

Universal properties of Kaluza-Klein gravitons

Kevin Dimmitt,^a Gabriel Larios,^b Praxitelis Ntokos^a and Oscar Varela^{a,b}

^a*Department of Physics, Utah State University,
Logan, UT 84322, U.S.A.*

^b*Departamento de Física Teórica and Instituto de Física Teórica UAM/CSIC,
Universidad Autónoma de Madrid, Cantoblanco, 28049 Madrid, Spain*

E-mail: kevin.dimitt@aggiemail.usu.edu, gabriel.larios@uam.es,
praxitelis.ntokos@usu.edu, oscar.varela@usu.edu

ABSTRACT: Vacua of different gaugings of $D = 4$ $\mathcal{N} = 8$ supergravity that preserve the same supersymmetries and bosonic symmetry tend to exhibit the same universal mass spectrum within their respective supergravities. For AdS_4 vacua in gauged supergravities that arise upon consistent truncation of string/M-theory, we show that this universality is lost at higher Kaluza-Klein levels. However, universality is still maintained in a milder form, at least in the graviton sector: certain sums over a finite number of states remain universal. Further, we derive a mass matrix for Kaluza-Klein gravitons which is valid for all the AdS_4 vacua in string/M-theory that uplift from the gaugings of $D = 4$ $\mathcal{N} = 8$ supergravity that we consider. The mild universality of graviton mass sums is related to the trace of this mass matrix.

KEYWORDS: AdS-CFT Correspondence, Extended Supersymmetry, Supergravity Models

ARXIV EPRINT: [1911.12202](https://arxiv.org/abs/1911.12202)

Contents

1	Introduction	1
2	Massive gravitons with at least SU(3) symmetry	4
2.1	M-theory	4
2.1.1	Massive gravitons with at least G_2 symmetry	4
2.1.2	Massive gravitons with at least $SU(4)_c$ symmetry	7
2.2	Type IIB	9
2.3	Individual spectra in M-theory, type IIA and type IIB	10
2.4	Universality of traces	12
3	Graviton spectrum on the $\mathcal{N} = 4$ SO(4) type IIB S-fold	16
4	Kaluza-Klein graviton mass matrix	19
4.1	The mass matrix	19
4.2	Mass matrix traces	21
5	Discussion	22
A	SU(3) invariance in $D = 4$ $\mathcal{N} = 8$ gauged supergravity	24

1 Introduction

It is now well established that maximal gauged supergravities in four dimensions typically come in one-parameter families [1–3]. All members in a given family share the same gauge group. The parameter, discrete or continuous, that characterises the family leaks in as a remnant of the freedom of the ungauged theory to select an electric/magnetic duality frame before the gauging is introduced. It also happens that different such gaugings usually have vacua, of anti-de Sitter (AdS), de Sitter or Minkowski type, that exhibit the same residual supersymmetries $0 \leq \mathcal{N} \leq 8$ and the same residual bosonic symmetry [1, 4–8]. Often (though not always), vacua with the same (super)symmetries have identical mass spectra within their corresponding gauged supergravities. An interesting question therefore arises for $D = 4$ $\mathcal{N} = 8$ gaugings with a higher dimensional origin: for vacua of different such $\mathcal{N} = 8$ gaugings with the same symmetry and the same spectrum within their corresponding $\mathcal{N} = 8$ theories, are the masses still the same up the corresponding Kaluza-Klein (KK) towers?

In this paper, we will address this question by focusing on three $\mathcal{N} = 8$ gaugings with AdS vacua: the purely electric SO(8) gauging [9], the dyonic ISO(7) gauging [10] and the dyonic $(SO(6) \times SO(1, 1)) \ltimes \mathbb{R}^{12}$ gauging as described in [4]. All three gaugings enjoy higher dimensional origins in, respectively, $D = 11$ supergravity [11] on S^7 [12], massive

(Super)symmetry	SO(8)	ISO(7)	$(\text{SO}(6) \times \text{SO}(1, 1)) \times \mathbb{R}^{12}$	same spectrum?
$\mathcal{N} = 8, \text{SO}(8)$	✓	×	×	—
$\mathcal{N} = 2, \text{SU}(3) \times \text{U}(1)$	✓	✓	×	✓
$\mathcal{N} = 1, \text{G}_2$	✓	✓	×	✓
$\mathcal{N} = 1, \text{SU}(3)$	×	✓	✓	✓
$\mathcal{N} = 0, \text{SO}(7)$	✓	✓	×	✓
$\mathcal{N} = 0, \text{SO}(6)$	✓	✓	✓	✓
$\mathcal{N} = 0, \text{G}_2$	×	✓	×	—
$\mathcal{N} = 0, \text{SU}(3)$	×	✓	✓	×

Table 1. Possible residual (super)symmetries, regardless of their $E_{7(7)}$ embedding, of AdS vacua in the $\text{SU}(3)$ -invariant sector of the three different gaugings that we consider.

type IIA supergravity [13] on S^6 [14, 15] and type IIB supergravity [16] on an S-fold geometry [17]. In order to narrow down and simplify the problem, we will further focus on the $\text{SU}(3)$ -invariant sector of these gaugings. This sector has been explicitly uplifted to the respective higher-dimensional supergravities [18–20], and the corresponding vacua have been charted [21, 10, 20]. The list of possible symmetries of AdS vacua in this sector across all three gaugings of interest has been summarised for convenience in table 1. Only the symmetries are indicated in the table, regardless of their actual embedding into the relevant gauge groups and ultimately $E_{7(7)}$. For example, the same table row accounts for the $\text{SU}(3) \times \text{U}(1)_c$ and $\text{SU}(3) \times \text{U}(1)_v$ solutions of $\text{SO}(8)$ and $\text{ISO}(7)$ supergravity, which differ in their embeddings into $E_{7(7)}$. The table also shows whether vacua with the same (super)symmetries in different theories exhibit the same mass spectrum within their corresponding $D = 4$ $\mathcal{N} = 8$ theories. All of them do, except for the solution with $\mathcal{N} = 0$ $\text{SU}(3)$ symmetry. For this reason, we will not be concerned with the latter in this paper.

It was already shown in [22] that the $\mathcal{N} = 2$ $\text{SU}(3) \times \text{U}(1)$ -invariant solutions of the $\text{SO}(8)$ [21] and $\text{ISO}(7)$ gaugings [14] do cease to have the same KK spectrum upon uplift to $D = 11$ [23] and type IIA [14], thus answering in the negative the question posed above. Fortunately, it was not necessary to compute the entire KK spectrum to elucidate that question in [22]. Computing the spectrum of gravitons following [24] was enough to give an answer. In this paper, we will extend this statement to all solutions with at least $\text{SU}(3)$ symmetry. Any pair of such solutions in table 1 with the same (super)symmetry and the same spectrum within their $\mathcal{N} = 8$ supergravities fails to have the same spectrum of KK gravitons. This is so regardless of the embedding of the residual symmetry group into the gauge group, and ultimately into the duality group $E_{7(7)}$ of the ungauged $\mathcal{N} = 8$ theory. For example, both the $\text{SO}(7)_v$ and $\text{SO}(7)_c$ solutions of the $\text{SO}(8)$ gauging have the same spectrum within the $\mathcal{N} = 8$ theory [25]. But, as we will show in this paper, the spectrum of gravitons for both solutions differ.

It was also observed in [22] that, despite their different KK mode structure, certain sums of masses up the Kaluza-Klein tower did still remain universal for the $\mathcal{N} = 2$ $\text{SU}(3) \times \text{U}(1)$ solutions of the $\text{SO}(8)$ and $\text{ISO}(7)$ gaugings. In other words, while the eigenvalue-by-eigenvalue equivalence of both spectra at the lowest KK level was lost up the KK tower,

certain combinations of higher KK modes did still remain universal. In section 2, we will find the same behaviour for solutions of the SO(8) and ISO(7) gauging within the SU(3)-invariant sector. The solutions of the $(\text{SO}(6) \times \text{SO}(1, 1)) \ltimes \mathbb{R}^{12}$ gauging behave similarly but slightly differently. For that reason, we relegate their discussion to the concluding section 5. In [22], the relevant combinations of KK modes were identified as traces of a KK graviton mass matrix. In section 4, we formalise this notion and introduce an SL(8)-covariant KK graviton mass matrix whose form is qualitatively similar to the bosonic mass matrices of $D = 4$ $\mathcal{N} = 8$ gauged supergravity (see [26]). Prior to discussing this, we complete in section 3 the KK graviton spectrum for all known solutions of the $(\text{SO}(6) \times \text{SO}(1, 1)) \ltimes \mathbb{R}^{12}$ gauging by computing the spin-2 spectrum of a solution with $\mathcal{N} = 4$ supersymmetry and SO(4) symmetry [17], which lies outside the SU(3) sector of section 2.

In this paper, we want to compute the spectrum of massive gravitons about the AdS₄ solutions of the ten- and eleven-dimensional supergravities specified above. For later reference, the relevant geometries are of the form

$$d\hat{s}_{d+4}^2 = e^{2A(y)} \left[(\bar{g}_{\mu\nu}(x) + h_{\mu\nu}(x, y)) dx^\mu dx^\nu + d\bar{s}_d^2(y) \right], \quad (1.1)$$

where $d = 7$ in M-theory and $d = 6$ in type II. The Einstein frame is used in the latter case. The external and internal coordinates have been collectively denoted (x, y) . The metrics $\bar{g}_{\mu\nu} dx^\mu dx^\nu \equiv ds^2(\text{AdS}_4)$ and $d\bar{s}_d^2(y)$ respectively denote the unit-radius four-dimensional anti-de Sitter metric, and a background metric on the internal d -dimensional space that will be specified below on a case-by-case basis. The warp factor $e^{2A(y)}$ takes values on the internal space. Finally, $h_{\mu\nu}$ is taken to be a spin-2 perturbation over AdS₄ that also depends on the internal coordinates. More concretely, the perturbation is assumed to take on the factorised form

$$h_{\mu\nu}(x, y) = h_{\mu\nu}^{[tt]}(x) \mathcal{Y}(y), \quad (1.2)$$

with $\mathcal{Y}(y)$ a function on the internal space only, and $h^{[tt]}$ transverse ($\bar{\nabla}^\mu h_{\mu\nu}^{[tt]} = 0$) with respect to the Levi-Civita connection corresponding to $\bar{g}_{\mu\nu}$, traceless ($\bar{g}^{\mu\nu} h_{\mu\nu}^{[tt]} = 0$), and subject to the Fierz-Pauli equation

$$\bar{\square} h_{\mu\nu}^{[tt]} = (M^2 L^2 - 2) h_{\mu\nu}^{[tt]}, \quad (1.3)$$

for a graviton of squared mass M^2 . Here, L is the effective AdS₄ radius (introduced in our context by the warping $e^{2A(y)}$), such that the combination $M^2 L^2$ is dimensionless.

While the computation of supergravity spectra in general is a very complicated problem, the calculation of massive graviton spectra is comparatively much simpler. The reason is that the linearised spin-2 equations decouple from the supergravity fluxes and become a boundary value problem involving only the warp factor and the internal background metric. Indeed, with the above assumptions, the linearised ten- and eleven-dimensional Einstein equations reduce to the eigenvalue problem [24]

$$-\frac{e^{-(d+2)A}}{\sqrt{\bar{g}}} \partial_M (e^{(d+2)A} \sqrt{\bar{g}} \bar{g}^{MN} \partial_N \mathcal{Y}) = M^2 L^2 \mathcal{Y}, \quad (1.4)$$

with \bar{g}^{MN} , $M, N = 1, \dots, d$, the inverse metric and \bar{g} the determinant of the internal metric $d\bar{s}_d^2$ in (1.1). For unwarped geometries, $A = 0$, (1.4) reduces to the spectral problem for the Laplacian on the internal space. Previous calculations of KK graviton spectra in related contexts include [27, 24, 28–34].

2 Massive gravitons with at least SU(3) symmetry

We now compute the KK graviton spectrum about the AdS₄ solutions of $D = 11$ supergravity, massive type IIA supergravity and type IIB supergravity that uplift from critical points of SO(8) supergravity, dyonic ISO(7) supergravity and $(\text{SO}(6) \times \text{SO}(1, 1)) \times \mathbb{R}^{12}$ supergravity with at least SU(3) symmetry. For convenience, some features of the SU(3)-invariant sector of these $D = 4$ $\mathcal{N} = 8$ gaugings are summarised in appendix A.

2.1 M-theory

The class of AdS₄ solutions of $D = 11$ supergravity [11] that we are interested in arises upon consistent uplift on S^7 [12] of critical points of $D = 4$ $\mathcal{N} = 8$ SO(8)-gauged supergravity [9]. For definiteness, we will restrict to critical points that preserve at least the SU(3) subgroup of SO(8). There are six such vacua [21]. The corresponding uplifts are given by the $D = 11$ solutions first found in [35, 23, 36–39]. These solutions are invariant, both in $D = 4$ and in $D = 11$, under a number of subgroups of SO(8) larger than SU(3), and display supersymmetries $\mathcal{N} = 0, 1, 2, 8$. See table 1 for a summary. The entire spectrum about the Freund-Rubin $\mathcal{N} = 8$ SO(8)-invariant AdS₄ solution [35] has long been known [40–42] (see also [43] for a review). The spectrum of gravitons about the $\mathcal{N} = 2$ SU(3) \times U(1)_c-invariant solution [23] is also known [27]. The graviton spectra that we will give below for the four other AdS₄ solutions in this sector are new.

A convenient starting point for our analysis is the local geometries recently presented in [18]. In that reference, the full, dynamical SU(3)-invariant sector of $D = 4$ $\mathcal{N} = 8$ SO(8)-gauged supergravity [9] was uplifted to $D = 11$ using the consistent truncation formulae of [44]. In particular, the results of [18] provide a unified treatment for all the $D = 11$ AdS₄ solutions that uplift from critical points of $D = 4$ $\mathcal{N} = 8$ SO(8)-gauged supergravity with at least SU(3) symmetry. In order to simplify the calculations, we will focus on two disjoint further subsectors with symmetries G₂ and SU(4)_c larger than SU(3). We will obtain the graviton spectra for arbitrary constant values of the $D = 4$ scalars in those sectors. Finding the actual spectra about each individual solution will simply entail an evaluation of those formulae at the corresponding scalar vevs.

