

The odd free surface flows of a colloidal chiral fluid

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In simple fluids, such as water, invariance under parity and time-reversal symmetry imposes that the rotation of constituent 'atoms' is determined by the flow and that viscous stresses damp motion. Activation of the rotational degrees of freedom of a fluid by spinning its atomic building blocks breaks these constraints and has thus been the subject of fundamental theoretical interest across classical and quantum fluids. However, the creation of a model liquid that isolates chiral hydrodynamic phenomena has remained experimentally elusive. Here, we report the creation of a cohesive two-dimensional chiral liquid consisting of millions of spinning colloidal magnets and study its flows. We find that dissipative viscous 'edge-pumping' is a key and general mechanism of chiral hydrodynamics, driving unidirectional surface waves and instabilities, with no counterpart in conventional fluids. Spectral measurements of the chiral surface dynamics suggest the presence of Hall viscosity, an experimentally elusive property of chiral fluids. Precise measurements and comparison with theory demonstrate excellent agreement with a minimal chiral hydrodynamic model, paving the way for the exploration of chiral hydrodynamics in experiment.

Hydrodynamic theories describe the flow of systems as diverse as water, quantum electronic states¹ and galaxies² over decades in scale³. Since hydrodynamic equations are built on symmetry principles and conservation laws alone, systems with similar symmetries have similar descriptions and flow in the same way.

For example, symmetry under parity and time reversal—conditions met by all conventional fluids at thermal equilibrium—constrains both the stress and viscosity tensors to be symmetric. These constraints are in principle alleviated in collections of interacting units that are driven to rotate^{4–10}. This seemingly innocent twist on an otherwise structureless fluid represents, however, an elemental change with rich hydrodynamic consequences common to quantum Hall fluids, vortex fluids and chiral condensed matter^{11–18}. Collections of spinning particles offer a natural opportunity to engineer and study the properties of such chiral fluids; experimental examples include rotating bacteria^{19,20}, colloidal and millimetre-scale magnets^{21–26}, ferrofluids in rotating magnetic fields^{27,28} and shaken chiral grains^{29,30}. Such systems have been shown to have non-trivial dynamics. For example, ferrofluids driven by a.c. fields can flow against external pressure³¹ and small numbers of spinning particles self-assemble into dynamic crystalline clusters^{21–26,32–35}.

A colloidal chiral fluid

We report the creation of a millimetre-scale cohesive chiral fluid (Fig. 1a) by spinning millions of colloidal magnets with a magnetic field (Fig. 1b,c), and we track its flows over hours (see Supplementary Videos 1 and 2). The macroscopic flow of our chiral fluid is reminiscent of free surface flows of Newtonian fluids: nearby droplets merge (Fig. 1d and Supplementary Video 3), fluid spreads on a surface under the influence of gravity (Fig. 1e and Supplementary Video 4), voids collapse (Fig. 1f and Supplementary Video 5) and thin streams

become unstable, as revealed by flowing fluid past a solid object (Fig. 1g and Supplementary Video 6). We demonstrate that these seemingly familiar features are accompanied by unique free surface flows. We then exploit the odd interfacial dynamics of this prototypical chiral liquid to infer its material constants, which remain out of reach of conventional rheology.

In contrast to Newtonian fluids, the surface of our fluid supports a spontaneous unidirectional edge flow in its steady state, as well as unusual morphological dynamics such as the rotation of asymmetric droplets. These features, illustrated in Fig. 1 and Supplementary Videos 1 and 3–6, follow from the breaking of parity through active rotation^{25,29,30,36}.

Chiral surface waves and 'edge-pumping'

To investigate these lively surface flows, we first look at surface excitations in a simple slab geometry, as shown in Fig. 2a and Supplementary Video 7. We measure the spectrum of surface fluctuations, $|h(k, \omega)|^2$, by tracing the height profile, $h(x, t)$, of the surface and Fourier-transforming it in space and time. We observe the spectrum to be peaked along a curve $\omega(k)$, revealing the existence of dispersive waves (see Fig. 2b). The curve has only one branch with odd parity, meaning that the waves are unidirectional. This behaviour contrasts with that of conventional surface waves that propagate in all directions.

These surface waves beg a hydrodynamic description. Chiral-fluid hydrodynamics follows from conservation of momentum and angular momentum, and thus includes both the spinning rate of individual fluid particles and the momentum and angular momentum of their flow^{6,29,37–39}. As our colloids are birefringent, we are able to measure their individual spinning rate by imaging through crossed polarizers. We find that all particles rotate at the same rate, Ω , which is set by the rotating magnetic field

