

Search for pair-produced three-jet resonances in proton-proton collisions at $\sqrt{s} = 13$ TeV

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A search has been performed for pair-produced resonances decaying into three jets. The proton-proton collision data used for this analysis were collected with the CMS detector in 2016 at a center-of-mass energy of 13 TeV and correspond to an integrated luminosity of 35.9 fb^{-1} . The mass range from 200 to 2000 GeV is explored in four separate mass regions. The observations show agreement with standard model expectations. The results are interpreted within the framework of R -parity violating SUSY, where pair-produced gluinos decay to a six quark final state. Gluino masses below 1500 GeV are excluded at 95% confidence level. An analysis based on data with multijet events reconstructed at the trigger level extends the reach to masses as low as 200 GeV. Improved analysis techniques have led to enhanced sensitivity, allowing the most stringent limits to date to be set on gluino pair production.

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I. INTRODUCTION

Multijet final states at hadron colliders provide a unique window into many possible extensions of the standard model (SM), albeit in the presence of large SM background processes. Many of these models predict resonances, such as heavy colored fermions transforming as octets under $SU(3)_c$ [1–4] or supersymmetric gluinos that undergo R -parity violating (RPV) decay into three quarks [5–7]. All analyses of data collected at the Fermilab Tevatron by CDF [8] and at run 1 of the CERN LHC by CMS [9,10] at $\sqrt{s} = 7$ and 8 TeV used the jet-ensemble method to suppress the large SM background. Searches for similar signals have been performed by ATLAS [11–13] at $\sqrt{s} = 7, 8$, and 13 TeV. These analyses provide limits that exclude gluinos undergoing RPV decays, for gluino masses below 144, 650, and 917 GeV for the Tevatron, CMS, and ATLAS results, respectively.

Presented here are the results of a dedicated search for pair-produced resonances, each decaying into three quarks (referred to as “three-jet resonances” hereafter) in multijet events in proton-proton (pp) collisions. The study is based on a data sample of pp collisions at $\sqrt{s} = 13$ TeV, corresponding to an integrated luminosity of $35.9 \pm 0.9 \text{ fb}^{-1}$ [14], collected in 2016 with the CMS detector [15]. Events with at least six jets, each with high transverse

momentum (p_T), are selected and investigated for the presence of three-jet resonances consistent with strongly coupled particle decays. The event selection criteria are optimized using a supersymmetric gluino model with the assumption that the gluinos decay with a 100% branching fraction to quarks. Compared to previous analyses, this search extends its reach to lower masses because of improvements in data acquisition. Additionally, improvements in analysis techniques such as use of Dalitz variables and new selection algorithms significantly enhance sensitivity over the entire mass spectrum. We observe an improvement in sensitivity by a factor of 6.2 (1.8) at 200 (2000) GeV compared to the previous best limits.

II. THE CMS DETECTOR

The central feature of the CMS apparatus [15] is a superconducting solenoid of 6 m internal diameter, providing a magnetic field of 3.8 T. Within the solenoid volume are a silicon pixel and strip tracker, a lead tungstate crystal electromagnetic calorimeter (ECAL), and a brass and scintillator hadron calorimeter (HCAL), each composed of a barrel and two end-cap sections. Forward calorimeters extend the pseudorapidity (η) coverage provided by the barrel and end-cap detectors. Muons are detected in gas-ionization chambers embedded in the steel flux-return yoke outside the solenoid. A particle-flow (PF) algorithm [16] aims to reconstruct and identify each individual particle in an event, with an optimized combination of information from the various elements of the CMS detector. The energy of photons is obtained from the ECAL measurement. The energy of electrons is determined from a combination of the electron momentum at the primary interaction vertex as

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determined by the tracker, the energy of the corresponding ECAL cluster, and the energy sum of all bremsstrahlung photons spatially compatible with originating from the electron track. The energy of muons is obtained from the curvature of the corresponding track. The energy of charged hadrons is determined from a combination of their momentum measured in the tracker and the matching ECAL and HCAL energy deposits, corrected for zero-suppression effects and for the response function of the calorimeters to hadronic showers. Finally, the energy of neutral hadrons is obtained from the corresponding corrected ECAL and HCAL energy. The physics objects are the jets, clustered with the tracks assigned to the vertex as inputs, and the associated missing transverse momentum, taken as the negative vector sum of the p_T of those jets. The reconstructed vertex with the largest value of summed physics-object p_T^2 is taken to be the primary pp interaction vertex.

Jets are reconstructed from the energy deposits in the calorimeter towers together with the tracks assigned to the vertex, clustered using the anti- k_T algorithm [17,18] with a distance parameter of 0.4 (referred as AK4 jets). Jet momentum is determined as the vectorial sum of all particle momenta in the jet, and is found from simulation to be within 5% to 10% of the true momentum over the whole p_T spectrum and detector acceptance. Additional proton-proton interactions within the same or nearby bunch crossings can contribute additional tracks and calorimetric energy depositions to the jet momentum. To mitigate this effect, tracks identified to be originating from pileup vertices are discarded, and an offset correction is applied to correct for remaining contributions. Jet energy corrections are derived from simulation to bring the measured response of jets to that of particle level jets on average. In situ measurements of the momentum balance in dijet, photon + jet, Z + jet, and multijet events are used to account for any residual differences in jet energy scale in data and simulation [19]. Additional selection criteria are applied to each jet to remove jets potentially dominated by anomalous contributions from various subdetector components or reconstruction failures [20]. The jet energy resolution amounts typically to 15% at 10 GeV, 8% at 100 GeV, and 4% at 1 TeV.

III. TRIGGERS

Events of interest are selected using a two-tiered trigger system [21]. The first level (L1), composed of custom hardware processors, uses information from the calorimeters and muon detectors to select events at an average rate of around 100 kHz within a time interval of less than 4 μ s. The second level, known as the high-level trigger (HLT), consists of a farm of processors running a version of the full event reconstruction software optimized for fast processing, and reduces the event rate to around 1 kHz before data storage. To keep the recorded data rate low, high thresholds

are imposed for the triggers used to study jet-based physics, such as requiring high- p_T jets and a large H_T (scalar sum of AK4 jet p_T values).

For the high-mass search, covering the signal mass region above 700 GeV, we use events collected by the OR of two different triggers: the first requires $H_T \geq 800$ GeV calculated with jet $p_T \geq 40$ GeV; and the second requires at least four jets with $p_T \geq 70$ GeV and $H_T \geq 750$ GeV. Hereafter, this set of triggers will be referred to as jets + H_T . In order to achieve full trigger efficiency for events passing the offline selection, the following selection is imposed: $H_T \geq 900$ GeV with jet $p_T \geq 50$ GeV and jet multiplicity (N_{jets}) ≥ 6 . All jets are required to be within $|\eta| < 2.4$. The high thresholds of this trigger makes it insensitive to physics at low mass scales (~ 200 GeV).

To probe new physics at low mass scales, the selection criteria for the trigger must be relaxed. The trigger used for the low-mass search is called the PF scouting trigger, which has an H_T requirement of ≥ 410 GeV calculated with jet $p_T \geq 20$ GeV. This results in an event record rate about 2 kHz. Owing to limitation on the available bandwidth, a minimal amount of information is stored per event, specifically: PF objects, comprising jets, leptons, and photons as reconstructed at the HLT. This yields an event size of 10 KB/event which, is significantly smaller than the 1 MB event size for normal triggers. The thresholds of the PF scouting trigger allow us to reconstruct the fully hadronic decay of the top quark, which provides a well understood three-jet resonance signal to validate both the PF scouting trigger and the search strategy. In order to achieve full trigger efficiency for events passing the offline selection, the following selection is imposed: $H_T \geq 650$ GeV with jet $p_T \geq 30$ GeV and $N_{\text{jets}} \geq 6$. All jets are required to be reconstructed within $|\eta| < 2.4$.

