0.06	-4.58	0.10	0.32
1.81	1.94	-2.59	0.06	-4.71	0.11	0.34
1.79	1.93	-2.56	0.06	-4.63	0.12	0.37
1.79	1.93	-2.60	0.06	-4.66	0.11	0.40
1.78	1.92	-2.68	0.06	-4.76	0.11	0.47
1.79	1.94	-2.68	0.06	-4.83	0.10	0.39
		-2.60	0.11	-4.70	0.17	0.37
1.80	1.92	-1.47	0.04	-2.54	0.04	0.32
1.80	1.92	-1.50	0.03	-2.59	0.03	0.33
1.78	1.90	-1.61	0.04	-2.79	0.03	0.35
1.79	1.91	-1.57	0.03	-2.73	0.03	0.32
1.78	1.91	-1.48	0.03	-2.56	0.04	0.31
1.78	1.91	-1.58	0.03	-2.76	0.03	0.32
1.78	1.91	-1.55	0.04	-2.69	0.04	0.32
1.78	1.91	-1.56	0.02	-2.67	0.03	0.38
1.81	1.95	-1.47	0.03	-2.45	0.03	0.41
1.81	1.95	-1.55	0.03	-2.62	0.04	0.40
1.81	1.95	-1.50	0.03	-2.47	0.04	0.45
1.82	1.96	-1.59	0.03	-2.65	0.03	0.46
1.83	1.97	-1.56	0.03	-2.60	0.04	0.44
1.79	1.93	-1.69	0.04	-2.80	0.04	0.50
1.81	1.95	-1.61	0.03	-2.74	0.03	0.41
		-1.55	0.13	-2.64	0.22	0.38
1.82	1.96	-2.69	0.06	-4.84	0.11	0.41
1.82	1.96	-2.57	0.05	-4.70	0.11	0.31
1.81	1.95	-2.66	0.07	-4.92	0.11	0.27
1.82	1.96	-2.61	0.05	-4.68	0.10	0.41
1.81	1.96	-2.75	0.06	-4.94	0.11	0.41
1.81	1.95	-2.74	0.06	-4.94	0.11	0.40
1.81	1.93	-2.72	0.06	-4.90	0.10	0.40
1.81	1.95	-2.71	0.06	-4.89	0.11	0.40
		-2.68	0.12	-4.85	0.21	0.38
1.78	1.96	-3.05	0.06	-5.54	0.10	0.41

1	.77	1.95	-3.00	0.06	-5.42	0.12	0.43
1	.76	1.93	-3.16	0.05	-5.74	0.12	0.42
1	.78	1.96	-3.01	0.06	-5.43	0.12	0.43
1	.78	1.96	-3.08	0.06	-5.55	0.11	0.44
1	.77	1.95	-3.19	0.07	-5.75	0.13	0.45
1	.77	1.95	-2.97	0.06	-5.37	0.11	0.41
1	.81	1.95	-2.94	0.06	-5.32	0.13	0.40
1	.80	1.95	-3.01	0.06	-5.39	0.12	0.49
			-3.05	0.17	-5.50	0.32	0.43
1	.76	1.90	-1.64	0.04	-2.80	0.04	0.40
1	.76	1.89	-1.74	0.03	-2.96	0.04	0.43
1	.77	1.91	-1.67	0.03	-2.87	0.03	0.39
1	.77	1.91	-1.62	0.03	-2.72	0.04	0.43
1	.75	1.89	-1.62	0.03	-2.75	0.03	0.41
1	.77	1.91	-1.69	0.03	-2.86	0.03	0.43
1	.76	1.90	-1.68	0.03	-2.82	0.04	0.46
1	.77	1.91	-1.63	0.03	-2.86	0.03	0.32
			-1.66	0.08	-2.83	0.15	0.41
1	.83	1.97	-2.93	0.08	-5.21	0.15	0.50
1	.83	1.97	-2.90	0.07	-5.20	0.15	0.45
1	.83	1.98	-2.87	0.08	-5.11	0.15	0.48
1	.83	1.97	-2.90	0.09	-5.13	0.15	0.53
1	.83	1.97	-2.99	0.07	-5.28	0.16	0.55
1	.83	1.98	-2.87	0.07	-5.14	0.15	0.45
1	.82	1.97	-2.99	0.08	-5.31	0.15	0.51
1	.83	1.97	-2.97	0.08	-5.26	0.15	0.52
			-2.93	0.10	-5.21	0.15	0.50
1	.71	1.84	-4.25	0.06	-7.82	0.11	0.45
1	.70	1.84	-4.32	0.05	-7.94	0.11	0.47
1	.73	1.87	-4.28	0.07	-7.75	0.13	0.58
1	.73	1.87	-4.29	0.07	-7.80	0.13	0.55
1	.73	1.86	-4.28	0.06	-7.78	0.11	0.55
1	.73	1.86	-4.31	0.07	-7.90	0.13	0.49
1	.74	1.88	-4.27	0.07	-7.84	0.13	0.48
		0.00	-4.29	0.05	-7.83	0.14	0.51
1	.88	2.03	-1.36	0.03	-2.23	0.04	0.41
1	.89	2.04	-1.34	0.03	-2.23	0.04	0.39
1	.88	2.03	-1.34	0.03	-2.22	0.04	0.39
1	.90	2.05	-1.30	0.03	-2.10	0.04	0.44
1	.89	2.04	-1.35	0.02	-2.17	0.03	0.46
1	.90	2.04	-1.31	0.03	-2.13	0.03	0.42
1	.91	2.06	-1.34	0.03	-2.18	0.03	0.44
1	.89	2.04	-1.28	0.03	-2.12	0.03	0.37
		0.04	-1.33	0.06	-2.18	<u>U.11</u>	0.41
1	.89	2.04	-1.11	0.03	-1./5	0.03	0.42

1.90	2.05	-1.14	0.03	-1.82	0.03	0.39
1.89	2.04	-1.11	0.03	-1.73	0.04	0.45
		-1.12	0.03	-1.77	0.10	0.42
1.82	1.96	-2.66	0.07	-4.79	0.12	0.40
1.83	1.97	-2.65	0.06	-4.77	0.12	0.40
1.83	1.97	-2.60	0.06	-4.67	0.11	0.40
 1.83	1.97	-2.63	0.07	-4.72	0.12	0.39
		-2.64	0.05	-4.74	0.10	0.40
1.53	1.64	-1.87	0.03	-3.29	0.04	0.36
1.54	1.66	-1.89	0.02	-3.36	0.04	0.32
1.55	1.67	-1.84	0.03	-3.26	0.04	0.32
1.56	1.68	-1.78	0.04	-3.08	0.04	0.39
1.53	1.65	-1.82	0.03	-3.19	0.03	0.35
1.51	1.63	-1.82	0.04	-3.21	0.05	0.34
1.51	1.63	-1.77	0.04	-3.14	0.03	0.31
1.51	1.63	-1.94	0.04	-3.46	0.04	0.33
1.52	1.63	-1.77	0.04	-3.20	0.04	0.26
 1.55	1.67	-1.75	0.03	-3.18	0.04	0.24
		-1.83	0.12	-3.24	0.22	0.32

2SE ‰	f ₂₅ *	2σ ^a	bias* _{25(RM-SCOI)}	2σ ^b	f 26 *	$2\sigma^{a}$
0.06	-2.54	0.11			-4.63	0.19
0.07	-2.50	0.11			-4.53	0.20
0.06	-2.44	0.11			-4.43	0.19
0.06	-2.52	0.11			-4.56	0.19
0.06	-2.49	0.11			-4.48	0.20
0.05	-2.53	0.11			-4.51	0.19
0.06	-2.61	0.11			-4.61	0.19
0.06	-2.61	0.11			-4.68	0.19
0.11	-2.53	0.11			-4.55	0.17
0.06	-1.51	0.18	1.03	0.18	-2.53	0.30
0.06	-1.54	0.18	0.99	0.18	-2.58	0.30
0.08	-1.65	0.18	0.88	0.18	-2.78	0.30
0.05	-1.61	0.18	0.93	0.18	-2.72	0.30
0.06	-1.52	0.18	1.02	0.18	-2.55	0.30
0.05	-1.62	0.18	0.91	0.18	-2.75	0.30
0.07	-1.59	0.18	0.94	0.18	-2.68	0.30
0.05	-1.60	0.18	0.93	0.18	-2.66	0.30
0.05	-1.51	0.18	1.02	0.18	-2.44	0.30
0.06	-1.59	0.18	0.94	0.18	-2.61	0.30
0.07	-1.54	0.18	0.99	0.18	-2.46	0.30
0.05	-1.63	0.18	0.90	0.18	-2.64	0.30
0.05	-1.60	0.18	0.93	0.18	-2.59	0.30
0.06	-1.73	0.18	0.80	0.18	-2.79	0.30
0.06	-1.65	0.18	0.88	0.18	-2.73	0.30
0.12	-1.59	0.13	0.94	0.13	-2.63	0.22
0.07	-2.60	0.14	-0.07	0.15	-4.67	0.22
0.06	-2.49	0.14	0.04	0.15	-4.53	0.22
0.07	-2.58	0.14	-0.05	0.15	-4.75	0.23
0.06	-2.53	0.14	0.00	0.15	-4.52	0.22
0.07	-2.66	0.14	-0.13	0.15	-4.77	0.22
0.06	-2.65	0.14	-0.12	0.15	-4.78	0.22
0.07	-2.63	0.14	-0.10	0.15	-4.74	0.22
0.07	-2.63	0.14	-0.09	0.15	-4.72	0.22
0.11	-2.60	0.12	-0.07	0.12	-4.69	0.21
0.06	-3.02	0.10	-0 45	0.13	-5 40	0 19

Corrected

			• • •	A 1 -		• • •
0.07	-2.97	0.10	-0.40	0.13	-5.27	0.20
0.07	-3.13	0.10	-0.56	0.13	-5.60	0.20
0.06	-2.97	0.10	-0.41	0.13	-5.28	0.20
0.06	-3.04	0.10	-0.47	0.13	-5.41	0.19
0.04	-3.15	0.10	-0.58	0.13	-5.61	0.21
0.06	-2.93	0.10	-0.37	0.13	-5.23	0.19
0.07	-2.90	0.10	-0.40	0.13	-5.18	0.21
0.07	-2.98	0.10	-0.47	0.13	-5.24	0.20
0.05	-3.01	0.17	-0.46	0.15	-5.36	0.32
0.08	-1.57	0.20	0.91	0.20	-2.71	0.30
0.06	-1.67	0.20	0.81	0.20	-2.86	0.30
0.07	-1.60	0.20	0.88	0.20	-2.77	0.30
0.06	-1.55	0.20	0.94	0.20	-2.62	0.30
0.06	-1.55	0.20	0.93	0.20	-2.65	0.30
0.06	-1.62	0.20	0.87	0.20	-2.76	0.30
0.08	-1.61	0.20	0.87	0.20	-2.72	0.30
0.06	-1.56	0.20	0.92	0.20	-2.77	0.30
0.08	-1.59	0.08	0.89	0.08	-2.73	0.15
0.06	-2.87	0.15	-0.37	0.15	-5.05	0.24
0.05	-2.84	0.15	-0.34	0.15	-5.04	0.24
0.07	-2.81	0.15	-0.31	0.16	-4.95	0.24
0.07	-2.84	0.16	-0.34	0.16	-4.97	0.24
0.07	-2.93	0.15	-0.43	0.15	-5.12	0.25
0.06	-2.81	0.15	-0.31	0.15	-4.98	0.24
0.06	-2.93	0.15	-0.42	0.16	-5.15	0.24
0.05	-2.91	0.15	-0.40	0.16	-5.10	0.24
0.07	-2.87	0.10	-0.36	0.10	-5.05	0.15
0.06	-3.88	0.10	-1.36	0.13	-7.13	0.17
0.06	-3.96	0.10	-1.39	0.13	-7.25	0.17
0.05	-3.91	0.11	-1.41	0.11	-7.06	0.19
0.05	-3.92	0.11	-1.42	0.11	-7.11	0.19
0.06	-3.91	0.10	-1.41	0.11	-7.09	0.17
0.07	-3.94	0.11	-1.44	0.12	-7.20	0.19
0.07	-3.91	0.11	-1.41	0.13	-7.15	0.19
0.10	-3.92	0.05	-1.41	0.05	-7.14	0.14
0.05	-1.33	0.14	1.20	0.15	-2.20	0.16
0.07	-1.32	0.14	1.22	0.15	-2.20	0.16
0.07	-1.32	0.15	1.22	0.15	-2.19	0.16
0.05	-1.28	0.15	1.26	0.15	-2.07	0.16
0.06	-1.33	0.14	1.21	0.15	-2.14	0.16
0.06	-1.29	0.15	1.25	0.15	-2.10	0.16
0.07	-1.32	0.14	1.22	0.15	-2.15	0.16
0.06	-1.25	0.14	1.28	0.15	-2.09	0.16
0.06	-1.31	0.06	1.23	0.06	-2.14	0.11
0.06	-1.04	0.14	1.50	0.15	-1.47	0.13

0.05	-1.07	0.14	1.47	0.15	-1.54	0.13
0.06	-1.04	0.14	1.50	0.15	-1.44	0.13
0.05	-1.05	0.03	1.49	0.03	-1.48	0.10
0.06	-2.53	0.12	-0.02	0.13	-4.55	0.23
0.05	-2.51	0.12	-0.01	0.13	-4.53	0.23
0.05	-2.47	0.12	0.04	0.13	-4.43	0.23
0.06	-2.49	0.12	0.02	0.13	-4.48	0.23
0.01	-2.50	0.05	0.01	0.05	-4.50	0.10
0.06	-1.86	0.07	0.67	0.13	-3.27	0.13
0.05	-1.87	0.07	0.63	0.13	-3.33	0.13
0.06	-1.83	0.07	0.68	0.13	-3.24	0.13
0.07	-1.77	0.07	0.74	0.13	-3.06	0.13
0.06	-1.81	0.07	0.70	0.13	-3.17	0.13
0.06	-1.81	0.07	0.70	0.13	-3.19	0.14
0.07	-1.75	0.07	0.76	0.13	-3.12	0.13
0.06	-1.93	0.07	0.58	0.13	-3.44	0.13
0.09	-1.76	0.08	0.75	0.13	-3.18	0.13
0.07	-1.74	0.07	0.77	0.13	-3.16	0.13
0.09	-1.81	0.12	0.70	0.12	-3.21	0.22

bias* _{26(RM-SCOI)}	2σ ^b	

2.06	0.31
2.00	0.31
1.80	0.31
1.86	0.31
2.03	0.31
1.83	0.31
1.90	0.31
1.92	0.31
2.03	0.31
1.86	0.31
2.01	0.31
1.83	0.31
1.88	0.31
1.68	0.31
1.74	0.31
1.90	0.23
 -0.12	0.25
0.02	0.25
-0.20	0.25
0.03	0.25
-0.22	0.25
-0.23	0.25
-0.18	0.25
-0.17	0.25
-0.13	0.21
 -0.81	0.23

-0.68	0.23
-1.01	0.23
-0.69	0.23
-0.82	0.23
-1.02	0.23
-0.63	0.23
-0.73	0.25
-0.79	0.25
 -0.80	0.27
 1.72	0.31
1.56	0.31
1.65	0.31
1.80	0.31
1.77	0.31
1.66	0.31
1.70	0.31
1.65	0.31
 1.69	0.15
 -0.61	0.25
-0.60	0.25
-0.51	0.25
-0.52	0.25
-0.68	0.25
-0.54	0.25
-0.71	0.25
-0.66	0.25
 -0.60	0.15
 -2.56	0.22
-2.67	0.23
-2.62	0.20
-2.67	0.20
-2.65	0.20
-2.77	0.20
-2.70	0.25
 -2.66	0.13
 2.28	0.19
2.28	0.19
2.29	0.19
2.41	0.19
2.34	0.19
2.38	0.19
2.33	0.19
2.39	0.19
2.34	0.11
3.02	0.19

2.95	0.19
3.04	0.19
3.00	0.10
-0.09	0.25
-0.07	0.25
0.02	0.25
-0.03	0.25
-0.04	0.10
1.31	0.22
1.13	0.25
1.22	0.25
1.40	0.25
1.29	0.25
1.27	0.25
1.35	0.25
1.02	0.25
1.29	0.25
1.31	0.25
1.26	0.22

Appendix EA6-2: SIMS Mg isotope data (session FI2, 2.6nA, O⁻ prima

^a 2 σ error (A) = $\sqrt{(2SE_{SIMS})^2 + (2SD_{ICPMS})^2}$ ^b 2 σ error = $\sqrt{(A)^2 + (2SD_{SIMS_bracket})^2}$ Erros of averaged values are 2SD.

File Analysis spot		Cycle#	Cycle#
THE	Analysis spot	(measured)	(used)
20180904@132.asc	WI-STD-91 SCOL	30	30
20180904@133.asc	WI-STD-91 SCOL	30	30
20180904@134.asc	WI-STD-91 SCOL	30	30
20180904@135.asc	WI-STD-91 SCOL	30	30
20180904@146.asc	WI-STD-91 SCOL	30	30
20180904@147.asc	WI-STD-91 SCOL	30	30
20180904@148.asc	WI-STD-91 SCOL	30	30
20180904@149.asc	WI-STD-91 SCOL	30	30
	Average (SC-OI)		
20180904@136.asc	WI-STD-91 HNOL	30	30
20180904@137.asc	WI-STD-91 HNOL	30	30
	Average (HN-OI)		
20180904@138.asc	WI-STD-91 OROL-2_O_1	30	30
20180904@139.asc	WI-STD-91 OROL-3_O_1	30	30
20180904@140.asc	WI-STD-91 OROL-6_O_1	30	30
20180904@141.asc	WI-STD-91 OROL-17_0_1	30	30
	Average (OR-OI)		
20180904@142.asc	WI-STD-91 HaKOL-g1-2_O_1	30	30
20180904@143.asc	WI-STD-91 HaKOL-g1-3_O_1	30	30
	Average (HaK-OI)		
20180904@144.asc	WI-STD-91 KN1-10_O_1	30	30
20180904@145.asc	WI-STD-91 KN1-10_0_2	30	30
	Average (KN-OI)		
20180904@150.asc	WI-STD-91 WNOL-1_O_1	30	30
20180904@151.asc	WI-STD-91 WNOL-1_O_2	30	30
20180904@152.asc	WI-STD-91 WNOL-4_O_1	30	30
	Average (WN-OI)		
20180904@153.asc	WI-STD-91 WKOL-10_0_1	30	30
20180904@154.asc	WI-STD-91 WKOL-1_2	30	30
20180904@155.asc	WI-STD-91 WKOL-5_O_1	30	30
20180904@156.asc	WI-STD-91 WKOL-8_0_1	30	30
	Average (WK-OI)		
20180904@159.asc	WI-STD-91 A77257-5_O_1	30	30

20180904@160.asc	WI-STD-91 A77257-5_O_2	30	30
	Average (A77257-OI)		
20180904@161.asc	WI-STD-91 UWOL1-42_O_1	30	30
20180904@162.asc	WI-STD-91 UWOL1-41_O_1	30	30
	Average (UW-OI)		
20180904@163.asc	WI-STD-91 FJOL-2_O_1	30	30
20180904@164.asc	WI-STD-91 FJOL-3_O_1	30	30
	Average (FJ-OI)		
20180904@157.asc	WI-STD-91 SWOL-1_O_1	30	30
20180904@158.asc	WI-STD-91 SWOL-3_O_1	30	30
20180904@165.asc	WI-STD-91 SWOL-15_O_1	30	30
	Average (SW-OI)		

ary ion beam)

		Measure	d			
24Mg (100M cps)	Yield (10 ⁸ cps/nA)	d25Mg-m ‰	2SE ‰	d26Mg-m ‰	2SE ‰	d26Mg*-m ‰
1.85	0.76	-3.18	0.12	-5.69	0.24	0.51
1.85	0.76	-3.18	0.12	-5.74	0.23	0.45
1.85	0.76	-3.21	0.12	-5.80	0.23	0.45
1.85	0.76	-3.21	0.12	-5.74	0.24	0.52
1.84	0.76	-3.23	0.13	-5.85	0.24	0.45
1.84	0.76	-3.19	0.12	-5.78	0.23	0.45
1.85	0.76	-3.21	0.12	-5.74	0.24	0.50
1.85	0.76	-3.23	0.12	-5.79	0.24	0.50
		-3.21	0.04	-5.77	0.10	0.48
1.81	0.75	-4.50	0.05	-8.31	0.09	0.46
1.82	0.75	-4.49	0.06	-8.26	0.10	0.47
		-4.50	0.02	-8.29	0.06	0.47
1.68	0.69	-1.31	0.04	-2.10	0.03	0.46
1.69	0.70	-1.27	0.03	-2.07	0.03	0.40
1.68	0.69	-1.28	0.04	-2.08	0.03	0.41
1.65	0.68	-1.26	0.04	-2.07	0.05	0.38
		-1.28	0.05	-2.08	0.03	0.41
1.88	0.77	-3.33	0.14	-6.05	0.25	0.44
1.89	0.78	-3.32	0.14	-5.93	0.25	0.54
		-3.33	0.01	-5.99	0.17	0.49
1.95	0.80	-1.00	0.03	-1.55	0.03	0.39
1.95	0.80	-1.06	0.04	-1.59	0.03	0.47
		-1.03	0.08	-1.57	0.05	0.43
1.90	0.78	-3.24	0.10	-5.78	0.19	0.53
1.90	0.78	-3.34	0.10	-6.02	0.19	0.48
1.90	0.78	-3.25	0.09	-5.86	0.18	0.48
		-3.28	0.11	-5.89	0.25	0.50
1.89	0.78	-3.16	0.10	-5.60	0.20	0.54
1.90	0.78	-3.10	0.10	-5.55	0.19	0.50
1.90	0.78	-3.14	0.10	-5.56	0.20	0.56
1.90	0.78	-3.12	0.10	-5.53	0.20	0.55
		-3.13	0.05	-5.56	0.06	0.54
1.95	0.80	-1.14	0.03	-1.82	0.05	0.40

