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## 3 Main Manuscript for

4 Extreme hydroxyl amounts generated by thunderstorm-induced  
5 corona on grounded metal objects

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28 This PDF file includes:

29       Main Text

30       Figures 1 to 5

31       Abstract

32 Atmospheric electrical discharges are now known to generate unexpectedly large amounts of the  
33 atmosphere's primary oxidant, hydroxyl (OH) in thunderstorm anvils, where electrical discharges  
34 are caused by atmospheric charge separation. The question is "Do other electrical discharges also  
35 generate large amounts of oxidants?" In this paper, we demonstrate that corona formed on  
36 grounded metal objects under thunderstorms produce extreme amounts of hydroxyl (OH),  
37 hydroperoxyl (HO<sub>2</sub>), and ozone (O<sub>3</sub>). Hundreds of parts per trillion to parts per billion of OH and  
38 HO<sub>2</sub> were measured during seven thunderstorms that passed over the rooftop site during an air  
39 quality study in Houston, TX in summer 2006. A combination of analysis of these field results and  
40 laboratory experiments shows that these extreme oxidant amounts were generated by corona on  
41 the inlet of the OH-measuring instrument and that corona are easier to generate on lightning rods  
42 than on the inlet. In the laboratory, increasing the electric field increased OH, HO<sub>2</sub>, and O<sub>3</sub>, with 14  
43 times more O<sub>3</sub> generated than OH and HO<sub>2</sub>, which were equal. Calculations show that corona on  
44 lightning rods can annually generate OH that is 10 to 100 times ambient amounts within centimeters  
45 of the lightning rod, and on high-voltage electrical power lines can generate OH that is 500 times  
46 ambient a meter away from the corona. Contrary to current thinking, previously unrecognized  
47 corona-generated OH, not corona-generated UV radiation, mostly likely initiates premature  
48 degradation of high-voltage polymer insulators.

49       Significance Statement

50 Corona generated on grounded metal objects by thunderstorms or on high-voltage power  
51 transmission lines are known to produce ozone. However, during an air quality study in Houston,  
52 TX, thunderstorms produced extreme amounts of hydroxyl (OH), which is the atmosphere's primary  
53 oxidant, and hydroperoxyl (HO<sub>2</sub>). Laboratory studies confirm that this extreme OH and HO<sub>2</sub> is  
54 generated by electrical discharges including corona. Corona generated 28 ozone molecules and  
55 one HO<sub>2</sub> molecule for each OH molecule. Calculations show that corona on grounded metal objects  
56 under thunderstorms and on high-voltage electrical power transmission lines can generate enough  
57 OH near the corona to accelerate the air chemistry and prematurely age materials, including  
58 insulators for the high-voltage transmission lines.

59

60

61       **Main Text**

62

63       **Introduction**

64

65 Underappreciated sources of hydroxyl (OH), the atmosphere's primary oxidant, have been recently  
66 discovered in thundercloud anvil clouds (1), adding to the existing known OH sources. Atmospheric  
67 OH is important because, when chemicals are emitted into Earth's atmosphere, OH reacts to  
68 remove many of them. For example, methane, a potent greenhouse gas, is removed from the  
69 atmosphere primarily by reaction with OH. However, the subsequent sequence of reactions for  
70 many of the chemicals emitted into the atmosphere produce the pollutants ozone (O<sub>3</sub>) and small  
71 particles, both of which are harmful to human health and important factors in climate change. Thus,  
72 any atmospheric phenomena that influences the amount of atmospheric OH also contributes  
73 simultaneously to air quality and climate change.

74

75 Extreme amounts of OH and another reactive atmospheric constituent hydroperoxyl ( $\text{HO}_2$ ) are  
76 generated not just by lightning flashes but also by weaker electrical discharges, even those too  
77 weak to be seen or heard. This conclusion came from a combination of airborne observations in  
78 thunderstorm anvil clouds during the Deep Convective Clouds and Chemistry (DC3) airborne study  
79 in 2012 and subsequent laboratory experiments (1,2). We call  $\text{HO}_x$  generated by any type of  
80 atmospheric electrical discharge  $\text{LHO}_x$ . Models of electrical discharges have long calculated that  
81 lightning flashes and corona discharges produce OH and  $\text{HO}_2$  (together called  $\text{HO}_x$ ), but these  
82 modeled OH lifetimes were too short for OH to have atmospheric relevance (3-5).  
83

84 However, contrary to expectations from the models, the OH and  $\text{HO}_2$  observed in the anvil clouds  
85 and the laboratory live 100 times longer than expected. Presumably, either water vapor is simply  
86 dissociated by corona ultraviolet radiation into OH and H, which quickly adds an  $\text{O}_2$  to make  $\text{HO}_2$ ,  
87 or neutral and ion chemistry accomplishes the same feat (6-8). OH generated in thunderstorm  
88 anvils could be responsible for as much as 2 to 16% of total global atmospheric oxidation (1).  
89

90 Corona discharges, which are stronger than subvisible discharges but weaker than lightning  
91 flashes, have been observed for millennia on pointed objects, such as sailing ship masts, when  
92 thunderstorms were overhead. It is also known that corona on high-voltage electrodes generate  
93 OH under certain conditions (9), but the amount of OH or  $\text{HO}_2$  generated was not quantified. If  
94 extreme OH and  $\text{HO}_2$  can be generated by subvisible discharges in thunderstorm anvils and by  
95 corona on high-voltage electrodes, then they should also be generated by coronas that occur when  
96 thunderstorms pass over grounded metal objects, such as lightning rods, or on high-voltage  
97 electrical power transmission lines (HVTL). Corona-generated OH would react with atmospheric  
98 constituents and organic materials near the corona. In this paper, we test this hypothesis using  
99 analysis of quantitative observations from the top of an 18-story building, laboratory studies,  
100 photochemical box model calculations, and estimates of potential oxidation in the atmosphere and  
101 on nearby materials.  
102

## 103 **Results**

104 Extremely large OH and  $\text{HO}_2$  signals were observed with the Ground-based Tropospheric  
105 Hydrogen Oxide Sensor (GTHOS) during times when thunderstorms were over the TexAQS II  
106 Radical Measurement Project (TRAMP) rooftop site in Houston, Texas (Figs. S1). TRAMP's focus  
107 was on Houston air quality (10), so these large signals were noted but not carefully analyzed until  
108 recently. Re-examining these signals was motivated by the discovery of extreme OH and  $\text{HO}_2$   
109 generated by electrical discharges in thunderstorm anvil clouds and enabled by new analysis  
110 methods that produce OH and  $\text{HO}_2$  measurements every 0.2 seconds.  
111

112 During the six-week long study, enhanced  $\text{HO}_x$ , which is the sum of OH and  $\text{HO}_2$ , was observed  
113 during seven periods when thunderstorms were overhead (Figure 1, Table S1). In these periods,  
114  $\text{HO}_x$  reached to over 3000 pptv on 23 September, exceeded 1000 pptv on four other occasions,  
115 and had a mean value of 120 pptv (Fig. 1A). Enhanced OH reached to over 1000 pptv, with a mean  
116 value of 47 pptv. Also plotted in Figures 1C and 1D are  $\text{HO}_2$  and OH, both observed and modeled  
117 using the RACM chemical mechanism (11). The observed and modeled OH and  $\text{HO}_2$  are  
118 consistent for times without thunderstorms, thus supporting the GTHOS calibration. Compared to  
119 the typical daytime peak values of 10-50 pptv for  $\text{HO}_2$  and 0.3-0.6 pptv for OH, these enhanced  
120  $\text{HO}_x$  values are one to two orders of magnitude higher for  $\text{HO}_2$  and two to three orders of magnitude  
121 higher for OH. These are the largest concentrations of OH and  $\text{HO}_2$  ever observed in the  
122 atmosphere, larger even than those observed in thunderstorm anvils clouds.  
123

124 These extreme  $\text{HO}_x$  amounts always occurred during thunderstorm activity, which included periods  
125 of rain, sometimes heavy (Table S1). All seven periods of enhanced  $\text{HO}_x$  occurred under overhead  
126 thunderstorms (cumulonimbus clouds, Cbs) with low brightness temperatures, which indicates  
127

128 towering convective clouds. Of these seven periods, three of the largest enhancements occurred  
129 primarily in the minutes just before heavy rain began, two occurred primarily during the rain, and  
130 two began minutes after the heaviest rain had fallen. For the six times that rain was indicated by  
131 OH Rayleigh scattering (called “rain” in Fig. 1) but no enhanced HO<sub>x</sub> was observed (Table S1), the  
132 satellite visible and water vapor infrared images showed that the rain fell from either edges of small  
133 cumulonimbus clouds or from widespread stratocumulus clouds.

134  
135 Three segments from three days show the detailed behavior of OH and HO<sub>2</sub> (Fig. 2). On 19 August  
136 (Fig. 2A), the enhanced HO<sub>x</sub> was a series of spikes lasting less than a second each, whereas on 1  
137 and 15 September (Fig. 2B, 2C), it was variable but more continuous. Enhanced HO<sub>x</sub> was observed  
138 only when the laser wavelength was on the OH absorption line, but the scattering signal in the OH  
139 detection axis due to rain occurred also when the laser wavelength was off the OH absorption line.  
140 On 15 September during an hourly test period (Fig. 2C), the HO<sub>2</sub> signal dropped to background  
141 when the reagent NO was turned off, and the OH spectrum can be seen in both the OH and HO<sub>2</sub>  
142 signals during a laser wavelength scan (please see SI for more detail), providing strong evidence  
143 that the extreme signals were atmospheric OH and HO<sub>2</sub>.

144  
145 We searched through the GTHOS data sets of a dozen other tower-based field studies. For most  
146 of these, no thunderstorms rumbled over the measurement site. Also, for the past 3½ field studies,  
147 a Teflon inlet has been placed over the usual metal inlet in order to add a second OH-detection  
148 method by scavenging OH with a chemical reactant. This scavenging inlet was added for the  
149 second half of the Study of Houston Atmospheric Radical Precursors (SHARP) field study in 2009  
150 (12). In SHARP, the GTHOS inlet was located within two meters of the same location it occupied  
151 during TRAMP and, for the first few weeks, had the same grounded aluminum inlet that was used  
152 during TRAMP(13) (Figs. S1, 3).

153  
154 Only one thunderstorm brought heavy rain to the site during those few weeks – on 24 April 2009  
155 between 19:00 and 22:00 CDT (Fig. S2). More than 40 individual sub-second enhanced HO<sub>x</sub> spikes  
156 were measured. They first occurred a few minutes before the rain began and continued throughout  
157 the rainy period. Median values were 1.1 ppbv for OH and for HO<sub>2</sub>, and 2.2 ppbv for HO<sub>x</sub>. Maximum  
158 HO<sub>x</sub> was 16 ppbv, more than 4 times larger than the largest peak measured during TRAMP. After  
159 the scavenging inlet was placed over the metal inlet, several thunderstorms occurred but no  
160 extreme HO<sub>x</sub> spikes were observed. These TRAMP and SHARP observations raise the question  
161 of the origin of the corona.

162  
163 Origin of the observed extreme OH and HO<sub>2</sub>  
164  
165 The DC3 study demonstrated that electrical discharges in thunderstorm anvils at 9 to 12 km altitude  
166 could produce extreme amounts of OH and HO<sub>2</sub> (1), and our previous laboratory studies showed  
167 that extreme HO<sub>x</sub> amounts could be generated by either visible flashes or subvisible discharges at  
168 any tropospheric pressure (2). In that research, visible corona was not studied.

169  
170 One possible origin is the measurement instrument itself. GTHOS had known interferences for both  
171 OH and HO<sub>2</sub> during TRAMP, which have since been remediated. However, laboratory and field  
172 experiments all show that the measured extreme OH and HO<sub>2</sub> are inconsistent with these known  
173 interferences. Please see the SI for more detail. The extreme OH and HO<sub>2</sub> are 100 times larger  
174 than can be produced by known sources. The spikes and variable signals are uncorrelated with the  
175 amounts and variation of known interference sources. Also, laboratory experiments demonstrate  
176 that measured extreme OH and HO<sub>2</sub> are interference free (2).

177  
178 The other known mechanisms for generating extreme OH and HO<sub>2</sub> are combustion and high-level  
179 energetic radiation, but neither of those sources existed near the GTHOS sampling inlet during  
180 TRAMP. Unfortunately, TRAMP had no instruments that could detect electric fields or electrical

181 discharges. Nevertheless, the coincidence of  $\text{LHO}_x$  and thunderstorms overhead indicates that  
182 these extreme OH and  $\text{HO}_2$  values were likely generated by electrical discharges near or on the  
183 GTHOS sampling inlet (Table S1). It is likely that either subvisible discharges or corona were the  
184 type of electrical discharge and not sparks because sparks cause noise on the GTHOS electronics  
185 channels for the microchannel plate detector high voltage (1) and this electronic noise was not  
186 observed during TRAMP nor in recent laboratory studies of corona.

