



Bistability of the large-scale dynamics in quasi-two-dimensional turbulence

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In many geophysical and astrophysical flows, suppression of fluctuations along one direction of the flow drives a quasi-two-dimensional upscale flux of kinetic energy, leading to the formation of strong vortex condensates at the largest scales. Recent studies have shown that the transition towards this condensate state is hysteretic, giving rise to a limited bistable range in which both the condensate state as well as the regular three-dimensional state can exist at the same parameter values. In this work, we use direct numerical simulations of thin-layer flow to investigate whether this bistable range survives as the domain size and turbulence intensity are increased. By studying the time scales at which rare transitions occur from one state into the other, we find that the bistable range grows as the box size and/or Reynolds number Re are increased, showing that the bistability is neither a finite-size nor a finite- Re effect. We furthermore predict a cross-over from a bimodal regime at low box size, low Re to a regime of pure hysteresis at high box size, high Re , in which any transition from one state to the other is prohibited at any finite time scale.

Key words: turbulence simulation, turbulent transition

1. Introduction

Ever since the seminal works of Batchelor (1969) and Kraichnan (1967), it has been known that in two-dimensional (2-D) turbulence, contrary to what is observed in three-dimensional (3-D) turbulence, kinetic energy cascades inversely, from the smaller scales at which it is injected to ever larger and larger scales. While the forward cascade that is observed in 3-D turbulence is always arrested once it arrives at the scales at which

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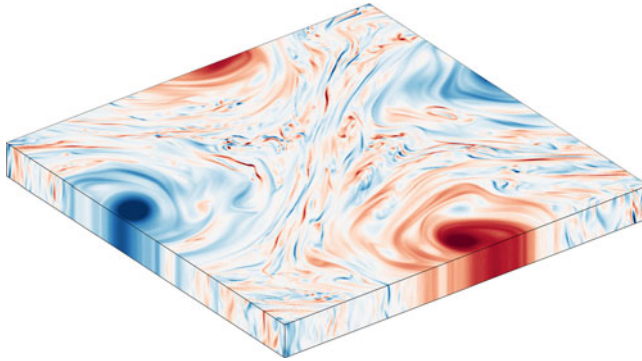


Figure 1. A vortex condensate in thin-layer flow, visualised through a snapshot of vertical vorticity.

viscosity becomes effective in dissipating the kinetic energy, such a stopping mechanism does not always exist at the large scales to saturate the inverse cascade. In that case, kinetic energy piles up at the largest available length scale of the flow system into what is referred to as a ‘condensate’. This condensate typically manifests as a strong vortex structure at the system size, also known as the large-scale vortex; see [figure 1](#).

Even in 3-D flow systems, quasi-2-D dynamics can be observed if fluctuations in one direction are strongly suppressed, allowing an inverse cascade to develop ([Alexakis & Biferale 2018](#)). In forced rotating turbulence ([Smith, Chasnov & Waleffe 1996](#); [Smith & Waleffe 1999](#); [Mininni, Alexakis & Pouquet 2009](#); [Biferale *et al.* 2016](#)), rotating convection ([Julien *et al.* 2012](#); [Favier, Silvers & Proctor 2014](#); [Guervilly, Hughes & Jones 2014](#); [Rubio *et al.* 2014](#)) or rotating stratified turbulence ([Marino *et al.* 2013](#); [Pouquet & Marino 2013](#); [Marino *et al.* 2014](#); [van Kan & Alexakis 2020](#)), such quasi-2-D dynamics develops as a consequence of the Coriolis force, which inhibits the transfer of energy to eddies varying along the axis of rotation. Alternatively, such suppression could occur through, for example, magnetic forces ([Favier *et al.* 2010](#); [Alexakis 2011](#); [Reddy, Kumar & Verma 2014](#); [Baker *et al.* 2018](#)) or plainly through geometric confinement as observed in thin-layer flow ([Celani, Musacchio & Vincenzi 2010](#); [Benavides & Alexakis 2017](#); [Musacchio & Boffetta 2017, 2019](#)). This type of constrained dynamics is of eminent importance to many geo- and astrophysical flow settings, where (a combination of) the aforementioned mechanisms render(s) the flow quasi-2-D. Examples can be found in our oceans ([Scott & Wang 2005](#); [King, Vogelzang & Stoffelen 2015](#)), in the atmosphere ([Nastrom, Gage & Jasperson 1984](#); [Byrne & Zhang 2013](#)) and on gas-giant planets such as Jupiter and Saturn ([Heimpel & Aurnou 2007](#); [Heimpel, Gastine & Wicht 2016](#); [Stellmach *et al.* 2016](#)).