2.1.1 Massive gravitons with at least G₂ symmetry

The G₂-invariant sector of the $D = 4$ SO(8) supergravity contains an SL(2, ℝ)/SO(2) dilaton-axion pair (φ, χ) , in the notation of appendix A. The $D = 11$ uplift of this sector was given in section 3.2.3 of [18]. The warp factor and internal $d = 7$ geometry that feature in (1.1) are given by

$$e^{2A} = e^{-\varphi} X^{1/3} \Delta_1^{2/3} L^2, \quad d\bar{s}_7^2 = g^{-2} L^{-2} \left(e^{3\varphi} X^{-3} d\beta^2 + e^\varphi \Delta_1^{-1} \sin^2 \beta ds^2(S^6) \right). \quad (2.1)$$

Here β is an angle on S^7 , with¹ $0 \leq \beta \leq \pi$, and $ds^2(S^6)$ is the round Einstein metric on the unit radius S^6 . The dilaton φ appears explicitly in (2.1) and the axion χ appears through the combinations X defined in equation (A.7) and

$$\Delta_1 = X(e^{2\varphi} \sin^2 \beta + e^{-2\varphi} X^2 \cos^2 \beta). \quad (2.2)$$

Finally, g and L are constants. The former is the gauge coupling of the $D = 4$ supergravity and the latter is related via (A.3) to the G_2 -invariant potential V , given by (A.4) with the identifications (A.8). The geometry (2.1) is in fact invariant under the $SO(7)_v$ that rotates the round S^6 . When $\chi \neq 0$, the symmetry of the full $D = 11$ configuration is broken to G_2 by the supergravity four-form field strength.

For the class of geometries (2.1), the differential equation (1.4) becomes

$$\left[e^{-3\varphi} X^3 (\partial_\beta^2 + 6 \cot \beta \partial_\beta) + e^{-\varphi} \Delta_1 \sin^{-2} \beta \square_{S^6} \right] \mathcal{Y} = -g^{-2} M^2 \mathcal{Y}, \quad (2.3)$$

where \square_{S^6} is the S^6 Laplacian. Using separation of variables,

$$\mathcal{Y} = f \mathcal{Y}_k, \quad (2.4)$$

where $f = f(\beta)$ depends only on β and \mathcal{Y}_k are the S^6 spherical harmonics,

$$\square_{S^6} \mathcal{Y}_k = -k(k+5) \mathcal{Y}_k, \quad (2.5)$$

the PDE (2.3) reduces to an ODE for $f(\beta)$,

$$e^{-3\varphi} X^3 (f''(\beta) + 6 \cot \beta f'(\beta)) - e^{-\varphi} \Delta_1 \sin^{-2} \beta k(k+5) f(\beta) = -g^{-2} M^2 f(\beta), \quad (2.6)$$

where a prime denotes derivative with respect to β . Finally, it is convenient to introduce a further change of variables,

$$u = \cos^2 \beta, \quad f(u) = (1-u)^{\frac{k}{2}} H(u). \quad (2.7)$$

The independent variable u now ranges in $0 \leq u \leq 1$, covering this interval twice given the range of β below (2.1). In the variables (2.7), the differential equation (2.6) takes on the standard hypergeometric form

$$u(1-u)H'' + (c - (1+a_+ + a_-)u)H' - a_+a_-H = 0, \quad (2.8)$$

with

$$a_\pm = \frac{1}{2}(k+3) \mp \frac{1}{2} e^{3\varphi/2} X^{-3/2} \sqrt{g^{-2} M^2 + 9e^{-3\varphi} X^3 + k(k+5)(e^{-3\varphi} X^3 - e^\varphi X)}, \quad c = \frac{1}{2}. \quad (2.9)$$

The two linearly independent solutions to (2.8) are given by the hypergeometric functions

$${}_2F_1(a_+, a_-, c; u) \quad \text{and} \quad u^{1-c} {}_2F_1(1+a_+ - c, 1+a_- - c, 2-c; u). \quad (2.10)$$

¹This range of β corrects a typo below (B.22) of [18].

Both solutions are regular at $u = 0$ for all values of the parameters (2.9). At $u = 1$, however, regularity imposes restrictions on the parameters. Regularity of the first solution in (2.10) demands $a_+ = -j$ with j a non-negative integer. Bringing this condition to (2.9), we find a first tower of KK graviton squared masses:

$$g^{-2}M_{(1)j,k}^2 = e^{-3\varphi}X^3(2j+k)(2j+k+6) + e^{-\varphi}X(e^{2\varphi} - e^{-2\varphi}X^2)k(k+5). \quad (2.11)$$

The corresponding eigenfunctions are given by (2.4), (2.7), with $H(u)$ given by the first choice in (2.10), namely

$$\mathcal{Y}_{(1)j,k} = \mathcal{Y}_k \sin^k \beta \sum_{s=0}^j (-1)^s \binom{j}{s} \frac{(j+k+3)_s}{(\frac{1}{2})_s} \cos^{2s} \beta \quad (2.12)$$

(no sum in k), where

$$(x)_s = \begin{cases} 1 & , \text{ if } s = 0 \\ x(x+1) \cdots (x+s-1) & , \text{ if } s > 0 \end{cases} \quad (2.13)$$

is the Pochhammer symbol. Regularity of the second solution in (2.10) at $u = 1$ in turn requires $1 + a_+ - c = -j$, with j again a non-negative integer. Bringing this condition to (2.9), we find a second tower of KK graviton squared masses:

$$g^{-2}M_{(2)j,k}^2 = e^{-3\varphi}X^3(2j+1+k)(2j+1+k+6) + e^{-\varphi}X(e^{2\varphi} - e^{-2\varphi}X^2)k(k+5). \quad (2.14)$$

The associated eigenfunctions are now given by (2.4), (2.7), with $H(u)$ given by the second choice in (2.10):

$$\mathcal{Y}_{(2)j,k} = \mathcal{Y}_k \sin^k \beta \sum_{s=0}^j (-1)^s \binom{j}{s} \frac{(j+k+4)_s}{(\frac{3}{2})_s} \cos^{2s+1} \beta. \quad (2.15)$$

The eigenvalues (2.11) and (2.14) actually correspond to a unique tower of KK graviton masses. This is made apparent by introducing a new quantum number n defined as

$$n = \begin{cases} 2j+k & , \text{ for the first branch} \\ 2j+1+k & , \text{ for the second branch.} \end{cases} \quad (2.16)$$

In terms of (n, k) , (2.11) and (2.14) can be combined into the single KK tower:

$$g^{-2}M_{n,k}^2 = e^{-3\varphi}X^3n(n+6) + e^{-\varphi}X(e^{2\varphi} - e^{-2\varphi}X^2)k(k+5), \quad (2.17)$$

which is our final result. The quantum numbers range here as

$$n = 0, 1, 2, \dots, \quad k = 0, 1, \dots, n. \quad (2.18)$$

Only n ranges freely over the non-negative integers, due to its definition (2.16) in terms of the non-negative but otherwise unconstrained integer j . The range of k is limited to $k \leq n$ by (2.16). At fixed n , the eigenvalue (2.17) occurs with degeneracy

$$D_{k,7} \equiv \dim [k, 0, 0]_{\text{SO}(7)}, \quad (2.19)$$

where, more generally, $D_{k,N}$ is the dimension of the symmetric traceless representation $[k, 0, \dots, 0]$ of $\text{SO}(N)$,

$$\begin{aligned} D_{k,N} &= \binom{k+N-1}{k} - \binom{k+N-3}{k-2} \\ &= \frac{1}{(N-2)!} (2k+N-2)(k+N-3)(k+N-4) \cdots (k+2)(k+1), \end{aligned} \quad (2.20)$$

for $k \geq 2$ and

$$D_{0,N} = 1, \quad D_{1,N} = N, \quad \text{for all } N = 2, 3 \dots \quad (2.21)$$

It is also useful to note that

$$D_{n,N-1} = D_{n,N} - D_{n-1,N}, \quad \text{for all } n = 1, 2, \dots \text{ and all } N = 2, 3 \dots \quad (2.22)$$

The eigenfunctions (2.12), (2.15) can be similarly combined into

$$\mathcal{Y}_{n,k} = \mathcal{Y}_k \sin^k \beta \sum_{s=0}^{\lfloor \frac{n-k}{2} \rfloor} (-1)^s \binom{\lfloor \frac{n-k}{2} \rfloor}{s} \frac{(\lfloor \frac{n-k}{2} \rfloor + k + 3 + h_{n,k})_s}{(\frac{1}{2} + h_{n,k})_s} \cos^{2s+h_{n,k}} \beta, \quad (2.23)$$

where $\lfloor \cdot \rfloor$ means integer part and we define the symbol $h_{n,k}$ as

$$h_{n,k} = n - k - 2 \left\lfloor \frac{n-k}{2} \right\rfloor = \begin{cases} 0, & n-k \text{ even (for the first branch)} \\ 1, & n-k \text{ odd (for the second branch)}. \end{cases} \quad (2.24)$$

At fixed n and k , the eigenfunctions (2.23) span the $[k, 0, 0]$ representation of $\text{SO}(7)_v$. Moreover, it can be checked that these eigenfunctions at fixed n actually span the full symmetric traceless representation $[n, 0, 0, 0]$ of $\text{SO}(8)$. In other words, the eigenfunctions (2.23) turn out to be simply the $\text{SO}(8)$ spherical harmonics of S^7 , branched out into $\text{SO}(7)_v$ representations through

$$[n, 0, 0, 0] \xrightarrow{\text{SO}(7)_v} \sum_{k=0}^n [k, 0, 0]. \quad (2.25)$$

This is consistent with the quantum number ranges (2.18). This is also compatible with the internal geometry (2.1) being topologically S^7 : it can be continuously deformed into the round $\text{SO}(8)$ -invariant geometry by setting $\varphi = \chi = 0$. These arguments suggest that the spectrum (2.17), (2.23) is in fact complete. Thus, the quantum number n can be regarded as the Kaluza-Klein level, as it coincides with the unique integer that characterises the KK spectrum of the $\mathcal{N} = 8$ $\text{SO}(8)$ -invariant Freund-Rubin solution: see e.g. table 9 of [43].

2.1.2 Massive gravitons with at least $\text{SU}(4)_c$ symmetry

The $\text{SU}(4)_c$ -invariant sector of $\text{SO}(8)$ -gauged supergravity contains three pseudoscalars: $\chi, \zeta, \tilde{\zeta}$ in the notation of appendix A. In the Iwasawa parametrisation of the appendix, the $\text{SU}(3)$ -invariant dilatons φ, ϕ become identified in terms of the pseudoscalars via equation (A.9). With the understanding that φ, ϕ depend on the independent fields χ, ζ ,

$\tilde{\zeta}$, the former can be conveniently used to parametrise the $SU(4)_c$ -invariant sector, as the resulting expressions are more compact. The embedding of this sector into the $D = 11$ warp factor and internal metric reads [25], in the notation of [18],

$$e^{2A} = e^{\frac{4}{3}\phi+\varphi} L^2, \quad d\bar{s}_7^2 = g^{-2} L^{-2} \left[e^{-2\phi-\varphi} ds^2(\mathbb{CP}^3) + e^{-3\varphi} (d\psi + \sigma)^2 \right]. \quad (2.26)$$

Here, $ds^2(\mathbb{CP}^3)$ is the Fubini-Study metric on the complex projective space, σ a one-form potential for the Kähler form on the latter, and $0 \leq \psi \leq 2\pi$ a coordinate on the Hopf fibre of S^7 . The constant g is again the coupling of $SO(8)$ supergravity and L is fixed through (A.3) in terms of the $SU(4)_c$ -invariant potential V , given by (A.4) with the identifications (A.9). Away from the $SO(7)_c$ -invariant locus, (2.39) of [18], where the symmetry is enhanced accordingly, the geometry (2.26) is invariant under $SU(4)_c \times U(1)$, with $U(1)$ generated by ∂_ψ . This $U(1)$ is broken by the $D = 11$ supergravity four-form.

The $D = 11$ embedding (2.26) of the $SU(4)_c$ -invariant sector is homogeneous: the warp factor depends only on the $D = 4$ scalars and not on the S^7 coordinates, and the metric $d\bar{s}_7^2$ corresponds to a homogeneous stretching of the S^7 geometry along its Hopf fibre. Therefore, the differential equation (1.4) simplifies for this geometry as

$$\left[(e^{3\varphi} - e^{2\phi+\varphi}) \partial_\psi^2 + e^{2\phi+\varphi} \square_{S^7} \right] \mathcal{Y} = -g^{-2} M^2 \mathcal{Y}, \quad (2.27)$$

with \square_{S^7} the Laplacian on the round, Einstein metric on S^7 . The solutions of (2.27) are accordingly given by the $SO(8)$ spherical harmonics on S^7 , branched out into representations of the $SU(4)_c \times U(1)$ symmetry group of (2.26) and (2.27) via

$$[n, 0, 0, 0] \xrightarrow{SU(4)_c \times U(1)} \sum_{r=0}^n [r, 0, n-r]_{2r-n}, \quad (2.28)$$

with the subindex indicating the $U(1)$ charge. More concretely, the S^7 spherical harmonics, in the $[n, 0, 0, 0]$ of $SO(8)$, split according to (2.28) as

$$\mathcal{Y}_{n,r}(z, \bar{z}) = c_{a_1 \dots a_r}{}^{b_1 \dots b_{n-r}} z^{a_1} \dots z^{a_r} \bar{z}_{b_1} \dots \bar{z}_{b_{n-r}}, \quad (2.29)$$

for

$$n = 0, 1, 2, \dots, \quad r = 0, 1, \dots, n. \quad (2.30)$$

In (2.29), $z^1 = \mu^1 + i\mu^2$, etc, are complexified embedding coordinates of \mathbb{R}^8 constrained as $\delta_{AB} \mu^A \mu^B = 1$, with $A, B = 1, \dots, 8$, and $c_{a_1 \dots a_r}{}^{b_1 \dots b_{n-r}}$ is a constant tensor in the $[n-r, 0, r]$ of $SU(4)$. The functions (2.29) obey

$$\square_{S^7} \mathcal{Y}_{n,r} = -n(n+6) \mathcal{Y}_{n,r}, \quad \partial_\psi^2 \mathcal{Y}_{n,r} = -(n-2r)^2 \mathcal{Y}_{n,r}, \quad (2.31)$$

and thus satisfy the differential equation (2.27) with eigenvalue

$$g^{-2} M_{n,r}^2 = e^{2\phi+\varphi} n(n+6) + (e^{3\varphi} - e^{2\phi+\varphi}) (n-2r)^2. \quad (2.32)$$

This occurs with multiplicity

$$d_{n,r} = \dim[r, 0, n-r]_{\text{SU}(4)} = \frac{1}{12}(n+3)(r+1)(r+2)(n-r+1)(n-r+2). \quad (2.33)$$

To summarise, the complete spectrum of the eigenvalue equation (2.27) is (2.32), (2.29), with the quantum numbers ranging as in (2.30). The eigenvalues (2.32) have multiplicity (2.33) and the eigenfunctions (2.29) are simply the S^7 spherical harmonics split into $\text{SU}(4)_c \times \text{U}(1)$ representations through (2.28). The eigenvalues have been given in terms of $D = 4$ scalars. The massive KK graviton spectra about $D = 11$ AdS₄ solutions in this sector are obtained by fixing the $D = 4$ scalars to the corresponding vevs. Like in the case discussed in section 2.1.1, the integer n is identified with the KK level by an argument similar to that put forward below (2.25).

