

Measurement of the $2\nu\beta\beta$ Decay Half-Life of ^{130}Te with CUORE

D. Q. Adams,¹ C. Alduino,¹ K. Alfonso,² F. T. Avignone III,¹ O. Azzolini,³ G. Bari,⁴ F. Bellini,^{5,6} G. Benato,⁷ M. Biassoni,⁸ A. Branca,^{9,8} C. Brofferio,^{9,8} C. Bucci,^{7,†} J. Camilleri,¹⁰ A. Caminata,¹¹ A. Campani,^{12,11} L. Canonica,^{13,7} X. G. Cao,¹⁴ S. Capelli,^{9,8} L. Cappelli,^{7,15,16} L. Cardani,⁶ P. Carniti,^{9,8} N. Casali,⁶ D. Chiesa,^{9,8} M. Clemenza,^{9,8} S. Copello,^{12,11} C. Cosmelli,^{5,6} O. Cremonesi,⁸ R. J. Creswick,¹ A. D'Addabbo,^{17,7} I. Dafinei,⁶ C. J. Davis,¹⁸ S. Dell'Oro,^{9,8} S. Di Domizio,^{12,11} V. Dompè,^{17,7} D. Q. Fang,¹⁴ G. Fantini,^{5,6} M. Faverzani,^{9,8} E. Ferri,^{9,8} F. Ferroni,^{17,6} E. Fiorini,^{8,9} M. A. Franceschi,¹⁹ S. J. Freedman,^{16,15,*} S. H. Fu,¹⁴ B. K. Fujikawa,¹⁶ A. Giachero,^{9,8} L. Gironi,^{9,8} A. Giuliani,²⁰ P. Gorla,⁷ C. Gotti,⁸ T. D. Gutierrez,²¹ K. Han,²² K. M. Heeger,¹⁸ R. G. Huang,¹⁵ H. Z. Huang,² J. Johnston,¹³ G. Keppel,³ Yu. G. Kolomensky,^{15,16} C. Ligi,¹⁹ L. Ma,² Y. G. Ma,¹⁴ L. Marini,^{15,16} R. H. Maruyama,¹⁸ D. Mayer,¹³ Y. Mei,¹⁶ N. Moggi,^{23,4} S. Morganti,⁶ T. Napolitano,¹⁹ M. Nastasi,^{9,8} J. Nikkel,¹⁸ C. Nones,²⁴ E. B. Norman,^{25,26} A. Nucciotti,^{9,8} I. Nutini,^{9,8} T. O'Donnell,¹⁰ J. L. Ouellet,¹³ S. Pagan,¹⁸ C. E. Pagliarone,^{7,27} L. Pagnanini,^{17,7} M. Pallavicini,^{12,11} L. Pattavina,⁷ M. Pavan,^{9,8} G. Pessina,⁸ V. Pettinacci,⁶ C. Pira,³ S. Pirro,⁷ S. Pozzi,^{9,8} E. Previtali,^{9,8} A. Puiu,^{17,7} C. Rosenfeld,¹ C. Rusconi,^{1,7} M. Sakai,¹⁵ S. Sangiorgio,²⁵ B. Schmidt,¹⁶ N. D. Scielzo,²⁵ V. Sharma,¹⁰ V. Singh,¹⁵ M. Sisti,⁸ D. Speller,²⁸ P. T. Surukuchi,¹⁸ L. Taffarello,²⁹ F. Terranova,^{9,8} C. Tomei,⁶ K. J. Vetter,^{15,16} M. Vignati,⁶ S. L. Wagaarachchi,^{15,16} B. S. Wang,^{25,26} B. Welliver,¹⁶ J. Wilson,¹ K. Wilson,¹ L. A. Winslow,¹³ S. Zimmermann,³⁰ and S. Zucchelli^{23,4}

¹Department of Physics and Astronomy, University of South Carolina, Columbia, South Carolina 29208, USA

²Department of Physics and Astronomy, University of California, Los Angeles, California 90095, USA

³INFN—Laboratori Nazionali di Legnaro, Legnaro (Padova) I-35020, Italy

⁴INFN—Sezione di Bologna, Bologna I-40127, Italy

⁵Dipartimento di Fisica, Sapienza Università di Roma, Roma I-00185, Italy

⁶INFN—Sezione di Roma, Roma I-00185, Italy

⁷INFN—Laboratori Nazionali del Gran Sasso, Assergi (L'Aquila) I-67100, Italy

⁸INFN—Sezione di Milano Bicocca, Milano I-20126, Italy

⁹Dipartimento di Fisica, Università di Milano-Bicocca, Milano I-20126, Italy

¹⁰Center for Neutrino Physics, Virginia Polytechnic Institute and State University, Blacksburg, Virginia 24061, USA