This work focuses on the transition towards the condensate state of such quasi-2-D systems. Remarkably, in spite of the inherently widely different nature of the considered flow systems, recent studies have revealed that all across forced rotating turbulence ([Alexakis 2015](#); [Yokoyama & Takaoka 2017](#); [Seshasayanan & Alexakis 2018](#)), thin-layer turbulence ([van Kan & Alexakis 2019](#)) and even the natural system of rotating convection ([Favier, Guervilly & Knobloch 2019](#); [de Wit *et al.* 2022](#)), the transition into the condensate is discontinuous and shows hysteresis. This gives rise to a limited bistable range in which both the quasi-2-D condensate state and the 3-D flow state can exist at the same parameters.

Since it is now known that this bistability can also survive in natural forcing conditions ([Favier *et al.* 2019](#); [de Wit *et al.* 2022](#)) and the condensate can also form, albeit at more extreme parameters, between realistic no-slip walls ([Aguirre Guzmán *et al.* 2020](#)), we aim

to investigate whether the bistable range in the condensate transition could also survive under parameter conditions that are relevant to geo- and astrophysical flows. Motivated by the remarkable similarities in the condensate transition across the different flow systems, we focus on the conceptually and computationally most basic system of forced thin-layer turbulence. Specifically, we are interested in the dependence on the system size and the strength of turbulent forcing, quantified through the injection-scale Reynolds number Re , in order to investigate whether the bistable range of the condensate transition shrinks or grows as system size and Re are increased. We focus on very moderate values of the system size and Re in order to be able to gather computationally demanding statistics about the bistable range and the rare transitions into and out of the condensate state. By identifying the system size and Re dependence, we can then obtain a first clue as to whether the bistable behaviour could possibly be observed in the limits of large system size and large Re that are relevant to the real-world natural geophysical and astrophysical flow settings.

2. Numerical approach

In order to study thin-layer turbulence, we consider the idealised case of forced incompressible 3-D flow in a triply periodic box of dimensions $L \times L \times H$, where the vertical direction is thin, $H \ll L$. The flow system is identical to that described in van Kan & Alexakis (2019). We consider a Cartesian coordinate system (x, y, z) with unit vectors $(\mathbf{e}_x, \mathbf{e}_y, \mathbf{e}_z)$, where the thin vertical direction is chosen along $\mathbf{e}_z$. The flow $\mathbf{u}(\mathbf{x}, t)$ is then governed by the incompressible forced Navier–Stokes equations

$$\frac{\partial \mathbf{u}}{\partial t} + (\mathbf{u} \cdot \nabla) \mathbf{u} = -\nabla P + \nu \nabla^2 \mathbf{u} + \mathbf{f}, \quad (2.1a)$$

$$\nabla \cdot \mathbf{u} = 0, \quad (2.1b)$$

where $P(\mathbf{x}, t)$ denotes the pressure divided by the constant density and ν represents the kinematic viscosity of the fluid. We consider a stochastic forcing $\mathbf{f}(\mathbf{x}, t)$ that is vertically invariant ($\partial \mathbf{f} / \partial z = 0$) and acts exclusively in the $(\mathbf{e}_x, \mathbf{e}_y)$ plane, i.e. in the two-dimensional, two-component (2D2C) manifold. Furthermore, the forcing is divergence-free and acts only sharply on wavenumber $k_f \equiv 2\pi/\ell$ (specifically, exclusively the modes $(k_x, k_y) = (\pm k_f, 0)$ and $(0, \pm k_f)$ are forced), where its random phase is white noise (delta-correlated) in time. This results in a fixed mean injection rate $\langle \mathbf{u} \cdot \mathbf{f} \rangle = \epsilon$ that is solely prescribed by the forcing amplitude (Novikov 1965). Here, $\langle \rangle$ is used to represent the ensemble average. The choice of forcing is motivated by simplicity and comparability with previous studies. In general, one may consider various 3-D forcing functions (Poujol, van Kan & Alexakis 2020).