1.93	0.79	-1.28	0.06	-1.99	0.06	0.50
		-1.21	0.19	-1.91	0.24	0.45
1.85	0.76	-3.28	0.12	-5.84	0.24	0.54
1.85	0.76	-3.35	0.13	-5.89	0.25	0.63
		-3.31	0.10	-5.87	0.07	0.59
1.90	0.78	-1.11	0.03	-1.70	0.03	0.47
1.91	0.78	-1.14	0.03	-1.74	0.03	0.49
		-1.13	0.04	-1.72	0.06	0.48
1.84	0.76	-2.35	0.07	-4.09	0.12	0.48
1.85	0.76	-2.32	0.07	-4.02	0.13	0.50
1.85	0.76	-2.39	0.06	-4.12	0.11	0.54
		-2.35	0.07	-4.08	0.11	0.51

2SE ‰	f ₂₅ *	2σ ^a	bias* _{25(RM-SCOI)}	$2\sigma^{b}$	f ₂₆ *	2σ ^a
0.06	-3.11	0.15			-5.53	0.28
0.07	-3.11	0.15			-5.59	0.28
0.05	-3.14	0.15			-5.65	0.28
0.06	-3.14	0.15			-5.59	0.29
0.06	-3.16	0.15			-5.70	0.28
0.06	-3.12	0.15			-5.62	0.28
0.06	-3.13	0.15			-5.59	0.29
0.05	-3.16	0.15			-5.64	0.28
0.06	-3.13	0.04			-5.61	0.10
0.06	-4.14	0.10	-1.01	0.11	-7.62	0.16
0.07	-4.12	0.10	-0.99	0.11	-7.57	0.17
0.02	-4.13	0.02	-1.00	0.02	-7.59	0.06
0.08	-1.30	0.07	1.84	0.10	-2.08	0.13
0.06	-1.26	0.07	1.88	0.10	-2.05	0.13
0.06	-1.26	0.07	1.88	0.10	-2.05	0.13
0.08	-1.24	0.07	1.90	0.10	-2.05	0.13
0.07	-1.26	0.05	1.87	0.05	-2.06	0.03
0.06	-3.27	0.19	-0.14	0.19	-5.89	0.32
0.06	-3.26	0.19	-0.13	0.19	-5.77	0.32
0.15	-3.27	0.01	-0.13	0.01	-5.83	0.17
0.06	-1.04	0.18	2.10	0.18	-1.54	0.30
0.07	-1.09	0.18	2.05	0.18	-1.58	0.30
0.11	-1.06	0.08	2.08	80.0	-1.56	0.05
0.06	-3.20	0.13	-0.03	0.13	-5.63	0.25
0.07	-3.30	0.13	-0.13	0.14	-5.88	0.25
0.06	-3.22	0.12	-0.05	0.13	-5.72	0.24
0.05	-3.24	0.11	-0.07	0.11	-5.74	0.25
0.00						
0.06	-3.07	0.16	0.10	0.16	-5.44	0.28
0.07	-3.02	0.16	0.16	0.17	-5.38	0.27
0.06	-3.06	0.16	0.12	0.16	-5.40	0.28
0.06	-3.04	0.16	0.14	0.17	-5.37	0.28
0.05	-3.04	0.05	0.13	0.05	-5.40	0.06
0.05	-1.07	0.14	2.11	0.15	-1.53	0.14

Corrected

0.07	-1.21	0.15	1.97	0.16	-1.71	0.14
0.14	-1.14	0.19	2.04	0.19	-1.62	0.24
0.07	-3.14	0.16	0.03	0.16	-5.60	0.31
0.07	-3.22	0.16	-0.04	0.17	-5.65	0.32
0.13	-3.18	0.10	-0.01	0.10	-5.63	0.07
0.06	-1.04	0.20	2.13	0.20	-1.60	0.30
0.06	-1.07	0.20	2.10	0.20	-1.64	0.30
0.02	-1.06	0.04	2.12	0.04	-1.62	0.06
0.07	-2.33	0.16	0.85	0.16	-4.06	0.20
0.08	-2.30	0.16	0.87	0.16	-3.99	0.20
0.05	-2.37	0.15	0.80	0.16	-4.09	0.19
0.06	-2.33	0.07	0.84	0.07	-4.05	0.11

bias* _{26(RM-SCOI)}	2σ ^b

-2.01	0.20
-1.97	0.20
-1.99	0.06
3.56	0.20
3.58	0.20
3.58	0.20
3.59	0.20
3.58	0.03
-0.28	0.32
-0.16	0.32
-0.22	0.17
4.09	0.30
4.06	0.31
4.08	0.05
0.02	0.25
-0.23	0.26
-0.07	0.25
-0.09	0.25
0.21	0.28
0.27	0.28
0.25	0.28
0.28	0.28
0.25	0.06
4.14	0.19

3 97	0 1 9
0.01	0.15
4.05	0.24
0.05	0.32
0.00	0.32
0.02	0.08
4.08	0.31
4.03	0.31
4.05	0.06
1.60	0.21
1.67	0.21
1.57	0.20
1.61	0.11

Appendix EA6-3: SIMS Mg isotope data (session FI3, 1nA, O_2^- prima

^a 2 σ error (A) = $\sqrt{(2SE_{SIMS})^2 + (2SD_{ICPMS})^2}$ ^b 2 σ error = $\sqrt{(A)^2 + (2SD_{SIMS_bracket})^2}$ Erros of averaged values are 2SD.

File	Analysis spot	Cycle# (measured)	Cycle# (used)
	Olivine	;	
20190122@21.asc	WI-STD-91 SCOL	30	30
20190122@22.asc	WI-STD-91 SCOL	30	30
20190122@23.asc	WI-STD-91 SCOL	30	30
20190122@24.asc	WI-STD-91 SCOL	30	30
20190122@41.asc	WI-STD-91 SCOL	30	30
20190122@42.asc	WI-STD-91 SCOL	30	30
20190122@43.asc	WI-STD-91 SCOL	30	30
20190122@44.asc	WI-STD-91 SCOL	30	30
	Average (SC-OI)		
20190122@9.asc	WI-STD-91 WNOL1E_1	30	30
20190122@10.asc	WI-STD-91 WNOL1E_2	30	30
20190122@11.asc	WI-STD-91 WNOL1C_1	30	30
20190122@12.asc	WI-STD-91 WNOL1C_2	30	30
20190122@13.asc	WI-STD-91 WNOL4C_1	30	30
20190122@14.asc	WI-STD-91 WNOL4C_2	30	30
20190122@15.asc	WI-STD-91 WNOL4A_1	30	30
20190122@16.asc	WI-STD-91 WNOL4A_2	30	30
	Average (WN-OI)		
20190122@17.asc	WI-STD-91 WKOL1_1	30	30
20190122@18.asc	WI-STD-91 WKOL1_2	30	30
20190122@19.asc	WI-STD-91 WKOL8_1	30	30
20190122@20.asc	WI-STD-91 WKOL8_2	30	30
	Average (WK-OI)		
20190122@27.asc	WI-STD-91 HaKOL g1-2_1	30	30
20190122@28.asc	WI-STD-91 HaKOL g1-2_2	30	30
	Average (HaK-OI)		
20190122@31.asc	WI-STD-91 SWOL15_1	30	30
20190122@32.asc	WI-STD-91 SWOL15_2	30	30
	Average (SW-OI)		
20190122@37.asc	WI-STD-91 KNOL1-2_1	30	30
20190122@38.asc	WI-STD-91 KNOL1-2_2	30	30
	Average (KN-OI)		
20190122@95.asc	WI-STD-15 Skye-X_1_spike25	30	29

20190122@96.asc	WI-STD-15 Skye-X_2	30	30
20190122@97.asc	WI-STD-15 Skye-W_1	30	30
20190122@98.asc	WI-STD-15 Skye-W_2	30	30
20190122@99.asc	WI-STD-15 Skye-V_1	30	30
20190122@100.asc	WI-STD-15 Skye-V_2	30	30
	Average (SK-OI)		
20190122@124.asc	WI-STD-92 CL0908-1_1	30	30
20190122@125.asc	WI-STD-92 CL0908-2	30	30
20190122@126.asc	WI-STD-92 CL0908-3	30	30
20190122@127.asc	WI-STD-92 CL0908-4	30	30
20190122@128.asc	WI-STD-92 CL0908-9	30	30
20190122@129.asc	WI-STD-92 CL0908-10	30	30
20190122@130.asc	WI-STD-92 CL0908-11	30	30
20190122@131.asc	WI-STD-92 CL0908-18	30	30
	Average (CL0908-OI)		
20190122@132.asc	WI-STD-92 CL0919-1_1	30	30
20190122@133.asc	WI-STD-92 CL0919-2	30	30
20190122@134.asc	WI-STD-92 CL0919-3	30	30
20190122@135.asc	WI-STD-92 CL0919-4	30	30
20190122@136.asc	WI-STD-92 CL0919-5	30	30
20190122@137.asc	WI-STD-92 CL0919-6	30	30
20190122@138.asc	WI-STD-92 CL0919-7	30	30
20190122@139.asc	WI-STD-92 CL0919-9	30	30
	Average (CL0919-OI)		
20190122@144.asc	WI-STD-92 CL0933-1	30	30
20190122@145.asc	WI-STD-92 CL0933-2	30	30
20190122@146.asc	WI-STD-92 CL0933-2	30	30
20190122@147.asc	WI-STD-92 CL0933-4_1	30	30
20190122@148.asc	WI-STD-92 CL0933-5	30	30
20190122@149.asc	WI-STD-92 CL0933-6	30	30
20190122@150.asc	WI-STD-92 CL0933-7	30	30
20190122@151.asc	WI-STD-92 CL0933-8	30	30
	Average (CL0933-OI)		
20190122@152.asc	WI-STD-92 IGOL1_1	30	30
20190122@153.asc	WI-STD-92 IGOL2	30	30
20190122@154.asc	WI-STD-92 IGOL3	30	30
20190122@155.asc	WI-STD-92 IGOL4	30	30
20190122@156.asc	WI-STD-92 IGOL5	30	30
20190122@157.asc	WI-STD-92 IGOL6	30	30
20190122@158.asc	WI-STD-92 IGOL7	30	30
20190122@159.asc	WI-STD-92 IGOL8	30	30
	Average (IG-OI)		
20190122@7.asc	WI-STD-91 A77257_1_spike19,20	30	28
20190122@8.asc	WI-STD-91 A77257_2	30	30
20190122@170.asc	WI-STD-92 A77257-4_1	30	30

20190122@171.asc	WI-STD-92 A77257-4_2	30	30
20190122@172.asc	WI-STD-92 A77257-4_3	30	30
20190122@173.asc	WI-STD-92 A77257-4_4	30	30
	Average (A77257-OI)		
20190122@25.asc	WI-STD-91 HNOL	30	30
20190122@26.asc	WI-STD-91 HNOL	30	30
20190122@101.asc	WI-STD-91 HNOL	30	30
20190122@102.asc	WI-STD-91 HNOL	30	30
20190122@174.asc	WI-STD-92 HNOL	30	30
20190122@175.asc	WI-STD-92 HNOL	30	30
	Average (HN-OI)		
20190122@29.asc	WI-STD-91 UWOL1-42_1	30	30
20190122@30.asc	WI-STD-91 UWOL1-42_2	30	30
20190122@93.asc	WI-STD-15 UWOL1 g1_1	30	30
20190122@94.asc	WI-STD-15 UWOL1 g1_2	30	30
20190122@107.asc	WI-STD-91 UWOL1-1	30	30
20190122@108.asc	WI-STD-91 UWOL1-2	30	30
20190122@109.asc	WI-STD-91 UWOL1-3	30	30
20190122@110.asc	WI-STD-91 UWOL1-5	30	30
20190122@111.asc	WI-STD-91 UWOL1-6	30	30
20190122@112.asc	WI-STD-91 UWOL1-7	30	30
20190122@113.asc	WI-STD-91 UWOL1-10	30	30
20190122@114.asc	WI-STD-91 UWOL1-11	30	30
20190122@115.asc	WI-STD-91 UWOL1-42	30	30
20190122@166.asc	WI-STD-92 UWOL1-1 1	30	30
20190122@167.asc	WI-STD-92 UWOL1-1 ²	30	30
	Average (UW-OI)		
20190122@33.asc	WI-STD-91 OROL17_1	30	30
20190122@34.asc	WI-STD-91 OROL17 ²	30	30
20190122@35.asc	WI-STD-91 OROL5 1	30	30
20190122@36.asc	WI-STD-91 OROL5 2	30	30
20190122@164.asc	WI-STD-92 OROL2 1	30	30
20190122@165.asc	WI-STD-92 OROL2 2	30	30
0	Average (OR-OI)		
20190122@39.asc	WI-STD-91 FJOL2 1	30	30
20190122@40.asc	WI-STD-91 FJOL2 2	30	30
20190122@168.asc	WI-STD-92 FJOL2 1	30	30
20190122@169.asc	WI-STD-92 FJOL2 2	30	30
0	Average (FJ-OI)		
20190122@190.asc	NWA7325 OL7 1	30	30
20190122@191.asc	NWA7325 OL7 2	30	30
20190122@192.asc	NWA7325 OL7_3	150	30
20190122@193.asc	NWA7325 OL7 4	150	30
	Average (N7325-OI)		

Pyroxene

20190122@66.asc	WI-STD-19 Sp79-11 G5	30	30
20190122@67.asc	WI-STD-19 Sp79-11 G3	30	30
20190122@68.asc	WI-STD-19 Sp79-11 G6	30	30
20190122@69.asc	WI-STD-19 Sp79-11 G7	30	30
	Average (Sp79-11 En)		
20190122@70.asc	WI-STD-19 JE En G3	30	30
20190122@71.asc	WI-STD-19 JE En G1	30	30
	Average (JE En)		
20190122@72.asc	WI-STD-19 IG Opx G2	30	30
20190122@73.asc	WI-STD-19 IG Opx G5	30	30
	Average (IG Opx)		
20190122@74.asc	WI-STD-19 IG Cpx G3	30	30
20190122@75.asc	WI-STD-19 IG Cpx G3	30	30
	Average (IG Cpx)		
20190122@76.asc	WI-STD-19 AK95-6 G1	30	30
20190122@77.asc	WI-STD-19 AK95-6 G4	30	30
	Average (AK95-6 Di)		

ry ion beam)

	0	Measure	d			
24Mg (100M cps)	Yield (10°	d25Mg-m ‰	2SE ‰	d26Mg-m ‰	2SE ‰	d26Mg*-m
(10011000)	000/11/1)	/00		,00		/00
2.35	2.22	-1.87	0.04	-3.09	0.08	0.56
2.33	2.21	-1.96	0.04	-3.26	0.09	0.55
2.34	2.22	-1.95	0.05	-3.20	0.10	0.59
2.34	2.22	-1.97	0.04	-3.25	0.09	0.59
2.28	2.22	-1.97	0.05	-3.24	0.09	0.61
2.30	2.21	-1.99	0.05	-3.33	0.09	0.55
2.31	2.22	-1.93	0.04	-3.21	0.08	0.55
2.27	2.20	-1.93	0.07	-3.15	0.09	0.61
		-1.95	0.08	-3.22	0.15	0.58
2.36	2.27	-2.35	0.05	-3.93	0.11	0.64
2.36	2.27	-2.51	0.07	-4.24	0.12	0.65
2.37	2.28	-2.36	0.06	-4.00	0.11	0.59
2.38	2.28	-2.43	0.05	-4.09	0.11	0.64
2.37	2.28	-2.33	0.05	-3.87	0.11	0.67
2.36	2.28	-2.28	0.06	-3.84	0.11	0.60
2.37	2.29	-2.28	0.05	-3.85	0.10	0.59
2.33	2.25	-2.27	0.05	-3.78	0.10	0.64
		-2.35	0.17	-3.95	0.30	0.63
2.35	2.27	-2.19	0.04	-3.60	0.09	0.67
2.35	2.27	-2.16	0.06	-3.62	0.11	0.60
2.31	2.23	-2.00	0.05	-3.31	0.09	0.59
2.38	2.27	-1.98	0.05	-3.28	0.10	0.58
		-2.08	0.22	-3.45	0.37	0.61
2.36	2.24	-2.25	0.06	-3.81	0.12	0.59
2.35	2.25	-2.30	0.06	-3.78	0.12	0.69
		-2.28	0.06	-3.79	0.03	0.64
2.34	2.23	-0.96	0.03	-1.42	0.03	0.46
2.31	2.22	-1.03	0.03	-1.45	0.03	0.55
		-0.99	0.09	-1.44	0.04	0.50
2.22	2.14	-1.10	0.03	-1.61	0.05	0.54
2.20	2.15	-1.18	0.04	-1.79	0.04	0.51
		-1.14	0.12	-1.70	0.27	0.53
2.25	2.20	-4.16	0.04	-7.12	0.06	0.97

2.23	2.19	-4.10	0.04	-7.07	0.07	0.91
2.24	2.20	-4.06	0.03	-6.98	0.06	0.93
2.24	2.20	-4.08	0.04	-6.98	0.06	0.96
2.22	2.19	-4.10	0.04	-7.02	0.06	0.97
2.23	2.21	-4.12	0.04	-7.09	0.07	0.94
		-4.10	0.07	-7.04	0.12	0.95
2.22	2.23	-1.08	0.05	-1.30	0.05	0.80
2.22	2.24	-0.92	0.03	-1.00	0.04	0.81
2.21	2.22	-1.11	0.04	-1.39	0.04	0.76
2.25	2.23	-1.17	0.04	-1.43	0.04	0.85
2.23	2.21	-1.16	0.03	-1.43	0.04	0.83
2.23	2.21	-1.12	0.04	-1.40	0.04	0.78
2.24	2.23	-0.92	0.03	-0.96	0.04	0.84
2.20	2.20	-1.08	0.04	-1.40	0.04	0.71
		-1.07	0.19	-1.29	0.39	0.80
2.20	2.21	-1.07	0.04	-1.31	0.04	0.78
2.18	2.20	-1.21	0.04	-1.52	0.04	0.85
2.19	2.20	-1.09	0.03	-1.39	0.04	0.73
2.19	2.21	-1.32	0.03	-1.78	0.04	0.79
2.20	2.22	-1.15	0.04	-1.46	0.04	0.78
2.19	2.20	-1.12	0.03	-1.43	0.04	0.75
2.17	2.18	-1.14	0.05	-1.40	0.05	0.83
2.18	2.20	-1.10	0.04	-1.35	0.04	0.79
		-1.15	0.16	-1.46	0.29	0.79
2.20	2.21	-1.64	0.03	-2.38	0.05	0.82
2.22	2.22	-1.56	0.03	-2.27	0.05	0.77
2.21	2.21	-1.58	0.03	-2.33	0.05	0.74
2.21	2.22	-1.53	0.03	-2.17	0.04	0.81
2.22	2.22	-1.55	0.03	-2.22	0.05	0.81
2.22	2.22	-1.62	0.04	-2.31	0.05	0.85
2.22	2.21	-1.45	0.04	-2.06	0.03	0.77
2.21	2.21	-1.59	0.03	-2.29	0.06	0.80
		-1.56	0.12	-2.25	0.20	0.80
2.24	2.24	-2.00	0.05	-3.10	0.09	0.79
2.22	2.21	-1.95	0.04	-3.03	0.09	0.78
2.25	2.22	-1.99	0.04	-3.05	0.08	0.82
2.26	2.23	-1.99	0.04	-3.10	0.08	0.79
2.24	2.20	-2.00	0.04	-3.09	0.08	0.80
2.24	2.21	-2.00	0.04	-3.13	0.07	0.77
2.24	2.22	-2.03	0.05	-3.18	0.10	0.79
2.24	2.22	-1.97	0.05	-3.09	0.08	0.75
		-1.99	0.05	-3.10	0.09	0.79
2.35	2.24	-0.88	0.03	-1.17	0.04	0.54
2.32	2.23	-0.94	0.03	-1.29	0.03	0.55
2.27	2.24	-0.68	0.03	-0.59	0.04	0.73