187  
188 These OH and  $\text{HO}_2$  mixing ratios of 100 to 2000 pptv translate into concentrations of  $2.3 \times 10^9$  to  
189  $4.6 \times 10^{10}$  molecules  $\text{cm}^{-3}$ . At these concentrations, OH and  $\text{HO}_2$  rapidly react with many chemical  
190 species, but they also react with each other to form  $\text{H}_2\text{O}$  and  $\text{O}_2$  with a reaction rate coefficient of  
191  $\sim 10^{-10}$   $\text{cm}^3 \text{molecule}^{-1} \text{s}^{-1}$ . Thus, the  $\text{HO}_x$  lifetime from this reaction alone is in the range of 0.22 to  
192 4.3 seconds. For the periods when extreme  $\text{HO}_x$  was observed, the typical wind speeds measured  
193 within 2 m of the GTHOS sampling inlet were 2 to 3  $\text{m s}^{-1}$ . As a result, the sources of the extreme  
194  $\text{HO}_x$  had to have been within 0.5 to 13 m of the GTHOS inlet – somewhere on top of the building.

195  
196 This estimate of corona origin was further refined by comparing observed ratios of OH to  $\text{HO}_2$  to  
197 those modeled with a photochemical box model for many  $\text{LHO}_x$  peaks. Our laboratory studies  
198 demonstrate that coronas produce equal amounts of OH and  $\text{HO}_2$ , and, because OH reacts with  
199 many chemical species and decreases more rapidly than  $\text{HO}_2$ , the OH/ $\text{HO}_2$  ratio can be used as a  
200 measure of the time since the  $\text{LHO}_x$  was generated. For seventeen cases selected as being  
201 representative, the observed OH/ $\text{HO}_2$  was 0.2 to 0.9. In fifteen of the seventeen cases, the  
202 observed ratios were achieved in less than 0.2 seconds, with twelve occurring in less than 0.1  
203 seconds (Fig. S3). Note that these reaction times are determined only by the calculated OH/ $\text{HO}_2$   
204 ratios with the assumption the OH equals  $\text{HO}_2$  initially at time 0. They are not determined by the  
205 actual GTHOS measurement times because the GTHOS time resolution is only 0.2 seconds and  
206 the exact time that the corona generated OH relative to the start and end of each 0.2 second data  
207 bin is unknown. When these times are multiplied by the windspeed, most corona must have  
208 originated either from the inlet itself, and, in a few cases, from the extension posts of the scaffolding  
209 shown in Fig. S1.

210  
211 The photochemical box modeling also enabled the calculation of the OH exposure and the OH  
212 reaction rates. OH exposure is the integral of OH as a function of time and determines the total  
213 amount of each OH reactant that is removed (1, 2). Thus, OH exposure is a more important quantity  
214 for discharge-generated OH than the OH concentration is. Generally, a substantial fraction (0.2-  
215 0.8, mean: 0.5) of OH is lost to the  $\text{HO}_x$ -terminating reactions  $\text{OH} + \text{HO}_2 \rightarrow \text{H}_2\text{O} + \text{O}_2$ ,  $\text{HO}_2 + \text{HO}_2 \rightarrow \text{H}_2\text{O}_2$ ,  
216 and  $\text{OH} + \text{NO}_2 + \text{M} \rightarrow \text{HNO}_3 + \text{M}$ , where M is  $\text{N}_2$ ,  $\text{O}_2$ , and  $\text{H}_2\text{O}$ . However, even with this large terminal  
217 loss, the mean and median OH exposure was  $3 \times 10^9$  molecules  $\text{cm}^{-3} \text{s}$ , with a range of 0.3 to  $9 \times 10^9$   
218 molecules  $\text{cm}^{-3} \text{s}$ . This OH exposure is 1000 times greater than the mean OH exposure generated  
219 by normal photochemistry during TRAMP, albeit for a smaller volume and duration.

220  
221 Laboratory simulations provide OH and  $\text{HO}_2$  amounts comparable to TRAMP  
222

223 To test the hypothesis that corona on the GTHOS inlet could cause the extreme OH and  $\text{HO}_2$   
224 values, we performed a series of laboratory experiments described in the Materials and Methods  
225 section. Coronas numbered from one to many on the GTHOS inlet depending on the applied  
226 electric field (Fig. 3). Positive (negative) coronas occur when the ground metal object is charged  
227 positive (negative) relative to the thunderstorm cloud in the atmosphere and the other flat electrode  
228 in the laboratory. Note that the structures of positive and negative coronas are quite different, with  
229 positive coronas in elongated corona, for which the electrons are accelerated toward the inlet, and  
230 negative coronas more like glowing balls, for which the electrons are repelled from the inlet (14, 15)

231  
232 When the laboratory electric field for negative coronas was increased, OH,  $\text{HO}_2$ , and  $\text{O}_3$  increased

233 proportionally with the laboratory electric field (Fig. 4). NO and NO<sub>2</sub>, which are produced in greater  
234 amounts in positive corona than in negative corona, were measured to be below ~50-100 pptv, the  
235 detection limit of the laboratory instrument. At the highest electric fields in this experiment, HO<sub>x</sub>  
236 reached more than 20 ppbv and O<sub>3</sub> reached 300 ppbv (Fig. 4). The R<sup>2</sup> for the correlation between  
237 HO<sub>x</sub> and O<sub>3</sub> is 0.96, indicating that the production of HO<sub>x</sub> and O<sub>3</sub> are related. The ratio of O<sub>3</sub> to HO<sub>x</sub>  
238 is 14±2 for this experiment in which the water vapor fraction was 0.012. This ratio is coincidentally  
239 the same as that obtained in thunderstorm anvils during the DC3 airborne study in 2012 (16).  
240

241 Mapping HO<sub>x</sub> generated by corona on a GTHOS inlet by sampling the OH, HO<sub>2</sub>, and O<sub>3</sub> downwind  
242 of the corona enabled a calculation of the OH production rate (Fig. 5). For the experiment described  
243 in the Materials and Methods section, the peak HO<sub>x</sub> was 4.2 ppbv and the average HO<sub>x</sub> production  
244 rate was calculated to be  $1.8 \pm 0.9 \times 10^{15}$  HO<sub>x</sub> (OH+HO<sub>2</sub>) molecules s<sup>-1</sup>, with OH and HO<sub>2</sub> in equal  
245 amounts. This value is a million times the OH production rates generated in a similar volume of  
246 Houston air by the well-known urban HO<sub>x</sub> sources (11).  
247

248 If the GTHOS inlet can generate LHO<sub>x</sub>, then other grounded metal objects should also be able to  
249 generate LHO<sub>x</sub> as well. In laboratory experiments, four types of lightning rods were tried: stainless  
250 steel static dissipators; pointed copper, rounded copper, and rounded aluminum. The amounts of  
251 LHO<sub>x</sub> generated on these lightning rods was comparable to that generated on the second GTHOS  
252 inlet. However, for both positive and negative coronas, it was easier to initiate coronas on any of  
253 the typical types of lightning rods than it was on the GTHOS inlet (Fig. S4). Thus, it is likely that  
254 lightning rods and other grounded metal objects on buildings also generate coronas and enhanced  
255 OH and HO<sub>2</sub> as thunderstorms pass overhead.  
256

## 257 **Discussion**

258 These results demonstrate that visible coronas on grounded metal objects generate extreme  
259 amounts of OH and HO<sub>2</sub> and that more intense coronas generate more extreme amounts. The  
260 amounts of OH and HO<sub>2</sub> generated in the laboratory are also consistent with the amounts observed  
261 under thunderstorms during two field campaigns and the amounts generated by subvisible  
262 discharges and flashes in thunderstorm anvil clouds.  
263

264 This observation that coronas can occur on the GTHOS inlet requires that we revisit the possibility  
265 that coronas on the Airborne Tropospheric Hydrogen Oxide Sensor (ATHOS) inlet were responsible  
266 for at least some of the LHO<sub>x</sub> observed in the anvil clouds during DC3 (1). Evidence from the field  
267 and laboratory observations ruled out sparks and subvisible discharges, but not weak coronas.  
268 However, when the analysis using modeling and the HO<sub>2</sub>/OH ratio described earlier and in the SI  
269 was applied to the DC3 measurements, 86% of the discharges occurred more than 0.3 seconds  
270 (~60 m) in front of the ATHOS inlet, with a median time of 0.56 seconds (~90 m) (Fig. S5). Even  
271 the three cases at times less than 0.2 seconds could have occurred in front of the ATHOS inlet  
272 because 0.2 seconds is the measurement time resolution. These results provide strong evidence  
273 that the enhanced OH and HO<sub>2</sub> observed in anvil clouds during DC3 was not due to coronas on  
274 the ATHOS inlet.  
275

276 The laboratory experiments capture the magnitude of the OH and HO<sub>2</sub> observed during TRAMP  
277 and SHARP but do not capture the observed behavior. In some cases, HO<sub>x</sub> was generated in ppbv-  
278 level sub-second spikes (Fig. 2A), but in other cases it was generated as variable continuum for  
279 minutes at a level of a few hundred pptv (Fig. 2B, 2C). We have no solid explanation for the corona  
280 generating large sub-second HO<sub>x</sub> spikes in some cases and in less amounts but more constant  
281 HO<sub>x</sub> in others because the characteristics of the electric field and the space charge were not  
282 measured during TRAMP. However, it is possible that these differences are due to coronas of  
283 different polarity, different electric field strengths, different charge distributions, or a combination of  
284 all three.  
285

286  
287 In the atmosphere, coronas can deplete the charge distributions, reducing the electric fields;  
288 charges can move to screen the thunderstorm electric field; or the electric fields can vary, even  
289 flipping polarity as the storm passes overhead. In the laboratory, the DC HV supplies continually  
290 provide current while maintaining the voltage, thus sustaining the corona. Nonetheless, the  
291 comparable amounts of extreme OH and HO<sub>2</sub> observed in TRAMP, SHARP, and the laboratory  
292 indicates that the quantities found in the laboratory – HO<sub>x</sub> production rates and OH exposures –  
293 apply to the TRAMP and SHARP observations as well.  
294  
295 Both lightning rods in a typical urban area and high-voltage electrical power transmission lines are  
296 considered as sources of corona-generated OH. In both cases, the corona-generated OH is  
297 compared to the OH produced by well-known ambient processes. Please see the SI section 2.1 for  
298 more detail.  
299  
300 Consider a lightning rod under the typically 50 thunderstorms that annually pass over Midwestern,  
301 Mid-Atlantic, and Southern cities in the United States (17). For the approximately one hour that  
302 corona are estimated to be generated on a lightning rod annually, the thunderstorm-induced  
303 corona-generated OH is greater than or equal to ambient total OH for ambient air volumes of ~100  
304 m<sup>3</sup> (Fig. S6). Annually, corona-generated OH increases the oxidation by OH as much as a factor  
305 of 10 over ambient within ~10 cm of the lightning rod, thus greatly accelerating the degradation of  
306 nearby atmospheric constituents and surface materials. For an entire city, the annual impact is  
307 confined to the immediate vicinity of the lightning rods (Fig. S7). However, this analysis considers  
308 neither downwind effects of corona-oxidized air, the possibility that other pointed metal objects  
309 produce corona, or OH generated by lightning. These three effects are difficult to estimate, but all  
310 would increase the impact of thunderstorms on atmospheric oxidation in a city.  
311  
312 HVTLs produce corona along the length of the lines and at connections to insulators. Estimates of  
313 OH production per meter of line come from a combination of the previously measured O<sub>3</sub> (18,19)  
314 and our laboratory results of the relative amounts of OH and O<sub>3</sub> produced. Please see the SI section  
315 2.1. The corona occur continuously, although ~10 times more occurs in foul weather than in fair.  
316 Although the estimated amount of corona-generated OH along an HVTL is small compared to total  
317 ambient OH generated throughout the planetary boundary layer at the HVTLs, it is more than  
318 50,000 times larger than ambient within 20 cm (volume of 0.1 m<sup>3</sup>) of the corona and 500 times  
319 larger in a volume of 1 m<sup>3</sup> per meter of line. The distance that enhanced OH is transported  
320 downwind of the HVTL can be estimated using a simple Gaussian model (please see the SI for  
321 details). This model indicates that trees and structures at the edges of HVTL right of ways are  
322 routinely subjected continuously to 10 to 100 times ambient OH. Thus, corona-generated OH could  
323 greatly enhance atmospheric oxidation in close proximity to the HVTLs and affect trees more than  
324 10 meters from the HVTLs.  
325  
326 Corona-generated OH initiates the degradation of polymers used for HVTL insulators by breaking  
327 a hydrogen-carbon bond (20). Please see the SI section 2.2 for more detail. Corona-generated O<sub>3</sub>  
328 is only ~30 times greater than corona-generated OH, based Figure 4, yet it reacts ~100,000 times  
329 slower than OH for hydrocarbons with double carbon bonds and not at all for hydrocarbons without  
330 double bonds. The comparison between corona-generated UV radiation and OH is more  
331 complicated. Based on our laboratory measurements of corona-generated OH and UV radiation  
332 with wavelengths less than 290 nm (21), we estimate that the OH flux for a polymer cylinder is  
333  $6 \times 10^9$  molecules cm<sup>-2</sup> s<sup>-1</sup>, 4 times larger than the UV flux of  $1.5 \times 10^9$  molecules cm<sup>-2</sup> s<sup>-1</sup>, with an  
334 uncertainty of a factor of ~4 at 68% confidence. If these OH and UV fluxes are equally efficient at  
335 breaking this bond, then corona-generated OH is more important than corona-generated UV  
336 radiation for the premature degradation of the preferred insulating material, polydimethylsiloxane.  
337

338 Our discovery of extreme OH generated by corona on pointed grounded surfaces under  
339 thunderstorms leads to rethinking the effects of corona on HVTL insulators. Up to now, it has been  
340 thought that corona UV radiation in the 200-300 nm wavelength range was responsible for initiating  
341 the observed premature degradation of polymer HVTL insulators, which are increasingly replacing  
342 old ceramic insulators (22). In the literature, the existence of corona-generated OH appears to be  
343 completely unknown. However, according to our estimates, corona-generated OH, not corona-  
344 generated UV radiation, initiates most of the premature degradation of polymer insulators used with  
345 high-voltage electrical power transmission lines.

