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A radioactive beam of <sup>24</sup>Na with 90% of its content in its 1<sup>+</sup> isomeric state ( $E_{ex} = 0.472 \text{ MeV}$ ,  $t_{1/2} = 20.18 \text{ ms}$ ) has been developed and used to perform a measurement of the <sup>24</sup>Na<sup>m</sup>(d,p)<sup>25</sup>Na reaction at the John D. Fox Accelerator Laboratory at Florida State University. This reaction selectively populated  $\ell = 0$  transfers, allowing the study of low-spin states in <sup>25</sup>Na. Mirror symmetry arguments were then used to investigate the effects of the isomeric state of <sup>24</sup>Al ( $E_{ex} = 0.426 \text{ MeV}$ ,  $t_{1/2} = 130 \text{ ms}$ ) on the astrophysical rate of the <sup>24</sup>Al<sup>m</sup>(p, $\gamma$ )<sup>25</sup>Si reaction. Experimental parameters were extracted to provide, for the first time, an experimental reaction rate for the destruction of <sup>24</sup>Al via proton captures in its isomeric state relevant to rp-process nucleosynthesis.

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I. INTRODUCTION

The rapid proton capture (rp)-process occurs in hot <sup>57</sup> hydrogen-rich environments at temperature in excess of <sup>58</sup> 0.1 Giga-Kelvin (GK) [1]. X-ray bursts [2–4], novae and <sup>59</sup> supernovae outbursts [1], and mergers between neutron <sup>60</sup> stars and main sequence stars [5], have been proposed as <sup>61</sup> sites for this nucleosynthesis process. <sup>62</sup>

The rp-process starts at the breakout from the hot <sup>63</sup> 22 CNO cycle into the Ne-Na region, proceeding up the <sup>64</sup> 23 proton-rich side of stability via a series of proton-capture <sup>65</sup> 24 reactions and  $\beta$ -decays [2]. One of the nuclear reactions <sup>66</sup> 25 along the rp-process path, out of the Ne-Na region is the <sup>67</sup> 26  $^{24}$ Al(p, $\gamma$ ) $^{25}$ Si reaction [6]. Variations in this rate affect <sup>68</sup> 27 relative end-point abundances of <sup>28,29,30</sup>Si, <sup>33,34</sup>S, and <sup>69</sup> 28 <sup>36</sup>Ar in ONe novae [7]. In particular, the abundances <sup>70</sup> 29 of <sup>29,30</sup>Si to <sup>28</sup>Si are important in the identification of <sup>71</sup> 30 72 presolar grains in comets and asteroids [8]. 31

The effect of nuclear isomers in astrophysical processes  $^{\rm 73}$ 32 is not well understood. A recent theoretical study con-  $^{74}\,$ 33 cluded that the presence of isomeric states in stellar nu-  $^{75}$ 34 cleosynthesis scenarios can significantly impact the cal-  $^{76}\,$ 35 culation of the reaction rates due to their unique nuclear  $^{77}\,$ 36 properties [9]. Such is the case of the  ${}^{24}\text{Al}(p,\gamma)^{25}\text{Si}$  reac-37 tion. The existence of a low-lying isomeric state in <sup>24</sup>Al <sup>79</sup> 38  $(^{24}\text{Al}^m, \text{E}_{ex} = 0.426 \text{ MeV}, \text{t}_{1/2} = 130 \text{ ms}, \text{J}^{\pi} = 1^+)$  with <sup>80</sup> 39 a large difference in spin from the ground state (  $^{24}\mathrm{Al}^{gs},\,^{\mathrm{st}}$ 40  $t_{1/2} = 2.053$  hr,  $J^{\pi} = 4^+$ ), complicates the calculation of <sup>82</sup> 41 this reaction rate. 42

<sup>43</sup> The main contribution to the <sup>24</sup>Al(p, $\gamma$ ) reaction rate <sup>84</sup> <sup>44</sup> proceeds through low-lying resonances above the proton <sup>85</sup> <sup>86</sup> separation energy in <sup>25</sup>Si. It is expected that proton-<sup>86</sup> <sup>87</sup> captures on <sup>24</sup>Al<sup>gs</sup> and <sup>24</sup>Al<sup>m</sup> proceed through different <sup>87</sup> <sup>87</sup> resonances in <sup>25</sup>Si, therefore contributing separately to <sup>88</sup> the rate of destruction of <sup>24</sup>Al via proton capture reac-<sup>49</sup> tions, as was shown experimentally to be the case for the <sup>88</sup> <sup>50</sup> <sup>26</sup>Al<sup>m</sup>(p, $\gamma$ )<sup>27</sup>Si reaction [10, 11].

In rp-process nucleosynthesis, <sup>24</sup>Al is reached <sup>89</sup> through the <sup>23</sup>Mg(p, $\gamma$ )<sup>24</sup>Al reaction as well as the <sup>90</sup> <sup>22</sup>Mg(p, $\gamma$ )<sup>23</sup>Al(p, $\gamma$ )<sup>24</sup>Si( $\beta$ )<sup>24</sup>Al reaction chain [2, 5]. The <sup>91</sup> correct calculation of the <sup>24</sup>Al<sup>m</sup>(p, $\gamma$ )<sup>25</sup>Si reaction rate is <sup>92</sup> particularly important in the latter branch since the isomeric state in <sup>24</sup>Al is strongly populated by the  $\beta$ -decay of <sup>24</sup>Si as shown by rate calculations of Ref. [9, 12].

The <sup>24</sup>Al<sup>gs</sup> $(p,\gamma)$ <sup>25</sup>Si reaction rate has been the object of few previous studies. In the recent work by Longfellow et al. [6], states in <sup>25</sup>Si were studied using  $\gamma$ -ray spectroscopy, refining the experimental information previously reported by Benenson et al. [13], and determining  $\gamma$ -decays and branching ratios for several excited states of <sup>25</sup>Si. Above the proton separation threshold ( $S_p =$ 3.414(10) MeV) in the region of astrophysical relevance, two states, a  $9/2^+$  at  $E_{ex} = 3.695(14)$  MeV and a  $1/2^+$  at  $E_{ex} = 3.802(11)$  MeV, were identified. The results of that study were used to constrain the rate of the  $^{24}\text{Al}^{g}(\mathbf{p},\gamma)^{25}\text{Si}$  reaction [6], showing that the contribution of the  $9/2^+$  state was a factor of 10 higher than the one used in the previous network calculations performed by Herndl et al. [14]. Knapton et al. [15] have studied the <sup>24</sup>Na(d,p)<sup>25</sup>Na reaction to infer spectroscopic information on the mirror nucleus <sup>25</sup>Si, with a beam that was 100% in the <sup>24</sup>Na ground state. No information was previously available on the reactions of the <sup>24</sup>Na isomeric state.