¹¹INFN—Sezione di Genova, Genova I-16146, Italy

¹²Dipartimento di Fisica, Università di Genova, Genova I-16146, Italy

¹³Massachusetts Institute of Technology, Cambridge, Massachusetts 02139, USA

¹⁴Key Laboratory of Nuclear Physics and Ion-beam Application (MOE), Institute of Modern Physics, Fudan University, Shanghai 200433, China

¹⁵Department of Physics, University of California, Berkeley, California 94720, USA

¹⁶Nuclear Science Division, Lawrence Berkeley National Laboratory, Berkeley, California 94720, USA

¹⁷Gran Sasso Science Institute, L'Aquila I-67100, Italy

¹⁸Wright Laboratory, Department of Physics, Yale University, New Haven, Connecticut 06520, USA

¹⁹INFN—Laboratori Nazionali di Frascati, Frascati (Roma) I-00044, Italy

²⁰Université Paris-Saclay, CNRS/IN2P3, IJCLab, 91405 Orsay, France

²¹Physics Department, California Polytechnic State University, San Luis Obispo, California 93407, USA

²²INPAC and School of Physics and Astronomy, Shanghai Jiao Tong University; Shanghai Laboratory for Particle Physics and Cosmology, Shanghai 200240, China

²³Dipartimento di Fisica e Astronomia, Alma Mater Studiorum—Università di Bologna, Bologna I-40127, Italy

²⁴IRFU, CEA, Université Paris-Saclay, F-91191 Gif-sur-Yvette, France

²⁵Lawrence Livermore National Laboratory, Livermore, California 94550, USA

²⁶Department of Nuclear Engineering, University of California, Berkeley, California 94720, USA

²⁷Dipartimento di Ingegneria Civile e Meccanica, Università degli Studi di Cassino e del Lazio Meridionale, Cassino I-03043, Italy

²⁸Department of Physics and Astronomy, The Johns Hopkins University, 3400 North Charles Street Baltimore, Maryland 21211, USA

²⁹INFN—Sezione di Padova, Padova I-35131, Italy

³⁰Engineering Division, Lawrence Berkeley National Laboratory, Berkeley, California 94720, USA



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We measured two-neutrino double beta decay of ^{130}Te using an exposure of 300.7 kg yr accumulated with the CUORE detector. Using a Bayesian analysis to fit simulated spectra to experimental data, it was possible to disentangle all the major background sources and precisely measure the two-neutrino contribution. The half-life is in agreement with past measurements with a strongly reduced uncertainty: $T_{1/2}^{2\nu} = 7.71^{+0.08}_{-0.06}(\text{stat})^{+0.12}_{-0.15}(\text{syst}) \times 10^{20}$ yr. This measurement is the most precise determination of the ^{130}Te $2\nu\beta\beta$ decay half-life to date.

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Introduction.—Two-neutrino double beta ($2\nu\beta\beta$) decay is a nuclear transition with the longest lifetime experimentally measured. This process occurs when two neutrons in a nucleus simultaneously decay emitting two antineutrinos and two electrons. This decay can act as background for a hypothetical process called neutrinoless double beta ($0\nu\beta\beta$) decay [1], which may occur if the neutrino was a Majorana fermion [2], and which violates lepton number conservation [3]. $0\nu\beta\beta$ decay would be new physics and could explain the origin and nature of the neutrino mass states [4–7].

Precision measurements of the $2\nu\beta\beta$ decay half-life and studies of the $2\nu\beta\beta$ decay spectral shape can provide important input for nuclear models [8–11]. Measurements are available in literature for the $2\nu\beta\beta$ decay of various isotopes such as ^{116}Cd (Aurora [12]), ^{76}Ge (GERDA [13]), ^{100}Mo (CUPID-Mo [14]), ^{150}Nd (NEMO-3 [15]), ^{82}Se (NEMO-3 [16], CUPID-0 [17]), ^{136}Xe (EXO-200 [18], KamLAND-Zen [19,20]), and ^{96}Zr (NEMO-3 [21]). This Letter will discuss the first measurement of the ^{130}Te $2\nu\beta\beta$ decay half-life performed with the unprecedented statistics of the (CUORE) experiment.