2.2 Type IIB

We now move on to compute the graviton spectrum about the AdS₄ solutions of type IIB supergravity recently obtained in [20]. These geometries arise upon consistent uplift [17] on an S-fold geometry of AdS₄ vacua of $D = 4$ $\mathcal{N} = 8$ gauged supergravity with dyonic $[\text{SO}(6) \times \text{SO}(1,1)] \times \mathbb{R}^{12}$ gauging [4, 2] (see also [45]). The resulting type IIB uplifts correspond to limiting Janus-type solutions [46–49]. As in section 2.1, we will focus in this section on solutions that preserve at least $\text{SU}(3)$ symmetry. These were classified in [20]. We will compute the generic graviton spectra for arbitrary constant values of the $\text{SU}(3)$ -invariant scalars of the $D = 4$ supergravity.

The type IIB geometries under consideration are of the form (1.1) with $d = 6$ and² [20]

$$e^{2A} = \sqrt{Y} e^\varphi L^2, \quad ds_6^2 = \frac{e^{-\varphi}}{\sqrt{Y}} g^{-2} L^{-2} \left[\sqrt{Y} e^{-2\varphi} d\eta^2 + \frac{1}{\sqrt{Y}} (ds^2(\mathbb{CP}^2) + Y(d\tau + \sigma')^2) \right]. \quad (2.34)$$

The geometry inside the last parenthesis extends globally over a topological S^5 , with $ds^2(\mathbb{CP}^2)$ the Fubini-Study metric on the complex projective plane within S^5 and $0 \leq \tau < 2\pi$ the Hopf fibre angle. The local one-form σ' is a potential for the Kähler form on \mathbb{CP}^2 . The sixth internal coordinate η will be taken to be periodic, $\eta \sim \eta + T$, with T a positive number. The ten-dimensional geometry (2.34) also depends on the $\text{SU}(3)$ -invariant scalars of appendix A both explicitly and through the combination Y defined in (A.7). Finally, g is the gauge coupling constant of the $D = 4$ supergravity, and L is fixed through (A.3) in terms of the scalar potential V given by (A.6) with $\chi = 0$. For general values of the scalars, the geometry (2.34) displays an isometry group $\text{SU}(3) \times \text{U}(1)_\tau \times \text{U}(1)_\eta$, with the $\text{U}(1)_\eta$ factor broken by the type IIB fluxes. In particular, the type IIB fields charged under the S-duality group $\text{SL}(2, \mathbb{R})$ undergo a monodromy transformation as η crosses through different periods [17]. The type IIB metric is neutral under S-duality and thus insensitive to this transformation.

Like in section 2.1.2, the type IIB embedding (2.34) is homogeneous. Accordingly, the differential equation (1.4) reduces for this geometry to

$$\left[e^{3\varphi} \partial_\eta^2 + e^\varphi (1 - Y) \partial_\tau^2 + e^\varphi Y \square_{S^5} \right] \mathcal{Y} = -g^{-2} M^2 \mathcal{Y}, \quad (2.35)$$

²We have conveniently rescaled the metric and warp factor with respect to [20].

where \square_{S^5} is the Laplacian on the round, Einstein metric on S^5 . The complete set of eigenfunctions $\mathcal{Y} \equiv \mathcal{Y}_{\ell,p,j}$ that solve (2.35) can be taken to satisfy

$$\square_{S^5} \mathcal{Y}_{\ell,p,j} = -\ell(\ell+4)\mathcal{Y}_{\ell,p,j}, \quad \partial_\tau^2 \mathcal{Y}_{\ell,p,j} = -(\ell-2p)^2 \mathcal{Y}_{\ell,p,j}, \quad \partial_\eta^2 \mathcal{Y}_{\ell,p,j} = -\left(\frac{2\pi}{T}j\right)^2 \mathcal{Y}_{\ell,p,j}, \quad (2.36)$$

for

$$\ell = 0, 1, 2, \dots, \quad p = 0, 1, \dots, \ell, \quad j = 0, \pm 1, \pm 2, \dots, \quad (2.37)$$

with ℓ and j unconstrained and p constrained by ℓ through $p \leq \ell$. In other words, the eigenfunctions $\mathcal{Y}_{\ell,p,j}$ come in representations of $SU(3) \times U(1)_\tau \times U(1)_\eta$, and are explicitly given by products of harmonics on the S^1 generated by ∂_η and spherical harmonics $[\ell, 0, 0]_{SO(6)}$ on S^5 branched out into representations of $SU(3) \times U(1)_\tau$ via

$$[\ell, 0, 0] \xrightarrow{SU(3) \times U(1)_\tau} \sum_{p=0}^{\ell} [p, \ell-p]_{\ell-2p}. \quad (2.38)$$

Bringing (2.36) to (2.35), we find the eigenvalues

$$g^{-2} M_{\ell,p,j}^2 = e^\varphi Y \ell(\ell+4) + e^\varphi (1-Y)(\ell-2p)^2 + e^{3\varphi} \left(\frac{2\pi}{T}j\right)^2, \quad (2.39)$$

occurring with degeneracy

$$d_{\ell,p,j} = \begin{cases} \dim[p, \ell-p]_{SU(3)} &= \frac{1}{2}(p+1)(\ell-p+1)(\ell+2), & \text{if } j=0 \\ 2 \dim[p, \ell-p]_{SU(3)} &= (p+1)(\ell-p+1)(\ell+2), & \text{if } j \neq 0. \end{cases} \quad (2.40)$$

In summary, the complete eigenvalue spectrum of equation (2.35) is (2.39) with the eigenfunctions $\mathcal{Y}_{\ell,p,j}$ described above and with the quantum numbers ranging as in (2.37). The eigenvalues (2.39) have multiplicity (2.40), and have been given in terms of $D=4$ scalars. The massive KK graviton spectra about $D=11$ AdS₄ solutions in this sector are obtained by fixing the $D=4$ scalars to the corresponding vevs, as we will see next.

2.3 Individual spectra in M-theory, type IIA and type IIB

Using the results of sections 2.1 and 2.2 as well as [43, 27, 22], we can write down the KK graviton spectra about the AdS₄ solutions of the ten and eleven-dimensional supergravities that uplift from critical points with at least $SU(3)$ symmetry of the three $D=4$ $\mathcal{N}=8$ gauged supergravities that we are considering in this paper.

In M-theory, the spectrum above the AdS₄ solutions with at least G_2 symmetry and at least $SU(4)_c$ symmetry can be obtained by particularising (2.17) and (2.32), respectively, to the scalar vevs given in [18]. We have brought these results to table 2. In order to exhaust the KK graviton spectra of AdS₄ solutions of $D=11$ supergravity that uplift from critical points of $D=4$ $\mathcal{N}=8$ $SO(8)$ -gauged supergravity with at least $SU(3)$ symmetry, the table also includes the spectrum [43] about the $\mathcal{N}=8$ Freund-Rubin solution [35] and the

Solution	Mass	Degeneracy
$\mathcal{N} = 8, \text{SO}(8)$	$L^2 M_n^2 = \frac{1}{4}n(n+6)$	$d_n = D_{n,8}$
$\mathcal{N} = 2, \text{SU}(3) \times \text{U}(1)_c$	$L^2 M_{n,p,\ell,r}^2 = \frac{1}{2}n(n+6) - \frac{1}{3}\ell(\ell+4) - \frac{1}{9}(\ell-2p)^2 + \frac{1}{18}[3(n-2r) + 4(\ell-2p)]^2$	$d_{n,p,\ell,r} = \frac{1}{2}(p+1)(\ell-p+1)(\ell+2)$
$\mathcal{N} = 1, \text{G}_2$	$L^2 M_{n,k}^2 = \frac{5}{8}n(n+6) - \frac{5}{12}k(k+5)$	$d_{n,k} = D_{k,7}$
$\mathcal{N} = 0, \text{SO}(7)_v$	$L^2 M_{n,k}^2 = \frac{3}{4}n(n+6) - \frac{3}{5}k(k+5)$	$d_{n,k} = D_{k,7}$
$\mathcal{N} = 0, \text{SO}(7)_c$	$L^2 M_n^2 = \frac{3}{10}n(n+6)$	$d_n = D_{n,8}$
$\mathcal{N} = 0, \text{SU}(4)_c$	$L^2 M_{n,r}^2 = \frac{3}{8}n(n+6) - \frac{3}{16}(n-2r)^2$	$d_{n,r} = \frac{1}{12}(n+3)(r+1)(r+2)(n-r+1)(n-r+2)$

Table 2. The KK graviton spectra of AdS₄ solutions of $D = 11$ supergravity that uplift from critical points of $D = 4$ $\mathcal{N} = 8$ SO(8)-gauged supergravity with at least SU(3) symmetry. See (2.20) for the notation $D_{k,N}$. The quantum numbers range as in (2.41).

spectrum [27] about the $\text{SU}(3) \times \text{U}(1)_c$ -invariant AdS₄ solution [21, 23]. The latter is given as in [22], with $n_{\text{here}} = n_{\text{there}}$, $r_{\text{here}} = r_{\text{there}}$, $p_{\text{here}} = p_{\text{there}}$ and $\ell_{\text{here}} = p_{\text{there}} + q_{\text{there}}$. The corresponding multiplicities are also given in the table, and the quantum numbers range as

$$n = 0, 1, 2, \dots, \quad r, k = 0, 1, \dots, n, \quad \ell = p, \dots, p+r, \quad p = 0, 1, \dots, n-r. \quad (2.41)$$

The only quantum number that is free to range unrestricted over the non-negative integers is n , all the others being bound by it. This is consistent with the interpretation of n as the SO(8) KK level, see below (2.25). At fixed KK level n , the degeneracy of the $\mathcal{N} = 8$ SO(8)-symmetric spectrum is broken into representations of the isometry group of the internal metric. This may be larger than the symmetry of each solution, as the fluxes will further break the isometry to the actual symmetry quoted in the table. Similarly, the eigenfunctions corresponding to each solution are simply the S^7 spherical harmonics branched out into the representations of the relevant group.

For convenience, table 3 imports from [22] the KK graviton spectra of AdS₄ solutions of massive IIA supergravity that uplift from critical points of $D = 4$ $\mathcal{N} = 8$ dyonic ISO(7)-gauged supergravity with at least SU(3) symmetry. The table includes the squared masses in units of the corresponding AdS radius L , as well as the multiplicities. In this case, the quantum numbers' ranges are

$$k = 0, 1, 2, \dots, \quad \ell = 0, 1, \dots, k, \quad p = 0, 1, \dots, \ell, \quad (2.42)$$

with $k_{\text{here}} = n_{\text{in}}$ [22]. Again, k is the only quantum number that is unrestricted. For this reason, k can be interpreted in this case as the SO(7) KK level. The eigenfunctions are now the S^6 spherical harmonics split into representations of the internal isometry group. This again may be larger than the symmetry of each solution given in table 3 because the fluxes may further break the isometry to the actual symmetry of the metric and fluxes.

Finally, we turn to the spectrum of gravitons corresponding to the type IIB AdS₄ S-fold solutions that uplift from critical points with at least SU(3) symmetry [20] of $D = 4$ $\mathcal{N} = 8$ supergravity with $(\text{SO}(6) \times \text{SO}(1, 1)) \ltimes \mathbb{R}^{12}$ gauging. These are found by bringing the corresponding vevs, collected in our conventions in table 6 in appendix A, to equation (2.39).

Solution	Mass	Degeneracy
$\mathcal{N} = 2, \text{SU}(3) \times \text{U}(1)_v$	$L^2 M_{k,\ell,p}^2 = \frac{2}{3}k(k+5) - \frac{1}{3}\ell(\ell+4) + \frac{1}{9}(\ell-2p)^2$	$d_{k,\ell,p} = \frac{1}{2}(p+1)(\ell-p+1)(\ell+2)$
$\mathcal{N} = 1, \text{G}_2$	$L^2 M_k^2 = \frac{5}{12}k(k+5)$	$d_k = D_{k,7}$
$\mathcal{N} = 1, \text{SU}(3)$	$L^2 M_{k,\ell,p}^2 = \frac{5}{6}k(k+5) - \frac{5}{12}\ell(\ell+4) - \frac{5}{36}(\ell-2p)^2$	$d_{k,\ell,p} = \frac{1}{2}(p+1)(\ell-p+1)(\ell+2)$
$\mathcal{N} = 0, \text{SO}(7)_v$	$L^2 M_k^2 = \frac{2}{3}k(k+5)$	$d_k = D_{k,7}$
$\mathcal{N} = 0, \text{SO}(6)_v$	$L^2 M_{k,\ell}^2 = k(k+5) - \frac{3}{4}\ell(\ell+4)$	$d_\ell = D_{\ell,6}$
$\mathcal{N} = 0, \text{G}_2$	$L^2 M_k^2 = \frac{1}{2}k(k+5)$	$d_k = D_{k,7}$

Table 3. The KK graviton spectra of AdS₄ solutions of massive IIA supergravity that uplift from critical points of $D = 4$ $\mathcal{N} = 8$ dyonic ISO(7)-gauged supergravity with at least SU(3) symmetry, taken from [22]. See (2.20) for the notation $D_{k,N}$. The quantum numbers range as in (2.42).

The results are summarised in table 4. In order to derive the generic scalar-dependent spectra in section 2.2, we assumed that the S-fold direction η is compactified to a $\text{U}(1)_\eta$ with period T . The KK graviton spectra are sensitive to this period. The eigenfunctions are products of S^5 harmonics, possibly branched out into $\text{SU}(3) \times \text{U}(1)_\tau$ representations, and $\text{U}(1)_\eta$ harmonics. This $\text{U}(1)_\eta$ is broken by the IIB fluxes.

2.4 Universality of traces

When regarded as vacua of their corresponding $D = 4$ $\mathcal{N} = 8$ gauged supergravities, the AdS solutions under consideration with at least SU(3) symmetry tend to exhibit the same mass spectrum of scalars, vectors and fermions within their $D = 4$ supergravities. This is the case for all these solutions, except for the two $\mathcal{N} = 0, \text{SU}(3)$ -invariant critical points of ISO(7) supergravity and the $\mathcal{N} = 0, \text{SU}(3)$ -invariant critical locus of $(\text{SO}(6) \times \text{SO}(1, 1)) \times \mathbb{R}^{12}$ supergravity. The question that we would like to address in this section is whether this situation persists for higher KK modes. The spectrum of gravitons computed for these solutions in section 2.3 shows that this universality is indeed lost at higher KK levels: the KK gravitons do have completely different masses for all the solutions considered.