2.26	2.24	-0.71	0.03	-0.55	0.03	0.83
2.27	2.25	-0.73	0.04	-0.62	0.04	0.80
2.27	2.25	-0.79	0.03	-0.75	0.03	0.79
		-0.79	0.21	-0.83	0.64	0.71
2.30	2.19	-3.81	0.03	-6.77	0.07	0.65
2.30	2.18	-3.81	0.04	-6.78	0.07	0.64
2.27	2.19	-3.95	0.04	-6.78	0.06	0.92
2.27	2.20	-3.92	0.04	-6.72	0.05	0.92
2.20	2.19	-3.82	0.03	-6.52	0.06	0.91
2.20	2.18	-3.81	0.04	-6.50	0.05	0.93
		-3.85	0.13	-6.68	0.27	0.83
2.32	2.21	-2.02	0.06	-3.34	0.10	0.61
2.34	2.23	-2.04	0.05	-3.39	0.09	0.59
2.28	2.21	-2.12	0.04	-3.25	0.09	0.89
2.28	2.21	-2.07	0.05	-3.16	0.08	0.88
2.30	2.24	-2.23	0.04	-3.51	0.09	0.84
2.29	2.25	-2.20	0.04	-3.47	0.09	0.82
2.30	2.26	-2.24	0.05	-3.55	0.09	0.82
2.30	2.26	-2.19	0.06	-3.49	0.10	0.78
2.30	2.26	-2.23	0.04	-3.47	0.10	0.87
2.28	2.24	-2.24	0.05	-3.52	0.09	0.85
2.28	2.24	-2.22	0.04	-3.47	0.10	0.86
2.31	2.26	-2.23	0.04	-3.48	0.09	0.88
2.29	2.24	-2.17	0.04	-3.43	0.08	0.81
2.21	2.19	-1.98	0.05	-3.01	0.08	0.86
2.20	2.18	-1.99	0.04	-3.03	0.08	0.85
		-2.15	0.20	-3.37	0.36	0.81
1.85	1.79	-1.51	0.05	-2.40	0.07	0.55
1.83	1.77	-1.50	0.05	-2.36	0.06	0.56
1.90	1.84	-1.44	0.04	-2.22	0.06	0.58
1.90	1.84	-1.44	0.05	-2.24	0.06	0.57
1.84	1.82	-1.42	0.04	-2.03	0.07	0.74
1.83	1.81	-1.42	0.07	-2.07	0.07	0.71
		-1.46	0.08	-2.22	0.30	0.62
2.19	2.12	-1.31	0.03	-1.98	0.04	0.57
2.14	2.09	-1.30	0.04	-1.98	0.05	0.57
2.08	2.08	-1.26	0.04	-1.72	0.04	0.74
2.11	2.10	-1.31	0.04	-1.77	0.04	0.79
		-1.30	0.05	-1.86	0.27	0.67
2.24	2.24	-1.97	0.04	-2.83	0.03	1.02
2.27	2.27	-2.05	0.03	-2.98	0.03	1.03
2.28	2.25	-2.03	0.04	-2.97	0.04	1.00
2.25	2.27	-2.14	0.03	-3.09	0.03	1.08
		-2.05	0.14	-2.97	0.22	1.03

1.96	1.89	0.68	0.06	2.12	0.08	0.80
1.96	1.90	0.58	0.06	2.00	0.09	0.86
1.97	1.91	0.52	0.05	1.92	0.07	0.91
1.97	1.91	0.61	0.08	2.04	0.09	0.86
		0.60	0.13	2.02	0.17	0.86
1.90	1.87	1.05	0.07	2.83	0.07	0.78
1.95	1.87	0.99	0.04	2.72	0.06	0.79
		1.02	0.08	2.77	0.15	0.78
1.87	1.80	1.37	0.05	3.47	0.07	0.79
1.88	1.82	1.38	0.04	3.47	0.08	0.79
		1.38	0.003	3.47	0.003	0.79
0.68	0.66	0.37	0.08	1.51	0.11	0.79
0.66	0.64	0.45	0.08	1.66	0.10	0.77
		0.41	0.12	1.58	0.21	0.78
 0.62	0.60	-0.72	0.09	-0.65	0.14	0.76
 0.62	0.60	-0.77	0.10	-0.81	0.15	0.69
		-0.74	0.06	-0.73	0.22	0.72

			Corrected			
2SE ‰	f ₂₅ *	2σ ^a	bias* _{25(RM-SCOI)}	$2\sigma^{b}$	f ₂₆ *	2σ ^a
0.06	-1.80	0.10			-2.93	0.18
0.06	-1.89	0.10			-3.11	0.18
0.06	-1.87	0.10			-3.05	0.19
0.05	-1.90	0.10			-3.10	0.19
0.06	-1.90	0.10			-3.08	0.18
0.07	-1.92	0.10			-3.18	0.18
0.05	-1.86	0.10			-3.06	0.18
0.08	-1.86	0.12			-3.00	0.18
0.05	-1.87	0.08			-3.07	0.15
0.06	-2.31	0.10	-0.50	0.16	-3.79	0.20
0.05	-2.47	0.11	-0.67	0.16	-4.09	0.20
0.05	-2.32	0.10	-0.51	0.16	-3.86	0.20
0.08	-2.39	0.10	-0.58	0.16	-3.94	0.19
0.05	-2.30	0.10	-0.49	0.16	-3.73	0.20
0.05	-2.24	0.10	-0.44	0.16	-3.70	0.19
0.06	-2.24	0.10	-0.44	0.16	-3.71	0.19
0.05	-2.23	0.10	-0.43	0.16	-3.64	0.19
0.06	-2.31	0.17	-0.51	0.17	-3.81	0.30
0.06	-2.11	0.13	-0.30	0.16	-3.44	0.22
0.07	-2.08	0.14	-0.27	0.16	-3.46	0.22
0.05	-1.92	0.13	-0.11	0.16	-3.15	0.22
0.06	-1.89	0.13	-0.08	0.16	-3.11	0.22
0.08	-2.00	0.22	-0.19	0.22	-3.29	0.37
0.04	-2.20	0.14	-0.32	0.15	-3.65	0.23
0.06	-2.24	0.14	-0.36	0.15	-3.62	0.22
0.15	-2.22	0.06	-0.34	0.06	-3.64	0.03
0.05	-0.94	0.14	0.93	0.15	-1.39	0.16
0.06	-1.00	0.14	0.87	0.15	-1.42	0.16
0.13	-0.97	0.09	0.90	0.09	-1.41	0.04
0.05	-1.14	0.18	0.74	0.18	-1.59	0.30
0.05	-1.22	0.18	0.65	0.18	-1.78	0.30
0.04	-1.18	0.12	0.70	0.12	-1.69	0.27
0.06	-3.27	0.20	-1.35	0.21	-5.41	0.54

0.07	0.04	0.04	4 00	0.04	F 00	0 5 4
0.07	-3.21	0.21	-1.29	0.21	-5.36	0.54
0.05	-3.17	0.20	-1.25	0.21	-5.27	0.54
0.07	-3.19	0.20	-1.27	0.21	-5.27	0.54
0.06	-3.21	0.21	-1.29	0.21	-5.31	0.54
0.06	-3.23	0.21	-1.32	0.21	-5.38	0.54
0.05	-3.21	0.07	-1.30	0.07	-5.33	0.12
0.07	-0.87	0.07	0.93	0.11	-0.89	0.09
0.05	-0.71	0.06	1.09	0.11	-0.59	0.08
0.05	-0.90	0.06	0.90	0.11	-0.98	0.08
0.05	-0.96	0.06	0.84	0.11	-1.02	0.08
0.05	-0.95	0.06	0.85	0.11	-1.02	0.08
0.06	-0.91	0.06	0.89	0.11	-0.99	0.08
0.05	-0.71	0.06	1.09	0.11	-0.55	0.08
0.07	-0.87	0.06	0.93	0.11	-0.99	0.08
0.09	-0.86	0.19	0.94	0.19	-0.88	0.39
0.06	-0.86	0.07	0.94	0.11	-0.94	0.07
0.05	-1.01	0.07	0.79	0.11	-1.15	0.07
0.06	-0.88	0.07	0.92	0.11	-1.02	0.07
0.07	-1.11	0.07	0.69	0.11	-1.41	0.07
0.06	-0.94	0.07	0.86	0.11	-1.09	0.07
0.06	-0.91	0.07	0.89	0.11	-1.06	0.07
0.06	-0.93	0.08	0.87	0.11	-1.03	0.08
0.07	-0.89	0.07	0.91	0.11	-0.98	0.07
0.08	-0.94	0.16	0.86	0.16	-1.09	0.29
0.05	-1.39	0.06	0.45	0.14	-1.91	0.08
0.06	-1.31	0.06	0.53	0.14	-1.80	0.08
0.07	-1.33	0.06	0.52	0.14	-1.87	0.08
0.06	-1.28	0.06	0.56	0.14	-1.70	0.07
0.04	-1.30	0.06	0.54	0.14	-1.75	0.08
0.07	-1.37	0.06	0.47	0.14	-1.84	0.08
0.07	-1.20	0.06	0.64	0.14	-1.59	0.07
0.06	-1.34	0.06	0.51	0.14	-1.82	0.08
0.07	-1.31	0.12	0.53	0.12	-1.78	0.20
0.07	-1.97	0.10	-0.13	0.14	-3.05	0.18
0.05	-1.93	0.10	-0.08	0.14	-2.98	0.18
0.05	-1.96	0.10	-0.12	0.14	-3.00	0.17
0.06	-1.97	0.10	-0.13	0.14	-3.05	0.17
0.06	-1.97	0.10	-0.13	0.14	-3.04	0.17
0.06	-1.98	0.10	-0.13	0.14	-3.08	0.17
0.06	-2.01	0.11	-0.17	0.14	-3.13	0.18
0.07	-1.94	0.10	-0.10	0.14	-3.04	0.17
0.04	-1.96	0.05	-0.12	0.05	-3.05	0.09
0.06	-0.81	0.14	1.00	0.16	-0.88	0.13
0.06	-0.87	0.14	0.94	0.16	-1.00	0.13
0.04	-0.60	0.14	1.22	0.18	-0.30	0.13

0.06	-0.64	0.14	1.19	0.18	-0.27	0.13
0.07	-0.65	0.15	1.17	0.18	-0.33	0.14
0.06	-0.72	0.14	1.11	0.18	-0.46	0.13
0.26	-0.72	0.21	1.10	0.23	-0.54	0.64
0.06	-3.45	0.09	-1.58	0.13	-6.08	0.15
0.06	-3.45	0.09	-1.57	0.13	-6.09	0.16
0.08	-3.59	0.09	-1.67	0.11	-6.09	0.15
0.06	-3.56	0.09	-1.64	0.11	-6.03	0.15
0.06	-3.45	0.09	-1.63	0.18	-5.83	0.15
0.06	-3.45	0.09	-1.63	0.18	-5.81	0.15
0.28	-3.49	0.13	-1.62	0.08	-5.99	0.27
0.07	-1.89	0.11	-0.01	0.13	-3.10	0.22
0.08	-1.91	0.11	-0.03	0.13	-3.14	0.22
0.06	-1.99	0.11	-0.07	0.11	-3.01	0.22
0.07	-1.94	0.11	-0.02	0.12	-2.92	0.21
0.07	-2.09	0.11	-0.19	0.11	-3.27	0.22
0.06	-2.07	0.11	-0.16	0.11	-3.23	0.22
0.06	-2.11	0.11	-0.20	0.12	-3.31	0.22
0.06	-2.06	0.12	-0.15	0.12	-3.25	0.22
0.05	-2.09	0.11	-0.18	0.11	-3.23	0.22
0.05	-2.10	0.11	-0.20	0.12	-3.28	0.22
0.06	-2.09	0.11	-0.18	0.12	-3.23	0.22
0.05	-2.10	0.11	-0.19	0.11	-3.24	0.22
0.06	-2.04	0.11	-0.13	0.11	-3.18	0.21
0.07	-1.85	0.11	-0.02	0.18	-2.77	0.21
0.05	-1.85	0.11	-0.03	0.18	-2.79	0.21
0.18	-2.01	0.20	-0.12	0.15	-3.13	0.36
0.07	-1.50	0.08	0.38	0.13	-2.38	0.14
0.07	-1.49	0.08	0.39	0.13	-2.34	0.14
0.08	-1.43	0.07	0.45	0.13	-2.20	0.14
0.06	-1.43	0.08	0.45	0.13	-2.22	0.14
0.07	-1.41	0.08	0.42	0.18	-2.01	0.14
0.10	-1.41	0.09	0.42	0.18	-2.04	0.14
0.16	-1.44	0.08	0.42	0.06	-2.20	0.30
0.06	-1.24	0.20	0.64	0.20	-1.88	0.30
0.04	-1.23	0.20	0.64	0.20	-1.88	0.30
0.06	-1.19	0.20	0.63	0.21	-1.62	0.30
0.05	-1.24	0.20	0.58	0.21	-1.67	0.30
0.23	-1.23	0.05	0.62	0.06	-1.76	0.27
0.07	-1.90	0.15	-0.14	0.17	-2.77	0.10
0.06	-1.98	0.14	-0.22	0.17	-2.92	0.10
0.06	-1.96	0.14	-0.20	0.17	-2.91	0.10
0.07	-2.07	0.14	-0.28	0.15	-3.04	0.10
0.07	-1.98	0.14	-0.21	0.12	-2.91	0.22

0.06	0.73	0.08	2.54	0.12	2.23	0.14
0.06	0.64	0.08	2.45	0.12	2.10	0.14
0.05	0.58	0.07	2.39	0.12	2.02	0.13
0.08	0.67	0.10	2.48	0.12	2.15	0.14
0.09	0.65	0.13	2.46	0.13	2.13	0.17
0.10	1.31	0.09	3.12	0.12	3.33	0.13
0.08	1.25	0.07	3.06	0.12	3.23	0.13
0.01	1.28	0.08	3.09	80.0	3.28	0.15
0.08	1.41	0.09	3.22	0.12	3.54	0.15
0.06	1.41	0.09	3.23	0.12	3.54	0.15
0.002	1.41	0.003	3.22	0.003	3.54	0.003
0.12	0.51	0.12	2.32	0.13	1.79	0.22
0.13	0.59	0.12	2.40	0.13	1.93	0.22
0.03	0.55	0.12	2.36	0.12	1.86	0.21
0.11	0.02	0.10	1.83	0.12	0.78	0.18
0.15	-0.03	0.12	1.78	0.12	0.62	0.19
0.10	0.00	0.06	1.81	0.06	0.70	0.22

bias* _{26(RM-SCOI)}	2σ ^b	

-0.80	0.26
-1.1	1 0.26
-0.87	7 0.26
-0.9	5 0.26
-0.74	4 0.26
-0.7	1 0.26
-0.72	2 0.26
-0.6	5 0.26
-0.82	2 0.30
-0.4	5 0.26
-0.47	7 0.26
-0.16	o 0.26
-0.12	2 0.26
-0.30	0.37
-0.58	3 0.24
-0.56	o 0.24
-0.57	7 0.03
1.68	3 0.23
1.65	5 0.23
1.60	3 0.04
1.48	3 0.31
1.29	<u>9 0.31</u>
1.38	3 0.27
-2.54	4 0.54

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-2.49	0.54
-2.39	0.54
-2.39	0.54
-2.43	0.54
-2.50	0.54
-2.46	0.12
1.80	0.19
2.10	0.19
1.71	0.19
1.67	0.19
1.67	0.19
1.70	0.19
2.14	0.19
1.69	0.19
1.81	0.39
1.74	0.19
1.54	0.19
1.66	0.19
1.28	0.19
1.60	0.19
1.62	0.19
1.00	0.19
1.71	0.19
0.86	0.23
0.00	0.27
0.00	0.27
1.06	0.27
1.01	0.27
0.92	0.27
1.17	0.27
0.94	0.27
0.98	0.20
-0.29	0.27
-0.22	0.27
-0.24	0.27
-0.29	0.27
-0.28	0.27
-0.32	0.27
-0.37	0.27
-0.28	0.27
-0.29	0.09
2.12	0.26
1.99	0.26
2.46	0.32

2.50	0.32
2.43	0.32
2.30	0.32
2.30	0.41
-3.03	0.23
-3.03	0.23
-3.21	0.20
-3.15	0.20
-3.08	0.32
-3.06	0.32
-3.09	0.15
-0.03	0.23
-0.08	0.23
-0.13	0.23
-0.04	0.22
-0.39	0.23
-0.36	0.23
-0.44	0.23
-0.38	0.23
-0.36	0.23
-0.41	0.23
-0.36	0.23
-0.37	0.23
-0.31	0.22
-0.01	0.32
-0.03	0.32
-0.25	0.34
0.69	0.23
0.73	0.23
0.87	0.23
0.85	0.23
0.75	0.32
0.71	0.32
0.77	0.15
1.19	0.31
1.19	0.31
1.14	0.32
1.09	0.32
1.15	0.10
-0.21	0.27
-0.37	0.27
-0.35	0.27
-0.19	0.24
-0.28	0.18

5.00	0.20
4.87	0.20
4.79	0.20
4.92	0.20
4.90	0.17
6.11	0.20
6.00	0.20
6.05	0.15
6.32	0.20
6.32	0.20
6.32	0.003
4.55	0.23
4.70	0.23
4.63	0.21
3.55	0.20
3.39	0.20
3.47	0.22

Appendix EA6-4: SIMS Mg isotope data (session FI4, 2.6nA, O⁻ prima

^a 2 σ error (A) = $\sqrt{(2SE_{SIMS})^2 + (2SD_{ICPMS})^2}$ ^b 2 σ error = $\sqrt{(A)^2 + (2SD_{SIMS_bracket})^2}$ Erros of averaged values are 2SD.

File	Analysis spot	Cycle#	Cycle#
FIIC	Analysis spot	(measured)	(used)
20190425@31.asc	WI-STD-91 SCOL	30	30
20190425@32.asc	WI-STD-91 SCOL	30	30
20190425@33.asc	WI-STD-91 SCOL	30	30
20190425@34.asc	WI-STD-91 SCOL	30	30
20190425@47.asc	WI-STD-91 SCOL	30	30
20190425@48.asc	WI-STD-91 SCOL	30	30
20190425@49.asc	WI-STD-91 SCOL	30	30
20190425@50.asc	WI-STD-91 SCOL	30	30
	Average (SC-OI)		
20190425@21.asc	WI-STD-91 UWOL1-42_1	30	30
20190425@22.asc	WI-STD-91 UWOL1-42_2	30	30
20190425@66.asc	WI-STD-15 UWOL1 g1_1	30	30
20190425@67.asc	WI-STD-15 UWOL1 g1_2	30	30
	Average (UW-OI)		
20190425@23.asc	WI-STD-91 SWOL15_1	30	30
20190425@24.asc	WI-STD-91 SWOL15_2	30	30
20190425@25.asc	WI-STD-91 SWOL8_1	30	30
20190425@26.asc	WI-STD-91 SWOL8_2	30	30
20190425@27.asc	WI-STD-91 SWOL3_1	30	30
20190425@28.asc	WI-STD-91 SWOL3_2	30	30
20190425@29.asc	WI-STD-91 SWOL1_1	30	30
20190425@30.asc	WI-STD-91 SWOL1_2	30	30
	Average (SW-OI)		
20190425@35.asc	WI-STD-91 OROL17_1	30	30
20190425@36.asc	WI-STD-91 OROL17_2	30	30
20190425@37.asc	WI-STD-91 OROL5_1	30	30
20190425@38.asc	WI-STD-91 OROL5_2	30	30
	Average (OR-OI)		
20190425@43.asc	WI-STD-91 WKOL1_1	30	30
20190425@44.asc	WI-STD-91 WKOL1_2	30	30
20190425@45.asc	WI-STD-91 WKOL8_1	30	30
20190425@46.asc	WI-STD-91 WKOL8_2	30	30
	Average (WK-OI)		
20190425@51.asc	WI-STD-91 WNOL1E_1	30	30
20190425@61.asc	WI-STD-91 WNOL1E_3	30	30

20190425@53.asc	WI-STD-91 WNOL1C_1	30	30
20190425@54.asc	WI-STD-91 WNOL1C_2	30	30
20190425@55.asc	WI-STD-91 WNOL4C 1	30	30
20190425@56.asc	WI-STD-91 WNOL4C ²	30	30
20190425@57.asc	WI-STD-91 WNOL4A 1	30	30
20190425@58.asc	WI-STD-91 WNOL4A 2	30	30
	Average (WN-OI)		
20190425@59.asc	WI-STD-91 A77257-5_1	30	30
20190425@60.asc	WI-STD-91 A77257-5 2	30	30
20190425@114.asc	WI-STD-92 A77257-4 1	30	30
20190425@115.asc	WI-STD-92 A77257-4 2	150	30
	Average (A77257-OI)		
20190425@68.asc	WI-STD-15 SKOL-W_1	30	30
20190425@69.asc	WI-STD-15 SKOL-W_2	30	30
20190425@70.asc	WI-STD-15 SKOL-V 1	30	30
20190425@71.asc	WI-STD-15 SKOL-V_2	30	30
20190425@72.asc	WI-STD-15 SKOL-X_1	30	30
20190425@73.asc	WI-STD-15 SKOL-X 2	30	30
20190425@74.asc	WI-STD-92 SKOL-X 3	150	30
	Average (SK-OI)		
20190425@17.asc	WI-STD-91 HNOL	30	30
20190425@18.asc	WI-STD-91 HNOL	30	30
	Average (HN-OI)		
20190425@19.asc	WI-STD-91 HaKOL g1-2_1	30	30
20190425@20.asc	WI-STD-91 HaKOL g1-2_2	30	30
	Average (HaK-OI)		
20190425@39.asc	WI-STD-91 KNOL1-2_1	30	30
20190425@40.asc	WI-STD-91 KNOL1-2_2	30	30
	Average (KN-OI)		
20190425@41.asc	WI-STD-91 FJOL2_1	30	30
20190425@42.asc	WI-STD-91 FJOL2_2	30	30
	Average (FJ-OI)		
20190425@101.asc	N7325 OL7_1	30	30
20190425@102.asc	N7325 OL7_2	150	30
~	Average (N7325-OI)		
20190425@116.asc	WI-STD-92 IGOL1_1	150	30
20190425@117.asc	WI-STD-92 IGOL1_2	30	30
	Average (IG-OI)		
20190425@111.asc	WI-STD-92 CL0908-1_1	30	30
20190425@118.asc	WI-STD-92 CL0908-1_2	150	30
	Average (CL0908-OI)		
20190425@112.asc	WI-STD-92 CL0919-1_1	30	30
20190425@119.asc	WI-STD-92 CL0919-1_2	150	30
	Average (CL0919-OI)		
20190425@113.asc	WI-STD-92 CL0933-4_1	30	30
20190425@120.asc	WI-STD-92 CL0933-4_2	150	30
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	Average (CL0933-OI)		

ary ion beam)