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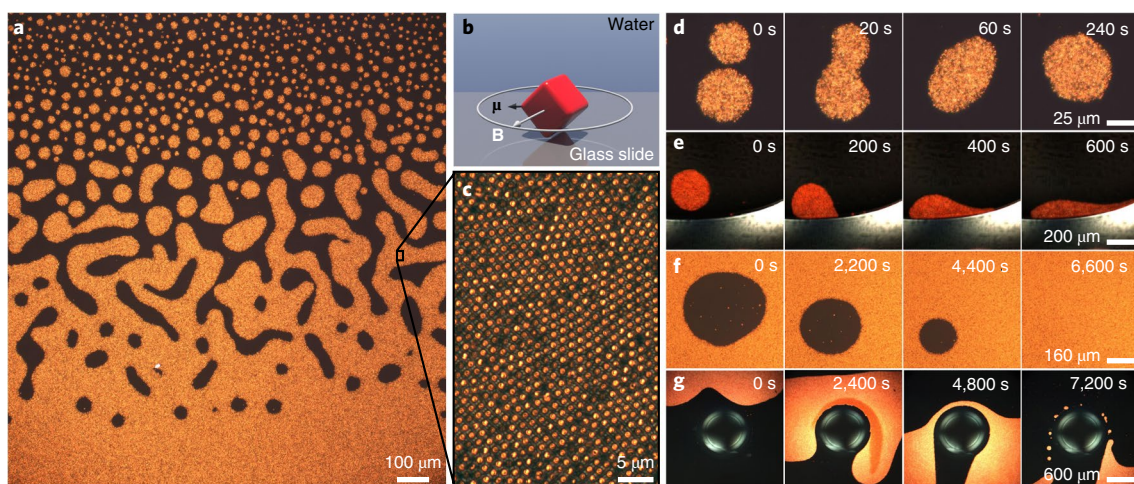


Fig. 1 | A chiral fluid of spinning colloidal magnets. **a**, An optical micrograph of the colloidal magnets in bulk, after a few minutes of spinning. **b**, A schematic diagram of one colloidal particle. The $\sim 1.6 \mu\text{m}$ haematite colloidal cubes have a permanent magnetic moment (μ , black arrow). They are suspended in water, sedimented onto a glass slide and spun by a rotating magnetic field (B , white arrow tracing the white circle). **c**, An optical micrograph of the colloidal magnets in bulk at increased magnification. **d–g**, The particles attract and form a cohesive material with an apparent surface tension that, over timescales from minutes to hours, behaves like a fluid: clusters coalesce (**d**) and spread like liquid droplets when sedimented against a hard wall (**e**); void bubbles collapse (**f**); and when driven past an obstacle, the fluid flows around it, thinning and eventually revealing an instability to droplet formation (**g**). All images were taken through crossed polarizers.

(see Fig. 3a and Supplementary Video 8). From this, it follows that the particles' rotational inertia is negligible; the torque exerted on each particle by the magnetic field instantly adjusts to balance the frictional torques exerted by the neighbouring particles, solid substrate and surrounding fluid layer. This fast response enables the decoupling of the angular momentum equation from the momentum equation. Nonetheless, a strong signature of the microscopic angular momentum manifests as an 'odd' stress. A minimal hydrodynamic theory to predict the chiral fluid velocity u_i then balances the force generated by viscous and odd hydrodynamic stresses, $\partial_j \sigma_{ij}$, against friction with the substrate, $\Gamma_{ij} u_j = (\Gamma_u \delta_{ij} + \Gamma_\perp \epsilon_{ij}) u_j$, and surface tension γ at the fluid interface. In this theory, which has been used to capture the bulk flows of chiral granular fluids, the hydrodynamic stress tensor is given by:

$$\sigma_{ij} = -p \delta_{ij} + \eta (\partial_i u_j + \partial_j u_i) + \eta_R \epsilon_{ij} (2\Omega - \omega) \quad (1)$$

σ_{ij} includes the pressure p and ordinary viscous stress also present in Newtonian fluids with a shear viscosity η . The additional term containing the Levi-Civita symbol ϵ_{ij} and the dissipative coefficient η_R , known interchangeably as 'rotational viscosity,' 'spin viscosity' or 'inter-rotor friction', captures the rotational friction between neighbouring particles^{6,13,27,29,37,38,40}. Such an odd stress builds up as the local spinning rate Ω deviates from half the local fluid vorticity $\omega = \hat{z} \cdot (\nabla \times \mathbf{u})$. In torque-free fluids, angular momentum conservation constrains these two quantities to be equal: odd stresses are unique to chiral fluids.

We finally simplify the model by assuming incompressibility and find this assumption to be supported by the agreement between theory and the experiments reported in this Article. We also note that there is no direct appearance of the magnetic field or its stresses in this hydrodynamic description unlike in conventional ferrofluids. In this respect, our colloidal chiral fluid can be seen as a special type of driven ferrofluid in which weakly Brownian particles are densely packed and magnetic forces provide cohesion and induce chirality (Supplementary Section 6).

To make a quantitative comparison between our model and the flows we observe, we require a measurement of the hydrodynamic and friction coefficients η , η_R , Γ_u and $\Gamma_\perp$. Fortunately, the prominent

effect of odd stress at the free surface of our chiral fluid can be effectively exploited to infer its bulk rheology. The homogeneous spinning motion of the colloidal particles gives rise to a net tangential edge flow even in the absence of pressure gradients. These tread-milling dynamics, characteristic of all chiral fluids^{4,9,19,25,29}, are illustrated in circular droplets in Fig. 3b–e and Supplementary Video 9. The tangential flow that is localized at the free surface is readily explained by expressing the hydrodynamic equation in terms of vorticity for an incompressible chiral fluid:

$$(\nabla^2 - \delta^{-2}) \omega = 0 \quad (2)$$

where $\delta = \sqrt{(\eta + \eta_R)/\Gamma_u}$. This Helmholtz equation indicates that the vorticity generated at the surface decays exponentially into the chiral fluid, with a characteristic penetration depth δ (see Fig. 3c,d,g). In this model, the loss of substrate friction causes the penetration depth to diverge, resulting in rigid-body rotation of the entire fluid, as observed in ferrofluid droplets³⁶. The magnitude of the vorticity at the free surface, $\omega_{\text{edge}} = 2\Omega\eta_R/(\eta + \eta_R)$, is set by the stress-free boundary condition for a flat strip and expresses the competition between the odd and viscous stresses (Supplementary Section 6). We point out that ω_{edge} is directly proportional to η_R , which demonstrates the importance of odd stress for the dynamics. Comparison between experiment and prediction (Fig. 3d) yields the values of η and η_R in terms of Γ_u . The latter is then measured by tilting the substrate and measuring the sedimentation rate of droplets (see Fig. 3f and Supplementary Section 3). Ultimately, we find $\eta = 4.9 \pm 0.2 \times 10^{-8} \text{ Pa s}$, $\eta_R = 9.1 \pm 0.1 \times 10^{-10} \text{ Pa s}$ and $\Gamma_u = 2.49 \pm 0.03 \times 10^3 \text{ Pa s m}^{-1}$. The sedimentation direction is aligned with gravity, which bounds $\Gamma_\perp \ll \Gamma_u$, so we take $\Gamma_\perp = 0$ in the following (Supplementary Section 3).

Equipped with the hydrodynamic coefficients, we can now investigate the origin of the surface waves within our model. The mass flux in the tangential surface flow provides significant insight. This flow, sketched in Fig. 2d and plotted in Fig. 2e,f, is determined by the balance of the tangential odd stress at the boundary, the shear stress and the substrate friction. In the presence of a perturbation to a free surface, the resistance to flow caused by the shear stress will be modulated in proportion to the curvature. For a sinusoidal

