

**International Ocean Discovery Program
Expedition 382 Preliminary Report
Iceberg Alley and Subantarctic Ice and Ocean Dynamics**

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382 Scientists

Publisher's notes

Core samples and the wider set of data from the science program covered in this report are under moratorium and accessible only to Science Party members until 20 May 2021.

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Abstract

International Ocean Discovery Program (IODP) Expedition 382, Iceberg Alley and Subantarctic Ice and Ocean Dynamics, investigated the long-term climate history of Antarctica, seeking to understand how polar ice sheets responded to changes in insolation and atmospheric CO₂ in the past and how ice sheet evolution influenced global sea level and vice versa. Five sites (U1534–U1538) were drilled east of the Drake Passage: two sites at 53.2°S at the northern edge of the Scotia Sea and three sites at 57.4°–59.4°S in the southern Scotia Sea. We recovered continuously deposited late Neogene sediment to reconstruct the past history and variability in Antarctic Ice Sheet (AIS) mass loss and associated changes in oceanic and atmospheric circulation.

The sites from the southern Scotia Sea (Sites U1536–U1538) will be used to study the Neogene flux of icebergs through “Iceberg Alley,” the main pathway along which icebergs calved from the margin of the AIS travel as they move equatorward into the warmer waters of the Antarctic Circumpolar Current (ACC). In particular, sediments from this area will allow us to assess the magnitude of iceberg flux during key times of AIS evolution, including the following:

- The middle Miocene glacial intensification of the East Antarctic Ice Sheet,
- The mid-Pliocene warm period,
- The late Pliocene glacial expansion of the West Antarctic Ice Sheet,
- The mid-Pleistocene transition (MPT), and
- The “warm interglacials” and glacial terminations of the last 800 ky.

We will use the geochemical provenance of iceberg-rafted detritus and other glacially eroded material to determine regional sources of AIS mass loss. We will also address interhemispheric phasing of ice sheet growth and decay, study the distribution and history of land-based versus marine-based ice sheets around the continent over time, and explore the links between AIS variability and global sea level.

By comparing north–south variations across the Scotia Sea between the Pirie Basin (Site U1538) and the Dove Basin (Sites U1536 and U1537), Expedition 382 will also deliver critical information on how climate changes in the Southern Ocean affect ocean circulation through the Drake Passage, meridional overturning in the region, water mass production, ocean–atmosphere CO₂ transfer by wind-induced upwelling, sea ice variability, bottom water outflow from the Weddell Sea, Antarctic weathering inputs, and changes in oceanic and atmospheric fronts in the vicinity of the ACC.

Comparing changes in dust proxy records between the Scotia Sea and Antarctic ice cores will also provide a detailed reconstruction of changes in the Southern Hemisphere westerlies on millennial and orbital timescales for the last 800 ky. Extending the ocean dust record beyond the last 800 ky will help to evaluate dust–climate couplings since the Pliocene, the potential role of dust in iron fertilization and atmospheric CO₂ drawdown during glacials, and whether dust input to Antarctica played a role in the MPT.

The principal scientific objective of Subantarctic Front Sites U1534 and U1535 at the northern limit of the Scotia Sea is to reconstruct and understand how ocean circulation and intermediate water formation responds to changes in climate with a special focus on the connectivity between the Atlantic and Pacific basins, the “cold water route.” The Subantarctic Front contourite drift, deposited be-

tween 400 and 2000 m water depth on the northern flank of an east–west trending trough off the Chilean continental shelf, is ideally situated to monitor millennial- to orbital-scale variability in the export of Antarctic Intermediate Water beneath the Subantarctic Front. During Expedition 382, we recovered continuously deposited sediments from this drift spanning the late Pleistocene (from ~0.78 Ma to recent) and from the late Pliocene (~3.1–2.6 Ma). These sites are expected to yield a wide array of paleoceanographic records that can be used to interpret past changes in the density structure of the Atlantic sector of the Southern Ocean, track migrations of the Subantarctic Front, and give insights into the role and evolution of the cold water route over significant climate episodes, including the following:

- The most recent warm interglacials of the late Pleistocene and
- The intensification of Northern Hemisphere glaciation.

Introduction

Several decades of research on the paleoceanography and paleoclimatology of the Scotia Sea using short cores and remote-sensing data have revealed a wealth of information on late Pleistocene ice sheet–ocean climate interactions. However, only deep drilling will enable us to understand the evolution and sensitivity of these interactions over longer timescales, in particular during periods of Earth’s history that were significantly warmer and characterized by higher sea levels than today. Drilling deep sites in the Pirie and Dove Basins allows us to decipher the geological history of the Antarctic region back to the Miocene and to test competing hypotheses in this important yet undersampled region. In particular, recovery of long records will improve our understanding of how the coupled Antarctic ice–ocean atmosphere system evolved to the present state and how it behaved across major climate transitions and during warmer-than-present times. We will also test how Seismic Reflectors a–c, which divide five main regional units, relate to key changes in Antarctic Ice Sheet (AIS) history, oceanic currents, and/or climate evolution. These changes could include the middle Miocene glacial intensification of the East Antarctic Ice Sheet (EAIS), the mid-Pliocene warm interval, the late Pliocene glacial enhancement of the West Antarctic Ice Sheet (WAIS), the mid-Pleistocene transition (MPT), or more recent glacial–interglacial changes. Using a variety of proxies and modeling approaches, the sediments recovered during International Ocean Discovery Program (IODP) Expedition 382 will be used to address key topics discussed in the following sections.

Past AIS behavior

The Scotia Sea sites are strategically located in two areas of “Iceberg Alley,” the major Antarctic gateway to lower latitudes where Antarctic icebergs generally exit to the Southern Ocean (Figure F1). Weber et al. (2014) revealed that times of enhanced iceberg-rafted debris (IBRD) recorded in Scotia Sea sediments correspond to times of AIS instability. Accordingly, with material collected during Expedition 382 we should be able to generate the first spatially integrated and long-term reconstruction of AIS variability and its relation to sea level and climate change. A recent proof-of-concept study revealed that, at least for the late Pleistocene, sedimentation rates at these sites are high enough to record climate and ice sheet changes on decadal-to-centennial timescales. The discovery that massive episodes of iceberg discharge begin abruptly, ramping up from background levels to peak values within decades (Weber et al.,

2014), has fundamental implications for projections of future AIS behavior in a warming world. In addition, IBRD provenance is able to discriminate circum-Antarctic events from regional AIS events, and sea level, ice sheet, and iceberg modeling will be used to improve our understanding of AIS dynamics, especially during warmer-than-present periods and glacial terminations. Also, expanded sea ice could enhance thermal isolation and AIS expansion and vice versa.

Regional atmospheric circulation and dust transport

Our two drilling areas in the Scotia Sea are located to monitor the history of prevailing Southern Hemisphere westerlies (SHW), located between Patagonian dust sources to the north and Antarctic ice-core dust records to the south (e.g., EPICA Community Members, 2006) (Figure F1). We will reconstruct dust transport history in the Pirie and Dove Basins with attendant implications for the SHW in more northerly (proximal; e.g., Sites U1534 and U1535) and southerly (distal; e.g., Antarctic ice cores) positions relative to the dust source. We will also investigate whether marine and ice-core dust proxies are coupled over the last 800 ky similar to the last glacial cycle (Weber et al., 2014). Such coupling will provide chronostratigraphic control and allow us to infer whether and how dust covaried with orbital cycles prior to the late Pleistocene.

Drake Passage throughflow and deepwater history

The Scotia Sea sites, complemented by the Subantarctic Front sites, are located to reconstruct the variability in the strength of the Drake Passage throughflow immediately east of this major geographic bottleneck constraining the flow of the Antarctic Circumpolar Current (ACC) (McCave et al., 2014). Using proxy records, we should be able to identify times of stronger ACC influence (higher proportions of Circumpolar Deep Water [CDW]) from times of stronger Weddell Sea influence (higher proportions of Weddell Sea Deep Water [WSDW]) (Figure F1), specifically at the southern Scotia Sea sites. Finally, because the sites stretch across major latitudinal ocean and atmosphere fronts (Figure F1), we should be able to reconstruct past changes in ocean productivity, sea ice extent, and regional wind patterns for the key periods described above, including late Quaternary glacial–interglacial and millennial-scale cycles.

Background

Geological setting

The formation of the Scotia Sea is linked to the opening of the Drake Passage since the Eocene. The Drake Passage is proposed to be the final stage in the Gondwana breakup, resulting in the onset of the ACC and the initiation of the thermal isolation of the Antarctic continent (e.g., Lawver and Gahagan, 2003). The Scotia Sea is underlain by two main tectonic plates, Scotia and Sandwich (e.g., Dalziel et al., 2013; Eagles and Jokat, 2014). In the Scotia plate, a series of small basins have developed that are bounded by the South Scotia Ridge on their southern part (e.g., Maldonado et al., 2006; Figure F1).

Sediments located in the Scotia Sea and its southern basins have recorded the history and evolution of the tectonics and oceanic currents in the area (e.g., Martos et al., 2013; Pérez et al., 2017). Five main seismic units (Figures F2, F3) separated by distinct stratigraphic discontinuities can be regionally traced among the small southern basins and north of the Weddell Sea (Maldonado et al., 2006). The ages of these regional units have not been established to

date; however, interpretation of the sedimentary record in the southern Scotia Sea basins (Pérez et al., 2017) and long-distance correlation of seismic records to the Weddell Sea continental margin (Lindeque et al., 2013) indicate an early Pleistocene age near the base of the top unit and a mid-Miocene age near the bottom of the lowermost unit. The seismic reflections are mostly subparallel in the three upper units, although internal transparent or chaotic facies indicate localized disturbances and downslope deposition (e.g., Pérez et al., 2016; Ruano et al., 2014; Martos et al., 2013), a general phenomenon present at all Expedition 382 drill sites.

In addition to tectonically induced downslope transport, the sedimentary record in the southern Scotia Sea basins is shaped by bottom currents along the pathway of the ACC and WSDW (e.g., Orsi et al., 1999; Naveira Garabato et al., 2002). Contourite deposits dominate the southern Scotia Sea, especially in the upper three units or from Reflector c upward (e.g., Maldonado et al., 2006; Martos et al., 2013; Pérez et al., 2014b, 2017). Previous studies have proposed that deposition in the southern Scotia Sea basins was influenced by the action of the ACC in the lower units and by the action of the WSDW flows in the upper three units (e.g., Martos et al., 2013; Pérez et al., 2014a). Reflector c (Maldonado et al., 2006) is regionally identified in the Scotia and Weddell Seas and shows a northward progradational pattern (Lindeque et al., 2013). Tectonic movements during the middle Miocene opened the Jane Basin, a gateway in the South Scotia Ridge that allowed northward flow of WSDW (Maldonado et al., 2006) and also initiated the uplift of the Shackleton Fracture Zone (Martos et al., 2013). These tectonic events may have forced the ACC to shift north, thereby contributing to the thermal isolation of Antarctica (Martos et al., 2013). These changes could also be the cause of EAIS expansion and establishment of full-glacial conditions in both the Scotia and Weddell Seas (Lindeque et al., 2013). Unit I, a contourite drift with basal Reflector a, represents the youngest regional sedimentary unit and was possibly deposited in concert with the intensification of Northern Hemisphere glaciation and/or postulated WAIS expansion (McKay et al., 2012). All of these hypotheses will be tested with material recovered during this expedition.

For Dove Basin, Galindo-Zaldívar et al. (2014) provide estimated ages for the basement of 24–21 Ma (late Oligocene to early Miocene) constrained by spreading magnetic anomaly models and $^{40}\text{Ar}/^{39}\text{Ar}$ dating from dredge samples from the spreading ridge. For Pirie Basin, estimated basement ages range from 43 (Barker et al., 2013) to 14 Ma (Galindo-Zaldívar et al., 2006) based on geothermal heat flow measurements and spreading magnetic anomalies, respectively.

Recent seismic interpretations for the Dove (Pérez et al., 2017) and Pirie Basins (Pérez et al., 2014b) added Reflector a' in Unit I. For Dove Basin, three seismic lines indicate a basin-like structure with several small-scale ridges and continuous reflectors in the central part (Figure F1). Site U1536 (proposed Site SCO-13) is located along Seismic Line 10/04, which runs to the deepest part of the basin in the north. This site provides the best compromise between sediment thickness, lack of disturbance, and acoustic basement at a relatively shallow depth (~900 meters below seafloor [mbsf]) (Figure F3). Site U1537 (proposed Site SCO-18) is located in the north of Dove Basin along Seismic Line 07/04 (Figure F3).

As in Dove Basin, drilling in Pirie Basin recovered structurally intact and thick Pliocene–Pleistocene deposits. Here, our aim was to (1) core Seismic Units I–III and (2) obtain a paleoclimate record from a northerly position relative to oceanic and atmospheric fronts that likely shifted during both the shorter term climate cycles and

the longer term climatic evolution of the region. East–west running Seismic Lines 13/13 and 11/08 indicate a north–south oriented basin. Site U1538 (proposed Site SCO-11) shows very thick and planar strata of Units I and II in the center of the small basin to the north.

Topographic parametric sonar (TOPAS) was used along profile Lines 2004_SCAN_TOPAS_L10C and 2004_SCAN_TOPAS_L07F in Dove Basin (Sites U1536 and U1537) to allow a high-resolution characterization of the two dominant sedimentary features. First, disturbed sedimentary sequences likely indicate mass transport deposits (MTDs) caused by continued tectonic activity (Ruano et al., 2014) (Figure F4). Second, interbedded layered deposits form mounds with wavy morphologies and elongated erosional features that were likely formed by bottom currents (Ruano et al., 2014).

In Pirie Basin, TOPAS profile Line 2013_SCAN_TOPAS_L13 over Site U1538 does not show evidence of MTDs, but it displays the same mounded and wavy morphologies that suggest oceanographic control of the sedimentation, which is in agreement with the interpretations from high-resolution studies in neighboring basins (Dove and Scan Basins; Ruano et al., 2014; García et al., 2016). A comparison of all three Scotia Sea sites is shown in Figure F4.

During Expedition 382, we also cored two sites in the Subantarctic Front, Sites U1534 and U1535 (proposed Sites SFSD-03A and SFSD-02A, respectively). The sediments targeted are part of a regional contourite drift largely composed of clay with silty horizons. Interpretation of seismic profiles suggests that the Subantarctic Front drift comprises four sediment packages that exhibit a parallel to subparallel stratigraphy with internal unconformities and that overlie a prominent unconformity representing the base of the drift (Figure F2). Koenitz et al. (2008) suggest that these four units were likely deposited under different oceanographic and climatic boundary conditions that occurred since the deep opening of Drake Passage and its resulting onset of deep ACC flow in the early Miocene.

Scientific objectives

1. Reconstructing the history of AIS growth and decay

Marine records of IBRD provide a nearly continuous history of variability in ice sheet mass loss. The majority of large icebergs (>5 km long) that calve from the AIS margin (Stuart and Long, 2011) are routed through Iceberg Alley after traveling counterclockwise around Antarctica in the ACC (Figure F1). Melt rates increase when icebergs reach the warmer ACC in our study area (Silva et al., 2006). Because icebergs currently account for ~50% of the total AIS mass loss (Gladstone et al., 2001), the Scotia Sea IBRD record will, for the first time, capture a spatially integrated signal that is representative of variability in the flux of icebergs released by the AIS over million-year timescales.

IBRD records of the last deglaciation demonstrate the outstanding opportunity for developing well-dated, long records of AIS discharge history. Based on X-radiograph counts, Weber et al. (2014) identified eight events of increased AIS discharge (AID1–AID8) that occurred between 19 and 9 ka. These records provide the first robust evidence for centennial- to millennial-scale variability in deglacial mass loss from the AIS. Specifically, AID6 (~15–13.9 ka) has the largest IBRD flux, peaking at 14.76–14.4 ka, synchronous with the timing (14.65–14.31 ka) and duration (340 y) of Meltwater Pulse 1A (MWP-1A) (Deschamps et al., 2012). It provides the first direct evidence for an Antarctic contribution to MWP-1A.

This interpretation is supported by novel modeling and ice-core studies (Fogwill et al., 2017; Gollledge et al., 2014) but is in marked contrast to reconstructions arguing for a minor and later (after ~14

ka) AIS retreat (e.g., Bentley et al., 2010; Conway et al., 1999; Mackintosh et al., 2011, 2014; Peltier, 2004). However, those studies rely on temporally and geographically restricted terrestrial and shallow-marine, near-field sequences (Anderson et al., 2014; Heroy and Anderson, 2007; Hillenbrand et al., 2013).

The contribution of the AIS to past sea level change remains poorly constrained, yet it is important for assessing the sensitivity of the AIS to climate change and its contribution to future sea level rise (Church et al., 2013). Thus, large uncertainties remain in our understanding of the interactions between the AIS, Southern and Northern Hemisphere climate, and global sea level. Studying the integrated variability in AIS volume change through critical periods of Earth's history has the potential to substantially advance our knowledge on several fronts. Even for the mid-to-late Holocene, a time of relative AIS stability, the IBRD record from Iceberg Alley reveals substantial decadal- to centennial-scale variations in ice mass loss that have far-field effects and significant implications for future AIS development (Bakker et al., 2016).

With the highly resolved and continuous sediment record obtained during Expedition 382, we should be able to address important questions about past AIS discharge history and associated changes in climate and sea level, related to the following:

- The mid-Miocene EAIS glacial intensification, which is the widespread discontinuity believed to mark the onset of WSDW flow into the Scotia Sea (Maldonado et al., 2006; Martos et al., 2013; Pérez et al., 2017) and the transition to full-glacial conditions in the Weddell Sea (Lindeque et al., 2013);
- AIS expansion and climate cooling during the late Pliocene as reconstructed for the WAIS (McKay et al., 2012);
- Potentially significant WAIS volume changes during mid-Pliocene warm intervals (Naish et al., 2009);
- The size of the AIS during warm periods (e.g., the mid-Miocene climate optimum and the mid-Pliocene warming); and
- The AIS contribution to sea level highstands during previous warm periods (Masson-Delmotte et al., 2013).

Is the AIS mass loss history involved in the MPT, and if so, will we observe abrupt or gradual changes in IBRD flux? How similar are previous glacial terminations to the last deglaciation? Does the more northerly located Pirie Basin contain a higher percentage of ice-rafted debris (IRD) from WAIS icebergs that escaped the generally counterclockwise circulation around Antarctica and instead were routed through the Drake Passage? Can we link the delivery of icebergs to Milankovitch orbital cycles, and if so, how do these links evolve with time?

X-radiographs obtained on board and grain size measurements that will be obtained postcruise will be used to identify periods of enhanced AIS mass loss throughout the Expedition 382 records. To simulate dynamics and changes in ice volume associated with circum-AIS instabilities, ice sheet climate models will also be utilized postcruise (e.g., Levermann et al., 2012; Winkelmann et al., 2012).

2. Chronology and dust-climate couplings

We will follow the approach of Cody et al. (2008) to establish a coherent low-resolution chronology by combining diatom biostratigraphy, magnetostratigraphy, and tephrostratigraphy. Also, oxygen isotopes ($\delta^{18}\text{O}$) on biogenic silica could help constrain the timing and related changes in global ice volume and deep-ocean circulation. Sediment physical and optical properties as well as X-ray fluorescence (XRF) scanner data will be used for orbital tuning and identification of tephra layers, followed by microscopic identifica-

tion and $^{40}\text{Ar}/^{39}\text{Ar}$ dating of the tephras. Prominent volcanic sources of ash layers are located to the west (e.g., Deception Island; Moreton and Smellie, 1998).