IV. GENERATION OF SIMULATED EVENTS

Pair-produced gluinos are used to model the signal. Gluino production is simulated using MADGRAPH 5_aMC@NLO 2.2.2 [22] and gluino decays are simulated using PYTHIA 8.212 [23], with each gluino decaying into three jets via the λ_{udd} quark RPV coupling. The coupling is set such that the branching fraction of the gluino to three jets is 100%. The masses of the generated gluinos range from 200 to 2000 GeV in steps of 100 GeV. For the generation of this signal, all superpartners except the gluino are decoupled [7] by setting the squark masses to high values. The natural width of the gluino resonance is assumed to be much smaller than the resolution of the detector, and no intermediate particles are produced in the gluino decay. Simulation of the CMS detector is performed using GEANT4 [24].

All simulated samples are produced with the parton distribution functions (PDF) NNPDF3.0 [25], with the precision (LO or NLO) set by the generator used.

V. EVENT SELECTION

Events, recorded with the PF scouting and jets + H_T triggers described above, are required to have at least one reconstructed primary vertex [26]. Since this analysis targets pair-produced three-jet resonances, we require events to contain at least six reconstructed jets.

To identify the three jets (triplet) produced by gluino decay in these multijet events, we extend the jet ensemble technique [8,27] by examining the internal dynamics of multijet events. This technique examines all possible triplets in each multijet event and applies selection criteria to the events, pairs of triplets, and individual triplets to maximize signal sensitivity. We find that restricting the set of considered triplets to the ones involving only the six jets of highest p_T in events with more than six jets, maximizes our sensitivity to the signal, while keeping the background manageable. From the combinatorics of 3 jets chosen from an ensemble of 6, we reconstruct 20 triplets per event, corresponding to two pairs of 10 triplets. For signal events, at most two triplets come from the pair-produced gluino decay, with the remaining triplets corresponding to incorrect jet combinations.

After the offline selection requirements mentioned above, we impose further selection criteria in two steps. In the first, we apply a selection based on event-level variables exploiting the kinematic features and decay topology of the event as a whole. In the second step, we impose selection requirements on variables defined by the features of the triplets and triplet pairs.

A. Dalitz variables

A very useful technique for studying three-body decays uses Dalitz plots, developed by R.H. Dalitz to study K meson decays [28]. Dalitz plots are used to study internal resonances in three body decays. We extend this formalism to construct Dalitz variables that contain information about the internal dynamics of the three-body decay, in order to differentiate between the gluino decays and QCD multijet backgrounds. To construct the Dalitz variables, we form the invariant masses of three dijet pairs inside the triplet, with masses m_{12} , m_{23} , m_{13} . Dalitz variables for a triplet are formed by normalizing these dijet invariant masses. They are defined as follows

$$\hat{m}(3,2)_{ij}^2 = \frac{m_{ij}^2}{m_{ijk}^2 + m_i^2 + m_j^2 + m_k^2}. \quad (1)$$

Here, m_i are the mass of the individual jets and m_{ijk} is the mass of the triplet. Indices here refer to jets in the triplet, where $i, j, k \in \{1, 2, 3\}$. There are three $\hat{m}(3,2)_{ij}^2$ in a triplet; we express this with the label (3,2), where the “3” refers to the overall object being a triplet and the “2” refers to pairs inside this triplet. The invariant mass of the dijet pairs is normalized such that their Dalitz variables sum up

to unity and are dimensionless. For signal triplets, the lack of an internal resonance and the evenly spread out jets make the Dalitz variables close to the value $1/3$, implying a symmetric decay, where the jets have uniform geometric separation in the center-of-mass frame of the gluino. Triplets made of jets arising from QCD multijets are more asymmetric, resulting in their $\hat{m}(3,2)_{ij}^2$ being closer to 0 or 1. This is illustrated in Fig. 1. The three $\hat{m}(3,2)_{ij}^2$ values per triplet are sorted from largest to smallest, and labeled $\hat{m}(3,2)_{\text{high}}^2$, $\hat{m}(3,2)_{\text{mid}}^2$, and $\hat{m}(3,2)_{\text{low}}^2$. We plot the three pairs of these $\hat{m}(3,2)^2$ s per event: $\hat{m}(3,2)_{\text{high}}^2$ vs $\hat{m}(3,2)_{\text{mid}}^2$, $\hat{m}(3,2)_{\text{high}}^2$ vs $\hat{m}(3,2)_{\text{low}}^2$, and $\hat{m}(3,2)_{\text{mid}}^2$ vs $\hat{m}(3,2)_{\text{low}}^2$. These three pairs occupy mutually exclusive

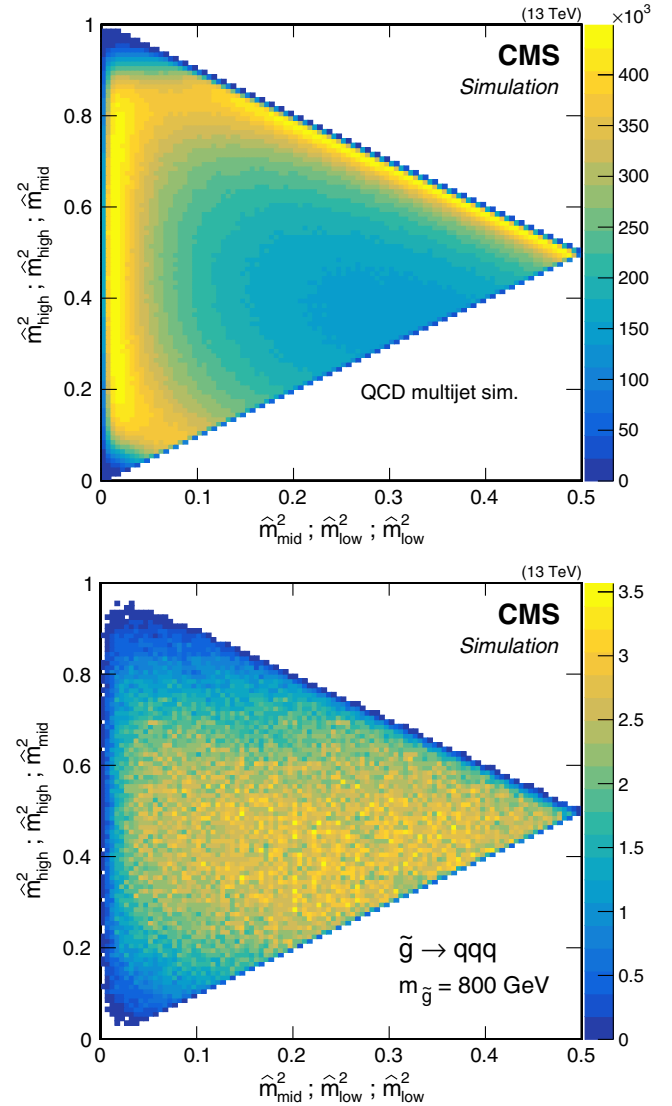


FIG. 1. Pair masses within the triplet as described in Eq. (1) plotting superimposed $\hat{m}(3,2)_{\text{high}}^2$ vs $\hat{m}(3,2)_{\text{low}}^2$, $\hat{m}(3,2)_{\text{high}}^2$ vs $\hat{m}(3,2)_{\text{mid}}^2$ and $\hat{m}(3,2)_{\text{mid}}^2$ vs $\hat{m}(3,2)_{\text{low}}^2$. QCD multijet triplets (left) cluster at the edge, while triplets from signal events ($m_{\tilde{g}} = 800$ GeV, right) fill the center.

regions in the $\hat{m}(3,2)^2$ vs $\hat{m}(3,2)^2$ plane, which combine to give a single overall distribution. This plot is referred to as a dimensionless Dalitz plot. When the variables are displayed in a Dalitz plot, the signal peaks in the center closer to the value $1/3$ while the QCD multijet background clusters around the edges.