The CUORE experiment primarily searches for neutrinoless double beta decay ($0\nu\beta\beta$) of $^{130}\text{Te} \rightarrow ^{130}\text{Xe} + 2e^-$ [22,23], however, other searches are possible [24–27]. First, this Letter will outline the CUORE detector, the data collection, and the analysis of the $2\nu\beta\beta$ decay of ^{130}Te . Thereafter, we provide a description of the technique used to fit the experimental data (comprised of events from both the $2\nu\beta\beta$ decay and the background sources), and finally, we present a discussion of the fit results.

CUORE detector.—CUORE is located underground at the Gran Sasso National Laboratory of INFN, Italy, with ~ 3600 meter water equivalent overburden to shield from cosmogenic backgrounds [28,29]. The CUORE detector consists of 988 TeO_2 crystals [30] arranged in 19 towers made of 13 floors of 4 crystals, read out as individual channels. Each crystal is $5 \times 5 \times 5 \text{ cm}^3$ in size and $\sim 750 \text{ g}$ in mass, with a natural abundance of $\sim 34\%$ for ^{130}Te . The crystals are cooled to $\sim 10 \text{ mK}$ at which point they have a low heat capacity and can be operated as cryogenic bolometers. The energy deposited in the crystal by particle interaction causes a temperature increase which is measured via a neutron transmutation doped Ge thermistor [31]. The signal rise time is $\sim 100 \text{ ms}$, and the high energy resolution is second only to that of Ge detectors. Si heaters

are used to inject regular reference pulses for a detector-based correction of thermal gain drifts from long term temperature variations.

The detector is housed in a large cryogen-free cryostat, cooled to $\sim 10 \text{ mK}$ by a dilution refrigerator [32]. The cryostat and the detector were constructed with strict radiopurity controls [30] in order to reduce the α and γ backgrounds seen in Cuoricino [33], and further reduce the γ background seen in CUORE-0 [34]. The cryostat is equipped with two lead shields: a 6-cm thick shield of ancient Roman lead [35] at $\sim 4 \text{ K}$ around and below the detector, and a 30-cm thick shield at $\sim 50 \text{ mK}$ located above the detector. An additional external shield is comprised of a 25-cm thick layer of lead surrounded by a 20-cm thick layer of polyethylene. A layer (2 cm thick) of boric acid that absorbs thermalized neutrons is located between the two shields.

When $2\nu\beta\beta$ decay occurs inside a bolometer the neutrinos escape without interacting, thus, we detect the two electrons sum kinetic energy forming a continuous, β -like spectrum from 0 keV up to the Q value of the decay ($Q_{\beta\beta} = 2527 \text{ keV}$ [36–38]). Background contributions to this spectrum originate from radioactivity in the detector and cryostat components. These backgrounds can be disentangled and quantified via careful analysis of the observed spectral shape and topological information in the segmented CUORE detector in comparison to a detailed background model [39,40].

Data collection.—CUORE began taking data in early 2017. The data collected through mid 2019 are analyzed in this work and are grouped into seven datasets (physics data bounded by ^{232}Th and ^{60}Co calibration data). The $2\nu\beta\beta$ decay analysis requires high quality data over the whole energy range, specifically, channels need to be both well performing and well calibrated. This led us to exclude two datasets from the analysis, given that the large majority of channels did not satisfy these criteria. With this choice of dataset channel, we have 300.7 kg yr of TeO_2 exposure (102.7 kg yr of ^{130}Te exposure).

The data itself are a collection of events, corresponding to a triggered waveform on a single bolometer. The modularity of the detector allows us to reconstruct the event topology via a time based coincidence analysis. Events are grouped into multiplets, $\mathcal{M}_i$ (i = number of triggered bolometers) if they occur within a $\pm 30 \text{ ms}$ window on bolometers that are $\leq 15 \text{ cm}$ apart from each

other, with a minimum energy of 70 keV. Given the extremely low trigger rate of CUORE bolometers (~ 1 mHz) and the distance requirement, multiplets with $i > 1$ contain practically no accidental coincidence events and are mainly induced by particles depositing energy in multiple crystals.