In this work, the  ${}^{24}\text{Al}^m(\mathbf{p},\gamma){}^{25}\text{Si}$  reaction was studied via the measurement of the  ${}^{24}\text{Na}^m(\mathbf{d},\mathbf{p}){}^{25}\text{Na}$  reaction using a  ${}^{24}\text{Na}$  beam with 90% of its content in its isomeric state. Spectroscopic information of states in  ${}^{25}\text{Na}$ populated by single-neutron transfer on the  ${}^{24}\text{Na}^m$  was extracted. These states are mirror to states in  ${}^{25}\text{Si}$  populated by the  ${}^{24}\text{Al}^m(\mathbf{p},\gamma){}^{25}\text{Si}$  reaction. The reported experimental information constraints for the first time the destruction rate of  ${}^{24}\text{Al}$  via proton captures on its isomeric state.

#### **II. EXPERIMENT**

The measurement of the  ${}^{24}$ Na ${}^m$ (d,p) ${}^{25}$ Na reaction was performed at the John D. Fox Accelerator Laboratory at Florida State University. A primary beam of stable  ${}^{23}$ Na was accelerated by the FN Tandem Van de Graff

accelerator followed by the linear accelerator (LINAC) 93 to an energy of 115 MeV. The primary <sup>23</sup>Na beam was 94 then sent to the RESOLUT radioactive beam facility [16]. 95 where it was incident on a production target filled with 96 deuterium gas to produce a radioactive beam of  $^{24}$ Na via 97 the  ${}^{23}Na(d,p){}^{24}Na$  reaction in-flight [17]. The production 98 target was 40 mm long, with 2.5 micron HAVAR entrance 99 and exit windows, was cooled with liquid nitrogen to 77K, 100 and was kept at 350 Torr. 101

The resultant <sup>24</sup>Na beam was then tuned using the focusing elements of RESOLUT and sent downstream to the reaction chamber. The 11<sup>+</sup> charge state of <sup>24</sup>Na arrived at the target position with 85.5 MeV. The main contaminant of the beam was the 10<sup>+</sup> charge state of the primary <sup>23</sup>Na at 73.8 MeV.

The beam was incident on a 517  $\mu g/cm^2$  CD<sub>2</sub> target. 108 In the reaction chamber, a double-sided 300  $\mu$ m thick 109 Micron S2 silicon detector was placed 10.5 cm upstream 110 from the target position to measure charged reaction par-111 ticles at backward angles. The angular coverage of the sil-112 icon detector was  $161.6^{\circ}$  to  $173.7^{\circ}$  in the lab frame. Out-113 side of the reaction chamber, two Sodium Iodide (NaI) 114 detectors were placed at close to  $90^{\circ}$  above and to the 115 side of the target position to monitor the isomeric con-116 tent of the beam via the detection of the 472 keV  $\gamma$ -rays<sub>151</sub> 117 from the decay of the isomeric state to the ground state<sub>152</sub> 118  $(t_{1/2} = 20.18 \text{ ms})$  of the <sup>24</sup>Na beam [18]. Downstream<sub>153</sub> 119 from the target position, an ionization chamber collected<sub>154</sub> 120 the unreacted beam as well as the heavy reactants. The<sub>155</sub> 121 ionization chamber had a 8 micron Kapton window, and 156 122 consisted of two 40 mm position sensitive sections, an  $80_{157}$ 123 mm section to measure energy loss ( $\Delta E$ ), and a 200 mm<sub>158</sub> 124 section to fully stop the beam (E). The ionization cham-<sub>159</sub> 125 ber was filled with isobutane, and was kept at a  $\text{pressure}_{160}$ 126 of 45 Torr. The two position sensitive sections were  $not_{161}$ 127 used in the analysis of this experiment. A schematics  $of_{162}$ 128 the experimental setup is shown in Fig. 1. 129 163

The <sup>24</sup>Na beam and its main contaminant, the primary<sub>164</sub> 130  $^{23}$ Na beam, were well separated in the ionization cham- $_{165}$ 131 ber by their energy losses as is shown in Fig. 2a. Time-166 132 of-flight information from the production target to the<sub>167</sub> 133 detectors was also used to differentiate the beam com-168 134 ponents. The ratio of <sup>24</sup>Na to <sup>23</sup>Na measured in the<sub>169</sub> 135 ionization chamber throughout the experiment was ap-170 136 proximately 1:1. 137 171

The isomeric content of the beam was determined us-172 138 ing sets of 2 minutes synchronized runs taken at various173 139 points during the experiment. For this purpose, a thick174 140 gold target was placed at the target position to fully stop<sub>175</sub> 141 the beam. The NaI detectors, placed close to  $90^{\circ}$  directly<sup>176</sup> 142 outside the reaction chamber, measured the 472 keV  $\gamma$ -177 143 rays characteristic of the decay of the isomeric state in178 144 <sup>24</sup>Na to its ground state. Fig. 2b shows a typical spec-179 145 trum obtained with one of the NaI detectors during a180 146 gold target run. A peak corresponding to the 472 keV<sub>181</sub> 147  $\gamma$ -ray is observed. A no-target run immediately followed<sub>182</sub> 148 the gold target measurement. The target was removed<sub>183</sub> 149 allowing the full beam to pass directly to the ionization<sub>184</sub> 150

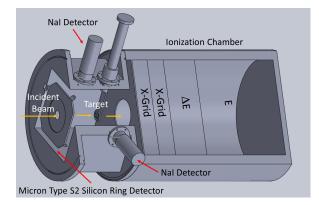


FIG. 1. Schematics of the experimental setup used during the present  $^{24}$ Na(d,p) $^{25}$ Na experiment. The beam enters the reaction chamber from the left, and is incident on a CD<sub>2</sub> target. Backward scattered protons from the interaction of the beam with the target are measured in a silicon S2 detector. The heavy products as well as the unreacted beam are measured downstream in an ionization chamber. Two NaI detectors were placed outside the reaction chamber at ~ 90° from the target position and are used to monitor the isomeric content of the beam via detection of the 472 keV  $\gamma$ -rays.

chamber to measure the total amounts of <sup>23</sup>Na and <sup>24</sup>Na. Fig. 2a shows a typical spectrum taken in the Ionization Chamber with no target. Additionally, during the notarget runs, the NaI detectors measured the background  $\gamma$ -rays in order to filter out any non-target related contribution to the 472 keV  $\gamma$ -ray spectrum. After the experiment, calibrated sources were placed at the target position to obtain the absolute efficiencies of the NaI detectors. From the sets of synchronized runs, it was determined that 90%  $\pm$  10% of the <sup>24</sup>Na beam was in the isomeric state.