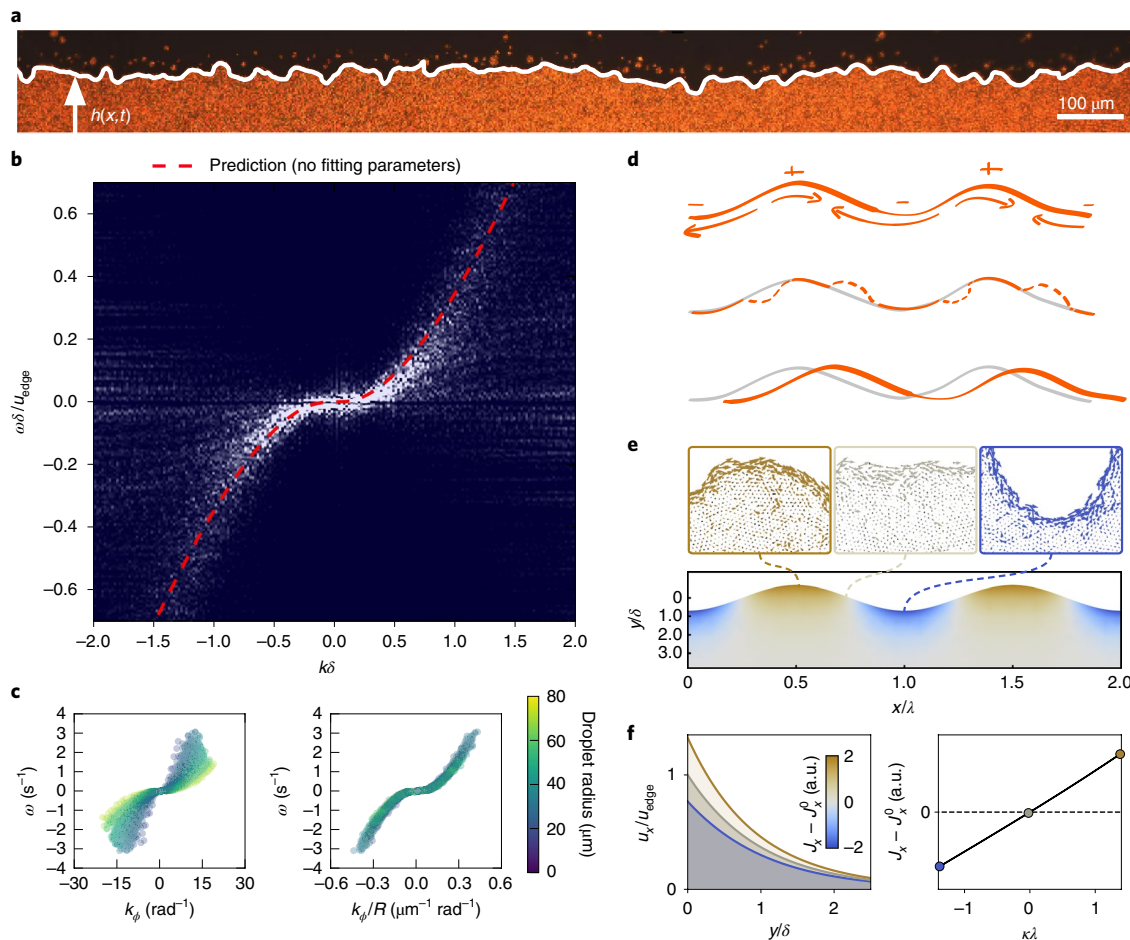


Fig. 2 | Surface waves in a chiral spinner fluid. **a**, Surface waves are excited by perturbing a strip of the spinner fluid. To characterize them, we track the height profile of the strip in time, $h(x, t)$. **b**, The resulting power spectrum from these waves $\langle |h(k, \omega)| \rangle$ is plotted versus the normalized wavevector $k\delta$ and frequency $\omega/(u_{\text{edge}}/\delta)$. The spectrum peaks on a curve corresponding to the dispersion relation of the waves. Shown with the red dashed line is the theoretical prediction for the dispersion relation, obtained with the hydrodynamic parameters that we measure in Fig. 3; its long-wavelength asymptotic form is given in equation (3). **c**, The power spectrum $\omega(k_\phi)$ for surface waves on a perturbed circular droplet of spinner fluid (left panel and Supplementary Section 4) can be collapsed (right panel) by rescaling the angular wavenumber k_ϕ by the droplet radius R . **d**, A sketch of the mechanism for wave propagation. The propagation of waves can be understood by considering the mass flux, plotted in **e**. The chiral fluid is displaced from the high-curvature to the low-curvature regions. This process explicitly breaks the left-right symmetry, thereby propagating surface waves along only one direction. **e**, Correction to the net mass flux along the interface due to a sinusoidal height perturbation, $J_x - J_x^0$, where J_x^0 is the mass flux in a flat strip and J_x is the mass flux in the presence of a perturbation of wavelength λ . This variation in mass flux tracks the variation in local curvature, as illustrated by the experimental images of regions of positive, neutral and negative curvature κ . **f**, The integrated flow, derived from our hydrodynamic theory, yields the mass flux (left), which scales with curvature (right).

perturbation, there is enhanced flow in positively curved regions (top of the wave) and decreased flow in negatively curved regions (bottom of the wave). This ‘edge-pumping’ mechanism is reminiscent of the phenomenon of shifting sand dunes⁴¹, in which an external wind moves material away from curved regions towards the flat wavefront, giving rise to unidirectional wave motion. As the chiral liquid produces an intrinsic surface wind, these free surface waves are distinctly self-shifting.

A linear stability analysis of the hydrodynamic equations (see Supplementary Section 6 for a detailed calculation) confirms this scenario and yields a prediction for the dispersion relation, dissipation rate and flow fields of surface waves, which we plot in Fig. 2b (red dashed curves). With no fitting parameters, our model shows excellent agreement with the experimentally measured dispersion relation. For surface waves $h \sim e^{i(kx + \omega t)}$ of long wavelength $k \ll 1/\delta$, the asymptotic dispersion relation is:

$$\omega(k) = 2\omega_{\text{edge}} \frac{\eta}{\eta + \eta_R} (k\delta)^3 = 2u_{\text{edge}} \frac{\eta}{\Gamma_u} k^3 \quad (3)$$

where $u_{\text{edge}} = 2\Omega\delta\eta_R/(\eta + \eta_R)$.

The wave dynamics are thus crucially sensitive to boundary-layer flows. A natural avenue for investigation, then, is to seek to increase the thickness of the boundary to increase its relative role. We now show how an increase of the penetration depth of the boundary layer amplifies chiral effects and suggests the presence of a long sought-after source of stress, commonly referred to as Hall viscosity.

Chiral wave damping and measurement of Hall viscosity

We reduce the surface friction by allowing our chiral liquid to sediment on an air/water interface (Fig. 4b), as opposed to a glass surface (Fig. 4a). Due to the difficulty in maintaining a slab geometry in this regime, we examine surface fluctuations on circular droplets.