We split the data into three types of spectra: a multiplicity 1 ($\mathcal{M}_1$) spectrum comprised of events where energy was deposited into a single bolometer, a multiplicity 2 ($\mathcal{M}_2$) spectrum comprised of the single energies detected by each of the two bolometers simultaneously triggered, and a Σ_2 spectrum comprised of the sum energy of the $\mathcal{M}_2$ events. The energy of a $2\nu\beta\beta$ decay event is deposited into a single bolometer with a probability obtained from Monte Carlo simulations of $\sim 90\%$. The majority of backgrounds deposit energy across two or more bolometers (such as γ 's that scatter from one crystal into another or α decays that occur on a surface between two crystals), making the $\mathcal{M}_2$ and Σ_2 spectra useful for understanding backgrounds. Events with multiplicity higher than two are not considered in this analysis since they do not add new information. The specific steps of the data processing and selection criteria are found in [23]. These steps include a dataset dependent evaluation of the energy calibration bias and of the signal efficiency. The former, defined as the difference between the reconstructed peak position and its nominal value, is measured in calibration data for γ lines that span from 511 keV to 2.6 MeV. The resulting bias is well below 0.5 keV for all the datasets. A similar result is obtained by fitting γ peaks due to background sources. The signal efficiency, defined as the probability of a signal being triggered, assigned to a correct energy and multiplicity, and finally passing data selection cuts, has an energy dependent behavior and is asymptotic to $\sim 95\%$.

Spectral fit.—We analyze the events with energies from a threshold of 350 keV to 2.8 MeV, where the $\mathcal{M}_1$ spectrum is dominated by $2\nu\beta\beta$ decay (between 900 and 2000 keV the contribution exceeds 50% of total $\mathcal{M}_1$ events) along with γ/β emissions from radioactive contaminants. To disentangle the $2\nu\beta\beta$ decay signal, we construct a background model (BM) that describes the data via a comprehensive list of possible sources. Guidelines for this work are taken from the CUORE-0 BM [39] and the CUORE background budget [41]. The background sources are radioactive contaminations located both in the bulk of the detector and cryostat components, on the surfaces of crystals, and materials with a line of sight to them. We also include cosmogenic muons.

We developed a GEANT4 [42] Monte Carlo (MC) simulation [39,41] which outputs the spectra produced by each source in the detector, reproducing all relevant features of the experimental data (e.g., multiplicity, time resolution, energy dependent trigger efficiencies, etc.). The elements of the CUORE experiment, including the cryostat, are grouped into nine geometric entities used in the model

as background source positions: the crystals, the copper structure holding the towers, the copper vessel enclosing the detector, the Roman lead, the internal and external shields made of modern lead, the cryostat thermal shields (grouped into two elements: the copper shields inside the Roman lead and those outside), and finally, the internal lead suspension system. The BM uses 62 simulated sources to fit the experimental data. One is the $2\nu\beta\beta$ decay in the crystals, and 60 others refer to different contaminants in the nine elements listed above. These include bulk and surface ^{238}U and ^{232}Th contaminations (allowing for secular equilibrium breaks), bulk ^{60}Co , ^{40}K , and a few other long lived isotopes, as indicated in Fig. 1. All these isotopes are identified from the presence of one or more characteristic γ lines in the observed spectra. The only exception is ^{90}Sr , a long-lived pure β emitter, that could be present due to a hypothesized contamination by radioactive fallout. The remaining simulation (number 62) is the cosmogenic muon flux. As described in [39], a variable binning is applied to all the spectra: the minimum bin size is 15 keV, and bins with less than 30 counts are merged. All counts belonging to a single γ line are combined into a single bin to avoid systematics from the modeling of the γ peak shapes in MC simulations. Finally, the trigger efficiency vs energy and the efficiency of quality cuts are included in the analysis as global parameters.

The observed spectrum is reconstructed by simultaneously fitting a linear combination of the 62 MC simulated spectra to the $\mathcal{M}_1$, $\mathcal{M}_2$, and Σ_2 data. The fit is done with a Bayesian approach using a Markov-Chain Monte Carlo, implemented in the JAGS software package, to sample the joint posterior probability density function (PDF) of the fit parameters [43–46]. The likelihood is a product, over bins and spectra, of Poisson distributions that give the probability of drawing the experimental counts as a function of the MC spectra normalizations. To prevent bias while tuning data quality cuts and setting the fitting procedure, the MC normalization coefficient was blinded to keep the extracted $2\nu\beta\beta$ decay half-life in terms of a nonphysical ratio that could not be compared to previous results.