Using this feature, we define a variable called mass distance squared (or D^2) to characterize the symmetry between the jets inside a triplet. This variable, which is plotted in Fig. 2, is defined as

$$D_{[3,2]}^2 = \sum_{i>j} \left(\hat{m}(3,2)_{ij} - \frac{1}{\sqrt{3}} \right)^2. \quad (2)$$

We extend this idea to the event-level to define a second variable, to estimate the angular spread of the 6 jets within a

pairs of triplets. This distance measure will have a low value for signal-like topologies, indicating well separated jets with similar momentum and a high value for dijetlike topologies such as QCD. For this purpose, new Dalitz variables are defined as normalized invariant mass of jet triplets constructed from the six highest p_T jets

$$\hat{m}(6,3)_{ijk}^2 = \frac{m_{ijk}^2}{4m_{ijklmn}^2 + 6\sum_i m_i^2}. \quad (3)$$

Here, m_{ijklmn} is the invariant mass of the top six jets, ordered in p_T . Indices here refer to the top six jets ordered in p_T , where $i, j, k, l, m, n \in \{1, 2, \dots, 6\}$. In the label (6,3), the first index refers to the overall object being a six-jet ensemble, and the second refers to triplets inside this six-jet ensemble. For a six-jet topology, we will have 20 such $\hat{m}(6,3)_{ijk}^2$ variables. Six-jet events from QCD multijets are largely due to a core dijet event with extra radiated jets. These jets tend to be grouped together. Jets from pair-produced gluino decays tend to be distributed more uniformly across the detector. This makes these variables close to 0 or 1 for QCD multijets and close to $1/20$ for jets coming from signal decay. The invariant mass of the triplet is normalized such that these 20 event-level Dalitz variables sum up to 1. Using the previously defined variables, we define the following six-jet distance measure in a similar way to $D_{[3,2]}^2$,

$$D_{[(6,3)+(3,2)]}^2 = \sum_{i<j<k} \left(\sqrt{\hat{m}(6,3)_{ijk}^2 + D_{[3,2],ijk}^2} - \frac{1}{\sqrt{20}} \right)^2. \quad (4)$$

This $D_{[(6,3)+(3,2)]}^2$ combines the $D_{[6,3]}^2$ and $D_{[3,2]}^2$ into a single event-level variable. Figure 2 shows the $D_{[(6,3)+(3,2)]}^2$ and $D_{[3,2]}^2$ distributions for QCD multijet background and gluino simulation after the selection criteria: $H_T \geq 650$ GeV, sixth jet $p_T \geq 50$ GeV and $N_{\text{jets}} \geq 6$. The small disagreement between QCD multijet simulation and data visible in Fig. 2 is due to imperfect modeling of the QCD multijet simulation. Since the QCD multijet simulation is not used for predicting the background, this discrepancy has a negligible effect on this search.

B. Other pair and triplet level selections

For each triplet pair, we calculate a variable called “mass asymmetry”, defined as

$$A_m = \frac{|m_{ijk} - m_{lmn}|}{m_{ijk} + m_{lmn}}. \quad (5)$$

Here, m_{ijk} and m_{lmn} are masses of the two unique triplets in a triplet pair. This variable shows discriminating power between signal and background.

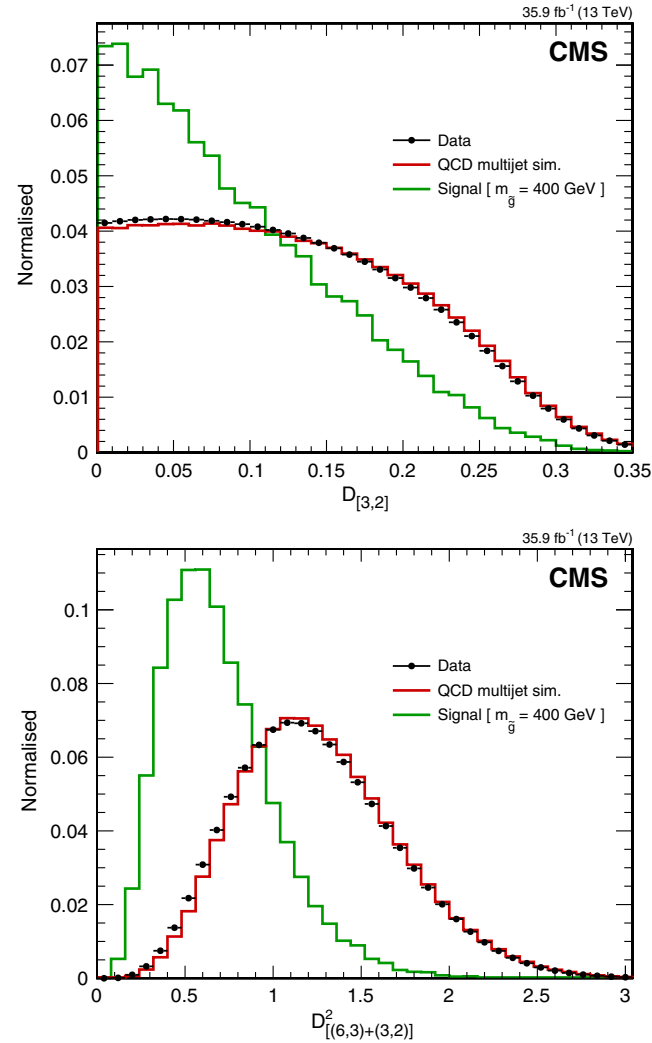


FIG. 2. right: The $D_{[3,2]}^2$ variable as described in Eq. (2) for signal (gluino of mass 400 GeV) and QCD multijet triplets. left: The $D_{[(6,3)+(3,2)]}^2$ distribution as described in Eq. (4), for signal (gluino of mass 400 GeV) and QCD multijet triplets. The distributions are made after nominal selection criteria.