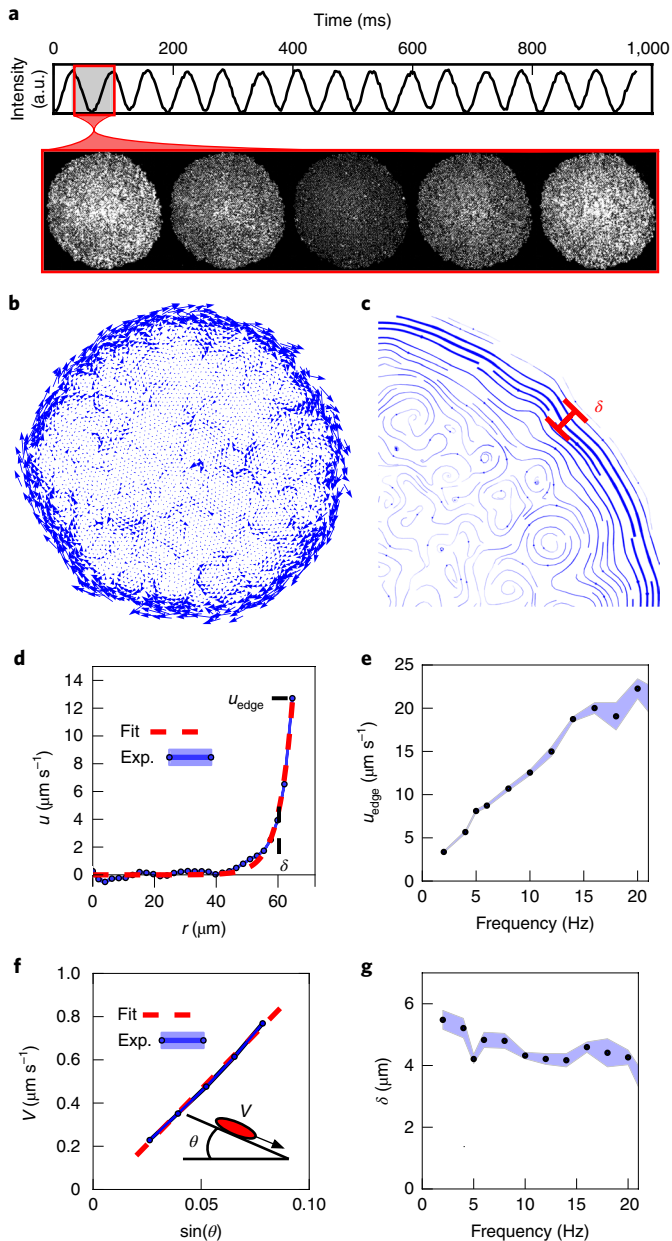


Fig. 3 | Characterization of a droplet of chiral spinner fluid. **a**, When viewed through crossed polarizers, the particles blink as they spin. This allows us to confirm that they all spin at the same frequency, set by the rotating magnetic field. **b**, By measuring the velocity of each particle within a cluster, we find a flow profile that is concentrated at the edge within a penetration layer δ shown in **c**, **d** and **g**. **c**, A zoomed-in view of the flow streamlines, obtained by averaging several instantaneous velocity profiles such as the one shown in **b**. **d**, By measuring the flow profile, the edge current u_{edge} and penetration depth δ are extracted. **e**, **g**, By measuring the flow profile $u(r)$ at a range of frequencies, we extract the shear viscosity, η , and rotational viscosity, η_r , in terms of the substrate friction, Γ_u . **f**, Finally, by tilting a sample by an angle θ and measuring the sedimentation velocity of a droplet, we extract the substrate friction. All error bars represent standard deviations.

As can be seen in Fig. 4a,b and Supplementary Video 10, the edge flow penetrates deeper into the chiral fluid as friction is reduced. The dispersion relations for high- and low-friction droplets display the same trend, although the range of accessible wavevectors normalized by the penetration length ($k\delta$) is larger in the low-friction case.

An extension of our theory to circular geometries (Supplementary Section 6) again accurately captures the dispersion relations for high friction (Fig. 4a) and low friction (Fig. 4b).

The remarkable agreement between experiment and theory is, however, challenged when investigating the damping dynamics of the chiral waves. Experimentally, the damping rate α of chiral waves of wavevector k is given by fitting a Lorentzian to the width of the power spectrum (Supplementary Section 4); the resulting damping rates are shown in Fig. 4c,d. Our hydrodynamic theory predicts this damping rate to be proportional to surface tension. This is natural since surface tension flattens interfacial deformation: in the absence of inertia, the relaxation does not overshoot and capillary waves are overdamped. In the long-wavelength limit ($k\delta \ll 1$), the damping rate $\alpha \sim (\gamma/\Gamma_u)|k|^3$ stems from the competition between surface tension and substrate friction. As seen in Fig. 4c, in the high-friction case, we again find excellent agreement between theory and experiment, which provides a direct measurement of surface tension. The value we find, $\gamma = 2.3 \pm 0.2 \times 10^{-13}$ N, is consistent with an estimate based on magnetic interactions between rotating dipoles (Supplementary Section 6).

In the case of low surface friction, however, we observe a distinct new feature in the dissipation rate: a levelling off of the dissipation rate at short wavelengths that cannot be accounted for by the hydrodynamic theory discussed thus far, suggesting the presence of an additional mechanism for surface wave dissipation in our chiral fluid. Seeking a hydrodynamic description, we recall that isotropic chiral fluids can in principle possess an additional stress in their constitutive relation, known interchangeably as ‘anomalous viscosity’, ‘odd viscosity’ or ‘Hall viscosity’^{12,14,15,42}. This non-dissipative, transverse stress is linked by Onsager relations to the breaking of time-reversal symmetry.

Theoretically, odd viscosity has indeed been shown to arise in the hydrodynamics of plasmas^{43–45}, systems of spinning molecules^{13,17}, and quantum Hall fluids and vortex fluids^{12,42,46}. A signature of Hall viscosity was further revealed in the transport properties of magnetized three-dimensional dilute gases¹⁸. Here we conjecture our dense chiral fluid to support an additional Hall stress $\sigma_{ij}^o = \eta_o (\partial_i \epsilon_{jk} u_k + \epsilon_{ik} \partial_k u_j)$. In incompressible fluids such as the one considered here, the effect of odd viscosity can be seen solely at the edge. This is because in the bulk flow Hall stress is merely absorbed into the fluid pressure. The signature of odd viscosity in our chiral fluid is thus an additional boundary stress. The component normal to the interface σ_{nn} , after absorbing a vortical component into the pressure, is given by

$$\sigma_{nn} = 2\eta_o \left(\partial_s u_n + \frac{u_s}{R(s)} \right) \quad (4)$$

where u_n (respectively, u_s) is the velocity normal (respectively, tangential) to the surface (see Fig. 4e), and $R(s)$ is the local radius of curvature.

