General Framework for Array Noise Analysis and Noise Performance of a Two-Element Interferometer with a Mutual-Coupling Canceler

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Abstract—This paper investigates the noise performance of a two-element phased array and interferometer containing a recently introduced self-interference canceler, which in the context of this work acts as a mutual-coupling canceler. To this end, a general framework is proposed to permit noise analysis of this network and a large variety of other networks. The framework-based numerical analysis for a two-element phased array shows that the addition of the canceler significantly increases the beam-equivalent noise temperature. For a two-element interferometer used in cosmology, this increase in noise temperature is still acceptable as the sky noise temperature in the 20-to-200-MHz band is high. When used in an interferometer, the canceler provides the ability to null mutual coherence at the interferometer output. The ability to provide matching to reduce the sensitivity of the null in mutual coherence to the phase of the 90° hybrids in the canceler is discussed.

Index Terms—Antenna arrays, noise coupling, coupling canceler

I. INTRODUCTION

NTENNA arrays are increasingly used in modern communication and remote-sensing systems, scientific instruments, military equipment, and biomedical devices [1]–[11]. In comparison to single-antenna transceivers, the design of antenna-array transceivers is challenging due to electromagnetic coupling between antennas [12], [13]. However, antenna arrays offer new opportunities, such as, for example, beam steering and noise shaping [14], [15]. Two-element arrays form an important subset of arrays that are both used on their own and also as tools for studying the performance of larger arrays on a smaller and tractable scale.

This paper investigates the impact of a recently introduced wideband self-interference canceler, which was originally aimed at fullduplex multiple-input multiple-output (MIMO) wireless systems [16] and is similar to an earlier work in [17]. The cancellation is based on replica antennas that are located in isolation within an electromagnetically shielded absorptive box. The replica antennas mimic mutual coupling and scattering of the main antennas, thus allowing passive microwave-based subtraction of coupled components from the signals of interest [16]. While an inspection of the canceler network suggests that this replica-antenna-based canceler is expected to increase the beam-equivalent noise temperature, $T_{\rm rec}$, of such an array, the canceler provides the possibility of decoupling the antennas and thus has potential uses in radio cosmology.

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Radio cosmology instruments typically rely on interferometric measurements to discern 10s to 100s mK spectrally rich perturbations in the Cosmic Microwave Background of 3 K, that is itself much weaker than the Galactic noise of $\sim 10^2$ to $10^4\,\rm K$. Fortunately, while Galactic signals are strong, their smooth spectra permit their separation from cosmological signals. A single-antenna radiometer has already been used for cosmology experiments between 50 and 100 MHz [18]. Interferometric instruments are now being investigated [19] to verify the outcomes of [18].

A cosmology interferometer is expected to consist of two closely spaced antennas connected to receiver circuitry followed by a correlator. The close antenna spacing ($< 0.4\lambda$, where λ is the wavelength) realizes correlation for isotropic noise of the surrounding medium [20] that is subsequently detected by the correlator for measuring the noise temperature of a uniform sky, $T_{\rm sky}$. Design equations for such a system, without the decoupling network, were recently presented in [20], where it was shown that the cross-correlation does indeed contain the desired signal, particularly when the antenna spacing is minimized. However, the mutual coupling of closely spaced antennas results in the noise from each receiver propagating to the output through multiple paths, thereby contaminating the output by the undesirable mutual coherence. To reduce this contamination, a decoupling network may be employed [21]-[26]. While isolators form an intuitive implementation of such a network, their bandwidths are limited to approximately 10% in the 20-to-200-MHz frequency range. Instead, a wideband decoupling network stemming from the work in [16] is considered here. As is shown, this network is able to decouple antennas and thereby null the mutual-coupling-induced correlation in two-element interferometer-based cosmology instruments.

In Section II, we derive key equations that form a general framework for describing the noise performance of an M-element array. This framework is used in the rest of the paper, which is divided into two sections: Section III discusses the beam-equivalent noise temperature, $T_{\rm rec}$, of the two-element array, and Section IV investigates the mutual coherence of the system. Finally, Section V discusses our results, and Section VI concludes the paper.

II. ANTENNA ARRAY WITH THE CANCELER

The noise of antenna arrays directly connected to receivers has been investigated in the past by many groups [27]–[31]. These investigations have adopted different approaches to deriving expressions for noise analysis but they are all specific to their particular systems and are not applicable to the network discussed here. Therefore, we start by developing a general framework that can be adopted for any other network while also avoiding the need for manual error-prone signal-flow-graph analysis. We use this framework to analyze the noise behavior of the system shown in Fig. 1. This block diagram is a receiver-only implementation of the system based on [16]. The main array consists of the M-antenna array, labeled as Antenna #1 to #M, and referred to as "antenna array" in the rest of the paper. The replica array is nominally identical to the main array but placed

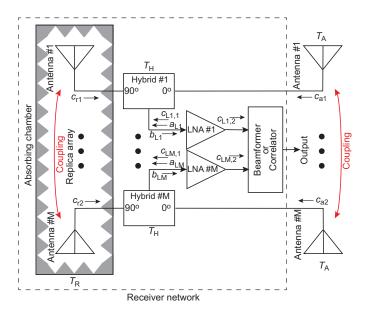


Fig. 1. Block diagram of an M-element array with the mutual-coupling canceler implemented with a replica array and 90° hybrids.

in an absorbing chamber. Unlike [16], two nominally identical 90° hybrids are used. In addition to preserving the cancellation property analyzed in [16], these hybrids have two advantages: the reflection coefficients as seen by the LNAs only depend on the hybrid and do not depend on the antenna array, and maintaining similar performance between two 90° hybrids is easier than between the splitter and 180° hybrid used in the system in [16]. The receiver network, as identified in the figure, consists of the replica array, hybrids, LNAs, and a beamformer or correlator. The noise of the complete system is modeled by noise waves c emanating out of input and output ports of all sub-components. In a conventional implementation, i.e. without the replica array and the hybrids, the noise emanating from the LNA inputs, such as $c_{L1,1}$ to $c_{LM,1}$, couples to adjacent antennas and propagates to the beamformer. Such coupling causes non-zero cross correlation [20]. With the replica array in place, $c_{Li,1}$, where $i = 1 \dots M$, propagate to the beamformer through two paths while experiencing a 180° phase shift, thus providing the possibility of cancellation. While cancellation of coupled $c_{Li,1}$ is therefore possible, the additional absorbing chamber and hybrids undoubtedly contribute additional noise.