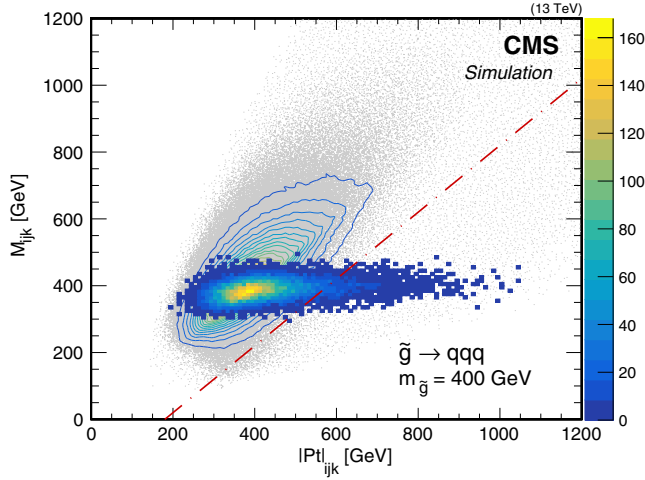


FIG. 3. Expected distribution of triplet invariant mass versus the scalar sum of jet p_T in the triplet for a gluino of mass 400 GeV decaying into jets. The filled color represents correctly reconstructed signal triplets, while the gray points and contour lines represent the QCD multijet background and combinatorial background from signal events. The red dashed line illustrates the Δ cut: triplets to the right of the line are accepted by the selection criterion.

For triplets from multijet QCD events or combinatorial background, the scalar sum p_T ($|p_T|_{ijk}$) will scale with the triplet mass (m_{ijk}). Whereas it is not case for signal triplets as they have constant invariant mass. We exploit this feature of signal triplets by constructing a selection, referred to as a “Delta cut”, defined as

$$m_{ijk} < |p_T|_{ijk} - \Delta, \quad (6)$$

where m_{ijk} is the triplet invariant mass, $|p_T|_{ijk}$ is the scalar sum of jet p_T in the triplet (triplet scalar p_T), and Δ is an adjustable offset. A scatter plot of the triplet invariant mass versus triplet scalar p_T for a gluino with a mass of 400 GeV is shown in Fig. 3, which clearly shows that by imposing this selection criterion we eliminate most of the background while retaining a significant fraction of the signal.

We optimize selection criteria in four separate mass ranges with a metric defined as the ratio of signal to the square root of the background obtained by integrating the triplet mass distribution from gluino and QCD multijet simulations in a window around the signal peak. We note that background can include triplets from QCD multijet, as well as combinatorial background from signal. The four resulting signal regions are defined in Table I and labeled from 1 to 4.

VI. BACKGROUND ESTIMATION

There are three sources of background that we consider: QCD multijets, fully hadronic decays of $t\bar{t}$ pairs, and combinatorial background from signal events. We find

that background due to the $t\bar{t}$ decays is only significant in the lowest mass region of the search. This background is estimated from events simulated with POWHEG [29–32] and their decay is simulated with PYTHIA. The $t\bar{t}$ production rate extracted from a background-only fit in region 1 agrees with the SM expectation within the statistical uncertainty of the measurement. The mass distributions of the QCD multijet and combinatorial backgrounds are estimated by fitting a smooth function to data. Studies of simulated QCD multijet events and combinatorial background from signal events show that the combined mass distribution can be described by a single smooth function. Except for the lowest mass region, the triplet invariant mass background distribution is smoothly falling (as we can see in Fig. 4), and we use three types of functions, fit directly to the data, to model this background in different regions.

The background distribution of triplets in region 1 shows features due to the turn-on of QCD multijet background and $t\bar{t}$ decays. For modeling this QCD multijet and combinatorial background, we use a function inspired by the formulation of Planck’s law of blackbody radiation with an added logarithmic correction to the tails, and this distribution models the background well. This function models the QCD multijet background turn-on better than the four-parameter function, used to fit triplet mass distributions in other regions:

$$\frac{dN}{dx} = \frac{1}{(x+c)^{5+d\ln\frac{x}{\sqrt{s}}}} \frac{a}{e^{\frac{b}{x+c}} - 1}, \quad (7)$$

where, a is the factor controlling the normalization of the fit, b is the “temperature” term in blackbody distribution, c controls the translation of the whole distribution, d controls the strength of the logarithmic term, and $\sqrt{s}$ is the center-of-mass energy of the proton-proton collisions.

For modeling the background in regions 2 and 3, we use the following four-parameter function

$$\frac{dN}{dx} = p_0 \frac{(1 - \frac{x}{\sqrt{s}})^{p_1}}{(\frac{x}{\sqrt{s}})^{p_2+p_3\ln\frac{x}{\sqrt{s}}}}, \quad (8)$$

and for region 4, we used the same parametrization, with p_3 set to zero, to model the background.

The functional form in Eq. (8) successfully models the steeply falling dijet mass distribution of QCD multijet production and has been used extensively in dijet resonance searches [34,35].

We test for possible bias introduced by the choice of background parameterization. We perform signal injection tests on pseudo-experiments generated from QCD multijet simulation. These pseudo-experiments are fit to alternative background parameterizations and the effect on the strength of the extracted signal is examined. These tests indicate a negligible bias. To further validate our background

