

## Phase transitions in $\epsilon$ -FeOOH at high pressure and ambient temperature

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### ABSTRACT

Constraining the accommodation, distribution, and circulation of hydrogen in the Earth's interior is vital to our broader understanding of the deep Earth due to the significant influence of hydrogen on the material and rheological properties of minerals. Recently, a great deal of attention has been paid to the high-pressure polymorphs of FeOOH (space groups  $P2_1nm$  and  $Pnnm$ ). These structures potentially form a hydrogen-bearing solid solution with AlOOH and phase H ( $MgSiO_4H_2$ ) that may transport water ( $OH^-$ ) deep into the Earth's lower mantle. Additionally, the pyrite-type polymorph (space group  $Pa\bar{3}$  of FeOOH), and its potential dehydration have been linked to phenomena as diverse as the introduction of hydrogen into the outer core (Nishi et al. 2017), the formation of ultralow-velocity zones (ULVZs) (Liu et al. 2017), and the Great Oxidation Event (Hu et al. 2016). In this study, the high-pressure evolution of FeOOH was re-evaluated up to  $\sim 75$  GPa using a combination of synchrotron-based X-ray diffraction (XRD), Fourier transform infrared spectroscopy (FTIR), and optical absorption spectroscopy. Based on these measurements, we report three principal findings: (1) pressure-induced changes in hydrogen bonding (proton disordering or hydrogen bond symmetrization) occur at substantially lower pressures in  $\epsilon$ -FeOOH than previously reported and are unlikely to be linked to the high-spin to low-spin transition; (2)  $\epsilon$ -FeOOH undergoes a 10% volume collapse coincident with an isostructural  $Pnnm \rightarrow Pnnm$  transition at approximately 45 GPa; and (3) a pressure-induced band gap reduction is observed in FeOOH at pressures consistent with the previously reported spin transition (40 to 50 GPa).

**Keywords:** Phase transitions, infrared-spectroscopy, optical absorption spectroscopy, X-ray diffraction, spin transition; Volatile Elements in Differentiated Planetary Interiors

### INTRODUCTION

The accommodation, distribution, and circulation of hydrogen in the deep Earth is key to understanding the evolution of Earth's interior due to the significant influence of hydrogen on the material and rheological properties of high-pressure phases (e.g., Sarafian et al. 2017; Karato 2010). Geophysical observations indicate that the Earth's upper mantle and transition zone, at least locally, host significant quantities of hydrogen (van der Meijde et al. 2003; Dixon et al. 2004; Pearson et al. 2014; Palot et al. 2016; Tschauner et al. 2018). Additionally, tomographic evidence supports the idea that subducting lithospheric plates pierce the transition zone, potentially ushering water into the Earth's lower mantle (van der Hilst et al. 1997). Yet while hydrogen in the upper mantle and transition zone is hosted primarily in nominally anhydrous phases, these phases are not stable at the high pressure and temperature conditions of the lower mantle (Hirschmann 2006). If hydrogen introduced into the lower mantle remains in the lower mantle, it is almost certainly accommodated primarily in minor hydrous phases, because the dominant lower mantle minerals (bridgmanite, ferropericlase, and calcium silicate perovskite) do not have the same capacity for water storage (Bolfan-Casanova et al. 2002,

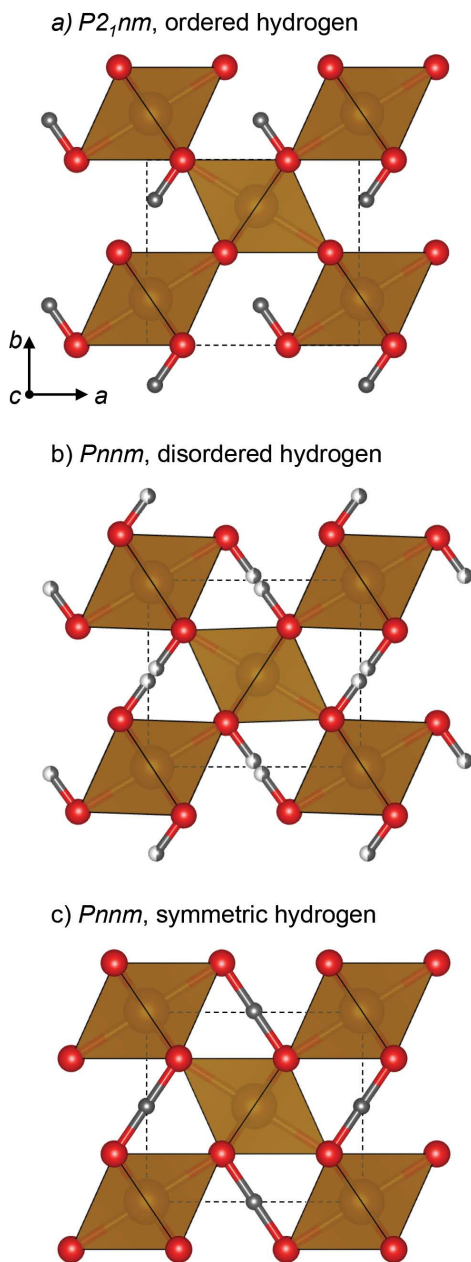
2003; Panero et al. 2015). A plausible lower mantle host is the isostructural FeOOH-AlOOH- $MgSiO_4H_2$  system.