In addition to magnetostratigraphic dating using the magnetic reversal history (Stoner et al., 2003, 2002), relative paleointensity offers the potential for suborbital magnetic stratigraphy between polarity boundaries (Channell et al., 2009; Stoner et al., 2000; Valet, 2003; Valet et al., 2005; Ziegler et al., 2011). Calibration of paleointensity with the dust/ice-core chronology could potentially be extrapolated to the sedimentary record beyond the Southern Ocean (Channell et al., 2016). A better understanding of magnetic susceptibility should be derivable through particle size specific (magnetic) measurements (Hatfield and Stoner, 2013; Hatfield et al., 2013).

Correlating dust proxies such as magnetic susceptibility, Ca, and Fe with the Antarctic ice-core dust flux should provide high-resolution chronological control (Weber et al., 2012) at least for the late Pleistocene and allow us to reconstruct atmospheric circulation. The correlation of marine dust records to well-dated ice-core dust records has been a major step forward in developing Southern Ocean chronologies for the subantarctic Pacific (Lamy et al., 2014), subantarctic Atlantic (Martínez-García et al., 2011, 2014), and Scotia Sea (Weber et al., 2012, 2014). This correlation is possible because dust deposition appears to be coherent and synchronous across much of the Southern Ocean and the AIS. Age models developed for Cores MD07-3133 and MD07-3134 followed this correlation strategy by synchronizing marine (magnetic susceptibility and XRF-based counts of Ca and Fe) to ice-core dust proxies (e.g., non-sea salt Ca fluxes) (McCave et al., 2014; Pugh et al., 2009; Weber et al., 2012).

We will correlate marine and ice core dust proxies for the past 800 ky (the maximum age of ice-core records) and evaluate the robustness of the strategy used by Martínez-García et al. (2011), who correlated ocean dust proxies with the LR04 stack (Lisiecki and Raymo, 2005) back to the base of the Pleistocene.

Martínez-García et al. (2011) suggested that dust plays a major role in intensification of Northern Hemisphere glaciation since the mid-Pliocene, with a distinct and abrupt increase in dust flux at the start of the MPT at 1.25 Ma. They posited an important role for dust in the subsequent intensification of glaciations (see also Rodríguez-Sanza et al., 2012). In contrast, Elderfield et al. (2012) inferred an abrupt increase in AIS volume at ~900 ka attributed to the onset of the MPT. Finally, Raymo et al. (2006) proposed that the MPT was associated with the establishment of largely marine-based ice margins around Antarctica. In any of these scenarios, drilling the Scotia Sea sites provides a unique opportunity to investigate iceberg-dust-climate interactions and their possible role in AIS behavior, as well as to help evaluate the role of dust in the evolution of Pliocene–Pleistocene glacial cycles, atmospheric CO_2 , and iron fertilization (Martínez-García et al., 2011, 2014).

Relative paleointensity and magnetic reversal determinations will be used in tandem to obtain further age control for the younger sections. Paleointensity has proven to work in the Southern Hemisphere (Channell et al., 2000; Mazaud et al., 2002; Stoner et al., 2002, 2003) and on the Antarctic margin (Guyodo et al., 2001; Sagnotti et al., 2001). The paleointensity record is well correlated with the LR04 stack for at least the last 2.2 My (NARPI-2200 [Channell et al., 2016], PISO-1500 [Channell et al., 2009], SINT-2000 [Valet et al., 2005], and PDAM2M [Ziegler et al., 2011]) and to 3 Ma (EPAPIS [Yamazaki and Oda, 2005]) based on tuning of rock magnetic variables to LR04 and other $\delta^{18}\text{O}$ records (e.g., Tian et al., 2002).

3. AIS mass loss and regional-to-global sea level change

Sea level changes originating from the AIS show a complex spatial variability that is sensitive to the geometry and timescale of mass loss. For example, a rapid mass loss causes local sea level fall adjacent to the former ice source, a fall that can be an order of magnitude greater than the associated globally averaged (eustatic) sea level rise. At progressively greater distances from the region of mass loss, the amplitude of sea level rise increases, peaking at ~15% above eustatic in the far field. Thus, any rapid AIS mass loss would drive a local, large-amplitude sea level fall that may stabilize the ice sheet to further retreat (Gomez et al., 2010, 2013). In contrast, rapid mass loss from Northern Hemisphere ice sheets will lead to a sea level rise adjacent to the AIS, with a strong geographic gradient and maximum amplitudes in the Atlantic sector. This far-field sea level rise will impact AIS stability, and the gradient will make the Atlantic sector particularly susceptible to collapse (Clark et al., 2002; Weber et al., 2011).

The peak IBRD flux of AID6 corresponds to the time of MWP-1A, with a six-fold increase in flux relative to the Holocene steady-state average (Weber et al., 2014). This peak flux suggests that the IBRD record of Iceberg Alley may provide a proxy for the relative amount of mass released by the AIS and its associated contribution to global sea level rise for previous warm times in the late Neogene. Through development of long IBRD records from Iceberg Alley, we will evaluate whether meltwater pulses occurred and estimate their possible contribution to global sea level as well as the possible role of regional sea level in modulating AIS mass loss.

For past warm periods detected at all sites (e.g., Marine Isotope Stages [MIS] 5.5, 11, and 31 and potentially others farther back in time), it would be of specific interest to identify the source and relative timing of AIS mass loss. For instance, was there a late sea level highstand during MIS 5.5 (O'Leary et al., 2013)? Is there physical evidence for an AIS trigger? Investigations will be accompanied by deglacial simulations for iceberg transport around Antarctica, focusing on the last two deglaciations (e.g., Siddall et al., 2006) and warmer-than-present periods to better understand high-latitude climate processes and AIS dynamics and their impact on sea level.

To explore the impact of ice mass changes on sea level in the near and far field, we will adopt a state-of-the-art sea level model that incorporates all deformational, gravitational, and rotational effects of glacial isostatic adjustment (Kendall et al., 2005). This model also takes account of shoreline migration associated with local onlap/offlap of water and changes in the extent of grounded, marine-based ice. Sensitivity studies will be performed using Maxwell viscoelastic Earth models in which Earth's structure varies with depth through the lithosphere. We will also perform calculations incorporating lateral variations in viscosity structure using a finite element numerical model (Latychev et al., 2005). This model may be particularly important given seismic inferences of large differences in mantle structure below the WAIS and EAIS (Danesi and Morelli, 2001) and their impact on dynamic topography (Austermann et al., 2015).

4. Water mass changes and ocean thermal forcing

The Southern Ocean plays an important role in the Atlantic Meridional Overturning Circulation (AMOC) because it controls, through upwelling, the rate of heat and carbon exchange with the surface (Marshall and Speer, 2012), which is a major driver of climate change. This interhemispheric teleconnection shows that a reduction in the AMOC and attendant decrease in northward heat transport leads to warming of the Southern Hemisphere and to

changes in the relative contributions of North Atlantic Deep Water (NADW) to Antarctic Bottom Water (AABW) and WSDW, as well as the amount of CDW that is formed through mixing. Our goal is to study the changes in water mass properties that lead to atmospheric and oceanic reorganizations and associated changes in wind fields, sea ice, marine productivity, atmospheric CO₂, surface ocean and atmosphere temperatures, and AIS mass loss.

The radiogenic isotope composition of authigenic neodymium (ϵ_{Nd}) provides important information on past changes in deepwater mixing (e.g., Gutjahr et al., 2008; Piotrowski et al., 2005, 2008; Goldstein and Hemming, 2003; van de Flierdt et al., 2006). The isotopic signal of seawater-derived ϵ_{Nd} will be separated from other sedimentary components following the method of Blaser et al. (2016), seawater-derived Nd will be extracted from fish teeth following the method of Scher and Martin (2004), or the method of Roberts et al. (2010, 2012) will be used for foraminifer-rich sediments. Where possible, past seawater Nd isotope compositions will also be extracted from discrete ferromanganese nodules or carbonate constituents to ensure their seawater origin. The Nd isotope composition in the modern water column is well resolved in the Drake Passage (Stichel et al., 2012; Struve et al., 2017) and the Pacific (Carter et al., 2012; Rickli et al., 2014) and Atlantic (Stichel et al., 2012) sectors of the Southern Ocean.

Based on Piotrowski et al. (2004), Robinson and van de Flierdt (2009), and Skinner et al. (2013), CDW was very likely much more radiogenic (higher ϵ_{Nd}) during glacials because of lower admixing of NADW relative to CDW. In combination with modeling efforts, we will (1) assess past changes in water mass composition using authigenic neodymium isotopes to test the ocean thermal forcing hypothesis for previous terminations (e.g., Böhm et al., 2015; Howe and Piotrowski, 2017) and (2) test whether changes in NADW, CDW, and WSDW are associated with the formation of Reflectors a, b, and c (Pérez et al., 2017).

Given the importance of ocean thermal forcing to the dynamics of AIS marine margins and grounding lines, we will also evaluate the relationship between sea-surface temperature (SST) and IBRD flux. Because the alkenone unsaturation index (U^k_{37}) is likely not applicable (Shevenell et al., 2011), we will extract isoprenoid and branched glycerol diether glycerol tetraether (GDGT) lipids to measure the TEX₈₆ ratio (Schouten et al., 2002) and convert it to SST using published calibrations (e.g., Kim et al., 2008).

5. Provenance studies

Identifying the regional sources of individual IBRD events is crucial to assess past AIS evolution. In particular, does the IBRD signal at any given time reflect an ice sheet-wide contribution, or were different iceberg discharges associated with specific sectors of the AIS? The diverse geology of the Weddell embayment and East Antarctica provides distinct isotope geochemical sources that will be distinguished in IBRD and the fine sediment fraction. For example, the $^{40}\text{Ar}/^{39}\text{Ar}$ age of a single hornblende (or mica) IBRD grain allows it to be traced to a source on the continent with a similar $^{40}\text{Ar}/^{39}\text{Ar}$ age, and a population of ~30 hornblende grains should show which sources are most actively exporting debris in icebergs, which is information critical to reconstructing the past history of the AIS (e.g., Williams et al., 2010; Pierce et al., 2014). Similarly, Pb-Pb and $^{40}\text{Ar}/^{39}\text{Ar}$ in feldspars (e.g., Flowerdew et al., 2013), U-Pb in zircons (e.g., Pierce et al., 2014), and other thermochronometers (e.g., reviewed in Licht and Hemming, 2017) can distinguish input from different source areas. Combining these mineral tracers should lead to more accurate identification of provenance and hence the spatial history of AIS glaciation.

The provenance of glacially eroded fine sediment will also be assessed using detrital radiogenic isotope systems. For example, Sm-Nd isotopes, which reflect the “crustal extraction” bedrock age, have been widely used for AIS provenance studies (e.g., Roy et al., 2007). Recent studies on the Wilkes Land margin also indicate extensive EAIS retreat during Pliocene sea level highstands (Cook et al., 2013). Overall, East and West Antarctic bedrock have distinctly different Nd isotopic compositions, so Sm-Nd isotope results should let us gauge the balance of sediment input between East and West Antarctica. Although some overlap occurs between the radiogenic isotopes from East and West Antarctica (Roy et al., 2007), broadly there is considerably older crust (and thus lower ϵ_{Nd} and ϵ_{Hf} and higher $^{87}\text{Sr}/^{86}\text{Sr}$) from East Antarctic sources. In some cases, the provenance information from sand and fine fractions can provide contrasting evidence (e.g., Cook et al., 2017; Licht and Hemming, 2017), so it is important to examine both. Other geochemical tracers, especially radiogenic isotopes systems such as Rb-Sr, U-Th-Pb, Lu-Hf, and K-Ar are likely to provide greater clarity in the assessment of mixing from difference sources.

Finally, provenance studies may help to resolve a longstanding issue by constraining the source of the high sea level during MIS 5, which was 5–9 m higher than today (Dutton and Lambeck, 2012; Kopp et al., 2009). Greenland contributed from 1 to 4 m (NEEM Community Members, 2013), implying an additional contribution from the AIS was necessary (Masson-Delmotte et al., 2013). Although the WAIS is commonly invoked as the most likely source, no data support this assumption, and additional sources from the EAIS should also be considered, particularly with the growing recognition that it also has substantial areas of marine-based ice that may play an increasing role in future AIS mass loss and global sea level scenarios.

6. Glacial–interglacial sea ice shifts and the carbon cycle

Changes in sea ice extent influences the flux of heat and carbon between the deep ocean and the atmosphere (Schmitt et al., 2012) and the degree of primary production (Anderson et al., 2009). Sea ice is recognized as an important amplifier of high-latitude climate anomalies, and of particular interest is the role of sea ice variability on glacial–interglacial climate. Expanded Antarctic sea ice during glacials increases export of AABW/WSDW and limits mixing and ventilation of the deep ocean, possibly causing ~20%–60% of the glacial–interglacial CO₂ change according to carbon cycle models (Ferrari et al., 2014; Sigman et al., 2010; Stephens and Keeling, 2000). However, there are few proxy constraints to test these models, and large uncertainties remain. Little is known about the areal extent and duration of summer and winter sea ice during previous glacials and interglacials. Similarly, the primary controls on changing sea ice distribution and associated feedbacks are also poorly constrained. In combination with modeling efforts, seasonal sea ice changes will be investigated over multiple glacial cycles to help shed light on the processes that cause changes in ocean circulation, ventilation, and CO₂ sequestration.

We hypothesize that changes in Antarctic winter and summer sea ice extent are driven by migration of the SHW and fluxes of glacial meltwater to the surface ocean. The sea salt Na⁺ flux record of the European Project for Ice Coring in Antarctica (EPICA) Dronning Maud Land (EDML) ice core suggests that sea ice extent underwent large fluctuations during the last deglaciation (Wolff et al., 2006), which has been simulated by modeling studies (Menviel et al., 2011). Similarly, models suggest that ocean thermal forcing during the last deglaciation induced increased freshwater at the onset of MWP-1A, causing consecutive sea ice expansion during the

Antarctic Cold Reversal and retreat thereafter at the beginning of the Younger Dryas (Weber et al., 2014). Based on findings for the last glacial (Allen et al., 2011; Collins et al., 2013), we hypothesize that previous glacial expansions of summer and winter sea ice in the Scotia Sea will reach their maximum extents at $\sim 59^\circ$ and $\sim 55^\circ$ S, respectively.

Fogwill et al. (2014) modeled a 9° poleward shift of the SHW and a 0.2° – 0.9° C warming in the upper 500 m of the subpolar gyres that persisted for millennia during deglaciation. Associated AIS draw-down might also have induced large-scale sea ice expansion and IBRD increase (Fogwill et al., 2017). Reconstructing these Antarctic sea ice conditions and dust-climate couplings will provide much needed empirical data to assess whether changes in sea ice cover are driven by the position of the southern SHW and/or flux of glacial meltwater.

The Scotia Sea core sites are ideally located to record past variations in both summer and winter sea ice boundaries. To determine the latter, we will use diatom assemblages (Allen et al., 2011) and a new biomarker-based method that uses highly branched isoprenoids (HBIs) (Collins et al., 2013). The biomarker IP₂₅, a mono-unsaturated HBI biomarker derived from sea ice diatoms, has been used to reconstruct sea ice conditions in Arctic regions for Quaternary to Pliocene timescales (Belt et al., 2007; Belt and Müller, 2013; Müller et al., 2009; Stein and Fahl, 2013). HBIs have been used as a corresponding proxy for seasonal sea ice in the Southern Ocean (Massé et al., 2011). Importantly, the source organism of HBIs found in Antarctica, the diatom *Berkeleya adeliensis*, has now been identified (Belt et al., 2016). *B. adeliensis* is endemic to Antarctic sea ice only and the HBI it produces has been named biomarker IPSO₂₅. Because *B. adeliensis* flourishes in platelet ice, its sedimentary abundance is sensitive to nearby glacial or ice shelf settings, likely explaining its occurrence in glacial Scotia Sea sediment (Collins et al., 2013). Also, Belt et al. (2015) and Smik et al. (2016) demonstrated enhanced production of a different HBI biomarker within a retreating ice edge environment. Combined analysis of IPSO₂₅ with other HBIs should therefore provide insights into seasonal sea ice dynamics.

Analysis of sea ice diatoms (*Fragilariopsis curta*, *Fragilariopsis cylindrus*, and *Fragilariopsis obliquecostata*) (Gersonde and Zielinski, 2000) has also been successfully used to reconstruct Pleistocene sea ice extent in the Southern Ocean (Allen et al., 2011; Gersonde et al., 2005), and the same approach will be implemented here to complement the HBI analysis. In addition, novel sedimentary ancient DNA (sedaDNA) analyses (Armbrecht et al., 2019) will be used to complement both diatom and HBI data to achieve a broader picture of sea ice-associated paleocommunities.

These methods (IPSO₂₅ and sea ice diatom analysis; depending on preservation also sedaDNA) will be applied to piston cores along the north–south transect of the southern Scotia Sea sites to determine the presence/absence of permanent and/or seasonal sea ice, the position of the marginal ice zone, and associated wind mixing and ocean stratification for the last 60 ky. We will apply these same methods to older intervals. To further strengthen our investigation of the silica inventory, we will also include $\delta^{30}\text{Si}$ and $\delta^{18}\text{O}$ studies on diatoms.

Reconstructions of sea ice and paleotemperature will be accompanied by transient climate model simulations (e.g., Timmermann et al., 2014). To study Southern Hemisphere climate changes throughout past glacial–interglacial cycles, the global earth system model for studying the long-term evolution of climate (LOVE-

CLIM) (Menviel et al., 2008, 2010; Timmermann et al., 2009) will be used, based on a dynamical atmosphere coupled to an ocean general circulation model and dynamic/thermodynamic sea ice and diagnostic vegetation models. This model has been used to simulate the transient evolution of the past four glacial–interglacial cycles (Timmermann et al., 2014). Numerical experiments are also underway with the Pennsylvania State University AIS model (Pollard and DeConto, 2009), forced by the LOVECLIM transient model output of the past 408 ky, which will allow future investigations of climate–ice sheet interactions on orbital timescales and comparisons with paleoproxy data generated from Iceberg Alley drill cores. Climate–ice sheet interactions can also be studied with the Consortium for Small Scale Modeling (COSMOS) climate model, a state-of-the-art comprehensive coupled general circulation model that has been extensively used to study past climate (Knorr et al., 2011; Wei et al., 2012; Zhang et al., 2013). The COSMOS model is now coupled to an ice sheet model that includes ice-shelf and grounding-line dynamics (Barbi et al., 2013; Thoma et al., 2014).

7. ACC flow speed reconstruction

The SHW are crucial for the Drake Passage throughflow (e.g., Marshall and Speer, 2012; Toggweiler and Samuels, 1995), atmospheric CO₂ release (Denton et al., 2010; Lamy et al., 2007; Sigman and Boyle, 2000), and paleoproductivity (Anderson et al., 2009; Lamy et al., 2014; Roberts et al., 2016; Spreng et al., 2013). For the last glaciation, a northward shift of oceanic fronts is suggested for the southeast Pacific (Lamy et al., 2007; Moreno et al., 1999), South Atlantic (e.g., Barker et al., 2009; Park and Latif, 2008), and southwest Pacific/Indian Ocean regions (e.g., De Deckker et al., 2012; Lorrey et al., 2012). Modeling studies suggest a reduction of ACC transport volume during glacial stages and an increase during glacial terminations (Knorr and Lohmann, 2003, 2007), during which the ACC and associated fronts shifted south.