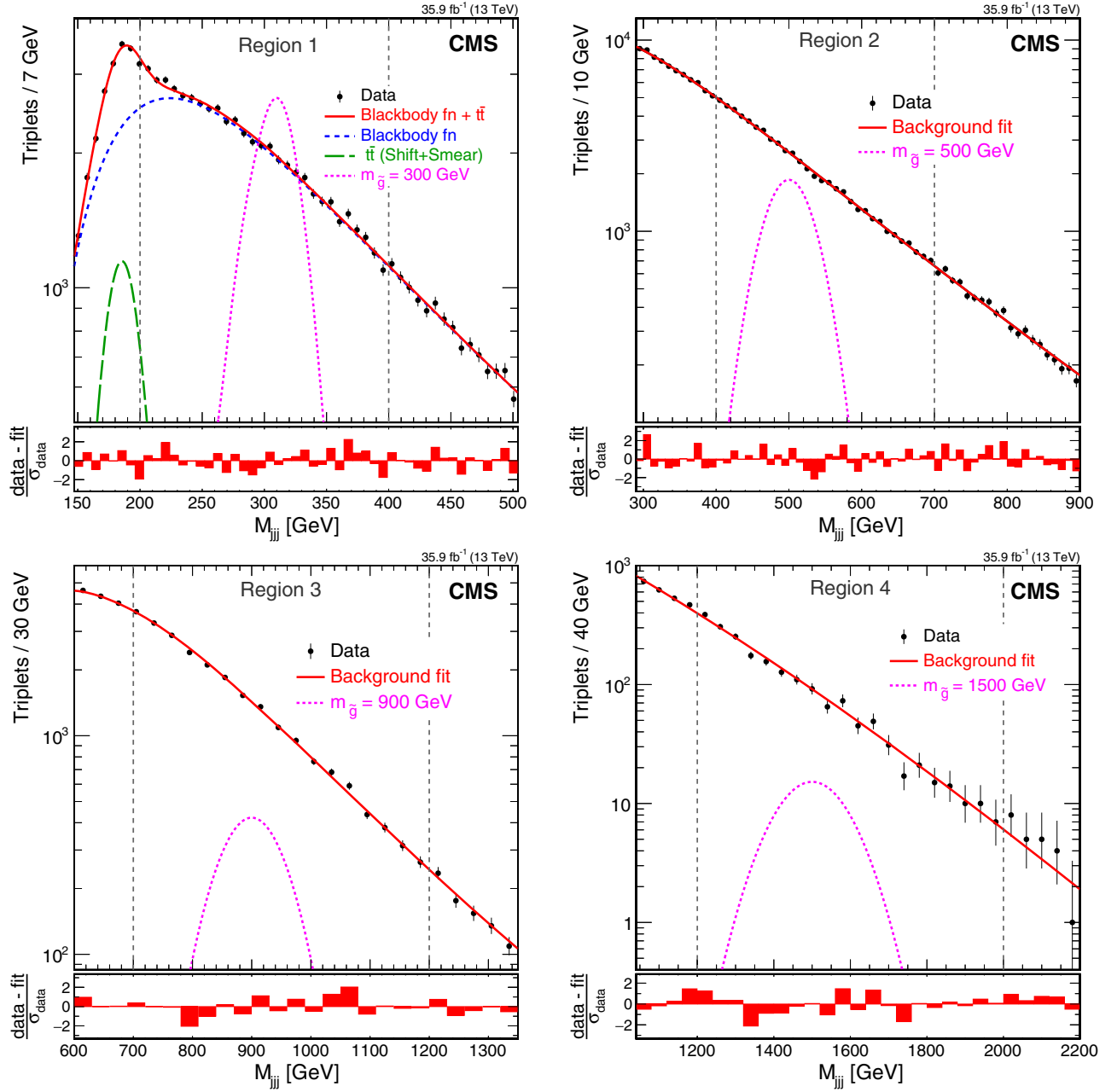


FIG. 4. Mass distributions and background-only fits for each of the mass regions. Region 1 (top left) is fit to a function that combines the blackbodylike term described in Eq. (7) with a simulated $\tilde{t}\bar{\tilde{t}}$ distribution, while region 2 and 3 (top right and bottom left) are fit to the four parameter function from Eq. (8), and region 4 (bottom right) is fit to three parameter function from Eq. (8) with p_3 set to zero. The vertical gray lines indicate the mass regions. The gluino signal normalized to the cross section expected from [33] is shown in magenta.

TABLE I. Gluino mass ranges used in this analysis, and selection criteria used. Note that the gluino mass ranges in the upper two rows in the table use events collected using the PF scouting trigger, while the lower two rows in the table use events collected using jets + H_T trigger. The symbols $>$ and $<$ represent the direction of the cut.

Region	Gluino mass range	Jet p_T	H_T	Sixth jet p_T	$D_{[(6,3)+(3,2)]}^2$	A_m	Δ	$D_{[3,2]}^2$
1	200–400 GeV	> 30 GeV	> 650 GeV	> 40 GeV	< 1.25	< 0.25	> 250 GeV	< 0.05
2	400–700 GeV	> 30 GeV	> 650 GeV	> 50 GeV	< 1.00	< 0.175	> 180 GeV	< 0.175
3	700–1200 GeV	> 50 GeV	> 900 GeV	> 125 GeV	< 0.9	< 0.15	> 20 GeV	< 0.2
4	1200–2000 GeV	> 50 GeV	> 900 GeV	> 175 GeV	< 0.75	< 0.15	> -120 GeV	< 0.25

TABLE II. Summary of the systematic uncertainties in the signal yield. For the uncertainty affecting the distribution (shape), the value represents the percentage difference in the nominal value of the systematic uncertainty. These systematic uncertainties are applied to the signal.

Data set	Source of systematic	Effect	Value
All	Luminosity	Yield	2.5%
	Acceptance	Yield	5%
PF scouting	Shift	Shape	3.5%
	Smear	Shape	4%
jets + H_T	Jet energy correction	Shape	2.5%
	Jet energy resolution	Shape	12%

parametrization, we performed pseudoexperiments generated using data from the PF scouting and jets + H_T event samples. We also perform statistical studies (F-tests) to determine the optimum number of parameters for the background function, to avoid over-constraint. The distributions of triplet mass in the four search regions are shown in Fig. 4, along with the results of fits to the background-only hypothesis. The mass distributions expected for a typical gluino decay is shown in magenta, with the rate normalized to that expected from [33]. The fits reproduce the data distributions well, indicating absence of a signal.

For the signal triplet-mass distribution, signal simulations parameterized with double Gaussian distributions are used. These parameterizations accurately describe the shape of signal triplet mass distribution. The acceptances for the search is defined as the number of correct triplets passing the selection, divided by the number of events generated. The selection criteria given in Table I result in signal acceptance of 2.6×10^{-4} , 8.4×10^{-2} and 1.7×10^{-1} for the resonance masses $m_{\tilde{g}} = 200, 900$, and 1600 GeV, respectively.

VII. SYSTEMATIC UNCERTAINTIES

The search in regions 1 and 2 uses PF scouting data. Jets in these events did not have the full offline set of corrections applied. We use the well-measured all-hadronic decay of the top quark to determine the corrections and corresponding systematic uncertainties for the PF scouting data. The triplet-mass distribution from $t\bar{t}$ simulation must be adjusted in order to agree with the data in region 1, with two transformations to the simulated triplet mass distribution required. The first is a translation of 6.6 GeV, referred to as the “shift” correction. The second is a convolution with a Gaussian distribution of width of 8.9 GeV, referred to as the “smear” correction. The shift and smear values determined from the top resonance measurement are also applied to the gluino simulation. We performed a separate study to investigate the dependence of the shift and smear on the triplet scalar p_T and found negligible correlation. The uncertainties associated with the shift and smear corrections are determined by observing the change in goodness-of-fit metric between simulation and data as these

parameters are varied. Corresponding systematic uncertainties for the shift and smear corrections are estimated to be 3.5% and 4%, respectively. These corrections are defined as a percentage of the mean of the signal distribution. For the jets + H_T data, adjustments analogous to shift and smear are applied to correct for the effects arising from uncertainties in the measurement of jet energy corrections (2.5%) and jet energy resolution (12%). These systematic uncertainties affect the shape of the signal triplet mass distributions.

The other systematic uncertainties affecting the yield from the signal samples are the integrated luminosity measurement (2.5%) and the uncertainty in the determination of acceptance (5%), which includes contributions from uncertainties in the PDF. We list the systematic uncertainties for both data sets in Table II.

VIII. LIMITS

The mass distribution of data is described well by the background parameterization, as illustrated in Fig. 4. We see no significant excess that could indicate the presence of signal, and place upper limits on the product of the cross section and branching fraction for the pair production of three-jet resonances. A modified frequentist approach, with the CL_s criterion as the figure of merit and a profile likelihood as the test statistic, is employed. Limits are calculated with the frequentist asymptotic approximation in RooStats [36–39]. The full CL_s calculator gives similar results. The data are fit using a binned maximum-likelihood function, based on the respective four-parameter function. In region 1, the rate for $t\bar{t}$ events is set to the value observed from the background-only fit and is allowed to float within the systematic uncertainty. The overall QCD scale is unconstrained and the nuisance parameters effecting the overall rate are introduced as log-normal constraints.