The input parameters of the thin-layer flow system are combined to give three dimensionless numbers. We define an injection-scale Reynolds number $Re \equiv (\epsilon \ell^4)^{1/3} / \nu$, the ratio between the forcing scale and the domain height $Q \equiv \ell / H$, and the ratio between the forcing scale and the width of the domain $K \equiv \ell / L$. Finally, we define a forcing time scale $\tau_f \equiv (\ell^2 / \epsilon)^{1/3}$ and energy scale $E_f \equiv (\epsilon \ell)^{2/3}$ that are used to non-dimensionalise the different temporal and energetic quantities reported in this work, respectively.

Equations (2.1a) and (2.1b) are solved numerically in the triply periodic domain using a pseudo-spectral code that is an adapted version of the Geophysical High-Order Suite for Turbulence (GHOST) as introduced by Mininni *et al.* (2011), employing 2/3 dealiasing. In order to investigate the dependence of the condensate transition and its bistable range on Re and the box size, we take the results in van Kan, Nemoto & Alexakis (2019) as a starting point and extend them to smaller and larger Re and K . For each value of K and Re , we vary

$1/K$	Re	Q	$N_x \times N_y \times N_z$	$1/K$	Re	Q	$N_x \times N_y \times N_z$
6	192	[1.53 : 1.70]	$96 \times 96 \times 16$	8	76	[1.00 : 1.13]	$64 \times 64 \times 16$
7	192	[1.53 : 1.69]	$112 \times 112 \times 16$	8	131	[1.29 : 1.48]	$96 \times 96 \times 16$
8	192	[1.44 : 1.81]	$128 \times 128 \times 16$	8	192	[1.44 : 1.81]	$128 \times 128 \times 16$
9	192	[1.44 : 1.73]	$144 \times 144 \times 16$	8	329	[1.74 : 2.05]	$192 \times 192 \times 16$

Table 1. The different series of input parameters used in this work for varying box size (left columns) and varying Re (right columns).

the thinness of the fluid layer Q as the principal control parameter in close vicinity to the condensate transition. An overview of the full set of input parameters that are considered in this work is provided in [table 1](#).

Resolutions $N_x \times N_y \times N_z$ are chosen such that we maintain the same ratio between the grid spacing $L/N_{x,y}$ and Kolmogorov length $\eta = (\nu^3/\epsilon)^{1/4}$ as used in van Kan *et al.* (2019) of $L/N_{x,y} \approx 3.2\eta$ in the horizontal directions, and we resolve finer than that in the vertical direction. For the vertical, we ensure that we keep 16 grid cells in order to maintain sufficient degrees of freedom in the thin direction.

For each unique set of input parameters (Q, K, Re), at least 40 independent runs are carried out by using a different random seed for the stochastic forcing in order to obtain statistics about the rare transitions from one state into the other. The combination of this need for a multitude of independent runs with the fact that the large-scale dynamics is very slow compared to the background smaller-scale 3-D dynamics renders only the rather moderate parameters that are considered here computationally accessible. The simulations in this work comprise more than ~ 1.0 million CPU hours.

The main diagnostic quantity that we use here to probe the strength of the condensate is the 2-D large-scale energy E_{ls} , defined from the Fourier components $\hat{\mathbf{u}}(\mathbf{k})$ of the flow as

$$E_{ls} = \frac{1}{2} \sum_{\substack{\mathbf{k} \\ |\mathbf{k}| \leq k_{max} \\ k_z=0}} [|\hat{\mathbf{u}}(\mathbf{k}) \cdot \mathbf{e}_x|^2 + |\hat{\mathbf{u}}(\mathbf{k}) \cdot \mathbf{e}_y|^2], \quad (2.2)$$

where the cut-off wavenumber $k_{max} = \sqrt{2}(2\pi/L)$.