In order to confirm the experimentally obtained isomeric content of the beam, Distorted Wave Born Approximation (DWBA) calculations were also performed for the <sup>23</sup>Na(d,p)<sup>24</sup>Na reaction. The DWBA code Fresco [19] was utilized to perform these calculations. The optical model parameters for the incoming <sup>23</sup>Na+d and outgoing <sup>24</sup>Na+p channels were taken from Ref. [20].

The overall DWBA calculated yields for the isomeric and ground states in <sup>24</sup>Na as well as the DWBA calculated isomeric content are shown in Fig. 3. The experimentally determined isomeric content is also shown. There is good agreement between the experimental measurement and the DWBA calculated yields of the isomeric ratio. In addition, the DWBA calculations show that as the energy varies, the isomeric content of the beam varies smoothly, thus small changes in the production energy will have no significant effect in the overall isomeric content of the beam. This observation contrasts that of the production of an isomeric beam in <sup>26</sup>Al (<sup>26</sup>Al<sup>m</sup>) via the <sup>26</sup>Mg(p,n) reaction [21].

The <sup>24</sup>Na(d,p)<sup>25</sup>Na reaction was measured using a 517  $\mu$ g/cm<sup>2</sup> thick CD<sub>2</sub> target which was bombarded with a 85.5 MeV <sup>24</sup>Na beam. The absolute normalization of the

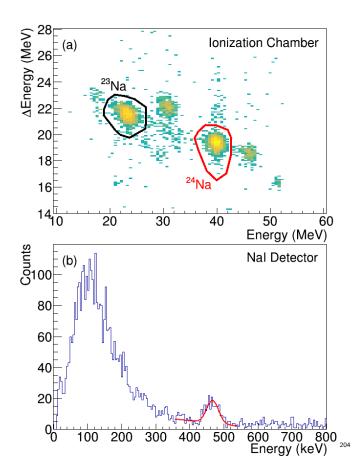


FIG. 2. (a) Ionization Chamber Spectrum for a no-target two<sup>205</sup> minute run. The <sup>24</sup>Na beam as well as its main contaminant,<sup>206</sup> the primary <sup>23</sup>Na beam, are indicated. The ratio of <sup>23</sup>Na to<sub>207</sub> <sup>24</sup>Na was about 1:1. (b) Typical spectrum of one of the NaI<sub>208</sub> detectors taken during a gold-target two minute run. The<sub>209</sub> peak indicated in red correspond to the 472 keV  $\gamma$ -ray in <sup>24</sup>Na which was used to determine and monitor the isomeric content of the beam.

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<sup>24</sup>Na beam was performed using the <sup>23</sup>Na(d,p)<sup>24</sup>Na reac-<sub>215</sub> 185 tion, which has been previously studied in Refs. [20, 22],216 186 and which was measured through the  $^{23}$ Na component<sub>217</sub> 187 of the beam. Two states strongly populated by the<sub>218</sub> 188 <sup>23</sup>Na(d,p) reaction in <sup>24</sup>Na at  $E_{ex} = 1.34$  MeV,  $J^{\pi} =_{219}$  1<sup>+</sup> and  $E_{ex} = 1.846$  MeV,  $J^{\pi} = 2^+$  were observed in the<sub>220</sub> 189 190 silicon detector when gating on the  $^{23}$ Na beam compo-<sub>221</sub> 191 nent. 192 222

Cross sections for these states were extracted and<sup>223</sup> 193 normalized using DWBA calculations with the optical<sup>224</sup> 194 model parameters and spectroscopic factors given in225 195 Refs. [20, 22]. The total amount of <sup>23</sup>Na beam was<sup>226</sup> 196 then obtained by taking into account the target thick-227 197 ness and solid angle coverage of the silicon detector in<sup>228</sup> 198 the present experiment for both states. The absolute<sup>229</sup> 199 <sup>24</sup>Na beam normalization was then calculated using the<sub>230</sub> 200 <sup>24</sup>Na to <sup>23</sup>Na ratio measured throughout the experiment<sub>231</sub> 201 in the ionization chamber. The typical intensity of the<sup>232</sup> 202 <sup>24</sup>Na beam was determined to be  $\sim 800$  pps. 233 203

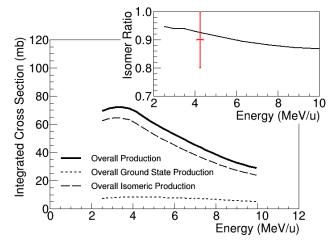


FIG. 3. DWBA calculations for the production of  $^{24}$ Na beam via the  $^{23}$ Na(d,p) $^{24}$ Na reaction. The production cross section as a function of the beam energy for the ground state (short dashed line), isomeric state (long dashed line), and total production of  $^{24}$ Na (solid line), are shown. In the inset, the DWBA calculated isomer ratio, or isomer to total ratio, is shown by the solid black line. The experimentally determined isomeric ratio is shown by the single point. Good agreement is observed between DWBA calculations and the experimental data point.

### III. RESULTS

States in <sup>25</sup>Na populated in the present experiment via the <sup>24</sup>Na(d,p) reaction with a beam of <sup>24</sup>Na with 90% of its content in its isomeric 1<sup>+</sup> state were measured in the silicon detector in the angular range of  $\theta_{lab} = 161.6^{\circ} -$ 173.7°. The energy of the measured protons was then converted to apparent excitation energy in <sup>25</sup>Na using the Q-value of the isomeric state in <sup>24</sup>Na.