The observed and expected 95% confidence level (C.L.) upper limits on the product of gluino pair-production cross section and branching fraction, as a function of gluino mass, are presented in Fig. 5. The solid red line in the figure show the next-to-leading order (NLO) plus next-to-leading-logarithm (NLL) cross sections for gluino pair production [33], and the shaded region around the solid red line

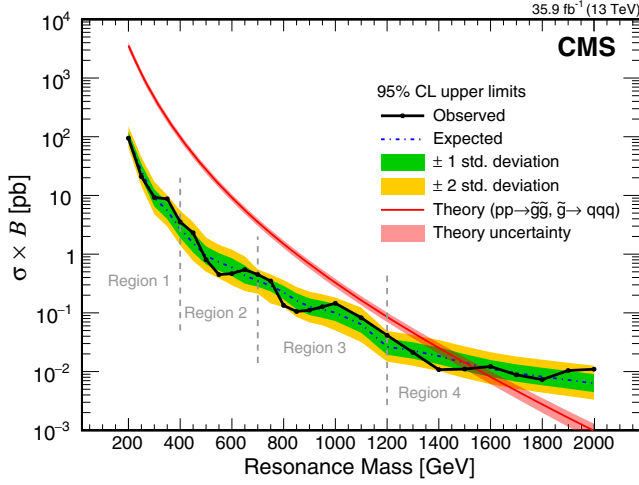


FIG. 5. Observed and expected frequentist CL_s limits on cross section times branching fraction are calculated in the asymptotic approximation. The solid red curve shows the prediction for the gluino pair productions from [33]. The band around the theory curve indicates the uncertainty associated with PDF and scale choices. The gray vertical lines indicate the boundaries between the mass regions.

represent the corresponding 1 standard deviation uncertainties, which range from 14% to 31%. We use the points where the 1 sigma uncertainty curve for the NLO + NLL cross section crosses the observed limit curve to obtain our final results.

The production of gluinos decaying by an R -parity violating interaction into jets is excluded at 95% C.L. for gluino masses below 1500 GeV. This is the most stringent mass limit to date on this model of RPV gluino decay, assuming a 100% branching fraction for gluinos decaying to quark jets.

IX. SUMMARY

A search has been performed for pair-produced resonances decaying into three jets. The proton-proton collision data used for this analysis were collected with the CMS detector in 2016 at a center-of-mass energy of $\sqrt{s} = 13$ TeV and correspond to an integrated luminosity of 35.9 fb^{-1} . The mass range from 200 to 2000 GeV is explored in four separate mass regions. The observations show agreement with standard model expectations. The results are interpreted within the framework of R -parity violating SUSY, where pair-produced gluinos decay to a six quark final state. Gluino masses below 1500 GeV are excluded at 95% confidence level. An analysis based on data with multijet events reconstructed at the trigger level extends the reach to masses as low as 200 GeV. Improved analysis techniques have led to enhanced sensitivity, allowing the most stringent limits to date to be set on gluino pair production.

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- [1] E. Farhi and L. Susskind, Grand unified theory with heavy color, *Phys. Rev. D* **20**, 3404 (1979).
 - [2] W. J. Marciano, Exotic new quarks and dynamical symmetry breaking, *Phys. Rev. D* **21**, 2425 (1980).
 - [3] P. H. Frampton and S. L. Glashow, Unifiable Chiral Color with Natural Glashow-Iliopoulos-Maiani Mechanism, *Phys. Rev. Lett.* **58**, 2168 (1987).
 - [4] P. H. Frampton and S. L. Glashow, Chiral color: An alternative to the Standard Model, *Phys. Lett. B* **190**, 157 (1987).
 - [5] R. S. Chivukula, M. Golden, and E. H. Simmons, Six jet signals of highly colored fermions, *Phys. Lett. B* **257**, 403 (1991).
 - [6] R. S. Chivukula, M. Golden, and E. H. Simmons, Multi-jet physics at hadron colliders, *Nucl. Phys. B* **363**, 83 (1991).
 - [7] R. Essig, Physics beyond the Standard Model: Supersymmetry, dark matter, and LHC phenomenology, Ph.D. thesis, Rutgers University, USA, 2008.
 - [8] T. Aaltonen *et al.* (CDF Collaboration), First Search for Multijet Resonances in $\sqrt{s} = 1.96$ TeV $p\bar{p}$ Collisions, *Phys. Rev. Lett.* **107**, 042001 (2011).
 - [9] CMS Collaboration, Search for three-jet resonances in pp collisions at $\sqrt{s} = 7$ TeV, *Phys. Lett. B* **718**, 329 (2012).
 - [10] CMS Collaboration, Searches for light- and heavy-flavour three-jet resonances in pp collisions at $\sqrt{s} = 8$ TeV, *Phys. Lett. B* **730**, 193 (2014).
 - [11] ATLAS Collaboration, Search for pair production of massive particles decaying into three quarks with the ATLAS detector in $\sqrt{s} = 7$ TeV pp collisions at the LHC, *J. High Energy Phys.* **12** (2012) 086.
 - [12] ATLAS Collaboration, Search for massive supersymmetric particles decaying to many jets using the ATLAS detector in pp collisions at $\sqrt{s} = 8$ TeV, *Phys. Rev. D* **91**, 112016 (2015); Erratum **93**, 039901 (2016).
 - [13] ATLAS Collaboration, Search for R-parity-violating supersymmetric particles in multi-jet final states produced in $p - p$ collisions at $\sqrt{s} = 13$ TeV using the ATLAS detector at the LHC, *Phys. Lett. B* **785**, 136 (2018).
 - [14] CMS Collaboration, CMS Luminosity Based on Pixel Cluster Counting—Summer 2012 Update, CMS Physics Analysis Summary Report No. CMS-PAS-LUM-12-001, 2012, <http://cdsweb.cern.ch/record/1482193>.
 - [15] CMS Collaboration, The CMS experiment at the CERN LHC, *J. Instrum.* **3**, S08004 (2008).
 - [16] CMS Collaboration, Particle-flow reconstruction and global event description with the CMS detector, *J. Instrum.* **12**, P10003 (2017).
 - [17] M. Cacciari, G. P. Salam, and G. Soyez, The anti- k_T jet clustering algorithm, *J. High Energy Phys.* **04** (2008) 063.
 - [18] M. Cacciari, G. P. Salam, and G. Soyez, Fastjet user manual, *Eur. Phys. J. C* **72**, 1896 (2012).
 - [19] CMS Collaboration, Jet energy scale and resolution in the CMS experiment in pp collisions at 8 TeV, *J. Instrum.* **12**, P02014 (2017).
 - [20] CMS Collaboration, Jet algorithms performance in 13 TeV data, CMS Physics Analysis Summary Report No. CMS-PAS-JME-16-003, 2017, <https://cds.cern.ch/record/2256875>.
 - [21] CMS Collaboration, The CMS trigger system, *J. Instrum.* **12**, P01020 (2017).
 - [22] J. Alwall, P. Demin, S. de Visscher, R. Frederix, M. Herquet, F. Maltoni, T. Plehn, D. L. Rainwater, and T. Stelzer, MadGraph/MadEvent v4: The new web generation, *J. High Energy Phys.* **09** (2007) 028.
 - [23] T. Sjöstrand, S. Ask, J. R. Christiansen, R. Corke, N. Desai, P. Ilten, S. Mrenna, S. Prestel, C. O. Rasmussen, and P. Z. Skands, An introduction to PYTHIA 8.2, *Comput. Phys. Commun.* **191**, 159 (2015).
 - [24] S. Agostinelli *et al.* (GEANT 4 Collaboration), Geant4—A simulation toolkit, *Nucl. Instrum. Methods Phys. Res., Sect. A* **506**, 250 (2003).
 - [25] R. D. Ball *et al.* (NNPDF Collaboration), Parton distributions for the LHC Run II, *J. High Energy Phys.* **04** (2015) 040.
 - [26] CMS Collaboration, Tracking and Primary Vertex Results in First 7 TeV Collisions, CMS Physics Analysis Summary Report No. CMS-PAS-TRK-10-005, 2010, <http://cdsweb.cern.ch/record/1279383>.
 - [27] CMS Collaboration, Search for Three-Jet Resonances in pp Collisions at $\sqrt{s} = 7$ TeV, *Phys. Rev. Lett.* **107**, 101801 (2011).
 - [28] R. H. Dalitz, Decay of τ mesons of known charge, *Phys. Rev.* **94**, 1046 (1954).
 - [29] S. Alioli, S.-O. Moch, and P. Uwer, Hadronic top-quark pair-production with one jet and parton showering, *J. High Energy Phys.* **01** (2012) 137.
 - [30] P. Nason, A new method for combining NLO QCD with shower Monte Carlo algorithms, *J. High Energy Phys.* **11** (2004) 040.