For each source, except cosmogenic muons, a uniform prior is used (the activity can span from zero to ten times the maximum activity compatible with the measured spectrum). For muons, additional information is gained from the high multiplicity spectra ($\mathcal{M} > 5$) where muons become dominant, and is used to extract a Gaussian prior for the BM fit. The fit result is a joint posterior PDF for the 62 parameters from which we extract the marginalized posterior PDF for the $2\nu\beta\beta$ decay rate [47]. The fitting procedure closely follows the description in [39] and a paper detailing the CUORE BM is in preparation.

Model systematics.—The background model is able to reproduce the major features of the observed spectra (see Fig. 1) with a global $\chi^2/\text{d.o.f}$ of 681/365 ($\chi^2_{\text{red}} = 1.87$). The suboptimal agreement between the data and the MC

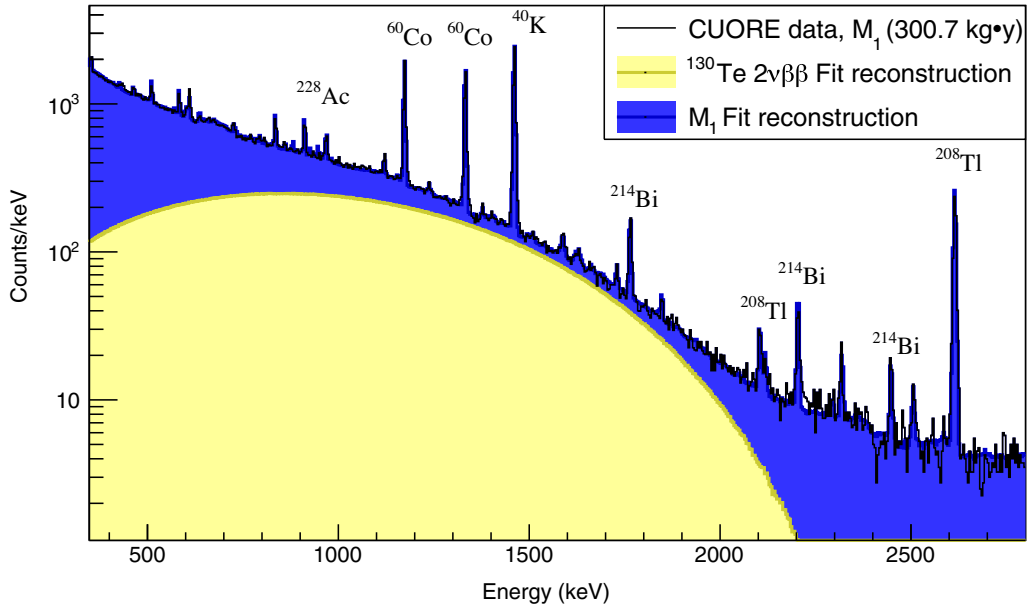


FIG. 1. The observed $\mathcal{M}_1$ spectrum (black) compared with its reconstruction as obtained by the background model (blue). The reconstructed $2\nu\beta\beta$ decay component is shown in yellow for comparison. We observe that, from 900 to 2000 keV, more than 50% of the $\mathcal{M}_1$ spectrum counts originate from the $2\nu\beta\beta$ decay process. The experimental data and the spectra reconstructed by the fit have been converted back to 1 keV binning for illustrative purposes. Selected gamma lines from background contaminants are labeled.

likely arises from an imperfect modeling of source position and distribution. Increased statistics from more data will allow for refinement of the background model by better identifying source locations or additional subdominant contaminants. This fit makes very limited use of the data, and it is based on a simplified description of sources, therefore, the result is not particularly informative on the specific position and intensity of a source. Overall the background composition matches very well the expectation discussed in [41] with a few exceptions. We see an excess of ^{238}U in the cryostat elements. The localization is not clear, but the $2\nu\beta\beta$ decay result is insensitive to the source position. We also have a quite evident ^{210}Pb surface contamination of the copper of the tower holding structure that is ~ 100 times higher than in CUORE-0. Its major contribution to the measured spectra is the α peak at 5.3 MeV due to ^{210}Po . The shape of the peak proves that the contamination is right at the surface of the copper, likely due to ^{222}Rn exposure. We observe an excess in the ^{60}Co crystal contamination compared with the expected 1 nBq/kg [41]. However, this is anticorrelated with ^{60}Co in the copper of the tower holding structure. Increased statistics and a more extensive use of coincidences will allow us to clarify this point.

