

# Settling and clustering of snow particles in atmospheric turbulence

Cheng Li<sup>1,2</sup>, Kaeul Lim<sup>1,2</sup>, Tim Berk<sup>1,3</sup>, Aliza Abraham<sup>1,2</sup>, Michael Heisel<sup>1,4</sup>, Michele Guala<sup>1,4</sup>, Filippo Coletti<sup>1,3</sup> and Jiarong Hong<sup>1,2,†</sup>

<sup>1</sup>St Anthony Falls Laboratory, University of Minnesota, Minneapolis, MN 55414, USA

<sup>2</sup>Department of Mechanical Engineering, University of Minnesota, Minneapolis, MN 55455, USA

<sup>3</sup>Department of Aerospace Engineering and Mechanics, University of Minnesota, Minneapolis, MN 55455, USA

<sup>4</sup>Department of Civil, Environmental, and Geo-Engineering, University of Minnesota, Minneapolis, MN 55455, USA

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The effect of turbulence on snow precipitation is not incorporated into present weather forecasting models. Here we show evidence that turbulence is in fact a key influence on both fall speed and spatial distribution of settling snow. We consider three snowfall events under vastly different levels of atmospheric turbulence. We characterize the size and morphology of the snow particles, and we simultaneously image their velocity, acceleration and relative concentration over vertical planes approximately 30 m<sup>2</sup> in area. We find that turbulence-driven settling enhancement explains otherwise contradictory trends between the particle size and velocity. The estimates of the Stokes number and the correlation between vertical velocity and local concentration are consistent with the view that the enhanced settling is rooted in the preferential sweeping mechanism. When the snow vertical velocity is large compared to the characteristic turbulence velocity, the crossing trajectories effect results in strong accelerations. When the conditions of preferential sweeping are met, the concentration field is highly non-uniform and clustering appears over a wide range of scales. These clusters, identified for the first time in a naturally occurring flow, display the signature features seen in canonical settings: power-law size distribution, fractal-like shape, vertical elongation and large fall speed that increases with the cluster size. These findings demonstrate that the fundamental phenomenology of particle-laden turbulence can be leveraged towards a better predictive understanding of snow precipitation and ground snow accumulation. They also demonstrate how environmental flows can be used to investigate dispersed multiphase

† Email address for correspondence: [jhong@umn.edu](mailto:jhong@umn.edu)

flows at Reynolds numbers not accessible in laboratory experiments or numerical simulations.

**Key words:** atmospheric flows, particle/fluid flow, turbulent boundary layers

## 1. Introduction

The fall speed of snow, and frozen hydrometeors in general, is a crucial parameter for meteorological prediction (Hong, Dudhia & Chen 2004). The spatio-temporal distribution of snow precipitation directly impacts the ground accumulation, which in turn influences local hydrology, road conditions, vegetation development, avalanche danger and mass balance of glaciers (Lehning *et al.* 2008; Scipión *et al.* 2013). At a global level, the rate at which ice and snow particles settle in the atmosphere is one of the most important determinants of climate sensitivity (IPCC 2014). Considering its importance, our understanding of snow particle settling is far from satisfactory, and the process remains poorly characterized (Heymsfield & Westbrook 2010). We will use the word settling as it is more common in fluid mechanics, although the atmospheric science literature often terms it sedimentation. Also, because we focus here on relatively small hydrometeors as opposed to large dendritic ones, we will generally refer to snow particles as opposed to snowflakes.

A common approach is to parameterize the vertical velocity of the snow particles (or generic hydrometeor)  $W_s$  as a power-law function of the characteristic diameter  $d_s$  (Locatelli & Hobbs 1974):

$$W_s = a_w d_s^{b_w}, \quad (1.1)$$

where  $a_w$  and  $b_w$  are empirical constants. A first difficulty lies in the specificity of the constants to the type of hydrometeors, which display a broad range of size, morphology, porosity and riming that affect the balance between drag and gravity (Pruppacher & Klett 1997). Fall speed relationships that include the object mass  $m_s$  and frontal area  $A_s$  enable the definition of a particle Reynolds number and drag coefficient, and show more generality (Böhm 1989; Mitchell 1996). However, even these models are ultimately similar to (1.1) in that they resort to a power-law dependence of  $m_s$  and  $A_s$  on  $d_s$ .

Small snow particles and ice crystals (millimetre-sized or smaller) form the vast majority of the frozen precipitation in the atmosphere (Pruppacher & Klett 1997). Field studies focused on these small particles report values of the coefficient  $b_w \approx 0.25$  (although with significant variability; see Locatelli & Hobbs 1974; Tiira *et al.* 2016; von Lerber *et al.* 2017). From the fluid mechanics standpoint, this would seem a very weak relation between the fall speed and the diameter. Stokes drag implies  $b_w = 2$ , and while nonlinear drag certainly affects the process, it is not expected to be dominant as the particle Reynolds number for those small hydrometeors is typically  $O(10)$ . Important factors that contribute to the trend include: particle bulk density, which varies significantly with the level of riming and porosity (Pruppacher & Klett 1997); particle anisotropy, which can be high for needles and crystal aggregates (Dunnavan *et al.* 2019); and in general the complexity of the snow particle morphology, especially for dendritic ice crystals, aggregates and open geometries (Westbrook 2008; Heymsfield & Westbrook 2010). Morphological factors, however, are not expected to play major roles for small hydrometeors of compact shape, while the weak dependence with the diameter is still observed (Tiira *et al.* 2016). Therefore, it is evident that other environmental factors can influence the settling process besides the snow particle properties.

The effect of atmospheric turbulence on the snow fall speed has only recently been recognized. Garrett & Yuter (2014) considered data from a field study and showed

that, in high turbulence, the fall speed was seemingly insensitive to the snow particle diameter. They simultaneously measured hydrometeor morphology and fall speed using a multi-camera system, and found that unrealistic density estimates were obtained by assuming that the observed vertical velocity coincided with the terminal velocity in quiescent air. While it is intuitive that air turbulence would broaden the distribution of snow particle velocities, as recently confirmed (Garrett *et al.* 2015), one may also expect the effectively Gaussian velocity fluctuations to cancel out, leaving the average fall speed unaffected. This view, however, does not account for well-known phenomena in particle-laden flows.

When small heavy particles fall through turbulence, their mean settling velocity can be significantly altered compared to still-fluid conditions (Nielsen 1993; Wang & Maxey 1993; Balachandar & Eaton 2010). The fall can be hindered, e.g. if weakly inertial particles are trapped in vortices (Tooby, Wick & Isaacs 1977), or if fast-falling particles are slowed down by nonlinear drag (Mei, Adrian & Hanratty 1991) or loiter in upward regions of the flow (Good, Gerashchenko & Warhaft 2012). More often, however, turbulence is found to enhance the settling through a process known as preferential sweeping (Maxey 1987; Wang & Maxey 1993): inertial particles favour downward regions of the flow, i.e. they oversample fluid with vertical velocity fluctuations aligned with the direction of gravity. This effect is especially strong (and dominant over other mechanisms that hinder the fall) when the particles have an aerodynamic response time  $\tau_p$  comparable to the Kolmogorov time scale  $\tau_\eta$ ; that is, when the Stokes number  $St \equiv \tau_p/\tau_\eta = O(1)$ . Laboratory experiments have shown that in this case the mean vertical velocity can see a multi-fold increase (Aliseda *et al.* 2002; Good *et al.* 2014; Huck *et al.* 2018; Petersen, Baker & Coletti 2019). Recently, Nemes *et al.* (2017) imaged and tracked snow particles in the atmospheric surface layer, observed a substantial increase in settling velocity and suggested that preferential sweeping was at play.

Another well-known behaviour exhibited by inertial particles in turbulence is the tendency to form clusters, especially when  $St = O(1)$  (Eaton & Fessler 1994; Balkovsky, Falkovich & Fouxon 2001; Chun *et al.* 2005; Monchaux, Bourgoïn & Cartellier 2012; Gustavsson & Mehlig 2016). Along with the ability of inertial particles to maintain significant relative velocity for vanishing separations, this effect is thought to enhance their collision rate (Sundaram & Collins 1997; Bewley, Saw & Bodenschatz 2013). As such, clustering is expected to be consequential for a variety of natural phenomena, from atmospheric cloud dynamics (Shaw 2003; Saw *et al.* 2008; Grabowski & Wang 2013) to dust agglomeration in circumstellar nebulae (Cuzzi *et al.* 2001). Precipitating snow can be largely composed of aggregates formed by successive collisions of ice crystals (Dunnavan *et al.* 2019). Thus, if clustering of frozen hydrometeors does occur, it is likely to play an important role in determining their particle shape, size and fall speed. To date, there is no direct evidence that snow particles cluster in the atmosphere, nor the properties that such clusters may possess, and the impact that turbulence may have on the evolution of frozen precipitation remains speculative.