Based on grain size analyses on 12 cores from a north–south transect across the Scotia Sea, McCave et al. (2014) examined the frontal system of the ACC and concluded that during the Last Glacial Maximum (LGM), near-bottom current velocities were essentially the same as they are today, except for the sea ice zone south of $\sim 56^\circ$ S where they were slower. A recent study found a glacial reduction and millennial timescale variability in the ACC throughflow (Lamy et al., 2015). Our drill sites are located in the southern half of this transect. Using a similar methodology, we will reconstruct paleocurrents to evaluate the long-term evolution of current speeds in the ACC and whether changes in the Drake Passage throughflow and shifts in frontal systems are associated with Seismic Units I–V in the long term and with glacial–interglacial and/or millennial timescales in the short term. Analyses will be made in conjunction with modeling efforts and compared to other Southern Ocean regions.

Drilling in the more southerly Dove Basin should also provide insights into the evolution of the WSDW overflow through the Orkney Passage. As a working hypothesis, this process could have started with Reflector c during the mid-Miocene and strengthened after Reflector a during the late Pliocene according to Maldonado et al. (2006). Both Dove Basin and isolated basins in the South Scotia Ridge show complex contourite systems (Maldonado et al., 2003). Could this imply that Dove Basin was generally more dominated by WSDW, whereas Pirie Basin was under higher influence of CDW? Today, WSDW flows northwest at a relatively high speed (≤ 0.6 m/s) at the southern limit of Dove Basin (Naveira Garabato et al., 2002).

8. Pliocene–Pleistocene ice age climate cycles and underlying causes

The global cooling trend observed over the last 3–5 My, with the initiation of Northern Hemisphere glaciation at ~2.6 Ma, has been attributed to the interplay of changes in orbital configuration (Lisiecki and Raymo, 2005), CO₂ decrease (Lunt et al., 2008; Raymo et al., 1988), tectonic development (Cane and Molnar, 2001; Raymo et al., 1988), intensification of Antarctic glaciation (McKay et al., 2012), and the emergence of the cold tongue in the eastern equatorial Pacific (Martinez-Garcia et al., 2010). Prevalent theories for subsequent glacial–interglacial changes propose that fluctuations in global ice volume are caused by variations in the amount of insolation received at critical latitudes and seasons due to variations in Earth's orbital parameters. In particular, it is widely agreed that variations in ice volume over the last 700,000 y at the precessional (~23 ky) and obliquity (41 ky) frequencies are directly forced by and coherent with Northern Hemisphere summer insolation, whereas the ~100 ky eccentricity component of the ice age cycle results from nonlinear amplification mechanisms (e.g., Abe-Ouchi et al., 2013; Imbrie et al., 1993, 1992). However, during the late Pliocene–early Pleistocene interval from ~3 to 1 Ma, only weak variance at 100 and 23 ky periods was observed in proxy ice-volume records such as benthic $\delta^{18}\text{O}$ (e.g., Raymo and Nisancioglu, 2003). Instead, the records are dominated by the 41 ky obliquity cycle (Figure F5). So why is there no strong precessional signal as predicted by the Milankovitch model?

Hypotheses presented to explain this conundrum point to the Southern Hemisphere and its important role in contributing to the observed global climate signal (Huybers and Denton, 2008; Raymo et al., 2006; Raymo and Nisancioglu, 2003). However, only a few records are available from the Southern Ocean (e.g., Ocean Drilling Program [ODP] Site 1090; Martínez-Garcia et al., 2011), East Antarctica (e.g., Integrated Ocean Drilling Program Site U1356 at Wilkes Land; Cook et al., 2013; Patterson et al., 2014), and West Antarctica (e.g., ANDRILL Core AND-1B; McKay et al., 2012) to test these various hypotheses. For instance, McKay et al. (2012) suggested that the intensification of Antarctic cooling strengthened SHW, invigorated circulation in the Southern Ocean, reduced the AMOC, and thereby contributed to Northern Hemisphere glaciation. On the other hand, Raymo et al. (2006) proposed that the precession cycle should be obvious in Antarctic ice-volume variations prior to the MPT but is absent in globally integrated records such as $\delta^{18}\text{O}$ because precession is out of phase between the Northern and Southern Hemisphere ice sheets. Because our proposed sites are located in a region key to reconstructing the evolution of the Southern Ocean and AIS, the climate and ice sheet signals we find should help in solving this Neogene puzzle.

Scientific objectives for the Subantarctic Front sites

It was anticipated that Sites U1534 and U1535 from the contourite drift deposited beneath the Subantarctic Front will allow us to extend reconstructions of Antarctic Intermediate Water (AAIW) variability and migrations of the Subantarctic Front (Figure F1) to at least 1 Ma. Existing short (~7 m) cores from the area demonstrate the potential of these sediments to yield paleoceanographic records that document past changes in surface and intermediate waters. Deposited beneath the jet of the Subantarctic Front, the drift sediments are sensitive to velocity changes associated with frontal migration. Roberts et al. (2017) compared sortable silt records from the same drift with records from Drake Passage to infer a more northerly position of the Subantarctic Front during the last glacial.

The Subantarctic Front is inferred to have shifted to its current position during the deglaciation, effectively opening up the “cold water route” and enhancing throughflow of Pacific-sourced intermediate waters into the South Atlantic. Stable isotope and elemental analyses of benthic foraminifers in these cores were also used to reconstruct the history of intermediate–deep density stratification in the Atlantic sector of the Southern Ocean since the LGM (Roberts et al., 2016).

The cores collected at Sites U1534 and U1535 will enable us to better understand the role of the cold water route on millennial, orbital, and longer timescales. Calcareous benthic foraminifers suitable for geochemical analysis were found in the drift sediments throughout the last glacial cycle. The persistent occurrence of calcareous foraminifers and tuning of benthic $\delta^{18}\text{O}$ to the LR04 stack will provide an orbital-scale age model for the longer drift records recovered during Expedition 382. Accurately resolving interhemispheric phasing will help identify interactions between the cold water route and the AMOC (Barker et al., 2011). In addition, new records from the Subantarctic Front drift will complement the “warm route”/Agulhas leakage records of the southeast Atlantic (e.g., Marino et al., 2013; Martínez-Méndez et al., 2010; Peeters et al., 2004), as well as the records generated from ODP Leg 177 (e.g., Caley et al., 2013; Martínez-Garcia et al., 2011) and emerging records from IODP Expedition 361 (Hall et al., 2017).

Additional objectives for coring the drift site include providing the first long-term marine perspective on the evolution of the Patagonian Ice Sheet (PIS) and its interactions with the southwest Atlantic and generating the first long-term SST records from the subantarctic southwest Atlantic.

Broader implications

Deep drilling in Iceberg Alley provides an outstanding opportunity to understand late Neogene AIS variability, to identify regional sources of ice mass loss, and to deliver an important historical context for future projections of AIS behavior in a warmer world. Recent novel findings (Bakker et al., 2016; Fogwill et al., 2017; Weber et al., 2014) and the hypotheses developed for this proposal address AIS stability and its relation to climate change. In particular, they challenge scenarios in which the AIS made only a relatively small contribution to the LGM sea level lowstand (e.g., Whitehouse et al., 2012) and made virtually no contribution to MWP-1A (e.g., Bentley et al., 2010, 2014; Mackintosh et al., 2011, 2014). A better understanding of the Antarctic sea level contribution is critical to understanding late Neogene changes and projecting sea level rise in a future warming world. Global mean sea level could rise >1 m by 2100, and large uncertainties remain in projecting AIS contributions to sea level rise beyond that time (Church et al., 2013). Recently appreciated ice-loss mechanisms such as ice-cliff failure and hydrofracturing (DeConto and Pollard, 2016) may increase the possible contribution of the AIS to future sea level rise substantially, specifically on the multicentennial timescale. Understanding the role of ocean thermal forcing and the resulting changes in water masses and sea ice extent, as well as their influence on past AIS behavior, will help constrain projections of future sea level rise.

By reconstructing past variability back to the mid-Miocene in sea ice, SSTs, water masses, dust-climate couplings, and deepwater current speeds in this major Antarctic gateway, our expedition holds the potential to provide important new understanding of the relation between AIS volume changes, shifting SHWs, ocean fertilization, and associated changes in CO₂.

Principal results

Site U1534

Background and objectives

Site U1534 is located at 53°11.38'S, 58°45.65'W, in 605 m of water (Table T1). The site is situated on the northern flank of a large trough ~660 km east of the Strait of Magellan at Common Depth Point (CDP) 10005 on Seismic Reflection Profile SGFI93. Site U1534 is the first of a pair of sites. The second site (U1535), which is 8 km to the east, targeted an expanded copy of the unit between Reflectors A and B of the Site U1534 stratigraphy, to 120 mbsf.

The primary objective at Site U1534 was to recover a continuous ~300 m sedimentary section from a contourite drift where the crest is currently being deposited beneath the Subantarctic Front in 600 m of water. The drift is bathed with AAIW as it begins its northward penetration into the Atlantic Ocean. The principal scientific motivation to core this drift is to better understand the role of AAIW in climate change on millennial, glacial–interglacial, and longer time-scales.

AAIW is one of the primary water masses characterizing the northward-flowing upper limb of the AMOC. Counterbalancing the southward flow of NADW, AAIW currently penetrates as far as 20°N in the North Atlantic Ocean. Despite a recent surge in research effort (e.g., Voigt et al., 2016; Roberts et al., 2017; McClymont et al. 2016), a consensus on contributions of AAIW to AMOC variability has not been reached.

Today, AAIW is primarily formed in the southeast Pacific and southwest Atlantic Oceans. AAIW and its precursor, Subantarctic Mode Water, are advected eastward through the Drake Passage along the Subantarctic Front. This injection of cold, low-salinity Pacific water into the South Atlantic Ocean is commonly referred to as the cold water route. Variability in the connectivity between the Pacific and Atlantic Oceans via this cold water route may modify the physical properties of AAIW forming in the southwest Atlantic Ocean, in turn affecting the heat and salinity budget of the AMOC. Proxy records from Sites U1534 and U1535 will provide the first constraints on past variability in end-member physical and geochemical properties of AAIW as it enters the upper limb of the AMOC.

Site U1534, located at the thickest part of the contourite drift, targeted the highest sediment accumulation rates to optimize the temporal resolution of subsequent reconstructions. Four laterally offset sedimentary units indicate that the drift was deposited under an evolving bottom current regime (Koenitz et al., 2008; Pérez et al., 2015). Seismic interpretation of drift deposits throughout the Scotia Sea area suggests that drift deposits began forming beneath the ACC in the early Miocene and that unconformity-bounded sedimentary units represent deposition under different oceanographic configurations. To test hypotheses about the conditions under which these units were deposited, a key objective at Site U1534 was to core all three of the reflectors and determine the ages of the unconformities. Furthermore, confirming whether these reflectors are synchronous with their proposed counterparts in the Scotia Sea will enable a regional reconstruction of oceanographic evolution of the ACC in this sector of the Southern Ocean.

Operations

Expedition 382 started on 20 March 2019, and the science party boarded on 21 March. The vessel left Mardones Pier at 2135 h to make way for a commercial vessel and moved to anchorage close to

Punta Arenas, Chile. On 25 March at 1630 h, we started the 19.5 km transit to the Cabo Negro Oil Terminal, arriving at 2015 h. Overnight, 1300 metric tons of arctic-blend diesel were loaded. However, the ship was not permitted to depart because winds remained above the pier's operating limit until the evening of 27 March. The ship pushed off at 2230 h, and the pilot disembarked at 0535 h on 28 March, beginning the sea voyage to Site U1534. The 461 km transit was completed in 39 h at an average speed of 11.8 kt. We arrived at Site U1534 at 1300 h on 29 March.

The first core of Expedition 382 was taken in Hole U1534A at 2250 h on 29 March 2019 in 605.1 m of water. We used the advanced piston corer (APC) system to take Cores 382-U1534A-1H through 14H (to 115.7 mbsf), drilled through a hard layer, and switched to half-length APC (HLAPC) coring for Cores 16F–44F (to 256.7 mbsf). The extended core barrel (XCB) system was used to take the final cores (46X and 47X). Overall, Cores 1H–47X penetrated from the seafloor to 266.3 mbsf and recovered 271.7 m (102%).

Hole U1534B started at 0245 h on 1 April. Core 382-U1534B-1H recovered 1.5 m of sediment, but the core liner shattered and no clear mudline was observed, so we ended the hole.

Hole U1534C started at 0345 h on 1 April. Cores 382-U1534C-1H through 19H penetrated from the seafloor (606.3 meters below sea level [mbsl]) to 168.0 mbsf and recovered 159.6 m (95%).

Hole U1534D started at 0045 h on 2 April. The purpose of this hole was to fill a stratigraphic gap and provide more material for sampling in the upper part of the stratigraphy. Cores 382-U1534D-1H through 3H penetrated from the seafloor to 28.5 mbsf and recovered 29.6 m (104%).

Principal results

Late Pleistocene strata (0–83 mbsf) are characterized by interbedded foraminifer-bearing greenish gray silty clay and clayey silt. A distinctly lighter colored sediment interval spanning several sections in Core 382-U1534C-5H contains almost 50% carbonate (composed of nannofossils and foraminifers) and likely represents interglacial MIS 11. Centered at 40 mbsf, this interval is interpreted to represent Reflector A in the seismic profile. Other intervals where planktonic foraminifers are more abundant are tentatively assigned to interglacials MIS 5e, MIS 9, and MIS 17, as supported by paleomagnetic and biostratigraphic data and by cyclicity in physical property data such as natural gamma radiation (NGR) and bulk density, with interglacials exhibiting lower NGR and bulk density values.

The most common biogenic components, sponge spicules and diatoms, generally compose between 10% and 25% of the sediment. Calcareous microfossils such as foraminifers, ostracods, and coccoliths are concentrated in meter-scale layers corresponding to interglacial periods to ~83 mbsf. Several well-preserved macrofossils were found throughout the cores, including solitary cold-water corals, gastropods, and mollusk shell fragments. Ice-rafted clasts are present but rare.

A downhole transition to early Pleistocene–late Pliocene radiolarian and diatom species below 83 mbsf (Section 382-U1534A-10H-3), the depth expected for Reflector B, indicates that this reflector represents a hiatus. The absence of a clear magnetic reversal from 0 to 83 mbsf suggests that the age above the unconformity is no older than 0.78 Ma. Below Reflector B, paleomagnetic normal polarity extends to 233 mbsf and may constrain most of this interval to Gauss Subchron C2An.1n (2.58–3.03 Ma), implying sedimentation rates of ~40 cm/ky. Green-gray silty clay continues to dominate

below Reflector B. Carbonate remains <1 wt% from 110 to 233 mbsf and increases to as high as 5 wt% below 233 mbsf. Smear slides suggest that foraminifers are rare to common below 233 mbsf.

Diatom and radiolarian biostratigraphy indicate a second hiatus at ~257 mbsf that is interpreted to be Reflector C. Below Reflector C, diatom and radiolarian biostratigraphic markers date the two lowermost cores of Hole U1534A (257–266 mbsf) as early Pliocene (4.1–4.7 Ma). Samples for the analysis of sedaDNA were collected from the upper 91 m, including high-resolution sampling throughout the Holocene in the upper 9 m.

Interstitial water and headspace gas samples were taken from all APC cores and every second HLAPC core from Hole U1534A and have a normal headspace gas depth profile. Salinity is relatively low from ~30 to 150 mbsf, and the highest alkalinity (as high as 30 mM) is at 20–30 mbsf, coinciding with the methane–sulfate transition zone.

Four formation temperature measurements were made with the advanced piston corer temperature tool (APCT-3) while taking Cores 382-U1534A-4H, 7H, 10H, and 13H. The geothermal heat flow was estimated using formation temperature measurements and laboratory thermal conductivity measurements.

In summary, cores collected at Site U1534 will give new insights into the evolution of the contourite drift deposit and ice sheet–ocean–climate interactions in the southwest Atlantic Ocean, an area where long paleoclimatic records are relatively sparse. Dating the seismic reflectors and the sediment units between these unconformities will allow us to reassess the local ice sheet and oceanic regimes under which sediment supply and ocean current configuration led to drift deposition. High sediment accumulation in the late Pliocene interval will provide insights into the history of the PIS in addition to AAIW export as Northern Hemisphere glaciation intensified. Late Pleistocene sedimentation rates of ~10 cm/ky will allow millennial-scale reconstructions of the position of the Subantarctic Front as well as AAIW export to be extended back to at least MIS 17. The prominence of MIS 11 at Site U1534 and all the other sites cored during Expedition 382 provides an opportunity to reconstruct a transect across the ACC during this superinterglacial in unprecedented temporal resolution. Comparison of the records from Site U1534 with recent records from Amundsen Sea and Ross Sea IODP Expeditions 379 and 374 will provide a critical link between processes on the Antarctic continent, South American glaciations, and the upper limb of the AMOC. Interhemispheric and interbasin comparison of these records will determine how variable transport via the cold water route may affect AAIW export in the South Atlantic and the rate of AMOC on multiple timescales.

Site U1535

Background and objectives

Site U1535 is located on the northern flank of a large trough ~668 km east of the Strait of Magellan at 53°11.49'S, 58°38.60'W, in 647 m of water (Table T1). The site is situated at CDP 9220 on Seismic Reflection Profile SGFI93-107. Site U1535 is the second of a pair of sites, with the first site (U1534) located 8 km to the west. The scientific motivation to core this drift is identical to that of Site U1534 (see [Site U1534](#) for detailed information about the objectives at this pair of sites). The more extensive coring program at Site U1534 targeted the more expanded units above Reflector A and below Reflector B, and Site U1535 was selected to exploit the thicker lateral extension of the sedimentary unit between Reflectors A and B. The primary objective at Site U1535 was, therefore, to recover a continuous 120 m sedimentary section from the upper sequences of

the contourite drift deposited on the northern flank of the east–west trending trough.

Operations

After completing the 8 km transit east from Site U1534 in dynamic positioning (DP) mode, we arrived at Site U1535 at 0930 h on 2 April 2019. Hole U1535A started at 1200 h on 2 April. The seafloor was determined by the Core 382-U1535A-1H mudline to be at 646.9 mbsl. Cores 9H–11H were difficult to remove from the core barrel, so we switched to HLAPC coring for Core 12F at 88.5 mbsf. Cores 1H–17F penetrated from the seafloor to 117.3 mbsf and recovered 112.4 m (96%). Having achieved sufficient stratigraphic coverage and considering the overall time constraints for the expedition, we concluded operations at Subantarctic Front Sites U1534 and U1535 and started pulling up the drill string at 0330 h on 3 April. The bit cleared the rotary table at 0840 h on 3 April, ending Hole U1535A, and the rig floor was secured for the sea voyage to the Scotia Sea sites.

Principal results

Hole U1535A contains several sequences of silty clays interbedded with clayey silts and clay layers. The most common biogenic components (sponge spicules and diatoms) generally compose between 10% and 25% of the sediment. Intervals containing high concentrations of calcareous nannofossils and other microfossils are interpreted to be interglacial sediments. Several well-preserved macrofossils were found throughout the cores, including solitary cold-water corals, gastropods, and mollusk shells and shell fragments. Ice-rafted clasts are present but rare.