Measured						
24Mg (100M cps)	Yield (10 ⁸ cps/nA)	d25Mg-m ‰	2SE ‰	d26Mg-m ‰	2SE ‰	d26Mg*-m ‰
2.36	. 0.88	-2.42	0.12	-4.32	0.21	0.39
2.37	0.88	-2.46	0.12	-4.37	0.21	0.43
2.37	0.88	-2.52	0.10	-4.39	0.21	0.51
2.37	0.88	-2.50	0.12	-4.43	0.21	0.44
2.36	0.88	-2.46	0.11	-4.34	0.22	0.46
2.37	0.88	-2.49	0.11	-4.39	0.21	0.46
2.37	0.88	-2.50	0.11	-4.41	0.22	0.45
2.37	0.89	-2.51	0.10	-4.44	0.22	0.46
		-2.48	0.07	-4.39	0.09	0.45
2.34	0.88	-2.52	0.12	-4.51	0.23	0.41
2.35	0.88	-2.48	0.12	-4.46	0.22	0.37
2.34	0.88	-2.55	0.12	-4.51	0.23	0.45
2.36	0.88	-2.55	0.12	-4.49	0.23	0.48
		-2.52	0.07	-4.49	0.05	0.43
2.32	0.87	-1.66	0.07	-2.94	0.11	0.29
2.33	0.87	-1.74	0.06	-3.01	0.11	0.38
2.32	0.86	-1.63	0.06	-2.82	0.10	0.37
2.33	0.87	-1.61	0.06	-2.80	0.11	0.34
2.32	0.86	-1.60	0.06	-2.77	0.10	0.35
2.34	0.87	-1.61	0.06	-2.79	0.11	0.35
2.32	0.87	-1.56	0.06	-2.68	0.11	0.36
2.34	0.87	-1.60	0.06	-2.77	0.10	0.35
		-1.63	0.11	-2.82	0.21	0.35
2.07	0.77	-0.54	0.03	-0.75	0.03	0.31
2.07	0.77	-0.57	0.04	-0.78	0.03	0.34
2.12	0.79	-0.45	0.04	-0.57	0.02	0.30
2.12	0.79	-0.43	0.03	-0.54	0.02	0.30
		-0.50	0.14	-0.66	0.25	0.31
2.41	0.90	-2.54	0.10	-4.51	0.19	0.45
2.41	0.90	-2.42	0.09	-4.24	0.19	0.47
2.41	0.90	-2.49	0.10	-4.39	0.19	0.46
2.42	0.90	-2.43	0.09	-4.30	0.19	0.43
		-2.47	0.12	-4.36	0.24	0.45
2.43	0.91	-2.69	0.10	-4.80	0.18	0.45
2.42	0.90	-2.79	0.10	-4.93	0.18	0.51

2.43	0.91	-2.68	0.09	-4.82	0.18	0.40
2.44	0.91	-2.72	0.09	-4.81	0.18	0.48
2.44	0.91	-2.62	0.10	-4.60	0.19	0.50
2.44	0.91	-2.59	0.08	-4.63	0.16	0.42
2.42	0.90	-2.61	0.09	-4.63	0.18	0.46
2.44	0.91	-2.52	0.08	-4.42	0.15	0.49
		-2.65	0.17	-4.70	0.33	0.46
2.39	0.89	-0.78	0.03	-1.13	0.03	0.39
2.42	0.91	-0.66	0.03	-0.82	0.04	0.47
2.47	0.92	-0.68	0.03	-0.69	0.03	0.64
2.45	0.91	-0.69	0.03	-0.73	0.04	0.62
		-0.70	0.11	-0.84	0.40	0.53
2.29	0.86	-4.13	0.05	-7.62	0.11	0.43
2.30	0.86	-4.23	0.07	-7.70	0.12	0.53
2.30	0.86	-4.27	0.05	-7.85	0.10	0.46
2.30	0.86	-4.25	0.06	-7.79	0.12	0.48
2.31	0.87	-4.26	0.07	-7.84	0.11	0.46
2.31	0.87	-4.27	0.07	-7.81	0.11	0.51
2.31	0.86	-4.28	0.05	-7.83	0.11	0.49
		-4.24	0.10	-7.78	0.17	0.48
2.28	0.86	-3.89	0.06	-7.20	0.11	0.39
2.30	0.86	-3.92	0.06	-7.24	0.10	0.39
		-3.91	0.04	-7.22	0.07	0.39
2.37	0.89	-2.68	0.13	-4.77	0.24	0.44
2.37	0.89	-2.66	0.13	-4.74	0.23	0.44
		-2.67	0.02	-4.76	0.04	0.44
2.45	0.91	-0.25	0.03	-0.15	0.03	0.34
2.43	0.90	-0.21	0.03	-0.08	0.03	0.33
		-0.23	0.06	-0.12	0.09	0.34
2.40	0.89	-0.34	0.03	-0.33	0.03	0.33
2.40	0.89	-0.34	0.03	-0.32	0.03	0.33
		-0.34	0.002	-0.32	0.01	0.33
2.44	0.91	-2.13	0.05	-3.43	0.10	0.73
2.47	0.92	-2.16	0.04	-3.48	0.09	0.72
		-2.15	0.03	-3.45	0.07	0.73
2.39	0.89	-2.70	0.12	-4.59	0.23	0.68
2.39	0.89	-2.69	0.12	-4.49	0.24	0.74
		-2.69	0.02	-4.54	0.13	0.71
2.34	0.87	-1.71	0.06	-2.75	0.11	0.57
2.35	0.88	-1.77	0.07	-2.77	0.10	0.69
		-1.74	0.10	-2.76	0.03	0.63
2.36	0.88	-1.46	0.04	-2.24	0.09	0.61
2.36	0.88	-1.47	0.05	-2.24	0.08	0.63
		-1.47	0.02	-2.24	0.004	0.62
2.37	0.88	-2.41	0.09	-3.98	0.19	0.72

2.37	0.88	-2.41	0.08	-4.04	0.17	0.66
		-2.41	0.002	-4.01	0.08	0.69

2SE ‰	f ₂₅ *	$2\sigma^{a}$	bias* _{25(RM-SCOI)}	2σ ^b	f ₂₆ *	2σ ^a
0.06	-2.34	0.15			-4.16	0.27
0.06	-2.39	0.15			-4.22	0.27
0.05	-2.44	0.14			-4.24	0.26
0.05	-2.43	0.15			-4.28	0.26
0.05	-2.39	0.15			-4.19	0.27
0.05	-2.41	0.14			-4.23	0.27
0.04	-2.42	0.14			-4.26	0.27
0.05	-2.44	0.14			-4.29	0.27
0.06	-2.41	0.07			-4.23	0.09
0.04	-2.39	0.15	-0.01	0.20	-4.26	0.30
0.06	-2.34	0.15	0.04	0.20	-4.21	0.30
0.04	-2.41	0.15	-0.06	0.16	-4.27	0.30
0.05	-2.42	0.16	-0.07	0.16	-4.25	0.30
0.10	-2.39	0.07	-0.03	0.10	-4.25	0.05
0.05	-1.64	0.16	0.74	0.20	-2.91	0.19
0.05	-1.72	0.15	0.66	0.20	-2.98	0.19
0.06	-1.61	0.15	0.77	0.20	-2.79	0.19
0.05	-1.59	0.15	0.79	0.20	-2.77	0.19
0.05	-1.58	0.15	0.80	0.20	-2.74	0.19
0.05	-1.59	0.15	0.79	0.20	-2.76	0.19
0.05	-1.54	0.15	0.84	0.20	-2.65	0.19
0.05	-1.58	0.15	0.80	0.20	-2.74	0.18
0.05	-1.60	0.11	0.77	0.11	-2.79	0.21
0.04	-0.53	0.07	1.88	0.12	-0.73	0.13
0.06	-0.56	0.07	1.85	0.12	-0.76	0.13
0.07	-0.43	0.07	1.98	0.12	-0.55	0.13
0.05	-0.42	0.07	2.00	0.12	-0.51	0.13
0.03	-0.49	0.14	1.93	0.14	-0.64	0.25
0.05	-2.46	0.16	-0.05	0.16	-4.34	0.27
0.05	-2.33	0.16	0.08	0.16	-4.07	0.27
0.06	-2.40	0.16	0.01	0.17	-4.22	0.27
0.05	-2.34	0.16	0.07	0.16	-4.13	0.27
0.04	-2.38	0.12	0.03	0.12	-4.19	0.24
0.05	-2.65	0.13	-0.26	0.14	-4.65	0.24
0.05	-2.76	0.13	-0.36	0.13	-4.79	0.24

Corrected

0.05	-2.64	0.12	-0.25	0.12	-4.67	0.24
0.05	-2.68	0.12	-0.29	0.13	-4.67	0.24
0.04	-2.58	0.13	-0.19	0.13	-4.46	0.25
0.05	-2.56	0.12	-0.16	0.12	-4.49	0.23
0.05	-2.57	0.12	-0.18	0.12	-4.48	0.24
0.05	-2.49	0.11	-0.09	0.12	-4.28	0.22
0.08	-2.62	0.17	-0.22	0.17	-4.56	0.33
0.05	-0.71	0.14	1.69	0.15	-0.85	0.13
0.06	-0.59	0.14	1.81	0.15	-0.53	0.13
0.05	-0.61	0.14	2.00	0.16	-0.40	0.13
0.07	-0.62	0.14	1.99	0.16	-0.44	0.14
0.24	-0.63	0.11	1.87	0.30	-0.56	0.40
0.07	-3.24	0.21	-0.90	0.21	-5.91	0.55
0.05	-3.34	0.21	-0.99	0.22	-5.99	0.55
0.04	-3.38	0.21	-1.03	0.21	-6.14	0.55
0.05	-3.36	0.21	-1.01	0.21	-6.08	0.55
0.06	-3.37	0.21	-1.03	0.22	-6.13	0.55
0.05	-3.38	0.21	-1.04	0.22	-6.10	0.55
0.06	-3.39	0.21	-1.04	0.21	-6.13	0.55
0.07	-3.35	0.10	-1.01	0.10	-6.07	0.17
0.05	-3.53	0.10	-1.15	0.20	-6.50	0.18
0.05	-3.56	0.11	-1.18	0.20	-6.55	0.17
0.005	-3.54	0.04	-1.17	0.04	-6.53	0.07
0.05	-2.62	0.18	-0.24	0.20	-4.61	0.30
0.05	-2.60	0.18	-0.22	0.20	-4.58	0.30
0.001	-2.61	0.02	-0.23	0.02	-4.60	0.04
0.05	-0.29	0.18	2.12	0.18	-0.14	0.30
0.06	-0.25	0.18	2.16	0.18	-0.07	0.30
0.02	-0.27	0.06	2.14	0.06	-0.10	0.09
0.05	-0.27	0.20	2.15	0.20	-0.23	0.30
0.05	-0.27	0.20	2.15	0.20	-0.22	0.30
0.01	-0.27	0.002	2.15	0.002	-0.23	0.01
0.05	-2.06	0.15	0.26	0.16	-3.37	0.13
0.06	-2.08	0.15	0.24	0.15	-3.42	0.13
0.01	-2.07	0.03	0.25	0.03	-3.40	0.07
0.04	-2.67	0.15	-0.07	0.16	-4.54	0.27
0.06	-2.66	0.15	-0.06	0.16	-4.44	0.28
0.09	-2.67	0.02	-0.07	0.02	-4.49	0.13
0.06	-1.50	0.07	1.11	0.16	-2.34	0.13
0.05	-1.56	0.09	1.04	0.16	-2.36	0.13
0.16	-1.53	0.10	1.07	0.10	-2.35	0.03
0.03	-1.25	0.07	1.36	0.16	-1.87	0.10
0.05	-1.26	0.08	1.34	0.16	-1.87	0.10
0.04	-1.26	0.02	1.35	0.02	-1.87	0.004
0.06	-2.16	0.10	0.44	0.16	-3.51	0.20

0.00 2.16 0.002 0.44 0.002 3.54			0.10	0.44	0.10	-2.16	0.05	
0.09 -2.10 0.002 0.44 0.002 -3.54	0.08	-3.54	0.002	0.44	0.002	-2.16	0.09	

-0.04	0.36
0.01	0.36
-0.17	0.31
-0.15	0.31
-0.09	0.18
1.32	0.36
1.25	0.36
1.45	0.36
1.46	0.36
1.49	0.36
1.48	0.36
1.59	0.36
1.50	0.36
1.44	0.21
3.52	0.19
3.49	0.19
3.70	0.19
3.74	0.19
3.61	0.25
-0.11	0.28
0.16	0.28
0.01	0.28
0.10	0.28
0.04	0.24
-0.46	0.25
-0.59	0.25

-0.48	0.25
-0.48	0.25
-0.26	0.26
-0.29	0.24
-0.29	0.25
-0.08	0.23
-0.37	0.33
3.37	0.21
3.68	0.21
3.99	0.30
3.95	0.30
3.75	0.58
-1.82	0.55
-1.90	0.56
-2.05	0.55
-1.99	0.55
-2.05	0.55
-2.01	0.55
-2.04	0.55
-1.98	0.17
-2.28	0.36
-2.33	0.36
-2.31	0.07
-0.39	0.36
-0.36	0.36
-0.37	0.04
4.11	0.31
4.18	0.30
4.15	0.09
4.02	0.31
4.03	0.31
4.02	0.01
0.49	0.25
0.44	0.25
0.47	0.07
-0.16	0.30
-0.07	0.30
-0.12	0.13
2.04	0.30
2.02	0.30
2.03	0.03
2.52	0.30
2.52	0.30
2.52	0.004
0.86	0.30

0.81	0.30
0.83	0.08

^a 2σ error (%	$(2SD_{SIMS})^2 -$	+(2SD _{ICPMS})) ² , where 2SD _{SIMS} r	epresents e	ither δ ²⁵ Mg hom
			02-		
			Session	FI1	Sessior
	RM name	Fo#	bias* _{25(RM-SCOI)}	2 σ ^a	bias* _{25(RM-SCOI)}
	OR-OI	59.27	0.70	0.14	0.42
	FJ-OI	73.36	0.89	0.24	0.62
	KN-OI	77.97	0.95	0.22	0.70
	CL09-19-OI	79.44			0.86
	CL09-08-OI	80.53			0.94
	SW-OI	81.81	1.23	0.19	0.90
	A77257-OI	85.70	1.49	0.27	1.10
	CL09-33-OI	86.63			0.53
	SC-OI	88.82	0.00	0.17	0.00
	UWOL-1	89.23	0.01	0.18	-0.12
	IG-OI	89.58			-0.12
	HaK-OI	91.85	-0.36	0.19	-0.34
	WK-OI	94.30	-0.07	0.18	-0.19
	WN-OI	95.52	-0.46	0.19	-0.51
	N7325-OI	97.42			-0.21
	SK-OI	99.81			-1.30
L	HN-OI	100.00	-1.41	0.16	-1.62

		O-					
	Session FI2						
RM name	Fo#	bias* _{25(RM-SCOI)}	2 σ ^a	bias* _{25(RM-SCOI)}			
OR-OI	59.27	1.87	0.14	1.93			
FJ-OI	73.36	2.12	0.23	2.15			
KN-OI	77.97	2.08	0.22	2.14			
CL09-19-OI	79.44			1.35			
CL09-08-OI	80.53			1.07			
SW-OI	81.81	0.84	0.19	0.77			
A77257-OI	85.70	2.04	0.27	1.87			
CL09-33-OI	86.63			0.44			
SC-OI	88.82	0.00	0.17	0.00			
UWOL-1	89.23	-0.01	0.18	-0.03			
IG-OI	89.58			-0.07			
HaK-Ol	91.85	-0.13	0.18	-0.23			
WK-OI	94.30	0.13	0.18	0.03			
WN-OI	95.52	-0.07		-0.22			
N7325-OI	97.42		0.00	0.25			
SK-OI	99.81			-1.01			

	HN-OI	100.00	-1.00	0.15	-1.17
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Id Fo content obtained by multiple sessions

ogeneity of each RM or long-term reproducibility of the running standard (SC-OI) durin



0.24

g each session, whichever is larger.





Data reported in Ushikubo et al. (2013).GCA

^a 2 σ error (A) = $\sqrt{(2SE_{SIMS})^2 + (2SD_{ICPMS})^2}$ ^b 2 σ error = $\sqrt{(A)^2 + (2SD_{SIMS_bracket})^2}$ Erros of averaged values are 2SD.

File	Analysis spot	24Mg (100M cps)	Yield (10 ⁸ cps/nA)
Olivine (L	Jshikubo et al., 2013.GCA)		
20090910-17	STD18 SC-OL	2.32	0.75
20090910-18	STD18_SC-OL	2.30	0.73
20090910-19	STD18_SC-OL	2.32	0.73
20090910-20	STD18_SC-OL	2.34	0.73
20090910-36	STD18_SC-OL	2.75	0.74
20090910-37	STD18_SC-OL	2.72	0.73
20090910-38	STD18_SC-OL	2.74	0.73
20090910-39	STD18_SC-OL	2.80	0.74
	Average (SC-OI)		
20090910-21	STD18_Fo100-4	2.42	0.75
20090910-22	STD18_Fo100-4	2.42	0.75
20090910-23	STD18_Fo100-4	2.44	0.74
20090910-24	STD18_Fo100-4	2.47	0.75
20090910-25	STD18_Fo100-4	2.49	0.75
20090910-26	STD18_Fo100-4	2.50	0.74
20090910-27	STD18_Fo100-4	2.53	0.75
	Average (HN-OI)		
20090910-28	STD18_Fo90-75	2.57	0.75
20090910-29	STD18_Fo90-75	2.56	0.74
20090910-30	STD18_Fo90-75	2.59	0.74
20090910-31	STD18_Fo90-75	2.64	0.75
	Average (IG-OI)		
20090910-32	STD18_Fo60-37	2.60	0.73
20090910-33	STD18_Fo60-37	2.61	0.73
20090910-34	STD18_Fo60-37	2.62	0.73
20090910-35	STD18_Fo60-34	2.66	0.73
	Average (OR-OI)		
Pyroxene (L	Jshikubo et al., 2013.GCA)		
20090910-42	STD19 SC-OL	2.96	0.77
20090910-43	STD19 SC-OL	2.99	0.77