Here we present and analyse data from three field studies where settling snow is illuminated and imaged over vertical planes  $\sim 30 \text{ m}^2$ , using previously introduced approaches (Hong *et al.* 2014; Nemes *et al.* 2017). We characterize the snow particle velocities and accelerations across a broad range of atmospheric conditions, and show that turbulence plays a dominant role in determining both mean and variance of the snow fall speed. Moreover, we document for the first time the appearance of clusters in the snow spatial distribution, describe their multi-scale geometry and assess their settling velocity. The paper is organized as follows: the experimental set-ups to characterize the atmospheric conditions, snow particle properties and large-scale velocity fields are described in § 2; the

| Dataset       | PIV/PTV set-up    |                  |   |   |                    | DIH set-up      |  |                              |
|---------------|-------------------|------------------|---|---|--------------------|-----------------|--|------------------------------|
|               | Duration<br>(min) | $z_{FOV}$<br>(m) | $\Delta z_{FOV} \times \Delta x_{FOV}$<br>(m <sup>2</sup> ) | Resolution<br>(mm pixel <sup>-1</sup> ) | $\theta$<br>(deg.) | $L_{CL}$<br>(m) | Resolution<br>( $\mu\text{m pixel}^{-1}$ ) | Volume<br>(cm <sup>3</sup> ) |
| January 2016  | 5                 | 10.8             | $7.1 \times 4.0$  | 5.6                                     | 21.1               | 25              | 24   | 18.8                         |
| November 2018 | 17                | 9.1              | $8.4 \times 4.7$  | 6.5                                     | 14.5               | 31              | 14   | 42                           |
| January 2019  | 15                | 20.2             | $14.7 \times 8.3$   | 12.0                                    | 19.9               | 53              | 14   | 42                           |

Table 1. Summary of key parameters of PIV, PTV and DIH measurement set-ups for each deployment dataset used in the present paper (see figure 1). All PIV/PTV datasets have the same acquisition rate of 120 fps.

results in terms of snow particle size, velocity, acceleration and concentration fields are reported in § 3; and in § 4 we draw conclusions and discuss future perspectives.

## 2. Methodology

### 2.1. Field experiment set-ups

The data presented in the current study were acquired in three field deployments conducted at the Eolos Wind Energy Research Field Station in Rosemount, MN, between 2016 and 2019. We refer to them as January 2016, November 2018 and January 2019. The station features a meteorological tower instrumented with wind velocity, temperature and humidity sensors installed at elevations ranging from 7 to 129 m. Four of these elevations (10, 30, 80 and 129 m) are instrumented with Campbell Scientific CSAT3 3D sonic anemometers with a sampling rate of 20 Hz. Detailed descriptions of the field station and instrument specifications are provided in Hong *et al.* (2014) and Toloui *et al.* (2014).

For each deployment, the spatial distribution and motion of the settling snow is captured using super-large-scale particle image velocimetry (PIV) and particle tracking velocimetry (PTV) described in Hong *et al.* (2014) and Nemes *et al.* (2017), respectively. The size and shape of snow particles are simultaneously obtained using an in-house digital in-line holographic (DIH) system introduced in Nemes *et al.* (2017). A minimum of 15 000 holograms are captured for each dataset. Because the DIH set-up is located just metres away from the PIV/PTV field of view (FOV), spatial variability between both measurements is deemed negligible. The key information for the PIV/PTV and DIH systems in each deployment is summarized in table 1, and further details of the set-ups are given in the following. All three sets have fairly constant snowfall and wind intensity.

The PIV/PTV set-up is similar to the one used in Nemes *et al.* (2017). The illumination is provided by a 5 kW searchlight with a divergence  $<0.3^\circ$  and an initial beam diameter of 300 mm, shining on a curved mirror that redirects the beam vertically and expands it into a light sheet. The system is attached to a trailer for mobility in aligning the sheet with the wind direction and minimizing out-of-plane motion. The illuminated snow particle images are recorded using a CMOS (Sony A7RII, Sony Corp.) camera at 120 fps and  $720 \times 1080$  pixel<sup>2</sup> for all three datasets. The camera is placed on a tripod at a distance  $L_{CL}$  from the light sheet with a tilt angle  $\theta$  from the horizontal. The coordinate system (streamwise  $x$ , spanwise  $y$  and vertical  $z$ , and the corresponding velocity components  $u$ ,  $v$  and  $w$ ) as well as the position and dimensions of the FOV are illustrated in figure 1.

Figure 2 shows sample images used for PIV/PTV in each deployment, which span a broad range of imaging conditions and have different FOVs (figure 2*a–c*). For consistency, in all datasets we analyse a sampling region of  $7.1 \text{ m} \times 4.0 \text{ m}$  (matching the January 2016 FOV), centred at  $z_{FOV}$  and  $x = 0$ , with the same  $32 \times 32$  pixel<sup>2</sup> PIV interrogation window

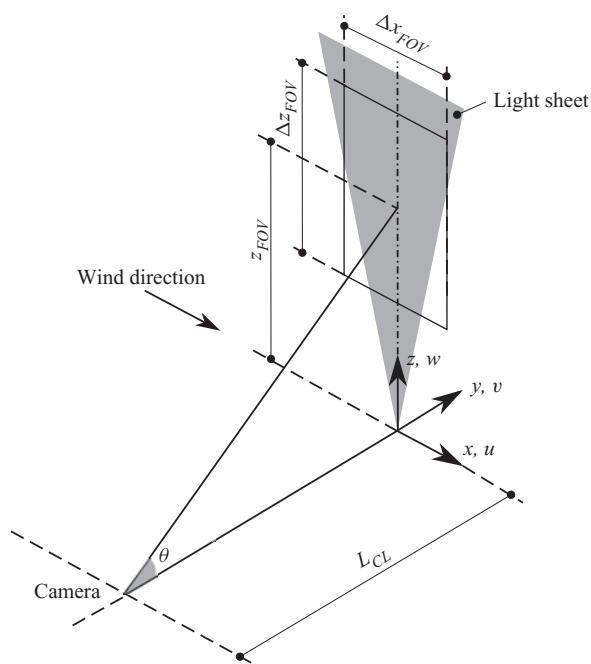


Figure 1. Schematic of the measurement set-up used in the deployments. The FOV has width  $\Delta x_{FOV}$ , height  $\Delta z_{FOV}$  and is centred at an elevation  $z_{FOV}$  (see table 1). The other symbols are defined in the text.

(figure 2d–f). Based on previous parameterization of the boundary layer at the Eolos site (Heisel *et al.* 2018), the selected regions are located in the logarithmic layer, well above the roughness sublayer and possible snow saltation layer (Guala *et al.* 2008b). The particle image density is similar for January 2016 and November 2018, and significantly higher for January 2019. Accordingly, both PIV and PTV are applied to January 2016 and November 2018, while only PIV (which can deal with high particle image densities) is performed for the January 2019 dataset. Particle image velocimetry provides Eulerian velocity fields (for atmospheric measurements, see Hong *et al.* (2014), Toloui *et al.* (2014) and Heisel *et al.* (2018)), which we use to measure the snow fall speed, while PTV provides Lagrangian trajectories, which we use to calculate snow particle accelerations (Nemes *et al.* 2017). The image and data processing consist of two phases, the first being image distortion correction and enhancement and the second being velocity extraction. During the first phase, the image distortion due to tilt of the camera is firstly corrected using the measured camera–light distance and tilt angle. The images are further enhanced through temporal background subtraction. The enhanced images are used to calculate both Eulerian fields based on PIV cross-correlation (Dasari *et al.* 2019) and Lagrangian fields using PTV (Nemes *et al.* 2017). The same images are also used to estimate the relative snow particle concentration, as described in § 3.3.

Two versions of the DIH system are employed to characterize the snow particle size, shape and number density. The earlier version is employed for January 2016 and is described in detail in Nemes *et al.* (2017). The later version, which has larger sampling volume and improved spatial resolution and data acquisition capabilities, is used for November 2018 and January 2019. It uses a diode laser (Roithner 5 mW, wavelength of 635 nm), a beam expander (Edmund Optics 9 mm plano-concave lens) and a collimating lens (Thorlabs 100 mm biconvex lens with anti-reflective coating) to generate a 50 mm beam.