Species assemblages and biostratigraphic age markers are similar to those at Site U1534. We identified 17 diatom, radiolarian, and palynomorph biostratigraphic age events. Mid- to early Pleistocene biostratigraphic markers were recovered above Reflector B, and late Pliocene–markers were found in the two lowermost cores, beneath Reflector B. A downhole paleomagnetic transition from reversed to normal polarity in the two lowermost cores may represent the transition from the Matuyama Chron (C2r; 2.58 Ma) to Gauss Chron (C2An.1n, which is consistent with biostratigraphic constraints and the seismic stratigraphic relationship to Site U1534.

Interstitial water analyses show a typical diagenetic profile similar to observations made at neighboring Site U1534, with the characteristic signal of sulfate reduction in the uppermost 20 m and increasing ammonium, alkalinity, and barium concentrations below. Total organic carbon (TOC), total nitrogen (TN), and calcium carbonate contents range from 0.4 to 1.4 wt%, from 0.02 to 0.15 wt%, and from 0.1 to 49 wt%, respectively.

Two formation temperature measurements were made with Cores 382-U1535A-4H and 7H. The geothermal heat flow was estimated using formation temperature measurements with thermal conductivity core measurements.

The late Pleistocene interval at Site U1535 is more condensed than at Site U1534, with a 3 m thick pale-colored carbonate ooze that we interpret to be MIS 11 occurring at 25 mbsf compared with 40 mbsf at Site U1534. This depth is consistent with seismic Reflector A representing the foraminifer ooze of MIS 11 as also observed at Site U1534. Visual core descriptions noted that transitions in and out of the MIS 11 foraminifer ooze and other interglacial intervals are abrupt relative to those at Site U1534, suggesting the presence of multiple nondeposition and/or erosional hiatuses. Biostratigraphic and paleomagnetic interpretation is limited by these hiatuses but is consistent with a mid- to early Pleistocene age of the sediments be-

tween Reflectors A and B. Reflector B occurs at 98 mbsf at Site U1535 based on the correlation to seismic profiles and changes in the physical properties records at this depth. At Site U1534, the oldest sediments recovered above Reflector B are <0.78 Ma, so the recovery of older sediments above Reflector B at Site U1535 shows that the uppermost unit of the drift onlaps Reflector B. It is also noted from the seismic data that the most distal part of the uppermost unit of the drift is truncated, which is consistent with the more common occurrence of hiatuses at Site U1535.

Records from Site U1535 will be most useful for understanding the dynamics of the evolution of the drift. The more continuous, generally higher resolution records from Site U1534 will be preferred for paleoceanographic studies from the drift.

Site U1536

Background and objectives

Site U1536 is located 235 km northwest of the South Orkney Islands at 59°26.46'S, 41°3.66'W, in 3220 m of water (Table T1). Site U1536 is the first of two sites drilled in Dove Basin in the southern Scotia Sea. This site was targeted to recover a complete Neogene record of Antarctic ice and ocean dynamics, examine the character and age of regionally correlative seismic reflectors, and date the acoustic basement to constrain the tectonic origin of Dove Basin.

Previous piston coring at Site MD07-3134 in a small subbasin in the northeastern part of Dove Basin recovered a 58 m long, high accumulation rate sediment record covering the last glacial cycle (Weber et al., 2012). Site U1536 is located 23 km east of that site in the deeper part of a broad basin and on the western flank of a north-south ridge.

Sediments in the southern Scotia Sea are primarily deposited by contourite currents along the pathway of the ACC. Specifically in Dove Basin, contourite deposition is also assumed to be influenced by WSDW flowing from the south through bathymetric gaps around the South Orkney Plateau after exiting the Weddell Sea (Maldonado et al., 2003). The contourites are lens-shaped in seismic profiles, as thick as 1 km in the center of small troughs (Maldonado et al., 2006), and thin toward the edges of the troughs.

Three seismic lines indicate a basin-like structure with several small-scale ridges and continuous reflectors in the central to northern part of Dove Basin. Site U1536 is located at shotpoint (SP) 1709 on Multichannel Seismic (MCS) Reflection Profile SCAN 10/04. The five seismic units that encompass deposition above acoustic basement (Maldonado et al., 2006) served as a framework for drilling. The seismic reflectors show parallel lamination with occasionally undulating structures, which is indicative of minor synsedimentary downslope transport (Figure F3). This downslope movement is a fairly common feature in the basin, and none of the Expedition 382 proposed sites were able to avoid these disturbances completely. Site U1536 provided the best compromise between a relatively thick and undisturbed stratigraphic sequence for climate studies and access to acoustic basement, which was located at a relatively shallow depth of ~900 mbsf.

At Site U1536, the main objective was to obtain a complete late Neogene record of ice and ocean dynamics in the center of Iceberg Alley in the more southerly of our two drilling areas in the Scotia Sea. Specific objectives include (1) the reconstruction of past variability in AIS mass loss and its related sea level history, (2) a study of the water mass composition of the Drake Passage throughflow and WSDW inflow, and (3) a study of north-south shifts of the frontal systems in response to changing climate conditions, including changes in water mass properties, ocean temperature, and sea ice

extent. An additional goal was to reconstruct changes in dust-climate couplings between Patagonia and Antarctica as well as related atmospheric circulation changes throughout the Pliocene–Pleistocene in a more distal location relative to the source (e.g., Patagonia). A final science objective at Site U1536 was to drill through the deeper reflectors, specifically the basin-wide Reflector c, to determine the age of this unconformity. This determination would allow us to date the regional change in bottom water flow, which probably led to the formation of the reflector. We were not able to reach acoustic basement at ~900 mbsf to determine its nature and age.

Operations

We arrived at Site U1536 at 0430 h on 6 April 2019 (Figure F6) after making two 3.5 kHz survey profiles over the site in a cross shape, each 4 nmi long along the site survey seismic lines. The start of coring was delayed by bad weather and mechanical difficulties with the APC bottom-hole assembly (BHA) that started when the “pig” pipe-cleaning device was not fully pumped out of the bottom of the pipe, which caused a series of misfires of the piston corer necessitating a pipe trip back to the ship to inspect the BHA.

Hole U1536A started at 0005 h on 9 April. Cores 382-U1536A-1F through 53F penetrated from the seafloor to 354.4 mbsf and recovered 364.3 m (103%) (Figure F7). At 2115 h on 11 April, an iceberg approached within 5.7 nmi of the ship, so we stopped coring and raised the drill string to 50 mbsf. When the iceberg was 1.7 nmi away, we raised the drill string above the seafloor and moved ~1 nmi east-northeast to avoid its path, ending operations in Hole U1536A.

Hole U1536B started at 0930 h on 12 April. Cores 382-U1536B-1H through 25H penetrated from the seafloor to 226.1 mbsf and recovered 230.7 m (102%). At 1130 h on 13 April, we started to raise the drill string because of a large iceberg and a flotilla of smaller icebergs around it. At 1325 h, the drill string cleared the seafloor, ending Hole U1536B.

Hole U1536C started at 2035 h on 13 April. Cores 382-U1536C-1H through 40F (and drilled intervals 31, 91, 111, 131, 161, and 271, which advanced 144.0 m without recovery) penetrated from the seafloor to 352.0 mbsf and recovered 187.4 m (90% of the cored interval). The aim of Hole U1536C was to spot core the upper section to fill gaps in the stratigraphy recovered in Holes U1536A and U1536B and then core continuously from 224 mbsf down. After taking Core 37F (341 mbsf) late on 15 April, coring was interrupted for 8 h by an iceberg passing through the red zone. We were able to advance to 352.0 mbsf, but then another iceberg entered the red zone and we had to end Hole U1536C.

Hole U1536D started at 1940 h on 16 April. Core 382-U1536D-1H penetrated from the seafloor to 6.9 mbsf. The core liner shattered, and the core had to be pumped out of the barrel, so it could not be used for the intended purpose of interstitial water sampling. We ended Hole U1536D and raised the drill string to the ship to change to the rotary core barrel (RCB) BHA.

Hole U1536E started at 2140 h on 17 April. We drilled without recovery to 312 mbsf, and at 0850 h on 18 April we deployed a free-fall funnel (FFF). We then continued to drill down to 340 mbsf, just shallower than the depth reached in Holes U1536A and U1536C.

We started coring Hole U1536E at 2000 h on 18 April. Cores 382-U1536E-2R through 33R penetrated from 340.0 to 645.4 mbsf and recovered 113.3 m (36%). Coring was interrupted four times, three times by icebergs and once by bad weather conditions. Two of the icebergs forced us to pull out of the hole and then reenter guided by the subsea camera. At 1715 h on 24 April, after recovering Core

33R (645.4 mbsf), we decided to stop coring and log the hole. The “quad combo” logging tool string consisted of tools to measure magnetic susceptibility (MS), NGR, electrical resistivity, sonic velocity, and density of the formation. Starting at 0830 h on 25 April, we made downlog and uplog passes with the tool string, reaching 643 mbsf (~3 m above the bottom of the hole). Ship heave was ~2.5 m during logging, which caused depth discrepancies of a similar magnitude in the logging data. The logging tools were disassembled by 1515 h, and the drill string was raised to the ship, which completed operations in Hole U1536E.

Principal results

APC cores from Holes U1536A–U1536C provide near-continuous stratigraphic coverage to 354 mbsf, and RCB cores from Hole U1536E provide ~36% coverage from 340 to 645 mbsf (Figures F5, F8). Based on the recovered sediments, three major lithostratigraphic units (I–III from top to bottom) were identified (Figure F8). Unit I (seafloor to ~250 mbsf) consists of interbedded diatom oozes and silty clays. Diatom oozes dominate the sediment from the seafloor to 120 mbsf. The terrigenous component in the oozes is generally below 25%, and the biogenic component (mainly diatoms) in the silty clays typically exceeds 25% of the sediment. Unit II (~250–550 mbsf) is almost exclusively silty clay with varying amounts of biosilica. Lithification increases downhole. Unit III (~550 mbsf to the base of Hole U1536E at 643 mbsf) consists of semi- to fully lithified mudstone, some with biosilica, and beds of limestone, some of which are nannofossil bearing to nannofossil rich. Throughout the cores, gravel- to pebble-sized IRD is rare to common and clearly visible in the core X-ray images. A noticeably higher abundance of dropstones was observed from ~435 to 568 mbsf.

Diatom, radiolarian, foraminifer, and palynomorph biostratigraphic results are consistent for all samples from Holes U1536A, U1536C, and U1536E. We recorded 81 biostratigraphic events, based on which we estimate sedimentation rates of ~15 cm/ky from the seafloor to 250 mbsf (~1.5 Ma) and ~5.4 cm/ky from 250 to 620 mbsf (~8.4 Ma). A hiatus between Samples 382-U1536E-30R-CC and 31R-CC (~620 mbsf) leaves a biostratigraphic gap spanning ~13.2–8.4 Ma. No hiatuses were observed in the Pleistocene (seafloor to 320 mbsf) or Pliocene (320–450 mbsf) stratigraphic sections. Samples for the analysis of sedaDNA were collected from the upper 178 m with high-resolution Holocene sampling in the upper 15 m.

All paleomagnetic polarity zones of the 2012 geomagnetic polarity timescale (GPTS12) were identified in the APC-cored intervals of Site U1536 through the lower Gauss Chron (C1n–C2An.3n; 3.6 Ma), and polarity transitions were recovered for all polarity zone boundaries except the Reunion Subchron (C2r.2n) (Figures F5, F9). Despite discontinuous recovery in the RCB cores, we used initial biostratigraphic data to make tentative ties to the GPTS12 for intervals between Chrons C2Ar and C3An (3.60–6.03 Ma). Below this range, we observed intervals of normal and reversed polarity that may potentially be correlated with the GPTS12 in postcruise work.

Samples for headspace gas, interstitial water chemistry, and bulk sediment geochemistry were analyzed from Holes U1536A and U1536E. Headspace methane concentrations are generally low (2–4.6 parts per million by volume [ppmv]) throughout the sedimentary sequence. Ethane, propane, and other higher molecular weight hydrocarbons are below the detection limit. Site U1536 is characterized by moderately reducing sedimentary conditions, as indicated by the disappearance of dissolved sulfate concentrations at ~100 mbsf and a reappearance of resolvable dissolved sulfate at

~300 mbsf. The reducing conditions associated with this microbially mediated sulfate reduction in the upper section exert strong control over the interstitial water profiles of several parameters measured on board, including alkalinity, Ca, PO₄, Ba, Sr, and Mn. Bulk sediment TOC and TN contents are generally low with concentrations ranging from 0.1 to 0.8 wt% and from 0.02 to 0.09 wt%, respectively.

Physical property records reflect the lithology of the sediments. Silty clay-bearing diatom ooze, interpreted to be interglacial sediment, shows high *P*-wave velocity (*P*-wave logger [PWL]) values and low gamma ray attenuation (GRA) density, NGR, and MS values. Also, the color reflectance component *b** (the variability between yellow and blue) shows more yellowish colors for the diatom-rich interglacial periods. Variability in *b** has been shown to be a good indicator for temporal variability of biogenic opal over the last glacial cycle in sediments from the Scotia Sea (Sprenk et al., 2013). The glacial–interglacial cyclicity of *b** reveals high-amplitude variations since the mid-Brunhes transition (MBT) around ~0.43 Ma, thus following a global pattern of higher amplitude climate variability post-MBT (Barth et al., 2018) above ~90 mbsf. *b** shows medium-amplitude variability between ~90 and 181 mbsf (~1.1–0.43 Ma) and low-amplitude, higher frequency variability below that back to 3.5 Ma.

The combination of robust biomagnetostratigraphy and high-resolution records of MS, GRA, NGR, and *b** is used to assign glacial and interglacial periods, providing a strong time-stratigraphic framework for this site. Furthermore, MS, a presumed dust proxy at the sites from Iceberg Alley (Weber et al., 2012, 2018) shows distinct 100 ky peaks representing glacial periods that can be correlated in great detail with the dust record of Antarctic ice cores (e.g., to the EPICA Dome C record of Lambert et al. [2008]) over the last 800 ky.

An excellent late Pliocene–middle Pleistocene section was recovered that will allow in-depth study of the impact of the initiation of Northern Hemisphere glaciation on the Southern Ocean. An extended record of the early Pleistocene will allow us to assess the provenance and flux of IRD prior to the MPT. Although limited carbonate material is available to work with, the rich record of IRD and biosiliceous fossils combined with an excellent magnetostratigraphy should make this sequence suitable for in-depth paleoceanographic studies. Although discontinuously cored and heavily disturbed, the Miocene sequence should also provide an important window into the earlier glaciation history of Antarctica.

A continuous spliced record was constructed from Holes U1536A–U1536C for the top 243 meters composite depth (mcd) (Figure F9). Three intervals appear to contain slumped sediment. Two occur at ~40–50 and 95–100 mbsf, and one occurs below the splice at about 299–301 mbsf.

Downhole logging in Hole U1536E used a quad combo tool string to obtain as much data as possible in a single tool string run because of deteriorating sea conditions and the risk of icebergs. The logging tools measured MS, gamma ray spectra, electrical resistivity, *P*- and *S*-wave velocities, and density. These in situ measurements successfully cover data gaps in core recovery and complement the core measurements made in the laboratory. Both core and downhole density and *P*-wave velocity values generally increase with depth, reflecting downhole sediment compaction.

Integrating the core stratigraphy with the MCS profiles using the core and log velocity data to transform depths to two-way traveltimes, we identified Reflector b at ~365 mbsf in Core 382-U1536E-4R. This reflector corresponds to relative minima in PWL and GRA. We were also able to identify Reflector c, a central objec-

tive of the expedition, at ~530 mbsf in Core 22R. It is characterized by a sharp increase in PWL values. Determining ages for these basin-wide reflectors will be important in the assessment of the geodynamic history of the Scotia Sea.

Site U1537

Background and objectives

Site U1537 is located 265 km northwest of the South Orkney Islands at 59°6.65'S, 40°54.37'W, in 3713 m of water (Table T1). The site lies in the northeast part of Dove Basin in a ~1 km thick contourite drift and is situated at SP 4900 on MCS Reflection Profile SCAN 2014/07 (Figure F3) 5 km west-northwest of Crossing Line SCAN 10/04. The objective at Site U1537 was to obtain a second sedimentary record in Dove Basin with a similar overall seismic structure (Units I–V) but with less stratigraphic disturbance than observed at Site U1536. This site was originally an alternate site but was selected for drilling because the top ~150 m of the stratigraphy appeared to be uninterrupted by the thin slumped intervals observed at the other possible sites in Dove Basin (as seen in the seismic and echo sounder profiles). The goals were to reach Reflector b at ~350 mbsf and to reconstruct the Pliocene–Pleistocene ice and ocean dynamics.

Sediments in the southern Scotia Sea are primarily deposited by contourite currents along the pathway of the ACC. Specifically in Dove Basin, contourite deposition is also assumed to be influenced by WSDW flowing from the south through bathymetric gaps around the South Orkney Plateau after exiting the Weddell Sea (Maldonado et al., 2003). The contourites at Site U1537 are lens shaped in seismic profiles, as thick as 800 m in the center of the small trough, and thin toward the eastern edge.

Interpretation of Seismic Line 2004/07 identified Reflectors a, b, and c (Pérez et al., 2017) (Figure F3). In Dove Basin, three seismic lines indicate a basin-like structure with several small-scale ridges and continuous reflectors in the central to northern part of the basin. The seismic reflectors show overall parallel lamination with occasionally undulating structures, indicative of minor synsedimentary downslope transport. This downslope movement is a fairly common feature in the basin, and none of the Expedition 382 proposed sites were able to avoid these disturbances completely.

At Site U1537, the main objective was to obtain a late Neogene record of ice and ocean dynamics from the center of Iceberg Alley in the more southerly of our two drilling areas in the Scotia Sea. Specific objectives include (1) the reconstruction of past variability in AIS mass loss and the related sea level history, (2) a study of the water mass composition of the Drake Passage throughflow and WSDW inflow, and (3) a study of north–south shifts of the frontal systems in response to changing climate conditions including changes in water mass properties, ocean temperature, and sea ice extent. An additional goal was to reconstruct changes in dust–climate couplings between Patagonia and Antarctica as well as related atmospheric circulation changes throughout the Pliocene–Pleistocene in a distal location relative to the source (e.g., Patagonia).

Operations

We arrived at Site U1537 at 0515 h on 26 April 2019 (Figure F6). We started to make up the APC BHA, but high heave and roll forced us to pause operations for ~7 h. Additionally, we had to move ~1 nmi northwest to allow an iceberg to pass over the site location. Operations resumed at 1430 h, and we lowered the drill string to the seafloor.

We started Hole U1537A at 0140 h on 27 April at 3712.9 mbsf, as calculated from the mudline. Cores 382-U1537A-1H through 31F penetrated from the seafloor to 264.0 mbsf and recovered 268.9 m (102%) (Figure F7). After Core 26H was difficult to pull out of the formation and had to be drilled over, we switched to HLAPC coring for Core 27F at 240.0 mbsf. An iceberg moved into the red zone at 1830 h on 28 April, so we raised the drill string to 38 mbsf. At 2020 h after further encroachment of the iceberg, we ended Hole U1537A. Although we could not core at depth in Hole U1537A, we could take a mudline core because this operation does not require the drill string to be below the seafloor, and the iceberg was not close enough to be in the termination zone.

Hole U1537B started at 0010 h on 29 April and consisted of a single core taken for high-resolution interstitial water sampling. Core 382-U1537B-1H penetrated from the seafloor to 7.4 mbsf.