Fig. 4a shows the <sup>25</sup>Na apparent excitation energy spectrum obtained in the present experiment. A large background peak can be seen at energies above 5 MeV (low measured energies). This background peak arises from the  $\beta$ -decay of the <sup>24</sup>Na<sup>gs</sup> to <sup>24</sup>Mg. Although most of the beam is in the <sup>24</sup>Na isomeric state, it decays to the ground state in <sup>24</sup>Na via the emission of a 472 keV  $\gamma$ -ray with t<sub>1/2</sub> = 20.18 ms, where it subsequently  $\beta$ -decays to <sup>24</sup>Mg with t<sub>1/2</sub> = 14.997 hr. Over the course of the experiment, <sup>24</sup>Na in the ground state accumulated in the reaction chamber, providing the source of this  $\beta$ -decay background.

A run with no beam on target was taken to measure the  $\beta$ -decay background in the silicon detector. The shape of the  $\beta$ -decay spectrum shows good agreement with the high-energy structure in the excitation energy spectrum. This background spectrum was then scaled to the peak observed in the high energy portion of the excitation energy spectrum (red solid line) and subtracted. Fig. 4b shows the apparent excitation energy spectrum with the  $\beta$ -decay background subtracted.

The contribution of the ground state component of the

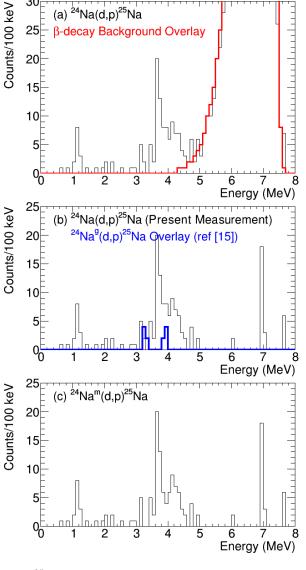
 $^{24}$ Na beam (10% of the total  $^{24}$ Na beam content) was 234 estimated using the results from Ref. [15], where the 235  $^{24}$ Na(d,p) reaction was measured using a pure  $^{24}$ Na<sup>gs</sup> 236 beam. From that work's results, we estimated that con-237 tributions from the ground state component of the beam 238 were negligible given the low-intensity of our <sup>24</sup>Na beam. 239 The estimated contribution of the  ${}^{24}$ Na<sup>gs</sup>(d,p)<sup>25</sup>Na to 240 the present experiment is also shown in Fig. 4b (blue 241 solid line). The <sup>25</sup>Na apparent excitation energy spec-242 trum from our present  ${}^{24}Na^m(d,p){}^{25}Na$  measurement is 243 shown in Fig. 4c, where both the contributions from the 244  $\beta$ -decay background and the ground state component of 245 the beam have been subtracted. 240

The results in Fig. 4 show the high selectivity in the 248 states populated in the present  ${}^{24}Na^m(d,p){}^{25}Na$  measure-249 ment. Three states are observed to be populated by the 250 isomeric <sup>24</sup>Na beam, plus a fourth possible state. The 251 three observed states are identified as the 1.069 MeV  $J^{\pi}$ 252  $= 1/2^{+}$  state, the 3.687 MeV J<sup> $\pi$ </sup> = 3/2<sup>+</sup> state, and the 253 the 4.289 MeV  $J^{\pi} = 1/2^+$  state. Angular distributions for 254 these strongly populated states were fitted using DWBA 255 calculations with the code Fresco [19]. For the incoming 256 and outgoing channels, optical model potential parame-257 ters were taken from Refs. [23, 24]. Fig. 5 shows the 258 angular distributions for these three states along with 259 the DWBA fits to the angular distributions. The spec-260 troscopic factors were extracted using a chi-square fit to 261 the experimental data and are listed in Table I. 262

Even with the limited statistics of the present experi-263 ment, the angular distributions confirm that  $\ell = 0$  trans-264 fers are selectively populated in the present reaction. 265

Shell model calculations were also performed using 266 the USDB interaction [25, 26]. From these calculations, 267 good agreement is found between the energies and spec-268 troscopic factors of states predicted by the shell model 269 and the three observed states in the present experiment. 270 Good agreement is also obtained between the experimen-271 tally extracted spectroscopic factors and the ones pre-272 dicted by the shell model. A comparison between states 273 and energies extracted from the experiment and with 274 shell model calculations is shown in Fig. 6 and listed 275 in Table I. The additional fourth possible state at  $E_{ex} =$ 276 3.950 MeV has previously been reported Refs. [18, 27]. 277 Using shell model calculations, this state is expected to 278 be a  $3/2^+$  state. Fig. 6 shows the apparent excitation en-279 ergy spectrum for states populated by the isomeric com-280 ponent of the beam with bars for the spectroscopic fac-281 tors overlayed. The yellow bars show the spectroscopic 282 factors for the experimentally observed  $\ell = 0$  transfers, 283 and the dark red and orange bars show the shell model 284 calculated spectroscopic factors for  $\ell = 0$  and  $\ell = 2$  trans-285 fers respectively. 286

The high selectivity of the present data allows us to 287 propose mirror level assignments for states in the <sup>25</sup>Na -288 <sup>25</sup>Si system. A diagram of the states in <sup>25</sup>Na, <sup>25</sup>Si, and 289 shell model calculations with mirror level assignments us-290 ing data from Refs. [6, 13, 18, 25, 27] is shown in Fig. 7. 291 For the shell model states, only those with  $C^2S > 0.075$ 292



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(a) <sup>24</sup>Na(d,p)<sup>25</sup>Na

FIG. 4. <sup>25</sup>Na apparent excitation energy spectrum from the  $^{24}\mathrm{Na}(\mathrm{d},\mathrm{p})^{25}\mathrm{Na}$  reaction using the Q-value from the isomeric state, in the angular range of  $\theta_{lab} = 161.6^{\circ} - 173.7^{\circ}$ , measured in the present work. (a) The spectrum for states in <sup>25</sup>Na populated in the present  $^{24}$ Na(d,p) reaction. The contribution of  $\beta$ -decay background from the decay of the <sup>24</sup>Na<sup>gs</sup> to <sup>24</sup>Mg, scaled to fit the data (red thick line) is observed at energies above 5 MeV. (b) <sup>25</sup>Na apparent excitation energy spectrum after the subtraction of the  $\beta$ -decay background contribution. The contribution from the ground state component of the beam is also shown (blue thick line). The ground state contribution was estimated from the work by Ref. [15]. Given the kinematics used, (Q-value of the isomeric state), states populated by the ground state appear shifted up in energy by 472 keV. (c) The  ${}^{25}$ Na excitation energy spectrum populated by the  ${}^{24}$ Na<sup>*m*</sup>(d,p) ${}^{25}$ Na reaction populated with only the isomeric component of the beam.