At moderate pressures ( $>6$  GPa) goethite ( $\alpha$ -FeOOH, space group  $Pbnm$ ), a widespread iron oxy-hydroxide, transforms into  $\epsilon$ -FeOOH, an orthorhombic (space group  $P2_1nm$ ,  $Z = 2$ ) phase composed of edge-sharing  $FeO_6$  units that are close-packed along the  $c$ -axis, with hydrogen atoms occupying the channels between these close-packed octahedra (Bendeliani et al. 1972; Pernet et al. 1975; Bolotina et al. 2008) (Fig. 1a). At increased pressure,  $\epsilon$ -FeOOH undergoes pressure-induced hydrogen-bond symmetrization resulting in a second-order phase transition (space group  $Pnnm$ ,  $Z = 2$ ) (Fig. 1c). However, the pressure at which this hydrogen bond symmetrization induced  $P2_1nm \rightarrow Pnnm$  transition occurs remains contested (Gleason et al. 2013; Xu et al. 2013; Thompson et al. 2017; Ikeda et al. 2019). At the pressures of the Earth's lower mantle,  $\epsilon$ -FeOOH forms a solid solution with  $\delta$ -AlOOH and phase H ( $MgSiH_2O_4$ ), creating a viable hydrogen reservoir with  $P$ - $T$  stability extending from lower mantle conditions to those of the core-mantle boundary (Sano et al. 2008; Nishi et al. 2015; Xu et al. 2019).

Hydrogen bond symmetrization is the process by which the hydrogen in an asymmetrical  $O-H\cdots O$  bonding unit becomes centered with respect to the two O atoms with increased pressure, as the longer hydrogen bridge bond ( $H\cdots O$ ) compresses more rapidly than the shorter but stiffer hydroxyl bond ( $O-H$ ) (Holzapfel 1972).

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**FIGURE 1.** Low and moderate pressure structures of  $\epsilon$ -FeOOH including the (a)  $P2_1nm$  structure with ordered hydrogen occupation and asymmetric hydrogen bonds, (b)  $Pnnm$  structure with disordered hydrogen (hydrogen positions are 50% occupied), and (c)  $Pnnm$  structure with symmetric hydrogen bonds. Dashed lines indicate unit cells, oxygen atoms are red spheres, hydrogen atoms are gray spheres, and  $FeO_6$  units are tan polyhedra, which are stacked in edge-sharing chains parallel to the  $c$ -axis. Image generated in VESTA (Momma and Izumi 2008).

Once the hydrogen bond is “symmetrized,” the two donor-acceptor distances are equivalent, and this shift in charge balance can lead to additional subtle structural changes (Fig. 1c). In  $\epsilon$ -FeOOH and other isostructural MOOH phases, the  $MO_6$  octahedra shift from a slightly offset position to one centered along the twofold axis,

leading to the increase in symmetry from  $P2_1nm$  to  $Pnnm$  (e.g., Sano-Furukawa et al. 2009, 2012). Alternatively, proton disorder in which both hydrogen and iron atoms are disordered over symmetric positions on the twofold axis may also produce MOOH with  $Pnnm$  symmetry (Bolotina et al. 2008) with the resulting diffraction patterns nearly indistinguishable from symmetrized H-bond structures (e.g., Fujihara et al. 2002) (Fig. 1b). Importantly, a recent neutron diffraction study by Sano-Furukawa et al. (2018), found that in  $\delta$ -AlOOH, which is isostructural to  $\epsilon$ -FeOOH, proton disordering is a precursor to hydrogen bond symmetrization. Accurately determining the pressure at which hydrogen bond symmetrization occurs is important because the phenomenon is linked to changes in compressibility (Vanpeteghem et al. 2003; Tsuchiya et al. 2005; Hushur et al. 2011) and because isotopic effects may influence H/D fractionation in the deep Earth (e.g., Sano-Furukawa et al. 2009).

In addition to the previously described  $P2_1nm \rightarrow Pnnm$  transition,  $\epsilon$ -FeOOH reportedly undergoes a high-spin to low-spin transition, determined on the basis of X-ray diffraction (XRD) and X-ray emission spectroscopy (XES) experiments (Gleason et al. 2013). The XES results from Gleason et al. (2013) are indicative of a high-spin to low-spin transition initiated at  $\sim 40$  GPa and completed at  $\sim 60$  GPa. However, that study did not use a pressure-transmitting medium, which introduced significant deviatoric stress into the sample, inhibiting the detection of a sharp transition. A contemporary study by Xu et al. (2013), which used a pressure-transmitting medium, pinpointed this electronic transition more closely to  $\sim 45$  GPa using a combination of Mössbauer spectroscopy, single-crystal diffraction, and resistivity measurements. In both studies, the authors propose causality between the hydrogen bond symmetrization and spin transition in  $\epsilon$ -FeOOH, although Gleason et al. (2013) report that hydrogen bond symmetrization induces the spin transition, whereas Xu et al. (2013) conclude the spin transition induces hydrogen-bond symmetrization. However, both studies agree that a large ( $\sim 11\%$ ) volume collapse occurs coincident with the high-spin to low-spin transition, attributable to an isostructural phase transition. Significant volume reductions have been previously reported due to high-spin to low-spin transitions without invoking hydrogen-bond symmetrization, although a volume reduction of  $\sim 11\%$  is higher than the volume reductions reported for either ferropentacite [(Mg,Fe)O] (1–3%) or ferromagnesite [(Mg,Fe)CO<sub>3</sub>] (6–10%) (Lin et al. 2013), possibly due to the larger relative volume occupied by iron atoms in FeOOH compared to these phases.

