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4 **Contributions of gravity waves in the ocean to  $T$ -phase  
5 excitation by earthquakes<sup>a)</sup>**

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14 Running title: Abyssal  $T$ -waves

15 Keywords: range-dependent waveguides; wave propagation theory; sound scattering; wind  
16 waves; sea swell; internal gravity waves

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18 Submitted to the *Journal of the Acoustical Society of America* on July 8, 2021  
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<sup>a)</sup> Parts of this work have been previously reported at the 178th Meeting of the Acoustical Society of America (San Diego, December 2019).

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22 **Abstract**

23 Generation of *T*-waves in a deep ocean by an earthquake in its epicentral region is often observed  
24 but the mechanism of the excitation of the acoustic waves travelling horizontally with the speed  
25 of sound remains controversial. Here, the hypothesis is investigated that the abyssal *T*-waves are  
26 generated by scattering of ballistic sound waves by surface and internal gravity waves in the  
27 ocean. Volume and surface scattering are studied theoretically in the small perturbation  
28 approximation. In the 3–50 Hz typical frequency range of the observed *T*-waves, linear internal  
29 waves are found to lack the necessary horizontal spatial scales to meet the Bragg scattering  
30 condition and contribute appreciably to *T*-wave excitation. In contrast, the ocean surface  
31 roughness has the necessary spatial scales at typical sea states and wind speeds. Efficiency of the  
32 acoustic normal modes' excitation at surface scattering of the ballistic body waves by wind seas  
33 and sea swell is quantified and found to be comparable to that of the established mechanism of  
34 *T*-wave generation at downslope conversion at seamounts. The surface scattering mechanism is  
35 consistent with key observational features of abyssal *T*-waves, including their ubiquity, low-  
36 frequency cutoff, presence on seafloor sensors, and weak dependence on the earthquake focus  
37 depth.

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40 PACS numbers: 43.30.Ma, 43.30.Dr, 43.30.Hw, 43.30.Ft, 43.30.Nb

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42 **I. INTRODUCTION**

43        The *T*-, or tertiary, phase of an underwater earthquake is composed of low-frequency  
 44    acoustics waves, which propagate to long ranges in underwater waveguide at speeds close to the  
 45    sound speed in water and arrive later than *P*-, or primary, and *S*-, or secondary phases, which are  
 46    due to compressional (*P*) and shear (*S*) body waves in the seabed, and later than seismo-acoustic  
 47    interface waves.<sup>1–4</sup> *T* waves weakly attenuate with range, travel over very large distances, and are  
 48    observed throughout the world ocean. They are the most common earthquake sounds in the  
 49    ocean and make strong but transient contributions to the ambient sound field.<sup>5, 6</sup> A  
 50    comprehensive review of *T*-wave research up to mid-2000s can be found in Refs. 2, 3, 7, and 8.

51        In addition to hydrophones at various depths in the water column,<sup>9–13</sup> *T* waves are  
 52    routinely observed by receivers on the seafloor in deep water,<sup>14–16</sup> which indicates, in agreement  
 53    with full-wave numerical modeling,<sup>8, 12, 17–19</sup> that *T*-waves are not confined in the SOFAR  
 54    channel. Because the wave speed and absorption in water are, respectively, smaller and much  
 55    smaller than in the earth crust, *T* waves prove to be the most sensitive and rather accurate means  
 56    to detect, characterize, and localize marine teleseismic events, including weak intraplate events.<sup>9,</sup>  
 57   <sup>20–23</sup> In addition, *T* waves carry information about the ocean. It was proposed to use  
 58    measurements of temporal variability of *T*-wave travel times to characterize internal tides and  
 59    associated ocean mixing<sup>24</sup> and, more recently, for ocean acoustic thermometry.<sup>25, 26</sup>

60        Numerous observations show that conversion of seismic energy into guided acoustic  
 61    waves in oceanic waveguide occurs in the vicinity of the earthquake epicenter and at prominent  
 62    bathymetric features, which may be located hundreds of kilometers away from the epicenter.<sup>3, 9,</sup>  
 63   <sup>13, 20, 27–31</sup> *T*-wave amplitudes remain significant for intermediate-depth earthquakes<sup>9, 32</sup> and are  
 64    insensitive to water depth.<sup>2</sup> *T* waves from deep-focus earthquakes, with hypocenter depths of

65 hundreds of km, have been also observed.<sup>3, 14</sup> The conversion mechanism and especially *T*-wave  
66 excitation in the immediate vicinity of the epicenter are not well understood.<sup>2, 8, 22</sup> Excitation of  
67 acoustic normal modes at large-scale bathymetric features can be explained in terms of the  
68 downslope conversion and diffraction of *P* and *S* body waves and/or seismo-acoustic interface  
69 waves by horizontally inhomogeneous bathymetry.<sup>8, 19, 33–36</sup> Ubiquitous “abyssal” *T* waves<sup>9, 32, 33,</sup>  
70<sup>37</sup> that are generated near the epicenter of earthquakes under flat abyssal planes, cannot be  
71 attributed to any of these generation mechanisms. Unlike the trapping of acoustic energy in the  
72 SOFAR channel by downslope conversion of steeply propagating sound, generation of abyssal *T*  
73 waves does not lend itself to a ray interpretation. It had been realized early on<sup>9, 32, 37</sup> that a wave  
74 scattering mechanism was required to explain abyssal *T*-wave observations. Johnson, Norris, and  
75 Duennebier discussed scattering at the ocean surface and seafloor and volume scattering of  
76 sound in the ocean among the conceivable generation mechanisms and favored scattering by the  
77 ocean surface.<sup>9, 32, 37</sup> However, their crude estimates of the generation efficiency were not  
78 encouraging. Keenan and Merriam<sup>38</sup> proposed sound scattering from keels on the undersurface  
79 of the ice cover as the mechanism of generation of abyssal *T* waves in the Arctic. The idea that  
80 sound scattering at the ocean surface could be an important mechanism of *T*-phase generation  
81 has been recently re-visited by Bottero<sup>8</sup> using full-wave, two-dimensional (2-D) numerical  
82 modeling in a scenario with strong, discrete scatterers located on the ocean surface.

83 Following Fox et al.<sup>20</sup> and De Groot-Hedlin and Orcutt,<sup>39, 40</sup> it is often implied in the  
84 current literature<sup>3, 6, 22</sup> that abyssal *T* waves are generated due to wave scattering by seafloor  
85 roughness, specifically due to coupling between the seismo-acoustic normal modes that are  
86 directly excited by the seismic source, and the normal modes comprising the *T*-phase.<sup>41, 42</sup> By  
87 modeling scattered waves as the field due to uncorrelated virtual sound sources distributed along

88 the seafloor, De Groot-Hedlin and Orcutt<sup>39, 40</sup> and Yang and Forsyth<sup>22</sup> successfully reproduced  
89 the shapes of envelopes of observed *T*-phase waveforms. However, detailed information about  
90 the seafloor roughness spectra is rarely if ever available around the epicenter of abyssal  
91 earthquakes with the granularity and at the spatial scales necessary for *T*-phase modeling. To our  
92 knowledge, the amplitude of the resulting *T* waves has never been related to actual seafloor  
93 roughness data or models in a quantitative manner and shown to be sufficient to explain the  
94 observed abyssal *T* waves.

95 Here, we examine an alternative hypothesis that sound waves coming at steep angles  
96 directly from the earthquake focus (ballistic body waves) are coupled to normal modes of the  
97 underwater acoustic waveguide by dynamic processes in the water column and on the ocean  
98 surface. Specifically, we investigate the generation of abyssal *T* waves at scattering of ballistic  
99 sound waves by the ocean surface roughness, which is due to surface gravity waves, and by  
100 volume inhomogeneities of the water column, which are caused by internal gravity waves. We  
101 view the ocean surface and volume scattering as either a complementary to the seafloor  
102 scattering or possibly an alternative mechanism of generation of abyssal *T* waves. Unlike the  
103 seafloor roughness data in the open ocean, extensive information on statistics of wind waves and  
104 sea swell<sup>43–45</sup> and internal gravity wave spectra<sup>46, 47</sup> is available, which allows one to reach  
105 definitive conclusions regarding significance of these generation mechanisms.

106 *T* waves are a seismo-acoustic phenomenon with representative wave frequencies being  
107 very high on the seismic scale and low for underwater sound. Typically, *T* waves are observed in  
108 the 1–100 Hz band.<sup>2, 3</sup> Lower frequencies dominate the signals from stronger and deeper  
109 earthquakes, while the highest frequencies are generated by the weakest detected seismic events.  
110 Abyssal *T* waves exhibit higher frequencies than the *T* waves generated at down-slope

111 conversion.<sup>3,32</sup> Therefore, this paper will focus on the 3–50 Hz frequency band that contains  
112 most of the abyssal *T*-wave energy. Observations indicate existence of a low-frequency cutoff in  
113 *T*-phase spectra, see, e.g., Refs. 13, 32, 48 and Ref. 8, p. 59. The low-frequency cutoff will be  
114 related to the *T*-phase generation process in this paper.

115 Mathematically, we describe the excitation of abyssal *T*-waves as scattering from the  
116 continuous spectrum into the discrete spectrum of the seismo-acoustic field. The continuous  
117 spectrum is represented here by the body waves, that are generated by an earthquake and reach  
118 the water column with a modest transmission loss at typical *T*-phase frequencies below about  
119 40–50 Hz. This process is reciprocal of scattering of the normal modes propagating in the  
120 oceanic waveguide by the rough ocean surface and/or volume inhomogeneities due to internal  
121 gravity waves (scattering from the discrete into the continuous spectrum of the acoustic field). In  
122 that problem, a part of the scattered energy is radiated into the seabed and carried away from the  
123 waveguide, leading to the well-known contribution to attenuation of the normal modes.<sup>49–52</sup>

124 The remainder of the paper is organized as follows. A theory of excitation of acoustic  
125 normal modes at scattering of a low-frequency body wave by rough ocean surface and random  
126 volume inhomogeneities is developed in Sec. II for underwater waveguides with either fluid or  
127 solid bottom. Efficiency of *T*-phase excitation by ballistic body waves is related to the spectral  
128 properties of the roughness and volume inhomogeneities. The theory is applied in Sec. III to  
129 surface scattering by wind seas with the Pierson-Moskovitz spectrum and wavetrains of sea swell  
130 to characterize the frequency spectra, directionality, and energy of the resulting *T* waves and the  
131 dependence of the *T*-phase properties on the earthquake focus depth. Simple, order-of-magnitude  
132 estimates of the *T*-phase energy are obtained in Sec. IV A and employed to argue, that surface  
133 scattering of ballistic body waves in the vicinity of the earthquake epicenter is a significant *T*-

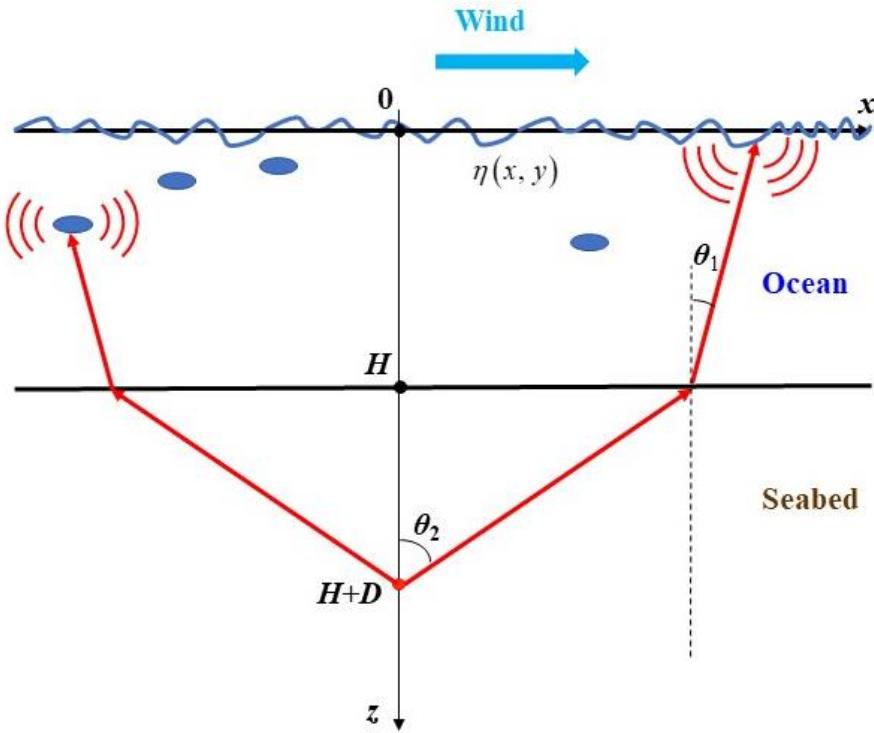
134 phase generation mechanism with a strength comparable to that of a seamount at a moderate  
135 distance from the epicenter. Section IV B discusses possible extensions of the theory to quantify  
136 other plausible mechanisms of generation of  $T$  waves and related waves in the atmosphere.  
137 Section V summarizes our findings.

138

## 139 **II. T-PHASE GENERATION BY SURFACE AND VOLUME SCATTERING**

### 140 **A. Scattering of low-frequency sound by the rough ocean surface**

141 Consider a horizontally stratified ocean of depth  $H$ . Introduce Cartesian coordinates  $x, y,$   
142  $z$  with the vertical coordinate  $z$  increasing downward. The mean position of the ocean surface is  
143 the horizontal plane  $z = 0$ ; the seafloor is located at  $z = H$  (Fig. 1). The epicenter of an  
144 earthquake, which generates  $T$  waves, is located in the vicinity of the origin  $x = 0, y = 0$  of the  
145 horizontal coordinates. In addition to the Cartesian coordinates, we will also use a cylindrical  
146 coordinate system  $\{r, \varphi, z\}$  with the same  $z$  axis. When averaged over perturbations due to  
147 internal gravity waves, sound speed  $c$  in the ocean and water density  $\rho$ , as well as the density and  
148 compressional and shear wave speeds in the seabed, are functions of  $z$ . We disregard the seafloor  
149 roughness and the effects of horizontal inhomogeneities of the water column and seabed when  
150 considering wave scattering by the ocean surface roughness.



151

152 **Figure 1. (Color online)** Geometry of the problem. Ballistic waves from the earthquake focus  
 153 scatter at the rough ocean surface and volume inhomogeneities in the water column, which act as  
 154 secondary sound sources and generate guided waves in the oceanic waveguide. The volume  
 155 inhomogeneities are symbolically represented by ovals in the figure. The ocean surface  
 156 roughness is described by the surface elevation  $\eta$ , which varies with the horizontal coordinates  $x$   
 157 and  $y$ . The earthquake focus is located at  $x = y = 0$  at the depth  $z = H + D$  under the seafloor  $z =$   
 158  $H$ .

159

160 Wave heights on the ocean surface are small compared to acoustic wavelengths at  $T$ -  
 161 phase frequencies (longer than 30 m for frequencies below 50 Hz). With a possible exception for  
 162 some breaking waves, slopes of the ocean surface are small compared to unity. Sound scattering  
 163 by such surfaces can be described by the small perturbation method.<sup>53, 54</sup> Consider scattering of

164 monochromatic acoustic waves of frequency  $\omega$  by a stationary (frozen) rough surface. We will  
 165 use complex notation for monochromatic wave fields, where the time dependence  $\exp(-i\omega t)$  of  
 166 the acoustic pressure and other quantities is assumed and suppressed. In the first approximation  
 167 of the small perturbation method, acoustic pressure  $p_{sc}$  in the wave scattered by a rough pressure-  
 168 release surface is

$$169 \quad p_{sc}(\mathbf{R}) = - \int \left[ \frac{\partial p_0}{\partial z_1} \frac{\partial G(\mathbf{R}; \mathbf{r}_1, z_1)}{\partial z_1} \right]_{z_1=0} \eta(\mathbf{r}_1) \frac{d\mathbf{r}_1}{\rho}. \quad (1)$$

170 Here integration is over the mean surface  $z = 0$ ;  $\mathbf{r}_1$  is a two-dimensional horizontal vector,  $\mathbf{R}$  is a  
 171 three-dimensional position vector;  $p_0$  is the acoustic pressure in the monochromatic wave in the  
 172 absence of surface roughness, i.e., in the “unperturbed” waveguide with the pressure-release  
 173 boundary  $z = 0$ . Acoustic pressure in the full acoustic field in water equals  $p_{sc} + p_0$ ;  $p_0$  contains

174 the incident wave and the wave reflected from the flat (horizontal) ocean surface. Surface  
 175 elevation  $\eta(\mathbf{r})$  is the vertical deviation of the rough surface from the mean plane  $z = 0$ .