346

347

## 348 Materials and Methods

349

### 350 The TRAMP study and measurements

351

352 The TRAMP air quality field study was conducted in Houston, TX in July and August, 2006 (10).  
353 The measurements were made from a scaffolding platform on top of the northernmost one of the  
354 twin Moody Tower dormitories on the University of Houston main campus (latitude: 29.7176°;  
355 longitude: 95.3413°W). The measurements were made at ~70 m above ground level. The campus  
356 is 5 km southeast of the tallest buildings in downtown Houston and 6-25 km west-southwest of the  
357 primary petrochemical facilities.

358

359 The OH and HO<sub>2</sub> measurements were made with the Ground-based Tropospheric Hydrogen  
360 Oxides Sensor (GTHOS), which is described in more detail in the SI. In addition to OH and HO<sub>2</sub>,  
361 the measurement suite included meteorological parameters, photolysis frequencies, O<sub>3</sub>, NO, NO<sub>2</sub>,  
362 CO, peroxydes, formaldehyde, and many volatile organic compounds, as shown in Table S2 (10).  
363 This measurement suite enabled the modeling of OH and HO<sub>2</sub> both for background conditions and  
364 for the rapid decay of OH and HO<sub>2</sub> that occurred following their generation by electrical discharges  
365 in the intensified electrical fields under thunderstorms.

366

### 367 Laboratory studies

368

369 In the laboratory, OH and HO<sub>2</sub> were measured with GTHOS. The corona was generated on an  
370 electrode of interest, which was used as the ground of a pair of electrodes, and the other electrode  
371 was a flat aluminum plate (17 cm×16 cm×0.5 cm) to which either negative or positive DC high  
372 voltage was applied. The primary electrode of interest was a second GTHOS inlet, the one that  
373 was used on TRAMP and SHARP, which was attached to a vacuum pump to simulate the airflow  
374 patterns of the actual GTHOS inlet. OH and HO<sub>2</sub> were actually sampled by an inlet attached to  
375 GTHOS. We were reluctant to use the actual GTHOS inlet as an electrode for more than a few  
376 minutes because of the potential for a spark to damage the GTHOS laser, electronics, or detectors.  
377 Four different lightning rods were also studied: stainless steel static dissipators; pointed copper,  
378 rounded copper, and rounded aluminum.

379

380 Note that this method of generating corona is different from the method used to generate subvisible  
381 discharges and sparks previously (2). Those previous discharges were generated by a Tesla coil  
382 as short electrical pulses across a pair of tungsten wire electrodes separated by one to two cm.

383 Corona experiments were run with the electrodes between 2.0 cm to 4.0 cm apart, with applied  
384 absolute voltages of ~7 kV to 32 kV with either positive or negative DC high voltage supplies  
385 (Information Unlimited HV350REG). A blower blew air past the corona to the GTHOS sampling  
386 inlet with a velocity profile that was flat to within 10% over a 10 cm by 6 cm cross sectional area  
387 and was set in the range of 1 to 10 m s<sup>-1</sup> toward the GTHOS inlet, which was located ~20 cm away.  
388 Located 1.5 cm to one side of the GTHOS inlet was a 1/4" Teflon tube, which was connected to an  
389 instruments to measure O<sub>3</sub> (Thermo Scientific 49i) and NO, NO<sub>x</sub>, and thus NO<sub>2</sub> (Thermo Scientific

42i and ECO PHYSICS nCLD 855Y). Other measurements include wind speed with a hot-wire  
391 anemometer (TSI Inc., 8455-09) and pressure (MKS Instruments Baratron® Type 222),  
392 temperature and relative humidity (Vaisala HMT310). Laboratory room air was used for most of  
393 these experiments, but, for a few experiments, dry-to-humidified purified air was used. The same  
394 results were found no matter the air source.

395  
396 In some experiments, the laboratory electric field was varied. OH and HO<sub>2</sub>, O<sub>3</sub>, NO, and NO<sub>2</sub> were  
397 simultaneously measured as the voltage applied to the plate was increased from ~7 kV to 32 kV.  
398

399 In another experiment, corona were generated on the second pumped GTHOS inlet used as the  
400 grounded electrode. A voltage of +22.5 kV was applied to the flat plate set 2.0 cm from the second  
401 GTHOS inlet. Corona was observed on the sharp edges on the outside of the inlet but never inside  
402 the inlet. OH and HO<sub>2</sub> generated by this corona were mapped by moving the corona relative to the  
403 sampling inlets. Measurements were taken for 1-2 minutes on positions in a 1.8 cm x 2.0 cm grid.  
404 The corona was 22 cm upwind of the GTHOS sampling inlet and the wind speed was 9 m s<sup>-1</sup>. The  
405 total OH production rate (OH molecules s<sup>-1</sup>) was found by multiplying the OH mixing ratio by the  
406 cross-sectional area surrounding each point, adding them together, and then multiplying by the air  
407 speed.

408  
409 Photochemical box modeling  
410

411 Photochemical box modeling was accomplished with the Framework for 0-D Atmospheric  
412 Modeling (F0AM) (23) using the Master Chemical Mechanism v3.3.1 (24,25). Details of the  
413 modeling can be found in Brune et al. (1). For this study, the photochemical box model was used  
414 for two primary purposes: determining either the OH exposure for the TRAMP LHO<sub>x</sub>  
415 measurements or the distance between the coronas and the GTHOS sampling inlet.

416  
417 The OH exposure was determined by multiplying, at each time step, the OH concentration by the  
418 width of the time bin, and then summing these products. The distance between the coronas and  
419 the GTHOS inlet was determined by a four-step process. First, OH and HO<sub>2</sub> mixing ratios were  
420 set equal to each other at half the value of the observed HO<sub>x</sub> because our laboratory studies  
421 show that OH and HO<sub>2</sub> are generated in equal amounts by coronas. Second, the decays of OH  
422 and HO<sub>2</sub> were modeled. Third, the time at which the modeled OH/HO<sub>2</sub> equaled the observed  
423 OH/HO<sub>2</sub> was found. Fourth, this time was multiplied by the measured wind speed to find the  
424 distance between the coronas and the GTHOS inlet.

425  
426 For both purposes, seventeen snapshots of OH and HO<sub>2</sub> were chosen amongst the seven  
427 periods that thunderstorms generated enhanced OH and HO<sub>2</sub> during the TRAMP study. These  
428 represent the range of initial HO<sub>x</sub> mixing ratios measured during TRAMP. The model was  
429 constrained to the O<sub>3</sub>, NO, NO<sub>2</sub>, CO, HCHO, volatile organic compounds, pressure, temperature,  
430 and relative humidity measured during the study (Table S2) and run for 100 seconds.

431  
432 Data and modeling access  
433

434 The data sets and information on the modeling code used in this study are publicly  
435 available: <https://doi.org/10.26208/k4ce-zw19>. Explanation of the data can be found in a txt  
436 file as well as in the SI, Section 3.

437  
438

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440

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446

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448

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520 **Figures legends**

521

522 **Figure 1.** TRAMP OH and HO<sub>2</sub> measurements versus Central Standard Time (CST). Shown are  
523 (A) rain-induced GTHOS scattering signal (arbitrary units), designated as “rain”; (B) periods with  
524 extreme HO<sub>x</sub> (gold) and OH (blue) mixing ratios (pptv), indicating LHO<sub>x</sub>; (C) ambient HO<sub>2</sub> from  
525 measurement (gold) and model (black); and (D) ambient OH from measurement (blue) and model  
526 (cyan).

527 **Figure 2.** Three examples of OH and HO<sub>2</sub> generated by electrical discharges versus time (CST).  
528 Three minutes were chosen for August 19 (A), Sept 1 (B), and Sept 15 (C). In A and B, dashed  
529 lines indicate times the laser wavelength is on an OH absorption line. In C, the dot-dashed line  
530 indicates an hourly instrument test period including a wavelength scan of the laser (15:31-15:33,  
531 with the OH Q<sub>1</sub>(2) fluorescence near 15:32) and the thicker dashed line indicates a time when the  
532 reagent NO was turned off, so that only OH was measured.

533

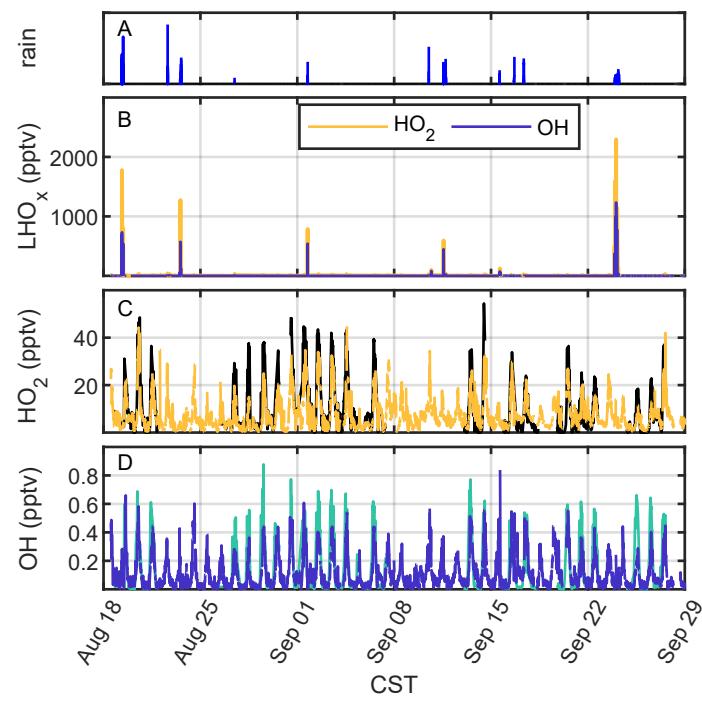
534 **Figure 3.** Coronas generated on the edges of the GTHOS inlet in the laboratory. In A are three  
535 positive elongated coronas and three faint wisps, with the inlet grounded and the overhead plate  
536 negatively charged. In B are negative ball-shaped coronas, with the inlet grounded and the  
537 overhead plate positively charged. For these pictures, the DC voltage was 25 kV and the inlet to  
538 plate separation was 2.5 cm. The flat part of the inlet is 19 mm in diameter.

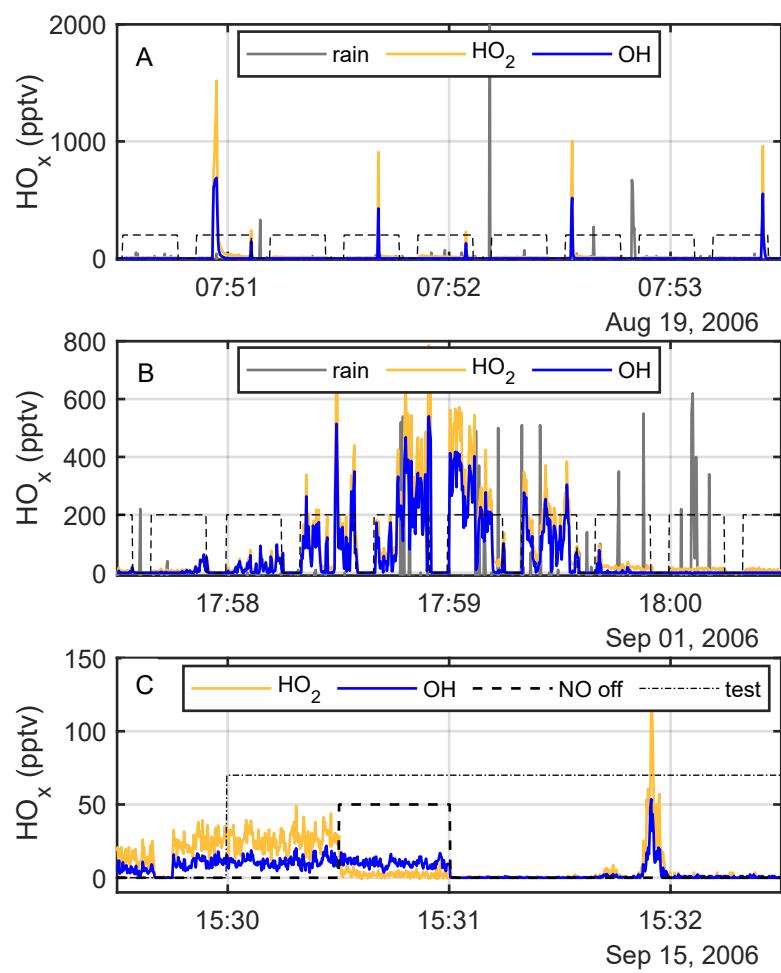
539

540 **Figure 4.** HO<sub>x</sub> and O<sub>3</sub> on a GTHOS inlet as a function of the laboratory electric field. A negative  
541 corona with an electrode separation of 2 cm was used in this experiment, but positive coronas give  
542 similar behavior, although different values as a function of electric field.