In this study, we present experimental evidence of two distinct transitions in  $\epsilon$ -FeOOH in the explored pressure range (0–70 GPa): a low-pressure ( $\sim 18$  GPa) second-order transition connected to hydrogen bond symmetrization and an independent first-order transition at  $\sim 45$  GPa. Additionally, we report evidence of a rapid reduction in the bandgap of  $\epsilon$ -FeOOH at pressures consistent with the previously reported spin transition.

## METHODS

Synthesis of  $\epsilon$ -FeOOH for this study was performed by Akio Suzuki and is described in Suzuki (2010). The high pressures necessary for this study were achieved by compressing  $\epsilon$ -FeOOH samples using symmetric-type diamond-anvil cells (DACs) with 250  $\mu m$  culet type II diamond anvils. Seventy-micrometer diameter sample chambers were laser ablated into stainless steel or rhenium gaskets pre-indenting to 22 or 28 GPa, respectively. Powdered  $\epsilon$ -FeOOH, kept in a desiccator

prior to loading, was pressed into platelets 3–5  $\mu\text{m}$  thick and  $\sim 30 \mu\text{m}$  in diameter. Platelets of  $\epsilon$ -FeOOH were loaded into the sample chamber, utilizing gas-loaded Ne or  $\sim 10 \mu\text{m}$  thick platelets of KBr as a pressure medium to reduce pressure gradients within the sample. Neon was loaded as a pressurized gas at the Advanced Photon Source using the COMPRES/GSECARS gas loading system (Rivers et al. 2008) and acted as a secondary pressure standard (Fei et al. 2007), with reported uncertainties based on the standard error of the (111), (200), and (220)  $d$ -spacings. Pressures were independently determined via the in situ monitoring of the  $R_1$  luminescence line (Dewaele et al. 2008) of 2–3 ruby grains placed in the sample chamber, with errors reflecting the larger of either the standard deviation or 3%.

Room-temperature X-ray diffraction experiments were performed at beamline 13-ID-D (GSECARS) at the Advanced Photon Source, Argonne National Laboratory. A monochromatic ( $\lambda = 0.3344$ ) incident X-ray beam was used, measuring 3  $\mu\text{m}$  by 4  $\mu\text{m}$  at full-width at half maximum of the focused spot. Sample-to-detector distances and tilt were calibrated using 1-bar diffraction of  $\text{LaB}_6$ . Diffraction patterns were integrated to produce 2 $\theta$  plots using DIOPTAS (Prescher and Prakapenka 2015), positions of individual diffraction peaks were determined using PeakFit (Systat Software), and lattice parameters were calculated from the fitted  $d$ -spacings using the author's own Mathematica script. Individual peaks were fit to single Gaussian curves, and unresolved overlapping peaks were not used in the calculation of lattice parameters. All samples prepared for X-ray diffraction experiments were loaded using Ne as a pressure medium.

Fourier transform infrared (FTIR) spectra in this study were collected using both synchrotron and global sources. Synchrotron experiments were performed at the 1.4.3 beamline of the Advanced Light Source at Berkeley Laboratory using a Nicolet 760 FTIR spectrometer with a custom microscope and HgCdTe detectors. Offline experiments were performed at Brookhaven National Laboratory using a Bruker Vertex 80v spectrometer and Hyperion 2000 microscope with an MCT detector. All FTIR spectra presented here reflect a spectral resolution of 4  $\text{cm}^{-1}$ , measured wavelength ranges of 500 to 8000  $\text{cm}^{-1}$ , and were recorded as 1024 or 512 scans. PeakFit was used for background subtraction and to obtain precise absorption peak positions using a least-squares refinement, and interference fringes in absorbance spectra were reduced mathematically in DatLab. Samples prepared for FTIR experiments used either Ne or KBr as pressure medium, and the three samples that went to the highest pressures used KBr.

Optical absorption measurements were collected at the University of Chicago. Spectra were recorded from 420 to 980 nm using a 0.3 m focal length Princeton Instruments SP-2300i spectrograph. Transmitted light ( $I_t$ ), generated with a tungsten filament bulb, was measured through the pressure medium (KBr) only, while incident light ( $I_0$ ) was collected through the sample and pressure medium. Background measurements were collected both with the light source off and with the light source on but measured at the gasket, which is opaque, with comparable results. Interference fringes in the absorbance spectra were removed mathematically in DatLab. To account for variation in sample thickness ( $h$ ), reported absorption values [ $\alpha = h^{-1} \ln(I_0/I_t)$ ] have been scaled such that  $I_0/I_t$  is equal for all samples at 10 GPa. Samples prepared for optical absorption measurements used KBr as pressure medium.