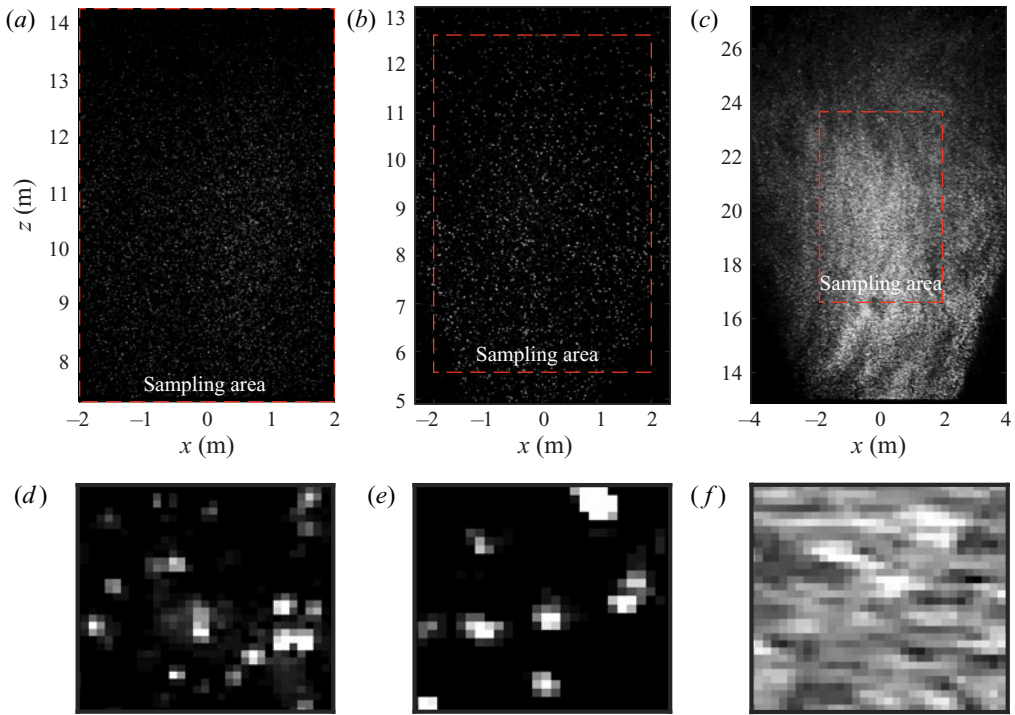


Figure 2. Samples of raw snow particle images at the full FOV (*a–c*) used for PIV/PTV, and close-up on the  $32 \times 32$  pixels<sup>2</sup> PIV interrogation window (*d–f*) from datasets January 2016 (*a,d*), November 2018 (*b,e*) and January 2019 (*c,f*).

A CMOS camera (PtGrey Blackfly  $2048 \times 1536$  pixels<sup>2</sup>,  $3.45 \mu\text{m pixel}^{-1}$ ), mounting a Fujinon 25 mm f/1.4 lens, captures the holograms resulting from the interference of the light scattered by the snow particles and the collimated beam. The camera is connected to a data acquisition system (Raspberry Pi 3, Model B), which interfaces with a laptop running FLIR SpinView software to control the camera and collect the images. Both versions of the DIH system are mounted approximately 2 m above ground level and allow the snow particles to settle through the sampling volume with minimal disturbance. Nemes *et al.* (2017) describe the processing steps through which detailed two-dimensional projections of the snow particle silhouette are obtained from the holograms. Briefly, the processing of the acquired hologram includes image enhancement, numerical reconstruction and snow particle sizing. The enhancement involves subtracting the temporal background to the intensity of the holograms, followed by adaptive histogram equalization. The snow particle holograms are then digitally reconstructed at multiple planes along the depth direction in the sample volume using the Rayleigh–Sommerfeld kernel (Katz & Sheng 2010). The particle boundaries are determined by standard segmentation using an in-house hologram sizing software, from which size and aspect ratio are determined for each hydrometeor. The position and size of the detected snow particles are used to determine the number concentration and volume fraction ( $\phi_V$ ).

## 2.2. Meteorological conditions and turbulence properties

Simultaneous measurements from the meteorological tower sensors provide a statistical description of the turbulence conditions during the PIV/PTV and DIH measurements. As shown in figure 3, the time series of the wind velocity ( $u$ ) from the 10 m sonic anemometer

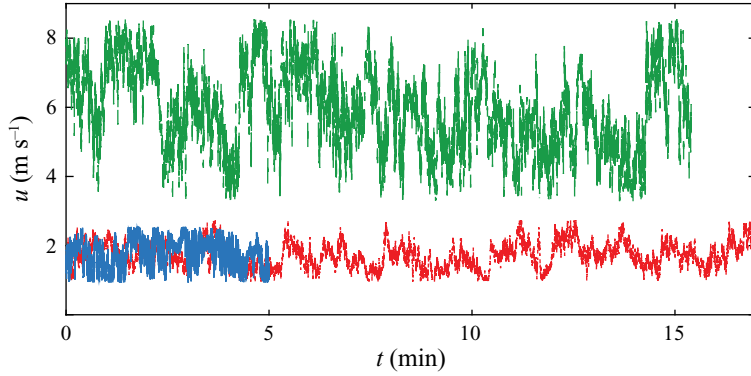


Figure 3. Time series of the sonic anemometer data at elevation  $z = 10$  m showing streamwise wind velocity for January 2016 (blue), November 2018 (red) and January 2019 (green).

(at an elevation comparable to the PIV/PTV FOV) indicate good stationarity for all datasets. Table 2 summarizes the key meteorological and turbulence parameters. They include mean and root-mean-square (r.m.s.) horizontal wind velocity ( $U$  and  $u_{rms}$ ), relative humidity (RH) and temperature ( $T$ ). Atmospheric stability conditions are estimated based on both the bulk Richardson number  $R_b$  and the Monin–Obukhov length  $L_{OB}$ :

$$R_b = -|g|\Delta\bar{\theta}_v\Delta z / \left( \bar{\theta}_v \left[ (\Delta V_N)^2 + (\Delta V_W)^2 \right] \right), \quad (2.1)$$

$$L_{OB} = -U_\tau^3 \bar{\theta}_v / \kappa g \overline{w'\theta'_v}. \quad (2.2)$$

Here  $g$  is the gravitational acceleration,  $\theta_v$  is the virtual potential temperature (calculated using 1 Hz temperature, pressure and relative humidity measurements),  $V_N$  and  $V_W$  are the average north and west wind velocity components, respectively,  $U_\tau$  is the shear velocity,  $\kappa$  is the von Kármán constant, prime indicates temporal fluctuations and the overbar indicates time averaging. The mean velocity differences  $\Delta V_N$  and  $\Delta V_W$  are the velocity differences measured from the sonic anemometers at 10 and 129 m ( $\Delta z = 119$  m). The length  $L_{OB}$  and all other turbulence quantities reported in this section are evaluated from the 20 Hz sonic sensor at  $z = 10$  m. We approximate the surface turbulent virtual potential heat flux with the measured  $\overline{w'\theta'_v}$ . The friction velocity  $U_\tau$  is estimated based on the Reynolds stresses (Stull 1988):

$$U_\tau = \left( \overline{u'w'^2} + \overline{v'w'^2} \right)^{1/4}. \quad (2.3)$$

For all deployments,  $R_b \ll 1$  and  $z/L_{OB} \ll 0.1$ , indicating that the boundary layer flow within the measuring domain can be approximated as neutrally stratified (Högström, Hunt & Smedman 2002). At least in our region of interest, this excludes the possibility of strong departures from canonical turbulent boundary layer flows due to stable stratification or to non-stationarities associated with gravity waves.