176 Mathematically, the rough surface is given by the equation  $z = \eta(\mathbf{r})$ . Note that  $p_{sc} \rightarrow 0$  in the limit  
 177  $\eta \rightarrow 0$  of vanishing roughness. In Eq. (1),  $G(\mathbf{R}; \mathbf{R}_1)$  is the acoustic Green’s function in the ocean  
 178 with the flat upper boundary  $z = 0$ . The Green’s function has the meaning of the acoustic  
 179 pressure at point  $\mathbf{R}$  due to a point sound source of volume velocity located at  $\mathbf{R}_1$ . In the water  
 180 column, the Green’s function satisfies the equation<sup>55</sup>

$$181 \quad \frac{\partial}{\partial \mathbf{R}} \left[ \frac{1}{\rho} \frac{\partial}{\partial \mathbf{R}} G(\mathbf{R}; \mathbf{R}_1) \right] + \frac{\omega^2}{\rho c^2} G(\mathbf{R}; \mathbf{R}_1) = -\delta(\mathbf{R} - \mathbf{R}_1) \quad (2)$$

182 as well as the appropriate boundary conditions on the ocean surface and the seafloor. Here  $\delta(\mathbf{R})$   
 183 is the Dirac delta function. The approximate solution Eq. (1) for the scattered wave describes  
 184 single scattering from the rough surface but accounts for all multiple reflections in the ocean  
 185 with the horizontal upper boundary.<sup>53-55</sup>

186 The physical meaning of Eq. (1) is that, in the first approximation of the small  
 187 perturbation method, the waves scattered from the rough ocean surface are described as the  
 188 waves generated by a known, distributed sound source in the ocean with the flat upper boundary.  
 189 Indeed, acoustic pressure in the field generated by monochromatic sound sources in an  
 190 inhomogeneous fluid satisfies the reduced wave equation<sup>55</sup>

$$191 \quad \nabla \cdot \left( \frac{\nabla p}{\rho} \right) + \frac{\omega^2}{\rho c^2} p = i\omega A + \nabla \cdot \left( \frac{\mathbf{F}}{\rho} \right), \quad (3)$$

192 where  $\mathbf{F}$  and  $A$  stand for the volume densities of the external force and volume velocity (i.e., the  
 193 volume injection rate), respectively. In terms of the acoustic Green's function  $G$  of the medium,  
 194 solution of the reduced wave equation is given by the equation<sup>55</sup>

$$195 \quad p(\mathbf{R}) = \int \left[ \frac{\mathbf{F}(\mathbf{R}_1)}{\rho(\mathbf{R}_1)} \cdot \frac{\partial G(\mathbf{R}; \mathbf{R}_1)}{\partial \mathbf{R}_1} - i\omega A(\mathbf{R}_1) G(\mathbf{R}; \mathbf{R}_1) \right] d\mathbf{R}_1, \quad (4)$$

196 where the integration is over the entire volume occupied by the sources. Comparison of Eq. (1)  
 197 and (4) shows that, in the first approximation of the small perturbation method, the scattered  
 198 wave coincides with the field that would be generated in the medium with horizontal upper  
 199 boundary by external forces with density

$$200 \quad \mathbf{F}(\mathbf{r}_1, z_1) = \left( 0, \quad 0, \quad -\eta(\mathbf{r}_1) \frac{\partial p_{in}}{\partial z_1} \delta(z_1) \right). \quad (5)$$

201 Equation (5) describes an effective vertical external force applied on the horizontal ocean  
 202 surface. The effective sound source depends on the incident wave and the roughness of the actual  
 203 ocean surface.

204 One can also reach the same conclusion that the scattered wave in an inhomogeneous  
 205 medium is equivalent to the sound field generated by the effective sound source Eq. (5) on the  
 206 horizontal boundary by comparing the boundary condition<sup>53, 54</sup>

207  $p_{sc}(\mathbf{r}_1, z = +0) = -\eta(\mathbf{r}_1)(\partial p_{in}/\partial z_1)_{z_1=0}$  for the scattered wave in the first approximation of the  
 208 small perturbation method with the discontinuity (jump)<sup>55</sup>  $p(\mathbf{r}_1, z = +0) - p(\mathbf{r}_1, z = 0)$   
 209  $= p(\mathbf{r}_1, z = +0) = F_{0z}(\mathbf{r}_1)$  of the acoustic pressure, which, according to Eq. (3), is caused by the  
 210 distribution of external vertical forces with volume density  $F_{0z}\delta(z)$  just below a pressure-  
 211 release boundary  $z = 0$ . Here  $z = +0$  denotes points situated below the boundary  $z = 0$   
 212 infinitesimally close to it.

213

214 **B. Excitation of normal modes at surface scattering**

215 In a horizontally stratified oceanic waveguide with a fluid seabed, the acoustic Green's  
 216 function is given by the sum of normal modes<sup>55, 56</sup>

$$217 \quad G(\mathbf{R}; \mathbf{R}_1) = \frac{i}{4} \sum_n f_n(z) f_n(z_1) H_0^{(1)}(\xi_n |\mathbf{r} - \mathbf{r}_1|) \\ = \sum_n f_n(z) f_n(z_1) \frac{\exp(i\xi_n |\mathbf{r} - \mathbf{r}_1| + i\pi/4)}{\sqrt{8\pi\xi_n |\mathbf{r} - \mathbf{r}_1|}} \left[ 1 + O\left(\frac{1}{\xi_n |\mathbf{r} - \mathbf{r}_1|}\right) \right] \quad (6)$$

218 plus a contribution of the continuous spectrum. The latter is usually negligible at long-range  
 219 propagation. Here  $H_0^{(1)}(\cdot)$  is a Hankel function of the first kind of order zero,  $\xi_n$  and  $f_n(z)$  are the  
 220 propagation constant and shape function of the  $n$ th normal mode,  $n = 1, 2, \dots$ . The shape  
 221 functions are normalized by the condition

$$222 \quad \int_0^\infty \frac{dz}{\rho(z)} f_n^2(z) = 1. \quad (7)$$

223 The shape function  $f_n(z)$  gives the vertical dependence of acoustic pressure in the  $n$ th normal  
 224 mode. When the horizontal separation of the points  $\mathbf{R} = (\mathbf{r}, z)$  and  $\mathbf{R}_1 = (\mathbf{r}_1, z_1)$  is large compared  
 225 to the wavelength, the Hankel function can be replaced by the dominant term of its asymptotic

226 expansion<sup>57</sup> leading to the right-most side in Eq. (6). With the points  $\mathbf{R}$  and  $\mathbf{R}_1$  located in water,  
 227 Eq. (6) remains valid in the waveguide with stratified solid seabed<sup>58</sup> but, instead of Eq. (7), the  
 228 normalization condition of the normal mode shape functions in the fluid-solid waveguide takes  
 229 the form

$$230 \quad \int_0^H \rho^{-1} f_n^2 dz + \frac{\omega}{\xi_n} \int_H^{+\infty} (\tau_{xz} v_z - \tau_{xx} v_x) \rho dz = 1, \quad (8)$$

231 where  $H$  is water depths,  $\tau_{xx}$  and  $\tau_{xz}$  are components of the stress tensor and  $v_x$  and  $v_z$  are  
 232 components of the particle velocity  $\mathbf{v} = (v_x, 0, v_z)$  in the seabed in the  $n$ th normal mode with the  
 233 dependence  $\exp(i\xi_n z)$  of its field on horizontal coordinates.<sup>58</sup> The shape functions  $f_n(z)$  are real-  
 234 valued in the absence of dissipation. The physical meaning of the normalization Eq. (8) is that  
 235 modes with the same amplitude carry the same power flux; the acoustic power flux  $J_n$  in a single  
 236 propagating normal mode with  $p(\mathbf{r}, z) = a f_n(z) H_0^{(1)}(\xi_n r)$ , where  $a$  is a constant, equals

$$237 \quad J_n = 2 |a^2| / \omega. \quad ^{55, 56, 58}$$

238 Substitution of the Green's function Eq. (6) into Eq. (1) for the scattered wave and  
 239 changing the order of the summation and integration gives

$$240 \quad p_{sc}(\mathbf{r}, z) = \sum_n \frac{f_n(z) \exp(-3i\pi/4)}{\sqrt{8\pi\xi_n} \rho(0)} \left. \frac{\partial f_n}{\partial z} \right|_{z=0} Q_n(\mathbf{r}), \quad (9)$$

$$241 \quad Q_n(\mathbf{r}) = \int d\mathbf{r}_1 \frac{\exp(i\xi_n |\mathbf{r} - \mathbf{r}_1|)}{\sqrt{|\mathbf{r} - \mathbf{r}_1|}} \eta(\mathbf{r}_1) \frac{\partial p_0}{\partial z_1}(\mathbf{r}_1, z_1 = 0), \quad (10)$$

242 provided  $\xi_n |\mathbf{r} - \mathbf{r}_1| \gg 1$ . Equation (9) represents the scattered wave in the waveguide as a sum  
 243 of normal modes, with  $f_n(z)$  being the dependence of the acoustic pressure on depth in the  $n$ th  
 244 normal mode. In the summand, the factor in front of  $Q_n$  is controlled by the waveguide's  
 245 properties and the receiver depth. Dependence on horizontal coordinates of the receiver, the

246 incident wave, and the properties of the rough surface is described by the factor  $Q_n$ , Eq. (10).

247 When discussing the scattered wave, we will refer to  $Q_n$  as the mode amplitude for brevity.

248 Equations (9) and (10) show that each normal-mode component of  $p_{sc}$  is a result of

249 interference of the contributions generated by scattering at different points on the rough surface.

250 A more intuitive derivation of the normal-mode representation, Eqs. (9) and (10), of the scattered

251 wave is obtained using the concept of the effective sources of the scattered wave. The surface

252 density of the effective vertical force on the flat surface of a horizontally stratified oceanic

253 waveguide is given by Eq. (5). A point source of the vertical force with

254  $\mathbf{F}(\mathbf{r}_1, z_1) = (0, 0, F_0 \delta(\mathbf{r}_1) \delta(z_1))$  generates the acoustic field<sup>56</sup>

$$255 \quad p(\mathbf{R}) = \frac{iF_0}{4\rho(z_1)} \sum_n f_n(z) \frac{\partial f_n(z_1)}{\partial z_1} H_0^{(1)}(\xi_n |\mathbf{r} - \mathbf{r}_1|) \quad (11)$$

256 in the waveguide. Here, as in Eq. (6) for the Green's function, we disregard the continuous

257 spectrum of the field. Adding the contributions (11) of elementary effective sources located at

258 different points on the boundary, i.e., by calculating the convolution of the field of a unit vertical

259 force with the source density Eq. (5), leads again to Eqs. (9) and (10).

260 Equation (10) can be further simplified in the far field of the distributed effective source

261 of the scattered wave. However, the far field assumption proves to be too restrictive to be useful

262 in the  $T$ -phase excitation problem. For orientation, with the effective source dimensions of  $L_T =$

263  $O(10 \text{ km})$  and sound frequency  $f \sim 20 \text{ Hz}$  the far-field condition  $r \gg \xi_n L_T^2$  requires the range  $r$

264 from the epicenter to be more than 10 Mm. Here, we will obtain more relevant and widely

265 applicable results by taking into account that the correlation scale of the ocean surface roughness

266 is much smaller than  $L_T$ .

267 As discussed in Sec. III C, extensive areas on the ocean surface can contribute to  $T$ -phase  
 268 generation, and we need to allow for variations of the surface roughness statistics within these  
 269 areas. Let the ocean surface elevation  $\eta(\mathbf{r})$  have zero mean and be a locally stationary random  
 270 function;<sup>53</sup> then  $\langle \eta(\mathbf{r}) \rangle = 0$  and

$$271 \quad \langle \eta(\mathbf{r}_1) \eta(\mathbf{r}_2) \rangle = C \left( \mathbf{r}_1 - \mathbf{r}_2; \frac{\mathbf{r}_1 + \mathbf{r}_2}{2} \right). \quad (12)$$

272 Here and below angular brackets  $\langle \cdot \rangle$  denote statistical average;  $C$  has the meaning of the  
 273 correlation function of the surface elevations. The characteristic spatial scales  $l$  and  $L$  of the  
 274 variation of the correlation function with respect to the difference  $\mathbf{r}_1 - \mathbf{r}_2$  and centroid  $0.5(\mathbf{r}_1 + \mathbf{r}_2)$   
 275 coordinates satisfy the condition  $l \ll L$ . In the particular case of wide-sense stationary random  
 276 elevations,  $L \rightarrow \infty$  and the correlation function  $C$  depends only on  $\mathbf{r}_1 - \mathbf{r}_2$ . In terms of the  
 277 correlation function, the root mean square (rms) surface elevation  $\sigma_\eta$  and the roughness spectrum  
 278 are given by the equation  $\sigma_\eta = \langle \eta^2(\mathbf{r}) \rangle^{1/2} = \sqrt{C(0; \mathbf{r})}$  and

$$279 \quad S_\eta(\mathbf{q}; \mathbf{r}) = (2\pi)^{-2} \int C(\mathbf{r}_1; \mathbf{r}) \exp(-i\mathbf{q} \cdot \mathbf{r}_1) d\mathbf{r}_1. \quad (13)$$

280 The spectrum and rms elevation of the surface roughness gradually vary with the position  $\mathbf{r}$ .

281 At reflection from the random rough surface, mode amplitudes Eq. (10) are also random,  
 282 and  $\langle Q_n(\mathbf{r}) \rangle = 0$ . For the mode amplitude variance, from Eqs. (10) and (12) we find

$$283 \quad \langle |Q_n(\mathbf{r})|^2 \rangle = \int d\mathbf{r}_1 d\mathbf{r}_2 \frac{\exp\left[i\xi_n(|\mathbf{r} - \mathbf{r}_1| - |\mathbf{r} - \mathbf{r}_2|)\right]}{\sqrt{|\mathbf{r} - \mathbf{r}_1||\mathbf{r} - \mathbf{r}_2|}} C\left(\mathbf{r}_1 - \mathbf{r}_2; \frac{\mathbf{r}_1 + \mathbf{r}_2}{2}\right) \frac{\partial p_0(\mathbf{r}_1, 0)}{\partial z} \left( \frac{\partial p_0(\mathbf{r}_2, 0)}{\partial z} \right)^*. \quad (14)$$

284 Here and below the asterisk \* denotes complex conjugation. The main contribution to the  
 285 integral is from such  $\mathbf{r}_1$  and  $\mathbf{r}_2$  that  $|\mathbf{r}_1 - \mathbf{r}_2|$  is of the order of or smaller than the roughness  
 286 correlation scale  $l$ . When the horizontal separation  $r$  from the epicenter is large compared to the

287 size  $L_T$  of the effective source of the scattered wave and  $r \gg \xi_n l^2$ , one can approximate the  
 288 product  $|\mathbf{r} - \mathbf{r}_1| |\mathbf{r} - \mathbf{r}_2|$  with  $r^2$  in the integrand in Eq. (14) and retain in the exponent only linear  
 289 terms of the developments

$$290 \quad \left| \mathbf{r} - \frac{\mathbf{r}_1 + \mathbf{r}_2}{2} \pm \frac{\mathbf{r}_1 - \mathbf{r}_2}{2} \right| = \left| \mathbf{r} - \frac{\mathbf{r}_1 + \mathbf{r}_2}{2} \right| \pm \left| \mathbf{r} - \frac{\mathbf{r}_1 + \mathbf{r}_2}{2} \right|^{-1} \left( \mathbf{r} - \frac{\mathbf{r}_1 + \mathbf{r}_2}{2} \right) \cdot \frac{\mathbf{r}_1 - \mathbf{r}_2}{2} + O\left( \frac{|\mathbf{r}_1 - \mathbf{r}_2|^2}{|2\mathbf{r} - \mathbf{r}_1 - \mathbf{r}_2|^2} \right) \quad (15)$$

291 of  $|\mathbf{r} - \mathbf{r}_j|, j = 1, 2$ , in powers of  $|\mathbf{r}_1 - \mathbf{r}_2|$ . We also assume that the unperturbed field  $p_0$  can be  
 292 represented as

$$293 \quad p_0(\mathbf{r}, z) = P(\mathbf{r}, z) \exp[i\mathbf{q}_{in}(\mathbf{r}) \cdot \mathbf{r}] \quad (16)$$

294 in the vicinity of the ocean surface in water. Here the complex amplitude  $P$  and the local  
 295 horizontal wave vector  $\mathbf{q}_{in}$  are gradually varying functions of  $\mathbf{r}$ , which are little changed over  
 296 distances  $O(l)$ .