543

544 **Figure 5.** Laboratory map of HO<sub>x</sub> mixing ratio for a corona on a second pumped GTHOS inlet used  
545 as the positive electrode. Peak HO<sub>x</sub> was 4.2 ppbv, with OH and HO<sub>2</sub> abundances equal to within  
546 10%. The calculated HO<sub>x</sub> production rate is  $1.8 \times 10^{15}$  molecules s<sup>-1</sup> (please see SI for details).

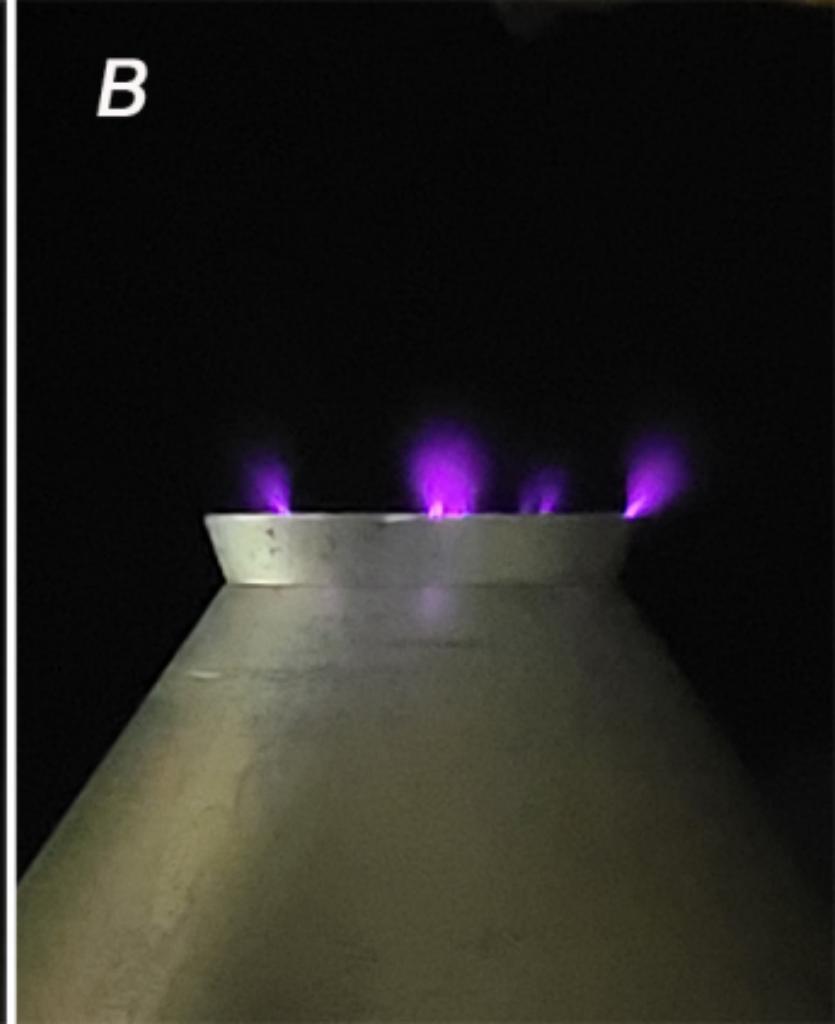


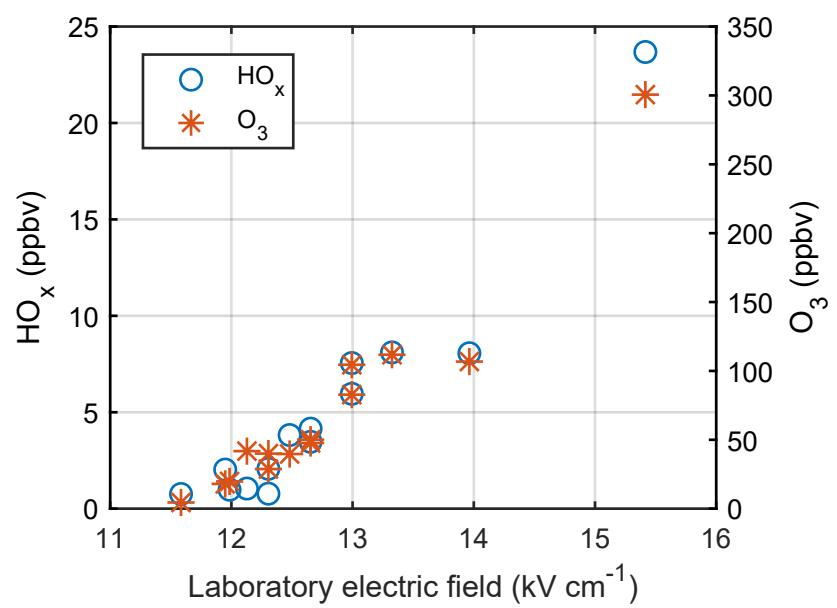


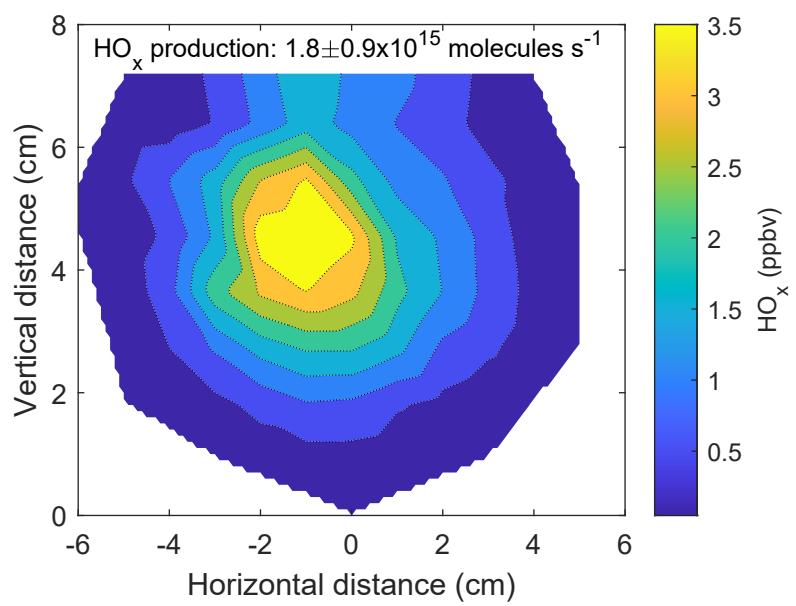
**A**



**B**









**Supplementary Information for**  
Extreme hydroxyl amounts generated by thunderstorm-induced  
corona on grounded metal objects

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**This PDF file includes:**

Supplementary text  
Figures S1 to S9  
Tables S1 to S2  
SI References

## Supplementary Information Text

### 1. TRAMP field study

The TexAQS II Radical Measurement Project (TRAMP) field study was conducted in Houston, TX in July and August, 2006. The goal of the study was to improve the understanding of the gaseous emissions and photochemistry that created the frequent air pollution in Houston, TX, particularly the ozone pollution (1). This site was chosen because it frequently sampled air from either downtown, the petrochemical facilities, or the cleaner air from the Gulf of Mexico to the south, and, because of its elevation, was less susceptible to nearby pollution sources. These months were chosen because peak ozone pollution occurs then.

A second study, the Study of Houston Atmospheric Radical Precursors (SHARP), occurred at the same location as TRAMP but in April and May 2009 (2). These months were chosen because plumes from biomass burning in Mexico and Central America reach Houston and a second annual ozone peak occurs then.

#### 1.1. Ground-based Tropospheric Hydrogen Oxides Sensor (GTHOS) measurement

OH and HO<sub>2</sub> were measured with the Ground-based Tropospheric Hydrogen Oxides Sensor (GTHOS), which is extensively described in Faloona et al. (3). The basic technique is laser induced fluorescence in a low-pressure detection chamber. Air is pulled through a ~1 mm diameter orifice with a vacuum pump, producing a low-pressure (~5 hPa) flow into a pair of detection axis. A pulsed laser beam (3 kHz frequency, 20 ns long) is sent into multipass optical cells in each detection axis and is monitored at the exit of the cells. Microchannel plate detectors, set at right angles to the laser beam and the air flow, detect scattering from the laser pulse in low-gain mode and, after the laser pulse has ended, the weak OH fluorescence in high-gain mode. The laser is alternately tuned to a wavelength near 308 nm that is not absorbed by OH and then to one that is absorbed. The difference between these two signals is proportional to OH. Between the first and second detection axes, reagent NO is added, reacting with HO<sub>2</sub> to form OH, which is then detected. GTHOS was calibrated before, after, and during the study. The absolute uncertainty is  $\pm 32\%$ , 95% confidence and the limit-of-detection is 0.04 pptv for OH and 0.2 pptv for HO<sub>2</sub> for a one-minute averaging time.

The GTHOS electronic and laser were located in an air-conditioned trailer on the building roof and the GTHOS detection system was located at the northeastern corner of the scaffolding (Fig. S1A). During SHARP, the measurement of OH and HO<sub>2</sub> was made from the same scaffolding platform as during TRAMP, but the instrument was located 2 meters to the south, in the nearest corner in Fig. S1A. Electrical cables and fiber optics connected the electronics and the laser to the detection axes. The detection axis were housed in a weatherproof box, with the inlet extending 40 cm above the top surface of the box (Fig. S1B). A sonic anemometer was located ~2 meters to the east and ~20 cm below the GTHOS inlet.

The GTHOs inlet was higher than all structures except a triangular radio tower holding meteorology and photolysis frequency instruments. It was ~6 m to the southwest of the GTHOS inlet and ~3 m higher. It was topped with three brush-type lightning rods.

#### 1.2. Evidence that the extreme OH and HO<sub>2</sub> are real

Two segments of the observed LHO<sub>x</sub> provide tests of the validity of these extreme OH and HO<sub>2</sub> values (Fig. 2). As the laser wavelength was tuned on and off resonance with an OH absorption line, the signals were detected only when the laser wavelength was on resonance. However, being on or off resonance makes no difference to the scattering signal caused by rain drops entering the inlet, as can be seen in the off-resonance gap just after 17:59:30 in Fig 2B.

Two other diagnostics provide more evidence for these signals being real HO<sub>2</sub> and OH (Fig. 2C).

First, just before 15:32:00, the laser wavelength was scanned to ensure that measurements were being made with the  $Q_1(2)$  transition in the  $A^2\varSigma^+(\nu'' = 0) \rightarrow X^2\Pi^+(\nu' = 0)$  OH electronic transition. The larger  $Q_1(2)$  and smaller  $Q_{21}(2)$  lines can be clearly seen, although their relative values are affected by the variability in the discharge-generated HO<sub>x</sub>. Second, when the NO reagent to convert HO<sub>2</sub> to detectable OH is turned off from 15:30:30 to 15:31:00, the HO<sub>2</sub> values drop down to those expected of ambient OH detected in the HO<sub>2</sub> detection axis. These extreme signals are definitely due to HO<sub>2</sub> and OH.

Interferences for both HO<sub>2</sub> and OH can occur in environments with greater amounts of volatile organic compounds. The HO<sub>2</sub> interference involves the rapid internal conversion of organic peroxy radicals to HO<sub>2</sub> after they react with the reagent NO used to convert HO<sub>2</sub> to detectable OH and is well understood (4). The OH interference appears to affect some OH detecting instruments more than others, with GTHOS, which was used in the TRAMP study, being one of the most sensitive (5). It occurs only in VOC-rich environments (4,5). While both interferences are now being minimized by instrument modifications, neither was for TRAMP or SHARP (6).

For these interferences to be responsible for the extreme OH and HO<sub>2</sub> observed under thunderstorms during TRAMP and SHARP, they would have to satisfy many criteria. First, they would have to produce, by two completely different mechanisms, hundreds of pptv to ppbv of OH and HO<sub>2</sub>. The greatest OH and HO<sub>2</sub> interferences ever observed, even with a hundred ppbv of O<sub>3</sub> and hundreds of ppbv of VOCs, are a few pptv for OH and ~10 ppbv for HO<sub>2</sub>. Second, the HO<sub>2</sub> interference mechanism is well-understood and has been successfully modeled when the initial VOCs are known (4). When the HO<sub>2</sub> interference is modeled for SHARP, even the most extreme of the hourly averages is less than 10 pptv (6).

Third, the OH interference was quantified with the addition of an OH-scavenging inlet during the second half of the SHARP study and was at most 0.5 pptv. After the scavenging inlet was added, no extreme OH was observed even under several thunderstorms of the same intensity as the one that caused extreme OH and HO<sub>2</sub> during the first half of SHARP. Fourth, since VOC-rich environments are shown to be necessary for the HO<sub>2</sub> and OH interferences, the observed extreme OH and HO<sub>2</sub> should correlate with VOCs in both the time series behavior, such as one-second spikes, and the variability. They do not. Fifth, we have been able to generate comparable amounts of extreme OH and HO<sub>2</sub> by electrical discharges in the laboratory and demonstrated that they cannot be due to interferences (7). The hypothesis that the extreme OH and HO<sub>2</sub> are caused by interferences fails to satisfy any of these criteria, lending more credibility that this extreme corona-generated OH and HO<sub>2</sub> is real.

## **2. Estimates of impact of corona-generated OH, O<sub>3</sub>, and UV radiation on the atmosphere and surfaces**

Extreme amounts of OH generated by corona can accelerate oxidation of atmospheric gases or of surface materials. The analysis of potential impacts rests on two conditions: the range of impact in time and space compared to that of ambient amounts of OH; the importance of OH oxidation compared to oxidation by other processes, particularly ultraviolet (UV) radiation and O<sub>3</sub> oxidation.