3. Bistability, bimodality and rare transitions

Bistability and bimodality are often met in dynamical systems. Here we refer to ‘bistability’ as the presence of two independent stable attractors, coexisting for the same value of parameters; whilst by ‘bimodality’ we refer to the case in which two attractors are linked by some trajectories such that, when followed, the system jumps from one attractor to the other and *vice versa*. In the former case, a hysteresis loop can exist when one of the parameters of the system is continuously varied. In an inherently fluctuating dynamical system like the one at hand, however, the precise extent of the bistable range in the hysteresis loop can be ambiguous, as rare sudden transitions from one hysteretic branch into the other may exist at very long time scales near both ends of the hysteresis loop. This raises the question as to how we can unambiguously define the precise extent of the bistable range in the hysteretic transition.

Earlier works that studied the bistable range in the quasi-2-D condensate transition would typically simulate up to a certain time scale that is constrained by computational

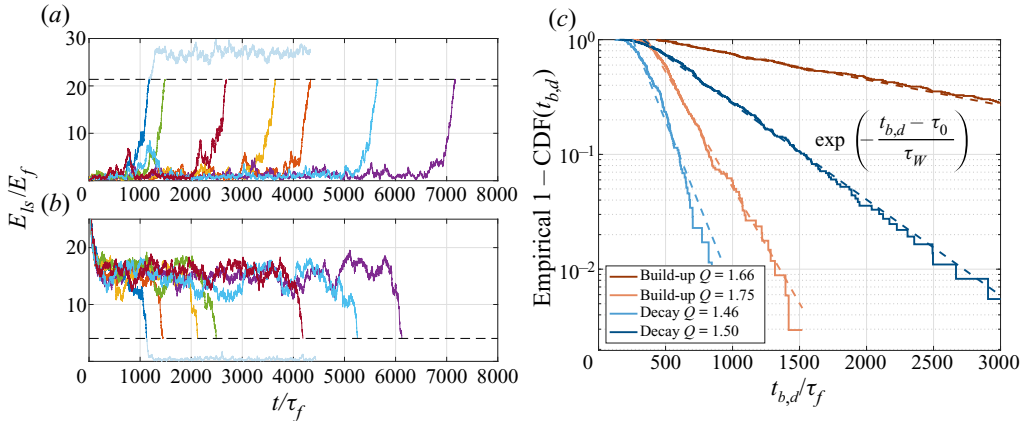


Figure 2. (a,b) Examples of time series of different realisations for (a) build-up at $Q = 1.67$ and (b) decay at $Q = 1.49$ (for $1/K = 9$ and $Re = 192$). Horizontal dashed lines represent the threshold energy at which the build-up time t_b or decay time t_d is defined and the simulation is terminated. The first blue run is continued for demonstration purposes. (c) Examples of distributions of $t_{b,d}$, shown through the empirical CDF (at $1/K = 8$ and $Re = 192$). Dashed lines represent fits of the exponential distribution.

limits and call a state ‘stable’ if no further transitions are observed (Yokoyama & Takaoka 2017; Favier *et al.* 2019; van Kan & Alexakis 2019; de Wit *et al.* 2022). However, in view of the rare transitions, this amounts to an in principle arbitrary cut-off at a certain time scale, neglecting any possible transitions occurring at larger time scales. We refer to this as *finite-time hysteresis*.

However, in the case of *pure hysteresis*, in the strict sense, all transitions from one state into the other are prohibited at any finite time scale, such that the system is absolutely bistable. This would require the time scales of the rare transitions to diverge at a certain asymptote. The existence of such asymptotes is the principal assumption in this work. As we will show in § 4, we can define such asymptotes based on the scaling of the time scales of the rare transitions that we observe, allowing us to study how these asymptotes shift as we vary the box size and Re in order to get a completely time-scale-independent and unambiguous method for comparing our results at these different parameters.

To analyse the rare transitions between the two states, we separately consider build-up events from the 3-D state into the condensate state and decay events *vice versa*. The build-up events are studied by initialising the simulations with a tiny perturbation onto a state of no flow and continuing until the condensate state is reached. For the decay events, we initialise the simulation with a snapshot from a condensate state at higher Q and we continue the run until the condensate has decayed and the 3-D state is obtained.