## RESULTS AND DISCUSSION

### X-ray diffraction

High-pressure, room-temperature X-ray diffraction (XRD) measurements were used to re-evaluate the phase diagram of FeOOH up to 75 GPa, as the existing literature exhibits a paucity of volume-pressure ( $V$ - $P$ ) data in the mid-pressure range (20–50 GPa) (Gleason et al. 2008, 2013; Suzuki 2010, 2016; Ikeda et al. 2019). Additionally, using a gas-membrane diaphragm to remotely pressurize the DAC between diffraction measurements without manual action, data could be collected with greater pressure resolution than in previous studies. As X-ray diffraction inherently relies on electron density, the detection of hydrogen using powder XRD is virtually impossible. The  $P2_1nm \rightarrow Pnnm$  transition is primarily defined by the pressure-induced symmetrization of the hydrogen bonds but associated with hydrogen bond disordering and symmetrization are subtle changes in the structure and strain accommodation, which enable the detection of these transitions indirectly (e.g., Sano-Furukawa et al. 2009, 2012; Kuribayashi et al. 2014).

In the lower pressure interval (0.5–43 GPa) the lattice parameters of  $\epsilon$ -FeOOH were determined using at a minimum the (110), (101), (011), (020), (210), (211), (121), (220), (002), and (301)  $hkl$  peaks. The resulting unit-cell volumes and lattice parameters are plotted in Figures 2a and 2b and are reported in tabulated form in Supplemental<sup>1</sup> Table S1. The measured lattice parameters and unit-cell volumes of  $\epsilon$ -FeOOH in this pressure interval are in excellent agreement with the results of Suzuki (2010, 2016) and Ikeda et al. (2019), albeit slightly lower than those reported by the Gleason et al. (2013). The diffraction data from Gleason et al. (2013) exhibit more scatter, likely due to the lack of pressure medium in that study, which likely contributed to the slight offset in measured  $P$ - $V$  values.

Based on the Bragg peaks indexed in this study, the  $P2_1nm$  and  $Pnnm$  structure cannot be distinguished from one another, as even at the lowest pressures of this study no peaks were observed that violated the additional systematic absences characteristic of the  $Pnnm$  symmetry ( $h + l = \text{odd}$  for  $h0l$ ,  $k + l = \text{odd}$  for  $0kl$ ). Additionally, based on the findings of Sano-Furukawa et al. (2009), who evaluated the  $P2_1nm \rightarrow Pnnm$  transition in isostructural  $\delta$ -AlOOH, the volume reduction associated with this second-order transition is likely unresolvable, rendering determinations of this phase boundary on this basis alone dubious. Thus, employing the approach of Sano-Furukawa et al. (2009, 2012), we evaluated the ratios of lattice parameters as a means of elucidating the subtle  $P2_1nm \rightarrow Pnnm$  transition.

The evolution of the  $a/b$  and  $b/c$  lattice parameter ratios (Figs. 2c and 2d) reveal a change in the axial compression of  $\epsilon$ -FeOOH with pressure, and the inflection points evident at  $18 \pm 1$  GPa indicate a shift in the accommodation of strain in this system. Similar changes in axial compression have been used previously as an indicator of second-order phase changes in similar materials (e.g., Sano-Furukawa et al. 2009, 2012; Kuribayashi et al. 2014). On the basis of these lattice parameter ratios, we confirm that  $\epsilon$ -FeOOH undergoes a second-order phase transition from the  $P2_1nm$  to  $Pnnm$  structure at  $18 \pm 1$  GPa, in good agreement with previous predictions based on density functional theory-based calculations (Thompson et al. 2017) and recent experiments (Ikeda et al. 2019). As any volume reduction across this presumed phase boundary is minimal (Fig. 2a), the  $V$ - $P$  data from the entire pressure interval (0 to 45 GPa) was fit to a third-order Birch-Murnaghan equation of state (EoS) (Birch 1978):