However, as we will now show, universality is still maintained, though in a milder form that is not apparent from the results of section 2.3. It turns out that certain sums of KK graviton masses weighted with their multiplicities do remain universal. This is the case at least for solutions in the same or different $\mathcal{N} = 8$ gaugings with the same symmetry and whose spectra within the $D = 4$ supergravity are the same. Specifically, if two AdS₄ solutions of $D = 11$ supergravity or massive IIA uplift from critical points with the same supersymmetry $\mathcal{N} \leq 8$, the same symmetry $G \supset \text{SU}(3)$ (possibly embedded differently into the gauge group) and the same spectrum within the $D = 4$ $\mathcal{N} = 8$ SO(8) or ISO(7) supergravities, then there exist infinitely many discrete combinations $L^2 \text{tr} M_{(n)}^2$, $n = 1, 2, 3, \dots$, of graviton masses weighted with their multiplicities that are the same for both solutions. This statement was proven for the $\mathcal{N} = 2$ $\text{SU}(3) \times \text{U}(1)$ -invariant solutions in [22]. Here we will extend that result to all other solutions with at least SU(3) symmetry in the SO(8) and ISO(7) gaugings, summarised in table 1 of the introduction. As discussed

Solution	Mass	Degeneracy
$\mathcal{N} = 1, \text{SU}(3)$	$L^2 M_{\ell,p,j}^2 = \frac{5}{6}\ell(\ell+4) - \frac{5}{36}(\ell-2p)^2 + \frac{5\pi^2}{T^2}j^2$	$d_{\ell,p,j}$
$\mathcal{N} = 0, \text{SO}(6)_v$	$L^2 M_{\ell,j}^2 = \frac{3}{4}\ell(\ell+4) + \frac{6\pi^2}{T^2}j^2$	$d_{\ell,j} = (2 - \delta_{j0}) D_{\ell,6}$
$\mathcal{N} = 0, \text{SU}(3)$	$L^2 M_{\ell,j}^2 = \frac{3}{4}\ell(\ell+4) + \frac{6\pi^2}{T^2}j^2$	$d_{\ell,j} = (2 - \delta_{j0}) D_{\ell,6}$

Table 4. The KK graviton spectra of AdS₄ S-fold solutions of type IIB supergravity that uplift from critical points of $D = 4$ $\mathcal{N} = 8$ $(\text{SO}(6) \times \text{SO}(1,1)) \times \mathbb{R}^{12}$ -gauged supergravity with at least SU(3) symmetry. See (2.20) for the notation $D_{k,N}$ and (2.40) for $d_{\ell,p,j}$. The quantum numbers range as in (2.37).

in [22] and further in section 4 below, the notation $L^2 \text{tr} M_{(n)}^2$ relates to the fact that the combinations in question correspond to traces of the (infinite-dimensional) KK graviton mass matrix at fixed KK level n .

More concretely, for the M-theory solutions we define $L^2 \text{tr} M_{(n)}^2$ to be the sum of the squared masses in units of the corresponding AdS radius L , weighted with the corresponding multiplicity as given in table 2. The sum is taken at fixed KK level n and over all other quantum numbers ranging as in (2.41). For example, using this prescription, one obtains for the $\mathcal{N} = 8$ SO(8) solution [22],

$$L^2 \text{tr} M_{(n)}^2 = L^2 M_n^2 d_n = 14 D_{n-1,10} . \quad (2.43)$$

In the last step, we have made use of the definition (2.20) as a shorthand for the resulting 8th degree polynomial in n . Similarly, for the $\mathcal{N} = 2$ SU(3) \times U(1)_c solution, we have [22]

$$L^2 \text{tr} M_{(n)}^2 = L^2 \sum_{r=0}^n \sum_{p=0}^{n-r} \sum_{\ell=p}^{p+r} M_{n,p,\ell,r}^2 d_{n,p,\ell,r} = \frac{56}{3} D_{n-1,10} . \quad (2.44)$$

Proceeding similarly, we compute the quantities $L^2 \text{tr} M_{(n)}^2$, $n = 1, 2, \dots$, for the KK graviton spectra summarised in table 2 for $D = 11$ AdS₄ solutions that uplift from critical points of $D = 4$ $\mathcal{N} = 8$ SO(8) supergravity with at least SU(3) symmetry. We obtain:

$$\begin{aligned}
 \mathcal{N} = 8, \text{SO}(8) & : L^2 \text{tr} M_{(n)}^2 = 14 D_{n-1,10} , \\
 \mathcal{N} = 2, \text{SU}(3) \times \text{U}(1)_c & : L^2 \text{tr} M_{(n)}^2 = \frac{56}{3} D_{n-1,10} , \\
 \mathcal{N} = 1, \text{G}_2 & : L^2 \text{tr} M_{(n)}^2 = \frac{35}{2} D_{n-1,10} , \\
 \mathcal{N} = 0, \text{SO}(7)_v & : L^2 \text{tr} M_{(n)}^2 = \frac{84}{5} D_{n-1,10} , \\
 \mathcal{N} = 0, \text{SO}(7)_c & : L^2 \text{tr} M_{(n)}^2 = \frac{84}{5} D_{n-1,10} , \\
 \mathcal{N} = 0, \text{SU}(4)_c & : L^2 \text{tr} M_{(n)}^2 = \frac{39}{2} D_{n-1,10} .
 \end{aligned} \quad (2.45)$$

In particular, the two $SO(7)$ -invariant solutions have their residual symmetry embedded differently into the $SO(8)$ gauge group as $SO(7)_v$ and $SO(7)_c$. They have the same mass spectrum within $D = 4$ $\mathcal{N} = 8$ $SO(8)$ supergravity, according to table 1. Their KK graviton spectra are different, though, according to table 2. But as can be seen from equation (2.45), the quantity $L^2 \text{tr} M_{(n)}^2$ is the same for both solutions for all n .

The quantities $L^2 \text{tr} M_{(k)}^2$ for the KK gravitons of massive IIA solutions with at least $SU(3)$ symmetry that uplift from critical points of dyonic $ISO(7)$ supergravity were computed similarly, for $k = 1, 2, \dots$, in [22]:

$$\begin{aligned}
 \mathcal{N} = 2, \text{SU}(3) \times \text{U}(1)_v & : L^2 \text{tr} M_{(k)}^2 = \frac{56}{3} D_{k-1,9} \quad , \\
 \mathcal{N} = 1, \text{G}_2 & : L^2 \text{tr} M_{(k)}^2 = \frac{35}{2} D_{k-1,9} \quad , \\
 \mathcal{N} = 1, \text{SU}(3) & : L^2 \text{tr} M_{(k)}^2 = \frac{65}{3} D_{k-1,9} \quad , \\
 \mathcal{N} = 0, \text{SO}(7)_v & : L^2 \text{tr} M_{(k)}^2 = \frac{84}{5} D_{k-1,9} \quad , \\
 \mathcal{N} = 0, \text{SO}(6)_v & : L^2 \text{tr} M_{(k)}^2 = \frac{39}{2} D_{k-1,9} \quad , \\
 \mathcal{N} = 0, \text{G}_2 & : L^2 \text{tr} M_{(k)}^2 = 21 D_{k-1,9} \quad .
 \end{aligned}
 \tag{2.46}$$

Here, we have again made use of the notation $D_{k,N}$ defined in (2.20) as a shorthand for the degree-7 polynomial in k that appears in the r.h.s.'s. Now, recall from section 2.3 that k and n can respectively be regarded as the KK levels in massive IIA and $D = 11$. At first KK level, the quantities $L^2 \text{tr} M_{(n=1)}^2$ in (2.45) and $L^2 \text{tr} M_{(k=1)}^2$ in (2.46) can be checked to match, by virtue of the first relation in (2.21), for solutions with the same symmetry group regardless of the embedding of the latter within the corresponding gauge group. For example, for the $D = 11$ $SU(3) \times U(1)_c$ solution [23] and the massive IIA $SU(3) \times U(1)_v$ solution [14], $[L^2 \text{tr} M_{(1)}^2]_{11D} = [L^2 \text{tr} M_{(1)}^2]_{\text{IIA}} = \frac{56}{3}$, at $n = k = 1$, as already noted in [22]. Inspection of (2.45) and (2.46) confirms that similar matches occur at KK level one, $n = k = 1$, for the $D = 11$ and massive IIA solutions with common (super)symmetry $\mathcal{N} = 1, \text{G}_2$, and $\mathcal{N} = 0, \text{SO}(7)$, and $\mathcal{N} = 0, \text{SU}(4) \sim \text{SO}(6)$.

Further, there is still matching at higher KK levels $n > 1$ in $D = 11$ and $k > 1$ in massive IIA, provided a prescription is adopted to relate n and k . An argument will be given in section 4 but, for now, these two quantum numbers can be thought of as being related as in (2.25), so that the $D = 11$ KK level n formally contains all IIA KK levels $k = 0, 1, \dots, n$. Using this prescription, it follows from (2.45) and (2.46) that

$$\sum_{k=0}^n [L^2 \text{tr} M_{(k)}^2]_{\text{IIA}} = [L^2 \text{tr} M_{(n)}^2]_{11D} \quad , \quad n = 0, 1, 2, \dots \quad , \tag{2.47}$$

for all the solutions that we are considering with the same symmetry and supersymmetry in massive IIA and $D = 11$. Here, $L^2 \text{tr} M_{(0)}^2 \equiv 0$ corresponds to the massless graviton, for both the $D = 11$ and type IIA cases, as well as for the IIB cases below. From (2.47) it

immediately follows that

$$\sum_{n=0}^m \sum_{k=0}^n [L^2 \text{tr} M_{(k)}^2]_{\text{IIA}} = \sum_{n=0}^m [L^2 \text{tr} M_{(n)}^2]_{11D}, \quad m = 0, 1, 2, \dots, \quad (2.48)$$

again for all solutions with the same (super)symmetry. The sums in (2.48) obviously run over repeated number of states, both in IIA and in $D = 11$. In (2.47), there are no repeated $D = 11$ states on the r.h.s., but the sum in the l.h.s. does run as well over repeated states in IIA. These overcounting issues can be avoided by subtracting two adjacent KK levels in $D = 11$: formally, the difference between KK levels n and $n - 1$ in $D = 11$ contains the same number of states as KK level $k = n$ in massive IIA. Using the identity (2.22), it follows from (2.45), (2.46) that

$$[L^2 \text{tr} M_{(n)}^2]_{\text{IIA}} = [L^2 \text{tr} M_{(n)}^2]_{11D} - [L^2 \text{tr} M_{(n-1)}^2]_{11D}, \quad n = 1, 2, \dots, \quad (2.49)$$

for solutions with the same (super)symmetry. This relation was already shown to hold in [22] for the $\mathcal{N} = 2$ $\text{SU}(3) \times \text{U}(1)$ invariant solutions. Here, we have extended this result to all other AdS solutions in the $\text{SU}(3)$ -invariant sectors of $\text{SO}(8)$ and $\text{ISO}(7)$ gauged supergravities with the same symmetry and supersymmetry.

The situation is similar, though slightly different, for the type IIB AdS₄ S-fold solutions that uplift from $D = 4$ $\mathcal{N} = 8$ $(\text{SO}(6) \times \text{SO}(1, 1)) \ltimes \mathbb{R}^{12}$ -gauged supergravity. According to table 1, this supergravity also has critical points with the same symmetry $G \supset \text{SU}(3)$ and supersymmetry as other critical points of the $\text{SO}(8)$ and $\text{ISO}(7)$ gauging: $\mathcal{N} = 0$ $\text{SO}(6)$, $\mathcal{N} = 1$ $\text{SU}(3)$ and $\mathcal{N} = 0$ $\text{SU}(3)$. The former two have the same spectrum within their corresponding $D = 4$ supergravities, while the latter does not. For this reason, we will only be interested in the former two vacua. Both for the $\mathcal{N} = 1$ $\text{SU}(3)$ and the $\mathcal{N} = 0$ $\text{SO}(6) \sim \text{SU}(4)$ solutions there are combinations, $[L^2 \text{tr} \tilde{M}_{(n)}^2]_{\text{IIB}}$, of the eigenvalues in table 4 that match the quantities $[L^2 \text{tr} M_{(k)}^2]_{\text{IIA}}$ and $[L^2 \text{tr} M_{(n)}^2]_{11D}$ for the solutions with the same symmetry for a certain choice of the period T . The tilde in $[L^2 \text{tr} \tilde{M}_{(n)}^2]_{\text{IIB}}$ is taken to signify that, in this case, the combinations also involve subtraction of eigenvalues. More concretely, consider the following quantities for the type IIB solutions with the quantum numbers fixed as indicated:

$$\begin{aligned} \mathcal{N} = 1, \text{SU}(3) : \\ L^2 \text{tr} \tilde{M}_{(1)}^2 &\equiv L^2 \left[\sum_{p=0}^{\ell} M_{\ell,p,j}^2 d_{\ell,p,j} \right]_{|\ell=1,j=0} \\ &\quad - L^2 [M_{\ell,p,j}^2 d_{\ell,p,j}]_{|\ell=p=0,j=-1} - L^2 [M_{\ell,p,j}^2 d_{\ell,p,j}]_{|\ell=p=0,j=+1}, \\ \mathcal{N} = 0, \text{SO}(6)_v : \\ L^2 \text{tr} \tilde{M}_{(1)}^2 &\equiv L^2 [M_{\ell,j}^2 d_{\ell,j}]_{|\ell=1,j=0} \\ &\quad - L^2 [M_{\ell,j}^2 d_{\ell,j}]_{|\ell=0,j=-1} - L^2 [M_{\ell,j}^2 d_{\ell,j}]_{|\ell=0,j=+1}. \end{aligned} \quad (2.50)$$

These quantities involve sums of mass eigenvalues, weighted with their degeneracies as given in table 4, and affected by a + or a - sign depending on whether $j = 0$ or $j \neq 0$.

Plugging in the expressions given in the table, the quantity $L^2 \text{tr} \tilde{M}_{(1)}^2$ for the $\mathcal{N} = 1$, $\text{SU}(3)$ solution evaluates to $\frac{65}{3}$ if $T = 2\pi$, matching the quantity $L^2 \text{tr} M_{(k=1)}^2$ for its counterpart type IIA solution at KK level $k = 1$, given in (2.46). Similarly, $L^2 \text{tr} \tilde{M}_{(1)}^2$ for the $\mathcal{N} = 0$, $\text{SO}(6)_v$ solution evaluated using the expressions given in table 4 gives $\frac{39}{2}$ for $T = 2\pi$. This again matches the quantity $L^2 \text{tr} M_{(k=1)}^2$ at KK level $k = 1$ given in (2.46) for the $\mathcal{N} = 0$, $\text{SO}(6)_v$ solution of massive IIA. It also matches $L^2 \text{tr} M_{(n=1)}^2$ at KK level $n = 1$ given in (2.45) for the $D = 11$ $\mathcal{N} = 0$, $\text{SU}(4)_c$ solution. Although it is not as clear cut in the type IIB case, it will be argued in section 5 that the states that enter the sums in (2.50) also belong to KK level $m = 1$ in an $\text{SL}(8)$ -covariant sense. The formal analytic continuation $j' = ij$, with $i^2 = -1$, removes the minus signs in (2.50). Under this analytic continuation, relations similar to (2.47) and (2.48) relate these formal sums at higher KK levels for these type IIB solutions to their $D = 11$ and type IIA counterparts. We will return to this point in section 5. In the next section, we will turn to compute the KK graviton spectrum of a different type IIB S-fold solution of particular interest.