Hole U1537C started at 0135 h on 29 April, and Cores 382-U1537C-1H through 2H penetrated from the seafloor to 15.3 mbsf and recovered 11.9 m (78%). After the Core 2H liner shattered and two large icebergs entered the red zone at 0400 h, we decided to end Hole U1537C. We raised the drill string clear of the seafloor and moved aside in DP mode to let the icebergs pass. By 1500 h we were able to move back to the site location.

Hole U1537D started at 1945 h on 29 April. Cores 382-U1537D-1H through 50F penetrated from the seafloor to 354.3 mbsf and recovered 349.0 m (99%). We took two HLAPC cores (20F and 21F) across a hard layer in Hole U1537A at ~182 mbsf before returning to full-length APC coring for Cores 22H–27H. Below 244 mbsf, the formation was too hard for APC coring, so we used HLAPC coring from that point to the bottom of the hole. Icebergs were monitored on 30 April and 1 May, but none of them posed a hazard to operations. From 1345 to 1430 h on 2 May, we paused coring while monitoring an iceberg that was at the edge of the red zone, but it turned away from the ship and we resumed coring. Because high winds and swell were forecast for 3 May, we decided to end Hole U1537D in the early evening of 2 May to leave enough time to raise the drill string to the ship and also leave sufficient time for the coring and logging program at the next site (U1538) in the Pirie Basin area. The bit cleared the seafloor at 2140 h and was raised to the ship by 1030 h on 3 May, which completed operations in Hole U1537D.

Principal results

Piston cores from Holes U1537A and U1537D provided near-continuous stratigraphic coverage to 264 and 354 mbsf, respectively (Figures F5, F8). Based on these cores, three major lithostratigraphic units were identified (Figure F8). Unit I (seafloor to ~180 mbsf) consists of interbedded diatom ooze and silty clay. Unit I is divided into three subunits depending on the relative abundance of diatom ooze and silty clay. Unit II (~180 to ~260 mbsf) is almost exclusively silty clay with varying amounts of biosilica and contains discrete color-banded intervals. Unit III (~260 to ~354 mbsf) is more lithified and consists of interbedded silty clay and diatom ooze. Unit III contains a series of slumps with inclined and folded beds interbedded with horizontally layered sediment. Throughout the cores, IBRD is rare to common and particularly visible in the core X-ray images as granule- to pebble-sized shapes with sharp edges. Bioturbation is also rare to common with abundant thin (millimeter thick), pyritized burrows visible in the X-ray images.

Diatom, radiolarian, foraminifer, and palynomorph biostratigraphic results are consistent for all samples from Holes U1537A and U1537D and show 61 biostratigraphic events. Based on these

events, we estimate sedimentation rates of ~ 14.8 cm/ky between the seafloor and 180 mbsf (~ 1.5 Ma) and ~ 6.1 cm/ky from 180 mbsf to the base of Hole U1537D at 354 mbsf (~ 4.45 to ~ 5.0 Ma). Diatom and radiolarian biostratigraphy indicates that Site U1537 has a continuous stratigraphy from the early Pliocene to the Holocene because no major hiatuses were detected. Diatoms and radiolarians are considerably reworked, especially from 287 to 311 mbsf. Overall preservation is relatively good in all samples, and fragmentation is more common than dissolution in diatoms. Palynomorphs (dinocyst, acritarch, prasinophytes, pollen, spores, copepod, and fungi remains) were identified in the 23 samples processed for palynology. Dinocyst diversity is generally low with relative abundance varying from abundant/common in the upper 155 m to few/barren in the lower half of Holes U1537A and U1537D. An exception is Samples 382-U1537D-38F-CC to 42F-CC (296–316 mbsf), which have a relatively high dinocyst abundance. A small number of samples (five) were collected for the analysis of sedaDNA at Site U1537 with the primary purpose of investigating spatial variability in the genetic signals between this site and nearby Site U1536.

All paleomagnetic polarity zones of the GPTS12 were identified in the combined paleomagnetic records from Holes U1537A and U1537D spanning the Brunhes Chron (C1n) to the base of the Mammoth Subchron (C2An.2r; 3.33 Ma) (Figures F5, F9). Deeper than Chron C2An.2r, we measured ~ 5 m of normal polarity, which we interpret to be the early Gauss Chron (C2An.3n). This interpretation places the base of this polarity zone between 3.33 and 3.60 Ma on the GPTS12. Intervals of slumped sediments underlie this normal polarity zone, but the intervening flat-lying sediments include a polarity transition and ~ 20 m of reversed directions, which we interpret to be the late Gilbert Chron (C2An and C2Ar). The two lowermost cores from Hole U1537D (Cores 382-U1537D-49F and 50F) have scattered paleomagnetic directions and cannot be assigned to a polarity zone at this time. This suggests that the base age for Hole U1537D above these two cores is between 3.6 and 4.2 Ma.

Samples were collected for headspace gas and interstitial water chemistry analyses in Hole U1537A. Downhole profiles of headspace gas and bulk inorganic and organic carbon at Site U1537 are very similar to those from Site U1536. Methane concentrations in headspace gas are low in all cores, and concentrations of ethane and propane are below the detection limit throughout the hole. Hole U1537A has relatively low contents of TOC, TN, and calcium carbonate. Site U1537 has comparable trends to Site U1536 but with some subtle differences. The sulfate-depleted zone is significantly less expanded than at Site U1536, encompassing <80 m. Dissolved Ba contents at those depths are significantly lower than at Site U1536, and alkalinities reach <30 mM (compared with almost 40 mM at Site U1536). Ammonia concentrations are also lower. Systematic Mg and K depletion downhole suggests that continuous authigenic clay formation occurs throughout the sediment column. Calcium is precipitated from interstitial water in the upper section and replenished below the sulfate reduction zone. Other major and trace elements, such as B, Si and Sr, increase downhole. Fe contents are generally low, whereas Mn is somewhat elevated in the uppermost 30 m. Overall, the observed geochemical patterns suggest only moderate diagenesis, which is encouraging for a variety of planned postcruise geochemical analyses.

Distinct cyclic changes were observed in physical property records in Holes U1537A and U1537D, especially for NGR and MS (Figure F9). These cycles are as long as 10 m in the upper 180 m. Below 180 mbsf, cycles are shorter, have lower amplitude, and are superimposed on longer scale changes. Cyclic changes are likely

controlled by insolation-driven climate variations and are useful for stratigraphic correlation. Downhole sediment compaction is reflected in the overall increasing trend in density and P -wave velocity with depth. Step changes in the physical property records line up with main discontinuities in seismic and echo sounder profiles. In particular, a baseline increase in GRA density, MS, and NGR at ~ 180 mbsf is associated with Reflector b in the preexpedition seismic interpretation. Additionally, an increase in velocity and density observed at 230 mbsf is coincident with an increase in reflectivity in the seismic data.

Site U1538

Background and objectives

Site U1538 is located 365 km north-northeast of the South Orkney Islands at $57^{\circ}26.52'S$, $43^{\circ}21.47'W$, in 3131 m of water (Table T1). The site lies in Pirie Basin and is situated at SP 4900 on Seismic Reflection Profile SCAN 2013 13 9 km east of Crossing Line SCAN 2013 11 (Figure F2). Site U1538 is located in the center of a small basin and provides the most undisturbed stratigraphy among the six potential drill sites in Pirie Basin. The seismic presite survey was conducted with a 3-channel SPARKER system (Maldonado et al., 2013), so the depth penetration and resolution of the profile is limited compared to the MCS records of Dove Basin Sites U1536 and U1537 and it is more difficult to follow prominent reflectors and assign seismic units. However, extended Pliocene–Pleistocene sections (Seismic Units I and II) are expected. Identification of Reflector d and acoustic basement is questionable.

Previous piston coring in Pirie Basin was conducted 9 km west of Site U1538 at Site MD07-3133 (alternate proposed Site SCO-01) along Seismic Line SCAN 2013 13. Site MD07-3133 recovered a 36 m long sediment core that reaches back to MIS 3, which implies that the high sedimentation rates observed at Dove Basin Sites U1536 and U1537 could be doubled in Pirie Basin. However, Site MD07-3133 is located at a rather marginal position of the small basin in the northern Pirie Basin with potential seismic discontinuities below ~ 160 mbsf. Therefore, Site U1538, in the central part of that basin, was chosen for drilling to achieve a mostly undisturbed recovery and the thickest Pleistocene section.

Sediments in the southern Scotia Sea are primarily deposited by contourite currents along the pathway of the ACC. Pirie Basin is located further north than Dove Basin closer to the center of the ACC and is expected to be less influenced by WSDW. The contourites are generally lens shaped and have a total thickness of as much as 1 km, similar to the sediment deposition observed above basement in Dove Basin (Maldonado et al., 2006).

At Site U1538, the main objective was to obtain a complete record of ice and ocean dynamics in the center of Iceberg Alley since the late Neogene in the more northerly of our two drilling areas in the Scotia Sea. Specific objectives in Pirie Basin include (1) the reconstruction of past variability in AIS mass loss and the related sea level history at a site with potentially more IRD derived directly from the Pacific side of the WAIS through clockwise transport in the ACC; (2) the capture of the northern, more CDW-dominated part of the Drake Passage throughflow and associated north–south shifts of frontal systems, water mass properties, ocean temperature, and sea ice extent; and (3) the reconstruction of changes in atmospheric circulation and dust-climate couplings between Patagonia and Antarctica and related atmospheric circulation changes throughout the Pliocene–Pleistocene in a more proximal location relative to the dust source (e.g., Patagonia).

Operations

We arrived at Site U1538 at 0240 h on 4 May 2019 after the 127 nmi northwest transit from the Dove Basin area (Figure F6).

We started Hole U1538A at 1615 h on 4 May. The seafloor depth was 3130.6 mbsf, as calculated from the mudline. Over the next 6 days, full-length APC/XCB Cores 382-U1538A-1H through 75X penetrated from the seafloor to 676.0 mbsf and recovered 476.4 m (70%) (Figure F7). Formation temperature measurements were made while taking Cores 6H, 9H, 12H, 15H, and 18H.

At 1130 h on 5 May, after taking Core 382-U1538A-17H, an iceberg and a bergy bit entered the red zone, so we raised the drill string to 45 mbsf and offset 35 m northwest to allow the bergy bit to pass. The larger of the two icebergs was clear of the red zone at 1530 h, and we lowered the drill string back to 159.8 mbsf and resumed coring. Because Core 30H was difficult to extract from the formation and required a drillover to retrieve it, we switched to HLAPC coring for Cores 31F–34F, reaching 294.4 mbsf.

At that point, the hole was deep enough that we would want to reenter it if forced off the site by an iceberg, so at 1720 h on 6 May we deployed a FFF to enable reentry. We resumed coring at 1830 h, taking Core 382-U1538A-35F, which also required a drillover, so we switched to XCB coring for the rest of Hole U1538A. From 1300 to 1500 h on 8 May, we waited for a large iceberg to pass the edge of the red zone (3 nmi from the ship) with the drill bit close to the bottom of the hole. After taking Core 66X (597.9 mbsf), coring was interrupted from 0415 to 0900 h on 9 May when an iceberg passed through the red zone. We raised the drill string to 45 mbsf and waited until after the iceberg passed, with a closest distance of 1.2 nmi from the ship, before resuming coring. From 1930 to 2100 h, ship heave was about 4 m, which was too high to maintain a consistent weight on bit, so we stood by and waited for heave to subside. The final cores from Hole U1538A (Cores 69X–75X) were recovered without interruption. We stopped coring at 1315 h on 10 May after reaching the maximum depth of 676 mbsf permitted by the IODP Environmental Protection and Safety Panel for this site.

We decided to forgo downhole logging in Hole U1538A because ship heave exceeded 3 m and was forecast to become higher, which would pose a risk to the downhole logging tools and increase uncertainty in the depth registration of the logging data. Because the hole has a FFE, it was possible that we could return to log it if sea and heave conditions became calmer before the end of operations at this site. We displaced the hole with 10.5 lb/gal heavy mud to help stabilize the borehole walls in case we were able to log.

We started Hole U1538B at 1900 h on 10 May. Core 382-U1538B-1H recovered 7.8 m and was sampled for interstitial water at high resolution. We then offset 20 m to the south.

We started Hole U1538C at 2135 h on 10 May. Cores 382-U1538C-1H through 12H penetrated from the seafloor to 105.9 mbsf and recovered 98.9 m (93%). One HLAPC core (6F) was taken to adjust the depth offset to splice over core breaks in Hole U1538A. At 1000 h on 11 May, the sea and wind conditions had deteriorated to the extent that we could not safely handle the core barrel on the rig floor. We waited with the drill string in the hole, but it was apparent that conditions would not improve soon, so at 1500 h we raised the drill string clear of the seafloor, ending Hole U1538C. The next 28.5 h were spent waiting on weather (WOW). The ship's heave (the maximum peak to peak heave over a 10 min interval) remained above 5 m until 1930 h on 12 May. At that time, we were able to resume rig floor operations.

Hole U1538D started at 2145 h on 12 May. Cores 382-U1538D-1H through 14H penetrated from the seafloor to 126.4 mbsf and re-

covered 114.6 m (91%). A formation temperature measurement was made while taking Core 12H. We had to stop coring to leave time to raise the drill string and secure the BHA for transit before stormy weather and high seas arrived early on 14 May. The last core came on deck at 1240 h on 13 May, and the bit was clear of the seafloor at 1430 h, ending Hole U1538D.

Having completed operations at Site U1538, we started the sea voyage to Punta Arenas, Chile, at 0000 h on 14 May. We had to leave Site U1538 earlier than originally planned because of deteriorating weather and sea conditions. Our course took us over the first two sites of the expedition (Sites U1534 and U1535), but headwinds slowed our progress to the extent that no time was available for further coring at Site U1534 (average speed was 3.7 kt on 14 May). However, we had time to make a single-line 3.5 kHz subbottom profiler survey east–west over the two sites on 17 May. We then continued to Punta Arenas. The 1088 nmi voyage took 5.6 days at an average speed of 7.7 kt. The pilot boarded at 0754 h on 19 May near the entrance to the Strait of Magellan, and we arrived at Prat Terminal in Punta Arenas, Chile, at 1530 h, ending Expedition 382.

Principal results

The lithology at Site U1538 is divided into two units (Figure F8). Lithostratigraphic Unit I (seafloor to 308 mbsf) consists of dark greenish gray silty clay–rich and silty clay–bearing diatom ooze alternating with diatom-rich and diatom-bearing silty clay (Figure F10). Some intervals of diatom ooze are over 20 m thick. IRD dropstones are small and rare on the split core surface, but more of them are evident in the core X-ray images, illustrating the value of the X-ray technique. From ~38 to 39 mbsf and from ~122 to 124 mbsf, dispersed diagenetic carbonate occurs in layers >50 cm thick. Similar diagenetic carbonate was found at Sites U1536 and U1537 as discrete millimeter- to centimeter-thick layers. Unit II is distinguished by increased compaction (semilithified silty clay and diatomite) and an increase in the abundance, size, and lithology of gravel- and pebble-sized IRD. In Unit II, distinct layers of enriched IRD were observed. Thin (<5 cm) diagenetic carbonate layers and lenses are present at ~351 and ~504 mbsf.

Diatoms and radiolarians were found in all core catcher samples, and additional samples were collected from selected depths (Figure F11), allowing us to identify 75 biostratigraphic events. Based on these events, we estimate sedimentation rates of ~22 cm/ky between the seafloor and 432 mbsf (~1.75 Ma), followed by ~2.5 cm/ky from 432 to 469 mbsf (~3.6 Ma), and finally ~12 cm/ky from 447 mbsf to the base of Hole U1538A. The deepest sample revealed that sediments recovered at this site date back to ~4.2–4.7 Ma. Overall preservation is relatively good in all samples, and only limited reworking of diatoms was detected.

Palynomorphs (dinocyst, acritarch, prasinophytes, pollen, spores, copepod, and fungi remains) were identified in all 30 samples processed for palynology. Dinocyst diversity is generally low with varying abundance throughout. Reworked pollen and spores were found throughout, whereas *in situ* specimens were only detected in certain intervals. Samples for the analysis of sedaDNA were also collected from the upper 84 m and are anticipated to record the Holocene and the last interglacial.

Paleomagnetic investigations at Site U1538 involved measurement of the natural remanent magnetization (NRM) of archive halves from all holes before and after demagnetization to a peak alternating field (AF) of 15 mT. All discrete samples from Site U1538 were subjected to AF demagnetization at 5, 10, and 15 mT to verify the archive-half measurements. Directional data agree well between

archive-half and discrete samples in APC and HLAPC cores. XCB cores often have a large amount of scatter in archive-half measurements, making those data difficult to interpret. However, discrete samples taken from large, intact pieces in the XCB cores appear to reliably record the polarity and can be used for magnetostratigraphy. We identified polarity zones from the Brunhes Chron (C1n) to the middle Gilbert Chron (C3n; tentatively the Sidufjall Subchron, 4.8–4.9 Ma) (Figure F9). However, not all polarity zones of the GPTS12 were identified, including Subchrons C2An (Gauss Chron) and C3n (middle Gilbert Chron). This is probably a result of discontinuous recovery and the slow sedimentation rate between ~1.75 and 3.6 Ma. The presence of the iron sulfide mineral greigite is suspected in some intervals.

In Hole U1538A headspace gas analyses, methane is the dominant hydrocarbon and is present only in low concentrations (2.0–5.4 ppmv), whereas ethane (C_2H_6) concentration is below the detection limit. Also, we found relatively low inorganic carbon (0.2–0.75 wt%), TN (0.04–0.23 wt%), and $CaCO_3$ (0.02–4.2 wt%) contents.

Interstitial water geochemical data were generated on 40 samples from Hole U1538A to 671 mbsf. The inorganic geochemical data from Site U1538 fall at the more extreme end of interstitial water properties observed during Expedition 382. Although the cores are not methanogenic, sulfate is depleted below 60 mbsf. Minor dissolved sulfate concentrations were detected below 500 mbsf ($SO_4 < 1.5 \mu M$). Significant organic matter degradation is apparent in the upper 120 m, with elevated phosphate concentrations and a pronounced Br/Cl gradient. Barite dissolution is evident at depths without resolvable sulfate between 60 and 445 mbsf, resulting in the highest dissolved barium concentrations of the Expedition 382 sites. The spatial evolution of Ba enrichment also differs from the previous sites in that Ba concentrations are above 600 μM from 250 to 500 mbsf before more rapidly decreasing within a few tens of meters underneath this zone. Dissolved Mn enrichment is observable in the uppermost 50 m at Site U1538, followed by Fe enrichment below, suggesting Mn-Fe oxide mobilization in the upper 260 m. Calcium concentrations decrease significantly from 11.7 to 1.9 mM in the sulfate reduction zone, recover to stable intermediate concentrations throughout the upper half of Site U1538, and become enriched deeper in the hole. Together with equally increasing B and Sr concentrations, this enrichment suggests increasing exchange reactions between the dissolved and solid phases in the older sediments. Although chloride concentrations decrease with depth together with interstitial water salinity, sodium concentrations increase. Authigenic clay formation is apparent in decreasing potassium and magnesium concentrations in the deeper parts of Site U1538. In comparison to the interstitial water profiles from Scotia Sea Sites U1536 and U1537, these data highlight the more exotic character of the profiles from Site U1538.