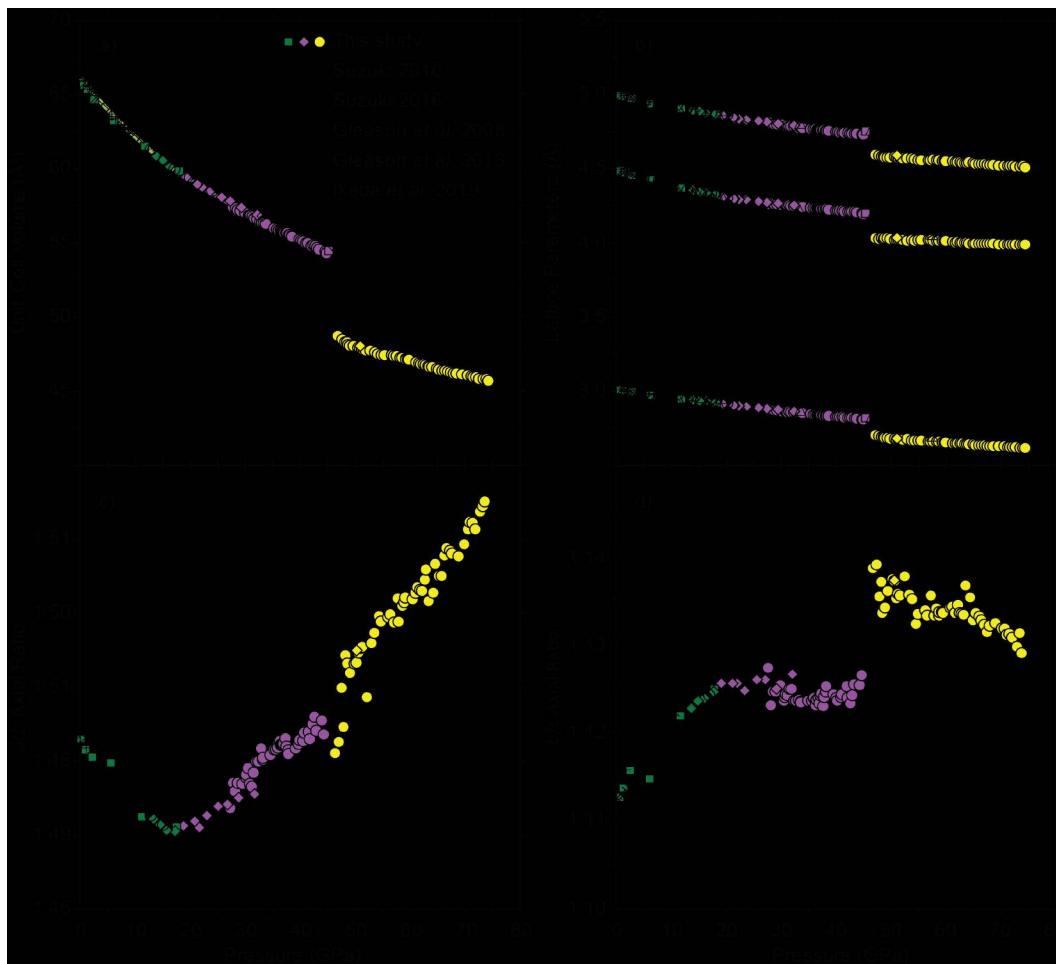
$$P(V, T) = 3K_0 f_E \left[ (1 + 2f_E)^{\frac{5}{3}} \left( 1 + \frac{3}{2} [K'_0 - 4] f_E \right) \right]$$

which relates pressure ( $P$ ), volume ( $V$ ), ambient pressure bulk modulus ( $K_0$ ), and its pressure derivative ( $K'_0$ ) in terms of finite Eulerian strain ( $f_E$ ):

$$f_E = \frac{1}{2} \left[ \left( \frac{V_0}{V} \right)^{\frac{2}{3}} - 1 \right]$$

which is a measure of the volume compression of a solid relative to its initial volume ( $V_0$ ). Residuals from this fit are shown in Supplemental<sup>1</sup> Figure S1.

To obtain EoS parameters with the most comprehensive data set available, the  $V$ - $P$  data from this study and Suzuki (2010, 2016) were combined, and the resultant equation of state pa-



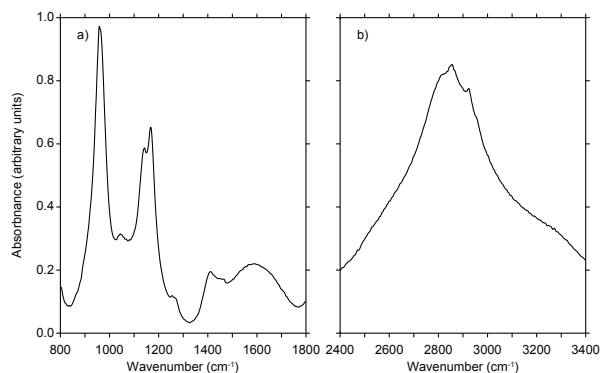
**FIGURE 2.** (a) Unit-cell volume, (b) lattice parameters, (c)  $a/b$ , and (d)  $b/c$  lattice parameter ratios of FeOOH as a function of pressure, including this study (solid symbols), Suzuki (2010) (black crosses), Suzuki (2016) (black X's), Gleason et al. (2008) (open diamonds), Gleason et al. (2013) (open triangles), and Ikeda et al. (2019) (open circles). Different solid symbols (circles, diamonds, squares) are used to show replicate samples in this study. Green, purple, and yellow symbols indicate that Bragg peaks have been indexed to the  $P2_1nm$  (green), high-spin  $Pnnm$  (purple), and low-spin  $Pnnm$  (yellow) structures. Light green and light purple lines in **a** represent the equations of state using the parameters from lines 3 and 4 of Table 1, respectively. Error bars reflect uncertainties in the lattice parameters from this study, and when not visible reflect that errors are smaller than the symbols.