In order to check the stability of the $2\nu\beta\beta$ decay half-life result, we run multiple fits over the whole dataset varying aspects of the background model. In particular, we test two different models for the $2\nu\beta\beta$ decay spectral shape, we alter the list of background sources used, and we remove the ^{90}Sr source. As additional probes of our sensitivity to various aspects of the BM sources, we fit subsets of data in which

we split the detector in half in different ways (see Fig. 2) and perform the fit on single datasets.

$2\nu\beta\beta$ Decay Model: There are two competing models for $2\nu\beta\beta$ decay, yielding slightly different spectral shapes [10,11,48]. The single state dominance (SSD) is the default used in this analysis. It results in a better fit quality and might be the first experimental hint for SSD dominance in ^{130}Te $2\nu\beta\beta$ decay. The alternate mechanism, higher state dominance (HSD), yields a slightly worse fit and a few

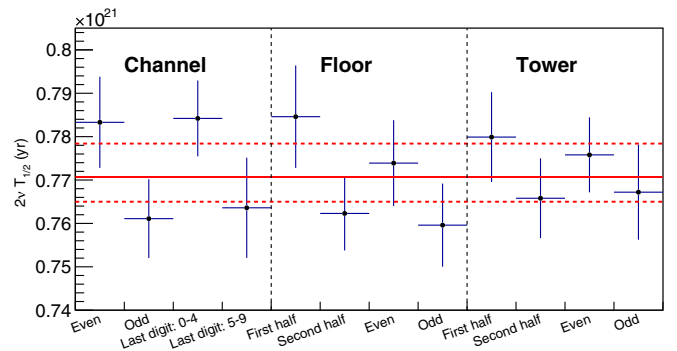


FIG. 2. Results of the $2\nu\beta\beta$ decay half-life from the tests in which data are fit by splitting the detector in half in different ways: Channel: Detector split based on channel numbers assigned to the individual crystals (even, odd, last digits: 0–4, 5–9). Floor: Detector split based on tower floor number (first and second halves, even, odd). Tower: Detector split based on tower number (first and second halves, even, odd). The results are compared to the one obtained in the full statistics fit (red solid line) and its statistical uncertainty range (red dashed lines).

percent increase of the $2\nu\beta\beta$ half-life. As this is the only shape test we perform on the $2\nu\beta\beta$ decay spectrum, we conservatively assume this systematic to be double sided and estimate the uncertainty as 68% of the difference between the fits with HSD and SSD: $\pm 1.3\%$.

Energy threshold: The energy threshold used in this analysis is 350 keV. If we vary this threshold in the range of 300–800 keV, we observe an increase of the $2\nu\beta\beta$ decay rate. We assume this systematic to be uniform between the best fit and the value that deviates the most. We symmetrize this uncertainty around the best fit to account for possible deviations given by untested threshold values, with a result of $\pm 0.4\%$.

Geometrical effect: All contaminants in the model are uniformly distributed in the nine simulated elements. To investigate possible biases, we compare fits done by splitting the detector according to crystal, floor, or tower number (see Fig. 2). Each pair of results are statistically compatible with each other to within two sigma. We take the pair with the largest splitting, subtract the statistical error, and interpret the result as the 1σ uncertainty of a flat distribution: $\pm 0.8\%$.

^{90}Sr : As mentioned, a background source due to possible crystal contamination with ^{90}Sr was introduced. This is the only long-lived pure β emitter produced by fission that produces a background (via its daughter ^{90}Y) extending up to 2.2 MeV, without any associated gamma emission that would allow us to constrain its activity [17]. The $2\nu\beta\beta$ decay result is weakly sensitive to this contaminant, as upon its removal the counts ascribed to $2\nu\beta\beta$ decay increase resulting in a slightly shorter half-life. Since ^{90}Sr has no clear signature, we use it as a proxy for the removal or addition of components to the BM. We take the systematic to be symmetric and 68% of the difference between the best fit and the fit without, giving $\pm 0.3\%$.

Datasets: We investigated the $2\nu\beta\beta$ decay result stability in time by fitting separately each of the five datasets used in this analysis and observed only statistical variations in the fit result. We also fit the two excluded datasets and use the result to quantify the bias introduced by their removal. This yields an asymmetric uncertainty of $+0.3\%$ and -1.1% .