In our system, where odd stress powers a boundary-layer edge flow, we thus expect odd viscosity to flatten surface deformation in a manner akin to surface tension, $\sigma^o \sim \eta_o u_s/R$. The excellent agreement between our measurements and predictions from a hydrodynamic theory with this addition confirms this simplified picture and supports the presence of Hall viscosity in our colloidal chiral fluid (see Fig. 4d,f,g). From the fit, we obtain $\eta_o = 1.5 \pm 0.1 \times 10^{-8}$ Pa m s.

The most visible suggestion of the presence of Hall viscosity is the decrease in slope in the damping relation, which can be understood on dimensional grounds. In the long-wavelength limit, the wave relaxation time is controlled by the competition of either surface tension or Hall stress with substrate friction. Dimensionally, this implies a scaling $\alpha \sim |k|^3$, since the ratios γ/Γ_u and $\eta_o u_s/\Gamma_u$ have dimensions of volume per unit time. In contrast, in the short-wavelength limit, surface friction plays no role and damping stems from

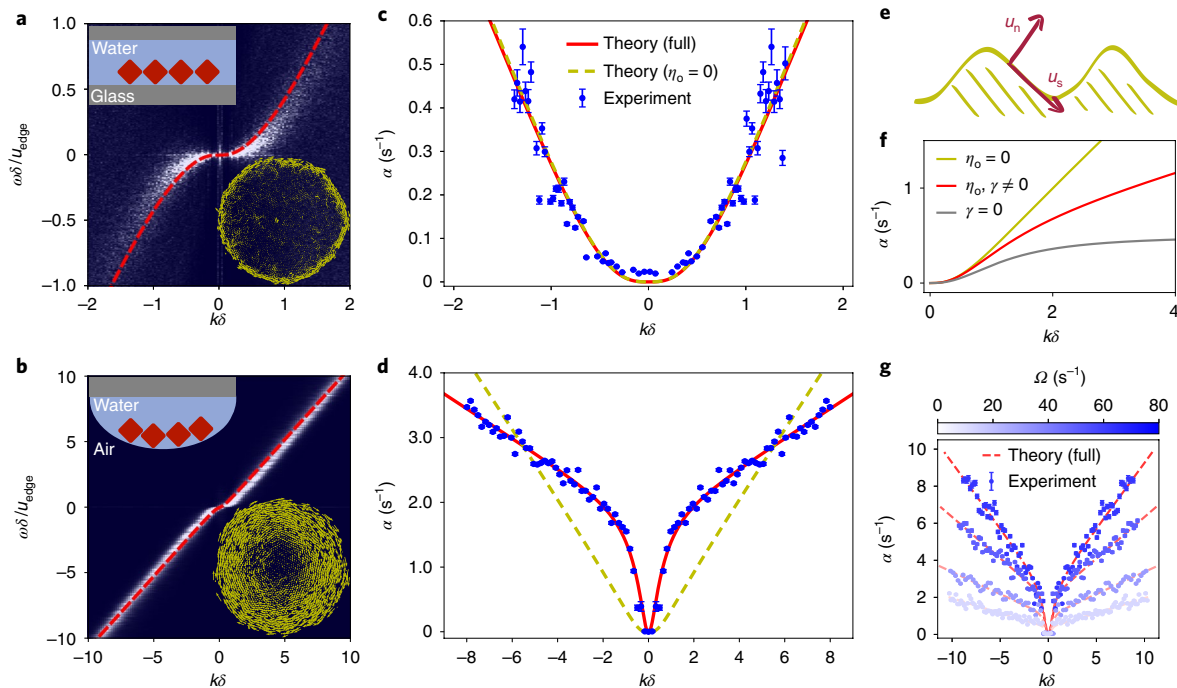


Fig. 4 | Wave dissipation and measurement of Hall viscosity. **a**, In the circular geometry, surface waves yield power spectra $\langle |R(k, \omega)| \rangle$, plotted here versus the normalized wavevector $k\delta$ and frequency $\omega\delta/u_{\text{edge}}$ (see Fig. 2c for a collection of spectra). **b**, Power spectrum at a low-friction air/water interface, for which the edge current is delocalized into the bulk when compared to a high-friction interface as in **a** (see insets). **c**, The dissipation rate of waves on the surface of a circular droplet can be used to extract the surface tension, and the shape of $\alpha(k)$ can be captured by a theory with no odd viscosity ($\eta_o = 0$). **d**, Lowering substrate friction causes the dissipation to level off for large $k\delta$, which can be captured only by a theory including η_o . **e**, The tangential and normal components of velocity at the boundary give rise to a normal Hall stress (equation (4)). **f**, The dissipation for a chiral fluid with η_o in the absence of surface tension, γ , versus the same for a fluid with γ in the absence of η_o . For small $k\delta$, the two curves are indistinguishable. For large $k\delta$, the η_o -dissipated fluid shows no k -dependence, while the γ -dissipated fluid shows linear k -dependence. Shown also for reference is the attenuation for a fluid with finite values of both γ and η_o . **g**, With η_o -induced attenuation, $\alpha(k)$ varies with frequency with all other parameters held constant, a trend that is not observed for γ -dissipated fluids. All error bars represent fit uncertainty from the determination of $\alpha(k)$ from the power spectrum (Supplementary Section 4).

the competition of surface tension or Hall stress and bulk viscosities alone. In this case, dimensional analysis requires linear scaling with wavenumber in the case of surface tension, and wavenumber independence in the case of Hall stress (Supplementary Section 6). This change in wavenumber dependence brings about a visible rollover to a decreased slope in the wave damping rate.