The integral time scale  $\tau_L$  and length scale  $L$  are calculated from the temporal autocorrelation function  $\rho$ :

$$\rho(\tau) = \overline{u'(t)u'(t+\tau)} / \overline{u'(t)^2}, \quad (2.4)$$

$$\tau_L = \int_0^{T_0} \rho(\tau) d\tau, \quad (2.5)$$

$$L = u_{rms}\tau_L, \quad (2.6)$$

| Dataset       | $U$<br>( $\text{m s}^{-1}$ ) | $u_{rms}$<br>( $\text{m s}^{-1}$ ) | RH<br>(%) | $T$<br>( $^{\circ}\text{C}$ ) | $R_b$<br>(—) | $L_{OB}$<br>(m) | $L$<br>(m) | $\tau_L$<br>(s) | $\varepsilon$<br>( $\text{cm}^2 \text{s}^{-3}$ ) | $\eta$<br>(mm) | $\tau_{\eta}$<br>(ms) | $Re_{\lambda}$<br>(—) | $G\tau_{\eta}$<br>(—) |
|---------------|------------------------------|------------------------------------|-----------|-------------------------------|--------------|-----------------|------------|-----------------|--|----------------|-----------------------|-----------------------|-----------------------|
| January 2016  | 1.98                         | 0.16                               | 98.3      | −0.1                          | 0.03         | 811             | ~4.9       | ~30.6           | ~8   | 1.29           | 126                   | 938                   | 0.003                 |
| November 2018 | 1.55                         | 0.38                               | 94.3      | −1.8                          | 0.12         | 1007            | 3.4        | 8.9             | 21   | 1.04           | 83                    | 3545                  | 0.003                 |
| January 2019  | 5.95                         | 1.18                               | 80.0      | −16.0                         | 0.03         | −2643           | 14.6       | 12.9            | 314  | 0.49           | 20                    | 9180                  | 0.008                 |

Table 2. The meteorological and turbulence parameters obtained using the sonic anemometer at  $z = 10$  m for all three datasets. See the text for the definitions of the symbols.



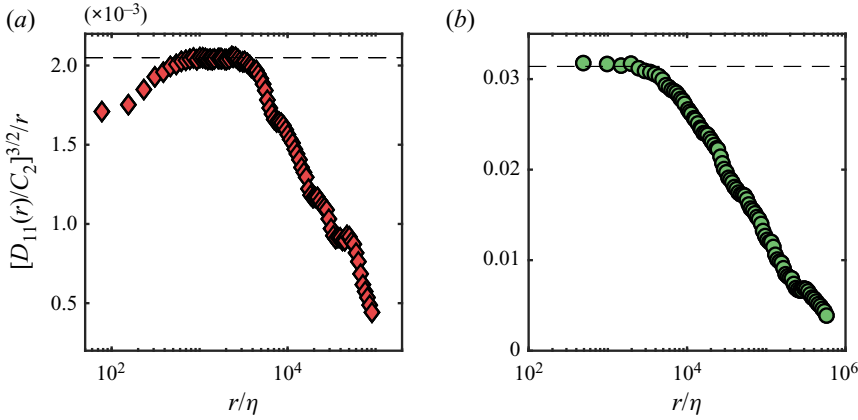


Figure 4. Estimation of energy dissipation using compensated second-order structure function of the streamwise velocity fluctuations calculated from the sonic anemometer for (a) November 2018 and (b) January 2019 at  $z = 10$  m. The dashed line indicates the inertial range prediction of turbulent dissipation rate according to (2.8) with  $C_2 = 2$ .

where  $t$  is time,  $\tau$  is the temporal separation and  $T_0$  is the first zero-crossing point of the autocorrelation function. Here and in the following, subscript ‘rms’ indicates r.m.s. fluctuation. The turbulent dissipation rate  $\varepsilon$  for November 2018 and January 2019 is estimated using the second-order temporal structure function of the streamwise velocity fluctuations:

$$D_{11}(\tau) \equiv \overline{[u'(t + \tau) - u'(t)]^2}. \quad (2.7)$$

To yield better convergence of  $D_{11}(\tau)$ , the velocity time series are divided into two-minute windows with 50 % overlap. Invoking the Taylor hypothesis, the temporal separation is converted to a spatial separation  $r = \tau U_{2\min}$ , where  $U_{2\min}$  is the mean velocity in each two-minute window. We then calculate  $\varepsilon$  using the Kolmogorov prediction for the spatial second-order structure function in the inertial range:

$$D_{11}(r) = C_2(\varepsilon r)^{2/3}, \quad (2.8)$$

where  $C_2$  is a constant close to 2 in high-Reynolds-number turbulence (Saddoughi & Veeravalli 1994). The compensated structure functions in figure 4 show good agreement with (2.8) for a broad range of separation time scales in both November 2018 and January 2019 datasets. For January 2016 the convergence of the structure function is less satisfactory and the dissipation is approximated from classic scaling arguments, i.e.  $\varepsilon = u_{rms}^3/L$ . We then obtain the Kolmogorov time and length scales,  $\tau_\eta = (\nu/\varepsilon)^{1/2}$  and  $\eta = (\nu^3/\varepsilon)^{1/4}$ , respectively, where  $\nu$  is the air kinematic viscosity. The Reynolds number  $Re_\lambda = u_{rms}\lambda/\nu$  (where  $\lambda = u_{rms}(15\nu/\varepsilon)^{1/2}$  is the Taylor microscale) spans a full decade across the three deployments, allowing us to investigate turbulence of vastly different intensity. The mean shear across the FOV  $G = \partial U/\partial z$  is much smaller than the small-scale velocity gradients, i.e.  $G(\nu/\varepsilon)^{1/2} = G\tau_\eta \ll 1$ , and we thus expect approximate small-scale isotropy (Saddoughi & Veeravalli 1994). This enables the comparison with previous laboratory experiments and simulations of particle–turbulence interactions performed in (nearly) homogeneous isotropic turbulence, consistent with the approach of Nemes *et al.* (2017).

At the scales considered here, the snow particle dynamics is strongly intermittent and influenced by low-frequency events in the atmospheric boundary layer. A substantial

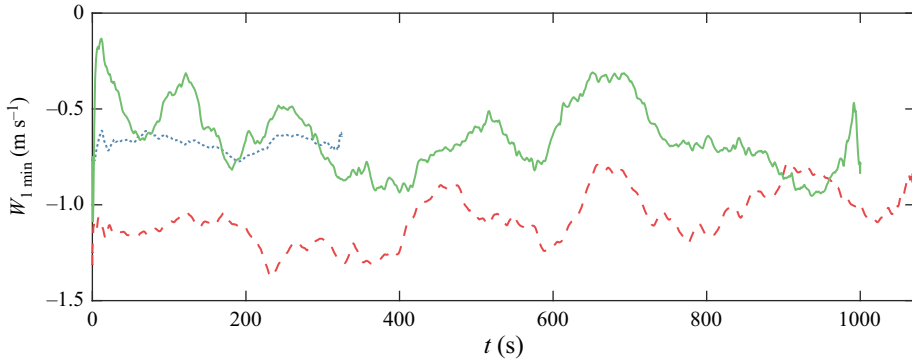


Figure 5. Temporal evolution of 1 min moving averaged settling velocity at the centre of the FOV for January 2016 (blue dotted line), November 2018 (red dashed line) and January 2019 (green solid line).

run time is needed to statistically characterize such processes. Figure 5 shows the 1 min moving average of the settling velocity ( $W_{1\text{ min}}$ ) at the centre of the FOV for the three datasets. Despite large oscillations, likely related to very-large-scale motions in the boundary layer, the trend is approximately stationary over the considered time windows. The time scale corresponding to the snow particle transport across the entire FOV can be estimated as  $T_{\text{FOV}} \sim \Delta x_{\text{FOV}}/U_s = 3.0, 2.8$  and  $1.1$  s for January 2016, November 2018 and January 2019, respectively, where  $U_s$  is the snow particle horizontal velocity. Thus, the length of recording covers at least  $100T_{\text{FOV}}$ . An estimate of  $T_{\text{FOV}}$  based on the snow settling velocity and vertical extent of the FOV leads to a similar conclusion. In terms of integral time scale, the recording covers at least  $10\tau_L$ . It was also verified that, for all three cases, the simultaneous measurements of vertical wind velocity from the 10 m sonic anemometer average out to negligibly small values compared to the snow settling velocities. The anemometer is located at approximately 50 m from the imaging site, for which it is expected to be representative of the same atmospheric conditions. Therefore, the present settling measurements are not significantly biased by up/downdrafts.

### 3. Results

#### 3.1. Snow particle size and settling velocity

Table 3 provides a summary of the average size, aspect ratio and concentration of the snow particles as measured by DIH. The particle size  $d_s$  is quantified using the projected-area diameter, corresponding to the diameter of the circle with the same projected area as the particle image. The aspect ratio  $s_2/s_1$  is defined as the ratio between minor and major axes of the ellipse fitted to each particle image. From the particle number concentration (average particle count per unit volume), the snow particle volume fraction  $\phi_V$  is estimated approximating each particle as a sphere of diameter  $d_s$ .