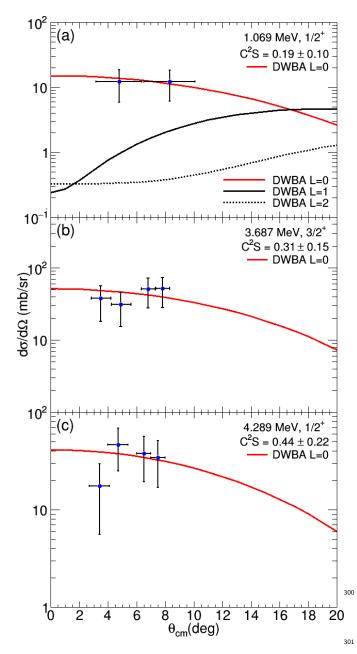


FIG. 5. Angular distributions for the states populated in the present <sup>24</sup>Na<sup>m</sup>(d,p)<sup>25</sup>Na reaction at (a)  $E_{ex} = 1.069 \text{ MeV}_{302}$  $(1/2^+)$ , (b)  $E_{ex} = 3.687 \text{ MeV} (3/2^+)$ , and (c)  $E_{ex} = 4.289_{303}$ MeV  $(1/2^+)$ . All three states show  $\ell = 0$  neutron transfers<sub>304</sub> from the 1<sup>+</sup> isomeric state in <sup>24</sup>Na. DWBA calculations were<sub>305</sub> used to fit the experimental data. A chi-square minimization was used to determine the best value for the spectroscopic factors (C<sup>2</sup>S).

were used. The observed  $1/2^+$  states in <sup>25</sup>Na at  $E_{ex} =_{311}$ 1.069 MeV, and  $E_{ex} = 4.289$  MeV are the mirror levels<sup>312</sup> of states in <sup>25</sup>Si at  $E_{ex} = 0.87$ , and  $E_{ex} = 3.802$  MeV<sup>313</sup> with spectroscopic factors for an  $\ell = 0$  transfer of C<sup>2</sup>S =<sub>314</sub> 0.19 and 0.44 respectively. The mirror level of the  $E_{ex^{315}}$ = 3.687 MeV state in <sup>25</sup>Na ( $\ell = 0$  C<sup>2</sup>S = 0.31) has not<sup>316</sup> been observed in <sup>25</sup>Si. These states are predicted by shell<sup>317</sup>

TABLE I. Spectroscopic factors for states observed in the  $^{24}$ Na<sup>m</sup>(d,p)<sup>25</sup>Na reaction. Both experimentally determined and USDB shell model spectroscopic factors are shown. For shell model states, only spectroscopic factors greater than 0.075 are considered.

Excitation		Experiment $C^2S$	USDB	$SM C^2S$
Energy (MeV)	$J^{\pi}$	$\ell = 0$	$\ell = 0$	$\ell = 2$
1.069	$1/2^{+}$	$0.19 \pm 0.10$	0.303	0.001
3.687	$3/2^+$	$0.31\pm0.15$	0.253	0.145
3.955	$(3/2^+)^a$		0.095	0.016
4.289	$1/2^{+}$	$0.44\pm0.22$	0.329	0.017

<sup>a</sup>Spin from shell model calculations.

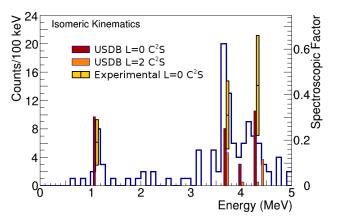


FIG. 6. <sup>25</sup>Na apparent excitation energy. The states in <sup>25</sup>Na populated in the present <sup>24</sup>Na<sup>m</sup>(d,p)<sup>25</sup>Na reaction in the energy range  $E_{cm} = 0$  - 5 MeV are compared with shell model predictions using the USDB interaction. Spectroscopic factors (right y-axis) extracted from fits to the experimental data (yellow bars) and predicted by shell model (dark red bars for  $\ell = 0$  transfers, orange for  $\ell = 2$ ) are overlaid.

model calculations [25].

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## IV. ASTROPHYSICAL IMPLICATIONS

In order to evaluate the contribution of the 1<sup>+</sup> isomeric state in <sup>24</sup>Al to the <sup>24</sup>Al(p, $\gamma$ )<sup>25</sup>Si reaction rate, we focused on the states in <sup>25</sup>Na that are mirrors to the states above the proton threshold in <sup>25</sup>Si (S<sub>p</sub> = 3.414 MeV) which are expected to dominate the astrophysical rate of the <sup>24</sup>Al(p, $\gamma$ )<sup>25</sup>Si reaction.

The observed  $1/2^+$  state at  $E_{ex} = 4.289$  MeV in <sup>25</sup>Na is mirror to the one at  $E_{ex} = 3.802$  MeV in <sup>25</sup>Si. After taking into account the energy of the isomeric state of <sup>24</sup>Al ( $E_{ex} = 0.426$  MeV), we find an energy with respect to the isomer of  $E_r^m = -0.038$  MeV, making it a subthreshold resonance. The contribution of this state to the reaction rate was calculated using a Breit-Wigner sub-threshold resonance formalism as described by Refs. [30, 31]. The effect of this  $1/2^+$  sub-threshold resonance is shown by the red line in Fig. 8a, where the ratio of

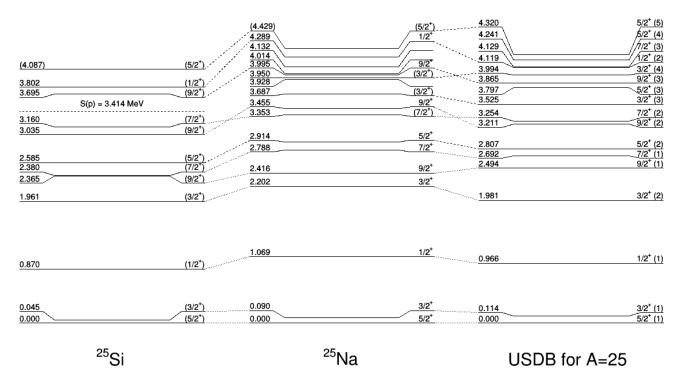


FIG. 7. Level schemes for <sup>25</sup>Si and <sup>25</sup>Na, and USDB shell model calculations for the A = 25 system. Level information for <sup>25</sup>Si and <sup>25</sup>Na taken from Refs. [6, 18, 27]. USDB level information from [25]. The dotted lines indicate mirror levels determined by Ref. [13]. The dashed lines indicate proposed mirror levels. The energies of the  $5/2^+$  state at 4.087 MeV in <sup>25</sup>Si and 4.429 MeV in <sup>25</sup>Na are based on the USDB shell model calculations with Thomas-Ehrman shifts of -0.479 MeV and 0.224 MeV respectively [28, 29].