20090910-44	STD19 SC-OL	3.00	0.77
20090910-45	STD19 SC-OL	3.03	0.77
20090910-65	STD19 SC-OL	3.45	0.74
20090910-66	STD19 ⁻ SC-OL	3.40	0.72
20090910-67	STD19 ⁻ SC-OL	3.40	0.72
20090910-68	STD19_SC-OL	3.42	0.72
	Average (SC-OI)		
20090910-46	STD19_SP79-11-G3	2.56	0.65
20090910-47	STD19_SP79-11-G3	2.60	0.65
20090910-48	STD19_SP79-11-G3	2.63	0.65
20090910-49	STD19_SP79-11-G3	2.65	0.65
20090910-50	STD19_SP79-11-G3	2.69	0.66
20090910-51	STD19_SP79-11-G3	2.61	0.63
20090910-52	STD19_SP79-11-G3	2.69	0.64
20090910-61	STD19_SP79-11-G5	2.89	0.64
20090910-62	STD19_SP79-11-G6	2.90	0.64
20090910-63	STD19_SP79-11-G4	2.94	0.64
20090910-64	STD19_SP79-11-G3	2.97	0.64
	Average (Sp79-11 En)		
20090910-53	STD19_IGOPX-G5	2.54	0.60
20090910-54	STD19_IGOPX-G4	2.59	0.61
20090910-55	STD19_IGOPX-G3	2.55	0.60
20090910-56	STD19_IGOPX-G1	2.59	0.60
	Average (IG-Opx)		
20090910-57	STD19_JE-En-G2	2.66	0.61
20090910-58	STD19_JE-En-G1	2.71	0.62
20090910-59	STD19_JE-En-G5	2.72	0.61
20090910-60	STD19_JE-En-G4	2.75	0.61
	Average (JE En)		

imary ion beam)

d25Mg-m ‰	2SE ‰	d26Mg-m ‰	2SE ‰	d26Mg*-m ‰	2SE ‰	f ₂₅ *
-2.98	0.04	-5.47	0.07	0.33	0.05	-2.91
-2.85	0.04	-5.21	0.06	0.34	0.05	-2.78
-2.88	0.04	-5.24	0.05	0.37	0.05	-2.81
-2.91	0.04	-5.35	0.07	0.32	0.05	-2.84
-2.96	0.04	-5.45	0.07	0.31	0.05	-2.89
-2.91	0.03	-5.31	0.05	0.36	0.06	-2.84
-2.91	0.03	-5.31	0.05	0.37	0.05	-2.84
-2.88	0.03	-5.24	0.06	0.37	0.05	-2.81
-2.91	0.08	-5.32	0.19	0.35	0.04	-2.84
-4.22	0.06	-7.79	0.10	0.41	0.04	-3.85
-4.15	0.05	-7.68	0.10	0.40	0.04	-3.79
-4.15	0.05	-7.70	0.09	0.38	0.06	-3.78
-4.19	0.05	-7.82	0.09	0.33	0.06	-3.82
-4.19	0.05	-7.83	0.09	0.32	0.06	-3.82
-4.13	0.04	-7.69	0.08	0.35	0.05	-3.76
-4.15	0.05	-7.69	0.09	0.38	0.04	-3.78
-4.17	0.06	-7.74	0.14	0.37	0.07	-3.80
-3.07	0.04	-5.64	0.05	0.34	0.06	-3.05
-3.12	0.03	-5.73	0.06	0.34	0.04	-3.09
-3.14	0.04	-5.77	0.06	0.35	0.04	-3.11
-3.14	0.04	-5.72	0.06	0.39	0.04	-3.11
-3.12	0.06	-5.71	0.10	0.36	0.05	-3.09
-0.66	0.02	-0.97	0.03	0.31	0.05	-0.65
-0.50	0.02	-0.69	0.02	0.28	0.04	-0.48
-0.47	0.02	-0.64	0.02	0.27	0.04	-0.45
-0.58	0.03	-0.87	0.02	0.26	0.06	-0.56
-0.55	0.17	-0.79	0.31	0.28	0.05	-0.54
-3.05	0.04	-5.60	0.07	0.34	0.05	-2.98
-3.09	0.05	-5.71	0.08	0.31	0.04	-3.02

-3	.03	0.05	-5.56	0.08	0.34	0.04	-2.96
-3	.09	0.04	-5.67	0.07	0.35	0.04	-3.02
-3	.15	0.04	-5.79	0.07	0.34	0.04	-3.07
-3	.04	0.04	-5.61	0.06	0.31	0.04	-2.97
-2	.97	0.03	-5.50	0.06	0.28	0.04	-2.90
-3	.01	0.03	-5.51	0.06	0.35	0.04	-2.94
-3	.05	0.11	-5.62	0.20	0.33	0.05	-2.98
-0	.21	0.05	-0.10	0.10	0.32	0.05	-0.16
-0	.25	0.05	-0.15	0.09	0.35	0.05	-0.20
-0	.28	0.05	-0.20	0.09	0.36	0.04	-0.23
-0	.30	0.05	-0.22	0.09	0.36	0.04	-0.24
-0	.38	0.05	-0.31	0.09	0.43	0.04	-0.32
-0	.45	0.04	-0.53	0.09	0.35	0.05	-0.40
-0	.35	0.05	-0.33	0.09	0.36	0.05	-0.30
-0	.31	0.05	-0.31	0.09	0.30	0.06	-0.25
-0	.36	0.04	-0.39	0.07	0.32	0.05	-0.30
-0	.30	0.04	-0.30	0.07	0.27	0.05	-0.24
-0	.41	0.05	-0.45	0.08	0.35	0.04	-0.35
-0	.33	0.14	-0.30	0.26	0.34	0.08	-0.27
0	.40	0.06	1.18	0.10	0.39	0.04	0.44
0	.18	0.05	0.77	0.08	0.43	0.04	0.22
0	.40	0.05	1.19	0.09	0.42	0.06	0.44
0	.41	0.05	1.12	0.09	0.32	0.04	0.45
0	.35	0.23	1.07	0.40	0.39	0.10	0.39
0	.17	0.05	0.70	0.08	0.37	0.05	0.43
0	.22	0.04	0.76	0.08	0.33	0.05	0.48
0	.16	0.04	0.70	0.08	0.39	0.05	0.42
0	.15	0.04	0.66	80.0	0.38	0.04	0.40
0	.17	0.06	0.70	0.08	0.37	0.05	0.43

		Correc	cted			
2σ ^a	bias* _{25(RM-SCOI)}	$2\sigma^{b}$	f ₂₆ *	2σ ^a	bias* _{26(RM-SCOI)}	2σ ^b
0.40			E 04	0.10		
0.10			-5.31	0.10		
0.10			-5.00	0.17		
0.10			-5.09	0.17		
0.10			-5.20	0.17		
0.10			-5.50	0.17		
0.03			-5.15	0.17		
0.03			-5.10	0.17		
0.00			-5.17	0.17		
0.00	-1 02	0.13	-7 10	0.10	-1 64	0.26
0.10	-0.95	0.13	-6.99	0.17	-1 53	0.26
0.10	-0.95	0.13	-7.00	0.17	-1.55	0.26
0.10	-0.99	0.13	-7.13	0.16	-1.67	0.26
0.10	-0.99	0.13	-7.14	0.17	-1.68	0.26
0.10	-0.93	0.13	-6.99	0.16	-1.54	0.26
0.10	-0.95	0.13	-7.00	0.17	-1.54	0.26
0.06	-0.97	0.06	-7.05	0.14	-1.59	0.14
0.10	-0.21	0.13	-5.59	0.16	-0.13	0.26
0.10	-0.25	0.13	-5.68	0.16	-0.22	0.26
0.10	-0.28	0.13	-5.72	0.16	-0.25	0.26
0.10	-0.27	0.13	-5.67	0.16	-0.20	0.26
0.06	-0.25	0.06	-5.66	0.10	-0.20	0.11
0.07	2.20	0.13	-0.95	0.13	4.54	0.26
0.07	2.36	0.13	-0.67	0.13	4.83	0.26
0.06	2.39	0.13	-0.61	0.13	4.88	0.26
0.07	2.28	0.13	-0.85	0.13	4.64	0.26
0.17	2.31	0.17	-0.77	0.31	4.72	0.32

0.10	-5.45	0.17
0.10	-5.56	0.18

0.10			-5.41	0.18		
0.10			-5.52	0.18		
0.10			-5.64	0.17		
0.10			-5.46	0.17		
0.09			-5.35	0.17		
0.09			-5.35	0.17		
0.11			-5.47	0.20		
0.08	2.83	0.15	0.01	0.15	5.50	0.27
0.07	2.79	0.15	-0.04	0.14	5.45	0.27
0.07	2.76	0.15	-0.09	0.14	5.40	0.27
0.07	2.75	0.15	-0.12	0.15	5.38	0.27
0.08	2.67	0.15	-0.21	0.14	5.29	0.27
0.07	2.59	0.15	-0.42	0.14	5.07	0.27
0.07	2.69	0.15	-0.22	0.14	5.27	0.27
0.07	2.73	0.15	-0.20	0.14	5.29	0.27
0.07	2.68	0.15	-0.28	0.13	5.21	0.27
0.07	2.75	0.15	-0.20	0.13	5.30	0.27
0.07	2.64	0.15	-0.34	0.14	5.15	0.27
0.14	2.72	0.14	-0.19	0.26	5.30	0.26
0.10	3.43	0.15	1.25	0.17	6.75	0.27
0.09	3.21	0.15	0.84	0.15	6.34	0.27
0.10	3.43	0.15	1.26	0.16	6.76	0.27
0.09	3.44	0.15	1.19	0.16	6.69	0.27
0.23	3.38	0.23	1.14	0.40	6.64	0.40
0.07	3.42	0.15	1.21	0.14	6.71	0.27
0.07	3.47	0.15	1.26	0.14	6.77	0.27
0.07	3.41	0.15	1.20	0.14	6.70	0.27
0.07	3.39	0.15	1.17	0.14	6.67	0.27
0.06	3.42	0.06	1.21	0.08	6.71	0.08

Data reported in Ushikubo et al. (2017).GCA

^a 2 σ error (A) = $\sqrt{(2SE_{SIMS})^2 + (2SD_{ICPMS})^2}$ ^b 2 σ error = $\sqrt{(A)^2 + (2SD_{SIMS_bracket})^2}$ Erros of averaged values are 2SD.

File	Analysis spot	24Mg (100M cps)	Yield (10 ⁸ cps/nA)
Pyroxene	(Ushikubo et al., 2017.GCA))	
20130430-25	WI-STD19_SCOI#1	1.60	0.69
20130430-26	WI-STD19_SCOI#2	1.59	0.69
20130430-27	WI-STD19_SCOI#3	1.61	0.69
20130430-28	WI-STD19_SCOI#4	1.60	0.69
20130430-37	WI-STD19_SCOI#5	1.61	0.69
20130430-38	WI-STD19_SCOI#6	1.62	0.69
20130430-39	WI-STD19_SCOI#7	1.61	0.68
20130430-40	WI-STD19_SCOI#8	1.58	0.67
	Average (SC-OI)		
20130430-29	WI-STD19_95AK-6-G1#1	0.48	0.21
20130430-30	WI-STD19_95AK-6-G2#1	0.48	0.21
20130430-31	WI-STD19_95AK-6-G4#1	0.48	0.21
20130430-32	WI-STD19_95AK-6-G5#1	0.48	0.21
	Average (95AK-6 Di)		
20130430-33	WI-STD19_IG-CPx-G3#1	0.52	0.22
20130430-34	WI-STD19_IG-CPx-G2#1	0.52	0.22
20130430-35	WI-STD19_IG-CPx-G4#1	0.52	0.22
20130430-36	WI-STD19_IG-CPx-G5#1	0.51	0.22
	Average (IG Cpx)		

imary ion beam)

d25Mg-m ‰	2SE ‰	d26Mg-m ‰	2SE ‰	d26Mg*-m ‰	2SE ‰	f ₂₅ *
-2.90	0.08	-5.67	0.14	-0.01	0.06	-2.83
-2.83	0.07	-5.53	0.13	-0.01	0.06	-2.76
-2.91	0.07	-5.73	0.13	-0.07	0.07	-2.83
-2.94	0.06	-5.66	0.13	0.07	0.07	-2.87
-2.99	0.06	-5.77	0.12	0.05	0.08	-2.92
-3.04	0.07	-5.84	0.13	0.08	0.07	-2.97
-2.99	0.06	-5.84	0.11	-0.01	0.07	-2.92
-3.02	0.07	-5.87	0.12	0.01	0.06	-2.95
-2.95	0.14	-5.74	0.23	0.01	0.10	-2.88
0.09	0.09	0.08	0.11	-0.09	0.19	0.83
0.17	0.08	0.20	0.10	-0.14	0.15	0.91
0.13	0.11	0.20	0.09	-0.05	0.21	0.87
-0.03	0.09	-0.20	0.11	-0.14	0.16	0.71
0.09	0.18	0.07	0.38	-0.10	0.08	0.83
1.13	0.08	2.29	0.10	0.09	0.15	1.27
1.23	0.06	2.39	0.10	-0.01	0.15	1.37
1.22	0.07	2.50	0.08	0.11	0.12	1.36
1.20	0.07	2.49	0.07	0.15	0.13	1.34
1.19	0.09	2.42	0.19	0.09	0.14	1.33

Appendix EA7-1: SIMS major and minor element analyses data (sessio

^a 2σ errors represent either 2SE of each RM or 2SD of the runnning standard (SC-OI), whiche ^b 2SD errors were deduced from 5 analyses of each grain.

File	Analysis spot	Fo#	Cycle#	
20190124@1.asc	WI-STD-92 SCOL	88.8	5	
20190124@5.asc	WI-STD-91 SCOL	88.8	5	
20190124@13.asc	WI-STD-91 SCOL	88.8	5	
20190124@22.asc	WI-STD-91 SCOL	88.8	5	
	Average (SC-OI)			
	2SD			
	2SD %			
20190124@21.asc	WI-STD-91 OROL17_1	57.9	5	
20190124@20.asc	WI-STD-91 OROL5_1	60.3	5	
20190124@19.asc	WI-STD-91 FJOL2_1	73.2	5	
20190124@18.asc	WI-STD-91 KNOL1-2_1	77.9	5	
20190124@3.asc	WI-STD-92 CL0919-1_1	79.5	5	
20190124@2.asc	WI-STD-92 CL0908-1_1	80.8	5	
20190124@17.asc	WI-STD-91 SWOL15_1	81.8	5	
20190124@24.asc	WI-STD-92 A77257-4_1	85.6	5	
20190124@6.asc	WI-STD-91 A77257-5_1	85.8	5	
20190124@4.asc	WI-STD-92 CL0933-4_1	86.5	5	
	Average (SC-OI)_Errors are 2SD	88.8		
20190124@16.asc	WI-STD-91 UWOL1-42_1	89.3	5	
20190124@23.asc	WI-STD-92 IGOL1_1	89.6	5	
20190124@15.asc	WI-STD-91 HaKOL g1-2_1	91.9	5	
20190124@12.asc	WI-STD-91 WKOL8_1	93.7	5	
20190124@11.asc	WI-STD-91 WKOL1_1	94.7	5	
20190124@7.asc	WI-STD-91 WNOL1E_1	94.9	5	
20190124@8.asc	WI-STD-91 WNOL1C_1	95.5	5	
20190124@9.asc	WI-STD-91 WNOL4A_1	96.0	5	
20190124@10.asc	WI-STD-91 WNOL4C_1	96.1	5	
20190124@25.asc	WI-STD-92 NWA7325 OL7_1	97.4	5	
20190124@26.asc	WI-STD-15 Skye-X_1	99.8	5	
20190124@28.asc	WI-STD-15 Skye-V_1_5 cycle	99.8	5	
20190124@27.asc	WI-STD-15 Skye-W_1_5 cycle	99.8	5	
20190124@14.asc	WI-STD-91 HNOL	100.0	5	
Averaged values of each olivien RMs				
	1 OR-OI	59.1		
	2 FJ-OI	73.2		

3	KN-OI	77.9
4	CL09-19	79.5
5	CL09-08	80.8
6	SW-OI	81.8
7	A77257-OI	85.7
8	CL09-33	86.5
9	Average (SC-OI)_Errors are 2SD	88.8
10	UWOL-1	89.3
11	IG-OI	89.6
12	HaK-OI	91.9
13	WK-OI	94.2
14	WN-OI	95.6
15	N7325-OI	97.4
16	SK-OI	99.8
17	HN-OI	100.0

1.55 1.4 1.45 1.45 1.45 1.45 1.2 1.2 1.2 1.15 0.2 0.19 0.18 0.17 0.16 0.2

L 55 0.45 0.4 0.35 0.3 0.25 0.2 0.15 0.15 0.05 0 L 55

In FE1, 6pA, O_2^- primary ion beam)

ver is larger

 24 Mg (10⁵ cps) 23Na 24Mg 27AI 28Si 40Ca 8.31 1.6.E-03 1.4.E+00 7.5.E-04 1.8.E-01 4.5.E-03 8.60 9.5.E-04 1.5.E+00 6.9.E-04 1.8.E-01 3.8.E-03 8.62 9.3.E-04 1.5.E+00 7.0.E-04 1.8.E-01 3.8.E-03 8.59 9.5.E-04 1.5.E+00 7.4.E-04 1.8.E-01 3.8.E-03 8.53 1.1.E-03 1.4.E+00 7.2.E-04 1.8.E-01 4.0.E-03 0.29 6.3.E-04 5.5.E-02 5.4.E-05 2.1.E-03 7.1.E-04 3% 57% 4% 8% 1% 18% 1.2.E+00 4.2.E-04 1.7.E-01 7.26 2.5.E-04 4.5.E-03 4.2.E-04 7.50 3.1.E-04 1.3.E+00 1.7.E-01 2.4.E-03 8.47 2.0.E-04 1.4.E+00 6.9.E-04 1.6.E-01 3.0.E-03 8.54 8.0.E-04 1.4.E+00 1.3.E-03 1.6.E-01 1.8.E-02 8.42 1.4.E-03 1.4.E+00 8.2.E-04 1.5.E-01 4.2.E-03 8.51 3.5.E-03 1.0.E-03 1.4.E+00 5.5.E-04 1.5.E-01 8.88 5.3.E-04 9.2.E-05 1.5.E+00 4.0.E-04 1.6.E-01 8.56 1.3.E-03 1.5.E+00 1.0.E-03 1.5.E-01 1.4.E-02 8.93 7.2.E-04 1.5.E+00 1.1.E-03 1.5.E-01 1.5.E-02 8.31 1.5.E-03 1.4.E+00 6.9.E-04 1.7.E-01 2.9.E-03 8.53 1.1.E-03 1.4.E+00 7.2.E-04 1.8.E-01 4.0.E-03 8.68 9.5.E-04 1.5.E+00 8.5.E-04 1.8.E-01 3.1.E-03 8.07 9.2.E-04 1.4.E+00 5.8.E-04 1.8.E-01 3.4.E-03 8.49 9.5.E-04 1.4.E+00 3.5.E-04 1.9.E-01 3.5.E-03 8.52 4.5.E-04 1.4.E+00 1.6.E-03 1.8.E-01 6.0.E-03 8.54 5.4.E-04 1.4.E+00 1.6.E-03 1.8.E-01 7.2.E-03 8.51 1.0.E-04 1.4.E+00 7.2.E-04 1.9.E-01 3.3.E-04 8.42 5.7.E-04 3.0.E-04 8.1.E-05 1.4.E+00 1.9.E-01 8.65 8.6.E-05 1.5.E+00 3.5.E-04 1.9.E-01 2.4.E-04 8.47 9.6.E-05 1.4.E+00 4.4.E-04 1.9.E-01 3.1.E-04 8.53 5.3.E-04 1.5.E+00 1.5.E-03 1.7.E-01 1.5.E-02 8.04 3.5.E-05 1.3.E+00 6.0.E-05 1.7.E-01 1.8.E-03 8.14 4.1.E-05 1.4.E+00 4.9.E-05 1.8.E-01 1.4.E-03 8.06 4.5.E-05 1.4.E+00 7.9.E-05 1.8.E-01 1.0.E-03 8.09 1.4.E+00 1.8.E-01 4.8.E-05 1.7.E-04 1.2.E-04 7.38 2.8.E-04 4.2.E-04 3.4.E-03 1.3.E+00 1.7.E-01 8.47 2.0.E-04 1.4.E+00 6.9.E-04 1.6.E-01 3.0.E-03

Yield (× 10⁸ cps n/

8.54	8.0.E-04	1.4.E+00	1.3.E-03	1.6.E-01	1.8.E-02
8.42	1.4.E-03	1.4.E+00	8.2.E-04	1.5.E-01	4.2.E-03
8.51	1.0.E-03	1.4.E+00	5.5.E-04	1.5.E-01	3.5.E-03
8.88	9.2.E-05	1.5.E+00	4.0.E-04	1.6.E-01	5.3.E-04
8.75	9.9.E-04	1.5.E+00	1.0.E-03	1.5.E-01	1.4.E-02
8.31	1.5.E-03	1.4.E+00	6.9.E-04	1.7.E-01	2.9.E-03
8.53	1.1.E-03	1.4.E+00	7.2.E-04	1.8.E-01	4.0.E-03
8.68	9.5.E-04	1.5.E+00	8.5.E-04	1.8.E-01	3.1.E-03
8.07	9.2.E-04	1.4.E+00	5.8.E-04	1.8.E-01	3.4.E-03
8.49	9.5.E-04	1.4.E+00	3.5.E-04	1.9.E-01	3.5.E-03
8.53	5.0.E-04	1.4.E+00	1.6.E-03	1.8.E-01	6.6.E-03
8.51	9.2.E-05	1.4.E+00	5.2.E-04	1.9.E-01	2.9.E-04
8.53	5.3.E-04	1.5.E+00	1.5.E-03	1.7.E-01	1.5.E-02
8.08	4.0.E-05	1.4.E+00	6.3.E-05	1.8.E-01	1.4.E-03
8.09	4.8.E-05	1.4.E+00	1.7.E-04	1.8.E-01	1.2.E-04