### **2.1. Range of impact**

OH-producing corona are generated indoors by several devices, including electric motors, high-voltage equipment, and even increasingly used in building and residential air purification systems. They are also generated outdoors by lightning rods and other grounded metal objects under thunderstorms, such as is demonstrated in this paper, and by high voltage electrical power transmission lines and power distribution stations (8). Indoor impacts are the subject of ongoing research; here we focus on outdoor effects.

#### *2.1.1. Comparison of corona-generated OH on lightning rods under thunderstorms to ambient OH*

This comparison is between corona-generated OH molecules produced in a small volume around the lightning rod for a limited time when the thunderstorm is overhead to the total number of ambient OH molecules produced by well-known atmospheric photochemistry for a range of volumes and time periods. Comparison of the volume of air with the total number of ambient-produced OH molecules to the total number of corona-produced OH molecules is particularly important to determine the relative impact of corona-generated OH on atmospheric oxidation.

To find the total number of OH molecules that can contribute to atmospheric oxidation per lightning rod, we need to estimate the number of thunderstorms that create a corona on the lightning rod, the length of time and amount of OH produced for that time, and the fraction of that extreme OH that does not react with corona-generated HO<sub>2</sub> or OH and is thus available to react with other atmospheric constituents. In the United States east of the Rocky Mountains, approximately 50 thunderstorms occur over more than 100 cities annually (9). Based on the measurements presented in Fig. 1 of this paper, each thunderstorm produces, on average, 100 pptv of OH for ~400 s. From the laboratory studies and Fig. 5, this average value translates to a corona-generated OH production rate of  $4 \times 10^{14}$  molecules s<sup>-1</sup>, of which a fraction of 35% will react with corona-generated HO<sub>2</sub> and OH, leaving 65% available to react in the atmosphere or on surfaces. This estimated total OH generated by corona on one lightning rod is  $4 \times 10^{18}$  molecules per year.

A typical ambient OH production rate per cm<sup>3</sup> of air is 10<sup>7</sup> molecules cm<sup>-3</sup> s<sup>-1</sup>, when averaged over a year. Its loss is due to the same gas-phase reactions that cause loss of corona-generated OH (excepting with corona-generated HO<sub>2</sub> and OH). By multiplying the ambient OH production rate per volume by the different ambient air volumes and times, we can produce a map of the ratio of total annual corona-generated OH to total ambient produced OH as a function of ambient OH volume and time (Fig. S6).

For shorter times and smaller volumes of ambient OH, the corona-generated OH is a million times more (lower left). Whereas for longer times and volumes, the contribution of corona-generated OH is negligible. The black line indicates the conditions for which corona-generated and ambient OH are equal. Of more interest is the time when the thunderstorms are generating corona, which is less than an hour total annually, marked by the red dashed line. The thunderstorm-induced corona-generated OH is greater than or equal to ambient total OH for ambient air volumes less than 100 m<sup>3</sup> and 1,000 times larger in a volume of 0.1 m<sup>3</sup>. The most interesting case is the annual impact of corona-generated OH, which increases the oxidation by OH as much as a factor of 10 over ambient within 10 cm (volume = 0.01 m<sup>3</sup>) of the lightning rod, thus greatly accelerating the degradation of nearby materials.

The same approach was applied to examine the potential impact of OH generated by corona on all lightning rods that might be in an urban area. Unfortunately, lightning rods are not required in most cities' building codes, so the only way to make an estimate is to assume that all buildings greater than 25 m (~6 stories) have lightning rods distributed every 6 m around the perimeter according to engineering guidance (10). Austin, TX is a rapidly growing mid-sized city and a good example of a typical urban area in a moderately lightning-prone area. According to the building data for 2013 (11), more than 400 buildings are taller than 25 m, requiring 1500 lightning rods to meet Underwriters Laboratories (UL) lightning protection guidance. For the Austin urban core with an area of ~5 km<sup>2</sup>, to get the impact of corona-generated OH on the urban scale, we multiply the estimates for a single lightning rod by the estimated number of lightning rods.

These estimates for the urban scale indicate that the corona-generated OH impacts occur over a volume of 10<sup>5</sup> m<sup>3</sup> during the thunderstorms, and for a year, only over 20 m<sup>3</sup>, which essentially is just near each lightning rod (Fig. S7). At the urban scale, given by the long dash line, the corona-generated OH is negligible for all time scales. In summary, impact of OH generated on lightning rods under thunderstorms is negligible for atmospheric processes on all scales greater than 1 m<sup>3</sup>, but can provide greater than 10 times the oxidation as ambient OH for materials near the lightning

rods each year.

### 2.1.2. Comparison of corona-generated OH on HVTL to ambient OH

High-voltage electric power transmission lines (HVTLs) produce corona along the length of the lines and electrical power substations, where the the HVTL voltage is reduced for distribution to users (7). Research from the 1970's determined that  $6 \times 10^{17}$  O<sub>3</sub> molecules m<sup>-1</sup> s<sup>-1</sup> were generated for a 765 kV line in rainy or snowy weather (12,13). The amount of corona produced scales with the voltage and is as much as 10 times less in fair weather. We can scale these numbers to determine the amount of OH that is generated by typical HVTLs of  $\sim 270$  kV in fair weather, which gives a typical O<sub>3</sub> production rate of  $3 \times 10^{16}$  O<sub>3</sub> molecules m<sup>-1</sup> s<sup>-1</sup>. From our paper, corona produce 28 O<sub>3</sub> molecules for each OH molecule, so that the OH production rate is estimated to be  $1 \times 10^{15}$  OH molecules m<sup>-1</sup> s<sup>-1</sup>. This value is consistent with the OH production rate measured in the laboratory for the GTHOS inlet (Fig. 5), giving us confidence in this estimate.

As was done in the previous case of lightning rods under thunderstorms, the corona-generated OH were compared to the ambient OH for different volumes and times, although in this case, the analysis was performed per meter of HVTL (Fig. S8). Unlike the cases of lightning rods under thunderstorms, the corona on the HVTL occur all the time, which is the reason that the line designating equal amounts of corona-generated and ambient OH is vertical on the plot.

Corona-generated OH equaling ambient OH occurs at a volume of 540 m<sup>3</sup> around and downwind of each meter of HVTL. For instance, for 1 m of HVTL, a plume 5 m high would have OH equal to or greater than ambient 100 m downwind. At a volume of 1 m<sup>3</sup> per meter of HVTL, corona-generated OH is over 500 times greater than ambient. In the 20 cm surrounding the HVTL, the exposure of the HVTL to OH is more than a 50,000 times that of ambient OH. Thus, corona-generated OH is a strong oxidizer of the materials near the HVTL.

Using a simple Gaussian plume model, we improve upon the plume estimate given above by comparing the downwind OH exposure along the plume centerline and comparing that to the ambient OH exposure. We assumed the atmosphere was moderately unstable (14) and the windspeed was 3 m s<sup>-1</sup> perpendicular to the HVTL. The corona-generated OH exposure was calculated using MCMv3.3.1 in the F0AM framework constrained by conditions and model parameters appropriate for a slightly polluted forest (OH reactivity: 5.5 s<sup>-1</sup>; NO: 0.47 ppbv; NO<sub>2</sub>: 0.8 ppbv) and then dispersed according the Gaussian plume equation for a line source (15).

The centerline plume corona-generated OH is more than  $10^{10}$  OH cm<sup>-3</sup> at the HVTL and does not become equal to the ambient OH until the plume is 100 m downwind. The clearance on either side of HVTLs is typically less than 25 m, which means that structures or trees at the edge of the right-of-way and the height of the HVTLs are receiving more than 10 times the ambient OH amounts continually. For stronger winds or stable conditions, this continual dosing of the nearby forest or structures increases this OH exposure to 100 times ambient. These estimates of impact do not include the possible impact of air that has flown past the corona, been oxidized, and, as a result, has a different chemical composition.

## 2.2. Comparison of impact of OH, O<sub>3</sub>, and UV on commonly used insulating materials

In the absence of corona-generated OH, degradation of materials is driven more by solar UV radiation than by ambient OH levels, which are a diurnally averaged  $1-4 \times 10^6$  molecules cm<sup>-3</sup>. A substantial amount of research has been conducted on this solar UV degradation and chemical degradation sequences have been developed and refined (16). They are all initiated by solar UV below  $\sim 360$  nm wavelength, but, as part of the sequence, OH is produced and participates in the degradation. Ozone can also play a role, but only if the polymer material has exposed double carbon bonds.

Polymers are formulated to be resistant to solar UV. This UV resistance for some more UV-susceptible polymers is being achieved by adding UV-blocking or absorbing chemicals, or by adding other chemicals that interrupt the reactive chemistry that causes the polymer decomposition (17). For use as HVTL and power substation insulators, two polymers have emerged as having the necessary mechanical properties and resistance to degradation by solar UV and O<sub>3</sub>: Ethylene propylene diene monomer (EPDM), and polydimethylsiloxane (PDMS). However, even these polymers degrade in sunlight, albeit much slower rate than others (18).

The polymers are also degraded by nearby corona more rapidly than expected (19). Degradation results in loss of hydrophobicity, increased electrical conduction, and mechanical failure with cracking and chalking. It is always assumed that corona-generated UV is involved, with possible contributions from corona-generated O<sub>3</sub>, although OH inevitably is produced in the oxidation process and participates in further degradation. OH directly produced by the corona appears to have never been considered. The question is “Is the corona degradation of these polymers initiated by UV radiation, O<sub>3</sub>, or OH?”

From the laboratory studies of corona presented in this paper and another paper (21), we can determine relative importance of these three. EPDM (SMILES chemical formula: C=CC1=CC2CC1CC2=CCC=CC) has three double bonds in each monomer unit, and is thus susceptible to attack by UV radiation, O<sub>3</sub>, and OH. However, the corona-generated O<sub>3</sub>-to-OH ratio is only 28, while the the reaction rate coefficients for the gas-phase alkene is in the range of 10<sup>-11</sup> cm<sup>3</sup> molecule<sup>-1</sup> s<sup>-1</sup> for OH and 10<sup>-17</sup> cm<sup>3</sup> molecule<sup>-1</sup> s<sup>-1</sup> for O<sub>3</sub>. Thus, degradation initiation is 30,000 times faster by OH than by O<sub>3</sub>. PDMS (SMILES chemical formula: C[Si](C)(Cl)Cl, CO[Si](C)(C)OC) has no double bonds and thus no reaction with O<sub>3</sub>. The reaction rates of OH with organosilicons is ~10<sup>-12</sup> cm<sup>3</sup> molec<sup>-1</sup> s<sup>-1</sup> (21), so reactions on surfaces should be fast and the uptake coefficient should be close to 1, making diffusion to the surface the rate-limiting step in the OH attack on these polymers. For these and all other polymers, corona-generated OH is the dominant degradation initiator and corona-generated O<sub>3</sub> is negligible.

How do the corona-generated UV flux and OH flux compare? From our laboratory studies (22), corona generate an OH production rate of 3×10<sup>14</sup> molecules s<sup>-1</sup> for a UV flux of 5×10<sup>5</sup> photons cm<sup>-2</sup> s<sup>-1</sup> for wavelengths < 200 nm at a distance of 22 cm. We can scale both this UV flux and this OH production rate to determine the UV flux and the OH flux close to the corona (~1 cm) so that they can be compared.

For the OH flux, scaling data from two laboratory studies can be used to produce the estimate. The OH production rate in Fig. 5 has a peak HO<sub>x</sub> mixing ratio of 3.5 ppbv, of which 1.7 ppbv is OH, with an OH production rate of 2×10<sup>15</sup> OH s<sup>-1</sup>. In a separate study (22), the OH production rate is compared to UV flux. For a UV flux that will be discussed in the next paragraph, the OH production rate was 3×10<sup>14</sup> OH s<sup>-1</sup>. Thus, the peak OH mixing ratio that corresponds to the OH production rate for the measured UV flux is (3×10<sup>14</sup>/2×10<sup>15</sup>) (1.7)(2.4×10<sup>10</sup>) = 6×10<sup>9</sup> OH cm<sup>-3</sup>. Assuming that the insulator is a cylinder of diameter 5 cm, we can use the derivation of molecular flux to or from a cylindrical surface (23) to estimate the OH flux to the surface. It is 1.4×10<sup>10</sup> OH cm<sup>-2</sup> s<sup>-1</sup>.

For the UV flux, the OH production rate of 3×10<sup>14</sup> OH s<sup>-1</sup> corresponded to a corona UV flux in the 180-200 nm wavelength band of 5×10<sup>5</sup> photons cm<sup>-2</sup> s<sup>-1</sup>, at a distance of 22 cm (22). This photomultiplier detection was calibrated against a NIST-calibrated Si photodiode. Scaling this number to 2 cm from the corona, which was the distance used to find the OH flux, gives 6×10<sup>7</sup> photons cm<sup>-2</sup> s<sup>-1</sup>. The measured UV flux in the 180-260 nm band was 1.5×10<sup>9</sup> photons cm<sup>-2</sup> s<sup>-1</sup> and for the 180-290 nm band was 6×10<sup>9</sup> photons cm<sup>-2</sup> s<sup>-1</sup>. The total UV corona flux in the 300-350 nm band is several hundred times greater (23).