rameters are presented alongside parameters fit to the data from only this study. In addition to fitting  $V$ - $P$  data from the entire 0–45 GPa pressure range, we also report EoS parameters derived from fitting the  $V$ - $P$  data of the  $P2_1nm$  and  $Pnnm$  structures separately. The resultant EoS parameters, as well as previously published results, are reported in Table 1. We find that EoS parameters derived from fitting the combined data set across the entire pressure range (0–45 GPa) are in reasonable agreement with those reported by Gleason et al. (2008) and Suzuki (2010, 2016). However, fitting  $V$ - $P$  data for the  $P2_1nm$  and  $Pnnm$  structures independently (using a fixed  $K'_0$  value of 4 because fewer fitted parameters are justified over the limited pressure ranges), reveals that the  $Pnnm$  structure has a bulk modulus ~20% higher than that of the  $P2_1nm$   $\epsilon$ -FeOOH, in good agreement with the ~19% predicted from DFT calculations (Thompson et al. 2017). These findings are also in good agreement with a prior study by Xu et al. (2013) that identified a reduction in the compressibility

of FeOOH at ~16 GPa but did not connect this change in material properties to the onset of pressure-induced hydrogen-bond symmetrization or proton disordering. Similar increases in bulk modulus have been reported for other hydrous phases due to the onset of pressure-induced hydrogen-bond symmetrization or proton disordering (Vanpeteghem et al. 2003; Tsuchiya et al. 2005; Hushur et al. 2011).

Additionally, XRD data from this study revealed a second, independent structural transition in  $\epsilon$ -FeOOH at ~45 GPa. The transition occurs over a small (<3 GPa) pressure interval and is evident in the appearance of multiple new Bragg peaks and a simultaneous reduction in the intensity, and eventual disappearance, of the Bragg peaks of the high-spin  $Pnnm$  structure (Fig. 3). This high-pressure (>45 GPa) structure of FeOOH is stable at ambient temperatures and pressures exceeding 70 GPa. While an earlier study identified a potential unit-cell volume reduction in this pressure range (Gleason et al. 2013), the



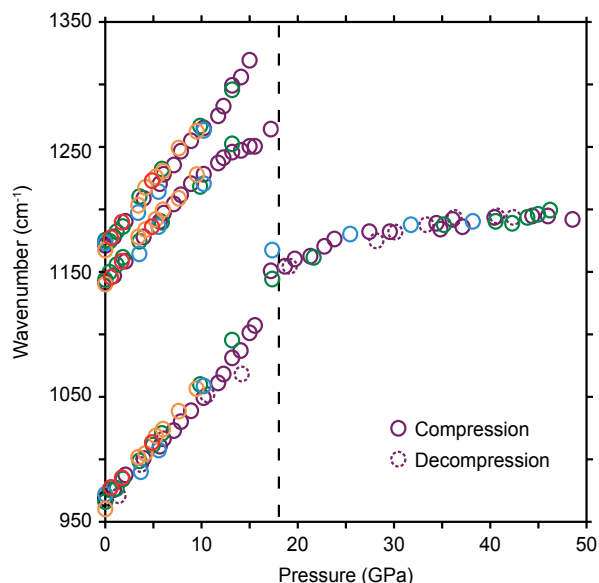


**FIGURE 4.** Ambient pressure infrared spectra of  $P2_1nm$   $\epsilon$ -FeOOH in the (a) lattice and OH-bending vibrations region and (b) OH-stretching region.

hydrogen-bond symmetrization. However, it is not feasible to predict the pressure of eventual hydrogen-bond symmetrization, as the relationship between the 1-bar stretching frequency and O–H pressure dependence is poorly understood in  $\epsilon$ -FeOOH and similarly structured phases, and the pressure range of available  $\epsilon$ -FeOOH O–H stretching frequency data is limited.

The ambient-pressure infrared spectrum of  $\epsilon$ -FeOOH also contains three distinct absorption bands at 965, 1145, and 1175  $\text{cm}^{-1}$  that were identified as O–H bending vibrations based on comparison to similar structures (Williams and Guenther 1996; Mashino et al. 2016; Pinney and Morgan 2013). Spectra at successive pressure steps were evaluated to determine the frequency shifts in the O–H bending as a function of pressure (Fig. 5). As with the OH-stretching region, these features broaden with increased pressure, and the two higher-frequency bands become indistinguishable above  $17.5 \pm 1$  GPa. The lowest-frequency bending band remains distinct to higher pressures, but the pressure dependence of this band changes dramatically above  $17.5 \pm 1$  GPa. The sudden change in the pressure dependence of this band suggests a reorientation of the O–H bonds, suggestive of the onset of pressure-induced hydrogen-bond symmetrization or proton disordering. A similar shift in the pressure dependence of the OH-absorption bands in isostructural  $\delta$ -AlOOH was observed by Kagi et al. (2010) at  $\sim 10$  GPa, consistent with the onset of proton disordering in that phase (Sano-Furukawa et al. 2018).