A. Key Equations

A set of linear equations that relate traveling waves at the ports of the complete network is

$$\mathbf{b} = \mathbf{S} \left(\mathbf{a} + \mathbf{a}_{\mathrm{s}} \right) + \mathbf{c} \tag{1}$$

with

$$\mathbf{S} = \begin{bmatrix} \mathbf{S}_{PASS} & \mathbf{0} \\ \mathbf{0} & \mathbf{S}_{ACT} \end{bmatrix}$$
 (2)

where in general S_{PASS} and S_{ACT} are S-parameter matrices of the passive and active sub-circuits, respectively; \mathbf{a} is a vector of all incident waves; \mathbf{b} is a vector of all reflected waves; \mathbf{a}_s is a vector of generated waves of sources connected to the network ports; and the noise waves are represented by the vector \mathbf{c} .

The connections between all internally interconnected ports in the system are identified by K, where a = Kb, such that (1) becomes

$$\mathbf{b} = \mathbf{S} \left(\mathbf{K} \mathbf{b} + \mathbf{a}_{\mathbf{s}} \right) + \mathbf{c} \tag{3}$$

and forms the main expression that is used throughout this work to determine the noise and gain of the system.

Whereas (1) to (3) are general, for the particular system in Fig. 1, the elements of (1) and (2) are

$$\mathbf{S}_{PASS} = \begin{bmatrix} \mathbf{S}_{A_{M \times M}} & & \mathbf{0} & \\ & \mathbf{S}_{R_{M \times M}} & & \mathbf{0} & \\ & & \mathbf{S}_{H1_{3 \times 3}} & & \\ & & \mathbf{0} & & \ddots & \\ & & & & \mathbf{S}_{HM_{3 \times 3}} \end{bmatrix}, \quad (4)$$

$$\mathbf{S}_{\text{ACT}} = \begin{bmatrix} \mathbf{S}_{\text{L1}_{2\times2}} & \mathbf{0} \\ & \ddots & \\ \mathbf{0} & \mathbf{S}_{\text{LM}_{2\times2}} \end{bmatrix}, \tag{5}$$

$$\mathbf{a} = \begin{bmatrix} \mathbf{a}_{A} & \mathbf{a}_{R} & \mathbf{a}_{H} & \mathbf{a}_{L} \end{bmatrix}_{1 \times 7M}^{T},$$

$$\mathbf{b} = \begin{bmatrix} \mathbf{b}_{A} & \mathbf{b}_{R} & \mathbf{b}_{H} & \mathbf{b}_{L} \end{bmatrix}_{1 \times 7M}^{T},$$

$$\mathbf{c} = \begin{bmatrix} \mathbf{c}_{A} & \mathbf{c}_{R} & \mathbf{c}_{H1} & \mathbf{c}_{H2} & \mathbf{c}_{L1} & \mathbf{c}_{L2} \end{bmatrix}_{1 \times 7M}^{T}$$

$$= \begin{bmatrix} \mathbf{c}_{PASS} & \mathbf{c}_{ACT} \end{bmatrix}^{T},$$
(6)

where subscripts A, R, H1(2), and L1(2) identify the antenna array, replica array, hybrids #1(#2), and LNAs #1(#2), respectively. The vector \mathbf{c} consists of individual noise contributions from each passive and active parts of the antenna network. For a two-antenna system, $\mathbf{c}_{A} = \begin{bmatrix} c_{a1} & c_{a2} \end{bmatrix}$, $\mathbf{c}_{R} = \begin{bmatrix} c_{r1} & c_{r2} \end{bmatrix}$, $\mathbf{c}_{H1} = \begin{bmatrix} c_{h1,1} & c_{h1,2} & c_{h1,3} \end{bmatrix}$, $\mathbf{c}_{H2} = \begin{bmatrix} c_{h2,1} & c_{h2,2} & c_{h2,3} \end{bmatrix}$, $\mathbf{c}_{L1} = \begin{bmatrix} c_{L1,1} & c_{L1,2} \end{bmatrix}$, and $\mathbf{c}_{L2} = \begin{bmatrix} c_{L2,1} & c_{L2,2} \end{bmatrix}$. If required, matrix \mathbf{S} in (4) can also include S-parameters of the beamformer/correlator to account for any impedance mismatch at the correlator/beamformer input and its gain. Note that (4) assumes a 3-port S-parameter model of hybrids. In the situations when hybrids are 4-port devices requiring a matched load at the extra port, S_H becomes a 4-port matrix in (4). In this case S_{PASS} would also include 1-port S-parameter matrices for each termination.

B. Noise

For noise analysis, the input signals are removed by setting $\mathbf{a}_s = [\mathbf{0}]$, and the noise waves, \mathbf{b}_n , emanating from each port of the network are found from (3) as

$$\mathbf{b}|_{\mathbf{a}_{s}=[\mathbf{0}]} \equiv \mathbf{b}_{n} = (\mathbf{I} - \mathbf{S}\mathbf{K})^{-1} \mathbf{c} = \mathbf{Q}\mathbf{c}, \tag{7}$$

where \mathbf{Q} is introduced for brevity and \mathbf{I} is the identify matrix. The noise-correlation matrix is described by

$$\overline{\mathbf{b}_{n}\mathbf{b}_{n}^{\dagger}} = \mathbf{Q}\overline{\mathbf{c}\mathbf{c}^{\dagger}}\mathbf{Q}^{\dagger} \tag{8}$$

where "†" indicates a Hermitian conjugate operator and, for the passive components, the terms of $\overline{cc^{\dagger}}$ are found from Bosma's theorem [32], [33] as

$$\overline{\mathbf{c}_{\text{PASS}}\mathbf{c}_{\text{PASS}}^{\dagger}} = kB\left(\mathbf{I} - \mathbf{S}_{\text{PASS}}\mathbf{S}_{\text{PASS}}^{\dagger}\right)\mathbf{T}_{\text{PASS}},\tag{9}$$

where \mathbf{T}_{PASS} is the diagonal matrix of the physical temperatures of all passive components. \mathbf{T}_{PASS} is a $5M \times 5M$ matrix for the system in Fig. 1. Assuming that the LNAs are identical (i.e., noise-correlation matrices $\mathbf{c}_{L1}\mathbf{c}_{L1}^{\dagger} = \ldots = \mathbf{c}_{LM}\mathbf{c}_{LM}^{\dagger} = \mathbf{c}_{L}$), the terms in $\mathbf{c}\mathbf{c}^{\dagger}$ due to active components (i.e. $\mathbf{c}_{ACT}\mathbf{c}_{ACT}^{\dagger}$) are found from active component noise parameters [34, p. 54] as

