

1 CO₂-rich rejuvenated volcanic rocks on Hawaiian islands

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13 Abstract

14 **Sedimentary carbonates are sent to the deep mantle if they are not completely destroyed at**
15 **subduction zones, and subducted carbonates may contribute to plume volcanism. To better**
16 **constrain the role of recycled carbonates in Hawaiian volcanism, we report high-precision**
17 **olivine and whole-rock geochemical compositions of shield and rejuvenated stage lavas**
18 **from Kauai, Oahu and Maui islands. The studied rejuvenated stage whole-rocks have low**
19 **SiO₂ and high CaO concentrations, and are depleted in HFSEs, such as, Nb and Zr,**
20 **consistent with a role of carbonated melt. Rejuvenated stage olivines have Ni abundance**
21 **lower than and CaO and MnO contents similar to those of shield stage at a given Fo. The**
22 **calculated partition coefficients of Ca (D_{Ca}^{Ol-melt}) and Mn (D_{Mn}^{Ol-melt}) between olivine and**
23 **shield melts are consistent with those of a dry melt system. However, the low D_{Ca}^{Ol-melt} and**
24 **D_{Mn}^{Ol-melt} for rejuvenated lavas can only be explained by a volatile-rich melt system. Based**
25 **on the observed D_{Ca}^{Ol-melt} and D_{Mn}^{Ol-melt} in rejuvenated lavas, and considering the effect of**
26 **H₂O, our modeling calculation shows that rejuvenated primary magmas contain up to ~10**

27 **wt% CO₂. Using olivine-spinel aluminum exchange thermometry, we show that the**
28 **rejuvenated primary magma temperatures are similar to/ slightly lower than those of shield**
29 **stage. We posit that the lower rejuvenated stage lavas were originated from melting of**
30 **carbonated peridotites in the plume. The carbonated melts that metasomatized the**
31 **peridotites were likely derived from subducted ancient carbonate-bearing lithospheric**
32 **mantle.**

33 **Keywords:** olivine, CO₂, metasomatism, rejuvenated lavas, mantle plume, Hawaiian islands

34 **1 Introduction**

35 The deep mantle may store most of the Earth's carbon (Dasgupta and Hirschmann, 2006;
36 Dasgupta and Hirschmann, 2010; Zhang et al., 2017; Plank and Manning, 2019). Plate
37 subduction is considered to be an important factor of carbon enrichment in the deep Earth's
38 interior (Dasgupta and Hirschmann, 2010; Plank and Manning, 2019). Deep-rooted mantle
39 plumes (e.g., Hawaiian plume) are good candidate for the study of deep Earth's carbon cycle
40 (Dixon et al., 2008). Since CO₂ can significantly change the chemistry of mantle-sourced melts,
41 e.g., lowering SiO₂ and elevating CaO of melts (Dasgupta et al., 2007a; Foley et al., 2009; Zhang
42 et al., 2017), it may play an important role in the origin of alkali basalts (Dixon et al., 2008;
43 Dasgupta et al., 2006, 2007a&b; Sisson et al., 2009).

44 The Emperor-Hawaii seamount chain is a type volcanic chain that may originate from near the
45 core-mantle boundary (Montelli et al., 2004; Huang et al., 2011; Weis et al., 2011; French and
46 Romanowicz, 2015). Shield stage tholeiitic basalts account for the majority (95 vol%) of
47 Hawaiian volcanic rocks, while pre-shield, post-shield and rejuvenated stages account for the
48 rest 5% (Sherrod et al., 2007). The role of an olivine-free lithology, pyroxenite or eclogite, in the
49 petrogenesis of Hawaiian shield lavas is highly debated. Specifically, Hawaiian shield tholeiitic
50 lavas have too high SiO₂ content to be produced by partial melting of garnet peridotite (e.g.,
51 Hauri, 1996; Wagner and Grove, 1998; Huang and Frey, 2005). This led to the suggestion that an
52 eclogite component play a role in producing the high SiO₂ contents in Hawaiian tholeiitic lavas
53 (Hauri, 1996; Huang and Frey, 2005). Alternatively, the high SiO₂ content may be a result of
54 melt-harzburgite reaction (Wagner and Grove, 1998). Sobolev et al. (2005; 2007) and Herzberg
55 (2006; 2011) noted that the high olivine Ni contents and low CaO and MnO in olivines from

56 Hawaiian shield lavas, as well as low whole rock CaO contents, required a pyroxenite-dominated
57 mantle source for Hawaiian shield lavas. Alternatively, the low CaO and MnO contents, and high
58 SiO₂ content of Hawaiian lavas may be explained if they represent mixtures of partial melts from
59 garnet peridotite and from eclogite (Huang et al., 2007). Matzen et al. (2013; 2017) showed that
60 the high olivine Ni content in Hawaiian shield stage olivines may simply reflect the temperature
61 difference between partial melting beneath the thick lithosphere and olivine crystallization in a
62 shallow magma chamber, and a pyroxenite-dominated mantle source is not required.