3 Graviton spectrum on the $\mathcal{N} = 4$ $\text{SO}(4)$ type IIB S-fold

The spin-2 spectrum about the AdS_4 solution [17] of type IIB supergravity that uplifts on a six-dimensional S-fold from the $\mathcal{N} = 4$ $\text{SO}(4)$ -invariant critical point [8] of $D = 4$ $\mathcal{N} = 8$ $[\text{SO}(6) \times \text{SO}(1, 1)] \times \mathbb{R}^{12}$ -gauged supergravity [4, 2] can be computed as in section 2. This solution has concrete field theory duals for specific choices of the period T of the S-fold coordinate η [50] (see also [51–55]). For this reason, it is particularly interesting to go in some detail about the corresponding spin-2 spectrum.

The relevant ten-dimensional geometry is (1.1) with $d = 6$ and [17]

$$d\tilde{s}_6^2 = g^{-2} L^{-2} \left[d\eta^2 + \frac{dr^2}{1-r^2} + \frac{r^2}{1+2r^2} ds^2(S_1^2) + \frac{1-r^2}{3-2r^2} ds^2(S_2^2) \right],$$

$$e^{2A} = L^2 \left[(1+2r^2)(3-2r^2) \right]^{1/4}, \quad (3.1)$$

with $L^2 = \frac{1}{2}g^{-2}$. The coordinate r ranges as $0 \leq r \leq 1$ and η is taken to be periodic, $\eta \sim \eta + T$, for some $T > 0$. Also, $ds^2(S_i^2)$, $i = 1, 2$, is the round, Einstein metric on each of two spheres S_i^2 . These are rotated by the $\text{SO}(4) = \text{SO}(3)_1 \times \text{SO}(3)_2$ isometry of the geometry (3.1). This $\text{SO}(4)$ isometry is also respected by the type IIB forms and is thus a symmetry, in fact the R-symmetry, of the full $\mathcal{N} = 4$ ten-dimensional solution. In contrast, the $\text{U}(1)_\eta$ isometry generated by ∂_η is broken by the supergravity forms. Topologically, the metric (3.1) extends over $S^5 \times S_\eta^1$ [17], with the S^5 directions corresponding to r and S_i^2 , $i = 1, 2$.

Next, we plug the geometry (3.1) into the PDE (1.4). It is natural to separate the eigenfunction \mathcal{Y} following the $\text{SO}(3)_1 \times \text{SO}(3)_2 \times \text{U}(1)_\eta$ isometry as

$$\mathcal{Y} = f(r) \mathcal{Y}_{(1)} \mathcal{Y}_{(2)} \mathcal{Y}_\eta, \quad (3.2)$$

where $\mathcal{Y}_{(1)}$, $\mathcal{Y}_{(2)}$ and \mathcal{Y}_η are the spherical harmonics on S_1^2 , S_2^2 and S_η^1 ,

$$\square_{S_1^2} \mathcal{Y}_{(1)} = -\ell_1(\ell_1 + 1) \mathcal{Y}_{(1)}, \quad \square_{S_2^2} \mathcal{Y}_{(2)} = -\ell_2(\ell_2 + 1) \mathcal{Y}_{(2)}, \quad \partial_\eta^2 \mathcal{Y}_\eta = -\left(\frac{2\pi}{T} j\right)^2 \mathcal{Y}_\eta, \quad (3.3)$$

with

$$\ell_1 = 0, 1, 2, \dots, \quad \ell_2 = 0, 1, 2, \dots, \quad j = 0, \pm 1, \pm 2, \dots \quad (3.4)$$

The separation of variables (3.2) turns the PDE (1.4) into the following ODE in r :

$$(1-r^2)f'' + \left(\frac{2}{r} - 5r\right)f' + \left[2M^2L^2 - \frac{4\pi^2}{T^2}j^2 - \frac{1+2r^2}{r^2}\ell_1(\ell_1+1) - \frac{3-2r^2}{1-r^2}\ell_2(\ell_2+1)\right]f = 0, \quad (3.5)$$

with a prime denoting derivative with respect to r . Finally, the change of variables

$$r^2 = u, \quad f = u^{\ell_1/2}(1-u)^{\ell_2/2}H(u), \quad (3.6)$$

where now $0 \leq u \leq 1$, brings the ODE (3.5) into hypergeometric form (2.8) with

$$a_{\pm} = \frac{1}{2}(\ell_1 + \ell_2 + 2) \pm \frac{1}{2}\sqrt{4 + 2M^2L^2 - 2\ell_1(\ell_1 + 1) - 2\ell_2(\ell_2 + 1) - \frac{4\pi^2}{T^2}j^2}, \quad c = \ell_1 + \frac{3}{2}. \quad (3.7)$$

The two linearly independent solutions to (2.8) are written in (2.10). In the present case, the second solution therein blows up at $u = 0$ since $\ell_1 \geq 0$ by (3.4), and is therefore excluded. The first solution in (2.10) is regular at $u = 0$ for all values of the parameters (3.7), and is also regular at $u = 1$ if $a_- = -h$ for a non-negative integer h . For this choice, the eigenfunction becomes a polynomial in u . Defining a new quantum number

$$\ell = 2h + \ell_1 + \ell_2, \quad (3.8)$$

the graviton masses then follow from (3.7) as

$$L^2M_{\ell, \ell_1, \ell_2, j}^2 = \frac{1}{2}\ell(\ell + 4) + \ell_1(\ell_1 + 1) + \ell_2(\ell_2 + 1) + \frac{2\pi^2}{T^2}j^2, \quad (3.9)$$

with the quantum numbers now ranging as

$$\ell = 0, 1, 2, \dots, \quad \ell_1 = 0, 1, \dots, \ell, \quad \ell_2 = 0, 1, \dots, \ell, \quad j = 0, \pm 1, \pm 2, \dots, \quad (3.10)$$

by (3.4), (3.8) and the fact that $h \geq 0$. The number of KK gravitons with mass (3.9) is

$$d_{\ell, \ell_1, \ell_2, j} = \begin{cases} (2\ell_1 + 1)(2\ell_2 + 1) & , \quad \text{if } j = 0 \\ 2(2\ell_1 + 1)(2\ell_2 + 1) & , \quad \text{if } j \neq 0. \end{cases} \quad (3.11)$$

To summarise, the KK gravitons about the $\mathcal{N} = 4$ SO(4)-invariant AdS₄ S-fold solution of type IIB supergravity reported in [17] have masses (3.9), with quantum numbers ranging as in (3.10) and degeneracy (3.11). The corresponding eigenfunctions are given by (3.2) with (3.6) and the first hypergeometric function in (2.10) for $H(u)$, which now becomes a polynomial. More precisely, the eigenfunctions (3.2) are products of S^5 spherical harmonics branched out in $\text{SO}(4) = \text{SO}(3)_1 \times \text{SO}(3)_2$ representations, and harmonics on S^1_{η} . A few of these modes for low values of the quantum numbers have been tabulated in table 5. The table includes the dimension Δ of the corresponding operators in the dual field theory [50], where

$$M^2L^2 = \Delta(\Delta - 3). \quad (3.12)$$

m	$(\ell, \ell_1, \ell_2, j)$	$d_{\ell, \ell_1, \ell_2, j}$	$L^2 M_{\ell, \ell_1, \ell_2, j}^2$	$\Delta_{\ell, \ell_1, \ell_2, j}$	Dual operator	Short?
0	(0, 0, 0, 0)	1	0	3	$\mathcal{T}_{\alpha\beta}^{(0)} _{s=2}$	✓
1	(1, 1, 0, 0), (1, 0, 1, 0)	6	$\frac{9}{2}$	$\frac{3}{2}(\sqrt{3}+1)$	$\mathcal{T}_{\alpha\beta}^{(0)}\mathcal{R}\mathcal{X}_{(1)}^a _{s=2}, S_{(1)} \leftrightarrow S_{(2)}$	
	(0, 0, 0, 1)	2	$\frac{2\pi^2}{7^2}$	$\frac{1}{2}(\sqrt{9+\frac{8\pi^2}{7^2}}+3)$	$\mathcal{T}_{\alpha\beta}^{(0)}\mathcal{Z} _{s=2}, \text{c.c.}$	
2	(2, 0, 0, 0)	1	6	$\frac{1}{2}(\sqrt{33}+3)$	$\mathcal{T}_{\alpha\beta}^{(0)}(1-2\mathcal{R}^2) _{s=2}$	
	(2, 1, 1, 0)	9	10	5	$\mathcal{T}_{\alpha\beta}^{(0)}\mathcal{R}(1-\mathcal{R}^2)^{1/2}\mathcal{X}_{(1)}^a\mathcal{X}_{(2)}^b _{s=2}$	✓
	(2, 2, 0, 0), (2, 0, 2, 0)	10	12	$\frac{1}{2}(\sqrt{57}+3)$	$\mathcal{T}_{\alpha\beta}^{(0)}\mathcal{R}^2(\mathcal{X}_{(1)}^a\mathcal{X}_{(1)}^b - \frac{1}{3}\mathcal{R}^2\delta^{ab}) _{s=2}, S_{(1)} \leftrightarrow S_{(2)}$	
	(1, 1, 0, 1), (1, 0, 1, 1)	12	$\frac{2\pi^2}{7^2} + \frac{9}{2}$	$\frac{1}{2}(\sqrt{27+\frac{8\pi^2}{7^2}}+3)$	$\mathcal{T}_{\alpha\beta}^{(0)}\mathcal{R}\mathcal{X}_{(1)}^a\mathcal{Z} _{s=2}, \text{c.c.}, S_{(1)} \leftrightarrow S_{(2)}$	
	(0, 0, 0, 2)	2	$\frac{8\pi^2}{7^2}$	$\frac{1}{2}(\sqrt{9+\frac{32\pi^2}{7^2}}+3)$	$\mathcal{T}_{\alpha\beta}^{(0)}\mathcal{Z}^2 _{s=2}, \text{c.c.}$	
	(0, 0, 0, 0) Redundant	1	0	3	$\mathcal{T}_{\alpha\beta}^{(0)} _{s=2}$	✓
3	(3, 1, 0, 0), (3, 0, 1, 0)	6	$\frac{25}{2}$	$\frac{1}{2}(\sqrt{59}+3)$	$\mathcal{T}_{\alpha\beta}^{(0)}\mathcal{R}(1-\frac{8}{5}\mathcal{R}^2)\mathcal{X}_{(1)}^a _{s=2}, S_{(1)} \leftrightarrow S_{(2)}$	
	(3, 2, 1, 0), (3, 1, 2, 0)	30	$\frac{37}{2}$	$\frac{1}{2}(\sqrt{83}+3)$	$\mathcal{T}_{\alpha\beta}^{(0)}\mathcal{R}^2(\mathcal{X}_{(1)}^a\mathcal{X}_{(1)}^b - \frac{1}{3}\mathcal{R}^2\delta^{ab})\mathcal{X}_{(2)}^c _{s=2}, S_{(1)} \leftrightarrow S_{(2)}$	
	(3, 3, 0, 0), (3, 0, 3, 0)	14	$\frac{45}{2}$	$\frac{3}{2}(\sqrt{11}+1)$	$\mathcal{T}_{\alpha\beta}^{(0)}\mathcal{R}^3(\mathcal{X}_{(1)}^a\mathcal{X}_{(1)}^b\mathcal{X}_{(1)}^c - \text{traces}) _{s=2}, S_{(1)} \leftrightarrow S_{(2)}$	
	(2, 0, 0, 1)	2	$\frac{2\pi^2}{7^2} + 6$	$\frac{1}{2}(\sqrt{33+\frac{8\pi^2}{7^2}}+3)$	$\mathcal{T}_{\alpha\beta}^{(0)}(1-2\mathcal{R}^2)\mathcal{Z} _{s=2}, \text{c.c.}$	
	(2, 1, 1, 1)	18	$\frac{2\pi^2}{7^2} + 10$	$\frac{1}{2}(\sqrt{49+\frac{8\pi^2}{7^2}}+3)$	$\mathcal{T}_{\alpha\beta}^{(0)}\mathcal{R}(1-\mathcal{R}^2)^{1/2}\mathcal{X}_{(1)}^a\mathcal{X}_{(2)}^b\mathcal{Z} _{s=2}, \text{c.c.}$	
	(2, 2, 0, 1), (2, 0, 2, 1)	20	$\frac{2\pi^2}{7^2} + 12$	$\frac{1}{2}(\sqrt{57+\frac{8\pi^2}{7^2}}+3)$	$\mathcal{T}_{\alpha\beta}^{(0)}\mathcal{R}^2(\mathcal{X}_{(1)}^a\mathcal{X}_{(1)}^b - \frac{1}{3}\mathcal{R}^2\delta^{ab})\mathcal{Z} _{s=2}, \text{c.c.}, S_{(1)} \leftrightarrow S_{(2)}$	
	(1, 1, 0, 2), (1, 0, 1, 2)	12	$\frac{8\pi^2}{7^2} + \frac{9}{2}$	$\frac{1}{2}(\sqrt{27+\frac{32\pi^2}{7^2}}+3)$	$\mathcal{T}_{\alpha\beta}^{(0)}\mathcal{R}\mathcal{X}_{(1)}^a\mathcal{Z}^2 _{s=2}, \text{c.c.}, S_{(1)} \leftrightarrow S_{(2)}$	
	(0, 0, 0, 3)	2	$\frac{18\pi^2}{7^2}$	$\frac{3}{2}(\sqrt{1+\frac{8\pi^2}{7^2}}+1)$	$\mathcal{T}_{\alpha\beta}^{(0)}\mathcal{Z}^3 _{s=2}, \text{c.c.}$	
	(1, 1, 0, 0), (1, 0, 1, 0) Redundant	12	$\frac{9}{2}$	$\frac{3}{2}(\sqrt{3}+1)$	$\mathcal{T}_{\alpha\beta}^{(0)}\mathcal{R}\mathcal{X}_{(1)}^a _{s=2}, S_{(1)} \leftrightarrow S_{(2)}$	
	(0, 0, 0, 1) Redundant	4	$\frac{2\pi^2}{7^2}$	$\frac{1}{2}(\sqrt{9+\frac{8\pi^2}{7^2}}+3)$	$\mathcal{T}_{\alpha\beta}^{(0)}\mathcal{Z} _{s=2}, \text{c.c.}$	

Table 5. The spin-2 spectrum of the $\mathcal{N} = 4$ S-fold solution. We have employed the notation $\mathcal{R}^2 = \delta_{ab}\mathcal{X}_{(1)}^a\mathcal{X}_{(1)}^b$ and $S_{(1)} \leftrightarrow S_{(2)}$ means exchange of the labels (1) and (2), and simultaneously $\mathcal{R}^2 \leftrightarrow 1 - \mathcal{R}^2$. The spectrum is organised in SL(8) KK levels (see section 5) $m = 0, 1, \dots$, and this leads to some redundant states as discussed in that section. The dual operators are single-trace. The tr symbol has been omitted.