During the course of these IR measurements, a reduction in the transmitted IR signal was observed in  $\epsilon$ -FeOOH samples compressed above  $\sim 40$  GPa (Supplemental<sup>1</sup> Fig. S4). In addition to the changes in IR transmission properties, changes in the optical properties of FeOOH were observed upon compression. The samples were translucent orange at ambient and low pressures, but reddened with increased pressure and became opaque to transmitted visible light at 45 GPa, slightly preceding the loss of IR transmission (Supplemental<sup>1</sup> Fig. S4). To quantify these observations, visible to near-infrared absorption measurements of  $\epsilon$ -FeOOH were collected up to pressures exceeding 70 GPa, to track the pressure dependence of the absorption edge. Representative high-pressure optical absorption spectra of  $\epsilon$ -FeOOH are shown in Figure 6a and indicate a systematic pressure-induced increase in absorption concordant with qualitative observations. Normalized data for five samples are plotted in Figure 6b, in which there are two key features. First, each sample underwent



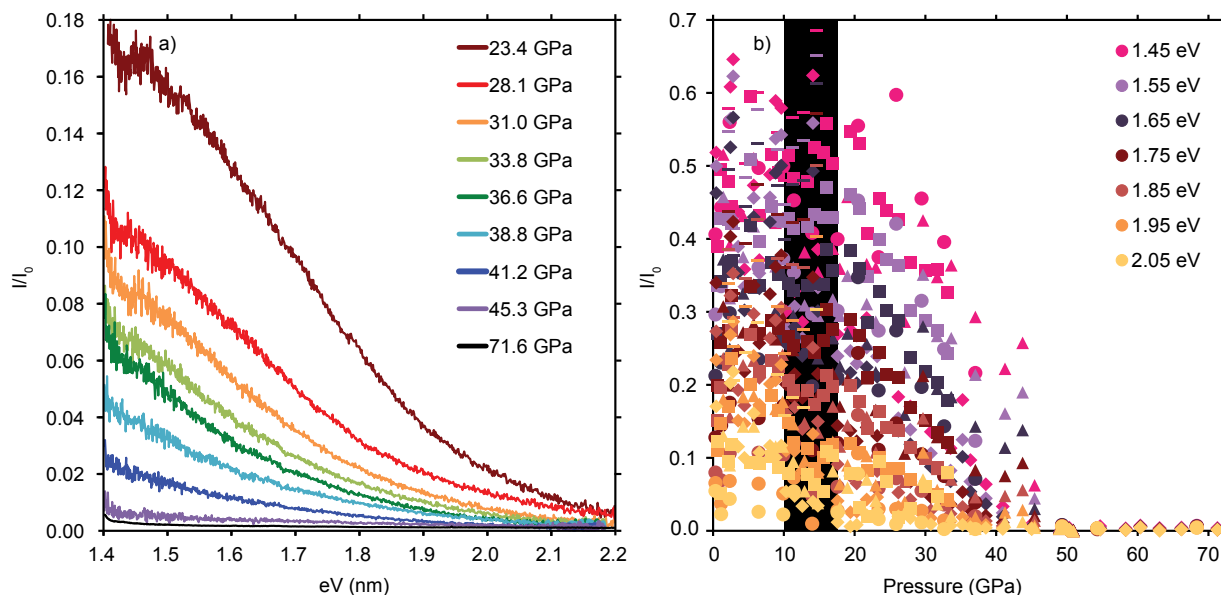
**FIGURE 5.** Evolution of the O–H bending frequencies of FeOOH as a function of pressure. The pressure of the  $P2_1nm \rightarrow Pnnm$  transition is indicated by a black dashed line. Different colors are used to distinguish the five replicate samples from this study using either Ne as a pressure medium (red, orange) or KBr as a pressure medium (green, blue, purple).

a transient minimum in optical transmission between 15–20 GPa before gradually recovering to previous values over the subsequent  $\sim 5$  GPa. Second, all five samples became opaque in the light range probed (1.45 to 2.05 eV) above 45.7 GPa (Fig. 6b), above which the ratio of transmitted light ( $I$ ) to incident light ( $I_0$ ) that can propagate through the sample decreases to zero. These vis/NIR absorption findings suggest a significant reduction in the bandgap in this pressure interval.

## IMPLICATIONS

Close evaluation of the high-pressure behavior of  $\epsilon$ -FeOOH at ambient temperature indicates two sequential, structural phase transitions in the 0–75 GPa pressure range. A second-order transition at  $18 \pm 1$  GPa was determined on the basis of three observations: (1) a change in axial compressibility evident in the lattice parameter ratios measured using powder XRD; (2) a change in the OH-bending modes observed with FTIR spectroscopy; and (3) a transient decrease in optical absorption. In accordance with the previous work of Sano-Furukawa et al. (2009, 2012, 2018) that evaluated isostructural MOOH phases, this  $18 \pm 1$  GPa transition is the  $P2_1nm \rightarrow Pnnm$  transition produced by either pressure-induced hydrogen-bond symmetrization or proton disordering, with future neutron diffraction work needed to differentiate between these mechanisms. Furthermore, this transition is observed to lead to a  $\sim 20\%$  increase in bulk modulus and could be linked to the increase in compressional velocity ( $V_p$ ) reported by Ikeda et al. (2019).