$$\mathbf{c}_{L} = kT_{L}B \begin{bmatrix} \overline{|c_{L,1}|^{2}} & \overline{c_{L,1}c_{L,2}^{*}} \\ \overline{c_{L,1}^{*}c_{L,2}} & \overline{|c_{L,2}|^{2}} \end{bmatrix}$$
(10)

where T_L is the LNA physical temperature¹ and

$$\begin{cases}
\overline{|c_{L,1}|^2} = \frac{T_{min}}{T_0} \left(|S_{L,11}|^2 - 1 \right) + 4N \frac{\left| 1 - S_{L,11} \Gamma_{opt} \right|^2}{1 - \left| \Gamma_{opt} \right|^2} \\
\overline{|c_{L,2}|^2} = |S_{L,21}|^2 \left(\frac{T_{min}}{T_0} + 4N \frac{\left| \Gamma_{opt} \right|^2}{1 - \left| \Gamma_{opt} \right|^2} \right) \\
\overline{c_{L,1} c_{L,2}^*} = \frac{S_{L,11}}{S_{L,21}} \overline{|c_{L,2}|^2} - 4N \frac{S_{L,21}^* \Gamma_{opt}^*}{1 - \left| \Gamma_{opt} \right|^2}.
\end{cases} (11)$$

The noise parameters in (11) are: the minimum noise temperature, T_{\min} ; the Lange invariant, N [35], [36]; and the optimum reflection coefficient for minimum noise, Γ_{opt} .

Having found $\mathbf{b}_{n}\mathbf{b}_{n}^{\dagger}$ in (8), we can calculate beam-equivalent receiver noise temperature, T_{rec} , with [37], [38]

$$T_{\text{rec}} = T_0 \frac{\tilde{\mathbf{w}}^{\dagger} \left. \overline{\mathbf{b}_{\mathbf{n}} \mathbf{b}_{\mathbf{n}}^{\dagger}} \right|_{T_A = 0} \tilde{\mathbf{w}}}{\tilde{\mathbf{w}}^{\dagger} \left. \overline{\mathbf{b}_{\mathbf{n}} \mathbf{b}_{\mathbf{n}}^{\dagger}} \right|_{\text{only } T_A \neq 0} \tilde{\mathbf{w}}}, \tag{12}$$

where $\tilde{\mathbf{w}}$ is the vector of complex gains of the beamformer/correlator in which all terms are zero except for those pertaining to the elements of \mathbf{b} that relate to power flow to the beamformer/correlator. The beam-equivalent noise temperature $T_{\rm rec}$ is defined with reference to the response of the antenna arrays to an isotropic external noise environment at temperature T_0 with which they are in thermal equilibrium.

Further, we find the cross-correlation between outputs i and j in the form of noise temperature by

$$T_{ij} = \frac{1}{kB} \tilde{\mathbf{w}}_i^{\dagger} \left. \overline{\mathbf{b}_n \mathbf{b}_n^{\dagger}} \right|_{\mathbf{a}_s = [\mathbf{0}]} \tilde{\mathbf{w}}_j. \tag{13}$$

C. Gain

To find gain, we set $\mathbf{c} = [\mathbf{0}]$ and set \mathbf{a}_s elements, which correspond to the antenna-array ports, non-zero to apply source waves at the outputs of array antennas in (3) to obtain

$$\mathbf{b} = \mathbf{QSa}_{s}.\tag{14}$$

With **b** from (14), in general the gain of the system in response to an input is

$$G = \frac{\tilde{\mathbf{w}}^{\dagger} \overline{\mathbf{b} \mathbf{b}^{\dagger}} \tilde{\mathbf{w}}}{\delta^{\dagger} \mathbf{a}_{\mathbf{s}} \mathbf{a}_{\delta}^{\dagger} \delta}, \tag{15}$$

where δ is the vector with non-zero terms pertaining to the array inputs. This gain definition considers any or all elements of \mathbf{a}_s as inputs and any or all elements of \mathbf{b} as outputs. In practice, δ and \mathbf{w} would only specify some of their elements, such as those pertaining to antenna ports as inputs and those pertaining to the waves traveling towards the beamformer as outputs. Further, the expression of the correlation gain can be obtained from (15) with

$$G_{ij} = \frac{\tilde{\mathbf{w}}_i^{\dagger} \overline{\mathbf{b} \mathbf{b}^{\dagger} \tilde{\mathbf{w}}_j}}{\delta_k^{\dagger} \overline{\mathbf{a}_s \mathbf{a}_s^{\dagger} \delta_l^{\dagger}}},\tag{16}$$

where subscripts k and $l=1\ldots M$ indicate elements of input vector \mathbf{a}_s that form the input to the system, and subscripts i and $j=1\ldots M$ indicate outputs pertaining to the correlation gain of interest.

III. TWO-ELEMENT PHASED-ARRAY $T_{ m REC}$

Conventionally, increasing array sensitivity requires the development of receivers such that their optimum reflection coefficient, $\Gamma_{\text{opt}} \in \mathbb{C}$, for minimum noise equals the active reflection coefficient, $\Gamma_{\text{act}} \in \mathbb{C}$, of the array [27]–[29], [37], [39]. For a given receiver, this equality can only be achieved for one scan angle, which results

 1 The assumption here is that the amplifier noise is linearly proportional to temperature around T_{0} = 290 K.

in only one beam whose noise can be fully minimized. For other scan angles, Γ_{act} deviates from Γ_{opt} , thereby increasing array noise above its minimum. The scan-dependence of Γ_{act} is due to mutual coupling of antennas in the array. Because of such coupling, some of the receiver noise couples to adjacent antennas and flows to the output through the beamformer, adversely impacting the output noise levels. Past methods of reducing mutual coupling included decoupling networks and low scattering antennas [21]–[25]. The authors of [26] demonstrated that decoupling is necessary for optimum noise matching of antenna arrays. Predictably, the addition of the decoupling network introduces noise, thus potentially negating the advantage of having a decoupled array.

In this section, we investigate the effect of the broadband mutual-coupling canceler on $T_{\rm rec}$ of a two-element antenna array and the possibility of using this canceler for making $\Gamma_{\rm act}$ scan independent.