The field theory is defined on a stack of D3-branes wrapped on S_η^1 . The form of the dual single-trace spin-2 operators can be inferred from the supergravity eigenfunctions. They correspond to bound states of the energy-momentum tensor $T_{\mu\nu}$, the six scalars $X_{(1)}^a, X_{(2)}^a$, $a = 1, 2, 3$, corresponding to the directions transverse to the D3-branes, and a complex coordinate Z on the D3-brane. In the table, all these have been promoted to superfields. The trace in the adjoint of the gauge group is understood.

The full KK spectrum about this $\mathcal{N} = 4$ solution must lie in representations of $\text{OSp}(4|4)$. A classification of these representations can be found in, for example, [56]. The massless graviton should belong to a short graviton multiplet, A_2 in the notation of [56], with $\ell_1 = \ell_2 = 0$. This multiplet contains gravitini and vectors in the fundamental and the adjoint, respectively, of the R-symmetry group $\text{SO}(4)$. It also contains spin-1/2 fields in the fundamental of $\text{SO}(4)$ and two singlet scalars, see e.g. equation (5.55) of [56]. By the results of [57–59], a consistent truncation of type IIB supergravity must exist at the full non-linear level on this S-fold solution to the pure $D = 4$ $\mathcal{N} = 4$ gauged supergravity with the field content just described. The $\text{SO}(4)$ -singlet scalars in this $D = 4$ $\mathcal{N} = 4$ supergravity parametrise an $\text{SL}(2, \mathbb{R})/\text{SO}(2)$ non-linear sigma model. The numerator group is inherited from the S-duality of type IIB supergravity.

The supergroup $\text{OSp}(4|4)$ also admits massive short representations. Some gravitons in the spectrum can be identified to belong to such representations, particularly A_2 in the notation of [56]. Spin-2 states in these short multiplets arise as Q^4 descendants of the superconformal primary therein, and thus saturate the unitarity bound

$$\Delta = \ell_1 + \ell_2 + 3. \tag{3.13}$$

Via (3.12), these states have masses

$$L^2 M^2 = (\ell_1 + \ell_2)(\ell_1 + \ell_2 + 3). \tag{3.14}$$

In the spectrum (3.9), short states of this type do indeed arise for any period T and all $\ell_1 = 0, 1, 2, \dots$, whenever $j = 0$ and the other quantum numbers ℓ, ℓ_1, ℓ_2 are related through $2\ell_1 = 2\ell_2 = \ell$ (so that $h = 0$ in (3.8) and ℓ is even). From (3.9) and (3.12), the masses and conformal dimensions of states with these quantum numbers are

$$L^2 M_{\ell_1}^2 = 2\ell_1(2\ell_1 + 3), \quad \Delta_{\ell_1} = 2\ell_1 + 3, \quad \ell_1 = 0, 1, 2, \dots, \tag{3.15}$$

which are indeed of the form (3.14), (3.13) and thus short.

All other gravitons belong to long multiplets. It can be checked that, away from the shortening relations among the quantum numbers, their Δ 's computed through (3.12) from (3.14) are always above the unitarity bound (3.13). We conclude, however, with the following tantalising observation. Consider the analytical continuation of the quantum number j into $j' = ij$, with $i^2 = -1$, as at the end of section 2.4, and fix the S-fold coordinate period to $T = 2\pi$. Then, it follows from (3.9), that states with $j' = \pm 1$, and $\ell_2 = \ell_1 + 1$, and $\ell = 2\ell_1 + 1$ (so that $h = 0$ in (3.8) and ℓ is odd) are also short with conformal dimensions $\Delta_{\ell_1} = 2(\ell_1 + 2)$ and masses $L^2 M_{\ell_1}^2 = 4\ell_1(\ell_1 + 2)$, for all $\ell_1 = 0, 1, 2, \dots$

4 Kaluza-Klein graviton mass matrix

We now switch gears to obtain a covariant expression for the infinite-dimensional KK graviton mass matrix and its associated trace formulae.

4.1 The mass matrix

We would like to determine the KK graviton mass matrix corresponding to string/M-theory AdS_4 solutions that uplift, at least, on the relevant spheres from the $\text{SO}(8)$ and $\text{ISO}(7)$ gaugings. In appendix A of [22], an $\text{SO}(7)$ -covariant mass matrix was derived for KK gravitons about solutions that uplift from the $\text{ISO}(7)$ gauging. Here, we would like to extend those results into a mass matrix that is formally $\text{SL}(8)$ covariant, in agreement with the formal, manifest covariance that the $\text{SO}(8)$ and $\text{ISO}(7)$ gauged supergravities take on using the embedding tensor formalism [60] particularised to gaugings contained in $\text{SL}(8) \subset \text{E}_{7(7)}$.

In order to do this, we start by assuming that the mass eigenfunctions,

$$\mathcal{Y}^{A_1 \dots A_m} = \mu^{(A_1} \dots \mu^{A_m)}, \quad m = 0, 1, 2, \dots, \tag{4.1}$$

are symmetric polynomials of the \mathbb{R}^8 coordinates μ^A , $A = 1, \dots, 8$. The latter are formally in the fundamental of $\text{SL}(8)$ and constrained as

$$\theta_{AB} \mu^A \mu^B = 1, \tag{4.2}$$

with $\theta_{AB} = \delta_{AB}$ for the $\text{SO}(8)$ gauging and $\theta_{AB} = \text{diag}(1, 1, 1, 1, 1, 1, 1, 0)$ for the $\text{ISO}(7)$ gauging. We then pose the KK graviton mass equation (1.4) for the consistent embedding metrics on the S^7 and S^6 as given in [44, 14]. With these assumptions, we follow similar steps to those in [22] to transform the PDE (1.4) into an algebraic eigenvalue problem by reading off an infinite-dimensional block-diagonal mass matrix,

$$\mathcal{M}^2 = \text{diag}(M_{(0)}^2, M_{(1)}^2, M_{(2)}^2, \dots, M_{(m)}^2, \dots). \tag{4.3}$$

Now, each block is an $\text{SL}(8)$ -covariant square matrix of size

$$\dim M_{(m)}^2 \equiv [m, 0, 0, 0, 0, 0, 0]_{\text{SL}(8)} = \begin{pmatrix} m+7 \\ m \end{pmatrix}. \tag{4.4}$$

These blocks are given explicitly by the following expressions: for $m = 0$,

$$M_{(0)}^2 = 0, \tag{4.5}$$

for $m = 1$,

$$(M_{(1)}^2)_{A^B} = -g^2 \mathcal{M}^{\text{MN}} \Theta_{\mathbb{M}}^{B C} \Theta_{\mathbb{N}}^C{}_{A^}, \tag{4.6}$$

for $m = 2$,

$$(M_{(2)}^2)_{A_1 A_2}{}^{B_1 B_2} = -2g^2 \mathcal{M}^{\text{MN}} \left[\Theta_{\mathbb{M}}^{(B_1| C} \Theta_{\mathbb{N}}^C{}_{(A_1} \delta_{A_2)}{}^{B_2)} + \Theta_{\mathbb{M}}^{(B_1}{}_{(A_1} \Theta_{\mathbb{N}}^{B_2)}{}_{A_2)} \right], \tag{4.7}$$

and for $m \geq 3$,

$$(M_{(m)}^2)_{A_1 \dots A_m}{}^{B_1 \dots B_m} = -m g^2 \mathcal{M}^{\text{MN}} \left[\Theta_{\mathbb{M}}^{(B_1| C} \Theta_{\mathbb{N}}^C{}_{(A_1} \delta_{A_2}{}^{B_2} \dots \delta_{A_m)}{}^{B_m)} + (m-1) \Theta_{\mathbb{M}}^{(B_1}{}_{(A_1} \Theta_{\mathbb{N}}^{B_2}{}_{A_2} \delta_{A_3}{}^{B_3} \dots \delta_{A_m)}{}^{B_m)} \right]. \tag{4.8}$$

Compared to [22], these expressions involve no trace removal within same-level indices. We have also restored the embedding tensors, $\Theta_{\mathbb{M}}^A{}_B$, with $\mathbb{M} = ([AB], [AB])$ a fundamental $\text{E}_{7(7)}$ index, and

$$\Theta_{[AB]}{}^C{}_D = 2 \delta_{[A}^C \theta_{B]D}, \quad \Theta^{[AB]C}{}_D = 2 \delta_D^{[A} \xi^{B]C}, \tag{4.9}$$

where θ_{AB} was defined below (4.2), and $\xi^{AB} = 0$ for the $\text{SO}(8)$ gauging while $\xi^{AB} = \text{diag}(0, 0, 0, 0, 0, 0, 0, m/g)$ for $\text{ISO}(7)$. Finally, g here and in (4.6)–(4.8) is the electric gauge coupling, m the magnetic coupling, and \mathcal{M}^{MN} is the inverse $\mathcal{N} = 8$ scalar matrix. Like the bosonic mass matrices of $D = 4$ $\mathcal{N} = 8$ gauged supergravity (see [26]), the KK graviton mass matrices (4.5)–(4.8) are quadratic in the $D = 4$ embedding tensor.

Since we have refrained ourselves from removing traces on same-level indices of the KK graviton mass matrices (4.5)–(4.8), the latter are manifestly $\text{SL}(8)$ -covariant. We can think

of these as blocks in the diagonal of the infinite dimensional graviton mass matrix (4.3). The integer m can be thought of as an $SL(8)$ KK level. Proceeding like this, though, the price one pays for the $SO(8)$ gauging is that the spectrum at fixed $SL(8)$ KK level $m \geq 0$ contains repeated physical modes: it includes modes of all $SO(8)$ KK levels n (as defined below (2.25)) such that $n = m - 2s$ according to

$$[m, 0, 0, 0, 0, 0, 0, 0]_{SL(8)} \xrightarrow{SO(8)} \sum_{s=0}^{\lfloor \frac{m}{2} \rfloor} [m - 2s, 0, 0, 0, 0]_{SO(8)}. \quad (4.10)$$

For the $ISO(7)$ gauging, there is an even larger overcounting. Every $SL(8)$ level $m \geq 0$ formally contains the $SO(8)$ levels n specified in (4.10), and each of these, in turn, includes all $SO(7)$ levels $k = 0, 1, \dots, n$ by (2.25). The repeated states can be projected out following (4.10) and (2.25), leaving only physical modes. We also remark that it is the full embedding tensor for the dyonic $ISO(7)$ gauging, including the magnetic contributions from ξ^{AB} , that enters (4.5)–(4.8) for this gauging.

Taking into account this overcounting, we have verified up to $SL(8)$ KK level $m = 3$ that, particularising the KK graviton mass matrices (4.5)–(4.8) to the corresponding critical points of the $SO(8)$ and $ISO(7)$ gaugings, their eigenvalues correctly reproduce the spectra given in tables 2 and 3. Note that, in order to obtain matching, g^2 here has to be traded for L^2 there by making use of (A.2) with the appropriate scalar potential (A.4) or (A.5). See section 5 for a discussion of the case corresponding to the $[SO(6) \times SO(1, 1)] \ltimes \mathbb{R}^{12}$ gauging.

4.2 Mass matrix traces

By (4.10), the first $SL(8)$ KK level $m = 1$ contains only the first $SO(8)$ KK level $n = 1$ (in our conventions, $\mathbf{8} \rightarrow \mathbf{8}_v$). For the $ISO(7)$ gauging, the first $SL(8)$ KK level contains the $SO(7)$ KK levels $k = 0$, which has zero mass by (4.5), and $k = 1$. At $SL(8)$ KK level $m = 1$, the eigenvalues of the $SL(8)$ -covariant mass matrix (4.6) thus reproduce the first KK-level eigenvalues with no overcounting for the $SO(8)$ gauging. For the $ISO(7)$ gauging, the first $SL(8)$ level also reproduces the $k = 1$ eigenvalues, together with an extra zero eigenvalue corresponding to $k = 0$. The trace of the $SL(8)$ level $m = 1$ mass matrix (4.6),

$$\text{tr } M_{(1)}^2 = -g^2 \mathcal{M}^{MN} \Theta_M^A{}_B \Theta_N^B{}_A, \quad (4.11)$$

must thus reproduce the KK level-one traces discussed in section 2.4, which indeed it does. Particularising (4.11) to each specific critical point with at least $SU(3)$ symmetry of the $SO(8)$ and $ISO(7)$ gaugings, making use of the relevant embedding tensors, and again trading g^2 for L^2 , all the r.h.s.'s of (2.45) and (2.46) with $n = 1$ and $k = 1$ are reproduced. For example, using the appropriate embedding tensors and vevs, we find that (4.11) evaluates to $\frac{56}{3}$, both for the $SU(3) \times U(1)_c$ point of the $SO(8)$ gauging and for the $SU(3) \times U(1)_v$ point of the $ISO(7)$ gauging, once that g^2 is replaced with the relevant L^2 . The trace relation (2.48) is a direct consequence of (4.11) and the overcounting feature mentioned in section 4.1.

In order to check consistency, it is also useful to evaluate (4.11) at generic invariant loci of the $D = 4$ $\mathcal{N} = 8$ scalar manifold. For example, on the $SU(3)$ -invariant sector of the $SO(8)$ gauging, (4.11) reduces to

$$\text{tr } M_{(1)}^2 = 2g^2 e^{-3\varphi - 2\phi} \left[X^3 e^{2\phi} + 6e^{2\varphi} X (Y^2 + Z^2) + 3(X + 4Y)e^{4\varphi + 2\phi} + 6Xe^{2\varphi + 4\phi} \right], \quad (4.12)$$

where the shorthand notations (A.7) have been employed. Further restricting (4.12) to the G_2 -invariant locus (A.8), we find

$$\text{tr } M_{(1)}^2 = 14g^2 (e^{-3\varphi} X^3 + 3e^\varphi X), \quad (4.13)$$

in agreement with (2.17). Alternatively, particularising (4.12) to the $SU(4)_c$ -invariant sector through (A.9), we get

$$\text{tr } M_{(1)}^2 = 8g^2 (e^{3\varphi} + 6e^{2\phi + \varphi}). \quad (4.14)$$

This expression can be retrieved from (2.32). Similarly, for the dyonic $ISO(7)$ gauging and $SU(3)$ -invariant scalars, (4.11) reduces to

$$\text{tr } M_{(1)}^2 = 6g^2 [e^\varphi (X + 4Y) + 2Xe^{2\phi - \varphi}], \quad (4.15)$$

in agreement with (2.34) of [22].