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52Cr	55Mn	56Fe	60Ni	23Na+/ 28Si+
1.5.E-04	1.9.E-03	8.7.E-02	5.3.E-04	8.8.E-03
1.6.E-04	2.0.E-03	8.9.E-02	5.3.E-04	5.3.E-03
1.7.E-04	2.0.E-03	8.9.E-02	5.6.E-04	5.2.E-03
1.7.E-04	1.9.E-03	8.8.E-02	5.3.E-04	5.3.E-03
1.6.E-04	1.9.E-03	8.8.E-02	5.4.E-04	6.2.E-03
1.6.E-05	5.2.E-05	1.5.E-03	2.8.E-05	3.6.E-03
10%	3%	2%	5%	58%
7.1.E-07	1.2.E-02	4.3.E-01	6.3.E-05	1.5.E-03
1.1.E-06	1.2.E-02	4.0.E-01	7.5.E-05	1.8.E-03
1.1.E-06	6.4.E-03	2.5.E-01	2.7.E-04	1.2.E-03
5.2.E-03	6.3.E-03	2.0.E-01	1.9.E-05	5.1.E-03
3.3.E-05	3.1.E-03	1.8.E-01	3.3.E-04	9.1.E-03
3.0.E-05	2.9.E-03	1.7.E-01	4.6.E-04	6.8.E-03
1.7.E-04	4.4.E-03	1.6.E-01	5.3.E-06	5.9.E-04
5.2.E-03	5.8.E-03	1.2.E-01	1.1.E-05	8.4.E-03
5.2.E-03	5.9.E-03	1.2.E-01	9.2.E-06	4.7.E-03
5.9.E-05	2.2.E-03	1.1.E-01	6.4.E-04	9.0.E-03
1.6.E-04	1.9.E-03	8.8.E-02	5.4.E-04	6.2.E-03
7.6.E-05	1.8.E-03	8.4.E-02	5.5.E-04	5.2.E-03
4.8.E-05	1.8.E-03	7.9.E-02	5.4.E-04	5.2.E-03
5.2.E-05	1.3.E-03	6.2.E-02	5.9.E-04	5.0.E-03
2.1.E-03	1.2.E-03	5.0.E-02	5.6.E-04	2.6.E-03
1.7.E-03	1.0.E-03	4.1.E-02	5.7.E-04	3.0.E-03
1.5.E-04	3.8.E-03	3.4.E-02	1.1.E-06	5.6.E-04
1.4.E-04	3.9.E-03	3.1.E-02	7.1.E-07	4.4.E-04
1.4.E-04	3.8.E-03	2.9.E-02	0.0.E+00	4.5.E-04
1.1.E-04	3.6.E-03	2.6.E-02	7.1.E-07	5.2.E-04
3.0.E-03	9.4.E-04	2.0.E-02	1.1.E-06	3.0.E-03
2.1.E-06	5.3.E-04	2.0.E-03	0.0.E+00	2.0.E-04
1.1.E-06	4.7.E-04	2.0.E-03	0.0.E+00	2.3.E-04
2.1.E-06	4.2.E-04	1.8.E-03	3.5.E-07	2.5.E-04
1.1.E-06	2.5.E-06	3.4.E-05	0.0.E+00	2.6.E-04
8.9.E-07	1.2.E-02	4.1.E-01	6.9.E-05	1.6.E-03
1.1.E-06	6.4.E-03	2.5.E-01	2.7.E-04	1.2.E-03

5.2.E-03	6.3.E-03	2.0.E-01	1.9.E-05	5.1.E-03
3.3.E-05	3.1.E-03	1.8.E-01	3.3.E-04	9.1.E-03
3.0.E-05	2.9.E-03	1.7.E-01	4.6.E-04	6.8.E-03
1.7.E-04	4.4.E-03	1.6.E-01	5.3.E-06	5.9.E-04
5.2.E-03	5.8.E-03	1.2.E-01	9.9.E-06	6.5.E-03
5.9.E-05	2.2.E-03	1.1.E-01	6.4.E-04	9.0.E-03
1.6.E-04	1.9.E-03	8.8.E-02	5.4.E-04	6.2.E-03
7.6.E-05	1.8.E-03	8.4.E-02	5.5.E-04	5.2.E-03
4.8.E-05	1.8.E-03	7.9.E-02	5.4.E-04	5.2.E-03
5.2.E-05	1.3.E-03	6.2.E-02	5.9.E-04	5.0.E-03
1.9.E-03	1.1.E-03	4.5.E-02	5.7.E-04	2.8.E-03
1.4.E-04	3.8.E-03	3.0.E-02	6.2.E-07	4.9.E-04
3.0.E-03	9.4.E-04	2.0.E-02	1.1.E-06	3.0.E-03
1.8.E-06	4.7.E-04	1.9.E-03	1.2.E-07	2.3.E-04
1.1.E-06	2.5.E-06	3.4.E-05	0.0.E+00	2.6.E-04







24Mg+/ 28Si+	27Al+/ 28Si+	40Ca+/28Si+	52Cr+/ 28Si+	55Mn+/ 28Si+
7.9.E+00	4.2.E-03	2.5.E-02	8.7.E-04	1.1.E-02
8.1.E+00	3.9.E-03	2.1.E-02	9.1.E-04	1.1.E-02
8.2.E+00	3.9.E-03	2.2.E-02	9.3.E-04	1.1.E-02
8.1.E+00	4.1.E-03	2.1.E-02	9.7.E-04	1.1.E-02
8.1.E+00	4.0.E-03	2.2.E-02	9.2.E-04	1.1.E-02
2.6.E-01	3.2.E-04	4.2.E-03	8.5.E-05	2.9.E-04
3%	8%	19%	9%	3%
7.2.E+00	2.4.E-03	2.6.E-02	4.1.E-06	7.2.E-02
7.5.E+00	2.5.E-03	1.4.E-02	6.2.E-06	6.9.E-02
9.0.E+00	4.3.E-03	1.9.E-02	6.6.E-06	4.0.E-02
9.3.E+00	8.5.E-03	1.2.E-01	3.3.E-02	4.1.E-02
9.3.E+00	5.3.E-03	2.7.E-02	2.1.E-04	2.0.E-02
9.5.E+00	3.6.E-03	2.3.E-02	2.0.E-04	1.9.E-02
9.6.E+00	2.5.E-03	3.4.E-03	1.1.E-03	2.8.E-02
9.7.E+00	6.6.E-03	9.4.E-02	3.4.E-02	3.8.E-02
9.9.E+00	7.1.E-03	9.5.E-02	3.4.E-02	3.9.E-02
8.4.E+00	4.1.E-03	1.7.E-02	3.5.E-04	1.3.E-02
8.1.E+00	4.0.E-03	2.2.E-02	9.2.E-04	1.1.E-02
8.1.E+00	4.7.E-03	1.7.E-02	4.2.E-04	1.0.E-02
7.8.E+00	3.3.E-03	1.9.E-02	2.8.E-04	1.0.E-02
7.6.E+00	1.8.E-03	1.9.E-02	2.8.E-04	7.0.E-03
8.1.E+00	8.9.E-03	3.4.E-02	1.2.E-02	6.7.E-03
8.0.E+00	8.9.E-03	4.0.E-02	9.5.E-03	5.6.E-03
7.6.E+00	3.8.E-03	1.7.E-03	8.1.E-04	2.0.E-02
7.7.E+00	3.1.E-03	1.6.E-03	7.6.E-04	2.1.E-02
7.7.E+00	1.8.E-03	1.2.E-03	7.2.E-04	2.0.E-02
7.7.E+00	2.4.E-03	1.7.E-03	6.0.E-04	1.9.E-02
8.4.E+00	8.7.E-03	8.3.E-02	1.7.E-02	5.4.E-03
7.8.E+00	3.5.E-04	1.1.E-02	1.2.E-05	3.0.E-03
7.8.E+00	2.8.E-04	8.1.E-03	6.0.E-06	2.6.E-03
7.6.E+00	4.4.E-04	5.6.E-03	1.2.E-05	2.3.E-03
7.5.E+00	9.2.E-04	6.5.E-04	5.9.E-06	1.3.E-05
7.4.E+00	2.4.E-03	2.0.E-02	5.1.E-06	7.0.E-02
9.0.E+00	4.3.E-03	1.9.E-02	6.6.E-06	4.0.E-02
9.3.E+00	8.5.E-03	1.2.E-01	3.3.E-02	4.1.E-02
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9.3.E+00	5.3.E-03	2.7.E-02	2.1.E-04	2.0.E-02
9.5.E+00	3.6.E-03	2.3.E-02	2.0.E-04	1.9.E-02
9.6.E+00	2.5.E-03	3.4.E-03	1.1.E-03	2.8.E-02
9.8.E+00	6.9.E-03	9.5.E-02	3.4.E-02	3.8.E-02
8.4.E+00	4.1.E-03	1.7.E-02	3.5.E-04	1.3.E-02
8.1.E+00	4.0.E-03	2.2.E-02	9.2.E-04	1.1.E-02
8.1.E+00	4.7.E-03	1.7.E-02	4.2.E-04	1.0.E-02
7.8.E+00	3.3.E-03	1.9.E-02	2.8.E-04	1.0.E-02
7.6.E+00	1.8.E-03	1.9.E-02	2.8.E-04	7.0.E-03
8.1.E+00	8.9.E-03	3.7.E-02	1.1.E-02	6.1.E-03
7.7.E+00	2.8.E-03	1.6.E-03	7.2.E-04	2.0.E-02
8.4.E+00	8.7.E-03	8.3.E-02	1.7.E-02	5.4.E-03
7.7.E+00	3.6.E-04	8.1.E-03	1.0.E-05	2.7.E-03
7.5.E+00	9.2.E-04	6.5.E-04	5.9.E-06	1.3.E-05





					Error (:
56Fe+/ 28Si+	60Ni+/28Si+	23Na/ 28Si 2SE%	24Mg/ 28Si 2SE%	27Al/ 28Si 2SE%	40Ca/ 28Si 2SE%
4.9.E-01	3.0.E-03	3.37	3.32	5.03	1.68
4.9.E-01	3.0.E-03	4.87	3.17	11.10	3.61
5.0.E-01	3.1.E-03	5.96	2.86	11.97	3.62
4.9.E-01	2.9.E-03	4.48	3.56	8.24	6.00
4.9.E-01	3.0.E-03				
6.9.E-03	1.7.E-04				
 1%	6%				
2.5.E+00	3.7.E-04	17.10	4.43	13.17	9.09
2.3.E+00	4.5.E-04	13.09	4.68	25.42	3.95
1.6.E+00	1.7.E-03	7.30	3.91	6.97	3.69
1.3.E+00	1.2.E-04	9.95	4.29	2.94	3.76
1.2.E+00	2.1.E-03	12.27	3.92	4.66	2.01
1.1.E+00	3.0.E-03	16.70	4.13	15.33	7.93
1.0.E+00	3.4.E-05	15.55	3.82	13.47	7.02
7.9.E-01	7.1.E-05	9.13	3.83	6.86	3.67
7.9.E-01	6.0.E-05	5.37	4.46	8.98	4.38
6.4.E-01	3.8.E-03	14.25	4.40	9.70	2.93
4.9.E-01	3.0.E-03	57.83	3.19	8.02	18.71
4.6.E-01	3.0.E-03	3.11	2.72	7.09	3.04
4.5.E-01	3.1.E-03	19.57	3.61	9.37	3.13
3.3.E-01	3.1.E-03	14.62	2.54	12.81	13.18
2.8.E-01	3.2.E-03	4.53	3.33	5.66	4.22
2.2.E-01	3.2.E-03	5.30	3.86	3.29	2.84
1.8.E-01	5.6.E-06	12.75	3.24	19.80	6.22
1.7.E-01	3.7.E-06	24.61	2.73	20.78	7.78
1.5.E-01	0.0.E+00	12.23	2.45	21.60	6.35
1.4.E-01	3.8.E-06	13.18	2.01	14.89	9.58
1.1.E-01	6.2.E-06	7.09	2.93	7.14	3.41
1.2.E-02	0.0.E+00	12.13	2.15	19.43	3.60
1.1.E-02	0.0.E+00	19.23	1.60	16.83	2.14
1.0.E-02	2.0.E-06	30.55	2.93	27.14	5.63
 1.9.E-04	0.0.E+00	38.97	2.65	21.09	5.73
2.4.E+00	4.1.E-04	15.09	4.56	19.30	6.52
1.6.E+00	1.7.E-03	7.30	3.91	6.97	3.69

1.3.E+00	1.2.E-04	9.95	4.29	2.94	3.76
1.2.E+00	2.1.E-03	12.27	3.92	4.66	2.01
1.1.E+00	3.0.E-03	16.70	4.13	15.33	7.93
1.0.E+00	3.4.E-05	15.55	3.82	13.47	7.02
7.9.E-01	6.5.E-05	7.25	4.14	7.92	4.03
6.4.E-01	3.8.E-03	14.25	4.40	9.70	2.93
4.9.E-01	3.0.E-03	28.92	1.60	4.01	9.36
4.6.E-01	3.0.E-03	3.11	2.72	7.09	3.04
4.5.E-01	3.1.E-03	19.57	3.61	9.37	3.13
3.3.E-01	3.1.E-03	14.62	2.54	12.81	13.18
2.5.E-01	3.2.E-03	4.91	3.59	4.48	3.53
1.6.E-01	3.3.E-06	15.69	2.61	19.27	7.48
1.1.E-01	6.2.E-06	7.09	2.93	7.14	3.41
1.1.E-02	6.6.E-07	20.64	2.22	21.13	3.79
1.9.E-04	0.0.E+00	38.97	2.65	21.09	5.73

2SE%)						
52Cr/ 28Si 2SE%	55Mn/ 28Si 2SE%	56Fe/ 28Si 2SE%	60Ni/ 28Si 2SE%	23Na/ 28Si 2σ	24Mg/ 28Si 2σ	27Al/ 28Si 2σ
8.69	3.62	3.71	11.81			
9.66	4.05	3.08	4.58			
5.48	3.05	2.96	8.04			
6.88	4.26	3.88	6.16			
122.48	4.73	4.85	13.86	8.6.E-04	3.2.E-01	3.2.E-04
131.66	4.66	5.09	24.17	1.0.E-03	3.5.E-01	6.2.E-04
200.00	4.73	4.60	10.26	7.1.E-04	3.5.E-01	3.5.E-04
1.98	5.08	4.30	28.75	3.0.E-03	4.0.E-01	6.8.E-04
18.24	5.04	4.30	1.82	5.3.E-03	3.6.E-01	4.2.E-04
17.21	3.70	4.35	10.32	3.9.E-03	3.9.E-01	5.5.E-04
7.93	3.66	4.06	21.31	3.4.E-04	3.7.E-01	3.4.E-04
2.83	4.62	4.07	43.56	4.8.E-03	3.7.E-01	5.3.E-04
3.79	5.12	3.90	46.11	2.7.E-03	4.4.E-01	6.4.E-04
24.83	4.31	4.39	3.46	5.2.E-03	3.7.E-01	4.0.E-04
9.21	2.71	1.39	5.65	3.6.E-03	2.6.E-01	3.2.E-04
25.66	3.07	2.27	7.13	3.0.E-03	2.6.E-01	3.8.E-04
26.45	4.18	3.83	5.22	3.0.E-03	2.8.E-01	3.1.E-04
18.25	1.69	1.84	4.92	2.9.E-03	2.4.E-01	2.4.E-04
3.98	3.98	3.28	7.40	1.5.E-03	2.7.E-01	7.2.E-04
4.49	6.85	3.25	3.81	1.7.E-03	3.1.E-01	7.1.E-04
12.67	4.22	3.63	131.60	3.2.E-04	2.5.E-01	7.6.E-04
14.71	2.46	3.58	200.00	2.5.E-04	2.4.E-01	6.4.E-04
11.28	3.05	2.72		2.6.E-04	2.5.E-01	4.0.E-04
7.26	2.31	2.54	122.48	3.0.E-04	2.5.E-01	3.6.E-04
4.58	4.83	2.26	81.66	1.7.E-03	2.7.E-01	7.0.E-04
96.58	2.90	0.68		1.2.E-04	2.5.E-01	6.8.E-05
81.67	3.93	3.01		1.3.E-04	2.5.E-01	4.7.E-05
97.83	7.35	3.92	200.00	1.5.E-04	2.4.E-01	1.2.E-04
133.84	71.55	18.73		1.5.E-04	2.4.E-01	1.9.E-04
127.07	4.69	4.97	19.01			
200.00	4.73	4.60	10.26			

2SE%)

1.98	5.08	4.30	28.75
18.24	5.04	4.30	1.82
17.21	3.70	4.35	10.32
7.93	3.66	4.06	21.31
3.31	4.87	3.99	44.83
24.83	4.31	4.39	3.46
4.61	1.36	0.70	2.83
25.66	3.07	2.27	7.13
26.45	4.18	3.83	5.22
18.25	1.69	1.84	4.92
4.24	5.42	3.26	5.60
11.48	3.01	3.12	
4.58	4.83	2.26	81.66
92.03	4.73	2.54	
133.84	71.55	18.73	

Error (2	2σ) ^a	

40Ca/	52Cr/ 28Si	55Mn/	56Fe/ 28Si	60Ni/ 28Si	MaO	aonb
28Si 2σ	2σ	28Si 2σ	2σ	2σ	IvigO	250

4.9.E-03	5.0.E-06	3.4.E-03	1.2.E-01	5.1.E-05	27.55	0.13
2.6.E-03	8.1.E-06	3.2.E-03	1.2.E-01	1.1.E-04	29.07	0.27
3.5.E-03	1.3.E-05	1.9.E-03	7.2.E-02	1.7.E-04	37.13	0.19
2.2.E-02	3.1.E-03	2.1.E-03	5.4.E-02	3.5.E-05	39.67	0.16
5.1.E-03	3.9.E-05	1.0.E-03	5.1.E-02	1.2.E-04	41.41	0.37
4.4.E-03	3.4.E-05	7.1.E-04	4.8.E-02	3.1.E-04	42.35	0.28
6.4.E-04	1.0.E-04	1.0.E-03	4.1.E-02	7.2.E-06	42.86	0.28
1.8.E-02	3.1.E-03	1.8.E-03	3.2.E-02	3.1.E-05	44.96	0.28
1.8.E-02	3.1.E-03	2.0.E-03	3.1.E-02	2.8.E-05	44.95	0.41
3.2.E-03	8.8.E-05	5.7.E-04	2.8.E-02	2.1.E-04	46.14	0.34
4.2.E-03	8.5.E-05	2.9.E-04	6.9.E-03	1.7.E-04	48.22	0.25
3.2.E-03	1.1.E-04	3.1.E-04	1.1.E-02	2.2.E-04	48.30	0.30
3.6.E-03	7.3.E-05	4.2.E-04	1.7.E-02	1.7.E-04	48.50	0.16
3.5.E-03	5.1.E-05	1.9.E-04	6.0.E-03	1.8.E-04	49.94	0.17
6.3.E-03	1.1.E-03	2.7.E-04	9.2.E-03	2.3.E-04	51.35	0.43
7.5.E-03	8.8.E-04	3.8.E-04	7.3.E-03	1.8.E-04	52.04	0.33
3.3.E-04	1.0.E-04	8.5.E-04	6.5.E-03	7.4.E-06	53.52	0.43
3.1.E-04	1.1.E-04	5.7.E-04	6.0.E-03	7.5.E-06	53.69	0.14
2.3.E-04	8.1.E-05	6.0.E-04	4.2.E-03		53.99	0.63
3.1.E-04	5.5.E-05	5.2.E-04	3.6.E-03	4.6.E-06	54.38	0.36
1.6.E-02	1.6.E-03	2.6.E-04	2.5.E-03	5.0.E-06	54.69	0.25
2.0.E-03	1.2.E-05	8.8.E-05	1.6.E-04		56.58	0.33
1.5.E-03	4.9.E-06	1.0.E-04	3.4.E-04		56.91	0.10
1.0.E-03	1.2.E-05	1.7.E-04	4.0.E-04	4.0.E-06	56.84	0.26
1.2.E-04	7.9.E-06	9.6.E-06	3.5.E-05		57.31	0.22

Oxide wt% (EPMA) Data were obtained by the sess

SiO2	2SD ^b	CaO	2SD ^b	Cr2O3	2SD ^b	FeO
------	------------------	-----	------------------	-------	------------------	-----

34.94	0.08	0.037	0.014			35.77
35.53	0.21	0.036	0.024			34.16
37.08	0.18	0.059	0.004			24.21
38.07	0.39	0.377	0.036	0.80	0.09	20.07
38.26	0.27	0.079	0.027			18.98
38.22	0.36	0.047	0.015			17.93
38.40	0.49					16.95
39.03	0.29	0.318	0.031	0.82	0.05	13.45
39.12	0.33	0.315	0.021	0.77	0.12	13.29
39.31	0.40	0.054	0.024			12.86
40.01	0.42	0.070	0.012			10.82
39.75	0.30	0.064	0.024			10.32
39.91	0.16	0.050	0.030			10.08
39.92	0.28	0.025	0.017			7.81
40.57	0.58	0.132	0.020	0.24	0.10	6.13
40.60	0.09	0.124	0.014	0.22	0.07	5.22
41.22	0.47					5.16
41.04	0.28					4.49
41.34	0.22					4.01
41.27	0.24					3.89
41.61	0.22	0.316	0.030	0.45	0.07	2.58
41.27	0.18	0.054	0.013			0.23
41.72	0.46	0.032	0.014			0.19
41.73	0.28	0.030	0.011			0.16
42.04	0.48					

sion_2019Jan.

2SD ^b	NiO	2SD ^b	MnO	2SD ^b	Total
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-	0.10			0.665	0.073	98.96
	0.31			0.645	0.057	99.44
	0.26	0.14	0.03	0.429	0.030	99.05
	0.25			0.446	0.075	99.43
	0.38	0.18	0.08	0.222	0.134	99.13
	0.27	0.25	0.07	0.207	0.050	99.01
	0.34			0.321	0.091	98.53
	0.27			0.455	0.044	99.03
	0.44			0.461	0.075	98.90
	0.27	0.42	0.07	0.163	0.066	98.94
	0.27	0.42	0.12	0.159	0.073	99.69
	0.27	0.41	0.09	0.166	0.075	99.01
	0.21	0.40	0.05	0.143	0.063	99.07
	0.13	0.51	0.17	0.128	0.073	98.34
	0.69	0.44	0.12	0.122	0.075	98.99
	0.27	0.49	0.09	0.097	0.021	98.80
	0.56			0.337	0.078	100.36
	0.16			0.346	0.116	99.56
	0.18			0.297	0.055	99.65
	0.23			0.345	0.101	99.88
	0.16			0.125	0.100	99.77
	0.09			0.072	0.002	98.20
	0.06			0.083	0.042	98.93
	0.03			0.072	0.002	98.83
_						99.35

Appendix EA7-2: SIMS major and minor element analyses data (sessio

^a 2σ errors represent either 2SE of each RM or 2SD of the runnning standard (SC-OI), whiche ^b 2SD errors were deduced from 5 analyses of each grain.