297 Changing integration variables in Eq. (14) from  $\mathbf{r}_1$  and  $\mathbf{r}_2$  to the difference and centroid  
 298 position vectors,  $\mathbf{r}_1 - \mathbf{r}_2$  and  $\mathbf{r}_3 = 0.5(\mathbf{r}_1 + \mathbf{r}_2)$ , and using Eqs. (13), (15), and (16), we obtain a  
 299 compact expression for the mode amplitude variance:

$$300 \quad \langle |Q_n(\mathbf{r})|^2 \rangle = \frac{4\pi^2}{r} \int d\mathbf{r}_3 \left| \frac{\partial P(\mathbf{r}_3, 0)}{\partial z} \right|^2 S_n(\xi_n \mathbf{e} - \mathbf{q}_{in}; \mathbf{r}_3), \quad \mathbf{e} = \frac{\mathbf{r} - \mathbf{r}_3}{|\mathbf{r} - \mathbf{r}_3|}. \quad (17)$$

301 Here  $\mathbf{e}$  has the meaning of the unit horizontal vector from an elementary scatterer to the  
 302 observation point, and  $\xi_n \mathbf{e}$  is the horizontal wave vector of the  $n$ th mode propagating from  $\mathbf{r}_3$  to  $\mathbf{r}$ .  
 303 For the distant observation points that we consider, it is close to the unit horizontal vector from  
 304 the epicenter to the observation point:  $\mathbf{e} = r^{-1}\mathbf{r} + O(L_T/r)$ . Inspection shows that Eq. (17) is  
 305 consistent with the more general result, Eq. (9) in Ref. 59, for the cross-correlation function of  
 306 the surface reverberation in the oceanic waveguide.

307        Integration in Eq. (17) is over the entire horizontal plane  $z = 0$ . The ocean surface area  
 308        that significantly contributes to normal mode excitation is controlled by the decrease of the  
 309        amplitude of the unperturbed field  $p_0$  with horizontal separation from the epicenter and is  
 310        affected by spatial distribution of the surface roughness. The integrand is proportional to the  
 311        average power scattered into the  $n$ th mode in the vicinity of the point  $(\mathbf{r}_3, 0)$  on the ocean surface.  
 312        The contributions of different points into the average mode's power are added incoherently,  
 313        according to Eq. (17). The first argument,  $\xi_n \mathbf{e} - \mathbf{q}_{in}$ , of the roughness spectrum  $S_\eta$  in the integrand  
 314        equals the change of the horizontal wave vector of sound at scattering and corresponds to  
 315        Bragg's scattering, as expected in the first approximation of the small perturbation method.<sup>53, 54</sup>  
 316        We will use Eq. (17) in Section III to investigate the effects on  $T$ -phase generation of the wind  
 317        speed, sea swell parameters, and depth of the earthquake focus.

318        Acoustic power flux in  $T$  waves can be calculated by integrating the power flux density  
 319        over the cylindrical surface  $r = \text{const.} > L_T$ ,  $0 < z < \infty$ . At distances  $r$  from the epicenter that are  
 320        large compared to the diameter  $L_T$  of the region, where  $T$  waves are excited,  $\nabla Q_n \approx i\xi_n r^{-1} Q_n \mathbf{r}$   
 321        according to Eq. (10). Using this equation and the normalization condition (8), for the power flux  
 322         $J_n$  in the  $n$ th mode we find

$$323 \quad J_n = \frac{r}{16\pi\omega} \left( \frac{1}{\rho} \frac{\partial f_n}{\partial z} \right)_{z=0}^2 \int_0^{2\pi} |Q_n^2(r \cos \varphi, r \cos \varphi)| d\varphi \quad (18)$$

324        from Eq. (9). The total power flux is given by the sum of the contributions  $J_n$ , Eq. (18), of all  
 325        propagating normal modes. For a random rough surface with the spectrum  $S_\eta$ , Eqs. (17) and (18)  
 326        give

$$327 \quad \langle J_n \rangle = \frac{\pi}{4\omega} \left( \frac{1}{\rho} \frac{\partial f_n}{\partial z} \right)_{z=0}^2 \int_0^{2\pi} \left[ \int d\mathbf{r}_3 \left| \frac{\partial P(\mathbf{r}_3, 0)}{\partial z} \right|^2 S_\eta(\xi_n \mathbf{e} - \mathbf{q}_{in}; \mathbf{r}_3) \right] d\varphi, \quad (19)$$

328 where  $\mathbf{e} = (\cos\varphi, \sin\varphi, 0)$ . As expected, the power flux is independent of  $r$  as long as the effect of  
 329 absorption on the propagating normal mode is negligible over ranges of the order of  $r$ .

330

331 **C. Excitation of normal modes at volume scattering by internal gravity waves**

332 Consider internal gravity waves propagating in otherwise horizontally stratified,  
 333 stationary ocean. The internal wave-induced currents  $\mathbf{u}$  and variations of the sound speed,  $\delta c$ , and  
 334 density,  $\delta\rho$ , from their unperturbed (background) values  $c(z)$  and  $\rho(z)$  are horizontally  
 335 inhomogeneous. The currents are slow and environmental perturbations are weak in the  
 336 following sense:  $|\delta c| + u \ll c$ ,  $\delta\rho \ll \rho$ . Neglecting terms of the second order in the small ratio  
 337  $u/c$ , monochromatic acoustic waves satisfy the following wave equation<sup>55, 60</sup> in the horizontally  
 338 inhomogeneous ocean with slow currents:

$$339 \quad \nabla \cdot \left( \frac{\nabla p}{\rho_0} \right) + \frac{\omega^2}{\rho_0 c_0^2} p + \frac{2i\omega}{\rho_0 c_0^2} \mathbf{u} \cdot \nabla p - \frac{2i}{\omega} \nabla \cdot \left( \frac{1}{\rho_0} \sum_{j=1}^3 \frac{\partial p}{\partial x_j} \frac{\partial \mathbf{u}}{\partial x_j} \right) = 0. \quad (20)$$

340 Here  $\rho_0 = \rho + \delta\rho$ ,  $c_0 = c + \delta c$ , and  $(x_1, x_2, x_3) = (x, y, z)$  are Cartesian coordinates. Acoustic  
 341 pressure  $p = p_0 + p_{sc}$  consists of the acoustic pressure  $p_0$  in the horizontally stratified ocean and  
 342 the perturbation (scattered wave)  $p_{sc}$ . In the water column,  $p_0$  satisfies Eq. (20) with  $\mathbf{u} = 0$  and  $\rho_0$   
 343 and  $c_0$  replaced with  $\rho$  and  $c$ , respectively.

344 The scattered wave vanishes when the environmental perturbations  $\mathbf{u}$ ,  $\delta c$ , and  $\delta\rho$  vanish.  
 345 Retaining only terms of the first order in the acoustic and environmental perturbations, from Eq.  
 346 (20) we find

$$347 \quad \nabla \cdot \left( \frac{\nabla p_{sc}}{\rho} \right) + \frac{\omega^2}{\rho c^2} p_{sc} = i\omega A_{sc} + \nabla \cdot \left( \frac{\mathbf{F}_{sc}}{\rho} \right), \quad (21)$$

348 where

349 
$$A_{sc} = \frac{-i\omega p_0}{\rho c^2} \left( \frac{\delta\rho}{\rho} + \frac{2\delta c}{c} \right) - \frac{2}{\rho c^2} \mathbf{u} \cdot \nabla p_0, \quad \mathbf{F}_{sc} = \frac{\delta\rho}{\rho} \nabla p_0 + \frac{2i}{\omega} \sum_{j=1}^3 \frac{\partial p_0}{\partial x_j} \frac{\partial \mathbf{u}}{\partial x_j}. \quad (22)$$

350 The above assumptions correspond to calculation of the scattered wave in the single-scattering,  
 351 or (first) Born, approximation. Comparison of Eqs. (3) and (21) shows, that in the Born  
 352 approximation the scattered wave can be viewed as the wave generated in horizontally stratified  
 353 ocean by distributed virtual sources with volume densities  $A_{sc}$  and  $\mathbf{F}_{sc}$ , Eq. (22), respectively, of  
 354 the volume velocity and external force. Using Eq. (4) for the field of distributed sources and Eq.  
 355 (6) for the Green's function, we find the scattered wave in the following form:

356 
$$p_{sc}(\mathbf{r}, z) = \sum_n \frac{f_n(z) \exp(-i\pi/4)}{\sqrt{8\pi\xi_n}} V_n(\mathbf{r}), \quad (23)$$

357 where

358 
$$V_n(\mathbf{r}) = \int d\mathbf{r}_1 \frac{\exp(i\xi_n |\mathbf{r} - \mathbf{r}_1|)}{\sqrt{|\mathbf{r} - \mathbf{r}_1|}} \int \frac{dz_1}{\rho} \left[ \left( \omega\rho A_{sc} + \xi_n \frac{\mathbf{r} - \mathbf{r}_1}{|\mathbf{r} - \mathbf{r}_1|} \cdot \mathbf{F}_{sc} \right) f_n + i \frac{\partial f_n}{\partial z_1} (\mathbf{F}_{sc})_z \right], \quad (24)$$

359 and  $(\mathbf{F}_{sc})_z$  stands for the vertical component of the vector  $\mathbf{F}_{sc}$  defined in Eq. (22). Equation (23)  
 360 represents the scattered wave as a sum of normal modes, with  $V_n$  describing the dependence of  
 361 the  $n$ th mode amplitude on horizontal coordinates.

362 In small-amplitude, or linear, internal waves, the sound speed and density perturbations  
 363 are proportional to the vertical displacement  $\zeta$  of fluid particles due to the internal wave:

364  $\delta c = \alpha_1(z) c \zeta, \quad \delta\rho = \alpha_2(z) \rho \zeta$ .<sup>46</sup> Vertical velocity  $u_3$  of fluid particles is given by time  
 365 derivative of  $\zeta$ , and horizontal components of the velocity are related to  $\zeta$  by the  
 366 incompressibility condition  $\nabla \cdot \mathbf{u} = 0$ .<sup>46</sup> In a random field of linear internal waves, let the vertical  
 367 displacement  $\zeta$  have zero mean and be a random function that is locally stationary in the  
 368 horizontal plane. Then the correlation function of vertical displacements is related to the spatial  
 369 spectrum  $S_\zeta$  of internal waves as follows:

370 
$$\langle \zeta(\mathbf{r}_1, z_1) \zeta(\mathbf{r}_2, z_2) \rangle = \int S_\zeta \left( \mathbf{q}; z_1, z_2; \frac{\mathbf{r}_1 + \mathbf{r}_2}{2} \right) e^{i\mathbf{q}(\mathbf{r}_1 - \mathbf{r}_2)} d\mathbf{q}, \quad (25)$$

371 Under these assumptions, the densities of the effective sources of the scattered sound wave are  
 372 also zero-mean random functions that are locally stationary in the horizontal plane. Using Eq.  
 373 (22), the spectra of the random sources can be related to the spectrum of the vertical  
 374 displacement of fluid particles; importantly, the source spectra have the same spatial scales as  $S_\zeta$ .

375 At scattering by random internal waves, the mode amplitudes  $V_n$  are random and have  
 376 zero mean. Calculation of the variance of the mode amplitude, and particularly the reduction of a  
 377 double integral over horizontal coordinate to a single integral, is similar to the calculation of  
 378  $\langle |Q_n|^2 \rangle$  in Sec. II B. From Eqs. (16), (22), (24), and (25) we find that

379 
$$\begin{aligned} \langle |V_n|^2(\mathbf{r}) \rangle &= \frac{4\pi^2}{r} \int d\mathbf{r}_3 dz_1 dz_2 \Phi(\mathbf{r}_3, z_1) \Phi(\mathbf{r}_3, z_2)^* S_\zeta(\xi_n \mathbf{e} - \mathbf{q}_{in}; z_1, z_2; \mathbf{r}_3), \\ \Phi(\mathbf{r}, z) &= [\alpha_2 \xi_n \mathbf{e} \cdot \mathbf{q}_{in} - k^2 (2\alpha_1 + \alpha_2)] \frac{f_m P}{\rho} + \frac{\alpha_2}{\rho} \frac{\partial P}{\partial z} \frac{\partial f_m}{\partial z}. \end{aligned} \quad (26)$$

380 Here, the unit horizontal vector  $\mathbf{e}$  is the same as in Eq. (17). For brevity, contributions of the  
 381 internal wave-induced currents into sound scattering are not included in Eq. (26). Equations (17)  
 382 and (26), which describe the variances of mode amplitudes that are proportional to the power  
 383 flux in respective normal modes resulting, respectively, from surface and volume scattering,  
 384 differ by additional integration over depths  $z_1$  and  $z_2$  of volume scatterers in Eq. (26). Note that  
 385 the spatial spectra  $S_n$  and  $S_\zeta$  of the surface elevation and the vertical displacement due to internal  
 386 waves in Eqs. (17) and (26) have the same vector argument  $\xi_n \mathbf{e} - \mathbf{q}_{in}$ , which equals the  
 387 difference of the horizontal wave vectors of the normal mode and the incident wave.

388 Because of the large values of the compressional and shear wave speeds around the  
 389 earthquake focus, earthquake-generated incident waves propagate at steep grazing angles in the  
 390 water column, see Sec. III C for details. Therefore,  $|\xi_n \mathbf{e} - \mathbf{q}_{in}| \sim \xi_n$ . The internal wave spectrum

391 peaks around 5 km horizontal wavelength, with minimum and maximum internal wave  
 392 wavelength in the ocean being about 0.5 km and 50 km, respectively.<sup>46</sup> In the 3–50 Hz frequency  
 393 range of observed *T* waves, horizontal wavelength  $2\pi/\zeta_n$  of acoustic normal modes ranges from  
 394 about 30–500 m. Hence, the internal wave spectrum in the integrand in Eq. (26) has negligibly  
 395 small values. The short-wave tail of the internal wave spectrum can possibly contribute to  
 396 generation of the lowest-frequency *T* waves away from the earthquake epicenter. In other words,  
 397 the internal wave field lacks the relatively short horizontal scales (< 500 m) that are required for  
 398 Bragg scattering of the earthquake-generated body waves into normal modes of the underwater  
 399 waveguide. As discussed below, ocean surface roughness spectrum is rich in the spatial scales  
 400 required for Bragg scattering into normal modes and, therefore, efficiently contributes to *T*-phase  
 401 generation.

402

### 403 **III. CONTRIBUTIONS OF WIND SEAS AND SEA SWELL INTO *T*-PHASE**

#### 404 **GENERATION**

##### 405 **A. *T*-phase excitation due to wind seas**

406 Dependence of the ocean surface roughness on wind speed and fetch have been studied  
 407 extensively, which allows for a reliable prediction of *T* waves generation at scattering by sea  
 408 surface roughness. Here, we use a simple Pierson-Moskovitz model<sup>43, 44</sup> of fully developed wind  
 409 seas to investigate the dependence of amplitudes of the normal mode components of the scattered  
 410 acoustic wave on its frequency, wind speed, and direction of propagation of the incident wave.