A. Circuit models

In the following numerical calculations, we assume LNA $T_{\rm min}=25~{\rm K},\,N=0.03,\,{\rm and}\,\,\Gamma_{\rm opt}=0.2\angle100^{\circ}$. Because both $T_{\rm min}$ and N are invariant upon lossless embedding, $\Gamma_{\rm opt}$ can be varied with a matching network without affecting $T_{\rm min}$ and N. We will use this fact to search for the optimum $\Gamma_{\rm opt}$ that minimizes $T_{\rm rec}$. The S-parameters of the LNA are assumed as

$$\mathbf{S}_{L} = \begin{bmatrix} 0.2\angle - 75^{\circ} & 0.01\angle 150^{\circ} \\ 3\angle - 150^{\circ} & 0.3\angle - 100^{\circ} \end{bmatrix}. \tag{17}$$

We based the antenna array S-parameters on a scaled version of a 30-mm spaced two-element array described in [40]:

$$\mathbf{S}_{A} = \mathbf{S}_{R} = \begin{bmatrix} 0.3\angle 100^{\circ} & 0.2\angle - 60^{\circ} \\ 0.2\angle - 60^{\circ} & 0.3\angle 100^{\circ} \end{bmatrix}.$$
 (18)

The ideal hybrid S-parameters are

$$\mathbf{S}_{H} = \frac{1}{\sqrt{2}} \begin{bmatrix} 0 & 1\angle P_{H} & 1\\ 1\angle P_{H} & 0 & 0\\ 1 & 0 & 0 \end{bmatrix}, \tag{19}$$

where $P_{\rm H}=90^{\circ}$, port 1 is the common port and ports 2 and 3 are the 90° and 0° coupled ports. For calculations with a non-ideal hybrid we use the manufacturer-specified S-parameters of a commercial component (Mini-Circuits JSPQW-100A+) at 100 MHz.

B. Numerical results

We first start with an ideal model in which we set LNA $\Gamma_{\rm opt}=0$ and $S_{L,11}=0$, use ideal hybrids, and sweep their phase $P_{\rm H}$ from 0° to 180° while keeping the physical temperatures of the antenna array and the LNA at 290 K and all other temperatures in $T_{\rm PASS}$ at zero. The outcome of this baseline calculation is shown in Fig. 2(a). As can be seen, $P_{\rm H}=90^{\circ}$ does indeed reduce $T_{\rm rec}$ as expected because mutual coupling between antennas is canceled. We also observe that the minimum $T_{\rm rec}$ is slightly higher than $2T_{\rm min}$. This is due to hybrid loss, $L_{\rm H}$, and the impedance mismatch between the antenna and the hybrid. As a result, the variation of $T_{\rm rec}$ with $P_{\rm H}$ is very insignificant. When $S_{\rm L,11}$ is set to its value in (17), the resultant $T_{\rm rec}$ remains unchanged as shown in Fig. 2(b).

Next, we assigned $\Gamma_{\rm opt}=0.2\angle-100^\circ$ while keeping everything else unchanged. Fig. 2(c) shows that $P_{\rm H}=90^\circ$ is no longer optimum. Instead, $P_{\rm H}\approx17.5^\circ$ now realizes low $T_{\rm rec}$. At $P_{\rm H}\approx17.5^\circ$, mutual coupling is not canceled and beamformer dependent $T_{\rm rec}$ results. However, $T_{\rm rec}$ is reduced slightly from that in Figs. 2(a) and (b) demonstrating that the removal of mutual coupling is not necessarily required for improving the overall noise of the receiver under investigation.

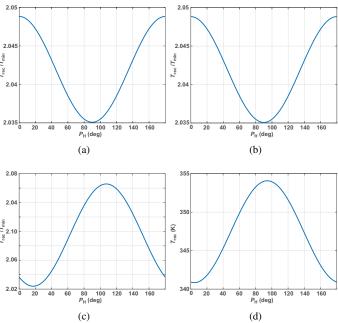


Fig. 2. $T_{\rm rec}$ for the system configured as: a) $\Gamma_{\rm opt}=0$, $S_{\rm L,11}=0$, ideal hybrids, and $T_{\rm H}=T_{\rm R}=0\,{\rm K}$; b) Same as a) but with $S_{\rm L,11}=0.2\angle-75^\circ$; c) Same as b) but $\Gamma_{\rm opt}=0.2\angle-100^\circ$; and d) same as c) but with all temperatures in ${\bf T}_{\rm PASS}$ set to 290 K. Since in subfigure d) $T_{\rm rec}$ is dominated by the ambient temperature rather than $T_{\rm min}$, the y-axis is not normalized.

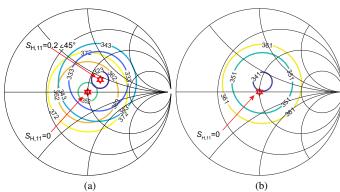


Fig. 3. Contours of constant $T_{\rm rec}$ as functions of $\Gamma_{\rm opt}$ at: a) $P_{\rm H}=90^\circ$ for $S_{\rm H,11}=0$ and $0.2\angle45^\circ$; and b) $P_{\rm H}=4.4^\circ$ from 2(d) and for $S_{\rm H,11}=0$.

Fig. 2(d) plots $T_{\rm rec}$ when temperatures of all components are set to 290 K. In this case, the replica array with the hybrids contributes 290 K in addition to $\sim L_{\rm H} \times T_{\rm min}$ from the LNA. As shown in these calculations, $T_{\rm rec}$ is significantly higher than $T_{\rm min}$, and, therefore, from the array sensitivity point of view, this type of canceler network does not provide the benefits of high sensitivity. Based on this analysis and the general expectation that the replica array contributes $T_{\rm R}$ to $T_{\rm rec}$, $T_{\rm rec}$ cannot be less than $T_{\rm R}$ even if the hybrids are designed such that $L_{\rm H}=0$ dB. Comparing Figs. 2(c) and (d) suggests that in order to maintain $T_{\rm rec}$ between $T_{\rm min}$ and $L_{\rm H}T_{\rm min}$, the passives should be cooled to temperature much lower than 290 K. Reducing $L_{\rm H}$ would further help reducing $T_{\rm rec}$.

In subsequent calculations, we set $P_{\rm H}=90^\circ$ to investigate the independence of optimum $\Gamma_{\rm opt}$ on the values of beamformer coefficients. As shown by the contours of constant noise temperature in Fig. 3(a), the value of $\Gamma_{\rm opt}$ that exhibits the lowest $T_{\rm rec}$ coincides with $S_{\rm H,11}$ for two different values of $S_{\rm H,11}$: $S_{\rm H,11}=0$ and $S_{\rm H,11}=0.2\angle 45^\circ$. Similar calculations show that the lowest $T_{\rm rec}$ always occurs when $\Gamma_{\rm opt}=S_{\rm H,11}$ for any scan angle, i.e.