the rate calculated of the  $1/2^+$  resonance state at  $E_r^m = _{346}$ -0.038 MeV in <sup>25</sup>Si to that of the current ground state<sub>347</sub> REACLIB rate is shown, for a temperature range from<sub>348</sub> 221 0.01 GK to 10 GK. 349

In addition to our experimentally observed states, the<sup>350</sup> 322 USDB shell model calculation [25, 26] predicts a  $5/2^{+351}$ 323 state that could contribute to the  ${}^{24}\text{Al}^m(p,\gamma){}^{25}\text{Si}$  reac- ${}^{352}$ 324 tion rate. Using a Thomas-Ehrman shift of -0.479 MeV<sup>353</sup> 325 on the proton single particle energy, we placed the un-<sup>354</sup> 326 observed state at  $E_{ex} = 4.087$  MeV in <sup>25</sup>Si [28, 29]. Ac-<sup>355</sup> 327 counting for the energy of the isomeric state of  ${}^{24}\text{Al}$  (E<sub>ex</sub><sup>356</sup> 328 = 0.426 MeV), we find a resonance energy with respect<sup>357</sup> 329 to the isomer of  $E_r^m = 0.247$  MeV. The mirror  $5/2^+$  state<sup>358</sup> 330 would be at  $E_{ex} = 4.429$  MeV in <sup>25</sup>Na. Due to the low<sup>359</sup> 331 statistics of the present experiment, and that this state<sup>360</sup> 332 would be populated by an  $\ell = 2$  transfer from the  $1^{+}_{361}$ 333 isomer, we are unable to identify this state which would<sup>362</sup> 334 be at the background level in our data. The shell model<sup>363</sup> 335 spectroscopic factor for  $\ell = 2$  transfer to this state was<sup>364</sup> 336 found to be  $C^2S = 0.115$  [25]. The resonance strength<sup>365</sup> 337 of such resonance was calculated according to Ref. [32],<sup>366</sup> 338 and was found to be  $\omega \gamma = 0.735$  meV. The effect this<sup>367</sup> 339 resonance as well as the uncertainty in the resonance en-368 340 ergy are indicated by the black lines in Fig. 8a, where<sub>369</sub> 341 the ratio of the rate calculated with the shell model pre-370 342 dicted  $5/2^+$  resonance state at  $E_r^m = 0.247$  MeV in <sup>25</sup>Si<sub>371</sub> 343 to that of the current ground state REACLIB rate as a<sub>372</sub> 344 function of temperature, is plotted. The energy of the<sub>373</sub> 345

predicted resonance has been varied by  $\pm 100$  keV to determine possible effects of a shift in its location and it is shown by the shaded region (the change in resonance energy also changes the calculated resonance strength). This variation in the energy was chosen based on the resonance energy uncertainty given in the network calculations of Ref. [33] for the same state when populated by the ground state.

The ratio of the total isomeric contributions of the  ${}^{24}Al(p,\gamma){}^{25}Si$  reaction rate determined in this work to the ground state contributions is shown in Fig. 8b. The upper and lower limits here are due to the uncertainty in the energy of the  $5/2^+$  shell model predicted resonance.

The present calculated rate for the <sup>24</sup>Al<sup>m</sup>(p, $\gamma$ )<sup>25</sup>Si reaction is shown in Fig. 9. The red lines show the rate for both the experimentally measured sub-threshold 1/2<sup>+</sup> state at  $E_r^m = -0.038$  MeV in <sup>25</sup>Si (red long dashed line), and the shell model predicted 5/2<sup>+</sup> state at  $E_r^m = 0.247$ MeV in <sup>25</sup>Si (red short dashed line), as well as their total combined contribution to the <sup>24</sup>Al<sup>m</sup>(p, $\gamma$ )<sup>25</sup>Si reaction rate (red solid line). The current REACLIB ground state rate [14, 35] with the recent work of Ref. [6] is also included for comparison (blue dotted line).

In the temperature range of interest to the rp-process  $0.1 \leq T_9 \leq 10$  [1, 2, 5], the contribution of the subthreshold state in <sup>25</sup>Si to the rate of the <sup>24</sup>Al(p, $\gamma$ )<sup>25</sup>Si rate is negligible. For the shell model predicted resonance state placed at  $E_r^m = 0.247$  MeV in <sup>25</sup>Si, it is observed

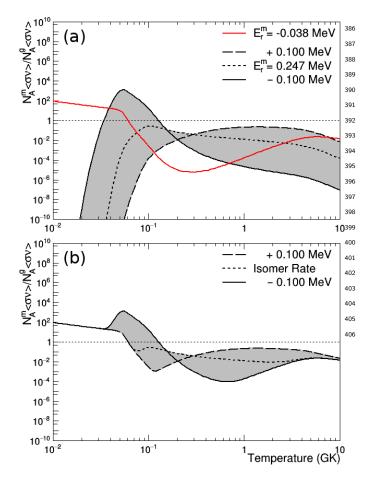


FIG. 8. Ratio of the isomeric rates extracted in this work to the current ground state REACLIB rate [34]. (a) Individual isomeric rate contributions. The  $1/2^+$  sub-threshold resonance at  $E_r^m = -0.038$  MeV (red line) and the  $5/2^+$  state at  $E_r^m = 0.247$  MeV (black lines) as function of the temperature. The resonance energy of the  $5/2^+$  resonance has been varied by  $\pm 100$  keV (shaded area). (b) Total isomeric contribution  $(1/2^+ + 5/2^+)$  determined in this work. The shaded area indicate the uncertainty in the energy of the  $5/2^+$  resonance.