File	Analysis spot	Fo content	Cycle#
20190427@5.asc	WI-STD-92 SCOL	88.8	5
20190427@6.asc	WI-STD-92 SCOL	88.8	5
20190427@14.asc	WI-STD-92 SCOL	88.8	5
20190427@15.asc	WI-STD-92 SCOL	88.8	5
20190427@21.asc	WI-STD-91 SCOL	88.8	5
20190427@27.asc	WI-STD-91 SCOL	88.8	5
20190427@33.asc	WI-STD-91 SCOL	88.8	5
20190427@43.asc	WI-STD-91 SCOL	88.8	5
20190427@44.asc	WI-STD-91 SCOL	88.8	5
20190427@45.asc	WI-STD-91 SCOL	88.8	5
	Average (SC-OI)		
	2SD		
	2SD %		
20190124@21.asc	WI-STD-91 OROL17_1	57.9	5
20190124@20.asc	WI-STD-91 OROL5_1	60.3	5
20190124@19.asc	WI-STD-91 FJOL2_1	73.2	5
20190124@18.asc	WI-STD-91 KNOL1-2_1	77.9	5
20190124@3.asc	WI-STD-92 CL0919-1_1	79.5	5
20190124@2.asc	WI-STD-92 CL0908-1_1	80.8	5
20190124@17.asc	WI-STD-91 SWOL15_1	81.8	5
20190124@24.asc	WI-STD-92 A77257-4_1	85.6	5
20190124@6.asc	WI-STD-91 A77257-5_1	85.8	5
20190124@4.asc	WI-STD-92 CL0933-4_1	86.5	5
	Average (SC-OI)_Errors are 2SD	88.8	
20190124@16.asc	WI-STD-91 UWOL1-42_1	89.3	5
20190124@23.asc	WI-STD-92 IGOL1_1	89.6	5
20190124@15.asc	WI-STD-91 HaKOL g1-2_1	91.9	5
20190124@12.asc	WI-STD-91 WKOL8_1	93.7	5
20190124@11.asc	WI-STD-91 WKOL1_1	94.7	5
20190124@7.asc	WI-STD-91 WNOL1E_1	94.9	5
20190124@8.asc	WI-STD-91 WNOL1C_1	95.5	5
20190124@9.asc	WI-STD-91 WNOL4A_1	96.0	5
20190124@10.asc	WI-STD-91 WNOL4C_1	96.1	5
20190124@25.asc	WI-STD-92 NWA7325 OL7_1	97.4	5
20190124@26.asc	WI-STD-15 Skye-X_1	99.8	5
20190124@28.asc	WI-STD-15 Skye-V_1_5 cycle	99.8	5

20190124@27.asc	WI-STD-15 Skye-W_1_5 cycle	99.8	5							
20190124@14.asc	WI-STD-91 HNOL	100.0	5							
Averaged values of each olivien RMs										
1	OR-OI	59.1								
2	FJ-OI	73.2								
3	KN-OI	77.9								
4	CL09-19	79.5								
5	CL09-08	80.8								
6	SW-OI	81.8								
7	A77257-OI	85.7								
8	CL09-33	86.5								
9	Average (SC-OI)_Errors are 2SD	88.8								
10	UWOL-1	89.3								
11	IG-OI	89.6								
12	HaK-Ol	91.9								
13	WK-OI	94.2								
14	WN-OI	95.6								
15	N7325-OI	97.4								
16	SK-OI	99.8								
17	HN-OI	100.0								

0.12 (Yu 0.115 0.11

500.0 50 ļ 0.2 0.18 0.16 0.14 0.12 0.0 0.04 0.04 0.02 0 0 55

n FE2, 16pA, O⁻ primary ion beam)

ver is larger

Yield (× 10⁸ cps n/

²⁴ Mg (10 ⁵ cps)	23Na	24Mg	27AI	28Si	40Ca
8.44	4.5.E-04	5.6.E-01	4.7.E-04	1.1.E-01	1.6.E-03
8.51	4.4.E-04	5.6.E-01	4.4.E-04	1.1.E-01	1.5.E-03
8.72	4.5.E-04	5.8.E-01	4.7.E-04	1.2.E-01	1.6.E-03
8.46	4.4.E-04	5.6.E-01	4.6.E-04	1.1.E-01	1.5.E-03
8.44	4.4.E-04	5.6.E-01	6.5.E-04	1.1.E-01	1.9.E-03
8.71	4.6.E-04	5.7.E-01	6.8.E-04	1.2.E-01	2.0.E-03
8.70	4.6.E-04	5.7.E-01	6.6.E-04	1.1.E-01	1.9.E-03
8.68	4.6.E-04	5.6.E-01	6.5.E-04	1.1.E-01	1.9.E-03
8.91	4.4.E-04	5.7.E-01	8.4.E-04	1.1.E-01	2.1.E-03
8.85	4.7.E-04	5.7.E-01	6.4.E-04	1.2.E-01	1.9.E-03
8.64	4.5.E-04	5.7.E-01	6.0.E-04	1.1.E-01	1.8.E-03
0.34	2.2.E-05	1.5.E-02	2.6.E-04	3.3.E-03	4.2.E-04
4%	5%	3%	44%	3%	23%
7.71	7.3.E-05	5.0.E-01	1.2.E-03	1.0.E-01	1.3.E-03
7.73	9.1.E-05	5.1.E-01	1.1.E-03	9.7.E-02	1.2.E-03
8.99	1.2.E-04	5.9.E-01	9.1.E-04	9.1.E-02	1.5.E-03
9.05	3.4.E-04	5.9.E-01	1.3.E-03	8.6.E-02	8.0.E-03
8.76	4.6.E-04	5.8.E-01	4.5.E-04	9.6.E-02	1.6.E-03
8.66	2.2.E-04	5.8.E-01	3.2.E-04	1.0.E-01	9.0.E-04
8.85	5.9.E-05	5.8.E-01	7.5.E-04	1.1.E-01	4.8.E-04
8.87	3.5.E-04	5.9.E-01	5.0.E-04	8.9.E-02	6.4.E-03
9.06	3.1.E-04	5.9.E-01	2.6.E-03	8.9.E-02	6.1.E-03
8.53	4.5.E-04	5.7.E-01	3.7.E-04	1.1.E-01	1.8.E-03
8.64	0.00	0.57	0.00	0.11	0.00
8.50	4.4.E-04	5.6.E-01	7.1.E-04	1.1.E-01	1.5.E-03
8.55	1.3.E-04	5.7.E-01	3.8.E-04	1.1.E-01	1.0.E-03
8.47	1.1.E-04	5.6.E-01	7.6.E-04	1.1.E-01	7.1.E-04
8.46	2.3.E-04	5.6.E-01	1.1.E-03	1.1.E-01	2.5.E-03
8.48	2.2.E-04	5.6.E-01	1.2.E-03	1.0.E-01	2.5.E-03
8.73	7.4.E-05	5.8.E-01	8.4.E-04	1.1.E-01	3.0.E-04
8.54	7.2.E-05	5.6.E-01	7.5.E-04	1.1.E-01	3.4.E-04
8.50	7.1.E-05	5.6.E-01	6.9.E-04	1.1.E-01	3.7.E-04
8.55	8.6.E-05	5.6.E-01	6.9.E-04	1.1.E-01	4.2.E-04
8.92	2.4.E-04	5.8.E-01	9.3.E-04	1.0.E-01	5.6.E-03
8.22	8.4.E-05	5.4.E-01	6.2.E-04	9.7.E-02	9.6.E-04
8.53	5.0.E-05	5.5.E-01	3.1.E-04	1.0.E-01	4.6.E-04

8.29	5.2.E-05	5.4.E-01	3.6.E-04	9.8.E-02	5.4.E-04
 8.43	5.7.E-05	5.5.E-01	5.3.E-04	1.0.E-01	2.7.E-04
7.72	8.2.E-05	5.1.E-01	1.2.E-03	9.8.E-02	1.3.E-03
8.99	1.2.E-04	5.9.E-01	9.1.E-04	9.1.E-02	1.5.E-03
9.05	3.4.E-04	5.9.E-01	1.3.E-03	8.6.E-02	8.0.E-03
8.76	4.6.E-04	5.8.E-01	4.5.E-04	9.6.E-02	1.6.E-03
8.66	2.2.E-04	5.8.E-01	3.2.E-04	1.0.E-01	9.0.E-04
8.85	5.9.E-05	5.8.E-01	7.5.E-04	1.1.E-01	4.8.E-04
8.97	3.3.E-04	5.9.E-01	1.5.E-03	8.9.E-02	6.2.E-03
8.53	4.5.E-04	5.7.E-01	3.7.E-04	1.1.E-01	1.8.E-03
8.64	4.5.E-04	5.7.E-01	6.0.E-04	1.1.E-01	1.8.E-03
8.50	4.4.E-04	5.6.E-01	7.1.E-04	1.1.E-01	1.5.E-03
8.55	1.3.E-04	5.7.E-01	3.8.E-04	1.1.E-01	1.0.E-03
8.47	1.1.E-04	5.6.E-01	7.6.E-04	1.1.E-01	7.1.E-04
8.47	2.2.E-04	5.6.E-01	1.2.E-03	1.0.E-01	2.5.E-03
8.58	7.6.E-05	5.7.E-01	7.4.E-04	1.1.E-01	3.5.E-04
8.92	2.4.E-04	5.8.E-01	9.3.E-04	1.0.E-01	5.6.E-03
8.35	6.2.E-05	5.4.E-01	4.3.E-04	9.9.E-02	6.5.E-04
8.43	5.7.E-05	5.5.E-01	5.3.E-04	1.0.E-01	2.7.E-04









52Cr	55Mn	56Fe	60Ni	23Na+/ 28Si+
8 9 E-05	7 8 E-04	35 E-02	1 7 E-04	4 1 E-03
9.1 E-05	80 E-04	3.5 E-02	1.7.E 04 1.6 E-04	3 9 F-03
9.1.E-05 9.9 E-05	83E-04	3.6 E-02	1.0.E-04	3.9.E-03
9.5.E 05	79E-04	3.6 E-02	1.7.E 04 1.7 E-04	3.8 F-03
87E-05	80 E-04	3.6 E-02	1.7.E 04 1.7 E-04	3 9 E-03
8.8 E-05	8.3 E-04	3.7 E-02	1.7.E 04 1 9 E-04	4 0 F-03
89E-05	8 0 E-04	3.6 E-02	1.3.E 04	4.0.E 00
8.6 E-05	8 1 E-04	3.6 E-02	1.6 E-04	4 0 E-03
9.5 E-05	84 E-04	37 E-02	17 E-04	3.9 F-03
9.7.E-05	8.4.E-04	3.6.E-02	1.8.E-04	4.1.E-03
 9.2.E-05	8.1.E-04	3.6.E-02	1.7.E-04	4.0.E-03
8.9.E-06	4.2.E-05	1.3.E-03	1.7.E-05	1.6.E-04
10%	5%	4%	10%	4%
 2.3.E-06	5.3.E-03	1.8.E-01	2.6.E-05	7.3.E-04
1.6.E-06	4.8.E-03	1.6.E-01	2.6.E-05	9.4.E-04
1.8.E-06	2.8.E-03	1.1.E-01	1.0.E-04	1.3.E-03
3.0.E-03	2.7.E-03	8.4.E-02	7.3.E-06	4.0.E-03
2.0.E-05	1.3.E-03	7.4.E-02	1.1.E-04	4.8.E-03
2.3.E-05	1.2.E-03	6.6.E-02	1.5.E-04	2.2.E-03
1.1.E-04	1.9.E-03	6.3.E-02	2.1.E-06	5.7.E-04
3.1.E-03	2.4.E-03	4.9.E-02	4.1.E-06	4.0.E-03
3.1.E-03	2.5.E-03	4.9.E-02	2.6.E-06	3.5.E-03
3.1.E-05	9.3.E-04	4.4.E-02	2.1.E-04	4.1.E-03
0.00	0.00	0.04	0.00	0.00
4.1.E-05	7.5.E-04	3.3.E-02	1.7.E-04	3.9.E-03
2.5.E-05	7.6.E-04	3.3.E-02	1.8.E-04	1.2.E-03
2.9.E-05	5.5.E-04	2.5.E-02	1.8.E-04	9.9.E-04
1.0.E-03	4.6.E-04	1.8.E-02	1.7.E-04	2.2.E-03
9.5.E-04	3.9.E-04	1.7.E-02	1.7.E-04	2.1.E-03
6.4.E-05	1.6.E-03	1.4.E-02	1.4.E-07	6.9.E-04
6.8.E-05	1.6.E-03	1.2.E-02	2.7.E-07	6.8.E-04
6.5.E-05	1.5.E-03	1.1.E-02	4.1.E-07	6.7.E-04
5.5.E-05	1.5.E-03	1.1.E-02	1.4.E-07	8.2.E-04
1.5.E-03	4.0.E-04	7.9.E-03	1.4.E-07	2.3.E-03
1.6.E-06	2.0.E-04	8.6.E-04	0.0.E+00	8.7.E-04
1.3.E-06	1.7.E-04	7.3.E-04	1.3.E-07	4.9.E-04

1.3.E-06	1.9.E-04	8.2.E-04	0.0.E+00	5.3.E-04
1.8.E-06	9.6.E-07	1.7.E-05	0.0.E+00	5.6.E-04
2.0.E-06	5.1.E-03	1.7.E-01	2.6.E-05	8.3.E-04
1.8.E-06	2.8.E-03	1.1.E-01	1.0.E-04	1.3.E-03
3.0.E-03	2.7.E-03	8.4.E-02	7.3.E-06	4.0.E-03
2.0.E-05	1.3.E-03	7.4.E-02	1.1.E-04	4.8.E-03
2.3.E-05	1.2.E-03	6.6.E-02	1.5.E-04	2.2.E-03
1.1.E-04	1.9.E-03	6.3.E-02	2.1.E-06	5.7.E-04
3.1.E-03	2.5.E-03	4.9.E-02	3.4.E-06	3.7.E-03
3.1.E-05	9.3.E-04	4.4.E-02	2.1.E-04	4.1.E-03
9.2.E-05	8.1.E-04	3.6.E-02	1.7.E-04	4.0.E-03
4.1.E-05	7.5.E-04	3.3.E-02	1.7.E-04	3.9.E-03
2.5.E-05	7.6.E-04	3.3.E-02	1.8.E-04	1.2.E-03
2.9.E-05	5.5.E-04	2.5.E-02	1.8.E-04	9.9.E-04
9.8.E-04	4.3.E-04	1.8.E-02	1.7.E-04	2.1.E-03
6.3.E-05	1.6.E-03	1.2.E-02	2.4.E-07	7.1.E-04
1.5.E-03	4.0.E-04	7.9.E-03	1.4.E-07	2.3.E-03
1.4.E-06	1.9.E-04	8.0.E-04	4.5.E-08	6.3.E-04
1.8.E-06	9.6.E-07	1.7.E-05	0.0.E+00	5.6.E-04





24Mg+/28Si+	27Al+/ 28Si+	40Ca+/28Si+	52Cr+/28Si+	55Mn+/28Si+	
5.0.E+00	4.2.E-03	1.4.E-02	8.0.E-04	7.0.E-03	
5.0.E+00	3.9.E-03	1.4.E-02	8.0.E-04	7.1.E-03	
4.9.E+00	4.0.E-03	1.4.E-02	8.4.E-04	7.2.E-03	
4.9.E+00	4.0.E-03	1.3.E-02	8.2.E-04	6.9.E-03	
4.9.E+00	5.8.E-03	1.7.E-02	7.7.E-04	7.1.E-03	
5.0.E+00	5.9.E-03	1.7.E-02	7.6.E-04	7.1.E-03	
5.0.E+00	5.8.E-03	1.6.E-02	7.8.E-04	7.0.E-03	
4.9.E+00	5.7.E-03	1.6.E-02	7.5.E-04	7.0.E-03	
5.1.E+00	7.5.E-03	1.9.E-02	8.4.E-04	7.5.E-03	
4.9.E+00	5.6.E-03	1.6.E-02	8.4.E-04	7.2.E-03	
5.0.E+00	5.3.E-03	1.6.E-02	8.0.E-04	7.1.E-03	
1.4.E-01	2.3.E-03	3.7.E-03	7.2.E-05	3.2.E-04	
3%	44%	24%	9%	5%	
5.1.E+00	1.2.E-02	1.3.E-02	2.3.E-05	5.4.E-02	
5.2.E+00	1.2.E-02	1.3.E-02	1.7.E-05	5.0.E-02	
6.5.E+00	1.0.E-02	1.6.E-02	2.0.E-05	3.0.E-02	
6.9.E+00	1.5.E-02	9.3.E-02	3.5.E-02	3.1.E-02	
6.0.E+00	4.7.E-03	1.6.E-02	2.0.E-04	1.4.E-02	
5.7.E+00	3.2.E-03	9.0.E-03	2.3.E-04	1.2.E-02	
5.5.E+00	7.1.E-03	4.6.E-03	1.1.E-03	1.8.E-02	
6.6.E+00	5.6.E-03	7.2.E-02	3.4.E-02	2.8.E-02	
6.6.E+00	2.9.E-02	6.8.E-02	3.5.E-02	2.8.E-02	
5.1.E+00	3.4.E-03	1.6.E-02	2.8.E-04	8.4.E-03	
4.96	0.01	0.02	0.00	0.01	
5.0.E+00	6.4.E-03	1.4.E-02	3.6.E-04	6.6.E-03	
4.9.E+00	3.3.E-03	8.8.E-03	2.2.E-04	6.6.E-03	
4.9.E+00	6.8.E-03	6.3.E-03	2.6.E-04	4.9.E-03	
5.3.E+00	1.1.E-02	2.4.E-02	9.7.E-03	4.4.E-03	
5.3.E+00	1.2.E-02	2.4.E-02	9.0.E-03	3.8.E-03	
5.3.E+00	7.8.E-03	2.8.E-03	5.9.E-04	1.5.E-02	
5.3.E+00	7.1.E-03	3.1.E-03	6.3.E-04	1.5.E-02	
5.3.E+00	6.5.E-03	3.5.E-03	6.2.E-04	1.5.E-02	
5.3.E+00	6.6.E-03	4.0.E-03	5.2.E-04	1.4.E-02	
5.7.E+00	9.2.E-03	5.5.E-02	1.5.E-02	3.9.E-03	
5.5.E+00	6.4.E-03	9.9.E-03	1.7.E-05	2.1.E-03	
5.4.E+00	3.1.E-03	4.5.E-03	1.3.E-05	1.6.E-03	

5.4.E+00	3.7.E-03	5.5.E-03	1.4.E-05	2.0.E-03
5.4.E+00	5.2.E-03	2.6.E-03	1.7.E-05	9.3.E-06
52 E+00	12 E-02	13E-02	2 0 E-05	52 E-02
6.5 E+00	1.2.E 02	1.6.E.02	2.0.E 05	3.0 E-02
69E+00	1.5.E 02	9.3 E-02	3.5 E-02	3.1 E-02
6.0 E+00	4 7 E-03	1.6 F-02	2.0 E-04	1 4 F-02
5.7 E+00	3 2 E-03	9.0 E-03	2.3 E-04	1.4.E 02
5.5 E+00	7 1 E-03	4.6 E-03	1 1 F-03	1.2.2.02 1.8 F-02
6.6 F+00	1 7 F-02	7 0 F-02	3 5 F-02	2 8 F-02
5.1.E+00	3.4.E-03	1.6.E-02	2.8.E-04	8.4.E-03
5.0.E+00	5.3.E-03	1.6.E-02	8.0.E-04	7.1.E-03
5.0.E+00	6.4.E-03	1.4.E-02	3.6.E-04	6.6.E-03
4.9.E+00	3.3.E-03	8.8.E-03	2.2.E-04	6.6.E-03
4.9.E+00	6.8.E-03	6.3.E-03	2.6.E-04	4.9.E-03
5.3.E+00	1.1.E-02	2.4.E-02	9.4.E-03	4.1.E-03
5.3.E+00	7.0.E-03	3.3.E-03	5.9.E-04	1.5.E-02
5.7.E+00	9.2.E-03	5.5.E-02	1.5.E-02	3.9.E-03
5.4.E+00	4.4.E-03	6.6.E-03	1.5.E-05	1.9.E-03
5.4.E+00	5.2.E-03	2.6.E-03	1.7.E-05	9.3.E-06