411 The Pierson-Moskovitz spectrum<sup>43, 44</sup> of the random surface elevation  $\eta$  is given by the  
 412 following equations:

$$413 \quad S_\eta(\mathbf{q}; \mathbf{r}) = W(q) D_w(q, \psi), \quad \int_{-\pi}^{\pi} D_w(q, \psi) d\psi = 1, \quad (27)$$

414

$$W(q) = \frac{0.024}{q^4} \exp\left(-\frac{0.74g^2}{q^2 U^4}\right). \quad (28)$$

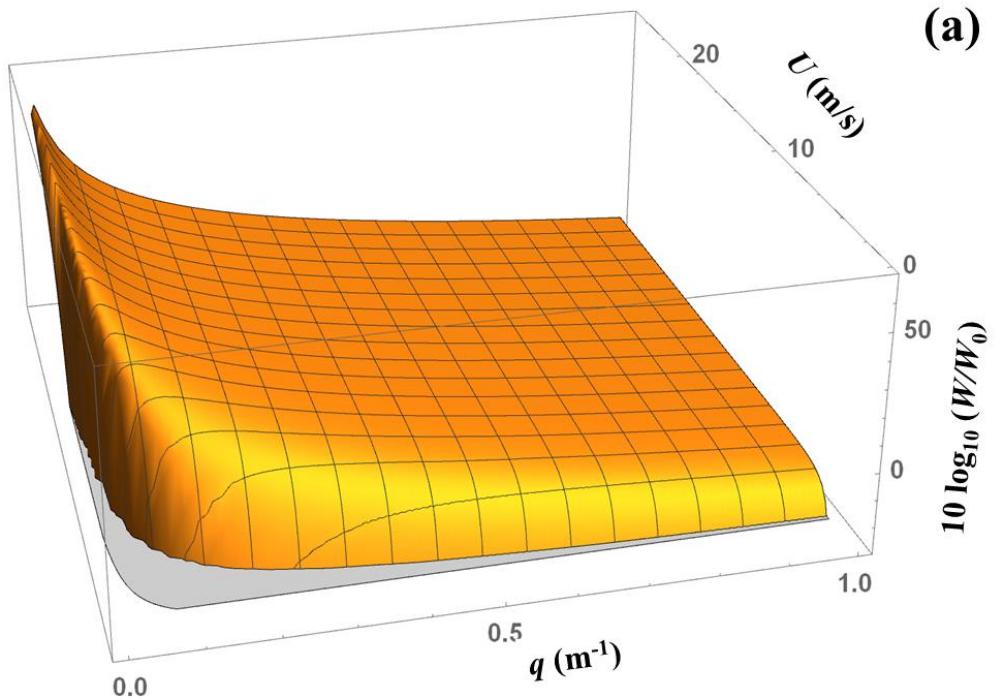
415 Here  $g$  is the acceleration due to gravity;  $U$  is the wind speed measured at height of 19.5 m above  
 416 the sea surface. The factor  $D_W$  describes the directionality of the surface waves;  $\mathbf{q} = q (\cos\psi,$   
 417  $\sin\psi, 0)$  is the wave vector of the waves, and angle  $\psi$  indicates the vector  $\mathbf{q}$  direction. The wind  
 418 speed may gradually change along the ocean surface:  $U = U(\mathbf{r})$ , and  $W$  and  $D_W$  in Eqs. (27) and  
 419 (28) depend on  $\mathbf{r}$  via  $U$ . In wind waves with the Pierson-Moskovitz spectrum, the spectral peak is  
 420 located at  $q_p = 0.70gU^{-2}$ ; and rms surface elevation  $\sigma_\eta = 0.13U^2/g$ . The wave height rapidly  
 421 increases, and the spectrum peak shifts towards longer waves, when the wind speed increases  
 422 (Fig. 2a). According to Eq. (28), the spectrum falls off very rapidly (exponentially) as the surface  
 423 wave wavelength becomes longer than at the spectrum peak, i.e., at  $q < q_p$ . The spectrum  
 424 decrease is much slower for short gravity waves, i.e., at  $q > q_p$  (Fig. 2a). Because of the Bragg  
 425 scattering condition, these properties of the wind wave spectrum are directly reflected in the  
 426 spectrum of abyssal  $T$ -waves and its wind dependence.

427 The rms amplitude  $\langle |Q_n^2| \rangle^{1/2}$  of the  $n$ th normal mode component of the  $T$ -phase field is  
 428 given by Eq. (17). Figure 2b illustrates the wind dependence of the  $T$ -phase energy in terms of  
 429 the contribution to the acoustic power flux in a normal mode from a unit area of the sea surface  
 430 above the earthquake focus. In this geometry, the horizontal wave vector of the incident wave  $\mathbf{q}_{in}$   
 431 = 0 in the right side of Eq. (17). Then, directionality of the  $T$ -phase radiation is given by the  
 432 factor  $D_W$  in the wind wave spectrum Eq. (27). Equation (18) shows that the wind speed  
 433 dependence of the acoustic power flux in the  $T$ -wave is obtained by integrating (or averaging) of  
 434  $|Q_n^2|$  over the  $T$ -wave propagation direction. In Fig. 2b we show the mode amplitude squared,

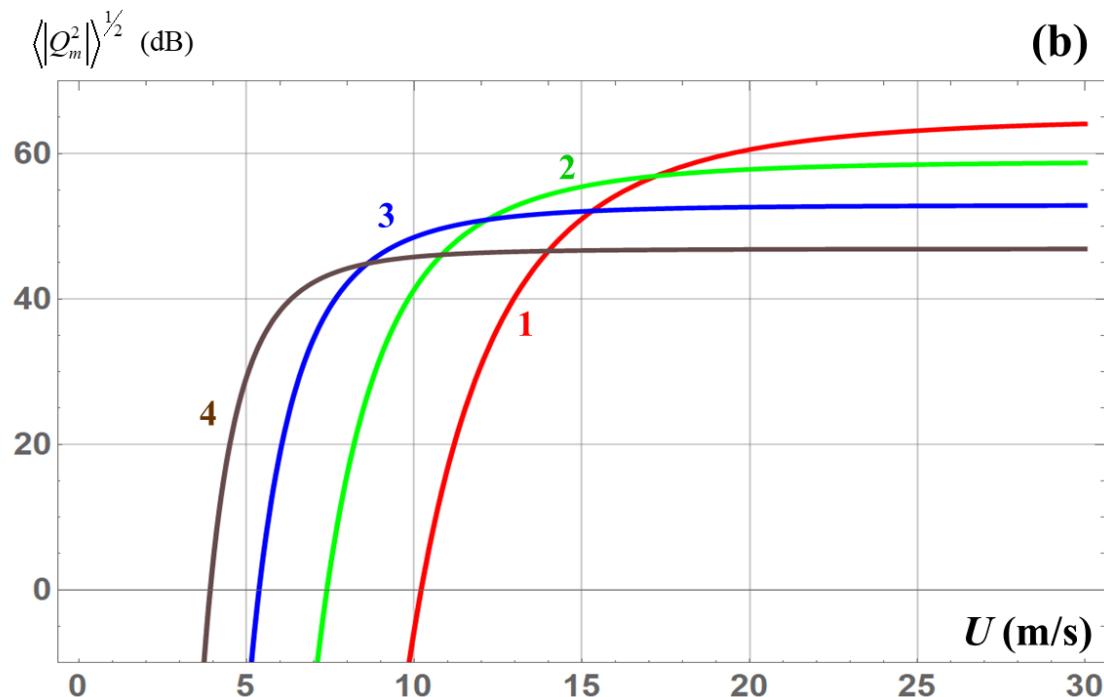
435  $|Q_n^2|$ , that is averaged over the statistical ensemble of fully developed wind waves. It is also  
 436 averaged over the  $T$ -wave propagation direction for a given wind direction or, equivalently, over  
 437 the wind direction for a given receiver position. On the other hand, it follows from Eq. (27) that,  
 438 after averaging over the wind direction,  $\langle |Q_n^2| \rangle$  is given by Eq. (17), where  $S_\eta(\xi_n \mathbf{e} - \mathbf{q}_{in}; \mathbf{r}_3)$  is  
 439 replaced with  $W(|\xi_n \mathbf{e} - \mathbf{q}_{in}|)$  in the integrand. Hence, the result is independent of the surface wave  
 440 directionality  $D_W$  and its dependence on  $q$  in Eq. (27). Since averaging over wind direction is  
 441 equivalent to integration over receiver azimuth, acoustic power flux in  $T$ -waves is also  
 442 independent of  $D_W$  at normal incidence of ballistic waves. Numerical values of the sound  
 443 frequency  $f$  indicated in Fig. 2b refer to the mode with the nominal phase speed  $c_n$  of 1500 m/s.  
 444 For a generic mode dispersion relation  $c_n = c_n(f)$ , the frequency  $f$  should be re-scaled to  $(1500$   
 445 m/s) $f/c_n(f)$ .

446  $T$ -phase amplitude rapidly increases with the wind speed for weak and moderate winds  
 447 and saturates at very high wind speeds (Fig. 2b). Higher acoustic frequencies are more readily  
 448 excited by weaker winds and saturate at smaller wind speeds. For an incident wave with a white  
 449 spectrum, higher acoustic frequencies dominate in the  $T$ -phase spectrum at low wind speeds,  
 450 while low frequencies prevail at strong winds. Abyssal  $T$ -phase energy and spectrum can be very  
 451 sensitive to the wind speed. Away from the saturation regime, a drastic, 40 dB increase in the  
 452 narrow-band mode amplitude requires an increase in the wind speed of just a few meters per  
 453 second (Fig. 2b).

454



455



456

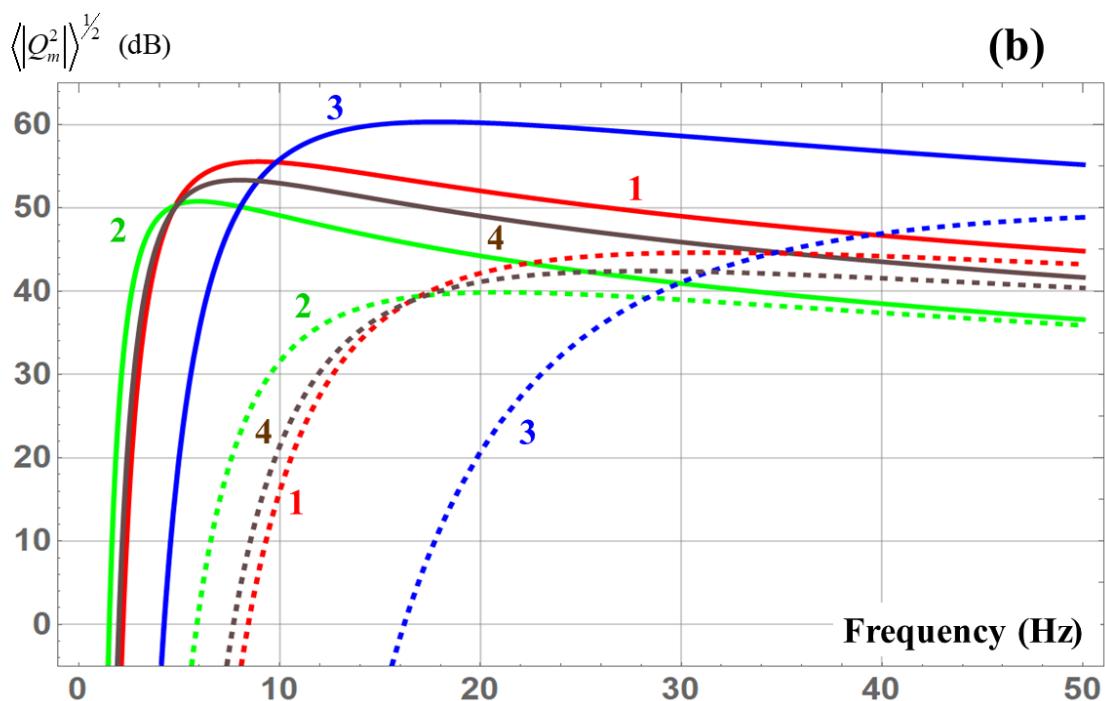
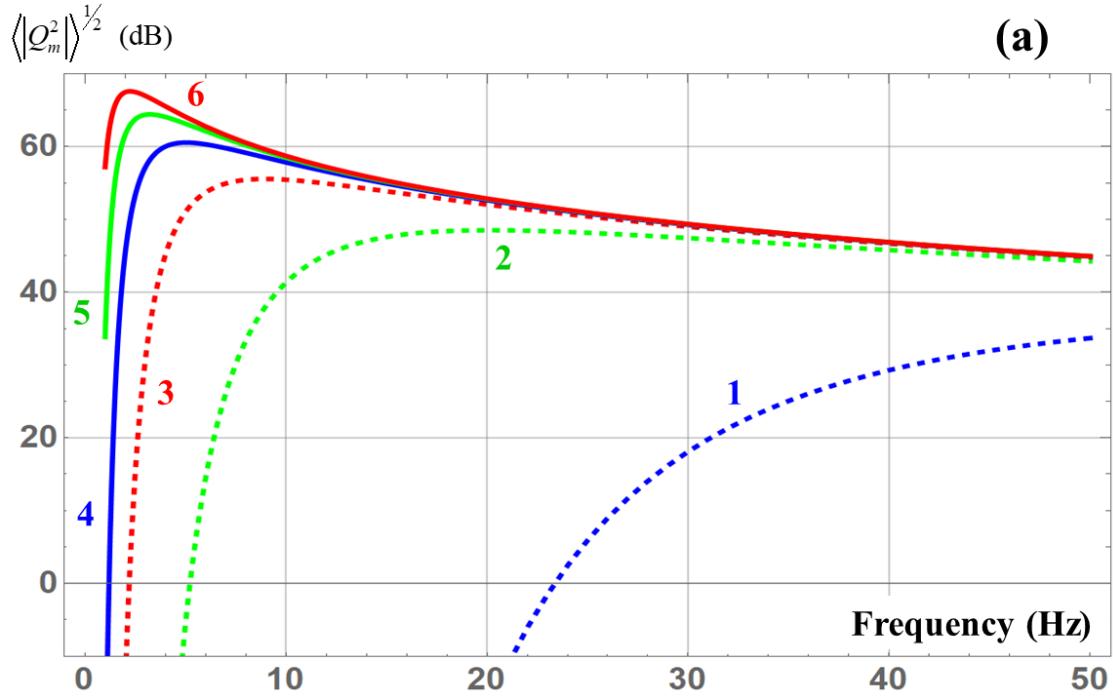
457 **Figure 2. (Color online)** Dependence of the abyssal  $T$ -phase mode amplitude on wind speed. (a)  
 458 Azimuthally averaged Pierson-Moskovitz spectrum of wind waves as described by Eq. (28) is  
 459 shown as a function of surface gravity wave wavenumber  $q$  and wind speed  $U$  at 19.5 m above

460 the sea surface;  $W_0 = 1 \text{ m}^4$ . (b) The rms amplitude of a normal mode of the  $T$ -waves generated by  
 461 scattering on wind seas in a unit area above the earthquake focus is shown for four frequencies: 5  
 462 Hz (1), 10 Hz (2), 20 Hz (3) and 40 Hz (4), and the mode phase speed of 1500 m/s. The mode  
 463 amplitude is arbitrarily normalized assuming a frequency-independent acoustic pressure  
 464 amplitude in the earthquake-generated incident wave.

465

466 The spectrum of  $T$ -waves at different wind speeds is further illustrated in Fig. 3. The  
 467 figure shows the mode amplitude squared,  $|\mathcal{Q}_n^2|$ , which is averaged over the statistical ensemble  
 468 of fully developed wind waves and over the wind direction. Therefore, the result is independent  
 469 of the wind waves directionality that is described by the factor  $Dw(q, \psi)$  in Eq. (27). Similar to  
 470 Fig. 2b, Fig. 3a refers to the  $T$ -phase generation at normal incidence of ballistic waves from the  
 471 earthquake focus. The figure shows a steady increase of normal mode amplitudes with wind  
 472 speed in the entire range of  $T$ -phase frequencies. The most distinctive feature of the predicted  $T$ -  
 473 phase spectra is a sharp low-frequency cutoff. At low acoustic frequencies, Bragg scattering into  
 474 proper normal modes of the underwater waveguide requires long wind waves, with their  
 475 wavevector matching the horizontal wave vector of the acoustic normal mode, see Eq. (17). For  
 476 instance, the resonance scattering into the modes at 5 Hz occurs at surface gravity waves with  
 477 wavelength of about 300 m. Thus, the low-frequency acoustic cutoff reflects the sharp drop in  
 478 the wind wave spectrum at  $q < q_p$ . The cutoff shifts to lower acoustic frequencies and the  $T$ -  
 479 phase spectrum broadens when the wind speed increases (Fig. 3a).