The OH and UV radiation fluxes are the same order of magnitude. It is not yet clear if these flux values are appropriate for corona on HTLVs, so they must be scaled once that information is

available. However, our laboratory studies show that they do scale together so that their relative value remains the same. Thus the following discussion will apply no matter how the amount of corona on HVTLS differs from our laboratory results.

Consider first a surface made of EPDM. Some of its C-C and C-H bonds can be broken by UV radiation at wavelengths less than  $\sim$ 350 nm, while most require radiation wavelengths less than 280 nm (25). Thus, corona-generated UV radiation can be responsible for EPDM degradation. OH readily reacts with ethylene and propylene in the gas-phase so that OH can also be responsible for EPDM degradation. However, the environmental degradation of EPDM is most likely initiated by solar UV radiation, with corona-generated UV radiation and OH playing a secondary role, because the solar UV flux in the 300-350 nm range ( $\sim 10^{15}$  photons  $\text{cm}^2 \text{ s}^{-1}$ ) is orders of magnitude larger than either the corona-generated UV flux or the OH flux.

Surfaces made of PDMS have been shown to be more resistant to degradation, including by UV radiation (20), which is one reason that PDMS is a polymer of choice for HVTL applications. This result is consistent with the stability of gas-phase organosilicon compounds under 20 minutes of maximum simulated solar radiation in an environmental chamber (21). The most exposed bonds in PDMS – the Si-O and  $\text{H}_2\text{C}-\text{H}$  bonds – require radiation wavelengths less than 260 nm for photodissociation. Thus, solar UV radiation is less effective for PDMS degradation than it is for EPDM degradation.

The appropriate UV flux for corona-generated UV less than 260 nm is  $1.5 \times 10^9$  photons  $\text{cm}^{-2} \text{ s}^{-1}$ , which is less than the simultaneous OH flux of  $6 \times 10^9$  molecules  $\text{cm}^{-2} \text{ s}^{-1}$ . Both UV radiation and OH initialize PDMS degradation by removing an H from a methyl group (16,21). After the H is abstracted, whether it be by UV radiation or OH reaction, the resulting degradation sequence leads to insulator failure (16).

For PDMS, the solar UV flux is negligible below 290 nm, so that the corona-generated UV flux and OH flux can be more important than solar UV flux for PDMS degradation. Since these two fluxes are the same order of magnitude, more information on the efficiency of each flux removing an H from the methyl groups is required to know whether corona-generated UV radiation or OH is more important. Several research studies have described PDMS degradation in detail, but we have not found one that quantifies the mechanism of photodissociation. Further, we found no studies that even mention that corona-generated OH might be involved, even though corona-generated OH is known to scientists developing methods to remove environmental pollution (26). The modeling or laboratory experiments to quantify the surface chemistry involving corona-generated UV radiation and OH, so that their relative importance can be compared, is well beyond the scope of this paper. However, the analysis presented here demonstrates that corona-generated OH is more likely to degrade PDMS and other polymers than corona-generated UV radiation is.

### 2.3. Uncertainty estimates for these impact calculations

The uncertainty of these estimates of the impact of corona-generated OH are based on our observations during TRAMP and SHARP on the GTHOS inlet and on our recent laboratory studies. All uncertainty estimates are given in terms of multiplicative factors and are at 68% confidence.

For the lightning rods, we estimate the following uncertainties as multiplicative factors: number of thunderstorms: 1.3; corona duration: 1.3; coronal OH production: 1.5; fraction reacting with corona-generated  $\text{HO}_2$ : 1.3; and annual ambient OH production: 2. The total uncertainty for the corona-to-ambient OH is a factor of 2.6, which we round up to 3 to account for unknown unknowns. For urban scale, our estimate of the uncertainty for total number of lightning rods in a mid-sized urban area is a factor of 3, giving the total uncertainty for ratio of corona-to-ambient OH of a factor of 4.2.

For the high-voltage electrical power transmission lines (HVTLS), we estimate the following uncertainties: ozone production: 3; O<sub>3</sub> to OH scaling: 1.3; ambient OH concentration: 2. The total uncertainty for HVTLS corona-to-ambient OH is a factor of 3.6, which we round up to 4 to account for unknown unknowns.

For surface oxidation and degradation, we estimate the following uncertainties for the OH flux: OH concentration: 2; boundary layer thickness: 2; number of exposed methyl groups: 3. The total uncertainty estimate in the rate of surface oxidation initiated by corona-generated OH is 4.1. For the corona UV flux at wavelengths below 200 nm, we estimate the following uncertainties as multiplicative factors: 187 nm flux: 2; ratio 200-300 nm flux to 187 nm flux: 4; scaling flux to a nearby surface: 2. The total uncertainty estimate in the corona UV flux is a factor of 5. We have no way to estimate the uncertainties for detailed mechanisms for the degradation initiation by corona-generated UV radiation or OH.

These uncertainty estimates are all in the range of a factor of 4 to 5 at 68% confidence. There are also unknown uncertainties associated with applying these uncertainty estimates to other locations and times. Planned laboratory and field measurements will cut these uncertainty estimates substantially.

### **3. Data access**

The data sets and information on the modeling code used in this study are publicly available at: <https://doi.org/10.26208/k4ce-zw19>

#### **3.1. TRAMP measurement files**

**1 file: TRAMP\_merge\_meas\_mod\_version20211230.csv**

The merge of 1-minute data from all TRAMP measurements for all times is a csv file.. The header gives the name of the variable and its units. Note that the extreme values for observed OH and HO<sub>2</sub> have been removed. The merge also contains the OH and HO<sub>2</sub> modeled with RACM (S4). “-9999” indicates no measurement. These data were used to plot Fig. 1C and 1D.

**18 files: TRAMP\_2006MMDD\_version\_20211211.csv**

The merges of 5-Hz data from TRAMP measurements for only a few hours around the times of the extreme OH and HO<sub>2</sub> are csv files. MMDD is the month and day that the file record starts, but each file typically runs from morning of the first day to the morning of the next day. The header gives the names of the variables and their units. TRAMP measurements in the 1-minute merge are interpolated to 5-Hz. In addition to these measurements and the 5-Hz OH and HO<sub>2</sub> measurements, are the following variables:

GTHOS\_Cal: “1” indicates an hourly period of instrument testing that started at 30 minutes past each hour.

GTHOS\_OHRay: Counts/0.2s is the OH scattering signal caused by rain entering the GTHOS inlet.

GTHOS\_OnLine: “1” indicates times when the laser wavelength is tuned to the OH absorption line enabling the measurements of OH and HO<sub>2</sub>. “0” indicates background measurements.

GTHOS\_Scan: “1” indicates periods when the laser wavelength is being scanned, producing the OH spectrum around the line that is being used.

“-9999” indicates no measurement.

#### **3.2. Cloud and weather information**

**2 GOES movies of Infrared Water Vapor and Visible for Houston area**

The two avi files contain movies of the Houston area. Intermixed are the Infrared Water Vapor and the Visible channels. The dates covered in each file are listed in the file name. These data are used in Table S1.

GOES\_IRWV\_Vis\_819\_822\_823\_911\_915\_916\_917\_924.avi  
GOES\_IRWV\_Vis\_826\_827\_901\_902\_910

### **3.3. Surface weather data from the NWS site, William Hobby Airport, can be found at the following website: <https://www.weather.gov/wrh/Climate?wfo=hgx>**

The following settings were used to retrieve the data:

Location: Houston, William, TX  
Date: first 2006-8; second 2006-9  
Product: Daily data for a month

### **3.4. Photochemical Box Modeling**

The code for the photochemical box modeling can be best obtained from <https://github.com/AirChem/F0AM>, a site maintained by Glenn Wolfe. The code is based on Matlab, so the reader will need to have access to Matlab..

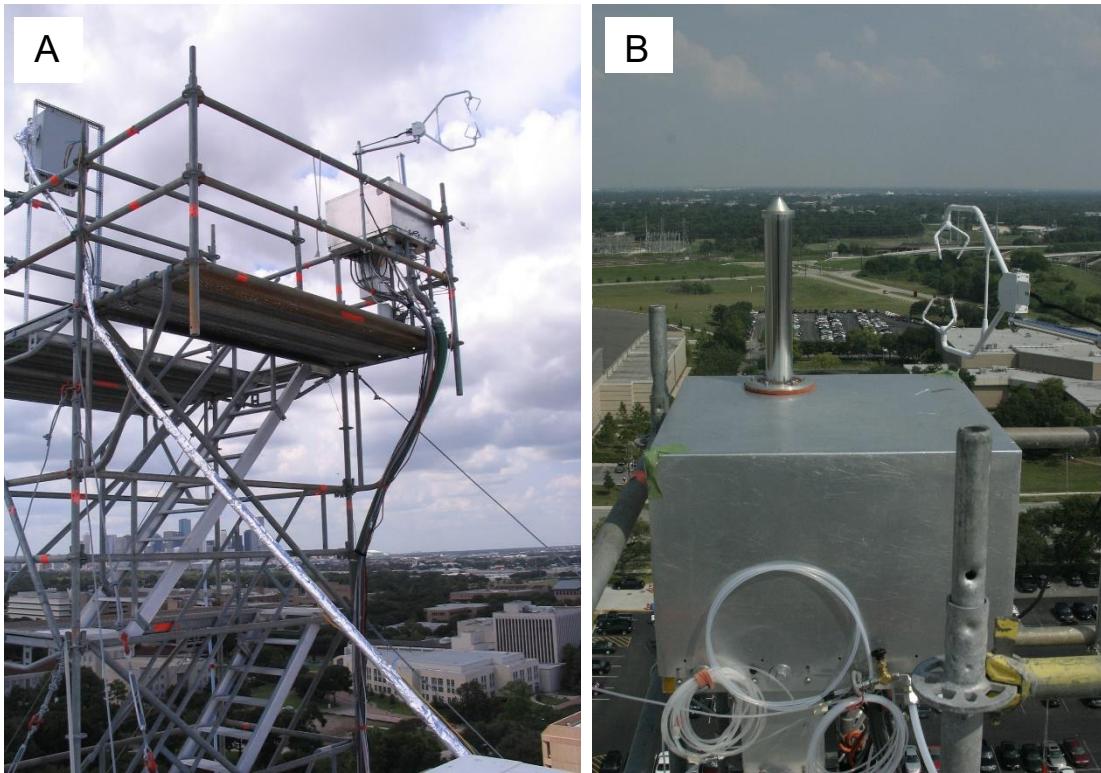
Model parameters were set as follows:

```
kdil = 10^-12 ; % Hz
ModelOptions.Verbose      = 1;
ModelOptions.EndPointsOnly = 0; % we want the calculations as a function of time
ModelOptions.LinkSteps    = 0;
ModelOptions.Repeat       = 1;
ModelOptions.IntTime      = 120; % seconds
ModelOptions.TimeStamp    = 0.0;
```

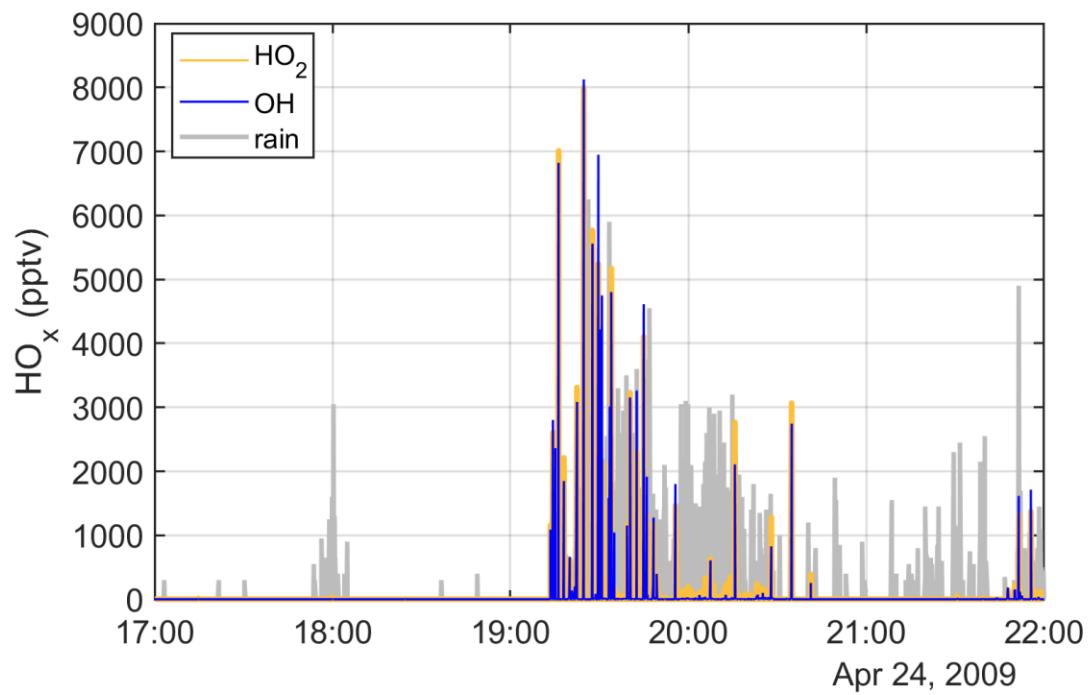
The input files to run the 17 different cases are in the file

TRAMP\_photochemical\_box\_model\_inputs.csv.