The particle sizes decrease significantly from January 2016 to November 2018 to January 2019, as illustrated by the probability density functions (p.d.f.s) of  $d_s$  (figure 6a). The distributions of the aspect ratio are similar for the three deployments and indicate relatively compact objects (figure 6b). This is confirmed by visual inspection of the DIH realizations: most detected hydrometeors are ice particles and crystals exhibiting moderate level of aggregation and relatively low shape complexity (Garrett *et al.* 2012). Given the limited elongation, the influence of the particle anisotropy on the motion dynamics is expected to be small (Voth & Soldati 2017). Consistent with the PIV/PTV images of figure 2, the DIH

| Dataset       | Mean diameter $d_s$<br>(mm) | Aspect ratio $s_2/s_1$<br>(—) | Number<br>concentration ( $\text{m}^{-3}$ ) | $\phi_V \times 10^{-7}$<br>(—) |
|---------------|-----------------------------|-------------------------------|---|--------------------------------|
| January 2016  | $1.09 \pm 0.45$             | $0.73 \pm 0.11$               | 816   | 7.4                            |
| November 2018 | $0.65 \pm 0.41$             | $0.65 \pm 0.16$               | 1644  | 6.3                            |
| January 2019  | $0.39 \pm 0.23$             | $0.57 \pm 0.17$               | 56 620                                      | 44                             |

Table 3. Snow particle properties (mean and standard deviation) as measured using DIH for all three datasets. Parameter  $\phi_V$  represents the snow particle volume fraction.

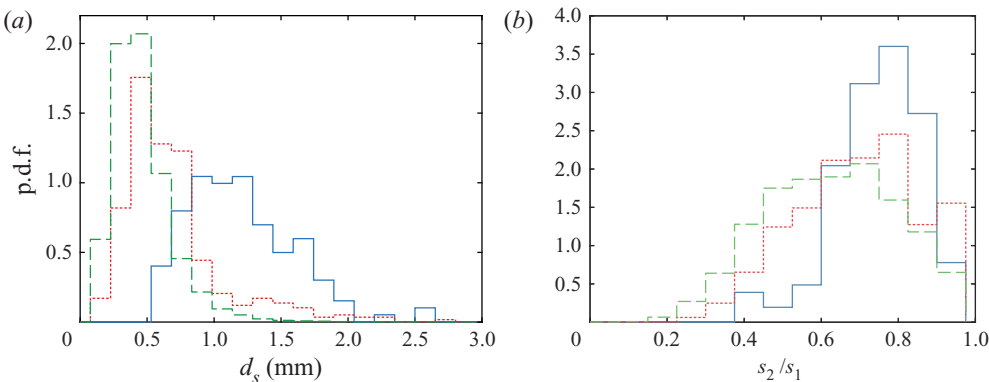


Figure 6. The p.d.f.s of (a) size and (b) aspect ratio of the snow particles for January 2016 (solid blue line), November 2018 (dotted red line) and January 2019 (dashed green line).

data for January 2016 and November 2018 yield comparable particle concentration, while January 2019 presents values one order of magnitude higher.

Table 4 reports key statistics for the snow particle vertical velocity  $w_s$  and acceleration  $a_s$  (where available) obtained from PIV and PTV, respectively. We focus first on the vertical velocities, of which figure 7 shows the p.d.f.s for the three datasets. The distributions are approximately normal, with only November 2018 displaying sizeable skewness ( $S_k$ ). January 2016 has a mean fall speed similar to that of January 2019, but the latter has almost double  $w_{s,rms}$ , reflecting the increased spread of the velocity distribution, with some particles reaching upward velocities. November 2018 shows a mean fall speed almost 60 % higher than for the other two cases, and an intermediate  $w_{s,rms}$ . A comparison between these velocity distributions and the size distributions in figure 6(a) is most interesting. The trends of both quantities are not reconcilable using classic velocity–diameter relationships; in particular, the fact that snow particles from November 2018 fall much faster than in January 2016, while being 40 % smaller on average. Also, the much larger diameter variance in January 2016 is at odds with its relatively narrow velocity distribution; vice versa, January 2019 has the largest spread of velocities while having the narrowest size distribution.

To explain these seemingly counterintuitive results, one may speculate that the snow particles in the datasets have significantly different densities, which could account for the incongruence between the distributions of sizes and fall speeds. However, widely accepted relations between hydrometeor mass  $m$  and diameter,  $m \sim d_s^{b_m}$  (obtained, for example, by combined *in situ* imaging and weighing gauges (Tiira *et al.* 2016; von Lerber *et al.* 2017)), do not support this view. For crystals and small aggregates in the present size range, the

| Dataset       | $W_s = \overline{w_s}$<br>(m s <sup>-1</sup> ) | $w_{s,rms}$<br>(m s <sup>-1</sup> ) | $S_k(w_s)$<br>(—) | $Re_s$<br>(—) | $ W_s /u_{rms}$<br>(—) | $\overline{a_s}$<br>(m s <sup>-2</sup> ) | $a_{s,rms}$<br>(m s <sup>-2</sup> ) |
|---------------|--|-------------------------------------|-------------------|---------------|------------------------|--|-------------------------------------|
| January 2016  | −0.68  | 0.21                                | −0.06             | 54.9          | 4.25                   | −0.012                                   | 0.505                               |
| November 2018 | −1.09  | 0.37                                | −0.23             | 54.5          | 2.87                   | −0.067                                   | 0.442                               |
| January 2019  | −0.71  | 0.55                                | −0.18             | 23.1          | 0.60                   | N/A                                      | N/A                                 |

Table 4. Snow particle properties as measured using DIH for all three datasets.

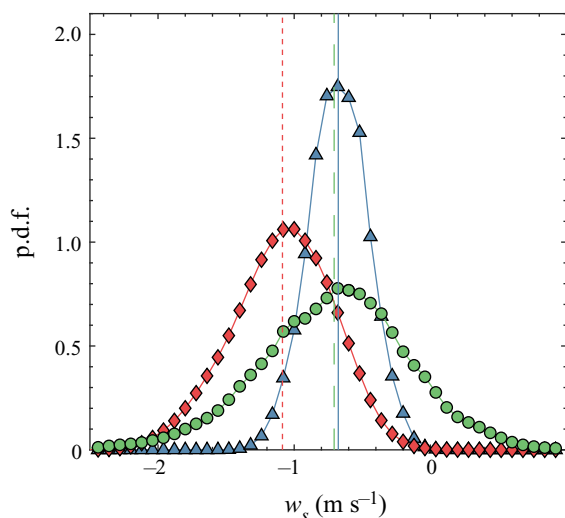


Figure 7. The p.d.f.s of the snow particle vertical velocity ( $w_s$ ) as measured by PIV for January 2016 (blue triangles), November 2018 (red diamonds) and January 2019 (green circles). Vertical solid, dotted and dashed lines mark the mean settling velocity ( $W_s$ ) corresponding to the three datasets, respectively.

exponent  $b_m$  is usually equal to or larger than 2 (Heymsfield *et al.* 2010; Tiira *et al.* 2016; von Lerber *et al.* 2017), implying that density ( $\sim m/d_s^3$ ) decreases less than linearly with size (Heymsfield *et al.* 2004). Neither can nonlinear drag explain the behaviour, because the snow particle Reynolds number  $Re_s = |W_s|d_s/\nu$  is of the same order for the three cases (and almost the same for January 2016 and November 2018).

Alternative scenarios about the nature of the hydrometeors may possibly account for the observations. For example, following the updated Best number formulation by Heymsfield & Westbrook (2010), the differences in measured fall speed could be explained if the November 2018 snow particles had (on average) a density more than four times higher than the January 2016 particles; or if they had correspondingly lower drag coefficients. Both scenarios, however, appear unlikely given the similar micrometeorological conditions, particle Reynolds number and particle aspect ratio of both datasets. While we acknowledge that the particle microphysics should be fully characterized to quantify their influence, it is at least questionable that the snow particle morphology can solely explain the observed fall speeds, through variation in density and/or aerodynamic properties.

We therefore hypothesize that the present behaviour is the result of the influence of air turbulence on snow particle settling. This is certainly consistent with the increasing variance of vertical velocities from January 2016 ( $Re_\lambda = 938$ ) to November 2018 ( $Re_\lambda = 3545$ ) to January 2019 ( $Re_\lambda = 9180$ ), as more intense turbulence is expected to result in

larger spread of the snow particle velocities. Importantly, the differences in mean fall speed can also be explained by the ability of turbulence to enhance the settling velocity of inertial particles. Specifically, the larger fall speed in November 2018 compared to January 2016 may be the consequence of strong (or stronger) preferential sweeping in the former case. Likewise, stronger preferential sweeping in January 2019 than in January 2016 would explain why the former case shows the same mean fall speed as the latter, despite significantly smaller snow particle sizes. Support for this hypothesis is lent by the particle acceleration data and the correlation between concentration and vertical velocity, presented in the following.