- [31] S. Frixione, P. Nason, and C. Oleari, Matching NLO QCD computations with parton shower simulations: The POWHEG method, *J. High Energy Phys.* **11** (2007) 070.
- [32] S. Alioli, P. Nason, C. Oleari, and E. Re, A general framework for implementing NLO calculations in shower Monte Carlo programs: The POWHEG BOX, *J. High Energy Phys.* **06** (2010) 043.
- [33] C. Borschensky, M. Krmer, A. Kulesza, M. Mangano, S. Padhi, T. Plehn, and X. Portell, Squark and gluino production cross sections in pp collisions at $\sqrt{s} = 13, 14, 33$ and 100 TeV, *Eur. Phys. J. C* **74**, 3174 (2014).
- [34] CMS Collaboration, Search for pair-produced resonances decaying to jet pairs in proton-proton collisions at $\sqrt{s} = 8$ TeV, *Phys. Lett. B* **747**, 98 (2015).
- [35] CMS Collaboration, Search for Narrow Resonances Decaying to Dijets in Proton-Proton Collisions at $\sqrt{s} = 13$ TeV, *Phys. Rev. Lett.* **116**, 071801 (2016).
- [36] T. Junk, Confidence level computation for combining searches with small statistics, *Nucl. Instrum. Methods Phys. Res., Sect. A* **434**, 435 (1999).
- [37] A. L. Read, Presentation of search results: The CL_s technique, *J. Phys. G* **28**, 2693 (2002).
- [38] G. Cowan, K. Cranmer, E. Gross, and O. Vitells, Asymptotic formulae for likelihood-based tests of new physics, *Eur. Phys. J. C* **71**, 1554 (2011); Erratum, *Eur. Phys. J. C* **73**, 2501(E) (2013).
- [39] L. Moneta, K. Belasco, K. S. Cranmer, A. Lazzaro, D. Piparo, G. Schott, W. Verkerke, and M. Wolf, The RooStats project, *Proc. Sci., ACAT2010* (**2010**) 057 [arXiv:1009.1003].

A. M. Sirunyan,¹ A. Tumasyan,¹ W. Adam,² F. Ambrogio,² E. Asilar,² T. Bergauer,² J. Brandstetter,² M. Dragicevic,² J. Erö,² A. Escalante Del Valle,² M. Flechl,² R. Frühwirth,^{2,b} V. M. Ghete,² J. Hrubec,² M. Jeitler,^{2,b} N. Krammer,² I. Krätschmer,² D. Liko,² T. Madlener,² I. Mikulec,² N. Rad,² H. Rohringer,² J. Schieck,^{2,b} R. Schöfbeck,² M. Spanring,² D. Spitzbart,² A. Taurok,² W. Waltenberger,² J. Wittmann,² C.-E. Wulz,^{2,b} M. Zarucki,² V. Chekhovsky,³ V. Mossolov,³ J. Suarez Gonzalez,³ E. A. De Wolf,⁴ D. Di Croce,⁴ X. Janssen,⁴ J. Lauwers,⁴ M. Pieters,⁴ H. Van Haevermaet,⁴ P. Van Mechelen,⁴ N. Van Remortel,⁴ S. Abu Zeid,⁵ F. Blekman,⁵ J. D'Hondt,⁵ J. De Clercq,⁵ K. Deroover,⁵ G. Flouris,⁵ D. Lontkovskyi,⁵ S. Lowette,⁵ I. Marchesini,⁵ S. Moortgat,⁵ L. Moreels,⁵ Q. Python,⁵ K. Skovpen,⁵ S. Tavernier,⁵ W. Van Doninck,⁵ P. Van Mulders,⁵ I. Van Parijs,⁵ D. Beghin,⁶ B. Bilin,⁶ H. Brun,⁶ B. Clerbaux,⁶ G. De Lentdecker,⁶ H. Delannoy,⁶ B. Dorney,⁶ G. Fasanella,⁶ L. Favart,⁶ R. Goldouzian,⁶ A. Grebenyuk,⁶ A. K. Kalsi,⁶ T. Lenzi,⁶ J. Luetic,⁶ N. Postiau,⁶ E. Starling,⁶ L. Thomas,⁶ C. Vander Velde,⁶ P. Vanlaer,⁶ D. Vannerom,⁶ Q. Wang,⁶ T. Cornelis,⁷ D. Dobur,⁷ A. Fagot,⁷ M. Gul,⁷ I. Khvastunov,^{7,c} D. Poyraz,⁷ C. Roskas,⁷ D. Trocino,⁷ M. Tytgat,⁷ W. Verbeke,⁷ B. Vermassen,⁷ M. Vit,⁷ N. Zaganidis,⁷ H. Bakhshiansohi,⁸ O. Bondu,⁸ S. Brochet,⁸ G. Bruno,⁸ C. Caputo,⁸ P. David,⁸ C. Delaere,⁸ M. Delcourt,⁸ A. Giammanco,⁸ G. Krintiras,⁸ V. Lemaître,⁸ A. Magitteri,⁸ K. Piotrkowski,⁸ A. Saggio,⁸ M. Vidal Marono,⁸ S. Wertz,⁸ J. Zobec,⁸ F. L. Alves,⁹ G. A. Alves,⁹ M. Correa Martins Junior,⁹ G. Correia Silva,⁹ C. Hensel,⁹ A. Moraes,⁹ M. E. Pol,⁹ P. Rebello Teles,⁹ E. Belchior Batista Das Chagas,¹⁰ W. Carvalho,¹⁰ J. Chinellato,^{10,d} E. Coelho,¹⁰ E. M. Da Costa,¹⁰ G. G. Da Silveira,^{10,e} D. De Jesus Damiao,¹⁰ C. De Oliveira Martins,¹⁰ S. Fonseca De Souza,¹⁰ H. Malbouisson,¹⁰ D. Matos Figueiredo,¹⁰ M. Melo De Almeida,¹⁰ C. Mora Herrera,¹⁰ L. Mundim,¹⁰ H. Nogima,¹⁰ W. L. Prado Da Silva,¹⁰ L. J. Sanchez Rosas,¹⁰ A. Santoro,¹⁰ A. Sznajder,¹⁰ M. Thiel,¹⁰ E. J. Tonelli Manganote,^{10,d} F. Torres Da Silva De Araujo,¹⁰ A. Vilela Pereira,¹⁰ S. Ahuja,^{11a} C. A. Bernardes,^{11a} L. Calligaris,^{11a} T. R. Fernandez Perez Tomei,^{11a} E. M. Gregores,^{11a,11b} P. G. Mercadante,^{11a,11b} S. F. Novaes,^{11a} Sandra S. Padula,^{11a} A. Aleksandrov,¹² R. Hadjiiska,¹² P. Iaydjiev,¹² A. Marinov,¹² M. Misheva,¹² M. Rodozov,¹² M. Shopova,¹² G. Sultanov,¹² A. Dimitrov,¹³ L. Litov,¹³ B. Pavlov,¹³ P. Petkov,¹³ W. Fang,^{14,f} X. Gao,^{14,f} L. Yuan,¹⁴ M. Ahmad,¹⁵ J. G. Bian,¹⁵ G. M. Chen,¹⁵ H. S. Chen,¹⁵ M. Chen,¹⁵ Y. Chen,¹⁵ C. H. Jiang,¹⁵ D. Leggat,¹⁵ H. Liao,¹⁵ Z. Liu,¹⁵ F. Romeo,¹⁵ S. M. Shaheen,^{15,g} A. Spiezia,¹⁵ J. Tao,¹⁵ Z. Wang,¹⁵ E. Yazgan,¹⁵ H. Zhang,¹⁵ S. Zhang,^{15,g} J. Zhao,¹⁵ Y. Ban,¹⁶ G. Chen,¹⁶ A. Levin,¹⁶ J. Li,¹⁶ L. Li,¹⁶ Q. Li,¹⁶ Y. Mao,¹⁶ S. J. Qian,¹⁶ D. Wang,¹⁶ Y. Wang,¹⁷ C. Avila,¹⁸ A. Cabrera,¹⁸ C. A. Carrillo Montoya,¹⁸ L. F. Chaparro Sierra,¹⁸ C. Florez,¹⁸ C. F. González Hernández,¹⁸ M. A. Segura Delgado,¹⁸ B. Courbon,¹⁹ N. Godinovic,¹⁹ D. Lelas,¹⁹ I. Puljak,¹⁹ T. Sculac,¹⁹ Z. Antunovic,²⁰ M. Kovac,²⁰ V. Brigljevic,²¹ D. Ferencek,²¹ K. Kadija,²¹ B. Mesic,²¹ A. Starodumov,^{21,h} T. Susa,²¹ M. W. Ather,²² A. Attikis,²² M. Kolosova,²² G. Mavromanolakis,²² J. Mousa,²² C. Nicolaou,²² F. Ptochos,²² P. A. Razis,²² H. Rykaczewski,²² M. Finger,^{23,i} M. Finger Jr.,^{23,i} E. Ayala,²⁴ E. Carrera Jarrin,²⁵ Y. Assran,^{26,j,k} S. Elgammal,^{26,j} S. Khalil,^{26,l} S. Bhowmik,²⁷ A. Carvalho Antunes De Oliveira,²⁷ R. K. Dewanjee,²⁷ K. Ehataht,²⁷ M. Kadastik,²⁷ M. Raidal,²⁷ C. Veelken,²⁷ P. Eerola,²⁸ H. Kirschenmann,²⁸ J. Pekkanen,²⁸ M. Voutilainen,²⁸ J. Havukainen,²⁹ J. K. Heikkilä,²⁹ T. Järvinen,²⁹ V. Karimäki,²⁹ R. Kinnunen,²⁹ T. Lampén,²⁹ K. Lassila-Perini,²⁹ S. Laurila,²⁹ S. Lehti,²⁹ T. Lindén,²⁹ P. Luukka,²⁹ T. Mäenpää,²⁹ H. Siikonen,²⁹ E. Tuominen,²⁹ J. Tuominiemi,²⁹ T. Tuuva,³⁰ M. Besancon,³¹