that its influence to the total reaction rate depends on 374 the exact location of its resonance energy as shown in Fig. 375 8b. Since variations in the energy of this state will cause 376 the overall contribution to change, further measurements 377 are needed to confirm the existence and location of this 378 state as well as the strength of the resonance to fully 379 determine its influence to the rate of the  ${}^{24}\text{Al}(p,\gamma){}^{25}\text{Si}_{407}$ 380 reaction when populated by the isomeric state. 381

## V. SUMMARY

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In summary, a radioactive beam of  $^{24}$ Na with  $90\%_{^{412}}$ of its content in the isomeric  $1^+$  state was developed,  $_{^{413}}$ characterized, and used to perform, for the first time,  $a_{^{414}}$  measurement of the <sup>24</sup>Na<sup>*m*</sup>(d,p)<sup>25</sup>Na reaction at Florida State University's John D. Fox Accelerator Laboratory. States in <sup>25</sup>Na up to  $E_{ex} = 5$  MeV in excitation energy, populated by  $\ell = 0$  transfers from the isomeric state in <sup>24</sup>Na, were selectively observed in this experiment. Spectroscopic information extracted from this experiment was compared with USDB shell model calculations and showed good agreement between experiment and theory.

Mirror symmetry arguments between <sup>25</sup>Na and <sup>25</sup>Si were used to provide spectroscopic information of states above the proton threshold in <sup>25</sup>Si and, for the first time, constrain the contribution of the isomeric 1<sup>+</sup> state in <sup>24</sup>Al to the rate of the <sup>24</sup>Al(p, $\gamma$ )<sup>25</sup>Si reaction. The contribution of an  $\ell = 0$  sub-threshold resonance was determined to be negligible. The presence of an additional  $\ell = 2$  resonance, predicted by the shell model but not observed in the present experiment, could have a role in the <sup>24</sup>Al(p, $\gamma$ )<sup>25</sup>Si reaction rate. Experimental information on the exact location of this state in <sup>25</sup>Si is needed to evaluate its impact to the <sup>24</sup>Al(p, $\gamma$ )<sup>25</sup>Si reaction rate.

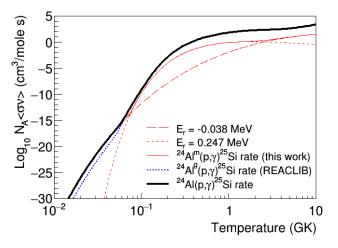


FIG. 9. The rate of the  ${}^{24}\text{Al}(p,\gamma){}^{25}\text{Si}$  reaction. The red lines show the contributions from the 1<sup>+</sup> isomeric state in  ${}^{24}\text{Al}$ presented in this work, while the blue line shows the current recommended REACLIB rate [34] with the state from the work of Ref. [6] added in. The black line denotes the total rate of the  ${}^{24}\text{Al}(p,\gamma){}^{25}\text{Si}$ .

## ACKNOWLEDGEMENTS

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- [1] R. Wallace and S. E. Woosley, Explosive hydrogen burn-476
   ing, The Astrophysical Journal Supplement Series 45,477
   389 (1981).
- [2] J. L. Fisker, H. Schatz, and F.-K. Thielemann, Explosive<sup>479</sup>
   hydrogen burning during type i x-ray bursts, The Astro-<sup>480</sup>
   physical Journal Supplement Series **174**, 261 (2008). <sup>481</sup>
- 421 [3] H. Schatz and K. Rehm, X-ray binaries, Nuclear Physics482
   422 A 777, 601 (2006).
   483
- [4] Z. Meisel, A. Deibel, L. Keek, P. Shternin, and J. Elfritz,484
   Nuclear physics of the outer layers of accreting neutron485
   stars, Journal of Physics G: Nuclear and Particle Physics486
   45, 093001 (2018).
- L. Van Wormer, J. Görres, C. Iliadis, M. Wiescher, and
   F.-K. Thielemann, Reaction rates and reaction sequences
   in the rp-process, The Astrophysical Journal 432, 326490 (1994).
- 431 [6] B. Longfellow, A. Gade, B. A. Brown, W. A. Richter, 492 432 D. Bazin, P. C. Bender, M. Bowry, B. Elman, E. Lun-493 433 derberg, D. Weisshaar, and S. J. Williams, Measurement 494 434 of key resonances for the  ${}^{24}\text{Al}(p, \gamma){}^{25}\text{Si}$  reaction rate us-495 435 ing in-beam  $\gamma$ -ray spectroscopy, Phys. Rev. C **97**, 054307496 436 (2018). 497
- 437 [7] C. Iliadis, A. Champagne, J. Jose, S. Starrfield, and 498
  438 P. Tupper, The effects of thermonuclear reaction-rate 499
  439 variations on nova nucleosynthesis: A sensitivity study, 500
  440 The Astrophysical Journal Supplement Series 142, 105501
  441 (2002). 502
- [8] D. D. Clayton and L. R. Nittler, Astrophysics with preso-503
   lar stardust, Annual Review of Astronomy and Astro-504
   physics 42, 39 (2004). 505
- [9] G. W. Misch, S. K. Ghorui, P. Banerjee, Y. Sun, and 506
   M. R. Mumpower, Astromers: Nuclear isomers in as-507
   trophysics, The Astrophysical Journal Supplement Series508
   252, 2 (2020). 509
- 449[10] C. M. Deibel, J. A. Clark, R. Lewis, A. Parikh, P. D.510450Parker, and C. Wrede, Toward an experimentally deter-511451mined  ${}^{26}\text{Al}^m(p,\gamma){}^{27}\text{Si}$  reaction rate in one novae, Phys.512452Rev. C 80, 035806 (2009).513
- [11]S. Almaraz-Calderon, K. E. Rehm, N. Gerken, M. L.514 453 Avila, B. P. Kay, R. Talwar, A. D. Avangeakaa, S. Bot-515 454 toni, A. A. Chen, C. M. Deibel, C. Dickerson, K. Hansel-516 455 man, C. R. Hoffman, C. L. Jiang, S. A. Kuvin, O. Nu-517 456 sair, R. C. Pardo, D. Santiago-Gonzalez, J. Sethi, and<sup>518</sup> 457 C. Ugalde, Study of the  ${}^{26}\text{Al}^m(d,p){}^{27}\text{Al}$  reaction and thesis 458 influence of the <sup>26</sup>Al 0<sup>+</sup> isomer on the destruction of <sup>26</sup>Al<sub>520</sub> 459 in the galaxy, Phys. Rev. Lett. 119, 072701 (2017). 460 521
- 461 [12] G. W. Misch, Private Communication.
- W. Benenson, J. Driesbach, I. D. Proctor, G. F. Trentel-523
   man, and B. M. Preedom, Energy levels of <sup>25</sup>Si from thes24
   reaction <sup>28</sup>Si(<sup>3</sup>He, <sup>6</sup>He)<sup>25</sup>Si at 70.4 mev, Phys. Rev. C 5,525
   1426 (1972). 526
- [14] H. Herndl, J. Görres, M. Wiescher, B. A. Brown, and 527
  L. Van Wormer, Proton capture reaction rates in the rp528
  process, Phys. Rev. C 52, 1078 (1995). 529
- [15] A. J. Knapton, Structure of  ${}^{25}Na$  measured using  $d_{70}$  d( ${}^{24}Na, p$ )  ${}^{25}Na$  with a radioactive  ${}^{24}Na$  beam, Ph.D. the-sai sis, University of Surrey (2017).
- 472 [16] I. Wiedenhöver, L. Baby, D. Santiago-Gonzalez, A. Ro-533
  473 jas, J. Blackmon, G. Rogachev, J. Belarge, E. Koshchiy,534
  474 A. Kuchera, L. Linhardt, *et al.*, Studies of exotic nucleis35
- 475 at the resolut facility of florida state university, in *Fission*