In lack of a more precise identification of the snow particle morphology (and hence density), such a hypothesis remains speculative in nature. The holography data, while not providing a direct density estimate, may be used to characterize the hydrometeor morphology in more detail than the shape and aspect ratio presented here. This would provide stronger evidence for any assumption on the microphysics of the tracked hydrometeors, and in turn strengthen the interpretation of the data. Such an analysis is beyond the scope of the present study and will be the basis of future investigations.

### 3.2. Snow particle acceleration

Figure 8 shows the p.d.f.s of the fluctuations of the horizontal acceleration for November 2018 and January 2016 as obtained by PTV, normalized by their r.m.s. values. These are compared to previous numerical and experimental studies of homogeneous turbulence laden with tracers (Mordant, Crawford & Bodenschatz 2004) and inertial particles of known  $St$  (Ayyalasomayajula *et al.* 2006; Bec *et al.* 2006). The long exponential tails of the p.d.f.s for all cases highlight the significant intermittency due to intense turbulence events (Porta *et al.* 2001; Voth *et al.* 2002), modulated by the inertia of the particles (Bec *et al.* 2006; Toschi & Bodenschatz 2009). As discussed in Nemes *et al.* (2017), we can leverage the high sensitivity of the acceleration p.d.f.s to  $St$  (Salazar & Collins 2012) and their low sensitivity to  $Re_\lambda$  and the specific flow configuration (Gerashchenko *et al.* 2008; Volk *et al.* 2008) to estimate the Stokes number of the snow particles. Without aiming for a precise value, we estimate  $St = O(0.1)$  for January 2016 and  $St = O(1)$  for November 2018. Particles with  $St = O(1)$  are known to display the strongest settling enhancement by preferential sweeping (Wang & Maxey 1993; Aliseda *et al.* 2002; Petersen *et al.* 2019). As such, these estimates of  $St$  for the snow particles are consistent with the observation of increased fall speeds in November 2018 compared to January 2019.

The values of the acceleration r.m.s. (table 4) provide a further indication that the snow particles captured in January 2016 are less sensitive to preferential sweeping compared to those in November 2018. In the latter case, the acceleration variance normalized by Kolmogorov scaling is  $a_0 = a_{s,rms}^2 / (\varepsilon^{3/2} \nu^{-1/2}) = 8.5$ . This value is close to the expectation for tracers at such high  $Re_\lambda$  (Ishihara *et al.* 2007; Ireland, Bragg & Collins 2016a), and it results from the action of two opposite effects: particle inertia, which reduces  $a_0$  compared to tracers (Bec *et al.* 2006; Ireland *et al.* 2016a); and the effect of particle trajectories crossing the trajectories of fluid elements due to gravitational drift, which increases  $a_0$  (Ireland, Bragg & Collins 2016b; Mathai *et al.* 2016). The gravitational drift is measured by the ratio of the fall speed to the air velocity fluctuations,  $|W_s|/u_{rms}$ . The January 2016 case has both lower particle inertia (according to our estimates of  $St$ ) and larger gravitational drift through the turbulence (table 4), and indeed it sees a multi-fold increase of the non-dimensional variance,  $a_0 = 37.9$ . A similarly marked rise of  $a_0$  was recently reported for both heavy particles (Ireland *et al.* 2016b) and bubbles

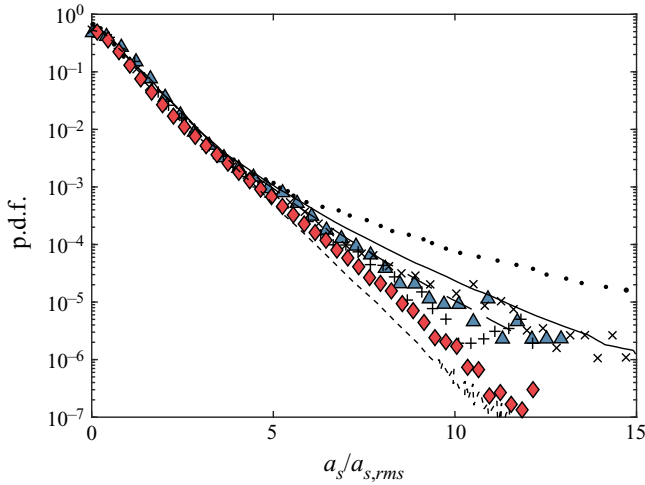


Figure 8. The p.d.f.s of horizontal component of snow particle accelerations for January 2016 (blue triangles) and November 2018 (red diamonds), compared to  $St = 0$  from Mordant *et al.* (2004) (dots), Ayyalasomayajula *et al.* (2006) ( $St = 0.09$ , crosses;  $St = 0.15$ , plus signs) and Bec *et al.* (2006) ( $St = 0.16$ , solid line;  $St = 0.37$ , dashed line;  $St = 2.03$ , dotted line).

(Mathai *et al.* 2016) in homogeneous turbulence, and was explained by the gravitational drift causing the particles to quickly decorrelate from the local turbulence structures, thus experiencing fast-changing fluid motions. This process limits the ability of the particles to obey preferential sweeping, and therefore this mechanism is expected to be less influential for January 2016 than November 2018.

### 3.3. Snow particle concentration

The snow particle concentration fields provide further support to the previous arguments. Because the January 2019 dataset does not allow for locating individual particles, we characterized the concentration using the local and instantaneous image intensity  $I(x, y, t)$ . This approach is based on the observation that scattered light intensity varies linearly with the particle number density  $N$  for monodisperse particles (Bernard & Wallace 2002) and with  $N\overline{a_s}^2$  for polydisperse particles (Raffel, Willert, Scarano, Kähler & Wereley 2018), and it is often used to measure relative concentration in particle-laden flows (e.g. Lai *et al.* 2016). We thus calculate the relative concentration as  $C^* = I/\bar{I}_{\min}$ , where  $\bar{I}_{\min}$  is the 1 min moving average of the image intensity at each location. This normalization helps attenuate temporal fluctuations of the lighting conditions (e.g. due to the power fluctuation of the searchlight) and does not affect the observed trends. We also note that using particle counting for January 2016 and November 2018 leads to the same conclusions for those datasets.

Figure 9(a) presents the p.d.f. of  $C^*$  for the three datasets, indicating different levels of spatio-temporal variability. January 2016 displays an approximately Gaussian distribution (with a kurtosis of 3.5, closely matched to the kurtosis of 3 from a Gaussian distribution), while January 2019 exhibits stretched exponential tails (kurtosis of 4.9), pointing to the significant likelihood of exceptionally low-/high-concentration events. The trend across the cases parallels that of  $Re_\lambda$ : the more intense the turbulence, the higher the variance and intermittency in the concentration fields. Figure 9(b) illustrates a sample  $C^*$  field from January 2019, for which the standard deviation of the concentration exceeds 10 %



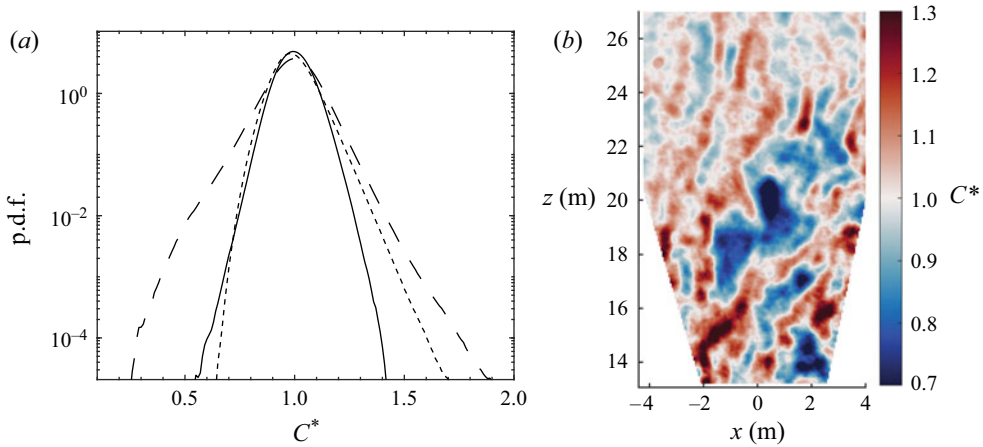


Figure 9. (a) The p.d.f. of snow particle relative concentration  $C^*$ . Black solid, dotted and dashed lines correspond to the January 2016 ( $Re_\lambda = 938$ ), November 2018 ( $Re_\lambda = 3545$ ) and January 2019 ( $Re_\lambda = 9180$ ) data, respectively. (b) Instantaneous  $C^*$  field from January 2019.

of the mean. This representative snapshot clearly shows relatively dense zones, vertically elongated and interleaved with more dilute ones. We characterize such spatial clustering in the next section. Here we stress that the more turbulent cases display stronger clustering, and that the latter is typically concurrent with settling enhancement by preferential sweeping (Aliseda *et al.* 2002; Baker *et al.* 2017; Petersen *et al.* 2019; Momenifar & Bragg 2020).