The Whole-Round Multisensor Logger (WRMSL) was used to measure GRA bulk density, MS, and *P*-wave velocity at 1 cm intervals, which is a higher spatial resolution than the 2.5 cm spacing used at the other sites. The core section halves were scanned on the X-ray imager. Discrete measurements collected in section halves included at least 3 *P*-wave velocity measurements per core and 150 thermal conductivity measurements (2–3 per core). Moisture and density measurements were made on 145 discrete samples from Hole U1538A. Light reflectance and MS point measurements were collected at 1 cm intervals on section halves from Cores 382-U1538A-1H through 35F and at 2 cm from the remaining cores from this site. Formation temperature measurements taken with

Cores 6H, 9H, 12H, 15H, and 18H with the APCT-3 yielded temperatures of 4.85°–13.8°C.

In the upper 220 m, physical property records have 10–20 m long amplitude cycles (e.g., bulk density values cycle from lows of 1.2 g/cm³ to highs of 1.7 g/cm³, and NGR values cycle from lows of 10 count/s to highs of 40 count/s) (Figure F9). Below 220 mbsf, the cycles have shorter periods and slightly lower amplitudes but are superimposed on longer scale baseline changes. An exception is the high-amplitude MS cycles from 310 to 450 mbsf underlain by low baseline MS values from 450 to 550 mbsf. The overall cycles and patterns are comparable in amplitude and thickness to those seen at Sites U1536 and U1537 in Dove Basin to the southwest. Downhole sediment compaction is reflected in increased density and *P*-wave velocity with depth. An increase in baseline values of NGR, density, and *P*-wave velocity is associated with a change in the seismic facies at ~370 mbsf.

A spliced record for Site U1538 was constructed for the uppermost 124 mcd from Holes U1538A, U1538C, and U1538D. It is based predominantly on correlation of MS and GRA density data.

Integrating the core stratigraphy with the seismic profiles using discrete sample caliper velocity data to transform depths to two-way traveltime, we identified Reflector a' at ~77 mbsf in Core 382-U1538A-10H. This reflector separates the higher variability of physical properties above from lower variability below. A more detailed acoustic to core correlation based on data obtained with the TOPAS subbottom profiler was possible in the upper ~135 m of Hole U1538A (Figure F4). Accordingly, besides identifying Reflector a', additional reflectors could be correlated with changes in physical properties at ~13, 45, and 90 mbsf. Assuming a depth to two-way traveltime conversion based on the caliper velocity measurements, Reflector a was identified by a decrease in GRA density at ~371 mbsf in Core 43X. Reflector b was observed at ~513 mbsf in Core 57X. We were also able to identify Reflector c by a sharp increase in WRMSL *P*-wave velocity at ~575 mbsf in Core 64X. Determining ages for these basin-wide reflectors will be important in the assessment of the geodynamic history of the Scotia Sea.

Preliminary scientific assessment

Operational considerations related to the science objectives

The operations plan for Expedition 382 laid out in the *Scientific Prospectus* (Weber et al., 2018) considered the maximum amount of possible coring if no hazards occurred. However, operating east of the Drake Passage in the center of Iceberg Alley and near the Antarctic sea ice edge involves a number of potential hazards, in particular icebergs and bad weather. Indeed, high waves and consequent high ship heave were a major challenge when drilling in both the Dove and Pirie Basins and prohibited coring for 62 h (2.6 days) while waiting on weather. This calculation, however, does not include the time necessary to reenter a hole. An overview of the actual operations times and activities is given in Figure F6.

Another potential hazard was sea ice. Detailed analyses of daily sea ice maps for the years 2012–2017 indicated that Antarctic sea ice advances north during the months of April and May. Reassessment during coring operations at the Subantarctic Front sites showed that the sea ice edge was ~300 km south of Site U1536 in early April 2019 and moving north at more than 10 km per day. Hence, we changed plans and went directly to the southernmost Dove Basin rather than drilling first in the more northerly located

Pirie Basin, knowing that we would need at least 16 operational days at Site U1536 and ~12 days to drill a second site in Dove Basin. In the end, the sea ice edge did not approach close enough to influence our operations in Dove Basin.

The biggest potential threat and also the most unpredictable were icebergs. Naturally, with the main focus on reconstructing Antarctic iceberg flux in the center of Iceberg Alley, this hazard was expected. However, icebergs do not break off the Antarctic continent at a preferred time of the year, nor do they travel in a seasonal pattern north. The fact that there is also interannual variability in iceberg shedding made it impossible to predict iceberg concentrations for the time of the drilling operations in Iceberg Alley before the expedition. Also, the real-time information available from satellite data during the expedition was not of sufficient resolution to allow for detailed planning. Therefore, during the expedition, iceberg observation relied almost entirely on visual and radar observations. This limited our ability to plan ahead more than 10–15 h when icebergs were in radar range and required us to adapt our operations continuously. When icebergs entered a red zone (less than 3 miles from the ship or less than twice the time required to raise the bit clear of the seafloor), drilling operations stopped and the drill string was pulled up to 50 mbsf. When an iceberg entered a 1 mile radius and was heading toward the ship, we pulled out of the hole—this happened three times. During Expedition 382, we spent 159 h (5.3 days) waiting on ice (WOI), which does not include the time required to reenter and return to the bottom of a hole (Figure F6). Additionally, there was a 2 day delay related to strong winds that prohibited fueling at the begin of the expedition.

Preliminary assessment of scientific objectives

The scientific objectives of Expedition 382 center around four thematic investigations of the late Neogene: the ice mass loss history of the AIS, the regional atmospheric circulation and dust transport, the flow of deep water through the Drake Passage, and the AAIW transport beneath the Subantarctic Front. Based on a preliminary analysis of the core material and data collected during the expedition, we assess the nine specific objectives that were formulated in the *Scientific Prospectus* (Weber et al., 2018).

The majority of these objectives require continuous and high-resolution sediment records for their investigation. We obtained continuous records reaching back to the Pliocene at three sites (Figure F8). The number of identifiable biostratigraphic events (first and last occurrences of microfossil species; Figure F11) ranged between 12 (Site U1534) and 81 (Sites U1536 and U1537), and the detection of all magnetic reversals since the mid-Pliocene indicate robust stratigraphic continuity at the Scotia Sea sites for this time interval. Sedimentation rates are also generally high, ranging from 5 to 25 cm/ky depending on the time interval and site (Figure F5). This means that the majority of high-resolution physical and optical properties measured during the expedition (Figure F5) have a temporal resolution of several decades to a few centuries. This combination of stratigraphic continuity and highly resolved sediment records obtained for the first time in an important yet undersampled region with respect to past changes in AIS dynamics and Southern Ocean paleoceanography is a very encouraging first result of our expedition. It is also an important prerequisite to fulfilling the majority of our scientific goals.

1. Reconstructing the history of AIS calving through time

A reconstruction of AIS calving through time requires continuous records of IBRD. Using the new X-ray imager, we took continuous images along all core sections, resulting in an impressive

collection of over 25,000 images. Initial counting of IBRD on these images confirmed that previous glacial terminations document AIS mass loss events similar to the last deglaciation (Weber et al., 2014). This is very promising for postcruise research on glacial terminations, over glacial–interglacial cycles, and with respect to the long-term, Pliocene–Pleistocene ice mass loss history. Also, IBRD was observed and documented in the lithologic descriptions at all sites. The X-ray images will be used to direct the postcruise sampling of IBRD-rich layers to study provenance (the potential source regions of IBRD and the icebergs that carried it). The available core material will also be used for detailed grain size studies to obtain a further indicator of IBRD variability. Equipped with such core material and data sets, we are confident we will be able to provide the first qualitative but detailed and long-term history of AIS mass loss via calving of icebergs.

2. Chronology and dust–climate couplings

The correlation of marine dust records to well-dated ice-core dust records has been a major step forward in developing high-resolution Southern Ocean chronologies (see **Scientific objectives**). According to Weber et al. (2012), MS should provide such high-resolution chronological control at least for the late Pleistocene. At all our Scotia Sea sites, highs in MS (i.e., more terrigenous material) occur consistently during glacial periods for at least the last 3 My. Because our sites extend much further back in time than Antarctic ice cores, we should have the ability to study dust–climate couplings also during the Pliocene–Pleistocene. At Site U1538, the MS signal is likely influenced by diagenetic overprint in deeper parts of the sedimentary sections. Further dust proxies such as Ca and Fe counts obtained through XRF scanning will be investigated postcruise for their stratigraphic potential. Also, higher values in color component b* (i.e., more yellowish colors) are systematically associated with diatom-rich interglacial periods in the late Pleistocene. This relation, first established by Sprenk et al. (2013), will be tested postcruise for its stratigraphic and paleoclimatic application.

An important stratigraphic approach will be to correlate the Scotia Sea sites before putting them on a common timescale. Variations in NGR provide reliable records that can be correlated in great detail between our sites in the given biomagnetic stratigraphic control (Figure F9). Also, GRA density records show distinct glacial–interglacial variability that were also used for site-to-site and hole-to-hole correlation. Time control provided by bio- and magnetic stratigraphy, possibly in combination with oxygen isotope stratigraphy, should allow us to put sediment physical property data older than the range of ice cores onto a robust chronology.

3. AIS mass loss and regional-to-global sea level change

This scientific objective requires the identification of glacial terminations that led to warmer-than-present times such as MIS 5, 11, 31, and so on. Our preliminary assessment of the core material shows that we have recovered a sedimentary record of these important times, and we should therefore be able to study the relation of AIS mass loss and global sea level change across terminations in great detail. However, this specific objective also requires evaluating IBRD records in the context of near- and far-field geophysical sea level effects, such as glacial isostatic adjustment (see **Scientific objectives**) and will be pursued postcruise in collaboration with sea level modelers.

4. Water mass changes and ocean thermal forcing

Because we obtained long and continuous sediment records in all three key regions (Subantarctic Front, Pirie Basin, and Dove Basin; Figure F1), we should be able to address past changes in the

AMOC and ocean temperature over the Pliocene–Pleistocene. However, these studies rely on a variety of biogeochemical investigations that will be conducted postcruise.

5. *Provenance of glacially eroded material*

The documentation of IBRD in X-ray images and sedimentologic descriptions provides IBRD concentration records to identify IBRD-rich intervals that are suitable for postcruise research. Sand- to pebble-sized IBRD will be characterized using a variety of mineralogical, geochemical, and isotopic techniques for the purpose of establishing its provenance, which will provide valuable knowledge of the sectors in Antarctica that previously underwent sustained ice mass loss via calving.

6. *Glacial–interglacial sea ice shifts and the carbon cycle*

A prerequisite to conducting investigations on past sea ice shifts and potentially associated changes in the carbon cycle was to obtain sediment records from different latitudes. We obtained sediment records that stretch from the Subantarctic Front in the north to Dove Basin in the south (spanning $\sim 6^\circ$ latitude). Also, the preservation of microfossils that can be used for such reconstructions is very good at our sites. In addition, we know rather precisely where glacial–interglacial transitions occur in the sediment record at least for the late Pleistocene. Therefore, we are optimistic that we will be able to reconstruct glacial–interglacial sea ice shifts.

7. *ACC flow speed reconstruction*

Core material has been obtained to address this topic at three different locations. At the Subantarctic Front sites, grain size analysis and orientation of the magnetic fabric will be used as proxy records of past changes in flow speed will inform us about changes in the AAIW and the position of the front. In Pirie Basin, cores will provide information on changes associated within the core region of the ACC. In the southern Dove Basin, the potentially higher influence of northward-flowing WSDW can be studied.

8. *Pliocene–Pleistocene ice age climate cycles and underlying causes*

All our sites, from the Subantarctic Front to the southern Scotia Sea, will provide material that can be used to study Pliocene–Pleistocene climate evolution in the Southern Hemisphere and allow us to test hypotheses about how orbital variations control global climate and ice volume. As described above, most of the physical property data sets obtained during the expedition have a sufficient resolution and show distinct cyclicity that is likely to be related to orbital changes. Initial bio- and magnetostratigraphy suggest that continuous high sedimentation rate sections were recovered.

Scientific objectives for the Subantarctic Front sites

Recovering the MPT was highly anticipated at these sites. However, biostratigraphic markers and the absence of a magnetic reversal indicate that the silty deposits above Reflector B at Site U1534 are no older than 0.78 Ma. Mid- to early Pleistocene sediments were recovered at deeper Site U1535, reflecting the onlapping nature of the uppermost sediment unit onto Reflector B; however, sediments were more condensed than at Site U1534, and abrupt lithologic boundaries suggest multiple hiatuses at this more distal site. Despite providing a shorter time interval, late Pleistocene sedimentation at Site U1534 appears to be continuous and on the order of 10 cm/ky, and it would be the preferred site for high-resolution paleoceanographic study through the late Pleistocene. Although time did not allow three full-depth holes to be cored, a fully spliced record was achieved above Reflector B at Site U1534. The benthic foraminifer *Uvigerina* spp. is persistent throughout the late Pleistocene and will

yield stable isotope and elemental ratio records to constrain an age model and reconstruct water mass properties at the site. Nannofossil-rich foraminifer oozes characterize MIS 5e and 11. There is high potential to yield multiproxy, millennial-scale paleoceanographic reconstructions throughout the last 700 ky at Site U1534.

Paleomagnetic and biostratigraphic markers indicate that the 170 m of silty clay recovered between Reflectors B and C at Site U1534 is of late Pliocene age. Seismic interpretation, paleomagnetism, and biostratigraphic markers indicate that sediments beneath Reflector B are contemporaneous at both Sites U1534 and U1535. With sedimentation rates as high as 40 cm/ky and prominent cyclicity evident in the physical properties, these late Pliocene sediments will provide key insights into the evolution of the PIS and dynamics of the cold water route while Northern and Southern Hemisphere glaciation was intensifying.

Assessment of seismic reflectors

During Expedition 382, we drilled through three structural seismic reflectors at the Subantarctic Front sites and through four of the six reflectors that were previously used to classify and describe the long-term history of the Scotia Sea. However, none of these reflectors were dated before. Postcruise work will determine the ages of the reflectors, compare them among the sites, and reevaluate the long-term depositional history.

Education and outreach

Expedition 382 had a team of two Education and Outreach Officers: Marlo Garnsworthy, an author, illustrator, editor, photographer, and science communicator from the United States/Australia, and Lee Stevens, a videographer, animator, photographer, and science communicator from the United States.

The officers communicated the goals and science of the expedition in simple terms and creative ways to students and the general public around the world via video broadcasts to schools and universities, social media, blog posts on the *JOIDES Resolution* website, Facebook live events, and articles published on external sites. The officers also developed educational material for students of all ages and the general public, such as animations, videos, infographics, illustrations, and two books (underway).

Professional development of the science team was also a significant focus of the Education and Outreach team, who provided science communication and social media workshops for the scientists and technicians on board and worked to develop scientists as writers and bloggers through developmental/substantive editing and consultation.

Live broadcasts

Via live video broadcast, the Education and Outreach and Expedition 382 science teams interacted with students all over the world to share the goals, techniques, and progress of the expedition. Audience sizes ranged between ten and several hundred students, and media were present at multiple events, generating further interest via newspaper articles and other media.

A typical session included an introduction to IODP, the ship, and the Expedition 382 goals, followed by a tour that included the derrick, rig floor, catwalk, and laboratories. Students met scientists in each laboratory and learned about the core flow process and specific areas such as physical properties, sedimentology, paleomagnetism, and biostratigraphy in relation to our expedition

research and goals. Tours ended with an extensive question and answer session that usually included at least one scientist.

Over the course of the expedition, the Education and Outreach officers completed 54 broadcasts. Students ranged from third grade to postgraduate university level, and broadcasts reached 2,655 people in 11 countries and 6 languages. Demand for these events was very high with between 2 and 4 broadcasts on most weekdays and occasional weekend events. Numerous broadcasts were recorded by teachers for future use.

Attendees received educational material and helpful links from <https://joidesresolution.org> prior to the broadcasts and a follow up and survey link after the broadcast.

Expedition activities

The Education and Outreach team produced an animation (5 Things You didn't Know About Icebergs), a video (Drilling for DNA), several pieces of informational artwork about polar ice (with scope to expand this project beyond the expedition), an infographic about how icebergs are formed and IBRD, and research and preliminary writing for two books (one for young adults and one for elementary-aged children), plus an additional video covering the scope of the expedition. A total of 29 blog posts were generated covering science, the expedition, shipboard life, and more.

Because of the loss of the imaging expert at the beginning of the expedition, the Education and Outreach team was requested to produce photographs during the expedition as an additional duty, and they created a significant photographic record of operations on the drill floor, catwalk, laboratories, and other parts of the expedition.

Social media

The Education and Outreach team maintained four websites during the expedition: the *JOIDES Resolution* Expedition 382 web page and blog, Facebook, Twitter, and Instagram, with daily postings on Twitter and Facebook. They shared responsibilities for these tasks and had a total of 241 posts with 53,969 shares, likes, or replies. Twitter engagement increased by 46% with 404,900 total impressions during the expedition. Facebook post reach increased by 60% from the previous 60 days with over 229,000 total impressions. The Education and Outreach team also hosted 6 Facebook Live events with 591 engagements and did a Twitter AMA thread on Earth Day.

Media

This expedition generated significant media attention with 13 external newspaper and magazine articles, 2 TV appearances, and 1 upcoming radio interview so far.

Videographer Stevens filmed and edited hours of footage for the CBS *This Morning* interview of Maureen Raymo for a piece featured on Earth Day and other footage to be used by German filmmakers MobyDok for a documentary about sea level rise featuring Mike Weber. Garnsworthy wrote articles for *EGU Cryosphere*, *Ocean-bites*, and *Envirobites* and is booked for postcruise presentations and talks.

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Table T1. Expedition 382 hole summary. mbsl = meters below sea level, DSF = drilling depth below seafloor. — = no data.

Hole	Latitude	Longitude	Water depth (mbsl)	Total cores (N)	Cored interval (m)	Recovered length (m)	Recovery (%)	Drilled interval (m)	Penetration DSF (m)	Time on hole (h)	Time on site (days)
U1534A	53°11.3777'S	58°45.6479'W	605.09	45	265.3	270.84	102.09	1	266.3	59.3	—
U1534B	53°11.3763'S	58°45.6298'W	605.07	1	1.5	1.51	100.67	—	1.5	4.1	—
U1534C	53°11.3865'S	58°45.6296'W	606.27	19	159.2	159.55	100.22	—	159.2	21.6	—
U1534D	53°11.3879'S	58°45.6259'W	606.07	3	28.5	29.64	104.00	—	28.5	3.6	—
Site U1534 totals:				68	454.5	461.54		1	455.5		3.69
U1535A	53°11.4899'S	58°38.6048'W	646.86	17	117.3	112.47	95.88	—	117.3	23.3	—
Site U1535 totals:				17	117.3	112.47			117.3		0.97
U1536A	59°26.4602'S	41°3.6596'W	3219.52	53	354.4	364.27	102.78	—	354.4	141.6	—
U1536B	59°26.4608'S	41°3.6399'W	3220.06	24	224.1	230.71	102.95	2	226.1	31.4	—
U1536C	59°26.4604'S	41°3.6191'W	3222.05	34	208.0	187.42	90.11	144	352.0	71.5	—
U1536D	59°26.4501'S	41°3.6198'W	3221.72	1	6.9	6.86	99.42	—	6.9	13.0	—
U1536E	59°26.4496'S	41°3.6392'W	3219.08	32	305.4	110.55	36.20	340	645.4	211.2	—
Site U1536 totals:				144	1098.8	899.81		486	1584.8		19.53
U1537A	59°6.6458'S	40°54.3725'W	3712.91	31	264.0	268.86	101.84	—	264.0	64.1	—
U1537B	59°6.6583'S	40°54.3516'W	3712.01	1	7.4	7.40	100.00	—	7.4	3.4	—
U1537C	59°6.5197'S	40°54.0870'W	3713.61	2	15.3	11.92	77.91	—	15.3	3.6	—
U1537D	59°6.6597'S	40°54.3677'W	3713.09	50	354.3	348.94	98.49	—	354.3	103.4	—
Site U1537 totals:				84	641.0	637.12			641.0		7.27
U1538A	57°26.5248'S	43°21.4691'W	3130.58	75	676.0	476.43	70.48	—	676.0	157.7	—
U1538B	57°26.5368'S	43°21.4454'W	3129.81	1	8.5	8.49	99.88	—	8.5	3.4	—
U1538C	57°26.5387'S	43°21.4521'W	3130.21	12	105.9	98.92	93.41	—	105.9	19.2	—
U1538D	57°26.5335'S	43°21.4723'W	3130.40	14	126.4	114.60	90.66	—	126.4	27.6	—
Site U1538 totals:				102	916.8	698.44			916.8		8.66
Expedition 382 totals:				415	3228.4	2809.38		487	3715.4		

Figure F1. Map of Scotia arc. Sites U1534 and U1535 are under the Subantarctic Front. Scotia Sea sites are in Dove Basin (U1536 and U1537) and Pirie Basin (U1538). Light blue dotted line and open arrows = Iceberg Alley after Anderson and Andrews (1999), large curved gray arrow = main wind direction of Southern Hemisphere westerlies (SHW), light blue solid arrows = Antarctic Circumpolar Current (ACC) flow direction, purple arrows = Weddell Sea Deep Water (WSDW) exit route (Maldonado et al., 2003), yellow dotted line = Subantarctic Front, yellow dashed line = polar front, green dotted line = Patagonian coastline during LGM (Iriondo, 2000), white dotted line = limit of Patagonian Ice Sheet (PIS) at LGM (Hein et al., 2010), white dashed line = winter sea ice limit. Satellite bathymetry is shown (Weatherall et al., 2015).