480



483 **Figure 3. (Color online)** Dependence of the amplitude of a modal component of the  $T$ -wave,  
484 which is generated by scattering on fully developed wind seas, on sound frequency and the mode  
485 propagation direction. (a) The frequency dependences of the rms amplitude of a normal mode,

486 which is generated by scattering in a unit area above the earthquake focus, are shown for six  
 487 wind speeds: 5 m/s (1), 10 m/s (2), 15 m/s (3), 20 m/s (4), 25 m/s (5) and 30 m/s (6). (b) The rms  
 488 amplitude of a normal mode is shown for scattering in a unit area above the earthquake focus (1)  
 489 and away from the epicenter (2–4), where the grazing angle of the earthquake-generated incident  
 490 wave is  $60^\circ$  at the depth, where  $c(z) = c_m$ . The horizontal propagation directions of the mode and  
 491 incident wave are either opposite (2), the same (3), or orthogonal (4). Solid and dashed lines  
 492 refer to the wind speeds of 15 m/s and 8 m/s, respectively. A nominal value of 1500 m/s is  
 493 assumed for the mode phase speed  $c_m$ .

494

495 The frequency dependence of the efficiency of *T*-wave generation by scattering of  
 496 obliquely incident waves is qualitatively similar to but quantitatively different from the case of  
 497 normal incidence. This is illustrated in Fig. 3b. At points on the ocean surface away from the  
 498 earthquake epicenter, *T*-phase is generated with different amplitudes in different horizontal  
 499 propagation directions, even after averaging over the wind direction (Fig. 3b). For obliquely  
 500 incident waves, wind waves of different wavelength are responsible for the *T*-waves propagating  
 501 in different azimuthal directions, see Eq. (17). When the incident wave and *T*-wave propagate in  
 502 opposite horizontal directions, the low-frequency cutoff shifts somewhat towards lower  
 503 frequencies; when the propagation directions are the same, there is a more significant shift  
 504 towards higher frequencies (Fig. 3b).

505 In addition to the frequency dependence of the generation efficiency of each normal  
 506 mode that is illustrated in Fig. 3, *T*-phase spectrum at a distant receiver is influenced by the  
 507 number of propagating modes, which increases with frequency, frequency-dependent  
 508 transmission losses due to sound attenuation, and the spectrum of the seismic source.

509

510 **B. *T*-phase excitation due to swell**

511 Statistically, wave height and surface gravity wave energy are dominated by sea swell, rather  
 512 than wind waves, almost everywhere in the World Ocean.<sup>45</sup> We argue below that swell is also  
 513 expected to dominate in generation of abyssal *T*-waves.

514 Sea swell is generated by very strong winds in distant storms. Because of the pronounced  
 515 dispersion of surface gravity waves in deep water, swell is observed at large distances from its  
 516 source as a wave train of long gravity waves with nearly identical wavelengths. A typical width  
 517 of the wavetrain is several tens of wavelengths across the wavefronts with even longer extent  
 518 along the wavefronts.<sup>61</sup> Thus, ocean surface elevations due to swell have much larger correlation  
 519 length than the surface roughness caused by wind waves. This difference has a major effect on  
 520 scattering of low-frequency sound. While wind waves can be modeled as a random wave field, it  
 521 is more appropriate to model a snapshot of sea swell in an area of several and perhaps a few tens  
 522 of km as a deterministic wave field.

523 Unlike wind waves, there are no widely accepted swell models. We will utilize the  
 524 following simple, idealized model to illustrate distinctive features of *T*-phase generation at sound  
 525 scattering by swell. At the time of an earthquake, let the surface elevation  $\eta$  in a swell wave train  
 526 be

$$527 \quad \eta(x, y) = \sqrt{2}\sigma_\eta(y) \sin(\mu x - \mu x_0), \quad |x - x_0| < L/2, \quad (29)$$

528 in a region of width  $L$  in the direction of swell propagation, which is chosen as the  $x$  coordinate  
 529 axis;  $\eta = 0$  at  $|x - x_0| \geq L/2$ . A large, integer number of swell wavelengths  $2\pi/\mu$  fits in the band  
 530  $|x - x_0| \leq L/2$ , and  $\eta(x, y)$  is a continuous function of horizontal coordinates. The rms surface  
 531 elevation  $\sigma_\eta$  is a gradually varying function of  $y$  and tends to zero at  $|y - y_0| \rightarrow \infty$ , so that the

532 energy of the wavetrain is finite. The center of the swell wavetrain is at the point  $(x_0, y_0, 0)$ ,  
 533 which can be located either at the earthquake epicenter  $(0, 0, 0)$  or away from it.

534 At scattering of ballistic sound waves [Eq. (16)] at the ocean surface with surface  
 535 elevations Eq. (29), Eq. (10) for the amplitude of a  $T$ -phase modal component becomes

$$536 \quad Q_n(\mathbf{r}) = \sqrt{2} \int_{x_0-L/2}^{x_0+L/2} dx_1 \sin(\mu x_1 - \mu x_0) \int_{-\infty}^{\infty} dy_1 \frac{\exp(i\xi_n |\mathbf{r} - \mathbf{r}_1| + i\mathbf{q}_{in} \cdot \mathbf{r}_1)}{\sqrt{|\mathbf{r} - \mathbf{r}_1|}} \sigma_\eta(y_1) \frac{\partial P(\mathbf{r}_1, 0)}{\partial z}, \quad (30)$$

537 where the two-dimensional horizontal position vector  $\mathbf{r}_1 = (x_1, y_1)$ . In the integral over  $y_1$  in Eq.  
 538 (30), the integrand contains a rapidly varying exponential and slowly varying functions  $\sigma_\eta, \mathbf{q}_{in} =$   
 539  $(q_{in1}, q_{in2}, 0)$ , and  $\partial P/\partial z$ . The integral can be calculated by the method of stationary phase.<sup>55</sup>  
 540 Disregarding small derivatives of  $q_{in2}$ , equation for the stationary point<sup>55</sup>  $y_1 = y_{1s}$  becomes

$$541 \quad \frac{y - y_{1s}}{\sqrt{(x - x_1)^2 + (y - y_{1s})^2}} = \frac{q_{in2}}{\xi_n}. \quad (31)$$

542 For any observation point at  $|x - x_0| > L/2$ , the integrand has a single stationary point. By  
 543 approximating the integral over  $y_1$  in Eq. (30) by contribution of the stationary point,<sup>55</sup> we obtain

$$544 \quad Q_n(\mathbf{r}) = 2\sqrt{\pi\xi_n} e^{i\pi/4} \times \int_{x_0-L/2}^{x_0+L/2} \frac{\sin(\mu x_1 - \mu x_0) \sigma_\eta(y_{1s}) \frac{\partial P(x_1, y_{1s}, 0)}{\partial z}}{\sqrt{\xi_n^2 - q_{in2}^2}} \exp\left(i\sqrt{\xi_n^2 - q_{in2}^2} |x - x_1| + iq_{in1}x_1 + iq_{in2}y\right) dx_1. \quad (32)$$

545 Assuming negligible variation of  $\sigma_\eta, \mathbf{q}_{in}$ , and  $\partial P/\partial z$  with  $x_1$  within the swell wave train,  
 546 the integral in the right side of Eq. (32) is easily calculated, and we obtain

$$547 \quad Q_n(\mathbf{r}) = 2e^{i\pi/4} \sqrt{\frac{\pi\xi_n}{\xi_n^2 - q_{in2}^2}} \frac{\partial P}{\partial z} \sigma_\eta L \left( \frac{\sin Y_2}{Y_2} - \frac{\sin Y_1}{Y_1} \right) \exp\left(i\sqrt{\xi_n^2 - q_{in2}^2} |x - x_0| + iq_{in2}y\right), \quad (33)$$

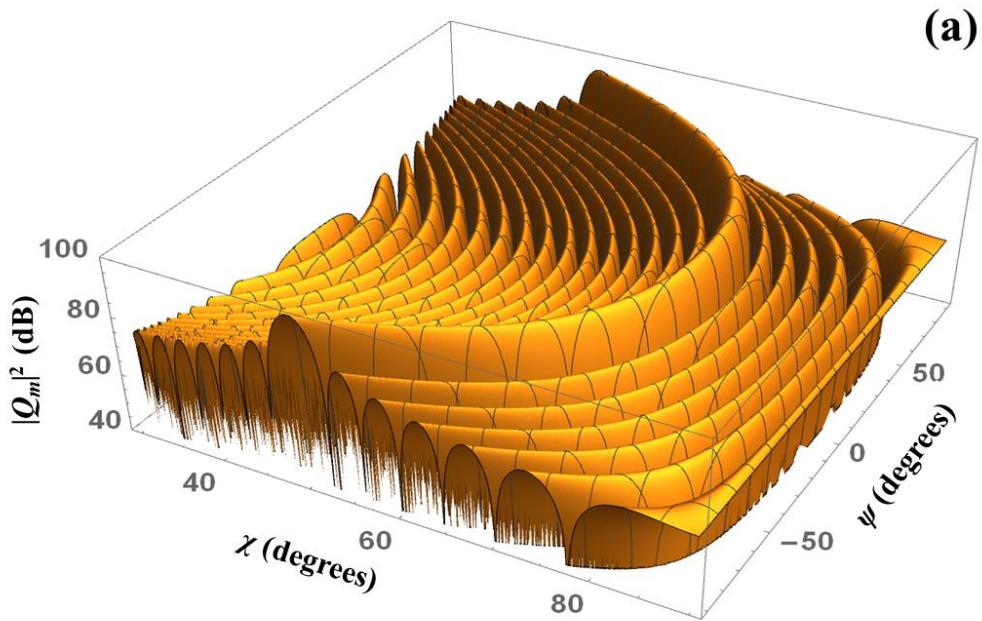
548 where

$$549 \quad Y_j = \left[ q_{in1} - \sqrt{\xi_n^2 - q_{in2}^2} \frac{x - x_0}{|x - x_0|} + (-1)^j \mu \right] \frac{L}{2}, \quad j = 1, 2. \quad (34)$$

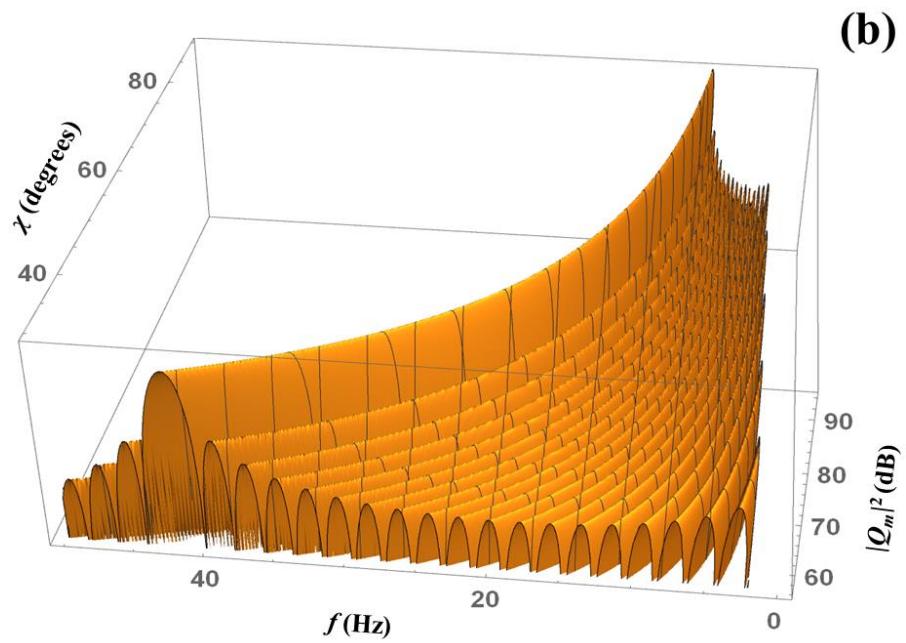
550 Equations (33) and (34) give the normal mode amplitudes in the abyssal  $T$  waves due to swell at  
551 the observation points at  $|x - x_0| > L/2$ , i.e., outside of the swell wavetrain.

552 The Bragg scattering condition and the narrow-band, quasi-periodic nature of surface  
553 elevation in swell wavetrains combine to produce rather different dependence of  $T$ -phase energy  
554 on the mode frequency and propagation direction than in the case of wind waves (cf. Figs. 2b  
555 and 3 with Fig. 4). Figure 4 illustrates predictions of Eqs. (33) and (34). At a given sound  
556 frequency and normal mode propagation direction, a swell wavetrain most efficiently generates  
557  $m$ th normal mode at a specific grazing angle  $\chi$  of the ballistic wave (Fig. 4a), with secondary  
558 peaks in  $\chi$  giving  $T$  waves that are weaker by tens of dB (Fig. 4a). The contrast between the main  
559 and subsequent peaks is controlled by the parameter  $\mu L \gg 1$ . The resonance value of the grazing  
560 angle  $\chi$  depends on the wavetrain position relative to the epicenter via the angle between  
561 azimuthal directions of the swell and ballistic wave propagation (Fig. 4a). For the sound  
562 frequency and swell wavelength (10 Hz and 200 m) in Fig. 4a, resonance excitation occurs for  
563 the wavetrains away from the epicenter, where  $\chi$  is between about  $47^\circ$  and  $78^\circ$ .

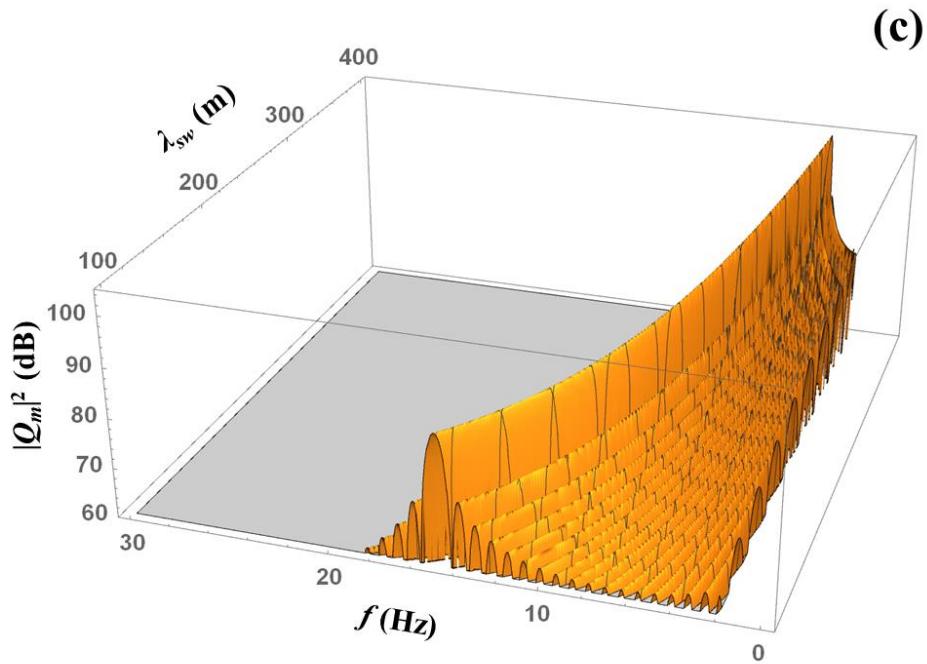
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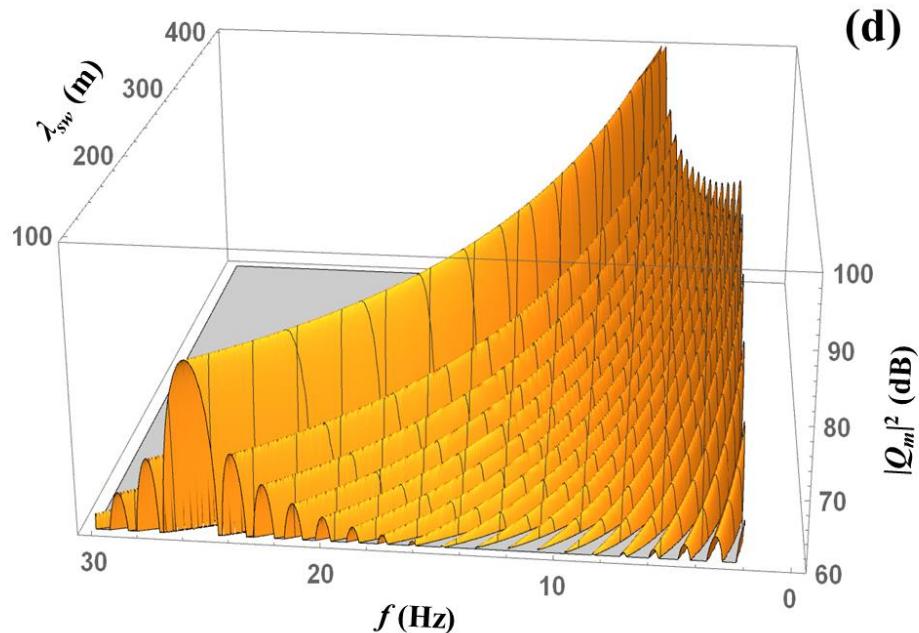
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566



567



568

569 **Figure 4. (Color online)** Generation of  $T$  waves at scattering of ballistic waves from an  
 570 earthquake by a wavetrain of sea swell. (a) Dependence of the amplitude of a normal-mode  
 571 component of the  $T$  wave on the grazing angle  $\chi$  of incident wave at the location of the swell

572 wavetrain and the angle  $\psi$  between the azimuthal directions of propagation of the incidence  
 573 waves and swell. Sound frequency is 10 Hz. Swell wavelength  $\lambda_{sw} = 200$  m. (b) Variation of the  
 574 normal mode amplitude with the grazing angle of incident waves and sound frequency, when the  
 575 angle between the azimuthal directions of propagation of the incidence waves and swell is  $45^\circ$ .  
 576 Swell wavelength  $\lambda_{sw} = 200$  m. (c) Dependence of the acoustic mode amplitude on sound  
 577 frequency and the wavelength of swell at normal incidence for vertically propagating ballistic  
 578 waves. (d) Same as in (c) but for the swell wavetrain located away from the earthquake  
 579 epicenter;  $\chi = 60^\circ$ ,  $\psi = 45^\circ$ . A common but otherwise arbitrary normalization of the acoustic  
 580 mode amplitude is used in all panels. The width of the swell wavetrain in the direction of its  
 581 propagation equals 20 swell wavelengths. A nominal value of 1500 m/s is assumed for the phase  
 582 speed  $c_m$  of the acoustic normal mode. Numerical values of the grazing angle of the earthquake-  
 583 generated incident waves refer to the depth, where  $c(z) = c_m$ .