List of background sources: In the reference fit, we use 62 sources, selected to be comprehensive, as extracted from the CUORE-0 background model [39]. Given the limited statistics used in this work and the choice of fitting only the γ region, some of the sources could be degenerate with each other while others cannot be identified easily. To check how the fit performs with a different background source list, all components with contributions compatible with 0, including ^{90}Sr , are removed. This reduces the number of distinct background source components down to 25. This has an impact on the $2\nu\beta\beta$ decay fit result compatible to that resulting from the removal of the ^{90}Sr alone.

As a result of this study, we can conclude that all the systematics we explored are, at most, in the range of 1%.

TABLE I. Chronology of $T_{1/2}^{2\nu}$ measurement in ^{130}Te . The relative uncertainty refers to statistical and systematic errors summed in quadrature.

	$T_{1/2}^{2\nu}$ (10^{20} yr)	Relative uncertainty	Reference
MiBeta	$6.1 \pm 1.4^{+2.9}_{-3.5}$	57%	2003 [51]
NEMO-3	$7.0 \pm 0.9 \pm 1.1$	20%	2011 [50]
CUORE-0	$8.2 \pm 0.2 \pm 0.6$	7.7%	2016 [39]
CUORE	$7.71^{+0.08+0.12}_{-0.06-0.15}$	2.0%	(This result)

The dominant contribution comes from the uncertainty in the decay model (SSD vs HSD) which may be improved with increased statistics or theoretical input. Finally, other sources of uncertainties, such as the efficiency of our coincidence selection, the chance of mixing up $\mathcal{M}_1$ with $\mathcal{M}_2$ events, or the efficiency of pulse shape cuts, have an overall impact on the final error that is lower than 0.1%. The Monte Carlo statistics, though optimized to yield negligible error, is properly accounted for in the fitting procedure.

$2\nu\beta\beta$ decay results and discussion.—To extract a robust estimate of the $2\nu\beta\beta$ decay half-life, we combine our systematics in quadrature. Through the unblinding of the correct normalization coefficient for the MC spectrum, we obtain the measurement of the $2\nu\beta\beta$ decay half-life of ^{130}Te . Though the posterior for ^{90}Sr is compatible with null activity, the insertion of ^{90}Sr in the BM does weakly distort the $2\nu\beta\beta$ decay posterior. Removal of the ^{90}Sr source results in a symmetric posterior, however, we choose to include this source due to the high anticorrelation with $2\nu\beta\beta$ decay.

We use an isotopic abundance of ^{130}Te of $(34.167 \pm 0.002)\%$ [49]. From this, and the systematic uncertainties described previously, we obtain a half-life of $T_{1/2}^{2\nu} = 7.71^{+0.08}_{-0.06}(\text{stat})^{+0.12}_{-0.15}(\text{syst}) \times 10^{20}$ yr, a value consistent with previous measurements (see Table I). This result is the most precise measurement of the $2\nu\beta\beta$ decay half-life of ^{130}Te to date and one of the most precise measurements of a $2\nu\beta\beta$ decay half-life. It represents a substantial improvement over previous measurements from NEMO-3 [50] and CUORE-0 [39] owing to the CUORE strict radiopurity controls, the improved signal-to-noise ratio, the increased statistics, and the robust background model.

Conclusion.—In this Letter, we described the analysis of the ^{130}Te $2\nu\beta\beta$ decay measured with CUORE. We exploit the geometry of the CUORE detector to tag single scatter and multiple scatter events to obtain separate spectra dominated by $2\nu\beta\beta$ decay and background events, respectively. The ^{130}Te $2\nu\beta\beta$ decay half-life is measured to be $T_{1/2}^{2\nu} = 7.71^{+0.08}_{-0.06}(\text{stat})^{+0.12}_{-0.15}(\text{syst}) \times 10^{20}$ yr. Compared to previous results (Table I) this is the most precise determination of the $2\nu\beta\beta$ decay half-life in ^{130}Te . The present result is dominated by a $\sim 2\%$ systematic

uncertainty. Further improvement will require a better understanding of the background sources localization, as indicated by the observed BM variations with different geometrical detector splittings. This refinement, as well as an improved study of the SSD vs HSD models, is feasible in the near future given the increased statistics being collected by CUORE.

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*Deceased.

†Corresponding author.

cuore-spokesperson@lngs.infn.it

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