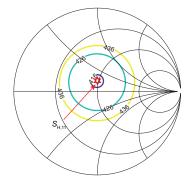


Fig. 4. Contours of constant $T_{\rm rec}$ as function of $\Gamma_{\rm opt}$ of the network with the non-ideal hybrid.

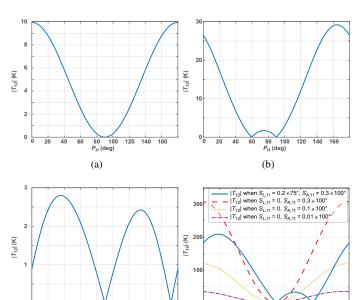


Fig. 5. $|T_{12}|$ for the system configured as: a) $\Gamma_{\rm opt}=0$, $S_{\rm L,11}=0$, ideal hybrids, and $T_{\rm A}=T_{\rm H}=T_{\rm R}=0$ K; b) Same as a) but with $S_{\rm L,11}=0.2\angle-75^\circ$; c) Same as b) but $\Gamma_{\rm opt}=0.2\angle-100^\circ$; and d) same as c) but with all temperatures in $T_{\rm PASS}$ set to 290 K with two different setting for $S_{\rm L,11}$.

 $\Gamma_{\rm act}=S_{\rm H,11}$. It is important to note that while $\Gamma_{\rm act}$ is scan-angle independent, $T_{\rm rec}$ does depend on the scan angle when the reflection coefficients of antennas are not zero. When $P_{\rm H}=4.4^{\circ}$ is selected from Fig. 2(d) to minimize $T_{\rm rec}$, the optimum $\Gamma_{\rm opt}\neq S_{\rm H,11}$, as seen from Fig. 3(b), and the array is no longer decoupled.

In the last step of this section, we replace the S-parameters of the ideal hybrid with the 100-MHz S-parameters of the commercial hybrid (Mini-Circuits JSPQW-100A+) . The outcome of this calculation in Fig. 4 shows the minimum of $T_{\rm rec}$ higher than in Fig. 3 due to the non-ideal hybrid while the network remains decoupled, and $\Gamma_{\rm opt} = S_{\rm H,11}$ still minimizes $T_{\rm rec}$.

IV. TWO-ELEMENT MUTUAL COHERENCE

A. Single-Frequency Two-Element Interferometer

1) Ideal Hybrids: In this section, we reuse the same calculation setup as in Fig. 2 and determine the mutual coherence, in terms of $|T_{12}|$, of the two-element array with circuit components as in Section III-A. Fig. 5 shows the results of this numerical analysis.

As can be seen from the sub-figures, it is possible to adjust the phase of hybrids to reduce cross correlation to zero. As expected, $P_{\rm H}=90^{\circ}$ is optimum, see Fig. 5. However, Fig. 5(b)

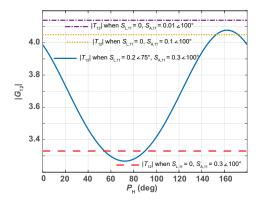


Fig. 6. Magnitude of correlation gain, $|G_{12}|$, (linear units) as a function of $P_{\rm H}$ for different $S_{\rm L,11}$ and $S_{\rm A,11(22)}$.

shows that due to non-zero $S_{L,11}$ a null in the cross-correlation also exists at $P_{\rm H} \approx 60^{\circ}$. The appearance of the second null can be explained with the following gedankenexperiment. Each LNA generates noise at both inputs, I_1 and I_2 , to the cross correlator. The noise contributions of LNA#1 and LNA#2 to I_1 are denoted by uncorrelated $s_{1,1}$ and $s_{2,1}$, respectively, and their contributions to I_2 are denoted by uncorrelated $s_{1,2}$ and $s_{2,2}$. The output of the correlator is $\overline{I_1^* I_2} = (s_{1,1} + s_{2,1})^* (s_{1,2} + s_{2,2})$, which is equivalent to $\overline{I_1^*I_2} = \overline{s_{1,1}^*s_{1,2}} + \overline{s_{1,1}^*s_{2,2}} + \overline{s_{2,1}^*s_{1,2}} + \overline{s_{2,1}^*s_{2,2}}$, where the first term, pertaining to the noise of LNA#1, and the last terms, pertaining to LNA#2, are not zero. The other two terms are cross-correlations of LNA#1 and LNA#2 generated noise and therefore are zero. The locations of the two LNAs relative to I_1 and I_2 are different. It is expected that since for identical LNAs $\overline{s_{1,1}^*s_{1,2}}=\overline{s_{2,1}s_{2,2}^*}$, then $\overline{I_1^*I_2} = 2\Re\left\{\overline{s_{1,1}^*s_{1,2}}\right\}$, and it is possible to null the cross correlation by arranging $P_{\rm H}$ such that $\overline{s_{1,1}^*s_{1,2}}$ is purely imaginary. At $P_{\rm H} \approx 60^\circ$, $\overline{s_{1,1}^* s_{1,2}}$ is in fact purely imaginary causing low cross correlation. While the decoupling at $P_{\rm H}=90^{\circ}$ due to hybrids is unique to the described system, the extra null in mutual coherence at $P_{
m H} pprox 60^{\circ}$ is not unique and, for example, is present in (11) from [20].

 Γ_{opt} is also responsible for moving the location of the additional null as can be seen by comparing Figs. 5(b) and (c). Further, having other system components contribute noise modifies the location of the extra null. The resultant P_{H} is dependent on $S_{\mathrm{L},11}$ and $S_{\mathrm{A},11(22)}$ as demonstrated in Fig. 5(d). As $S_{\mathrm{A},11(22)}$ approaches zero, the maximum of $|T_{12}|$ reduces because noise waves emanating from matched array ports connected to ideal hybrids exhibit no cross-correlation. However, the extra null still exists in this case albeit it is not very visible in the figure. Moreover, the array mutual coupling affects $|T_{12}|$. As the amount of coupling increases the peaks of the cross correlation function increase. Finally, these simulations show that a well-matched LNA with $|S_{\mathrm{L},11}|\approx 0$ avoids sharp nulls in $|T_{12}|$ and creates a smooth trough instead thereby reducing the sensitivity of $|T_{12}|$ to P_{H} .