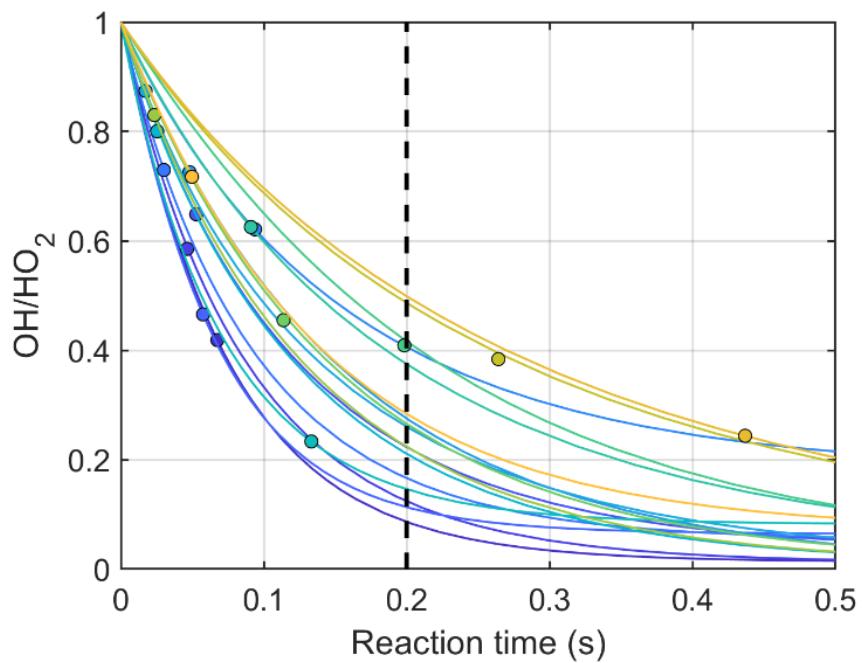
The chemical mechanism was MCM v331.



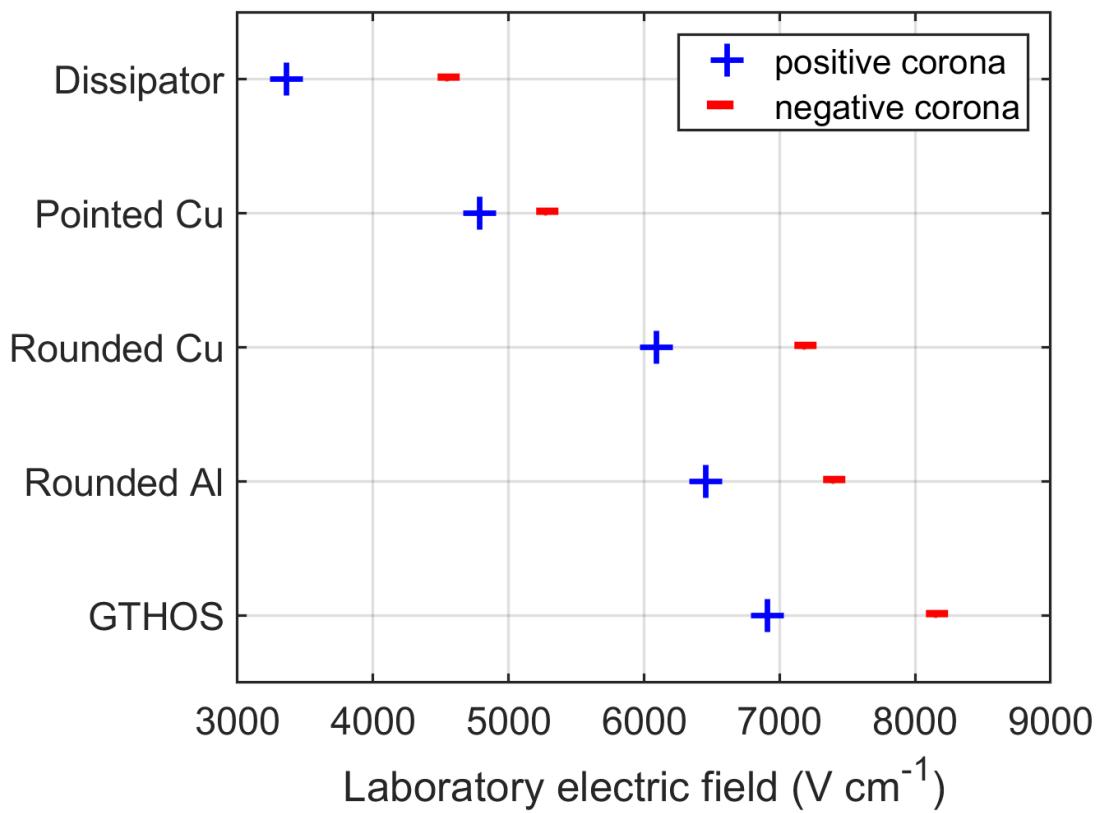
**Figure S1.** GTHOS on top of Moody Tower I during TRAMP in 2006. (A) The GTHOS sampling inlet was 5 m above the roof surface on construction scaffolding and located in the northeast corner of the scaffolding, which was at the north edge of the building. The sonic anemometer was oriented to the east (90°). (B) The GTHOS sampling inlet (5.5 cm dia.) is 40 cm above the flat surface of the GTHOS detection system. Nearby vertical structures include the extension tubes for the scaffolding and the sonic anemometer, although the GTHOS inlet was higher than all of them.



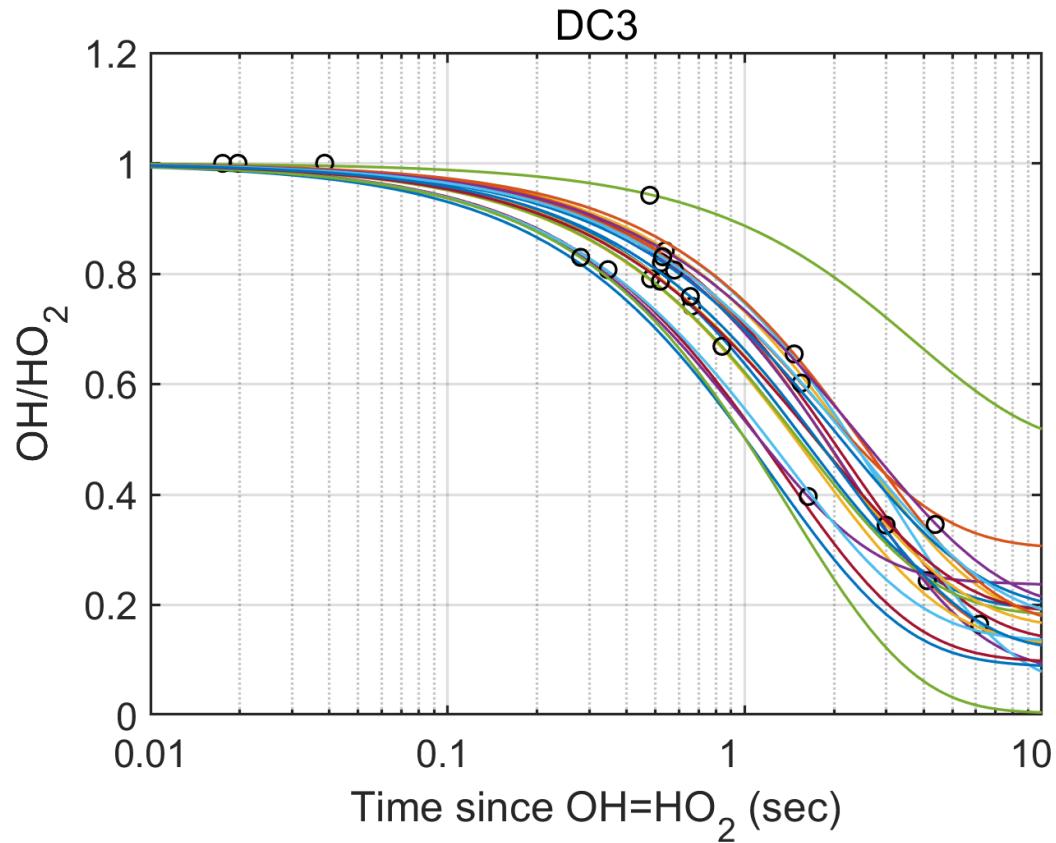
**Figure S2.** Enhanced HO<sub>x</sub> observed during the 2009 SHARP study. HO<sub>2</sub> (gold), OH (blue), and scattering indicating rain (gray) all consist of individual 0.2 s spikes. HO<sub>2</sub> and OH spikes always coincide and these spikes rarely coincide with the OH Rayleigh scattering spikes.



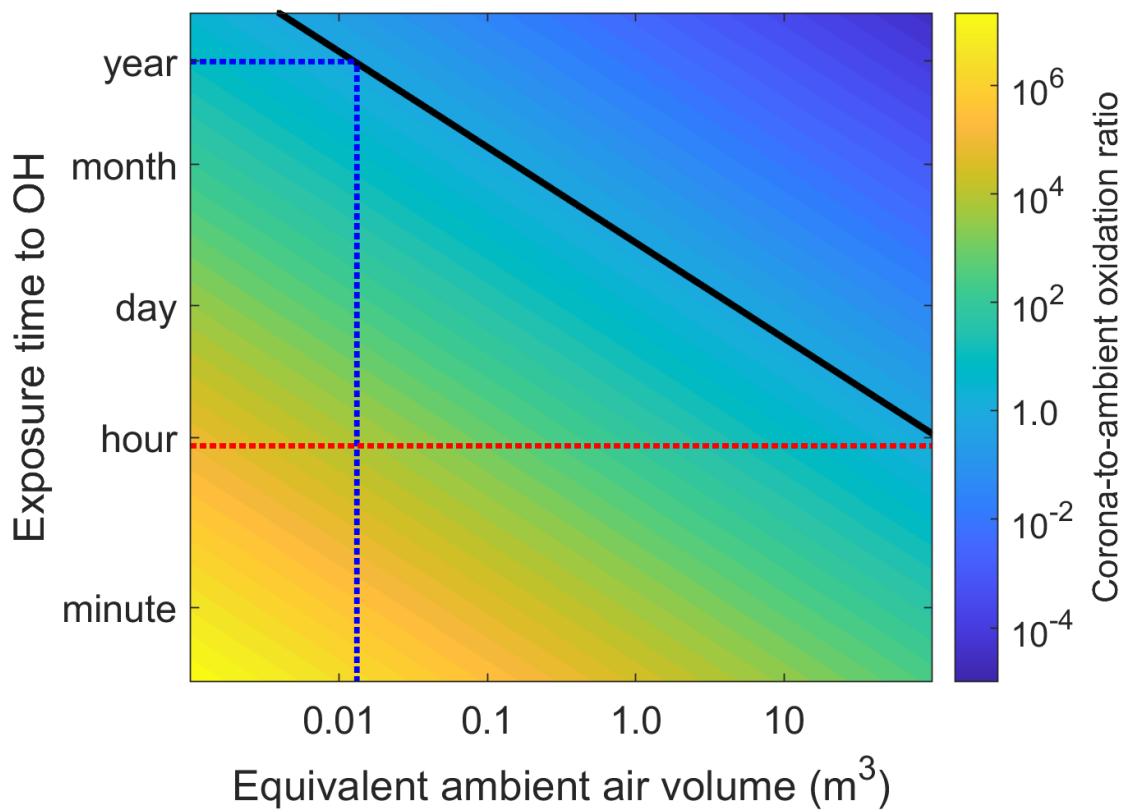
**Figure S3.** Modeled change in OH/HO<sub>2</sub> for 17 LHO<sub>x</sub> spikes during TRAMP. Lines are the modeled ratios as a function of time, assuming [OH]=[HO<sub>2</sub>] initially. Dots are the measured ratios of the LHO<sub>x</sub> spikes. Fifteen are achieved in less than 0.2 seconds, which is the time resolution of the measurements. The different colors indicate whether the measurements were early (cooler color) or late (warmer color) in the field study.