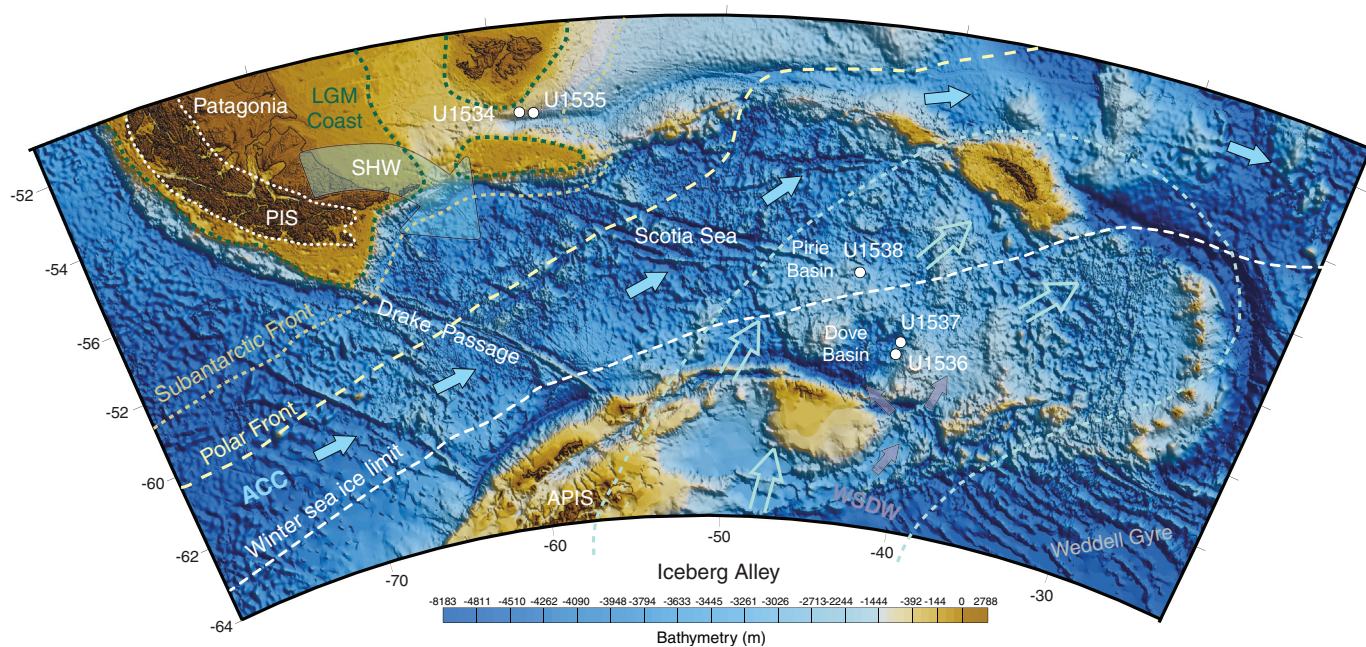


Figure F2. Top: Seismic profiles for Subantarctic Front drift Sites U1534 and U1535 along Seismic Line SGF193_107 (Koenitz et al., 2008). Main discontinuities are marked by Reflectors A (top of Subunit 2A), B (top of Subunit 1B), and C (top of Subunit 1A). Bottom: Seismic profile for Site U1538 in Pirie Basin (Seismic Line SCAN 13/13), with seismic units from Pérez et al. (2017).

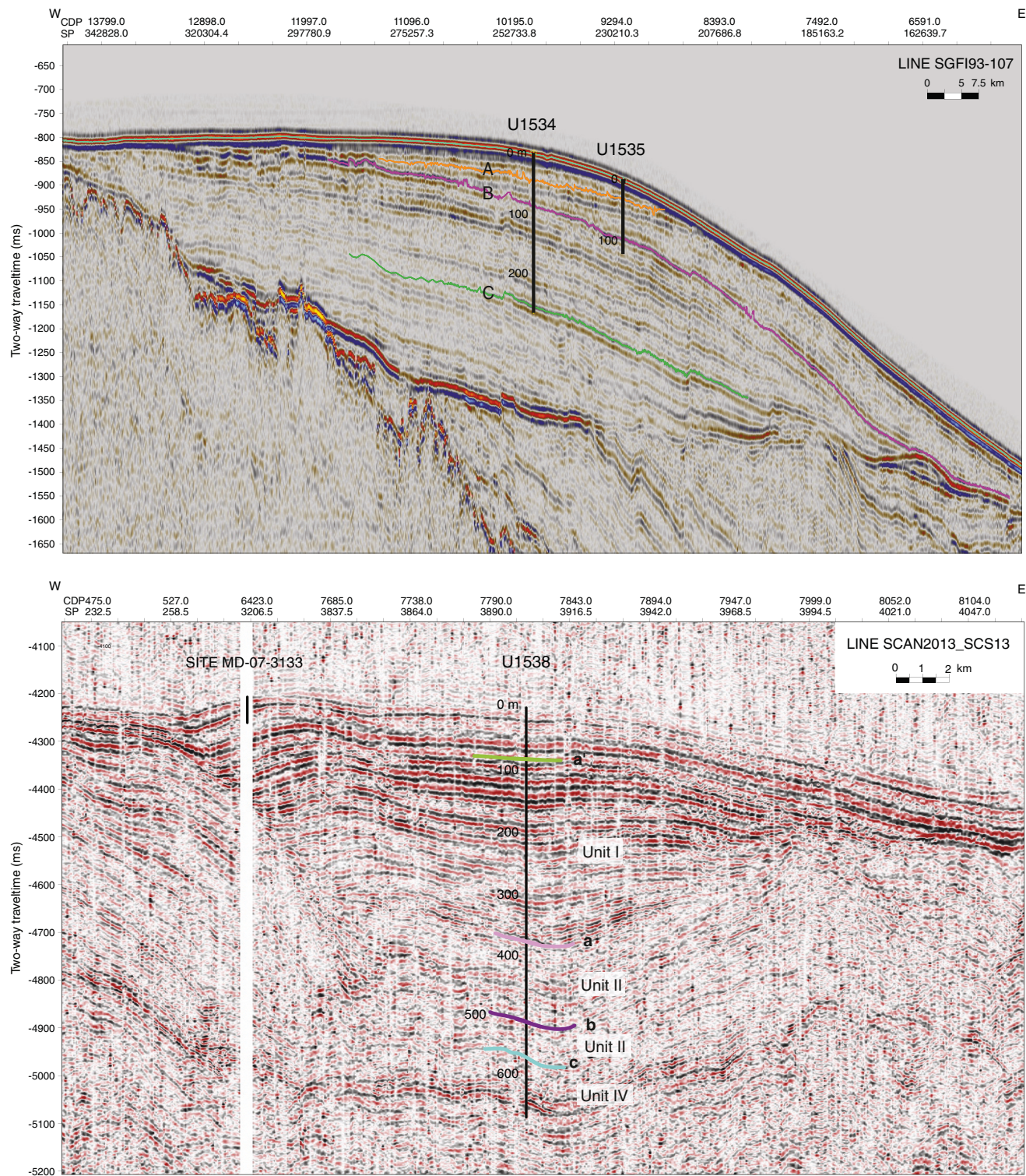


Figure F3. MCS profiles from Dove Basin (Maldonado et al., 2004). Top: Site U1536 along Line SCAN2004-L10. Bottom: Site U1537 along seismic Line SCAN2004-L07. Seismic units from Pérez et al. (2017).

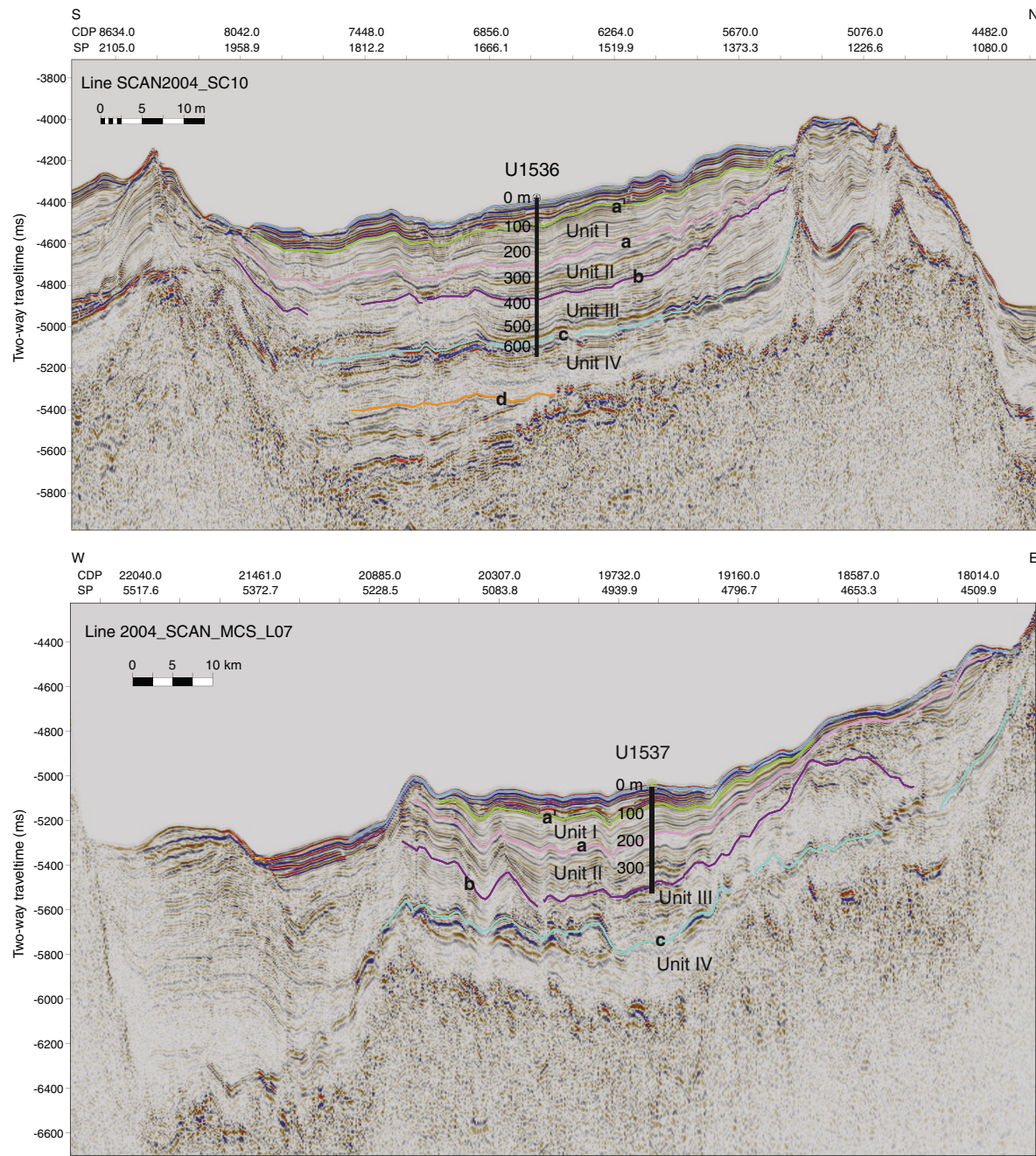


Figure F4. TOPAS profiles along three sites in Dove and Pirie Basins showing high-resolution seismic record formed by layered reflectors of medium–high acoustic amplitude forming wavy or mounded morphologies. MTDs are frequent in Dove Basin, interbedded with the layered seismic packages.

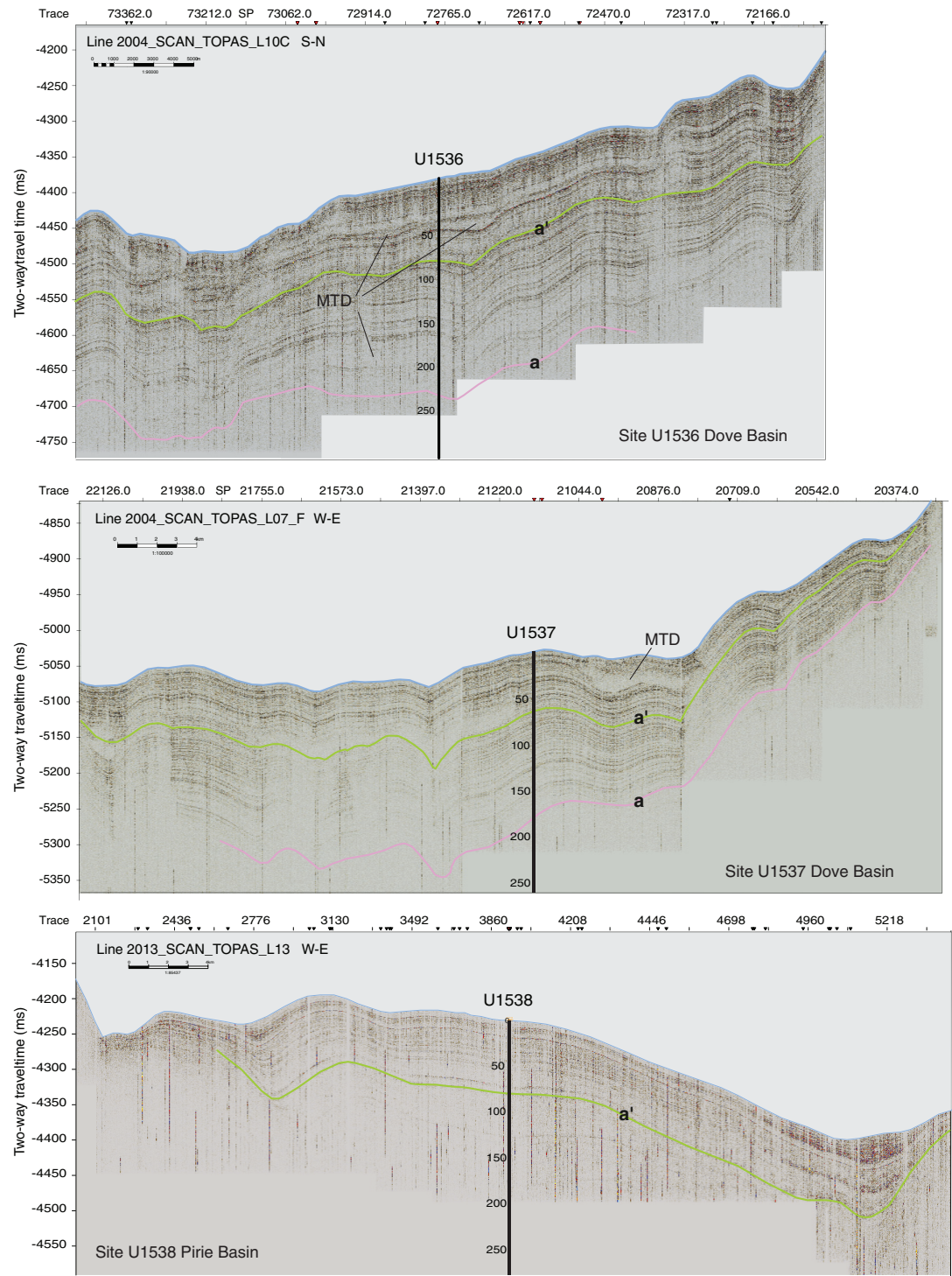


Figure F5. Age-depth relationship for the last 6 My, Expedition 382. Except for Site U1534, all sites show continuous deposition in this interval based on preliminary biomagnetostratigraphy. Site U1536 extends to middle Miocene (not shown in figure). Sedimentation rates are highest in Pirie Basin (Site U1538), decrease slightly to southern Dove Basin (Sites U1536 and U1537), and lower substantially to northern Subantarctic Front (Sites U1534 and U1535). Ages according to GPTS12 (Gradstein et al., 2012). Lower curves are $\delta^{18}\text{O}$ records (Lisiecki and Raymo, 2005). Colored bars in lower panel indicate temporal extent of mid-Pleistocene transition, Northern Hemisphere ice sheets, and mid-Pliocene warm period.

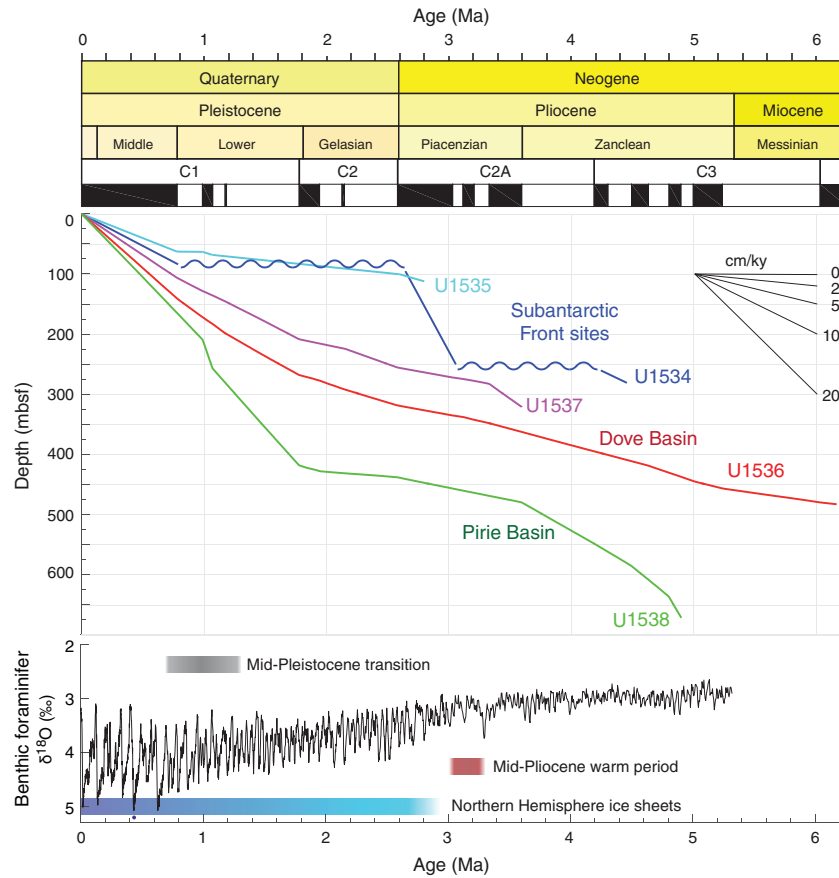


Figure F6. Operations schedule, Expedition 382. We tracked 153 icebergs over the course of the expedition. We had a total of 56 operational days.