584

585 *T*-phase spectrum and, in particular, the frequency, at which a normal mode is resonantly  
 586 generated, depend on the propagation directions of the ballistic wave and sea swell. It is  
 587 illustrated in Fig. 4b, where the mode amplitude is shown as a function of frequency and grazing  
 588 angle of the ballistic wave, when the sea swell travels at  $45^\circ$  angle to  $\mathbf{q}_{in}$ . In terms of variables  $Y_j$   
 589 introduced in Eq. (34), a resonance occurs when either  $Y_1 = 0$  or  $Y_2 = 0$ . The *T*-phase spectrum  
 590 and resonance frequency for each normal mode also depend on the swell wavelength  $\lambda_{sw} = 2\pi/\mu$ .  
 591 Longer  $\lambda_{sw}$  favors excitation of lower-frequency *T*-waves (Figs. 4c and 4d). The same swell  
 592 wavetrain generates lower-frequency *T*-waves, when it is located around the epicenter (Fig. 4c)  
 593 than away from it (Fig. 4d). If sea swell with the same wavelength and propagation direction is  
 594 present in a large area with dimension comparable to the hypocenter depths, the resonantly

595 excited normal mode is received at different frequencies at the observation points that are located  
 596 at different azimuthal directions from the epicenter. Note also that, according to Eq. (34), any  
 597 swell wavetrain resonantly scatters the ballistic waves of the compressional and shear-wave  
 598 origin in different azimuthal directions and at different frequencies.

599 According to Eqs. (33) and (34), the magnitude squared of the amplitude of  $n$ th normal  
 600 mode generated at scattering by sea swell is

$$601 \quad |Q_n^2| = \frac{32\pi^3 \xi_n}{\xi_n^2 - q_{in2}^2} \left| \frac{\partial P}{\partial z} \right|^2 \left| \Phi \left( q_{in1} - \sqrt{\xi_n^2 - q_{in2}^2} \frac{x - x_0}{|x - x_0|} \right) \right|^2, \quad (35)$$

602 where

$$603 \quad \Phi(q_1) = \frac{iL\sigma_\eta}{2^{3/2}\pi} e^{-iq_1 x_0} \frac{\sin Y}{Y} \Big|_{Y=(q_1 - \mu)L/2}^{Y=(q_1 + \mu)L/2} \quad (36)$$

604 is the one-dimensional wavenumber spectrum of the surface elevation due to swell, Eq. (29),  
 605 viewed as a function of  $x$ . We show below that same result for  $|Q_n^2|$  can be formally obtained  
 606 from the results that have been derived in Sec. II B for random sea surface roughness, if one uses

$$607 \quad S_\eta(q_1, q_2) = \frac{8\pi}{L} |\Phi(q_1)|^2 \delta(q_2) \quad (37)$$

608 for the swell power spectrum in Eq. (17). Here,  $\delta(\cdot)$  denotes the Dirac delta function. It originates  
 609 from the surface elevation being independent of coordinate  $y$ . We assume here that  $\sigma_\eta$  is  
 610 independent of coordinates. We will also assume for simplicity that variations of  $\mathbf{q}_{in}$  and  $\partial P/\partial z$  in  
 611 the incident wave are negligible within the swell wave train.

612 In the integrand in the right side of Eq. (17)  $q_2 = \xi_n (y - y_3) |\mathbf{r} - \mathbf{r}_3|^{-1} - q_{in2}$ , and  $q_2 = 0$   
 613 when  $y_3 = y_{1s}$ , see Eq. (31). Then,  $|\mathbf{r} - \mathbf{r}_3| = |x - x_3| (1 - \xi_n^{-2} q_{in2}^2)^{-1/2}$  and

614 
$$\delta(q_2) = \left| \frac{\partial}{\partial y_3} \left( \xi_n \frac{y - y_3}{|\mathbf{r} - \mathbf{r}_3|} - q_{in2} \right) \right|^{-1} \delta(y_3 - y_{1s}) = \frac{\xi_n^2 |x - x_3|}{(\xi_n^2 - q_{in2}^2)^{3/2}} \delta(y_3 - y_{1s}). \quad (38)$$

615 Inserting Eqs. (37) and (38) in the integrand in Eq. (17) and integrating first over  $y_3$  and then  
 616 over  $x_3$  gives Eq. (35). Note that this derivation of Eq. (35), like Eq. (17), apply in the far field  
 617 with respect to the correlation scale of the sea surface roughness. This is a very significant  
 618 limitation in the case of swell. No such assumption was made in the derivation of Eq. (33), which  
 619 is applicable everywhere outside the swell wavetrain itself.

620 To elucidate the relative significance of wind seas and swell in the abyssal  $T$  wave  
 621 problem, let us compare the acoustic power fluxes  $J_n$  in the normal modes generated at sound  
 622 scattering by two types of ocean surface roughness in the same area  $|x - x_0| \leq L/2$  of the ocean  
 623 surface. For simplicity, we will disregard dependence of  $\partial P/\partial z$  on  $x$  and variation of  $\mathbf{q}_{in}$  and  
 624 wind wave spectrum with coordinates within the area that contributes the most to the scattering.  
 625 Then, Eq. (19) gives

626 
$$\langle J_n \rangle = \frac{\pi^2 E}{2\omega} L \tilde{S}_\eta, \quad E = \left( \frac{1}{\rho} \frac{\partial f_n}{\partial z} \right)_{z=0}^2 \int_{-\infty}^{+\infty} \left| \frac{\partial P}{\partial z} \right|^2 dy \quad (39)$$

627 for wind seas. Here,  $\tilde{S}_\eta$  is the value of the wind wave spectrum at some point within the  
 628 integration domain in Eq. (19). [Equation (39) follows immediately from application of the first  
 629 mean value theorem for integrals to the right side of Eq. (19).]  $\tilde{S}_\eta$  in Eq. (39) can be viewed as a  
 630 weighted average of the spectrum  $S_\eta$  over the horizontal direction of the mode propagation  
 631 within the interval  $|\xi_n - q_{in}| < q < \xi_n + q_{in}$  of wavenumbers  $q$  of wind seas. This interval contains  
 632 all possible  $q = |\xi_n \mathbf{e} - \mathbf{q}_{in}|$  in the integrand in Eq. (19). When the peak  $q = q_p$  of the wave

633 spectrum lies within the interval  $|\xi_n - q_{in}| < q < \xi_n + q_{in}$ ,  $\tilde{S}_\eta \sim q_p^{-2} \sigma_\eta^2 / 2\pi$  according to Eqs. (27)

634 and (28);  $\tilde{S}_\eta$  is small otherwise.

635 For scattering by swell, acoustic power flux in a normal mode can be calculated by

636 integrating the  $x$  component of the acoustic power flux density along the vertical planes  $x - x_0 =$

637 const.  $> L/2$  (toward increasing  $x$ ) and  $x - x_0 = \text{const.} < -L/2$  (toward decreasing  $x$ ). Similar to

638 derivation of Eq. (18), from Eqs. (9) and (33) we find

$$639 J_n = \frac{\xi_n L^2}{16\omega(\xi_n^2 - q_{in2}^2)} \left( \frac{1}{\rho} \frac{\partial f_n}{\partial z} \right)_{z=0}^2 \int_{-\infty}^{+\infty} \sigma_\eta^2 \left( \frac{\sin Y_2}{Y_2} - \frac{\sin Y_1}{Y_1} \right)^2 \left| \frac{\partial P}{\partial z} \right|^2 dy \quad (40)$$

640 for the power flux toward increasing  $x$ .  $Y_{1,2}$  in Eq. (40) are given by Eq. (34) with  $(x - x_0)/|x - x_0|$

641 = 1. The power flux toward decreasing  $x$  is given by the same Eq. (40) but now with  $(x - x_0)/|x -$

642  $x_0| = -1$  in Eq. (34) for  $Y_{1,2}$ .

643 Resonant excitation of  $n$ th acoustic normal mode at scattering by swell occurs when one

644 of the four conditions,  $q_{in1} \pm \sqrt{\xi_n^2 - q_{in2}^2} \pm \mu = 0$ , is met. Then, one of the  $Y_{1,2}$  values in Eq. (40) is

645 zero. Near the resonance frequency [more specifically, as long as  $|Y_{1,2}|$  is either small or  $O(1)$ ],

646 the term in parenthesis in the integrand in the right side of Eq. (40) is  $O(1)$ , and

647  $J_n \sim EL^2 \sigma_\eta^2 / 8\omega \xi_n$ . Note that  $J_n$  is proportional to  $L^2$  due to coherent scattering of sound by the

648 swell wavetrain. Away from the resonance frequencies, when all  $|Y_j| \gg 1$ ,  $J_n$  decreases by the

649 factor of the order of  $\mu^2 L^2 \gg 1$ . For the contribution of wind seas, Eq. (39) gives

650  $\langle J_n \rangle \sim \pi EL \sigma_\eta^2 / 2\omega \xi_n^2$ , when the peak of the wind wave spectrum fully contributes to generation

651 of  $n$ th normal mode. As expected,  $\langle J_n \rangle$  is proportional to the area occupied by surface roughness

652 and, hence, to  $L$  at incoherent scattering of sound by random surface waves.

653            Aside from the roughly estimated numerical factors, in the vicinity of resonance  
 654    frequencies the energy of swell contribution to  $T$ -phase exceeds the maximum contribution of  
 655    wind waves with the same wave height by the factor  $\zeta_n L \gg 1$ . Thus,  $T$  waves due to swell can  
 656    dominate over the wind-wave contribution in narrow frequency bands not only in specific  
 657    directions but also in the azimuthally integrated power flux, even when the local winds are strong  
 658    and the peak of the wind wave spectrum  $q_p \sim \mu$ . However, according to Eq. (40), only a narrow  
 659    vicinity  $\delta f \sim c/L$  of the resonant frequency contributes significantly to the energy of sound  
 660    scattered by swell, and the broadband acoustic power fluxes due to scattering by wind waves and  
 661    swell with the same wave height prove to be comparable.

662

663    **C. Dependence of  $T$ -phase energy and duration on the hypocenter depth**

664    Calculation of the  $T$ -wave spectrum with Eqs. (17) and (33) requires knowledge of the  
 665    distribution of wind speed and sea swell in an area around earthquake epicenter as well as a  
 666    model of the ballistic waves generated by the earthquake. In this section, we use a basic model of  
 667    the seabed and simplified, semi-quantitative versions of the theoretical results for mode  
 668    amplitudes in order to estimate the dimensions of the area of the ocean surface, where  $T$  waves  
 669    are generated, and understand the variation of the abyssal  $T$ -phase duration and energy with the  
 670    depth of earthquake focus. For these estimates, the seabed is modeled as a homogeneous solid  
 671    half-space with the density and elastic parameters of the Earth's crust near the earthquake focus,  
 672    and a compact, directional seismic source is supposed to be located at the focus. For orientation,  
 673     $c_l = 8 \text{ km/s}$ ,  $c_t = 4 \text{ km/s}$ , and  $M = 3$  can serve as representative values of the compressional and  
 674    shear wave speeds and the ratio of the densities of earth's crust and sea water, respectively. The  
 675    hypocenter (focus) of the earthquake is at the point  $(0, 0, H + D)$  at depth  $D$  below the seafloor

676 (Fig. 1). The source will be characterized by the frequency-dependent amplitudes  $A_P$  and  $A_{SV}$  and  
 677 corresponding directional factors  $B_P(\theta, \varphi)$  and  $B_{SV}(\theta, \varphi)$  of compressional (P) and vertically  
 678 polarized shear (SV) waves that are radiated by the earthquake. Horizontally polarized shear  
 679 waves in the crust do not contribute to acoustic field in water.<sup>62</sup> By definition,  $|B_P| \leq 1$  and  $|B_{SV}|$   
 680  $\leq 1$ . When considering the incident waves that are scattered at the ocean surface, we focus on the  
 681 ballistic waves arriving directly from the source and disregard the weaker arrivals, which reach  
 682 the ocean surface and are scattered after previously undergoing surface and bottom reflections.

683 Parameters of the incident acoustic wave, which is scattered by the rough ocean surface,  
 684 affect the wind wave contribution to  $T$ -phase mode amplitudes, Eq. (17), via  $\partial P/\partial z$  and  $\mathbf{q}_{in}$ . The  
 685 amplitude and the angle of incidence of the incident wave vary along the ocean surface. With  
 686 wind waves being independent from the focal depth and the other earthquake properties, after  
 687 averaging over wind speeds and directions, Eq. (17) can be written as follows:

688  $\langle |Q_n^2| \rangle = 4\pi^2 r^{-1} \langle S_\eta \rangle \Psi$ , where

689 
$$\Psi = \int \left| \frac{\partial P(\mathbf{r}_3, 0)}{\partial z} \right|^2 d\mathbf{r}_3. \quad (41)$$

690 The average  $\langle S_\eta \rangle$  of the wind wave spectrum is largely insensitive to the angle of incidence of  
 691 the ballistic waves from the earthquake. For instance, it follows from Eqs. (27) and (28) that  
 692  $\langle S_\eta \rangle \sim q_p^{-2} \sigma_\eta^2 / 2\pi$  and is controlled by the representative wind speed alone, when the peak of the  
 693 wind wave spectrum contributes to  $T$ -phase generation. Hence, the effect of the earthquake  
 694 parameters on  $T$ -phase generation is characterized by the surface integral  $\Psi$  in Eq. (41).

695 Averaging Eq. (40) over the swell wavelength and wave trains' location and propagation  
 696 direction shows that  $\Psi$  Eq. (41) also encapsulates the effect of the incident wave on  $T$ -phase  
 697 generation due to sound scattering by sea swell.