63 Hawaiian rejuvenated stage volcanism occurred 0.5-2 Myr after shield stage and consists of
64 silica-undersaturated alkaline-rich rocks, e.g., alkali basalt, hawaiite, nephelinite and melilitite
65 (e.g., Clague and Dalrymple, 1987; Garcia et al., 2010; Phillips et al., 2016). Hawaiian
66 rejuvenated lavas are among the most silica-undersaturated and alkaline-rich in oceanic
67 islands/seamounts (Chauvel et al., 1997; Zhang et al., 2020). It remains unclear on the origin of
68 rejuvenated stage low-SiO₂ alkali-rich volcanic rocks (Reiners and Nelson, 1998; Bianco, 2005;
69 Garcia et al., 2010; Ballmer, 2011; Hofmann and Farnetani, 2013; Phillips et al., 2016).
70 Hawaiian rejuvenated lavas are more depleted in Sr-Nd-Hf isotope compositions than those of
71 shield stage (e.g., Chen and Frey, 1983; Frey et al., 2000; 2005 G-cubed; Yang et al., 2003;
72 Hofmann and Farnetani, 2013; Beguelin et al., 2019; DeFelice et al., 2019; Harrison et al., 2020),
73 indicating different mantle sources. Earlier studies suggested an origin of Hawaiian rejuvenated
74 lavas from the underlying metasomatically enriched Pacific lithospheric mantle (Chen and Frey,
75 1983; Garcia et al., 2010). However, there are increasing studies suggesting the isotopically
76 depleted rejuvenated stage source component is intrinsic of the mantle plume (Ribe and
77 Christensen, 1999 EPSL 171, 517-531; Hofmann and Farnetani, 2013; DeFelice et al., 2019;
78 Harrison et al., 2020; but see also Beguelin et al., 2019 for a different opinion).

79 Silica under-saturated alkali lavas could be explained by involvement of CO₂ in the mantle
80 source (Mallik and Dasgupta, 2012; 2013). Several studies have indicated that CO₂ played an
81 important role in the origin of Hawaiian lavas (Dixon et al., 2008; Huang et al., 2009&2011;
82 Sisson et al., 2009; Borisova and Tilhac, 2021). Tucker et al (2019) showed that the mantle
83 sources of Hawaiian shield stage lavas contain 380-480 ppm CO₂ based on measurement of
84 olivine hosted melt inclusions, while the source CO₂ contents of rejuvenated lavas were not
85 constrained. Dixon et al (2008) explained Hawaiian rejuvenated lavas with elevated Ba/Th by

86 carbonatite metasomatism, but the content of CO₂ has not been directly constrained. CO₂ not
87 only affects magma chemical compositions, but also element partitioning between olivine and
88 melts (Gavrilenko et al., 2016). Gavrilenko et al (2016) proposed that magma CO₂ concentration
89 could be calculated based on its influence on Ca partition between olivine/melt. In this study, we
90 have analyzed high-precision olivine chemistry and whole-rock major and trace elements of
91 Hawaiian rejuvenated stage lavas, aiming to evaluate the role of CO₂ in Hawaiian magmatism.

92 **2 Samples**

93 Shield and rejuvenated stage rock samples in this study were collected from Kauai, Oahu and
94 Maui Islands (Fig. 1), and their information is in [Supplementary Table 1](#). We have collected 26
95 rock samples, including 16 rejuvenated stage and 10 shield stage, from these islands (Fig. 1;
96 [Supplementary Table 1](#)). At Kauai Island, there are two sampling sites from the shield stage
97 Waimea volcanics (5.5-4 Ma), and 5 sites from the rejuvenated stage Koloa volcanics (2.6-0.15
98 Ma). At Oahu Island, there are 2 sampling sites from the shield stage Wai'anae volcanics and 1
99 site from the shield stage Ko'olau volcano (3.0-1.8 Ma); 2 sampling sites from the rejuvenated
100 stage Honolulu volcanics (0.8-<0.1 Ma). At Maui Island, there is 1 sampling site from the shield
101 stage Wailuku volcanics (2.0-1.3 Ma); 2 sites from the rejuvenated stage Hana volcanics (0-1.5
102 Ma) and 1 site from the rejuvenated stage Kula volcanics (0.93-0.15 Ma). The detailed
103 information on division of volcanics can be referred to [Sherrod et al \(2007\)](#). Despite the absence
104 of distinct differences in formation age between the Hana and Kula volcanics, we classify the
105 Hana and Kula volcanics as rejuvenated stage because of their low-silica and high alkaline
106 characteristics. The detailed information on sampling can be found in [Supplementary Table 1](#).