Direct evidence of preferential sweeping in the more turbulent datasets is provided by the correlation between the local concentration and the simultaneous vertical velocity. In figure 10 we plot the PIV-based vertical velocity of the snow particles conditioned by the local concentration. Because the latter is available at every pixel, we use its spatial average in each PIV interrogation window. The level of correlation between concentration and fall speed is marginal for January 2016, significant for November 2018 and the strongest for January 2019. Analogous trends were reported since the first demonstration of preferential sweeping by Wang & Maxey (1993), and recently confirmed in various simulations and laboratory experiments (among others, Aliseda *et al.* (2002), Baker *et al.* (2017), Huck *et al.* (2018) and Petersen *et al.* (2019)). This is a strong indication that the relatively large settling velocity of November 2018 and January 2019 (compared to diameter-based expectations) is due to the snow particles preferentially sampling downward air flow regions.

### 3.4. Snow particle clustering

In the above we have shown evidence that suggests how snow particles respond to air velocity fluctuations similarly to small inertial particles in turbulence. It is then of interest to characterize the appearance of one of the most striking effects of particle–turbulence interaction: spatial clustering. In the following we describe the properties of clusters identified using approaches analogous to those in previous laboratory studies. While quantitative comparisons are thwarted by the uncertainty on the snow particles Stokes number, we show that the concentration fields display the hallmark features repeatedly reported in laboratory studies.

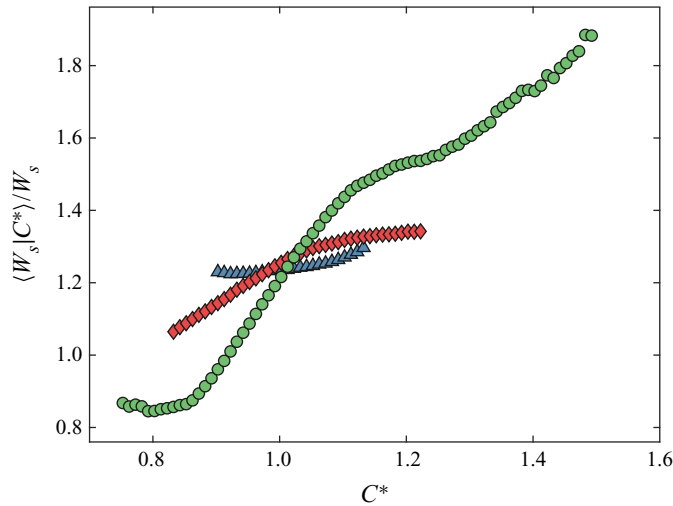


Figure 10. Ensemble-averaged snow particle settling velocity conditioned on value of the local relative concentration  $C^*$  and normalized by the unconditional mean settling velocity. Symbols as in figure 7.

We focus specifically on the January 2019 case, which presents the most intense turbulence and the most inhomogeneous concentration. We follow Aliseda *et al.* (2002) and identify clusters as connected regions where the concentration  $C^*$  is above a prescribed threshold. This approach is standard in image object segmentation, and it has been applied to passive scalars, enstrophy and velocity fluctuations to detect coherent flow structures in turbulence (Catrakis & Dimotakis 1996; Moisy & Jiménez 2004; Lozano-Durán, Flores & Jiménez 2012; Carter & Coletti 2018). In order to select an objective threshold  $C^*_{thold}$ , we analyse the percolation behaviour of the identified objects (Moisy & Jiménez 2004). For higher values of the threshold only a few small clusters are detected, which grow in size and number up to a maximum as the threshold is lowered. They then start to merge, their number decreasing until a single macro-cluster occupies the entire domain. This process is illustrated in figure 11, highlighting the chosen threshold which corresponds to the maximum number of identified clusters (Lozano-Durán *et al.* 2012; Carter & Coletti 2018). We disregard those that touch the image border, as their full spatial extent could be underestimated. The planar nature of the measurement limits our ability to fully characterize the cluster geometry. However, comparison between previous planar measurements characterizing coherent turbulence structures (Carter & Coletti 2018) and inertial particle clusters (Petersen *et al.* 2019) indicates that two-dimensional imaging successfully and quantitatively captures the important features identified in full three-dimensional datasets.

Similar to previous imaging studies of particle-laden turbulence, we consider the p.d.f. of the cluster area  $A_c$  normalized by the Kolmogorov scale (figure 12a). The slope displays a marked change above a length scale corresponding to the light sheet thickness (below which the likelihood of imaging overlapping objects prevents their characterization). Larger clusters show a power-law decay of the area distribution over more than a decade. This suggests self-similarity between clusters of different sizes, pointing to their origin from turbulent eddies (which are also self-similar (Moisy & Jiménez 2004; Baker *et al.* 2017)). The power-law exponent is consistent with the previously reported value of  $-2$  (Monchaux, Bourgoïn & Cartellier 2010; Petersen *et al.* 2019), while applying the same method to a random intensity field would yield a much narrower distribution

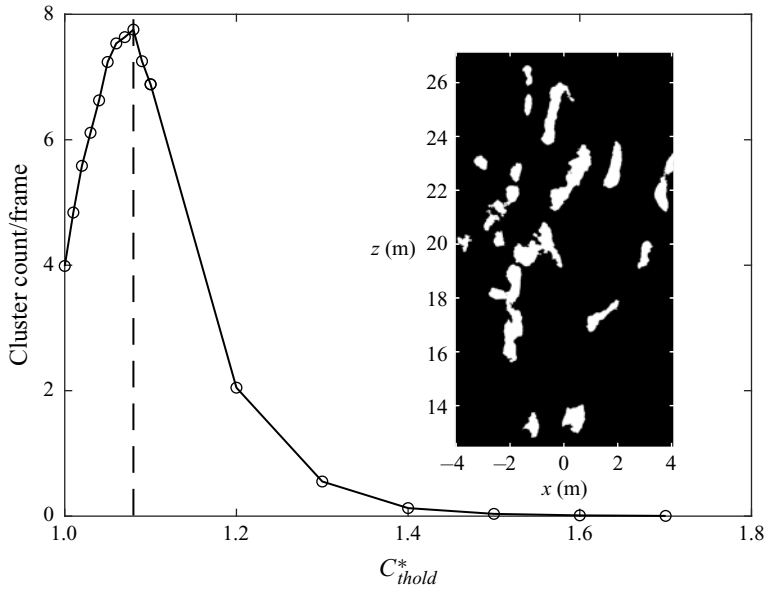


Figure 11. Average number of clusters per image as a function of relative concentration threshold  $C_{thold}^*$ . The inset shows clusters in a binarized concentration field (corresponding to the field shown in figure 9b) using the threshold that maximizes the number of detected clusters (vertical dashed line in the plot).

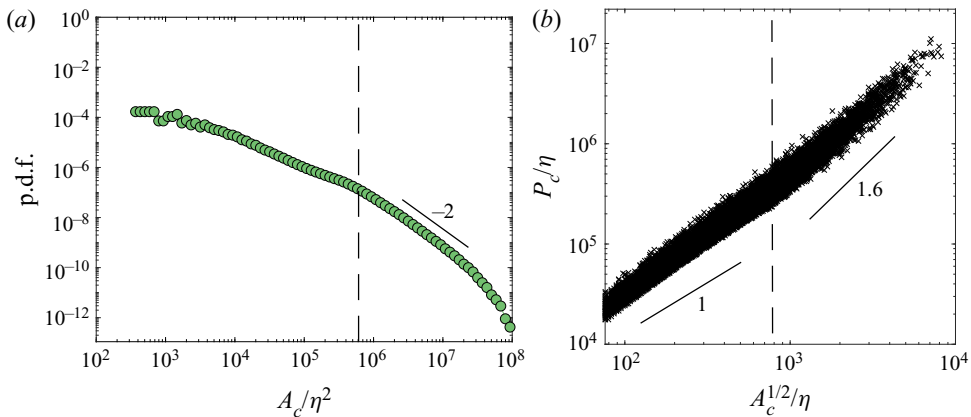


Figure 12. (a) The p.d.f. of snow particle cluster area normalized by Kolmogorov scaling, showing a power-law decay with an exponent close to  $-2$  for sizes larger than the light sheet thickness (vertical dashed line). (b) Scatter plot of cluster perimeter versus square root of the cluster area, both normalized by Kolmogorov scaling.