698 For the steep angles, at which ballistic waves from the earthquake propagate in the water  
 699 column, variations of the sound speed in water with depth are insignificant. Sound speed  $c$  and  
 700 density  $\rho$  in water will be assumed constant in the analysis of the ballistic waves. Then, using the  
 701 results for spherical wave transmission through a plane interface of two homogeneous media,<sup>55</sup>  
 702 we find

$$703 \quad \frac{\partial P(\mathbf{r}, 0)}{\partial z} = -\frac{i\omega}{c} A_P B_P(\theta_l, \varphi) T_l(\theta_l) \left[ \frac{\sin \theta_l}{r} \left/ \left( \frac{D}{\cos^3 \theta_l} + \frac{cH}{c_l \cos^3 \theta} \right) \right. \right]^{1/2} \\ \times \exp \left[ i\omega \left( \frac{D}{c_l} \cos \theta_l + \frac{H}{c} \cos \theta \right) - \frac{\alpha_l D}{\cos \theta_l} \right] \quad (42)$$

704 at the point  $\mathbf{r} = r(\cos \varphi, \sin \varphi, 0)$  on the ocean surface. Equation (42) describes the contribution of  
 705 compressional waves in the seabed and is obtained in the ray approximation. Here  $\theta$  and  $\theta_l$  are  
 706 the incidence angles (i.e., the angle ray makes with the  $z$  axis) in the ocean and seabed,  
 707 respectively;  $\alpha_l$  denotes attenuation coefficient of compressional waves, and  $T_l$  is the plane-wave  
 708 transmission coefficient of compressional waves at the seafloor. The incident angles are related  
 709 by Snell's law and can be found from the equations

$$710 \quad c^{-1} \sin \theta = c_l^{-1} \sin \theta_l, \quad r = H \tan \theta + D \tan \theta_l. \quad (43)$$

711 When  $r$  increases from 0 to infinity,  $\theta_l$  increases from 0 to  $\pi/2$  according to Eq. (43), while  $\theta$   
 712 increases from 0 to  $\arcsin(c/c_l)$ . The horizontal wave vector  $\mathbf{q}_{in}$ , which enters Eqs. (16), (17), and  
 713 (33), is  $\mathbf{q}_{in} = \omega c^{-1} \sin \theta \cdot (\cos \varphi, \sin \varphi, 0)$ .

714 Contribution of shear waves in the seabed into  $\partial P/\partial z$  at the ocean surface is given by  
 715 equations similar to Eqs. (42) and (43), except the  $SV$  wave source amplitude  $A_{SV}$ , directional  
 716 factor  $B_{SV}$ , and attenuation coefficient  $\alpha_l$  should be used instead of  $A_P$ ,  $B_P$ , and  $\alpha_l$ . Transmission  
 717 coefficient  $T_l$  of  $SV$  waves replaces  $T_l$  in Eq. (42). In addition, the shear wave speed  $c_t$  and  
 718 incidence angle  $\theta_l$  should be used instead of  $c_l$  and  $\theta_l$  in Eqs. (42) and (43). Since  $c_l > c_t$ , it

719 follows from Eq. (43) that at any  $r > 0$  the ballistic waves due to compressional waves in the  
 720 seabed arrive at the sea surface at steeper angles than the ballistic waves due the shear waves  
 721 radiated by the earthquake.

722 In the case of fluid-fluid interfaces, the transmission coefficient<sup>62</sup>

723 
$$T_l(\theta_l) = 2c \cos \theta_l / (c \cos \theta_l + M c_l \cos \theta). \quad (44)$$

724 At a solid-fluid interface,  $T_l$  and  $T_t$  are given by more cumbersome equations,<sup>62</sup> but, as in Eq.  
 725 (44),  $T_l$  is proportional to  $\cos \theta_l$  and vanishes when  $\theta_l \rightarrow \pi/2$ , while  $T_t$  is proportional to  $\cos \theta_t$  and  
 726 vanishes when  $\theta_t \rightarrow \pi/2$ , see equations (4.2.37)–(4.2.42) in Ref. 62. These properties of the  
 727 transmission coefficients ensure that areas far from the epicenter contribute little to  $T$  wave  
 728 generation. Transmission coefficients  $T_l(\theta_l)$  and  $T_t(\theta_t)$  have  $O(1)$  values for all real  $\theta_l$  and  $\theta_t$ ,  
 729 respectively;  $T_t(0) = 0$  and  $T_l(0)$  is nonzero.

730 Since the ballistic waves originating from compressional and shear waves in the seabed  
 731 have distinct horizontal wave vectors  $\mathbf{q}_{in}$ , the integral  $\Psi$  in Eq. (41) should be calculated  
 732 separately for these incident waves. [The  $\mathbf{q}_{in}$  values are close at near-normal incidence of ballistic  
 733 waves, which occurs in the vicinity  $r \ll H + D$  of the epicenter. However, since  $T_t(0) = 0$ , the  
 734 amplitude is then negligible of the incident wave due to  $SV$  waves in the bottom, and interference  
 735 of the two incident waves has no effect on  $T$  wave generation.] For the compressional wave  
 736 contribution, from Eqs. (41)–(43) we find

737 
$$\Psi_P = \left| \frac{\omega}{c} A_P \right|^2 \int_{-\pi}^{\pi} \frac{d\varphi}{2} \int_0^{\pi/2} |B_P(\theta_l, \varphi) T_l(\theta_l)|^2 \exp\left(-\frac{2\alpha_l D}{\cos \theta_l}\right) \sin 2\theta_l d\theta_l. \quad (45)$$

738 Equation (43) has been used to change the integration variable in Eq. (41) from  $r_3$  to  $\theta_l$ . The  
 739 result for the contribution  $\Psi_{SV}$  of the shear waves in the seabed differs from Eq. (45) by the  
 740 obvious change of notations, which has been discussed above for Eq. (42).

741 Note that Eq. (45) does not contain ocean depth  $H$ . Hypocenter depth  $D$  enters Eq. (45)  
 742 only via the exponential term that describes wave attenuation in the solid bottom. Thus, our  
 743 estimates show that the energy of abyssal  $T$  waves is independent of the ocean depth and is  
 744 insensitive to the hypocenter depth at such frequencies that wave energy dissipation is weak.  
 745 This finding is not restricted to the basic ocean and earth's crust model we consider and, by  
 746 changing the integration variables to ray launch angles, can be extended to stratified seabed as  
 747 long as the ray-theoretical description of the ballistic waves remains applicable.

748 The independence or lack of sensitivity of the abyssal  $T$ -wave energy to  $H$  and  $D$  appears  
 749 counter-intuitive at first. Indeed, according to Eq. (42), amplitudes of the incident waves on the  
 750 ocean surface rapidly decrease with increasing  $H$  and  $D$ . However, the decrease in amplitude is  
 751 compensated by an increase in the ocean surface area that contributes to  $T$  wave generation. For  
 752 instance, if  $H$  and  $D$  are increased by the same factor  $\beta > 1$  and the ray launch angle  $\theta_l$  (or  $\theta_t$ ) is  
 753 kept constant,  $r$  in Eq. (43) increase the same factor  $\beta$ . Incident wave amplitude in Eq. (42) is  
 754 decreased by the factor  $\beta$  as long as the wave dissipation is negligible. The decrease of the  
 755 integrand in the surface integral for  $\Psi$  in Eq. (41) by the factor  $\beta^2$  is exactly compensated by the  
 756 increase in  $d\mathbf{r}_3 = r_3 dr_3 d\varphi$ . This is closely related to the fact that, as long as dissipation is  
 757 negligible, the energy of body waves (as opposed to interface seismo-acoustic waves) reaching  
 758 the ocean surface remains unchanged, when depth of a compact seismic source varies.

759 In addition to  $T$ -phase energy, signal duration is another important characteristic of  $T$   
 760 waves. At distant receivers,  $T$ -phase duration is controlled by the seismic event (rupture)  
 761 duration in the earthquake focus, normal mode dispersion in the oceanic waveguide, and linear  
 762 dimensions of the region, where  $T$  waves are generated. Generation of  $T$  waves due to sound  
 763 scattering occurs with different efficiency at various points on the ocean surface and tends to

764 gradually decrease with distance from the epicenter. Assuming spatially uniform statistics of  
 765 surface gravity waves, the effective radius  $r_g$  of the area around the epicenter, where  
 766 abyssal  $T$  waves are generated, can be estimated as follows [cf. Eq. (41)]:

$$767 \quad r_g = \Psi^{-1} \int r |\partial P(\mathbf{r}, 0) / \partial z|^2 d\mathbf{r}. \quad (46)$$

768 Much like  $\Psi_P$  and  $\Psi_{SV}$  above,  $r_g$  needs to be estimated separately for the incident waves due to  $P$   
 769 and  $SV$  waves in the seabed. In terms of  $r_g$ , the lower bound of the  $T$ -phase duration can be  
 770 roughly estimated as the difference  $2r_g/c$  of acoustic travel times from the opposite margins of  
 771 the region, where  $T$  waves are generated. Similarly,  $r_g/c$  provides an estimate of the rise (onset)  
 772 time of the envelope of the  $T$ -phase waveform.

773 For the ballistic wave due to  $P$  waves in the seabed, from Eqs. (42), (43), and (46) we  
 774 find

$$775 \quad r_g = \left| \frac{\omega}{c} A_P \right|^2 \Psi_P^{-1} \int_{-\pi}^{\pi} d\varphi \int_0^{\pi/2} |B_P(\theta_l, \varphi) T_l(\theta_l)|^2 \left( \frac{H \cos \theta_l}{\sqrt{c_l^2 c^{-2} - \sin^2 \theta_l}} + D \right) \exp\left(-\frac{2\alpha_l D}{\cos \theta_l}\right) \sin^2 \theta_l d\theta_l. \quad (47)$$

776 Derivation of Eq. (47) is quite similar to that of Eq. (45). For the ballistic wave due to  $SV$  waves  
 777 in the seabed, the result follows from Eq. (47) after the previously discussed change in notation.  
 778 The integral in the right side of Eq. (47) and  $\Psi_P$  depend on the source directionality and  
 779 environmental parameters. In the case of an omnidirectional source in a homogeneous medium ( $c$   
 780  $= c_l$ ,  $T_l \equiv 1$ ) without dissipation, Eqs. (45) and (47) give  $r_g = 0.5\pi(H + D)$ . We now show that  $r_g$   
 781 remains of the order of  $H + D$  in the general case, with a possible exception for high frequencies.

782 Note that the integrands in Eqs. (45) and (47) are small, when either  $\sin \theta_l \ll 1$  (because  
 783 of the factors  $\sin 2\theta_l$  and  $\sin^2 \theta_l$ , respectively) or  $\cos \theta_l \ll 1$  (because of the transmission  
 784 coefficient). Hence,  $\tan \theta_l = O(1)$  in the range of  $\theta_l$  that contributes most to the integrals. The  
 785 integrand in Eq. (47) differs from the integrand in Eq. (45) by the factor  $r = H \tan \theta_l + D \tan \theta_l$ ,

786 which is of the order of  $H + D$ , when  $\tan\theta_l = O(1)$ . Thus,  $r_g = O(H + D)$  generally, and our  
 787 estimates indicate longer abyssal  $T$ -phase duration for deeper earthquakes. At sufficiently high  
 788 frequencies, i.e., when waves are strongly dissipated in the seabed over the path of length  $D$ , the  
 789 exponential factor  $\exp(-2\alpha_l D/\cos\theta_l)$  in the integrands of Eqs. (45) and (47) favors small  $\theta_l$ . It  
 790 results in smaller  $r_g$  values at higher  $T$ -wave frequencies than at lower ones.

791 Our results indicate, in agreement with observations,<sup>63–66</sup> that the  $T$ -phase rise (onset)  
 792 time increases with the hypocenter depth  $D$ . Furthermore,  $r_g$  and the rise (onset) time increase  
 793 with the water depth  $H$ . This prediction is opposite to that of the seafloor scattering model by de  
 794 Groot-Hedlin and Orcutt<sup>40</sup> but agrees with the observations analyzed by Williams et al.<sup>2</sup>

795

796 **IV. DISCUSSION**

797 **A. Comparison to other mechanisms of  $T$ -phase generation**

798 For scattering of ballistic waves by rough ocean surface to be a significant mechanism of  $T$ -  
 799 phase generation, the resulting  $T$  waves should have a sufficiently large amplitude. At the very  
 800 least, surface scattering should excite acoustic normal modes much more efficiently than these  
 801 are excited in a horizontally stratified ocean with plane, horizontal boundaries and interfaces.

802 The direct excitation of the  $T$  waves, which have phase and group speeds close to the  
 803 sound speed  $c$  in water, by seismic sources in layered media is very weak because of the  
 804 exponential attenuation of shape functions of the corresponding normal modes in the seabed.<sup>1,3,</sup>

805 <sup>41</sup> For a rough semi-quantitative estimate of the direct excitation, we model the seabed as a  
 806 homogeneous fluid half-space with the sound speed  $c_b > c$ . The seismic wave source is modeled  
 807 as a point monopole acoustic source with  $A = A_0\delta(x)\delta(y)\delta(z - D)$  in Eq. (3). (The conclusions  
 808 remain essentially unchanged for the more complicated dipole or quadrupole sources.) From Eqs.

809 (2), (3), and (6), we find the power flux  $J_n^{(D)} = \omega |A_0|^2 |f_n^2(H+D)|/8$  in the  $n$ th mode, generated in  
 810 a layered medium by a point source at the earthquake focus. Here  $A_0$  is the source amplitude.  
 811 Acoustic pressure is evanescent in the seabed:  $f_n(H+D) = f_n(H) \exp(-\omega D \sqrt{c_n^{-2} - c_b^{-2}})$ , where  
 812  $f_n(H)$  can be estimated from Eq. (7):  $f_n^2(H) \lesssim 2\rho(0)/H$ . When estimating  $J_n^{(D)}$ , one has to use  
 813 shear wave speed, rather than the larger compressional wave speed, for  $c_b$  because evanescent  
 814 shear waves attenuate more slowly below the seafloor and provide stronger coupling of the  
 815 seismic source to the normal modes we consider [i.e., a larger value of  $f_n(H+D)$ ].

816 The resulting expression for the power flux in the normal mode directly excited by the  
 817 seismic source should be compared to the power flux in the same mode excited due to scattering  
 818 of ballistic waves at the rough ocean surface. To estimate the average power flux  $J_n^{(W)}$  due to  
 819 scattering by wind waves on the ocean surface, we employ Eq. (19) and the estimates of the  
 820 spatial average of the surface roughness spectrum  $\langle S_\eta \rangle \sim q_p^{-2} \sigma_\eta^2 / 2\pi = (0.091)^2 \sigma_\eta^4 / 2\pi$  (Secs. III  
 821 A and III C) and the radius of the contributing region on the ocean surface  $r_g \sim H + D$  (Secs. III  
 822 C). For the modal power flux due to scattering by the wind waves, we arrive at the estimate

$$823 J_n^{(W)} \sim \frac{\pi^2 \omega \sigma_\eta^4 (H+D)^2 \sin^2 \chi_n}{2(0.091)^2 \rho(0) c^2(0) H} \left| \frac{\partial P}{\partial z} \right|_{z=0}^2. \quad (48)$$

824 In terms of the amplitude  $A_0$  of omnidirectional point source, for ballistic waves on the ocean  
 825 surface at the epicenter we have  $|\partial P/\partial z| \approx \omega^2 \rho_b T(0) |A_0| / 4\pi (H+D)$ , where  $\rho_b = M\rho(0)$  is the  
 826 seabed density and  $T$  is the transmission coefficient Eq. (44).

827 Combining the above estimates, we find

828

$$F_1 = \frac{J_n^{(W)}}{J_n^{(D)}} \sim \frac{\sin^2 \chi_n}{2(0.091)^2} \left( \frac{\omega \sigma_\eta}{c(0)} \right)^4 \left[ \frac{Mc(H)}{c(H) + Mc_b} \right]^2 \exp\left(2\omega D \sqrt{c_n^{-2} - c_b^{-2}}\right) \quad (49)$$

829 for the ratio of the acoustic power fluxes in  $T$  waves at surface scattering and direct excitation in  
 830 layered waveguide. The ratio  $F_1$  characterizes the relative significance of scattering by wind  
 831 waves compared to the direct excitation. Note that  $F_1$  rapidly increases with sound frequency,  
 832 roughness amplitude, and the earthquake focus depth. With  $\chi_n \cong 0.1$  rad,  $c_n \cong 1500$  m/s, and  $c_b \cong$   
 833 4000 m/s, Eq. (49) predicts that scattering due to wind waves generates  $T$  waves *hundreds of dBs*  
 834 stronger, than the direct excitation, at frequencies as low as 1 Hz and rms surface elevations as  
 835 small as  $\sigma_\eta = 0.3$  m even for rather shallow earthquakes with  $D = 10$  km (or at 2 Hz with even  
 836 smaller  $D = 5$  km). Thus, excitation due to surface scattering of ballistic body waves dominates  
 837 over the direct excitation at all  $T$ -phase frequencies, as expected.