Based on the same model settings as in Fig. 2(d), Fig. 6 also confirms that the correlation gain, G_{12} , of the network to a correlated input signal is not zero even at $P_{\rm H}$ where receiver-network-noise cross-correlation is nulled. To generate the input signal for this analysis, we make use of the results in [41], which showed that noise correlation matrix of an array due to thermal radiation from the warm scene can be found via Bosma's theorem. Therefore, to find G_{12} , the array is excited with an input \mathbf{a}_s whose noise-correlation matrix $\mathbf{a}_s\mathbf{a}_s^\dagger = \begin{bmatrix} \mathbf{c}_{\rm PASS}\mathbf{c}_{\rm PASS}^\dagger & \mathbf{c}_{\rm L}\mathbf{c}_{\rm L}^\dagger \end{bmatrix}$. We set $T_{\rm R} = T_{\rm H} = T_{\rm L} = 0$. By doing so, only the antenna array excites the network, and other network components are prevented from adding to the output power of the interferometer and inadvertently increasing the computed gain. As the result $\mathbf{a}_s\mathbf{a}_s^\dagger = \begin{bmatrix} \mathbf{c}_{\rm A}\mathbf{c}_{\rm A}^\dagger & \mathbf{0} \end{bmatrix}$, and the correlation of the

input signal received by the array is captured through the cross-correlation terms of $\mathbf{a_s} \mathbf{a_s^{\dagger}}$. While there is some phase dependence due to $S_{\rm L,11}$, the gain G_{12} , as found from (16), remains non-zero and approximately near $|S_{\rm L,21}|^2/L_{\rm H}$ for all $P_{\rm H}$. As $|S_{\rm A,11(22)}|$ decreases the hybrid insertion loss decreases towards its minimum value of 2, thereby increasing G_{12} until its maximum value.

While the correlated noise from the replica array is canceled, $G_{12} \neq 0$ for the correlated output from the antenna array. To explain this, we consider a situation in which both arrays are placed in separate absorbing chambers at thermal equilibrium with all other passive components such that $T_{\rm A} = T_{\rm R} = T_{\rm H}$. In this case, the noises at the hybrid outputs are uncorrelated. If we now remove the absorbing chamber surrounding the antenna array, $\mathbf{c}_{\rm A}\mathbf{c}_{\rm A}^{\dagger}$ changes from $\mathbf{c}_{\rm A}\mathbf{c}_{\rm A}^{\dagger} = kB\left(\mathbf{I} - \mathbf{S}_{\rm A}\mathbf{S}_{\rm A}^{\dagger}\right)\mathbf{T}_{\rm A}$ to $\left[kB\left(\mathbf{I} - \mathbf{S}_{\rm A}\mathbf{S}_{\rm A}^{\dagger}\right)\mathbf{T}_{\rm A} - \mathbf{C}_{\rm ext}\left(T_{\rm A}\right)\right] + \mathbf{C}_{\rm ext}\left(T_{\rm sky}\right)$, where $\mathbf{C}_{\rm ext}$ is the contribution from the external isotropic thermal noise and the first two terms form the noise contribution due to ohmic losses of the antenna array [41]. Consequently, $\overline{\mathbf{b}}\mathbf{b}^{\dagger}$ leading to (16) is written as

$$\overline{\mathbf{b}\mathbf{b}^{\dagger}}\Big|_{T_{L}=0} = \mathbf{Q}\mathbf{S}\overline{\mathbf{a}_{s}}\overline{\mathbf{a}_{s}^{\dagger}}\mathbf{S}^{\dagger}\mathbf{Q}^{\dagger}$$

$$= \mathbf{Q}\mathbf{S}\begin{bmatrix} \mathbf{C}_{\text{ext}}(T_{\text{sky}}) - \mathbf{C}_{\text{ext}}(T_{\text{A}}) & \mathbf{0}_{2M \times 5M} \\ \mathbf{0}_{5M \times 2M} & \mathbf{0}_{5M \times 5M} \end{bmatrix}\mathbf{S}^{\dagger}\mathbf{Q}^{\dagger}$$
(20)

to show that, when $T_{\rm sky} \neq T_{\rm A}$, the output of the networks is non-zero, resulting in a non-zero gain. In conclusion, the accuracy of detecting $T_{\rm sky}$ therefore relies on the accuracy of measuring $T_{\rm A}$.

2) Non-Ideal Hybrids: Next, we used the 100-MHz S-parameters for Mini-Circuits JSPQW-100A+ hybrid in calculations depicted in Fig. 7. As we observed above, the optimum $P_{\rm H}$ depends on system components. Therefore, it is expected that in the final system a means of fine tuning $P_{\rm H}$ is required. In the calculation results depicted in Fig. 7(a) and (b), we include two identical phase shifters for each hybrid with adjustable phases $\Delta P_{\rm H,1} = \Delta P_{\rm H,2} = \Delta P_{\rm H}$ to determine that for the given system $\Delta P_{\rm H} \approx 98^{\circ}$ and $\Delta P_{\rm H} \approx 3^{\circ} {\rm or} \ 21.5^{\circ}$ are required to minimize T_{rec} and $|T_{12}|$, respectively. Figs. 7(c) and (d) show the contours of constant gain magnitude, $|G_{12}|$, in response to a correlated input as functions of $S_{\rm L,11}$ at $\Delta P_{\rm H} \approx 21.5^{\circ}$ and 3° . The maximum $|G_{12}|$ of ~ 3 is near the edges of the Smith chart with $|G_{12}| > 0$ throughout the Smith chart. $|G_{12}|$ can be increased further by increasing LNA $S_{L,21}$. From these figures, it is concluded that the network does not null the correlated input signal while nulling noise generated in the network circuit components. Note that not all values of $S_{L,11}$ in Figs. 7(c) and (d) realize nulls in $|T_{12}|$, and, therefore, not all $S_{L,11}$ identified in these figures may be desirable.

The contour plots of constant $|T_{12}|$ as functions of $\Gamma_{\rm opt}$ and $S_{\rm L,11}$ are presented in Figs. 7(e) and (f) at $\Delta P_{\rm H} \approx 21.5^{\circ}$ and 3° . The two different phase settings result in two distinct contour plots. These calculations show that the phases, rather than magnitudes, of $\Gamma_{\rm opt}$ and $S_{\rm L,11}$ are more significant to minimizing $|T_{12}|$. It is also observed that the decoupling condition of $\Delta P_{\rm H} \approx 3^{\circ}$ is less sensitive to $\Gamma_{\rm opt}$ and $S_{\rm L,11}$ than the noise canceling $\Delta P_{\rm H} \approx 21.5^{\circ}$.