F. Couderc,³¹ M. DeJardin,³¹ D. Denegri,³¹ J. L. Faure,³¹ F. Ferri,³¹ S. Ganjour,³¹ A. Givernaud,³¹ P. Gras,³¹ G. Hamel de Monchenault,³¹ P. Jarry,³¹ C. Leloup,³¹ E. Locci,³¹ J. Malcles,³¹ G. Negro,³¹ J. Rander,³¹ A. Rosowsky,³¹ M. Ö. Sahin,³¹ M. Titov,³¹ A. Abdulsalam,^{32,m} C. Amendola,³² I. Antropov,³² F. Beaudette,³² P. Busson,³² C. Charlot,³² R. Granier de Cassagnac,³² I. Kucher,³² A. Lobanov,³² J. Martin Blanco,³² C. Martin Perez,³² M. Nguyen,³² C. Ochando,³² G. Ortona,³² P. Paganini,³² P. Pigard,³² J. Rembser,³² R. Salerno,³² J. B. Sauvan,³² Y. Sirois,³² A. G. Stahl Leiton,³² A. Zabi,³² A. Zghiche,³² J.-L. Agram,^{33,n} J. Andrea,³³ D. Bloch,³³ J.-M. Brom,³³ E. C. Chabert,³³ V. Cherepanov,³³ C. Collard,³³ E. Conte,^{33,n} J.-C. Fontaine,^{33,n} D. Gelé,³³ U. Goerlach,³³ M. Jansová,³³ A.-C. Le Bihan,³³ N. Tonon,³³ P. Van Hove,³³ S. Gadrat,³⁴ S. Beauceron,³⁵ C. Bernet,³⁵ G. Boudoul,³⁵ N. Chanon,³⁵ R. Chierici,³⁵ D. Contardo,³⁵ P. Depasse,³⁵ H. El Mamouni,³⁵ J. Fay,³⁵ L. Finco,³⁵ S. Gascon,³⁵ M. Gouzevitch,³⁵ G. Grenier,³⁵ B. Ille,³⁵ F. Lagarde,³⁵ I. B. Laktineh,³⁵ H. Lattaud,³⁵ M. Lethuillier,³⁵ L. Mirabito,³⁵ S. Perries,³⁵ A. Popov,^{35,o} V. Sordini,³⁵ G. Touquet,³⁵ M. Vander Donckt,³⁵ S. Viret,³⁵ T. Toriashvili,^{36,p} D. Lomidze,³⁷ C. Autermann,³⁸ L. Feld,³⁸ M. K. Kiesel,³⁸ K. Klein,³⁸ M. Lipinski,³⁸ M. Preuten,³⁸ M. P. Rauch,³⁸ C. Schomakers,³⁸ J. Schulz,³⁸ M. Teroerde,³⁸ B. Wittmer,³⁸ A. Albert,³⁹ D. Duchardt,³⁹ M. Erdmann,³⁹ S. Erdweg,³⁹ T. Esch,³⁹ R. Fischer,³⁹ S. Ghosh,³⁹ A. Güth,³⁹ T. Hebbeker,³⁹ C. Heidemann,³⁹ K. Hoepfner,³⁹ H. Keller,³⁹ L. Mastrolorenzo,³⁹ M. Merschmeyer,³⁹ A. Meyer,³⁹ P. Millet,³⁹ S. Mukherjee,³⁹ T. Pook,³⁹ M. Radziej,³⁹ H. Reithler,³⁹ M. Rieger,³⁹ A. Schmidt,³⁹ D. Teyssier,³⁹ S. Thüer,³⁹ G. Flügge,⁴⁰ O. Hlushchenko,⁴⁰ T. Kress,⁴⁰ T. Müller,⁴⁰ A. Nehr Korn,⁴⁰ A. Nowack,⁴⁰ C. Pistone,⁴⁰ O. Pooth,⁴⁰ D. Roy,⁴⁰ H. Sert,⁴⁰ A. Stahl,^{40,q} M. Aldaya Martin,⁴¹ T. Arndt,⁴¹ C. Asawatangtrakuldee,⁴¹ I. Babounikau,⁴¹ K. 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