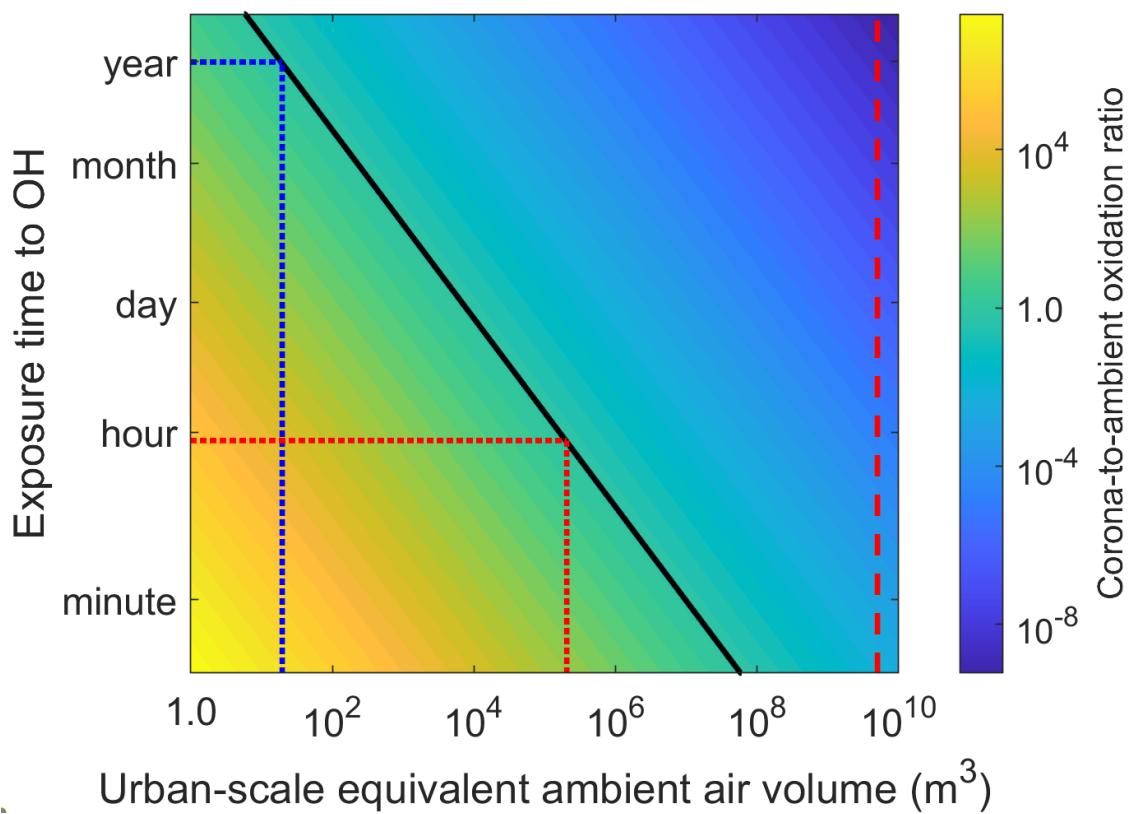
**Figure S4.** Electric field required to initiate coronas for different electrodes. The electrodes are a static dissipator, pointed copper lightning rod, rounded copper lightning rod, rounded aluminum lightning rod, and the GTHOS inlet. The same trends hold for positive (+) and negative (-) coronas.



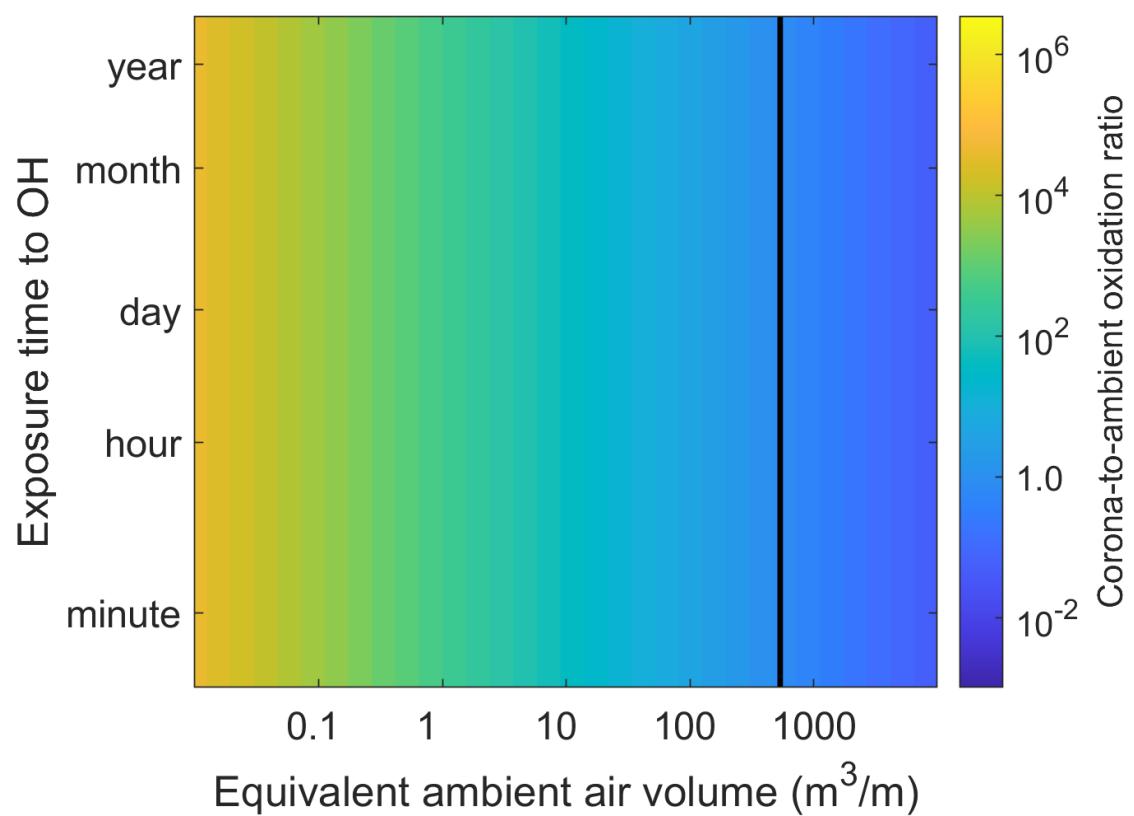
**Figure S5.** Modeled change in  $\text{OH}/\text{HO}_2$  for 22  $\text{LHO}_x$  spikes during DC3. Lines are the modeled ratios as a function of time, assuming  $[\text{OH}]=[\text{HO}_2]$  initially. Open circles are the measured ratios of the  $\text{LHO}_x$  spikes. The different colors indicate whether the measurements were early (blue) or late (yellow) in the field study. Nineteen encounters occur more than 0.2 seconds after the electrical discharge occurred, with the median time of 0.56 seconds.



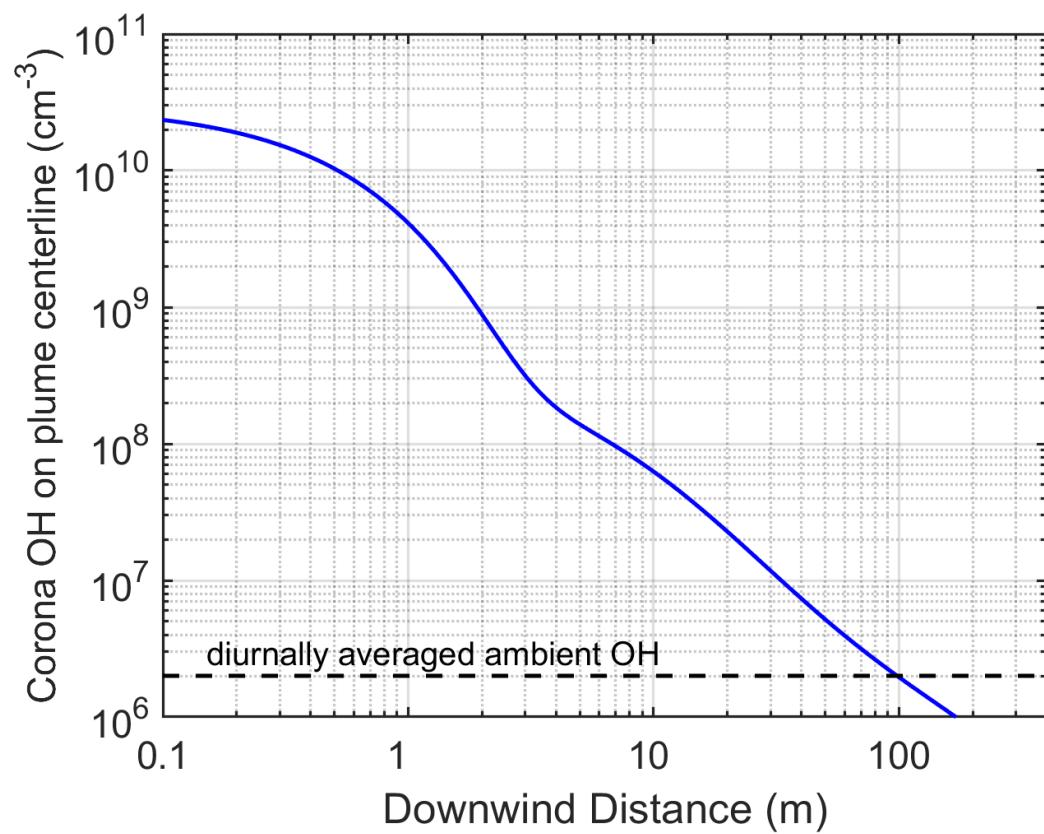
**Figure S6.** Comparison of corona-to-ambient ratio of total OH molecules for one lightning rod during a total of 50 storms. The total number of OH molecules produced during the storms by corona in the small corona volume can be compared to the total number of OH molecules produced in different air volumes and for different times. At small volumes and short times (lower left), corona produce much more OH than ambient conditions. The ambient and corona OH become equal (black line). At larger volumes and times (upper right), OH from corona are insignificant compared to ambient OH. During the storms, corona-produced OH is greater than ambient OH produced in  $100 \text{ m}^3$  of air (red dotted line), but on an annual basis, is greater only near the lightning rod (blue dotted line).



**Figure S7.** Comparison of corona-to-ambient ratio of total OH molecules for all lightning rods in a mid-sized urban center under 50 storms. The total number of OH molecules produced during the storms by corona in the small corona volume can be compared to the total number of OH molecules produced in different air volumes and for different times. At small volumes and short times (lower left), corona produce much more OH than ambient conditions. The ambient and corona OH become equal (black line). At larger volumes and times (upper right), OH from corona are insignificant compared to ambient OH. During the storms, corona-produced OH is greater than ambient OH produced in  $100 \text{ m}^3$  of air (red dotted line), but on an annual basis, is greater only near the lightning rod (blue dotted line). The air volume for the urban center is the dashed red line.



**Figure S8.** Comparison of corona-to-ambient ratio of total OH molecules for continuous corona on a high-voltage electrical power transmission line (HVTL) per meter of line. The total number of OH molecules produced by corona in the small corona volume on and downwind of the HVTL can be compared to the total number of OH molecules produced in different air volumes and for different times. For smaller volumes (left), corona produce much more OH than ambient conditions. The ambient and corona OH become equal (black line). At larger volumes (right), OH from corona are insignificant compared to ambient OH.



**Figure S9.** Centerline OH concentration in a Gaussian plume downwind of corona along a high-voltage electrical power transmission line. Corona-produced OH decay is due to both chemistry and dispersion. Assumed was a windspeed of  $3 \text{ m s}^{-1}$  and Pasquill dispersion class of moderately unstable. Corona-produced equals ambient OH at 100 meters downwind, where the Gaussian dispersion vertical standard deviation is  $\sim 10 \text{ m}$ .

	Hobby Airport <sup>1</sup>					
Date	Rain total (in)	LHO <sub>x</sub> time (GMT = CST + 6 hours)	OH Rayleigh scattering time (GMT)	overhead cloud brightness T (K±5K)	cloud type Cb: cumulonimbus Sc: stratocumulus	enhanced HO <sub>x</sub> behavior <sup>2</sup>
11-18 Aug	0					
19-Aug	1.93	13:38-14:18 15:48-16:00	13:55-14:11 15:39-16:00	195; 205	extensive Cb; Cb	spikes; spikes
20-21 Aug	0	0	0			
22-Aug	0	0	21:01-21:20	235	edge of small Cb	
23-Aug	0.16	17:52-19:24	18:20-21:38	205	persistent Cb	spikes
24-25 Aug	0	0	0			
26-Aug	Trace	0	17:18-17:46	225	edge of small Cb	
27-Aug	0.02	0	17:11-17:25	220	line of small Cbs	
28-31 Aug	0	0	0			
1-Sep	Trace	23:58-24:00	23:55-24:00	215	Cb	variable continuum
2-8 Sep	0	0	0			
9-Sep	0.09	0	0			
10-Sep	0.23	21:44-22:30	21:30-22:30	225	small Cb	weak spikes
11-Sep	0.08	19:43-19:49	19:12-20:26 22:13-24:00	215	Cb	weak spikes then variable continuum
12-Sep	0.01		0			
13-14 Sep	0		0			
15-Sep	0.77	21:07-21:45	20:40-22:00	205	Cb	variable continuum
16-Sep	0	0	21:50-22:30	230	small Cb to north	
17-Sep	0.59	0	14:24-14:53 17:24-17:35	220	widespread Sc	
18-Sep	0.06	0	0			
19-22 Sep	0	0	0			
23-Sep	0.78	0	0			
24-Sep	0.47	4:17-8:43	4:40-11:45	195	widespread large Cbs	spikes, then continuum; then spikes
25-28 Sep	0	0	0			

<sup>1</sup> Hobby Airport located 9.4 km southeast of the site

<sup>2</sup> Spikes are short enhancements (e.g., Fig 2A); variable continuum is longer unsteady enhancements (e.g., Fig. 2B and 2C).

**Table S1.** Weather and times with LHO<sub>x</sub> and GTHOS Rayleigh scattering for TRAMP.

Measurements used to constrain the photochemical box model			
Environment	Inorganics	VOCs	VOCs
Pressure	O <sub>3</sub>	HCHO	Ethene
Temperature	NO	CH <sub>3</sub> OOH	Propene
Water vapor	NO <sub>2</sub>	Acetaldehyde	Butenes
Photolysis frequencies	CO	Ethane	Pentenes
	HNO <sub>3</sub>	Propane	Hexenes
	HOOH	Butanes	Benzenes
		Pentanes	Xylenes
		Hexanes	Toluenes
		Acetylene	Isoprene

**Table S2.** Measurements used to constrain the photochemical box model

## SI References

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