For the $SO(8)$ and dyonic $ISO(7)$ gaugings, only the electric embedding tensor θ_{AB} actually participates in the trace formula (4.11), while its magnetic counterpart ξ^{AB} drops out. In order to see this, we expand (4.11) using the generic form (4.9) of the embedding tensor, to obtain

$$\text{tr } M_{(1)}^2 = -g^2 (\mathcal{M}^{ACDE} \theta_{CD} \theta_{EA} + \mathcal{M}_{ACDE} \xi^{CD} \xi^{EA} - 2\mathcal{M}^{CD}{}_{CF} \theta_{DA} \xi^{FA}). \quad (4.16)$$

For the $SO(8)$ gauging, the statement is immediate because $\xi^{AB} = 0$, and only the first term is non-vanishing. For the dyonic $ISO(7)$ gauging, the only relevant term in (4.16) is again the first one, because $\mathcal{M}_{8888} = 0$ and $\theta\xi = 0$.

5 Discussion

We have verified that all solutions with the same residual supersymmetry, the same bosonic symmetry containing $SU(3)$ and the same spectrum within the $SO(8)$ [9], $ISO(7)$ [10] and $(SO(6) \times SO(1,1)) \ltimes \mathbb{R}^{12}$ [4] gaugings all have different spectra of KK gravitons. The universality of the lowest KK level spectra is thus lost at higher KK levels. However, following [22], we have found that this universality still persists up the different KK graviton towers in a weaker sense. Certain sums of KK graviton masses remain the same for all such solutions. We emphasise that this is the case if the gauged supergravity solutions with the same (super)symmetries also had the same spectrum at lowest KK level, namely, within their corresponding $\mathcal{N} = 8$ supergravities.

For example, the $\mathcal{N} = 0$ $SU(3)$ -invariant points of dyonic $ISO(7)$ [10] and $(SO(6) \times SO(1,1)) \ltimes \mathbb{R}^{12}$ supergravities [20] do not have the same spectrum within the $D = 4$ supergravities. The relevant sums of KK graviton masses are also different. A similar observation

holds for the $U(1)$ -invariant $\mathcal{N} = 1$ points of $ISO(7)$ supergravity recently reported in [61]. Although they have the same residual symmetry, these points have different spectra within $ISO(7)$ supergravity. Using the formula (4.11), we have verified at KK level one that the relevant KK graviton mass sums are also different for these two solutions. It is also worth stressing that the trace universality property works for solutions that check out the above requirements independently of how the common residual symmetry group is embedded into the corresponding gauge group and ultimately $E_{7(7)}$. This was already noted in [22], where these observations were made for the $\mathcal{N} = 2$ $SU(3) \times U(1)$ -invariant points of $SO(8)$ and $ISO(7)$ supergravity. In the former case, the solution is embedded as $SU(3) \times U(1)_c$ and in the latter as $SU(3) \times U(1)_v$.

The relevant sums of KK graviton masses have been argued to be related to traces of the KK graviton mass matrix at fixed KK level. We have provided an $SL(8)$ -covariant expression for the latter in section 4. This mass matrix has qualitatively the same form as the bosonic mass matrices of $D = 4$ $\mathcal{N} = 8$ gauged supergravity in that it is quadratic in the embedding tensor and depends on the $E_{7(7)}/SU(8)$ $\mathcal{N} = 8$ (inverse) scalar matrix \mathcal{M}^{MN} . The $SL(8)$ -covariant KK graviton mass matrix (4.5)–(4.8) reproduces the KK graviton spectra of tables 2 and 3 for the uplifts of $SO(8)$ and $ISO(7)$ gauged supergravity critical points, with redundancies introduced by the $SL(8)$ representations (versus branchings of $SO(8)$ and $SO(7)$ representations, as explained in section 4.1). Interestingly, the mass matrix (4.5)–(4.8) also reproduces the KK graviton spectrum of table 4 and equation (3.9) for type IIB S-folds with period $T = 2\pi$ that uplift from vacua of the $(SO(6) \times SO(1, 1)) \times \mathbb{R}^{12}$ gauging, *provided* the $U(1)_\eta$ quantum number j is analytically continued as $j' = ij$, with $i^2 = -1$. The origin of this analytic continuation can be put down to the fact that the $SL(8)$ -covariant graviton mass matrix formula (4.5)–(4.8) actually sees the compactified $U(1)_\eta$ as the non-compact $SO(1, 1)$ factor in the $D = 4$ gauge group $(SO(6) \times SO(1, 1)) \times \mathbb{R}^{12}$. From a IIB perspective, this factor is associated to a hyperboloid uplift [17]. In any case, the (analytically continued) spectra of the type IIB S-folds can be also organised in $SL(8)$ KK levels $m = 0, 1, \dots$ through the branching

$$[m, 0, 0, 0, 0, 0, 0]_{SL(8)} \xrightarrow{SO(6) \times SO(1,1)} \sum_{s=0}^{\lfloor \frac{m}{2} \rfloor} \sum_{\ell=0}^{m-2s} \sum_{p=0}^{m-2s-\ell} [\ell, 0, 0]_{m-2s-\ell-2p}. \quad (5.1)$$

This approach contains redundant states that can be projected out as discussed below (4.10). It would be interesting to determine if (4.5)–(4.8) can be modified in such a way that the KK graviton spectra of vacua of the $SO(8)$ and $ISO(7)$ gauging are still obtained, and the physical spectra of the compactified S-folds is recovered as well.

The $SL(8)$ -covariant KK graviton mass matrix should have an $E_{7(7)}$ -covariant extension. From our analysis, it is not immediate to deduce what that extension should be. The trace formula (4.11) does admit a naive straightforward $E_{7(7)}$ -covariant extension, though:

$$\text{tr } M_{(1)}^2 = -g^2 \kappa_{\alpha\beta} \mathcal{M}^{MN} \Theta_M^\alpha \Theta_N^\beta, \quad (5.2)$$

with $\kappa_{\alpha\beta}$ the Killing-Cartan form of $E_{7(7)}$. This formula reduces to (4.11) for gaugings contained in $SL(8)$. It would be interesting to test (5.2) for gaugings not contained in

SL(8) that descend from higher dimensions. An example of such gaugings, which has Minkowski vacua though, is given in [62].

Acknowledgments

KD, PN and OV are supported by the NSF grant PHY-1720364. GL is supported by an FPI-UAM predoctoral fellowship. GL and OV are partially supported by grants SEV-2016-0597 and PGC2018-095976-B-C21 from MCIU/AEI/FEDER, UE.

A SU(3) invariance in $D = 4$ $\mathcal{N} = 8$ gauged supergravity

We find it useful to collect here some facts about the SU(3)-invariant sector of the three different gaugings of $D = 4$ $\mathcal{N} = 8$ supergravity considered in the main text. The field content is of course the same for all the gaugings considered but the interactions differ.

The SU(3)-invariant sector contains three scalars, φ , ϕ , a , and three pseudoscalars χ , ζ , $\tilde{\zeta}$. All these are coordinates on a submanifold

$$\frac{\text{SU}(1,1)}{\text{U}(1)} \times \frac{\text{SU}(2,1)}{\text{SU}(2) \times \text{U}(1)} \tag{A.1}$$

of $E_{7(7)}/\text{SU}(8)$, with the first factor parametrised by (φ, χ) , and the second by $(\phi, a, \zeta, \tilde{\zeta})$. The Lagrangian in this sector is

$$\begin{aligned} \mathcal{L} = & R \text{vol}_4 + \frac{3}{2}(d\varphi)^2 + \frac{3}{2}e^{2\varphi}(d\chi)^2 + 2(D\phi)^2 + \frac{1}{2}e^{4\phi} \left(Da + \frac{1}{2}(\zeta D\tilde{\zeta} - \tilde{\zeta} D\zeta) \right)^2 \\ & + \frac{1}{2}e^{2\phi}(D\zeta)^2 + \frac{1}{2}e^{2\phi}(D\tilde{\zeta})^2 + \frac{1}{2}\mathcal{I}_{\Lambda\Sigma} H_{(2)}^\Lambda \wedge *H_{(2)}^\Sigma + \frac{1}{2}\mathcal{R}_{\Lambda\Sigma} H_{(2)}^\Lambda \wedge H_{(2)}^\Sigma - V \text{vol}_4 . \end{aligned} \tag{A.2}$$

The scalar kinetic terms are given by a standard metric on (A.1). The precise form of the minimal ($D\phi$, etc.) and non-minimal ($\mathcal{R}_{\Lambda\Sigma}$, $\mathcal{I}_{\Lambda\Sigma}$) couplings of the scalars to the vectors are not needed in this paper. We do need the expression of the scalar potential V , which fixes the radius L of its AdS₄ vacua (for which $V_0 < 0$ at a critical point) as

$$L^2 = -\frac{6}{V_0} . \tag{A.3}$$

The potential is different for each gauging. For the SU(3)-invariant sector of the purely electric SO(8) gauging [9], the potential is [21], in the conventions of [18],

$$\begin{aligned} g^{-2}V = & -12e^\varphi - 6e^{-2\phi-\varphi}XY(e^{4\phi} + Y^2 + Z^2) - 12e^\varphi(Y-1) \left(1 + Y - \frac{3}{2}XY \right) \\ & + 6e^{-2\phi-\varphi}(Y-1)(e^{4\phi} + Y^2 + Z^2)X^2 + e^{-3\varphi} \left[\frac{1}{2}e^{-4\phi} + a^2 - 1 + \frac{1}{2}e^{4\phi}(1+a^2)^2 \right. \\ & \left. + \frac{1}{2}e^{-4\phi}(Y-1)(1+2Z^2 - 2e^{4\phi} + Y(1+2e^{4\phi} + 2Z^2) + Y^2 + Y^3) \right] X^3 . \end{aligned} \tag{A.4}$$

\mathcal{N}	G_0	χ	$e^{-\varphi}$	$e^{-\phi}$	a	ζ	$\tilde{\zeta}$	V_0	$L^2 M^2$
1	SU(3)	0	$\frac{\sqrt{5}m}{3g}$	$\sqrt{\frac{5}{6}}$	0	$\frac{1}{\sqrt{3}}$	$\frac{1}{\sqrt{3}}$	$-\frac{648g^3}{25\sqrt{5}m}$	$(0, 0, 4 - \sqrt{6}, 4 - \sqrt{6}, \sqrt{6} + 4, \sqrt{6} + 4)$
0	SO(6) _v	0	$\frac{m}{\sqrt{2}g}$	1	0	0	0	$-\frac{8\sqrt{2}g^3}{m}$	$(0, 0, 6, 6, -\frac{3}{4}, -\frac{3}{4})$
0	SU(3)	χ	$\frac{m}{\sqrt{2}g}$	$(1 - a^2)^{1/4}$	a	0	0	$-\frac{8\sqrt{2}g^3}{m}$	$(0, 0, 6, 6, \frac{27g^2\chi^2}{2m^2} - \frac{3}{4}, \frac{27g^2\chi^2}{2m^2} - \frac{3}{4})$

Table 6. All critical loci of $D = 4$ $\mathcal{N} = 8$ [SO(6) \times SO(1, 1)] \times \mathbb{R}^{12} -gauged supergravity with at least SU(3) invariance. All of these are AdS. For each point we give the residual supersymmetry \mathcal{N} and bosonic symmetry G_0 within the full $\mathcal{N} = 8$ theory, their location in the parametrisation that we are using and the cosmological constant V_0 and the masses of the scalars in units of the AdS radius. The $\mathcal{N} = 0$ SO(6)_v vacuum is the $\chi = a = 0$ point of the $\mathcal{N} = 0$ SU(3) critical locus.

For the dyonic ISO(7) gauging the, SU(3)-invariant potential reads [10]

$$\begin{aligned}
 V = \frac{1}{2}g^2 \left[X^3 e^{4\phi-3\varphi} + 12e^{2\phi-\varphi} \left(X^2(Y-1) - XY \right) + 12e^\varphi \left(3XY(Y-1) - 2Y^2 \right) \right] \\
 - gm\chi e^{3\varphi+2\phi} \left[6(Y-1) + e^{2\phi-2\varphi}(X-1) \right] + \frac{1}{2}m^2 e^{3\varphi+4\phi}. \quad (\text{A.5})
 \end{aligned}$$

Finally, the SU(3)-invariant potential for the [SO(6) \times SO(1, 1)] \times \mathbb{R}^{12} gauging is

$$\begin{aligned}
 V = 6g^2 e^\varphi \left[3XY(Y-1) - 2Y^2 \right] + 6gm\chi e^{3\varphi+2\phi}(Y-1) \left[1 - e^{-4\phi}(Y^2 + Z^2) \right] \\
 + \frac{1}{2}m^2 e^{3\varphi} \left[e^{4\phi} + e^{-4\phi}(Y^2 + Z^2)^2 - 2(Y^2 - 2Y + Z^2) \right], \quad (\text{A.6})
 \end{aligned}$$

as follows from [20]. In (A.4)–(A.6), g and m are the electric and magnetic gauge couplings of the parent $\mathcal{N} = 8$ supergravities. For the latter two gaugings at hand, these can be set equal, $m = g$, without loss of generality [2]. This is in fact what we have done, following [17, 20], to write the type IIB uplifts of sections 2.2 and 3. We have also employed the shorthand notations [10]

$$X \equiv 1 + e^{2\varphi}\chi^2, \quad Y \equiv 1 + \frac{1}{4}e^{2\phi}(\zeta^2 + \tilde{\zeta}^2), \quad Z \equiv e^{2\phi}a. \quad (\text{A.7})$$

The AdS vacua of the SO(8), ISO(7) and [SO(6) \times SO(1, 1)] \times \mathbb{R}^{12} $\mathcal{N} = 8$ gaugings that preserve at least the SU(3) subgroup of those gauge groups were respectively investigated in [21, 10, 20]. In our conventions, these correspond to extrema of the scalar potentials (A.4)–(A.6). The location of these vacua in scalar space, in the notation that we are using, can be respectively found in table 2 of [18], table 3 of [10] (with labels + there replaced with labels v here), and table 6 above.