At  $45 \pm 2$  GPa,  $Pnnm$   $\epsilon$ -FeOOH undergoes a first-order structural phase transition, identifiable due to the appearance of new Bragg peaks coincident with the disappearance of the high-spin  $Pnnm$  peaks. This high-pressure structure of FeOOH



**FIGURE 6.** Optical absorption data, including (a) representative optical transmission spectra, indicating the pressure-dependent increase in the absorption of  $\epsilon$ -FeOOH, and (b) data from five samples of  $\epsilon$ -FeOOH, each indicated with a different symbol (diamonds, squares, circles, triangles, and horizontal dashes), showing the change in the ratio of transmitted light ( $I$ ) to incident light ( $I_0$ ) as a function of pressure (X-axis) and the wavelength of the incident light, which is indicated by color.

can also be indexed to  $Pn\bar{m}$  with a  $\sim 10\%$  volume reduction, in good agreement with previous studies (Gleason et al. 2013; Xu et al. 2013). Therefore, we propose that the first-order transition at  $45 \pm 2$  GPa is a nearly isostructural S volume collapse (i.e.,  $Pn\bar{m} \rightarrow Pn\bar{m}$ ), accompanied by a small ( $\sim 1\%$ ) increase in the  $a/b$  ratio in the higher-pressure structure (Fig. 2d). As such, low-spin  $Pn\bar{m}$  FeOOH is stable across a wide pressure range (more than 30 GPa). However, the phase diagram of FeOOH at high pressures and moderate temperatures should be assessed to determine the relevance of low-spin  $Pn\bar{m}$  FeOOH to the Earth's interior, as the formation of the pyrite ( $Pa\bar{3}$ ) structure is reported at comparable pressures and high temperatures (Nishi et al. 2017; Hu et al. 2016). In the future, single-crystal X-ray diffraction structure refinements may be beneficial in verifying the symmetry of this higher pressure, low-spin structure.

In addition to the first-order structural transition identified in FeOOH at  $\sim 45$  GPa, a color change (translucent orange to black) was observed at high pressure, and the optical opacity of  $Pn\bar{m}$   $\epsilon$ -FeOOH was quantified using visible/NIR transmission and infrared absorption. Visible to near-infrared absorption measurements of  $\epsilon$ -FeOOH also identified a systematic pressure-induced increase in absorption consistent with band gap reduction in the high-spin  $Pn\bar{m}$  phase, although complete band gap collapse (i.e., metallization) of this phase cannot be established on the basis of these experiments alone. FeOOH becomes opaque in the 1.45–2.05 eV range and also in the mid-IR at  $\sim 45$  GPa, indicating a change in optical properties along with the high-spin to low-spin transition in this material, as both phenomena are tied to changes in the d orbitals of the iron atoms. In fact, a similar increase in optical opacity has been previously linked to the high-spin to low-spin transition in siderite (Lobanov et al. 2015) and ferroperricite (Keppler et al. 2007). The apparent width of this transition, when compared to the

sharpness of the structural transition observed using diffraction, may be attributable to the difference in the hydrostaticity of the pressure medium (KBr and Ne, respectively).

These results demonstrate that bandgap reduction, the previously observed high-spin to low-spin spin transition (Gleason et al. 2013; Xu et al. 2013), and the first-order phase transition in  $\epsilon$ -FeOOH all occur at a pressure ( $45 \pm 2$  GPa) greatly exceeding the  $P2_1nm \rightarrow Pn\bar{m}$  symmetry transition associated with a change in hydrogen bonding ( $18 \pm 1$  GPa). Therefore, hydrogen-bond symmetrization is not the driver of spin transition in  $\epsilon$ -FeOOH or vice versa. Rather, FeOOH undergoes a second-order  $P2_1nm \rightarrow Pn\bar{m}$  symmetry change caused by a shift in hydrogen bonding at 18 GPa, and a subsequent, unrelated first-order transition at 45 GPa that involves a significant volume change, spin transition, and optical opacity. Furthermore, although end-member  $\epsilon$ -FeOOH is expected to transition to the pyrite ( $Pa\bar{3}$ ) structure at the concurrent pressures and high temperatures of the Earth's mantle (Nishi et al. 2017; Hu et al. 2016), aluminum substitution is expected to increase the thermodynamic stability of (Fe,Al)OOH such that the low-spin  $Pn\bar{m}$  structure may be stable at the conditions of the deep Earth (e.g., Nishi et al. 2015; Xu et al. 2019). We anticipate that the high-spin to low-spin spin transition and associated volume collapse in intermediate compositions in the FeOOH-AlOOH-MgSiH<sub>2</sub>O<sub>4</sub> system are likely to occur at lower pressures than in the Fe-end-member, as has been observed in ferroperricite (Fei et al. 2007). Additional experiments are needed to assess the influence of cation substitution on the phase boundaries of the FeOOH-AlOOH-MgSiO<sub>4</sub>H<sub>2</sub> system, as this may play a pivotal role in the transport of water (OH<sup>-</sup>) deep into the Earth's lower mantle (Ohira et al. 2014; Xu et al. 2019) and may contribute to geophysical heterogeneities observed in the deep Earth (Thompson et al. 2017; Ohira et al. 2019).

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## Endnote:

<sup>1</sup>Deposit item AM-20-127468, Supplemental Figures and Tables. Deposit items are free to all readers and found on the MSA website, via the specific issue's Table of Contents (go to [http://www.minsocam.org/MSA/AmMin/TOC/2020/Dec2020\\_data/Dec2020\\_data.html](http://www.minsocam.org/MSA/AmMin/TOC/2020/Dec2020_data/Dec2020_data.html)).