(Sanada 1991). The detected objects can reach linear dimensions of several metres. While this challenges the classic estimates of inertial particle clusters being  $O(10\eta)$  in size (Eaton & Fessler 1994), there is mounting experimental evidence, from Voronoi tessellation analysis, that the cluster size increases with the flow Reynolds number (Sumbekova *et al.* 2017). Moreover, whereas radial distribution functions (RDFs) calculated from three-dimensional simulations display high levels mostly for separations below  $O(10\eta)$ , this is influenced by the fact that the clusters are typically much thinner in one dimension than in the other two (Baker *et al.* 2017).

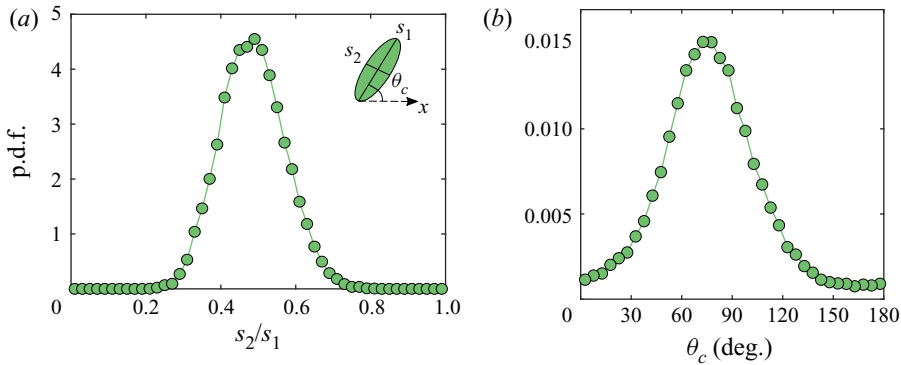


Figure 13. The p.d.f.s of (a) aspect ratio and (b) orientation angle of snow particle clusters.

To further describe the cluster topology, [figure 12\(b\)](#) shows a scatter plot of their perimeters versus the square root of their areas. For small clusters, the data points follow a power law with unit exponent, as expected for regular two-dimensional objects. Such a trend is inherently impacted by the light sheet thickness. For larger ones (especially for sizes far larger than the light sheet thickness) the exponent is significantly higher, indicating a convoluted structure of the cluster borders. This trend was observed in several previous imaging studies of particle-laden turbulence ([Aliseda \*et al.\* 2002](#); [Monchaux \*et al.\* 2010](#); [Petersen \*et al.\* 2019](#)) and is consistent with the view of inertial particle clusters as fractal sets ([Calzavarini \*et al.\* 2008](#)).

Clusters of heavy particles settling in turbulence are known to be elongated and preferentially aligned with the vertical direction ([Woittiez, Jonker & Portela 2009](#); [Dejoan & Monchaux 2013](#); [Ireland \*et al.\* 2016b](#); [Baker \*et al.\* 2017](#); [Petersen \*et al.\* 2019](#)). [Figure 13\(a\)](#) reports the p.d.f. of the cluster aspect ratio (obtained by ellipse-fitting as for the snow particle images from DIH). The peak slightly below 0.5 is consistent with the laboratory study of [Petersen \*et al.\* \(2019\)](#), who found the distribution to be robust for a range of physical parameters. The p.d.f. of the angle  $\theta_c$  made by the ellipse major axis with the horizontal ([figure 13b](#)) confirms a strong prevalence of vertically oriented clusters. The weak prevalence of less than  $90^\circ$  clusters is likely the consequence of the significant wind speed and mean shear.

Finally, we consider the cluster fall speed  $W_c$ , obtained by averaging the vertical velocity at all locations belonging to a given cluster. This is then ensemble-averaged over all clusters of a certain size, normalized by the mean settling velocity  $W_s$  and plotted against the cluster area ([figure 14](#)). Overall, clusters fall significantly faster than  $W_s$ , in keeping with the velocity–concentration correlation shown above. Also, there is an apparent trend of increasing fall speed with cluster size, as seen in laboratory experiments ([Huck \*et al.\* 2018](#); [Petersen \*et al.\* 2019](#)). While this could be merely due to the preferential sampling of downward flow regions, the sharp increase of  $W_c$  above a certain cluster size suggests that a more complex interaction between the snow and the turbulent air may be taking place.

The estimated particle volume fraction for January 2019 is above  $10^{-6}$  ([table 3](#)) and it may approach  $10^{-5}$  in the denser clusters. According to widespread criteria for gas–solid flows ([Elghobashi 1994](#); [Balachandar & Eaton 2010](#)), in this range of concentrations the dispersed phase is expected to exert a significant back-reaction on the carrier fluid (so-called two-way coupling). Therefore, the larger clusters, by virtue of the simultaneous action of large numbers of particles, may be able to collectively drag air down with them, enhancing their fall speed beyond that granted by one-way coupling mechanisms



## 4. Conclusions

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simulations: power-law size distribution, fractal-like silhouette, vertical elongation and large fall speed that increases with size.

Taken together, these results confirm and extend the conclusion of Nemes *et al.* (2017): the phenomenology of inertial particles in turbulence, built over decades of canonical flow studies, is largely applicable to the dynamics of snow settling in air. Presently, none of these well-known concepts are incorporated in weather forecasting models. Other recent studies, also imaging-based, demonstrated that those concepts are in fact directly applicable to atmospheric flows: within clouds, clustering of droplets was recently observed using airborne holographic instruments (Beals *et al.* 2015; Larsen *et al.* 2018), and elongated regions devoid of droplets were identified at a mountain-top station (Karpińska *et al.* 2019).

Environmental flows give access to much larger ranges of scales than laboratory or numerical experiments, enabling the exploration of fundamental fluid mechanics questions. To our knowledge, the present field work represents the highest Reynolds number flow measurements in which inertial particle clustering is observed and quantified. The fact that we observe very large clusters (up to the integral scales of the turbulence) lends support to recent claims that these clusters grow larger with  $Re_\lambda$  (Sumbekova *et al.* 2017), and that increasing  $Re_\lambda$  extends the range of scales to which the particles respond (Tom & Bragg 2019). Of course, the non-canonical aspects of naturally occurring particle-laden flows have to be considered. In particular, the morphology of snow particles is expected to play an important role, especially for complex-shaped and elongated snow particles (Westbrook & Sephton 2017). Our understanding of the interaction of anisotropic particles with turbulence has seen tremendous progress in recent years (Voth & Soldati 2017), and imaging studies capable of testing these dynamics in the field are warranted. This will require three-dimensional imaging at high spatial and temporal resolution, and will benefit from novel capabilities now available for Lagrangian tracking (see e.g. Guala *et al.* (2008a) and the recent review by Discetti & Coletti (2018)).

An important aspect that the present study cannot directly address is represented by the hydrometeor collision rate, and the impact that turbulence has on it. This is expected to be strongly related to the polydispersity of the snow particles. Polydispersity drives the classic (gravitational) mechanism by which larger particles fall faster than and collide with smaller ones (Pruppacher & Klett 1997). Recent simulations, however, show that turbulence enhances the relative velocity of polydisperse particles also in the horizontal direction (Dhariwal & Bragg 2018). Fundamental studies in this area are needed to extend the applicability of particle–turbulence dynamics to environmental flows. Additionally, field studies are warranted to investigate the correlation between the level of collision-driven aggregation and atmospheric turbulence.

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#### Author ORCIDs.

- ID Tim Berk <https://orcid.org/0000-0002-7159-2360>;
- ID Aliza Abraham <https://orcid.org/0000-0002-6584-3661>;
- ID Michael Heisel <https://orcid.org/0000-0002-4200-5550>;
- ID Filippo Coletti <https://orcid.org/0000-0001-5344-2476>;
- ID Jiarong Hong <https://orcid.org/0000-0001-7860-2181>.



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