The possibility of adding a matching network to realize coincidental $\Delta P_{\rm H}$ for the two nulls in the mutual coherence is explored in the next experiment whereby identical lossless single-stub matching networks are placed between the hybrids and LNAs. The matching network consists of a transmission line with a maximum length of 1λ that is manipulated in 0.1λ steps and a shunt capacitor stepped by $10~{\rm pF}$ from $1~{\rm pF}$ to $1~{\rm nF}$. Four results of this analysis, when the two nulls in the mutual coherence are located within 4° of each other, are shown in Fig. 8 where the resultant T_{12} as well as the realized $\Gamma_{\rm opt}$ and $S_{\rm L,11}$ are presented.

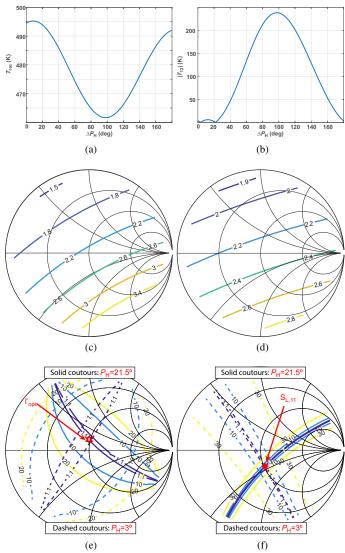


Fig. 7. Simulated a) $T_{\rm rec}$ and b) $|T_{12}|$ as functions of $\Delta P_{\rm H}$ with a non-ideal hybrid. The contour of constant gain $|G_{12}|$ as functions of $S_{\rm L,11}$ at $\Delta P_{\rm H} \approx 21.5^\circ$ and $P_{\rm H} \approx 3^\circ$ are shown in c) and d), respectively. Also shown the contours of constant $|T_{12}|$ at $\Delta P_{\rm H} \approx 21.5^\circ$ and 3° as functions of e) $\Gamma_{\rm opt}$ and f) $S_{\rm L,11}$.

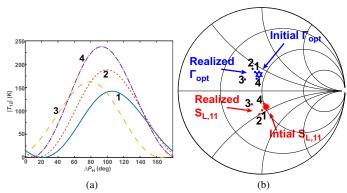


Fig. 8. a) Simulated $|T_{12}|$ with nearly coincidental zeros (< 4° difference) in mutual coherence for lower sensitivity to $\Delta P_{\rm H}$. b) Locations 1 to 4 of realized $\Gamma_{\rm opt}$ s and $S_{\rm L,11}$ s corresponding to curves identified in a) with corresponding numbers. Also the initial $\Gamma_{\rm opt}$ and $S_{\rm L,11}$ of the LNAs are shown.

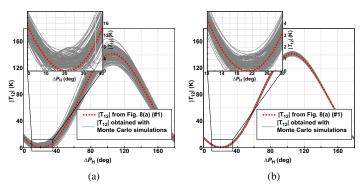


Fig. 9. Monte-Carlo simulation results of $|T_{12}|$ for the system configured as with the matching network realizing Location #1 of $S_{\rm L,11}$ and $\Gamma_{\rm opt}$ in Fig. 8(b). In the Monte-Carlo simulations, all S-parameters and $\Gamma_{\rm opt}$ s are varied so that their maximum variations for each pair of the corresponding network components were limited to (a) 10% and (b) 2%.

3) Monte-Carlo Sensitivity Analysis: Sections IV-A1 and IV-A2 progressed from ideal networks to networks that included practical component models. When non-ideal hybrids were used, two identical phase shifters, with phases $\Delta P_{\rm H.1} = \Delta P_{\rm H.2} = \Delta P_{\rm H}$, were employed to adjust hybrid phases to realize nulls in mutual coherence. Section IV-A2 also investigated the possibility of using a matching network ahead of LNAs so as to make the two nulls nearly coincidental for relaxed sensitivity to hybrid phase errors. While these sections assumed that the arrays of antennas, hybrids, and LNAs were perfectly identical, in practical systems, this assumption will be violated. For example, imperfect replica-array absorber, fabrication inaccuracies, process variations in the LNA circuit components, and so on will cause the loss of identicalness and will require some degree of tuning and possibly sorting of parts to identify the most similar ones. This section analyzes the impact on $|T_{12}|$ of mismatch between the corresponding components of the network. Monte-Carlo simulations are used for this analysis whereby all magnitudes and phases of all S-parameters and the LNA Γ_{opt} s are randomly varied within 5% and 1% of their original values, thereby creating maximum variations of 10% and 2%, respectively, between the corresponding components.

For these simulations, the network realizing Location #1 for Γ_{opt} and $S_{L,11}$ in Fig. 8(b) was selected. The outcome of these simulations in Fig. 9 show that even the 2% variations in S-parameters and Γ_{opt} s may result in the minimum of $|T_{12}|$ exceeding the 10s to 100s mK range required to discern the perturbations in the Cosmic Microwave Background. Although an analysis of the appropriate tuning is not the intent of this work, additional simulations were still performed to investigate the likelihood of tuning the minima in the $|T_{12}|$ curves below 10 mK. In these simulations, the phase shifter phases $\Delta P_{\rm H,1}$ and $\Delta P_{\rm H,2}$ were tuned independently. Fig. 10 shows the results of such analysis, where, for the case of the 10% variations, 34% of simulations returned minima in $|T_{12}|$ below 10 mK. This percentage increases to 52% when component mismatches are limited to 2%. These simulations suggest that two independently adjustable phase shifters and pre-selected components may be sufficient to null mutual coherence even when non-identical components are used.

B. Wideband interferometer

We next return to identical components and investigate the possibility of developing a wideband interferometer, such as, for example, would be used in cosmology instruments, by simulating the network in Fig. 1 in which the S-parameters of the antenna arrays are taken from calculations presented in [20], Mini-Circuits PMA2-43LN+ (biased with 60 mA current) is used for the LNA, and Mini-Circuits JSPQW-100A+ is used for the hybrid. This LNA

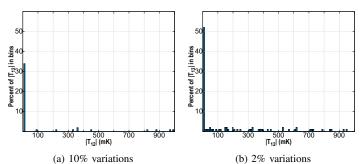


Fig. 10. Histogram of min $\{|T_{12}|\}$ distribution obtained with 100-iteration Monte-Carlo simulations with (a) 10% and (b) 2% variations between S-parameters and the LNA $\Gamma_{\rm opt}$ s of corresponding network components. In these simulations, $\Delta P_{\rm H1}$ and $\Delta P_{\rm H2}$ were independently swept. Each bin is 10 mK wide. Only bins below 1 K are shown.

was selected due to the availability of S-parameters and noise parameters in its datasheet. Still, the noise parameters are only given for frequencies above 800 MHz and the gain is low in the 50-to-100-MHz range. To rectify this, the noise parameters were linearly extrapolated to low frequencies as: $T_{\rm min} \approx T_0 \times 0.06 f_{\rm GHz}$, $N \approx 0.34 - 0.3 \times f_{\rm GHz}$, $Y_{\rm opt} \approx (0.01 f_{\rm GHz} + 0.004) - j0.005 f_{\rm GHz}$, where $Y_{\rm opt} = Z_0^{-1} \left(1 - \Gamma_{\rm opt}\right) / \left(1 + \Gamma_{\rm opt}\right)$. All temperatures in Tpass and $T_{\rm L}$ are set to 290 K. The S-parameters were used directly, as the low gain does not affect the results since the interferometer is assumed noiseless. In a practical implementation, an LNA developed specifically for this application would have sufficient gain to minimize the impact of the interferometer noise on the system performance.