and Properties of Neutron-Rich Nuclei (World Scientific, 2014) pp. 144–151.

- [17] B. Harss, R. C. Pardo, K. E. Rehm, F. Borasi, J. P. Greene, R. V. F. Janssens, C. L. Jiang, J. Nolen, M. Paul, J. P. Schiffer, R. E. Segel, J. Specht, T. F. Wang, P. Wilt, and B. Zabransky, Production of radioactive ion beams using the in-flight technique, Review of Scientific Instruments **71**, 380 (2000), https://doi.org/10.1063/1.1150211.
- [18] National nuclear data center, https://www.nndc.bnl. gov/.
- [19] Fresco coupled reaction channels calculations, http:// www.fresco.org.uk/.
- [20] C. Daum, The  $^{\overline{23}}$ Na $(d, p)^{24}$ Na reaction and the nuclear structure of  $^{24}$ Na, Nuclear Physics **45**, 273 (1963).
- [21] B. Asher, S. Almaraz-Calderon, O. Nusair, K. Rehm, M. Avila, A. Chen, C. Dickerson, C. Jiang, B. Kay, R. Pardo, D. Santiago-Gonzalez, and R. Talwar, Development of an isomeric beam of <sup>26</sup>Al for nuclear reaction studies, Nuclear Instruments and Methods in Physics Research Section A: Accelerators, Spectrometers, Detectors and Associated Equipment **899**, 6 (2018).
- [22] C. Daum, The nuclear structure of <sup>24</sup>Na and some remarks on the nuclear structure of <sup>25</sup>Mg, Nuclear Physics 51, 244 (1964).
- [23] Y. Han, Y. Shi, and Q. Shen, Deuteron global optical model potential for energies up to 200 mev, Phys. Rev. C 74, 044615 (2006).
- [24] A. Koning and J. Delaroche, Local and global nucleon optical models from 1 kev to 200 mev, Nuclear Physics A 713, 231 (2003).
- [25] A. Volya, Private Communication.

522

- [26] W. A. Richter, S. Mkhize, and B. A. Brown, sd-shell observables for the usda and usdb hamiltonians, Phys. Rev. C 78, 064302 (2008).
- [27] J. M. VonMoss, S. L. Tabor, V. Tripathi, A. Volya, B. Abromeit, P. C. Bender, D. D. Caussyn, R. Dungan, K. Kravvaris, M. P. Kuchera, R. Lubna, S. Miller, J. J. Parker, and P.-L. Tai, Higher-spin structures in <sup>21</sup>F and <sup>25</sup>Na, Phys. Rev. C **92**, 034301 (2015).
- [28] R. G. Thomas, On the determination of reduced widths from the one-level dispersion formula, Phys. Rev. 81, 148 (1951).
- [29] J. B. Ehrman, On the displacement of corresponding energy levels of C<sup>13</sup> and N<sup>13</sup>, Phys. Rev. 81, 412 (1951).
- [30] C. E. Rolfs and W. S. Rodney, *Cauldrons in the cosmos: Nuclear astrophysics* (University of Chicago press, 1988).
- [31] C. Angulo, M. Arnould, M. Rayet, P. Descouvemont, D. Baye, C. Leclercq-Willain, A. Coc, S. Barhoumi, P. Aguer, C. Rolfs, R. Kunz, J. Hammer, A. Mayer, T. Paradellis, S. Kossionides, C. Chronidou, K. Spyrou, S. Degl'Innocenti, G. Fiorentini, B. Ricci, S. Zavatarelli, C. Providencia, H. Wolters, J. Soares, C. Grama, J. Rahighi, A. Shotter, and M. Lamehi Rachti, A compilation of charged-particle induced thermonuclear reaction rates, Nuclear Physics A 656, 3 (1999).
- [32] R. Longland, C. Iliadis, A. Champagne, J. Newton, C. Ugalde, A. Coc, and R. Fitzgerald, Charged-particle thermonuclear reaction rates: I. monte carlo method and statistical distributions, Nuclear Physics A 841, 1 (2010).

- [33] C. Iliadis, R. Longland, A. Champagne, and A. Coc,542
   Charged-particle thermonuclear reaction rates: Iii. nu-543
- <sup>538</sup> clear physics input, Nuclear Physics A 841, 251 (2010). <sup>544</sup>
   <sup>539</sup> [34] R. H. Cyburt, A. M. Amthor, R. Ferguson, Z. Meisel, <sup>545</sup>
- 540 K. Smith, S. Warren, A. Heger, R. D. Hoffman, 546
- 541 T. Rauscher, A. Sakharuk, H. Schatz, F. K. Thielemann, 547 548

and M. Wiescher, The JINA REACLIB database: Its recent updates and impact on type-i x-ray bursts, The Astrophysical Journal Supplement Series **189**, 240 (2010).

[35] C. Iliadis, R. Longland, A. Champagne, A. Coc, and R. Fitzgerald, Charged-particle thermonuclear reaction rates: Ii. tables and graphs of reaction rates and probability density functions, Nuclear Physics A 841, 31 (2010).