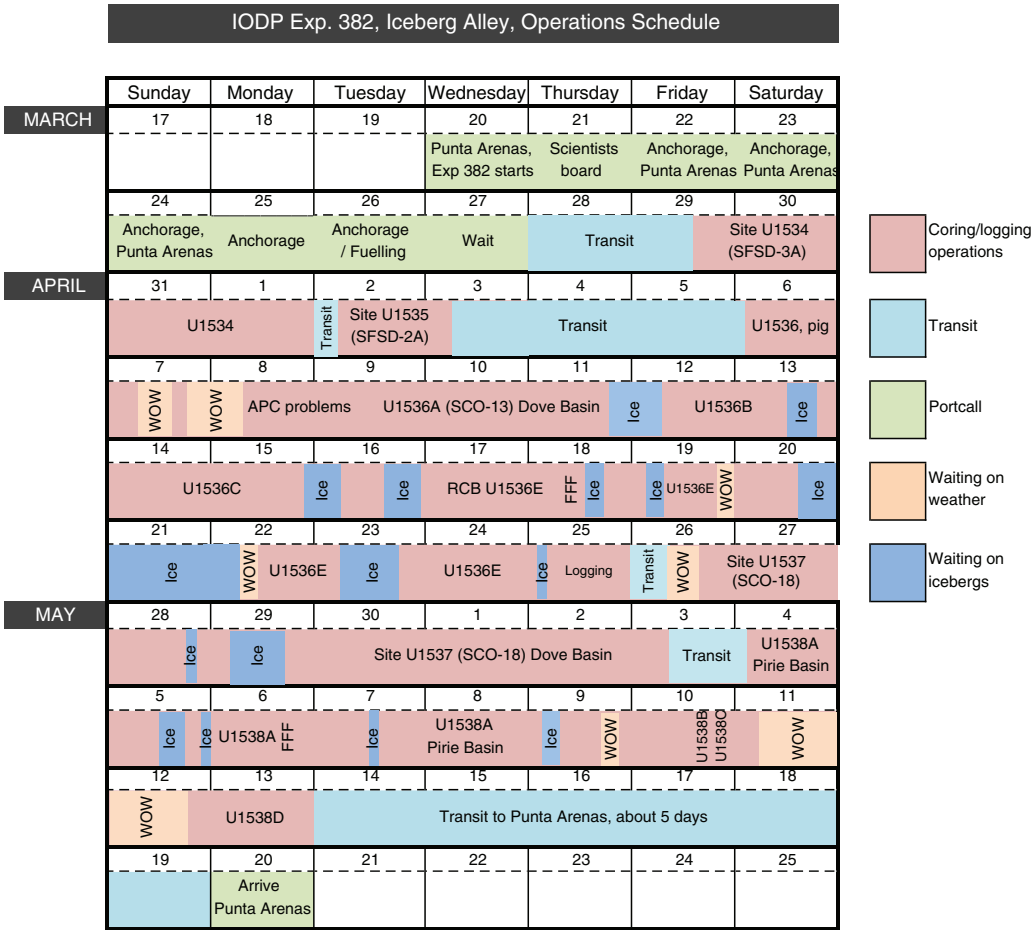


Figure F7. Drilling plan, Expedition 382. After transiting from Punta Arenas to Subantarctic Front, Sites U1534 and U1535 were drilled over 5 operational days at beginning of expedition. After transiting to southernmost working area at Dove Basin, Site U1536 was drilled. It took 21 days to achieve APC/HLAPC and RCB coring as well as logging. Site U1537 was then drilled in 7 days. After transiting north to Pirie Basin, Site U1538 was drilled over 10 days. A total of 3228 m over 18 holes was cored during Expedition 382 with an overall core recovery of 2810 m.

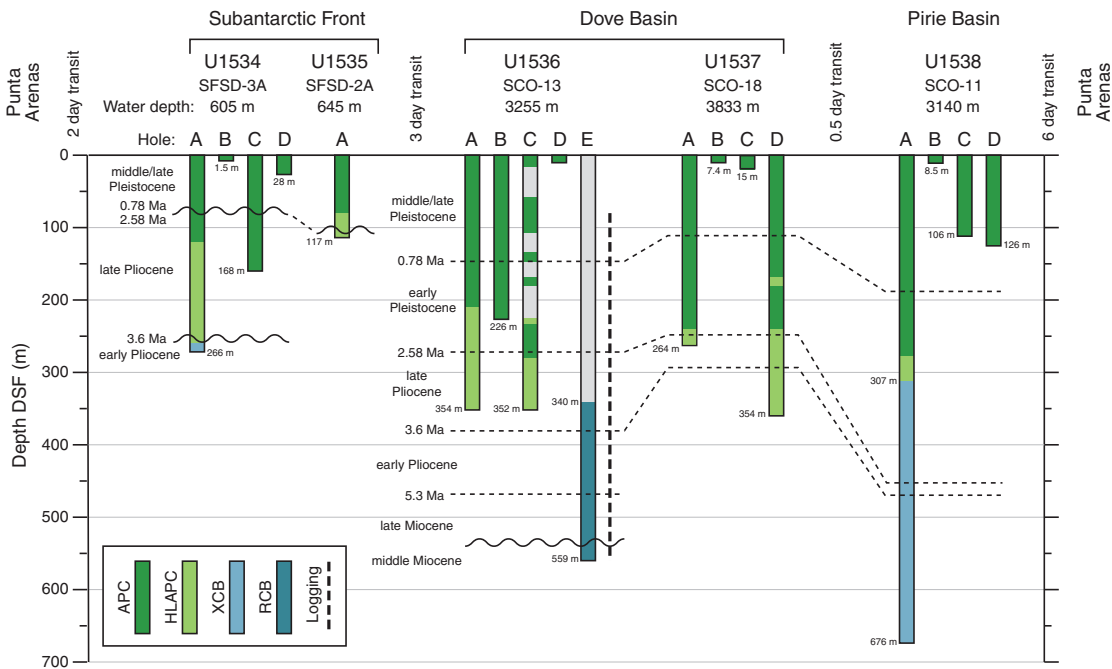


Figure F8. Principal lithology, Expedition 382. Hole U1536E reached middle Miocene strata. All other holes cover Pliocene–Pleistocene sediments. Lithology is dominated by silty clays with interbedded diatom ooze (Scotia Sea) or nannofossil ooze (Subantarctic Front). APC/HLAPC core recovery was almost complete.

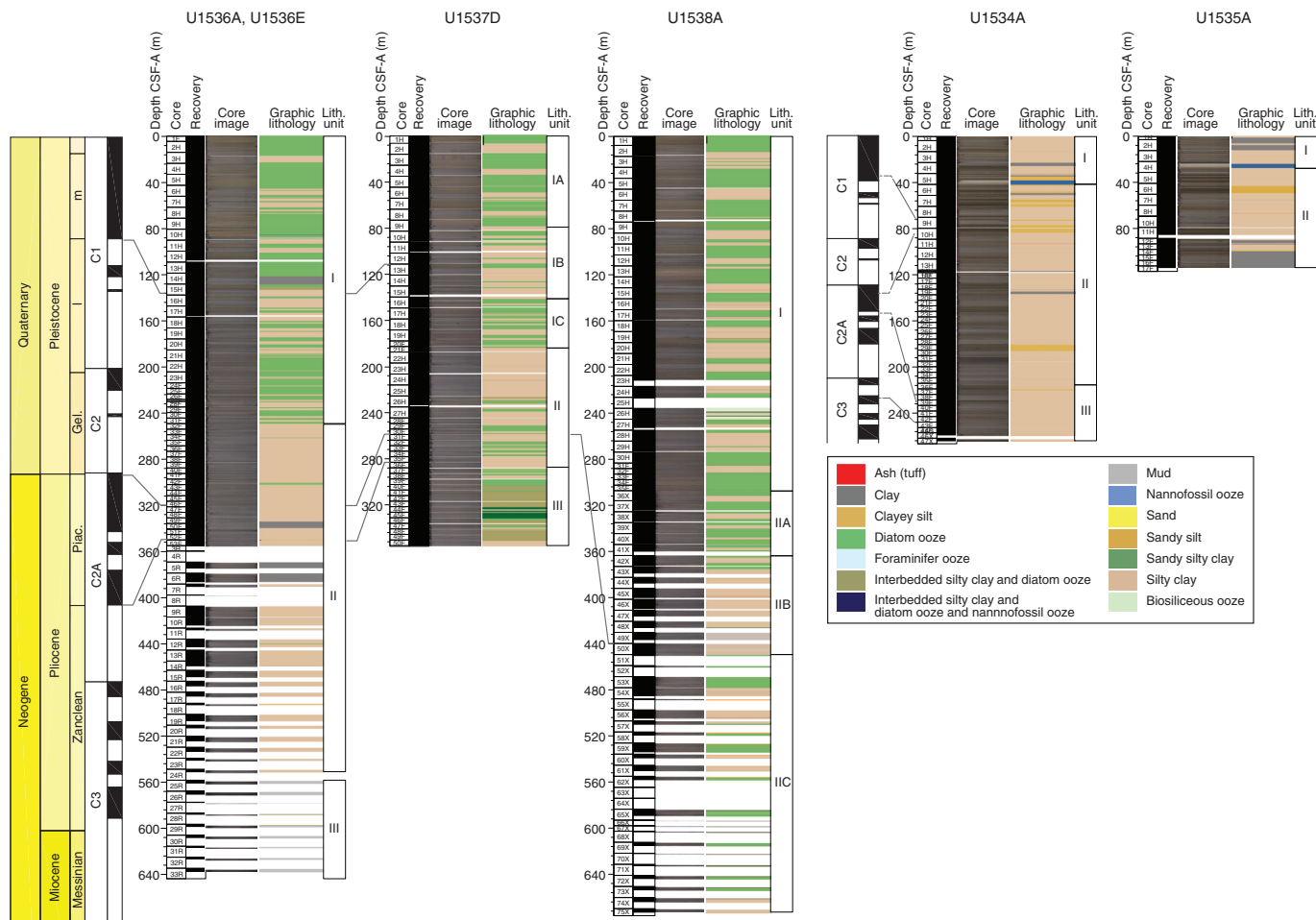


Figure F9. Lithology, paleomagnetic reversal pattern, and NGR measurement variability, Expedition 382. Note that Sites U1536 and U1537 are plotted on the composite depth scale (core composite depth below seafloor, Method D [CCSF-D]), whereas Holes U1538A and U1534A are plotted on the mbsf scale because the splice covers only the upper part of the record. cps = counts per second.

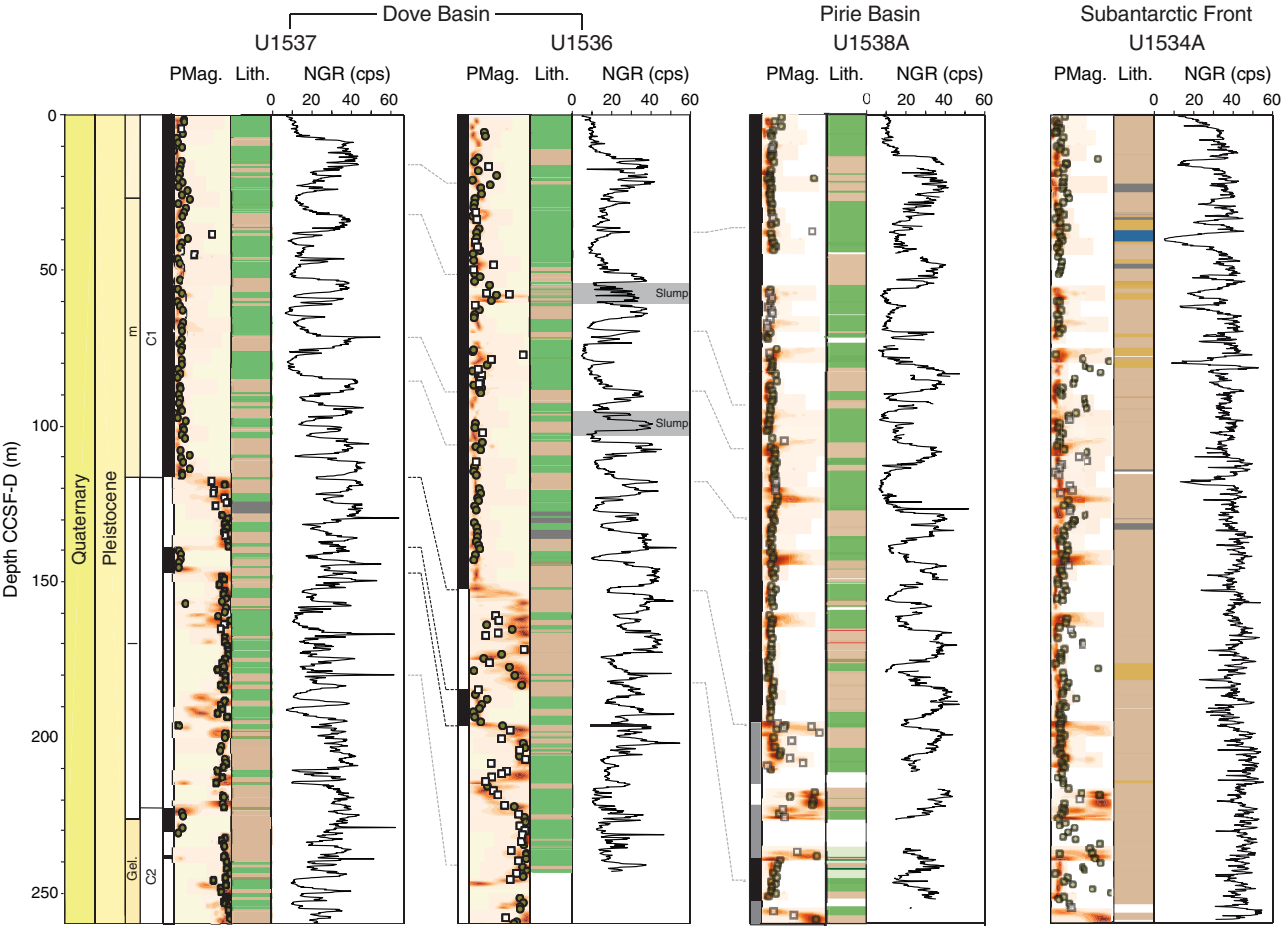


Figure F10. Examples of major lithologies of sediments from Iceberg Alley, Site U1538. (C) Interbedded silty clays and (A) diatom oozes change with lithification to (B) diatomite and (D) lithified silty clay. Pliocene deposits contain more (E) IBRD layers and (F) laminations.

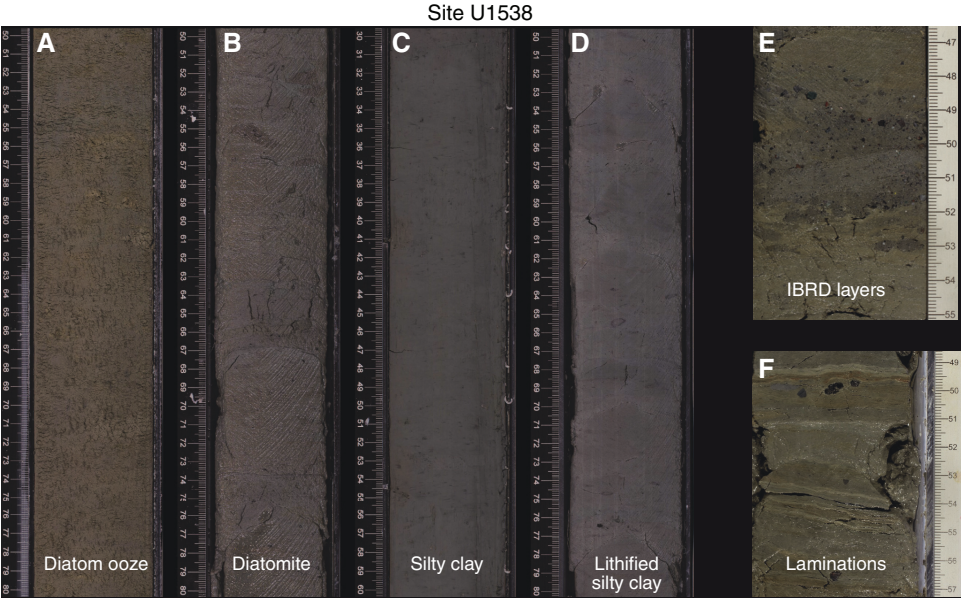


Figure F11. Biostratigraphically important diatoms and radiolarians and selected palynomorphs, Expedition 382. Diatoms (scale bars = 10 μ m): 1. *Fragilariopsis kerguelensis* (382-U1538A-16H-CC), 2. *Fragilariopsis obliquecostata* (382-U1537A-3H-CC), 3. *Rouxia constricta* (382-U1536A-14H-CC), 4. *Fragilariopsis rhombica* (382-U1537A-3H-CC), 5. *Fragilariopsis separanda* (382-U1536A-10H-1, 75 cm), 6. *Actinocyclus fasciculatus* (382-U1538A-49X-3, 75 cm), 7. *Shionodiscus oestrupii* var. *reimeri* (45X-CC), 8. *Fragilariopsis barronii* (382-U1536A-40F-CC), 9. *Rouxia antarctica* (382-U1538A-51X-1), 10. *Fragilariopsis aurica* (382-U1536C-40F-CC), 11. *Thalassiosira miocenica* (382-U1536E-22R-CC). Radiolarians (scale bars = 50 μ m): 12. *Antarctissa cylindrica* (382-U1537A-10H-CC), 13. *Stylatractus universus* (382-U1536A-14H-CC), 14. *Phormospyris antarctica* (382-U1538A-6H-CC), 15. *Cycladophora pliocenica* (382-U1534A-29F-CC), 16. *Cycladophora davisiana* (382-U1537A-13H-CC), 17. *Helotholus? vema* (382-U1537D-40F-CC), 18. *Lampromitra coronata* (382-U1536E-4R-CC), 19. *Larcopyle polyacantha titan* (382-U1538A-73X-CC), 20. *Lychnocanium grande* (73X-CC), 21. *Acrosphaera labrata* (382-U1536E-18R-CC). Palynomorphs (scale bars = 20 μ m): 22. *Selenopemphix* sp. 1 sensu Esper and Zonneveld (2007) (382-U1536A-3H-CC), 23. *Impagidinium* spp. (382-U1534A-9H-CC), 24. *Selenopemphix dionaeacysta* (16F-CC), 25. *Nothofagidites flemingii* (382-U1538A-60X-CC), 26. *Operculodinium? eirikianum* (382-U1536E-26R-CC).