					Error (2
56Fe+/ 28Si+	60Ni+/28Si+	23Na/ 28Si 2SE%	24Mg/ 28Si 2SE%	27Al/ 28Si 2SE%	40Ca/ 28Si 2SE%
3.1.E-01	1.5.E-03	1.71	3.93	17.07	6.55
3.1.E-01	1.4.E-03	3.60	2.63	8.60	2.87
3.1.E-01	1.4.E-03	4.69	3.38	9.48	4.50
3.1.E-01	1.5.E-03	6.77	3.91	13.54	5.00
3.2.E-01	1.5.E-03	5.97	4.20	14.60	4.60
3.2.E-01	1.7.E-03	6.85	4.20	15.65	4.09
3.2.E-01	1.5.E-03	2.40	4.51	11.55	3.11
3.2.E-01	1.4.E-03	7.75	2.31	11.84	1.55
3.3.E-01	1.5.E-03	3.89	2.58	15.01	1.45
3.2.E-01	1.6.E-03	3.07	2.17	11.65	1.02
3.2.E-01	1.5.E-03				
9.7.E-03	3 1.5.E-04				
3%	b 10%				
1.8.E+00	2.6.E-04	19.47	10.27	16.44	4.57
1.7.E+00) 2.6.E-04	10.83	7.54	13.20	4.74
1.2.E+00) 1.1.E-03	9.11	3.33	12.04	2.99
9.7.E-01	8.4.E-05	5.96	4.42	7.53	2.72
7.7.E-01	1.1.E-03	8.57	4.68	7.76	3.99
6.6.E-01	1.5.E-03	7.86	4.89	8.85	6.05
6.0.E-01	2.0.E-05	12.73	2.44	15.31	11.50
5.5.E-01	4.7.E-05	13.20	4.30	4.84	2.64
5.5.E-01	2.9.E-05	9.48	1.90	15.12	2.28
4.0.E-01	1.9.E-03	8.51	2.53	8.40	9.89
0.32	2. 0.00	4.15	2.73	44.18	23.62
3.0.E-01	1.5.E-03	6.72	5.03	14.99	3.67
2.9.E-01	1.6.E-03	12.31	2.86	12.66	5.19
2.2.E-01	1.6.E-03	10.82	1.68	17.97	4.09
1.7.E-01	1.6.E-03	11.71	2.92	6.38	2.35
1.6.E-01	1.6.E-03	5.04	2.11	8.36	1.69
1.3.E-01	1.2.E-06	20.63	2.20	14.23	13.03
1.2.E-01	2.6.E-06	17.06	5.47	16.41	10.24
1.1.E-01	3.9.E-06	28.13	3.72	22.52	14.09
1.1.E-01	1.3.E-06	14.35	2.40	20.04	10.07
7.8.E-02	2 1.3.E-06	7.98	4.65	10.10	4.09
8.8.E-03	0.0.E+00	24.14	3.17	15.30	5.20
7.1.E-03	1.3.E-06	26.21	5.74	15.76	4.75

8.3.E-03	0.0.E+00	22.48	4.43	10.89	5.10
1.6.E-04	0.0.E+00	36.59	4.33	18.20	11.96
17 E+00	26 F-04	15 15	8 90	14 82	4 65
1 2 E+00	1 1 F-03	9 11	3 33	12 04	2 99
97 F-01	84 E-05	5 96	4 42	7 53	2 72
7.7.E-01	1.1.E-03	8.57	4.68	7.76	3.99
6.6.E-01	1.5.E-03	7.86	4.89	8.85	6.05
6.0.E-01	2.0.E-05	12.73	2.44	15.31	11.50
5.5.E-01	3.8.E-05	11.34	3.10	9.98	2.46
4.0.E-01	1.9.E-03	8.51	2.53	8.40	9.89
3.2.E-01	1.5.E-03	2.08	1.36	22.09	11.81
3.0.E-01	1.5.E-03	6.72	5.03	14.99	3.67
2.9.E-01	1.6.E-03	12.31	2.86	12.66	5.19
2.2.E-01	1.6.E-03	10.82	1.68	17.97	4.09
1.7.E-01	1.6.E-03	8.38	2.52	7.37	2.02
1.1.E-01	2.2.E-06	20.04	3.45	18.30	11.86
7.8.E-02	1.3.E-06	7.98	4.65	10.10	4.09
8.1.E-03	4.5.E-07	24.28	4.44	13.99	5.02
1.6.E-04	0.0.E+00	36.59	4.33	18.20	11.96

52Cr/ 28Si 2SE%	55Mn/ 28Si 2SE%	56Fe/ 28Si 2SE%	60Ni/ 28Si 2SE%	23Na/ 28Si 2σ	24Mg/ 28Si 2σ	27Al/ 28Si 2σ
9.43	2.61	4.06	5.61			
5.85	6.86	2.69	9.71			
7.75	5.52	3.99	7.23			
6.16	4.69	3.61	8.40			
10.87	6.41	3.95	7.60			
11.56	6.40	4.07	8.27			
8.77	5.22	4.89	8.51			
5.46	3.20	1.71	5.42			
7.48	6.24	2.65	4.17			
6.46	4.23	2.27	4.81			

2SE%)

51.76	12.88	10.68	30.04	1.4.E-04	5.2.E-01	5.5.E-03
42.17	7.60	6.65	17.50	1.0.E-04	3.9.E-01	5.2.E-03
18.25	3.59	3.43	12.77	1.2.E-04	2.2.E-01	4.4.E-03
1.30	5.21	4.21	8.81	2.4.E-04	3.0.E-01	6.7.E-03
18.12	6.43	3.47	11.49	4.1.E-04	2.8.E-01	2.1.E-03
6.96	5.50	4.38	6.26	1.7.E-04	2.8.E-01	1.4.E-03
9.63	4.04	2.01	65.99	7.2.E-05	1.5.E-01	3.2.E-03
6.41	6.29	4.78	29.05	5.2.E-04	2.8.E-01	2.5.E-03
4.82	3.45	1.82	27.48	3.3.E-04	1.8.E-01	1.3.E-02
10.69	3.17	2.19	4.70	3.5.E-04	1.4.E-01	1.5.E-03
8.97	4.52	3.07	9.79	1.6.E-04	1.4.E-01	2.3.E-03
5.53	7.30	4.39	4.13	2.7.E-04	2.5.E-01	2.8.E-03
17.09	4.06	3.83	6.44	1.4.E-04	1.4.E-01	1.5.E-03
12.76	2.65	1.97	6.06	1.1.E-04	1.3.E-01	3.0.E-03
3.13	4.08	2.42	9.01	2.5.E-04	1.5.E-01	4.7.E-03
3.17	3.19	1.49	4.70	1.1.E-04	1.5.E-01	5.2.E-03
10.56	1.72	1.69	200.00	1.4.E-04	1.5.E-01	3.4.E-03
16.78	3.87	4.24	200.00	1.2.E-04	2.9.E-01	3.1.E-03
6.88	7.94	3.79	81.77	1.9.E-04	2.0.E-01	2.9.E-03
10.07	3.68	1.94	200.00	1.2.E-04	1.5.E-01	2.9.E-03
4.65	4.20	3.81	200.00	1.8.E-04	2.7.E-01	4.0.E-03
55.73	8.69	2.73		2.1.E-04	1.7.E-01	2.8.E-03
70.66	3.60	4.52	200.00	1.3.E-04	3.1.E-01	1.3.E-03

30.57	2.88	2.18		1.2.E-04	2.4.E-01	1.6.E-03
70.89	96.63	18.76		2.1.E-04	2.3.E-01	2.3.E-03
46.96	10.24	8 66	23 77			
18 25	3 50	3 4 3	10 77			
10.25	5.59	5.45	12.11			
1.30	5.21	4.21	8.81			
18.12	6.43	3.47	11.49			
6.96	5.50	4.38	6.26			
9.63	4.04	2.01	65.99			
5.61	4.87	3.30	28.26			
10.69	3.17	2.19	4.70			
4.48	2.26	1.54	4.89			
5.53	7.30	4.39	4.13			
17.09	4.06	3.83	6.44			
12.76	2.65	1.97	6.06			
3.15	3.63	1.96	6.85			
11.07	4.30	2.92	170.44			
4.65	4.20	3.81	200.00			
52.32	5.06	3.14				
70.89	96.63	18.76				

Error (2σ) ^a

40Ca/	52Cr/ 28Si	55Mn/	56Fe/ 28Si	60Ni/ 28Si	MaO	aopb
28Si 2σ	2σ	28Si 2σ	2σ	2σ	MgO	25D

3.1.E-03	1.2.E-05	6.9.E-03	1.9.E-01	7.8.E-05	27.55	0.13
3.0.E-03	7.1.E-06	3.8.E-03	1.1.E-01	4.6.E-05	29.07	0.27
3.8.E-03	3.6.E-06	1.4.E-03	4.0.E-02	1.4.E-04	37.13	0.19
2.2.E-02	3.2.E-03	1.6.E-03	4.1.E-02	8.2.E-06	39.67	0.16
3.8.E-03	3.7.E-05	8.9.E-04	2.7.E-02	1.3.E-04	41.41	0.37
2.1.E-03	2.0.E-05	6.6.E-04	2.9.E-02	1.5.E-04	42.35	0.28
1.1.E-03	1.1.E-04	8.0.E-04	1.8.E-02	1.3.E-05	42.86	0.28
1.7.E-02	3.1.E-03	1.7.E-03	2.6.E-02	1.4.E-05	44.96	0.28
1.6.E-02	3.1.E-03	1.2.E-03	1.7.E-02	7.9.E-06	44.95	0.41
3.8.E-03	3.0.E-05	3.8.E-04	1.2.E-02	1.8.E-04	46.14	0.34
3.7.E-03	7.2.E-05	3.2.E-04	9.7.E-03	1.5.E-04	48.22	0.25
3.2.E-03	3.3.E-05	4.8.E-04	1.3.E-02	1.5.E-04	48.30	0.30
2.1.E-03	3.7.E-05	3.0.E-04	1.1.E-02	1.5.E-04	48.50	0.16
1.5.E-03	3.3.E-05	2.2.E-04	6.7.E-03	1.6.E-04	49.94	0.17
5.6.E-03	8.7.E-04	2.0.E-04	5.4.E-03	1.6.E-04	51.35	0.43
5.7.E-03	8.1.E-04	1.7.E-04	4.9.E-03	1.6.E-04	52.04	0.33
6.5.E-04	6.3.E-05	6.7.E-04	4.0.E-03	2.5.E-06	53.52	0.43
7.4.E-04	1.1.E-04	6.9.E-04	4.9.E-03	5.2.E-06	53.69	0.14
8.3.E-04	5.5.E-05	1.2.E-03	4.0.E-03	3.2.E-06	53.99	0.63
9.4.E-04	5.3.E-05	6.5.E-04	3.2.E-03	2.5.E-06	54.38	0.36
1.3.E-02	1.3.E-03	1.8.E-04	3.0.E-03	2.6.E-06	54.69	0.25
2.3.E-03	9.3.E-06	1.8.E-04	2.7.E-04		56.58	0.33
1.1.E-03	9.4.E-06	7.3.E-05	3.2.E-04	2.7.E-06	56.91	0.10
1.3.E-03	4.2.E-06	8.9.E-05	2.6.E-04	56.84	0.26	
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6.3.E-04	1.2.E-05	9.0.E-06	3.1.E-05	57.31	0.22	

Oxide wt%	(EPMA) D	ata were	obtained	by the	sess
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SiO2	2SD ^b	CaO	2SD ^b	Cr2O3	2SD ^b	FeO
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34.94	0.08	0.037	0.014			35.77
35.53	0.21	0.036	0.024			34.16
37.08	0.18	0.059	0.004			24.21
38.07	0.39	0.377	0.036	0.80	0.09	20.07
38.26	0.27	0.079	0.027			18.98
38.22	0.36	0.047	0.015			17.93
38.40	0.49					16.95
39.03	0.29	0.318	0.031	0.82	0.05	13.45
39.12	0.33	0.315	0.021	0.77	0.12	13.29
39.31	0.40	0.054	0.024			12.86
40.01	0.42	0.070	0.012			10.82
39.75	0.30	0.064	0.024			10.32
39.91	0.16	0.050	0.030			10.08
39.92	0.28	0.025	0.017			7.81
40.57	0.58	0.132	0.020	0.24	0.10	6.13
40.60	0.09	0.124	0.014	0.22	0.07	5.22
41.22	0.47					5.16
41.04	0.28					4.49
41.34	0.22					4.01
41.27	0.24					3.89
41.61	0.22	0.316	0.030	0.45	0.07	2.58
41.27	0.18	0.054	0.013			0.23
41.72	0.46	0.032	0.014			0.19

41.73	0.28	0.030	0.011	0.16
42.04	0.48			

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2SD ^b	NiO	2SD ^b	MnO	2SD ^b	Total

0.1	0		0.665	0.073	98.96
0.3	1		0.645	0.057	99.44
0.20	6 0.14	0.03	0.429	0.030	99.05
0.2	5		0.446	0.075	99.43
0.3	8 0.18	0.08	0.222	0.134	99.13
0.2	7 0.25	0.07	0.207	0.050	99.01
0.34	4		0.321	0.091	98.53
0.2	7		0.455	0.044	99.03
0.44	4		0.461	0.075	98.90
0.2	7 0.42	0.07	0.163	0.066	98.94
0.2	7 0.42	0.12	0.159	0.073	99.69
0.2	7 0.41	0.09	0.166	0.075	99.01
0.2	1 0.40	0.05	0.143	0.063	99.07
0.13	3 0.51	0.17	0.128	0.073	98.34
0.69	9 0.44	0.12	0.122	0.075	98.99
0.2	7 0.49	0.09	0.097	0.021	98.80
0.5	6		0.337	0.078	100.36
0.10	6		0.346	0.116	99.56
0.18	8		0.297	0.055	99.65
0.23	3		0.345	0.101	99.88
0.10	6		0.125	0.100	99.77
0.0	9		0.072	0.002	98.20
0.0	6		0.083	0.042	98.93

0.03	0.072	0.002	98.83
			99.35

Appendix EA7-3: depth dependence of Mg⁺/Si⁺ ratios

We conducted longer major and minor element analyses (50 cycle O⁻). Present data are averaged by each 5 cycle to see depth deper

	Fo content	1-5 c	6-10 c	11-15 c
$6pA, O_2$ primary ion beam				
1 OR-OI	59	7.35	8.03	8.31
2 FJ-OI	73	8.97	9.81	10.16
3 KN-OI	78	9.28	10.24	10.60
4 CL09-19-OI	79	9.28	10.24	10.60
5 CL09-08-OI	81	9.48	10.26	10.62
6 SW-OI	82	9.65	10.42	10.96
7 A77257-OI	86	9.73	10.65	11.24
8 CL09-33-OI	87	8.40	9.07	9.52
9 SC-OI	89	8.13	8.80	9.10
10 UWOL-1	89	8.08	8.49	8.98
11 IG-OI	90	7.84	8.43	8.75
12 HaK-Ol	92	7.61	8.16	8.42
13 WK-OI	94	8.08	8.45	8.83
14 WN-OI	96	7.69	8.21	8.43
15 N7325-OI	97	8.35	8.73	8.90
16 SK-OI	100	7.77	8.13	8.43
17 HN-OI	100	7.55	7.92	8.25



16pA, O [−] pri	imary ion beam				
1 (OR-OI	59	5.16	6.21	6.89
2	FJ-OI	73	6.48	7.64	8.68
3	KN-OI	78	6.90	8.25	9.00
4 (CL09-19-OI	79	6.04	6.25	6.80
5 (CL09-08-OI	81	5.74	6.11	6.67
6 \$	SW-OI	82	5.54	6.10	6.36
7 /	A77257-OI	86	6.63	7.16	7.27
8 (CL09-33-OI	87	5.09	5.44	5.68
9 3	SC-OI	89	4.96	5.33	5.63
10	UWOL-1	89	4.96	5.34	5.66
11	IG-OI	90	4.95	5.44	5.54
12	HaK-OI	92	4.95	5.56	5.57
13 \	WK-OI	94	5.30	5.81	6.09
14 \	WN-OI	96	5.30	5.91	6.14
15 I	N7325-OI	97	5.72	6.27	6.64
16 \$	SK-OI	100	5.68	6.07	6.57
17	HN-OI	100	5.40	6.18	6.54



s) with two primary ion beam conditions (6 pA, O_2^- and 16 pA, ndences on Mg⁺/Si⁺ ratios. Errors are not shown.

	raw 24M	lg+/28Si+				
16-20 c	21-25 с	26-30 c	31-35 c	36-40 c	41-45 c	46-50 c
8.44	8.48	8.56	8.53	8.55	8.57	8.54
10.36	10.47	10.39	10.45	10.48	10.41	10.32
10.89	10.90	10.92	10.89	10.92	10.90	10.83
10.70	10.78	10.92	10.84	10.73	10.62	10.66
10.95	10.89	10.94	10.98	11.00	10.81	10.98
11.09	11.30	11.20	11.16	11.26	10.93	11.05
11.41	11.30	11.34	11.45	11.47	11.30	11.33
9.93	9.97	10.24	10.43	10.21	10.38	10.25
9.26	9.56	9.75	10.00	9.89	9.98	9.99
9.17	9.54	9.70	9.83	9.96	10.07	10.01
9.10	9.27	9.52	9.78	9.72	9.76	9.81
8.84	8.98	9.31	9.30	9.29	9.42	9.63
9.06	9.36	9.52	9.58	9.80	9.77	9.76
8.70	8.89	9.11	9.27	9.30	9.44	9.43
9.01	9.28	9.37	9.38	9.41	9.48	9.41
8.56	8.65	8.97	8.85	8.88	9.03	8.89
8.49	8.63	8.83	8.87	8.84	8.95	9.01



7.23	7.52	7.68	7.78	7.78	7.88	7.85
8.91	9.17	9.45	9.58	9.57	9.59	9.61
9.38	9.49	9.84	9.99	10.16	10.07	10.01
7.19	7.44	7.52	7.80	7.94	8.08	8.50
6.91	7.18	7.17	7.40	7.58	7.82	7.83
6.71	6.91	7.18	7.27	7.45	7.58	7.72
7.70	7.88	8.25	8.05	8.18	8.32	8.43
5.91	6.08	6.32	6.46	6.65	6.92	7.02
5.86	5.98	6.14	6.40	6.59	6.80	6.99
5.80	6.00	6.17	6.46	6.63	6.85	6.99
5.82	6.10	6.15	6.41	6.65	6.69	6.95
5.85	6.00	6.31	6.44	6.52	6.68	6.82
6.24	6.46	6.69	6.83	6.93	7.13	7.31
6.28	6.57	6.63	6.82	7.00	7.08	7.19
6.85	7.18	7.05	7.25	7.34	7.42	7.41
6.62	6.85	7.09	6.86	7.10	7.07	7.31
6.62	6.70	6.87	6.89	7.07	7.10	7.09

·3-3) raw data (16pA, O-)



Fo content

			24Mg	24Mg+/28Si+ normalized to SC-OI		
1-5 c	6-10 c	11-15 c	16-20 c	21-25 с	26-30 c	31-35 c
0.9	0 0.91	0.91	0.91	0.89	0.88	0.85
1.1	0 1.11	1.12	1.12	1.10	1.07	1.04
1.1	4 1.16	5 1.16	1.18	1.14	1.12	1.09
1.1	4 1.16	5 1.16	1.15	1.13	1.12	1.08
1.1	7 1.17	' 1.17	1.18	1.14	1.12	1.10
1.1	9 1.18	1.20	1.20	1.18	1.15	1.12
1.2	0 1.21	1.24	1.23	1.18	1.16	1.15
1.0	3 1.03	1.05	1.07	1.04	1.05	1.04
1.0	0 1.00	1.00	1.00	1.00	1.00	1.00
0.9	9 0.97	0.99	0.99	1.00	1.00	0.98
0.9	6 0.96	0.96	0.98	0.97	0.98	0.98
0.9	4 0.93	0.93	0.95	0.94	0.95	0.93
0.9	9 0.96	0.97	0.98	0.98	0.98	0.96
0.9	5 0.93	0.93	0.94	0.93	0.93	0.93
1.0	0.99	0.98	0.97	0.97	0.96	0.94
0.9	6 0.92	0.93	0.92	0.90	0.92	0.89
0.9	3 0.90	0.91	0.92	0.90	0.91	0.89



Appendix EA8: Calculation of residual biases for Cr-bearing olivine RMs

The modelled value (bias*_{25 (RM-SCOI)_cale}) is calculated by assuming linear correlation between bias*_{25 (RM-SCOI)} value and Fo content among olivine RMs with similar Fo contents.

$$bias^{*}_{25 (RM-SCOl)_calc} = \frac{(Fo_{high}) \times bias^{*}_{25 (RM-SCOl)_low} + (Fo_{low}) \times bias^{*}_{25 (RM-SCOl)_high}}{(Fo_{high}) - (Fo_{low})}$$

where Fo_{high} and Fo_{low} are Fo contents of each of two olivine RMs with higher and lower Fo contents relative to an olivine RM to be modelled. For example, if we estimate the bias*_{25 (RM-SCOI)_calc} value for KN-Ol (Fo_{78.0}), Fo_{high} and Fo_{low} are 73.4 (FJ-Ol) and 79.4 (CL09-19-Ol), respectively. Likewise, bias*_{25 (RM-SCOI)_high} and bias*_{25 (RM-SCOI)_low} are bias*_{25 (RM-SCOI)} values of each of two olivine RMs.