The selected array is wideband and consists of two closely spaced short antenna dipoles over a ground plane for smooth response. The close spacing enables a response to the isotropic noise temperature component of the surrounding medium. As the result, the antennas are highly reflective. The $S_{\rm L,11}$ and $\Gamma_{\rm opt}$ of the selected LNA are also very reflective within the 50-to-100-MHz band. Numerical calculation results in Fig. 7 illustrate that large $|S_{\rm L,11}|$ and $|\Gamma_{\rm opt}|$ are acceptable for minimizing $|T_{12}|$, although $|S_{\rm L,11}|\approx 0$ is preferred for the smooth trough in $|T_{12}|$ (see Fig. 5(d)) and $\Gamma_{\rm opt}=S_{\rm H,11}$ results in the lowest $T_{\rm rec}$.

Fig. 11 shows the results of the numerical analysis in which two identical ideal phase shifters (i.e. $\Delta P_{\rm H,1} = \Delta P_{\rm H,2} = \Delta P_{\rm H}$) are inserted at the 90° terminals of the hybrids. The two phaseshifter phase curves in Fig. 11(a) correspond to the conditions of noise cancellation and decoupling. The required phase shifts to null mutual coherence exhibit monotonic behavior (see Fig. 11(a)). At low frequencies, the coupling is dominated by the hybrid (Fig. 11(b)). The decoupling phase-shifter phase is able to significantly increase the isolation between the LNAs beyond what is afforded by the antenna array itself. At higher frequencies, the array coupling dominates. The network is still able to null mutual coherence but at the cost of larger changes in the required phase shifter phase and higher sensitivity to the phase. When $S_{L,11}$ and Γ_{opt} are reduced to nearly zero (1% of their original values), the optimum $\Delta P_{\rm H}$ for the two conditions come close to each other and remain near-zero over most of the bandwidth, as shown in Fig. 11(c), indicating that with careful engineering of the LNA, wideband nulling in mutual coherence and coincidental nulls for less sensitivity to hybrid phase is likely possible.

V. DISCUSSIONS

The inclusion of a replica array and 90° hybrids is shown to provide a means of nulling the mutual coherence at the output of a two-element interferometer. Two conditions for nulling are identified: decoupling and noise cancellation. Matching networks can be used to reduce the sensitivity of mutual coherence nulls to hybrid phase. The

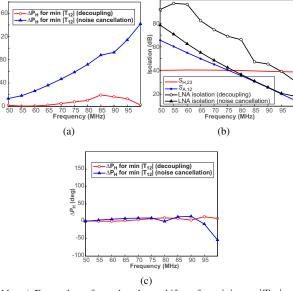


Fig. 11. a) Extra phase from the phase shifters for minimum $|T_{12}|$ versus frequency; b) Simulated isolation between antennas $(S_{\rm A,12})$ and between the coupled ports of the hybrid $(S_{\rm H,23})$, as well as the effective isolation of the LNAs realized with the network when $\Delta P_{\rm H}$ as in (a); and c) Extra phase from the phase shifters for minimum $|T_{12}|$ versus frequency when $S_{\rm L,11}$ and $\Gamma_{\rm opt}$ are near zero.

noise temperature of the interferometer is dominated by the ambient temperature of the replica array unless the latter is cooled. However, when nulling is achieved, the two-element interferometer becomes sensitive to even low-level correlated inputs such as isotropic noise temperature components of the surrounding medium.

For cosmology experiments in the 50-to-100-MHz range, the sky noise temperature $T_{\rm sky}\approx 60\lambda^{2.55}$ [42], which corresponds to $T_{\rm sky}$ between 5800 and 1000 K. These values are larger than $T_{\rm rec}$, the beam-equivalent receiver noise temperature of the interferometer, even without cooling (see Fig. 4), and are much larger in the limiting case when the replica array and the hybrids are cooled to nearly 0 K (see Figs. 2(a)-(c)). Note that for cosmology experiments, $T_{\rm rec}$ affects the integration time, and therefore keeping it below $T_{\rm sky}$ would result in minimum integration penalty. As a result, the proposed mutual-coupling canceler is of potential interest for radio cosmology.

VI. CONCLUSION

This paper investigated the use of a previously-proposed selfinterference canceler on the noise performance of two-element phased arrays and interferometers. In this work, the canceler acts as a mutualcoupling canceler. When used in phased arrays, the canceler makes LNA noise matching independent of the beamformer coefficients but increases the beam-equivalent noise temperature by the physical temperature of the canceler. In an interferometric applications, the addition of the canceler nulls mutual coherence in the output of the interferometer while realizing gain in response to correlated inputs to the antennas. It was shown that the accuracy of detecting the noise temperature of a uniform sky, $T_{\rm sky}$, with the proposed network relies on the accuracy of measuring its physical temperature. With nonideal, realistic, and wideband hybrid and antenna models, the low input reflection coefficient and Γ_{opt} of LNAs realize two nulls in the mutual coherence that are nearly coincidental even without additional matching networks or means of tuning the hybrid phases. However, when there are practical mismatches in network components, Monte-Carlo analysis demonstrated that to lower mutual coherence, there is a need for adjusting phases of each hybrid independently.

The analysis in this work employed a general framework, which can also be used for a large variety of other networks, such as an implementation-ready network of the two-element interferometer that includes tuning networks, additional amplifiers, any cables, and a non-ideal interferometer.

Future work will include thorough analysis of the effect of mismatches between the system components in a wideband system. The proposed framework for analyzing noise will be used to determine the advantageous combinations of the component S-parameters that are both realizable and able to create wideband nulls in mutual coherence. This will lead to identifying the minimum sufficient set of variable/tunable network parameters required to null mutual coherence even with not-exactly-identical components and to ultimately enable the experimental validation of the results of this work.

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