We note that for small ranges of $k\delta \sim [-1, 1]$, characteristic of spectral measurements in the presence of high surface friction, the levelling off cannot be seen and the relative roles of Hall viscosity and surface tension become hard to separate in our model. This is the case for the damping shown in Fig. 4c, which can be fitted well by both a non-zero and zero value of Hall viscosity (Supplementary Section 4).

Having established that Hall viscosity affects the damping of waves in a simple chiral fluid, it is natural to ask whether it has an effect on the dispersion of waves. The first term in equation (4) suggests that Hall viscosity and surface tension could act together to support wave propagation. Surface tension acts on a sinusoidal surface deformation by pulling down peaks and pushing up troughs, generating an in-phase normal velocity component. The normal Hall stress $\partial_x u_n$ would then act out of phase on the inflection points of the sinusoidal perturbation to propagate it in a chiral fashion. Our full theory confirms that this additional wave-driving mechanism indeed exists and generates waves even in the absence of edge currents. However, for our hydrodynamic parameters, their effect on the dispersion is minimal.

An odd instability

In much of the phenomenology we have discussed, surface dynamics are essentially boundary-layer dynamics. Another natural question,

then, is what happens when two boundary layers meet? Draining fluid past a curved obstacle brings about the progressive thinning of a curved strip of chiral fluid, as shown in Fig. 1g and Supplementary Video 6. The flow is smooth until the strip thickness becomes comparable to the penetration depth δ ; at that point, the flow becomes unstable, resulting in the formation of circular droplets. We study this novel pearling mechanism in experiment by creating a sequence of strips of decreasing thickness, as shown in Fig. 5a and Supplementary Video 11. We find that over a period of 10 min the strips of chiral fluid are stable for thicknesses above $\sim 32 \mu\text{m}$ and unstable below.

Although visually reminiscent of the Rayleigh–Plateau instability of a thin fluid cylinder jet⁴⁷, this instability is fundamentally different. In our two-dimensional system, surface tension is a purely stabilizing force, as seen in the wave analysis discussed above. Instead, the instability originates from the chiral surface dynamics of our fluid. A visual signature of this origin is the consistent offset in the phase between top and bottom perturbations at the moment the instability occurs in all strips: Fig. 5b shows one such example.

A linear stability analysis of a thin strip of chiral fluid quantitatively predicts the existence of unstable modes that consist of wave-like perturbations on the top and bottom surfaces that have a relative phase offset, as sketched in Fig. 5d. These are accompanied by a stable mode with an opposite relative phase. The associated stability diagram is shown in Fig. 5e, together with our experimental observations. As the Hall stress has little effect on the stability of modes for small δ (Supplementary Section 6), here we set $\eta_o = 0$.

An intuitive picture for the mechanism driving the instability is illustrated in Fig. 5d. The geometry of a thin slab with out-of-phase