Figures 2(a) and 2(b) show examples of different realisations of such rare transitions for one choice of parameters. The works of van Kan *et al.* (2019) and de Wit *et al.* (2022) have revealed that the waiting time that is spent until these sudden transitions from one state into the other commence is exponentially distributed, signifying that the transition process is memoryless. The typical mean waiting time τ_W can be obtained by defining representative thresholds in the large-scale energy and analysing the distribution of times $t_{b,d}$ after which these thresholds are crossed. Examples of the obtained empirical cumulative distributions are depicted in figure 2(c), showing that it closely follows the aforementioned exponential distribution. We can then obtain τ_W by fitting the empirical

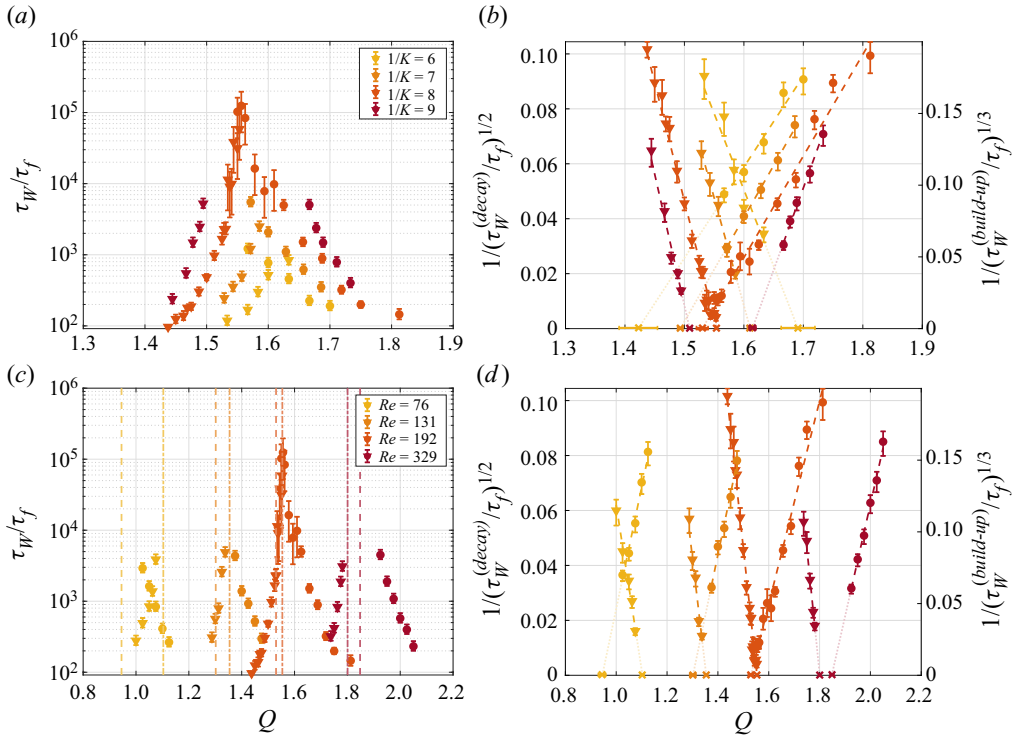


Figure 3. (a,c) Waiting times τ_W for build-up (circles) and decay (triangles) events and (b,d) their power-law transformation for varying box size $1/K$ at $Re = 192$ (a,b) and varying Re at $1/K = 8$ (c,d). Crosses on the horizontal axis in (b,d) denote the estimates for the asymptotes Q_0 . These asymptotes are also depicted in panel (c) by the vertical lines for build-up (dashed) and decay (dashed-dotted), but are omitted in panel (a) for readability.

cumulative distribution function (CDF) with a (shifted) exponential as

$$\text{CDF}(t_{b,d}) = 1 - \exp\left(-\frac{t_{b,d} - \tau_0}{\tau_W}\right). \quad (3.1)$$

This process can be repeated to obtain the waiting time scales τ_W for build-up events and decay events at different values of Q , varying it across the full extent of the hysteretic transition. This results in a series of typical waiting times τ_W as a function of Q , which can in turn be repeated for different box sizes and Re .