For the SO(8) and ISO(7) gaugings, the G₂-invariant scalar sector (employed in the former context in section 2.1.1 of the main text) is reached from the SU(3)-invariant sector through the identifications [18, 10]

$$\phi = \varphi, \quad \tilde{\zeta} = -2\chi, \quad a = \zeta = 0. \quad (\text{A.8})$$

The $SU(4)_c$ -invariant sector of the $SO(8)$ gauging is retrieved via [18]

$$e^{-2\phi} = 1 - \frac{1}{4}(\zeta^2 + \tilde{\zeta}^2), \quad a = 0, \quad e^{-2\varphi} = 1 - \chi^2. \quad (\text{A.9})$$

Open Access. This article is distributed under the terms of the Creative Commons Attribution License ([CC-BY 4.0](https://creativecommons.org/licenses/by/4.0/)), which permits any use, distribution and reproduction in any medium, provided the original author(s) and source are credited.

References

- [1] G. Dall’Agata, G. Inverso and M. Trigiante, *Evidence for a family of $SO(8)$ gauged supergravity theories*, *Phys. Rev. Lett.* **109** (2012) 201301 [[arXiv:1209.0760](https://arxiv.org/abs/1209.0760)] [[INSPIRE](#)].
- [2] G. Dall’Agata, G. Inverso and A. Marrani, *Symplectic deformations of gauged maximal supergravity*, *JHEP* **07** (2014) 133 [[arXiv:1405.2437](https://arxiv.org/abs/1405.2437)] [[INSPIRE](#)].
- [3] G. Inverso, *Electric-magnetic deformations of $D = 4$ gauged supergravities*, *JHEP* **03** (2016) 138 [[arXiv:1512.04500](https://arxiv.org/abs/1512.04500)] [[INSPIRE](#)].
- [4] G. Dall’Agata and G. Inverso, *On the vacua of $N = 8$ gauged supergravity in 4 dimensions*, *Nucl. Phys. B* **859** (2012) 70 [[arXiv:1112.3345](https://arxiv.org/abs/1112.3345)] [[INSPIRE](#)].
- [5] A. Borghese, A. Guarino and D. Roest, *All G_2 invariant critical points of maximal supergravity*, *JHEP* **12** (2012) 108 [[arXiv:1209.3003](https://arxiv.org/abs/1209.3003)] [[INSPIRE](#)].
- [6] A. Borghese et al., *The $SU(3)$ -invariant sector of new maximal supergravity*, *JHEP* **03** (2013) 082 [[arXiv:1211.5335](https://arxiv.org/abs/1211.5335)] [[INSPIRE](#)].
- [7] A. Borghese, A. Guarino and D. Roest, *Triality, periodicity and stability of $SO(8)$ gauged supergravity*, *JHEP* **05** (2013) 107 [[arXiv:1302.6057](https://arxiv.org/abs/1302.6057)] [[INSPIRE](#)].
- [8] A. Gallerati, H. Samtleben and M. Trigiante, *The $\mathcal{N} > 2$ supersymmetric AdS vacua in maximal supergravity*, *JHEP* **12** (2014) 174 [[arXiv:1410.0711](https://arxiv.org/abs/1410.0711)] [[INSPIRE](#)].
- [9] B. de Wit and H. Nicolai, *$N = 8$ supergravity*, *Nucl. Phys. B* **208** (1982) 323 [[INSPIRE](#)].
- [10] A. Guarino and O. Varela, *Dyonic $ISO(7)$ supergravity and the duality hierarchy*, *JHEP* **02** (2016) 079 [[arXiv:1508.04432](https://arxiv.org/abs/1508.04432)] [[INSPIRE](#)].
- [11] E. Cremmer, B. Julia and J. Scherk, *Supergravity theory in eleven-dimensions*, *Phys. Lett.* **76B** (1978) 409 [[INSPIRE](#)].
- [12] B. de Wit and H. Nicolai, *The consistency of the S^7 truncation in $D = 11$ supergravity*, *Nucl. Phys. B* **281** (1987) 211 [[INSPIRE](#)].
- [13] L.J. Romans, *Massive $N = 2a$ supergravity in ten-dimensions*, *Phys. Lett.* **169B** (1986) 374 [[INSPIRE](#)].
- [14] A. Guarino, D.L. Jafferis and O. Varela, *String theory origin of dyonic $N = 8$ supergravity and its Chern-Simons duals*, *Phys. Rev. Lett.* **115** (2015) 091601 [[arXiv:1504.08009](https://arxiv.org/abs/1504.08009)] [[INSPIRE](#)].
- [15] A. Guarino and O. Varela, *Consistent $\mathcal{N} = 8$ truncation of massive IIA on S^6* , *JHEP* **12** (2015) 020 [[arXiv:1509.02526](https://arxiv.org/abs/1509.02526)] [[INSPIRE](#)].
- [16] J.H. Schwarz, *Covariant field equations of chiral $N = 2$ $D = 10$ supergravity*, *Nucl. Phys. B* **226** (1983) 269 [[INSPIRE](#)].

- [17] G. Inverso, H. Samtleben and M. Trigiante, *Type II supergravity origin of dyonic gaugings*, *Phys. Rev. D* **95** (2017) 066020 [[arXiv:1612.05123](#)] [[INSPIRE](#)].
- [18] G. Larios, P. Ntokos and O. Varela, *Embedding the SU(3) sector of SO(8) supergravity in $D = 11$* , *Phys. Rev. D* **100** (2019) 086021 [[arXiv:1907.02087](#)] [[INSPIRE](#)].
- [19] O. Varela, *AdS₄ solutions of massive IIA from dyonic ISO(7) supergravity*, *JHEP* **03** (2016) 071 [[arXiv:1509.07117](#)] [[INSPIRE](#)].
- [20] A. Guarino and C. Sterckx, *S-folds and (non-)supersymmetric Janus solutions*, *JHEP* **12** (2019) 113 [[arXiv:1907.04177](#)] [[INSPIRE](#)].
- [21] N.P. Warner, *Some new extrema of the scalar potential of gauged $N = 8$ supergravity*, *Phys. Lett.* **128B** (1983) 169 [[INSPIRE](#)].
- [22] Y. Pang, J. Rong and O. Varela, *Spectrum universality properties of holographic Chern-Simons theories*, *JHEP* **01** (2018) 061 [[arXiv:1711.07781](#)] [[INSPIRE](#)].
- [23] R. Corrado, K. Pilch and N.P. Warner, *An $N = 2$ supersymmetric membrane flow*, *Nucl. Phys. B* **629** (2002) 74 [[hep-th/0107220](#)] [[INSPIRE](#)].
- [24] C. Bachas and J. Estes, *Spin-2 spectrum of defect theories*, *JHEP* **06** (2011) 005 [[arXiv:1103.2800](#)] [[INSPIRE](#)].
- [25] N. Bobev, N. Halmagyi, K. Pilch and N.P. Warner, *Supergravity instabilities of non-supersymmetric quantum critical points*, *Class. Quant. Grav.* **27** (2010) 235013 [[arXiv:1006.2546](#)] [[INSPIRE](#)].
- [26] M. Trigiante, *Gauged supergravities*, *Phys. Rept.* **680** (2017) 1 [[arXiv:1609.09745](#)] [[INSPIRE](#)].
- [27] I.R. Klebanov, S.S. Pufu and F.D. Rocha, *The squashed, stretched and warped gets perturbed*, *JHEP* **06** (2009) 019 [[arXiv:0904.1009](#)] [[INSPIRE](#)].
- [28] J.-M. Richard, R. Terrisse and D. Tsimpis, *On the spin-2 Kaluza-Klein spectrum of $AdS_4 \times S^2(\mathcal{B}_4)$* , *JHEP* **12** (2014) 144 [[arXiv:1410.4669](#)] [[INSPIRE](#)].
- [29] A. Passias and A. Tomasiello, *Spin-2 spectrum of six-dimensional field theories*, *JHEP* **12** (2016) 050 [[arXiv:1604.04286](#)] [[INSPIRE](#)].
- [30] A. Passias and P. Richmond, *Perturbing $AdS_6 \times_w S^6$: linearised equations and spin-2 spectrum*, *JHEP* **07** (2018) 058 [[arXiv:1804.09728](#)] [[INSPIRE](#)].
- [31] M. Gutperle, C.F. Uhlemann and O. Varela, *Massive spin 2 excitations in $AdS_6 \times S^2$ warped spacetimes*, *JHEP* **07** (2018) 091 [[arXiv:1805.11914](#)] [[INSPIRE](#)].
- [32] K. Chen, M. Gutperle and C.F. Uhlemann, *Spin 2 operators in holographic $4d \mathcal{N} = 2$ SCFTs*, *JHEP* **06** (2019) 139 [[arXiv:1903.07109](#)] [[INSPIRE](#)].
- [33] S. Speziali, *Spin 2 fluctuations in 1/4 BPS AdS_3/CFT_2* , [arXiv:1910.14390](#) [[INSPIRE](#)].
- [34] D. Andriot and D. Tsimpis, *Gravitational waves in warped compactifications*, [arXiv:1911.01444](#) [[INSPIRE](#)].
- [35] P.G.O. Freund and M.A. Rubin, *Dynamics of dimensional reduction*, *Phys. Lett.* **97B** (1980) 233 [[INSPIRE](#)].
- [36] B. de Wit, H. Nicolai and N.P. Warner, *The embedding of gauged $N = 8$ supergravity into $d = 11$ supergravity*, *Nucl. Phys. B* **255** (1985) 29 [[INSPIRE](#)].

- [37] B. de Wit and H. Nicolai, *A new SO(7) invariant solution of $d = 11$ supergravity*, *Phys. Lett.* **148B** (1984) 60 [INSPIRE].
- [38] F. Englert, *Spontaneous compactification of eleven-dimensional supergravity*, *Phys. Lett.* **119B** (1982) 339 [INSPIRE].
- [39] C.N. Pope and N.P. Warner, *An SU(4) invariant compactification of $d = 11$ supergravity on a stretched seven sphere*, *Phys. Lett.* **150B** (1985) 352 [INSPIRE].
- [40] F. Englert and H. Nicolai, *Supergravity in eleven-dimensional space-time*, in the proceedings of *Group Theoretical Methods in Physics: 12th international colloquium*, September 5–11, Trieste, Italy (1983).
- [41] E. Sezgin, *The spectrum of the eleven-dimensional supergravity compactified on the round seven sphere*, *Phys. Lett.* **B 138** (1984) 57.
- [42] B. Biran et al., *The fluctuating seven sphere in eleven-dimensional supergravity*, *Phys. Lett.* **B 134** (1984) 179.
- [43] M.J. Duff, B.E.W. Nilsson and C.N. Pope, *Kaluza-Klein Supergravity*, *Phys. Rept.* **130** (1986) 1 [INSPIRE].
- [44] O. Varela, *Complete $D = 11$ embedding of SO(8) supergravity*, *Phys. Rev.* **D 97** (2018) 045010 [arXiv:1512.04943] [INSPIRE].
- [45] N. Bobev, F.F. Gautason, K. Pilch, M. Suh and J. Van Muiden, *Janus and J-fold solutions from Sasaki-Einstein Manifolds*, *Phys. Rev.* **D 100** (2019) 081901 [arXiv:1907.11132] [INSPIRE].
- [46] D. Bak, M. Gutperle and S. Hirano, *A Dilatonic deformation of AdS_5 and its field theory dual*, *JHEP* **05** (2003) 072 [hep-th/0304129] [INSPIRE].
- [47] E. D'Hoker, J. Estes and M. Gutperle, *Ten-dimensional supersymmetric Janus solutions*, *Nucl. Phys.* **B 757** (2006) 79 [hep-th/0603012] [INSPIRE].
- [48] E. D'Hoker, J. Estes and M. Gutperle, *Exact half-BPS type IIB interface solutions. I. Local solution and supersymmetric Janus*, *JHEP* **06** (2007) 021 [arXiv:0705.0022] [INSPIRE].
- [49] E. D'Hoker, J. Estes and M. Gutperle, *Exact half-BPS type IIB interface solutions. II. Flux solutions and multi-Janus*, *JHEP* **06** (2007) 022 [arXiv:0705.0024] [INSPIRE].
- [50] B. Assel and A. Tomasiello, *Holographic duals of 3d S-fold CFTs*, *JHEP* **06** (2018) 019 [arXiv:1804.06419] [INSPIRE].
- [51] B. Assel, C. Bachas, J. Estes and J. Gomis, *Holographic duals of $D = 3$ $N = 4$ superconformal field theories*, *JHEP* **08** (2011) 087 [arXiv:1106.4253] [INSPIRE].
- [52] B. Assel, C. Bachas, J. Estes and J. Gomis, *IIB duals of $D = 3$ $N = 4$ circular quivers*, *JHEP* **12** (2012) 044 [arXiv:1210.2590] [INSPIRE].
- [53] I. Garozzo, G. Lo Monaco and N. Mekareeya, *The moduli spaces of S-fold CFTs*, *JHEP* **01** (2019) 046 [arXiv:1810.12323] [INSPIRE].
- [54] I. Garozzo, G. Lo Monaco and N. Mekareeya, *Variations on S-fold CFTs*, *JHEP* **03** (2019) 171 [arXiv:1901.10493] [INSPIRE].
- [55] I. Garozzo, G. Lo Monaco, N. Mekareeya and M. Sacchi, *Supersymmetric Indices of 3d S-fold SCFTs*, *JHEP* **08** (2019) 008 [arXiv:1905.07183] [INSPIRE].

- [56] C. Cordova, T.T. Dumitrescu and K. Intriligator, *Multiplets of superconformal symmetry in diverse dimensions*, *JHEP* **03** (2019) 163 [[arXiv:1612.00809](#)] [[INSPIRE](#)].
- [57] J.P. Gauntlett and O. Varela, *Consistent Kaluza-Klein reductions for general supersymmetric AdS solutions*, *Phys. Rev. D* **76** (2007) 126007 [[arXiv:0707.2315](#)] [[INSPIRE](#)].
- [58] E. Malek, *Half-maximal supersymmetry from exceptional field theory*, *Fortsch. Phys.* **65** (2017) 1700061 [[arXiv:1707.00714](#)] [[INSPIRE](#)].
- [59] D. Cassani, G. Josse, M. Petrini and D. Waldram, *Systematics of consistent truncations from generalised geometry*, *JHEP* **11** (2019) 017 [[arXiv:1907.06730](#)] [[INSPIRE](#)].
- [60] B. de Wit, H. Samtleben and M. Trigiante, *The maximal $D = 4$ supergravities*, *JHEP* **06** (2007) 049 [[arXiv:0705.2101](#)] [[INSPIRE](#)].
- [61] A. Guarino, J. Tarrío and O. Varela, *Flowing to $\mathcal{N} = 3$ Chern-Simons-matter theory*, [arXiv:1910.06866](#) [[INSPIRE](#)].
- [62] E. Malek and H. Samtleben, *Ten-dimensional origin of Minkowski vacua in $N = 8$ supergravity*, *Phys. Lett. B* **776** (2018) 64 [[arXiv:1710.02163](#)] [[INSPIRE](#)].