838 In a full-wave, 2-D SPECFEM simulation, Bottero<sup>8</sup> compared  $T$ -phase generation at a  
 839 large-scale bathymetric feature (a six kilometer-long, 12° bottom slope centered on the  
 840 earthquake epicenter) with contributions due to sound scattering by a compact scatterer on the  
 841 ocean surface. The scatterer was intended to roughly represent a large commercial vessel.  
 842 Bottero found that, in his model, the compact surface scatterers (“ships”) were as strong a  $T$ -  
 843 wave source as the downslope conversion on the large bathymetric feature.<sup>8</sup> While the target  
 844 strength of the scatterer in Ref. 8 is much larger than that of actual ships of the same  
 845 dimensions,<sup>67</sup> the full-wave simulation results<sup>8</sup> are extremely valuable as the first rigorous  
 846 comparison of the efficiency of surface scattering and downslope conversion as  $T$ -phase sources.  
 847 By analytic evaluation of  $T$ -phase generation by the compact scatterer considered in Ref. 8 and  
 848 by wind waves, the numerical results<sup>8</sup> have been used to demonstrate<sup>67</sup> that sound scattering by  
 849 wind waves dwarfs the contribution of scattering by ships in 3-D and can generate  $T$  waves at

850 least as efficiently as the presumably dominant<sup>3</sup> generation mechanism of the downslope  
 851 conversion on large bathymetric features.

852 We now provide a direct, semi-quantitative comparison of the energy of the *T* waves that  
 853 are generated in a 3-D ocean by either a large bathymetric feature (a seamount) or sound  
 854 scattering due to gravity waves on the ocean surface. Let an isolated seamount or a small island  
 855 be located at distance *R* from the epicenter. The seamount rises from the otherwise horizontal  
 856 seafloor to the ocean surface. Width of the seamount in the azimuthal direction is *l*. It is small  
 857 compared to *R* and large compared to water depth *H* and acoustic wavelengths in the *T*-wave  
 858 frequency band. The surface of the seamount makes angle  $\gamma$  with the horizontal plane. The  
 859 amplitude of the normal component of the oscillatory velocity of the surface of the seamount  
 860 differs from the velocity amplitude in the ballistic waves at the ocean surface at the epicenter by  
 861 the factor  $w > 0$ , which includes the effects of the geometric spreading and wave attenuation in  
 862 the bottom. For a seamount at range  $R \gg D + H$  from the epicenter, the ratio of the ballistic wave  
 863 amplitudes at the seamount and on the ocean surface at the epicenter  $w \sim \exp(-\alpha R)(H + D)/R$ ,  
 864 where  $\alpha$  stands for the attenuation coefficient of *P* or *S* waves in the seabed.

865 Consider the vertical cross-section of the ocean from its surface to the foot of the  
 866 seamount, where it meets horizontal seafloor. In this cross-section, the horizontal component of  
 867 the particle velocity due to seismic waves of frequency  $\omega$  in the seamount can be estimated as

$$868 v_1 = \frac{2w\sin\gamma}{\omega\rho} \left| \frac{\partial P}{\partial z} \right| \exp \left[ i\Phi(z) + i\sqrt{\omega^2 c^{-2} - \beta^2} (H - z) \cot\gamma \right] \sin\beta z, \text{ where factor } 2i\sin\beta z \text{ accounts}$$

869 for interference of incident and surface reflected acoustic waves with the vertical wavenumber  $\beta$ ,  
 870  $\Phi$  describes variation of the phase of seismic waves along the seamount slope, and  $\partial P/\partial z$  is  
 871 evaluated on the ocean surface at the earthquake's epicenter. Using normal mode orthogonality  
 872 to find modal components of the horizontal velocity, we obtain

873 
$$J_n^{(SM)} = \frac{4w^2 l \sin^2 \gamma}{\omega \xi_n \rho(0)} \left| \frac{\partial P}{\partial z} \right|^2 |U_n^2|, \quad (50)$$

874 
$$U_n = \sqrt{\rho(0)} \int_0^H dz \frac{f_n \sin \beta z}{\rho} \exp \left[ i\Phi + i\sqrt{\omega^2 c^{-2} - \beta^2} (H - z) \cot \gamma \right] \quad (51)$$

875 for acoustic power flux in the  $n$ th mode generated by oscillations of the seamount surface.

876 Here, we disregarded guided acoustic mode penetration into the seabed and used the mode

877 normalization condition Eq. (7).

878 Using the Cauchy–Schwarz inequality and the normalization condition Eq. (7), the upper

879 bound of the integral  $U_n$  Eq. (51) can be estimated as follows:  $|U_n^2| \leq \rho(0) \int_0^H dz \rho^{-1} \sin^2 \beta z \approx H/2$ .

880 A more accurate estimate of  $U_n$ , which accounts for oscillations of the integrand with  $z$ , is

881 
$$|U_n| \sim 2^{-1/2} (\omega^2 c^{-2}(0) - \xi_n^2)^{-1/4} = [2\omega c^{-1}(0) \sin \chi_n]^{-1/2}, \quad (52)$$

882 where  $\chi_n$  has the meaning of grazing angle at the ocean surface. The estimate Eq. (52) refers to

883 modes with significant amplitudes throughout the water column. At higher frequencies, there

884 may be modes with deep turning points, which are very weakly manifested at the ocean surface

885 and the seafloor. These normal modes are not considered here.

886 From Eqs. (48), (50) and (52), we find

887 
$$F_2 = \frac{J_n^{(W)}}{J_n^{(SM)}} \sim \frac{\pi^2 R^2 \exp(2\alpha R) \sin^3 \chi_n}{4(0.091)^2 l H \sin^2 \gamma} \left( \frac{\omega \sigma_\eta}{c(0)} \right)^4 \quad (53)$$

888 for the ratio of the modal power fluxes due to surface scattering and due to the seamount. The

889 ratio increases with the range  $R$ , surface roughness, and, in agreement with observations,<sup>32</sup> with

890  $T$ -wave frequency. It is larger for steeper normal modes (larger  $\chi_n$ ) and smaller for bigger (larger

891  $l$ ) and steeper (larger  $\gamma$ ) seamounts.

892        Depending on environmental parameters and wave frequency,  $F_2$  can be large (i.e.,  
 893        surface scattering dominates) or small (i.e., contribution of surface scattering is negligible)  
 894        compared to unity. Let  $\chi_n = 0.1$ ,  $\gamma = 0.4$ ,  $H = 4$  km, the angular azimuthal dimension of the  
 895        seamount as seen from the epicenter  $L/R = 0.1$ , and the rms surface elevation  $\sigma_\eta = 1$  m. (All  
 896        angles are in radian). To estimate the attenuation coefficient, we use compressional wave speed  
 897        of 8 km/s and  $Q$ -factor of 400.<sup>68, 69</sup> [Attenuation coefficient equals  $27.3 Q_P^{-1}$  dB per wavelength  
 898        in a wave with the quality factor  $Q_P$ .] Then, according to Eq. (53), surface scattering creates  $T$   
 899        waves as strong as those due to a seamount at the range  $R = 400$  km from the epicenter at the  
 900        frequency of about 5.0 Hz, with the surface scattering been the stronger  $T$ -wave source at higher  
 901        frequencies. For  $R = 600$  km, 300 km, 200 km and 100 km, the transition frequency, at which  $F$   
 902        = 1, shifts to about 3.7 Hz, 6.2 Hz, 8.3 Hz, and 13.5 Hz, respectively.

903        Because of their shorter wavelength and smaller quality factors, attenuation in the seabed  
 904        plays a bigger role for shear than compressional waves. Therefore, the ratio  $F_2$  Eq. (53) is larger  
 905        for the shear-wave contributions of the seamount oscillations. Let the shear wave speed and  $Q$ -  
 906        factor be 4 km/s and 200. Then, Eq. (53) gives rather low transition frequencies of 5.6 Hz, 3.3  
 907        Hz, and 2.4 Hz for  $R = 100$  km, 200 km, and 300 km, respectively.

908        It should be emphasized that Eq. (53) provides an estimate, rather than an accurate  
 909        prediction, of the relative significance of the surface scattering and a large topographic feature as  
 910         $T$ -wave sources. On the other hand, our estimates of the contribution of the surface scattering are  
 911        conservative in the sense that sea swell is expected to contribute to  $T$ -wave generation at least as  
 912        much as wind waves (Sec. III B), and that typical values of  $\sigma_\eta$  are larger for most of the world  
 913        ocean<sup>45</sup> than the 1 m assumed in our estimates.

914        Thus, scattering by surface gravity waves is expected to provide a significant contribution  
 915        to  $T$ -phase energy, which is comparable to the contribution due to a downslope conversion on a  
 916        seamount. In addition, being generated around the earthquake epicenter, the surface scattering  
 917        contribution will generally separate from the bathymetric contributions by its arrival time and  
 918        azimuth.

919

920        **B. Extensions of the theory**

921        We have assumed in Secs. II and III that the ocean is range-independent when averaged over  
 922        time-dependent variations due to surface and internal gravity waves. This assumption may be too  
 923        restrictive for the entire propagation path to distant receivers from the abyssal  $T$ -wave generation  
 924        site in the vicinity of the earthquake epicenter. However, the assumption is sufficient to evaluate  
 925        the acoustic energy of abyssal  $T$  waves and its modal distribution in the real ocean. Indeed,  
 926        outside of the relatively small region, where the  $T$  waves are generated, acoustic energy of the  
 927        scattered wave is conserved and is the same in the near field as in the far field, as long as  
 928        acoustic dissipation is negligible. Normal-mode distribution of the  $T$ -phase energy also remains  
 929        unchanged in horizontally inhomogeneous ocean as long as the adiabatic approximation<sup>55</sup> is  
 930        applicable. After the normal mode amplitudes in the  $T$ -phase spectrum are calculated as  
 931        described in Secs. II–III, the field can be readily propagated to long ranges with full account of  
 932        sound absorption using either the adiabatic approximation, the coupled-mode or parabolic-  
 933        equation propagation models.

934        We have focused on contributions of gravity waves in the ocean into  $T$ -phase generation.  
 935        However, the theory of excitation of normal modes of the oceanic waveguide by scattering of  
 936        body waves, as expressed by Eqs. (10), (17), and (26), can be applied to other types of surface

937 and volume scatterers. One important application is to *T*-phase generation at scattering by  
938 volume inhomogeneities within the seabed and roughness of the seafloor and sediment layer  
939 interfaces. This *T*-phase excitation mechanism has been previously considered<sup>41, 42</sup> for coupling  
940 within the discrete spectrum of the seismo-acoustic field. Arguably, the continuous spectrum  
941 (ballistic body waves) make a stronger contribution to *T* wave excitation by bottom scattering  
942 than the directly excited discrete spectrum modes, especially for earthquakes with deeper foci.  
943 Application of the theory developed in this paper would allow one to better constrain the  
944 effective sources of *T* waves on the seafloor and within the seabed (including their spatial  
945 distribution, directionality, and frequency dependence), which were either not related  
946 quantitatively to environmental properties<sup>41</sup> or arbitrarily assigned<sup>22, 39, 40</sup> in previous work.

947 Our finding that the contribution of the ballistic waves scattering by internal gravity  
948 waves into *T*-phase generation is negligible compared to the contribution of the ocean surface  
949 roughness does not necessarily mean that volume scattering in the water column plays no role in  
950 this problem. At long-range propagation, internal waves contribute to coupling of the modes  
951 generated by surface scattering to the modes confined in the SOFAR channel. Furthermore, the  
952 water column contains many different types of inhomogeneities in a wide range of spatial scales.  
953 Scattering of the infrasound generated by air guns from the thermohaline fine structure is  
954 successfully utilized in seismic oceanography to measure physical parameters of the water  
955 column.<sup>70, 71</sup> The frequency band and propagation directions of incident waves that are exploited  
956 in the seismic oceanography experiments<sup>70, 71</sup> are comparable to those of the ballistic infrasound  
957 waves in the ocean due to underwater earthquakes. Thus, seismic oceanography observations  
958 suggest that contributions of the fine structure inhomogeneities into scattering of ballistic waves

959 from the earthquakes are non-negligible. Further research is needed to evaluate this mechanism  
960 of volume scattering and its possible contribution to *T*-phase generation by volume scattering.

961 Evers et al.<sup>13</sup> reported observations of *T* waves in the ocean and their atmospheric  
962 counterpart, guided infrasonic waves in the atmosphere, which were generated by the same  
963 underwater earthquake. Quantitative explanation of the atmospheric observations remains  
964 elusive. We hypothesize that, akin to the abyssal *T*-phases, guided infrasonic waves in the  
965 atmosphere were excited by the scattering of the earthquake-generated body waves on the rough  
966 ocean surface and/or turbulence and internal gravity waves in the atmospheric boundary layer.  
967 Although quantitative analysis of the observations<sup>13</sup> is beyond the scope of this paper, it should  
968 be noted that Eqs. (10), (17), and (26) can be employed to assess the scattering hypothesis. A  
969 distinctive feature of the atmospheric observations by Evers et al.<sup>13</sup> is the low-frequency cutoff in  
970 the spectrum of the earthquake-generated infrasound. Observations of the low-frequency cutoff  
971 are consistent with predictions of Eqs. (10) and (17), as illustrated in Fig. 3 for *T* waves in the  
972 ocean, and provide a strong support for application of the surface scattering hypothesis to  
973 atmospheric manifestations of underwater earthquakes.

974

## 975 VI. CONCLUSION

976 The theory, which is developed in this paper from first principles, offers a quantitative  
977 explanation of ubiquitous observations of efficient generation of *T* waves in the vicinity of the  
978 earthquake epicenter, including the earthquakes under abyssal plains with relatively smooth  
979 seafloor. Wind waves and sea swell on the ocean surface have sufficient amplitudes for *T*-phase  
980 excitation and are rich in the spatial scales needed for Bragg scattering of ballistic body waves  
981 from the earthquake focus into the acoustic normal modes of the oceanic waveguide.

982        Surface scattering favors the acoustic modes, which span most of the water column, and  
983        is consistent with *T*-wave observations by receivers on the seafloor. Observations of low-  
984        frequency cutoff in the *T*-wave spectra find their natural explanation in the spectral properties of  
985        the sea surface roughness. Weak correlation between *T*-phase amplitude and hypocentral depth  
986        follows directly from a ray representation of ballistic waves in horizontally stratified fluid-solid  
987        environment. Ocean surface scattering also offers a simple explanation for observations of the  
988        increase of the *T*-phase onset time with the water depth and hypocentral depth.

989        Contributions of scattering by internal gravity waves into *T*-wave generation are found to  
990        be negligible compared to the contributions of surface gravity waves, among which the sea swell  
991        is expected to be the biggest contributor. Calculation of the wind-wave contribution to the  
992        conversion of the ballistic waves into *T* waves at surface scattering gives the lower bound of the  
993        abyssal *T*-wave energy.

994        Our focus on the gravity wave contributions to *T*-phase generation is not meant to imply  
995        that other, previously identified mechanisms are weak or unimportant. To understand the *T*-wave  
996        excitation, we suggest to consider sound scattering at the ocean surface in addition to the seafloor  
997        scattering and the seismic wave interaction with large bathymetric features. Presumably,  
998        depending on the local conditions, either the ocean surface scattering or the seafloor scattering  
999        may be the dominant mechanism of abyssal *T*-phase generation or the two mechanisms may  
1000        provide comparable contributions. The theory developed in this paper is expected to help in  
1001        identifying the surface scattering contributions in the appropriate *T*-phase data.

1002        Rigorous 3-D, full-wave numerical modeling (e.g., using the SPECFEM approach<sup>12, 17-19</sup>)  
1003        of *T*-phase in an ocean model, which combines a large bathymetric feature with a realistic  
1004        representation of the rough ocean surface, appears to be the logical next step in investigation of

1005 the ocean surface scattering as a *T*-wave generation mechanism and ascertaining its significance.  
1006 Further research is also needed to evaluate the significance of sound scattering by the  
1007 thermohaline fine structure and other water-column inhomogeneities as possible additional  
1008 sources of abyssal *T* waves and to extend the theory to the atmospheric counterpart<sup>13</sup> of the *T*-  
1009 phase phenomenon.

1010

## 1011 **ACKNOWLEDGMENTS**

1012 This work was supported in part by the Office of Naval Research, awards N00014-  
1013 17WX00773 and N00014-20WX01312. Helpful discussions with A. Bottero, P. Cristini, L.  
1014 Evers, and R. A. Stephen are gratefully acknowledged. The author thanks the Associate Editor  
1015 and three anonymous reviewers for their advice on improving the presentation.

1016

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