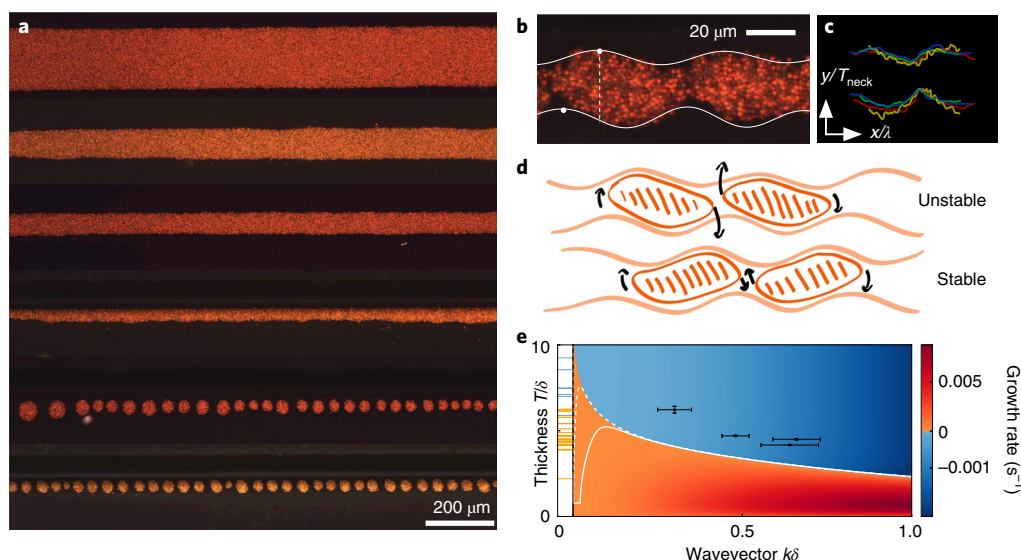


Fig. 5 | A hydrodynamic instability. **a**, Strips of chiral fluid with different thicknesses. Above $32\ \mu\text{m}$, the strips are stable, as observed over the course of ≥ 10 min. Below $32\ \mu\text{m}$, the strips break into droplets within 1 min. **b**, A chiral-fluid strip approaching instability. The continuous white lines represent the sum of the most prominent Fourier modes of the strip outline. The relative phase difference between interfaces is emphasized by the two white dots and vertical dashed line. **c**, An overlay of strip outlines at four breakup points; each colour corresponds to a different instability occurrence. The x axis is rescaled by the most prominent wavelength, λ , and the y axis is rescaled by the thickness at the narrowest point, T_{neck} . The relative phase between the top and bottom interface is consistent with theory. **d**, A schematic of the instability mechanism. Thin strips of chiral fluid are like a collection of elongated droplets rotating in the direction of the edge current. This leads to the breakup (top) or stabilization (bottom) of the strip. **e**, A stability diagram, calculated with linear stability analysis using our experimentally extracted values for the hydrodynamic coefficients, with $\eta_b = 0$ (Supplementary Section 5). The thinner the strip, the larger the range of unstable wavelengths. A surface fluctuation at an unstable wavelength will grow exponentially: orange denotes a positive growth rate and blue denotes a negative growth rate, namely damping. The contour lines mark growth rates corresponding to 10 min (continuous) and 1 day (dashed). The black points represent experimental data from unstable strips; wavelengths were measured by Fourier-transforming the strip outline. The error bars in thickness correspond to the standard deviation in the measurement at various points. The error bars in wavelength correspond to the half-width of the Fourier peaks. The horizontal lines on the y axis mark the recorded strip thicknesses: the orange and blue lines correspond to unstable and stable strips, respectively.

perturbations on the top and bottom surfaces can be approximated by a collection of elongated droplets of chiral fluid all canted in the same direction. Droplets of this kind rotate in the direction of the edge current, in this case clockwise (see Fig. 1d and Supplementary Video 3). Depending on the phase difference between the two interfaces, the rotation of these effective droplets will either increase the amplitude of the perturbation, resulting in the breakup of the strip (top), or decrease the amplitude of the perturbation and restore the flat interface (bottom). The consistent observation of this phase relation between the top and bottom perturbations across many experiments of strips becoming unstable (Fig. 5c) further corroborates our theoretical picture of the instability.

We have broken parity symmetry at the microscopic level in a colloidal chiral fluid, resulting in the emergence of an odd stress that in turn generates lively surface flows. Through these flows, we have observed a signature of Hall viscosity, an experimentally elusive and dissipationless transport coefficient that arises from the breaking of time-reversal symmetry. The combination of these features drives rich interfacial dynamics with no analogues in conventional fluids. These dynamics include the unidirectional propagation and anomalous attenuation of surface waves, and an asymmetric pearling instability. In principle, these chiral phenomena can be tuned (for instance, by altering the colloidal particles' shape and their effective interactions). Colloidal chiral fluids enable the study of universal aspects of a new class of hydrodynamics, and provide a platform for engineering active materials with so far untapped 'odd' behaviours.

Note added in proof: In the final stages of the editorial process, we became aware of a contemporaneous measurement of Hall viscosity in graphene's electron liquid⁴⁸.

Data availability

The data that support the plots within this paper and other findings of this study are available from the corresponding author upon request.

Online content

Any methods, additional references, Nature Research reporting summaries, source data, statements of code and data availability and associated accession codes are available at <https://doi.org/10.1038/s41567-019-0603-8>.

Received: 3 May 2018; Accepted: 25 June 2019;

Published online: 9 September 2019

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Acknowledgements

We would like to acknowledge discussions with P. Wiegmann, A. Abanov, V. Vitelli and A. Souslov. We also thank M. Fruchart for discussions and pointing us to the review of kinetic theory of Hall viscosity presented in the supplementary information of ref. ⁴⁶. We finally thank J. Simon for designing our current control circuits and R. Morton for the rendering in Fig. 1b. This work was primarily supported by the University of Chicago Materials Research Science and Engineering Center, which is funded by the National Science Foundation under award number DMR-1420709. Additional support was provided by NSF EFRI NewLaw grant 1741685 and the Packard Foundation. M.J.S. acknowledges the support from NSF grants DMR-1420073 (NYU-MRSEC) and DMS-1463962. S.S. acknowledges support from NSF award DMR-1653465. D.B. and W.T.M.I. gratefully acknowledge the Chicago–France FACCTS programme. The Chicago MRSEC (US NSF grant DMR 1420709) is also gratefully acknowledged for access to its shared experimental facilities.

Author contributions

W.T.M.I. initiated research. W.T.M.I., D.B. and M.J.S. designed and supervised research. V.S., E.S.B. and S.M. designed and performed experiments and analysed data. D.B. and W.T.M.I. analysed data. M.J.S. and W.T.M.I. developed theory. S.S. and V.S. synthesized particles. S.M. built the magnetic control system. All authors discussed the results and wrote the manuscript.

Competing interests

The authors declare no competing interests.

Additional information

Supplementary information is available for this paper at <https://doi.org/10.1038/s41567-019-0603-8>.

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Peer review information: *Nature Physics* thanks Petia Vlahovska and the other, anonymous, reviewer(s) for their contribution to the peer review of this work.

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