4. Time-scale statistics

The results for the series of transition time scales for varying box size and varying Re are provided in figure 3(a,c). As the transition is approached, τ_W increases faster than exponentially (see van Kan *et al.* 2019). This implies either that τ_W increases in a non-diverging super-exponential fashion – e.g. $\tau_W \propto \exp[\exp(Q)]$, as is typical for certain transitions controlled by extreme events (Goldenfeld, Guttenberg & Gioia 2010; Nemoto & Alexakis 2018, 2021; Gomé, Tuckerman & Barkley 2022) – or that it diverges at some critical value Q_0 . To determine which of the two holds for the present system is beyond the scope of this work. We will thus assume that the latter case applies, although one could alternatively interpret Q_0 as the value at which super-exponential behaviour starts in the

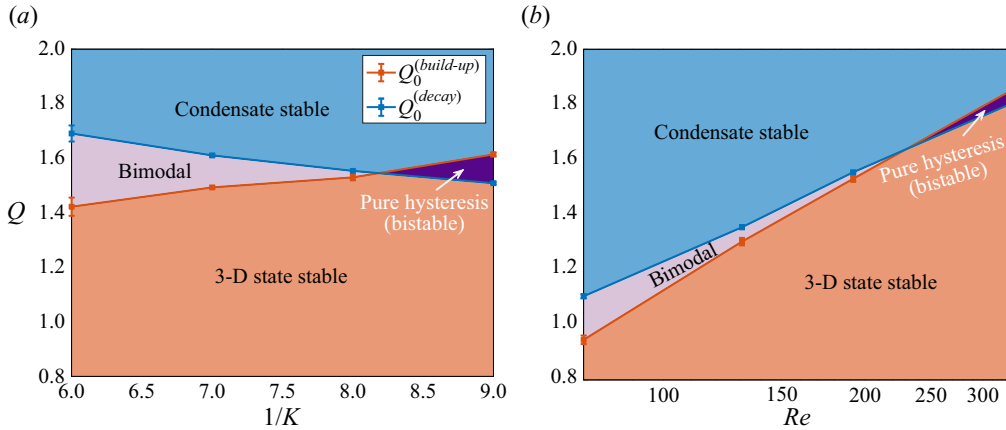


Figure 4. Phase diagrams of the condensate for (a) varying box size at $Re = 192$ and (b) varying Re at $1/K = 8$. Red and blue lines denote the asymptotes of waiting times for build-up and decay events, respectively.

former case. To determine Q_0 , we fit the transition time to a power-law divergence

$$\tau_W \propto \frac{1}{|Q - Q_0|^p}. \quad (4.1)$$

By plotting $1/\tau_W^{1/p}$ as a function of Q , we can then obtain Q_0 from a linear fit to the data (figure 3b,d). Empirically, we find that $p^{(build-up)} = 3$ and $p^{(decay)} = 2$ result in a satisfactory linearisation of our data. These asymptotes then predict the location in Q where the transition time becomes infinite, such that we can say that, beyond the asymptote, the transition cannot occur at any finite time scale.

Comparing the results at different box sizes and different Re , we find first of all that the transition is observed in a similar range of Q for the varying box sizes, while it clearly shifts as Re is varied, in agreement with the observations in van Kan & Alexakis (2019). More importantly, we observe that the branches of build-up and decay times move further apart as Re and box size are increased: the build-up branch moves to larger Q , while the decay branch moves to (relatively) smaller Q . The branches cross for the runs with $Re \leq 192$ or for $K \leq 8$, such that, at small box size and/or small Re , a *bimodal* range of Q exists for $Q_0^{(build-up)} < Q < Q_0^{(decay)}$ where the build-up and decay time scales are simultaneously finite (and, in fact, computationally accessible). Hence, in this range, the flow continually transitions back and forth between the 3-D state and the condensate state.

Conversely, for the largest box size and largest Re that we consider, the branches of the build-up and decay branches never cross, as the asymptotes reside on opposite ends. This indicates a profoundly different regime, where the decay times have diverged before the build-up times become finite, such that, in the range $Q_0^{(decay)} < Q < Q_0^{(build-up)}$, both states are absolutely stable, as no transition from one state into the other can occur in any finite time. This corresponds to a regime of *pure hysteresis* in which the system is (absolutely) bistable. Indeed, it is this range that is arguably the most unambiguous time-scale-independent definition of the bistable range of the system. These results are summarised in figure 4. It is thus evident from our results that the bistable range grows as the box size and/or Re are independently increased.

We argue that the strengthening of the bistability for large box sizes and Re is intuitive from the increase of the condensate energy level (compared to the 3-D-state energy) as Re or $1/K$ is increased, making it harder to jump from one state into the other. Our results thus indicate that bistability is not a finite-size, finite- Re effect, and such states could possibly be found in the geophysical limit where both Re and domain size are large.

5. Conclusions and outlook

In this work, we have demonstrated that the bistability observed in the transition to the quasi-2-D condensate state persists as the box size and Re are increased. By studying the time scales at which rare transitions from one state into the other occur, we quantified the precise extent of the bistable range. Fitting the mean time scales of these transitions with a diverging power law, we measured the locations of the asymptotes beyond which the transition is prohibited at any finite time scale. Since these asymptotes show a cross-over as we vary the box size and/or Re , this predicts a profound regime change from a bimodal regime at small box size and/or Re to a regime of pure hysteresis at large box size and/or Re , as summarised in [figure 4](#). Since we find that the branches of time scales of build-up transitions into the condensate and decay transitions out of the condensate at both ends of the bistable range only separate further and further as the box size and/or Re is increased, we conclude that this bistability is not a finite-size or finite- Re effect, but that the bistable range grows as we progress towards the limit of increasing system size and/or Re .

We remark that the method proposed here for quantifying the precise extent of the bistable range in a hysteretic transition using the time scales of rare transitions is entirely general, and a similar procedure can be followed in the context of any other fluctuating hysteretic dynamical system within or beyond fluid dynamics. However, we must note that the motivation of (4.1) is *ad hoc* here. Although the agreement with our data as shown in [figure 3](#) is satisfactory, a more fundamental physical motivation, supported by a larger dynamic range of parameters, would be needed to rigorously prove the validity of (4.1) as well as our choice of exponents. Indeed, although the waiting time increases faster than exponentially, the existence of an asymptote for τ_W in the first place is ultimately an assumption in itself and the possibility of the relation being any other super-exponential relation without divergence can in principle not be ruled out by numerical simulations alone. Nonetheless, while the existence of pure hysteresis certainly constrains the underlying physical mechanism of the transition, one may argue that it does not hold immediate implications in geophysical practice whether the time scales of transitions are strictly infinite or merely beyond any practical finite time scale. Moreover, we argue that the satisfactory agreement of (4.1) with our data in itself does convincingly prove our central result: that the bistable range is not an effect of finite box size or finite Re , but that it grows with increasing system size and/or Re . Indeed, this holds either in the strict terms of absolute bistability, or in terms of exceeding a certain finite super-exponential time scale.

While recent investigations have started to unveil different aspects of this peculiar type of transition between turbulent flow states, much of the underlying physical mechanism still remains poorly understood. In particular, which specific physical events trigger the flow to commence the transition – for example, either a series of vortex merging events, or rare fluctuations directly at the largest scale – remains an open question. Answering such questions would contribute greatly to our understanding of this flow phenomenon, and our work may act as a numerical inspiration as well as a quantitative benchmark to such theoretical studies. Specifically, understanding the physical mechanism behind the

transition may motivate the theoretical validity of relation (4.1), which we have motivated only empirically here.

The high computational demands of the presented analysis limit the range of accessible parameters in this work to rather moderate values at small grids. It is known from van Kan & Alexakis (2019) that the observed increase of the critical values Q_0 with Re in the examined range will eventually saturate at very large Re . The persistence of bistability at this asymptotic regime thus needs to be formally validated. Furthermore, we also limited ourselves to the simplest case of a thin layer forced by a 2-D body force. More natural 3-D forcing as well as effects such as rotation and stratification would also be needed to make contact with geophysical flows that would require a larger set of simulations to cover the high-dimensional parameter space, posing yet higher computational demands.

A promising solution may lie in the application of rare-event algorithms (C  rou & Guyader 2007; Lestang *et al.* 2018). Such algorithms are more efficient in probing rare transitions and have been successfully applied in various other flow contexts (Rolland 2018; Bouchet, Rolland & Simonnet 2019; Gom   *et al.* 2022). By progressing further towards the extreme geophysical conditions, we may, for example, investigate whether the growth of the bistable range of the condensate transition saturates at some point, which cannot be studied from the moderate parameters considered in our work. Finally, we remark that it also seems attractive now to study the condensate transition and its bistable behaviour from experiments in which more extreme parameters may be more easily accessible, or perhaps even from observations in real-world geophysical or astrophysical flows.

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