107 **3 Methods and results**

108 **3.1 Methods**

109 The 26 rock samples were analyzed for bulk-rock major and trace elements, and their olivine
110 phenocrysts were analyzed for major and trace elements using in situ high-precision Electron
111 Probe Microanalyzer (EMPA) technique. Bulk-rock major elements were analyzed using fused
112 glass discs with an Axios sequential X-ray Fluorescence Spectrometer at Institute of Geology
113 and Geophysics, Chinese Academy of Sciences (IGGCAS), Beijing, China. Samples were fused

114 at 1050 °C using a lithium tetraborate flux ($\text{Li}_2\text{B}_4\text{O}_7$) in a mixture consisting of 0.5 g of sample
115 and 5 g of lithium tetraborate. Loss on ignition (LOI) was determined at a temperature of 1000°C
116 in air for 3 hours. Basalt standards BCR-2, BHVO-2 and GSR-3 were analyzed as unknowns and
117 the results are shown in [Supplementary Method Table 1](#). Bulk-rock trace elements were analyzed
118 by a Perkin-Elmer Sciex ELAN DRC-e ICP-MS at the State Key Laboratory of Ore Deposit
119 Geochemistry (SKLODG), Institute of Geochemistry, Chinese Academy of Sciences (IGCAS).
120 The powdered samples (50 mg) were dissolved with $\text{HF} + \text{HNO}_3$ mixture in high-pressure Teflon
121 Bombs at ~190 °C for 48 h ([Qi et al., 2000](#)). Rh was used as an internal standard to monitor
122 sensitivity drift during measurement. How do you measure the trace element abundances?
123 BCR-2 and BHVO-2 solutions were used to monitor analytical accuracy. The analytical
124 precision was generally better than 10 % (2 sigma) based on replication of basalt standards
125 analyses.

126 Quantitative *in situ* analyses of olivine and spinel major and minor elements were conducted
127 on JXA-8230 EMPA equipped with 5 wavelength dispersive spectrometers at Institute of
128 Oceanology, Chinese Academy of Sciences. The operating conditions were: 20 kV accelerating
129 voltage, 60 nA beam current, and 1-5 μm beam diameter. For olivine analyses, the counting time
130 was 30 seconds for Si $\text{k}\alpha$, Mg $\text{k}\alpha$ and Fe $\text{k}\alpha$, and 80 seconds for Mn $\text{k}\alpha$, Ni $\text{k}\alpha$, Ca $\text{k}\alpha$, Cr $\text{k}\alpha$, Al
131 $\text{k}\alpha$. The off peak counting time was 80 seconds for Mn $\text{k}\alpha$, Ni $\text{k}\alpha$, Ca $\text{k}\alpha$, Cr $\text{k}\alpha$, Al $\text{k}\alpha$. Standards
132 used were olivine (Si, Mg), corundum (Al), MnO (Mn), wollastonite (Ca), NiO (Ni), and apatite
133 (P), magnetite (Fe), and Cr_2O_3 (Cr). For spinel analyses, the counting time was 30 seconds for Si
134 $\text{k}\alpha$, Mg $\text{k}\alpha$, Al $\text{k}\alpha$, Mn $\text{k}\alpha$, Ni $\text{k}\alpha$, Cr $\text{k}\alpha$, Ti $\text{k}\alpha$ and Fe $\text{k}\alpha$. The off peak counting time was 20
135 seconds for Si $\text{k}\alpha$, Mg $\text{k}\alpha$, Al $\text{k}\alpha$, Mn $\text{k}\alpha$, Ni $\text{k}\alpha$, Cr $\text{k}\alpha$, Ti $\text{k}\alpha$ and Fe $\text{k}\alpha$. Standards used were
136 spinel (Al, Mg), quartz (Si), MnO (Mn), rutile (Ti), NiO (Ni), magnetite (Fe) and Cr_2O_3 (Cr).
137 Unknown and standard intensities were corrected for dead time. All data were corrected with the
138 standard ZAF correction procedures. The olivine standard MongOLSh11-2 ([Batanova et al.,](#)
139 [2019](#)) was replicated during the analyses to monitor the reproducibility and accuracy, and the
140 results are shown in [Supplementary Method Table 3](#).

141 **3.2 Results**

142 **3.2.1 Whole-rock major and trace elements**

143 Whole-rock major and trace element compositions are shown in [Supplementary Table 2](#). As

144 shown in the plot of TAS vs. SiO₂ ([Supplementary Fig. 1](#)), shield stage lavas are tholeiitic basalts
145 with two plotting close to the division line between alkali basalt and tholeiitic basalts. Shield
146 stage lavas have SiO₂ between 46.4-50.0 wt%, MgO between 11.5-23.4 wt%, CaO between
147 5.1-10.7 wt%, Na₂O between 1.44-2.51 wt%, and P₂O₅ between 0.18-0.35 wt%. Shield stage
148 lavas show slight enrichment of light over heavy rare earth elements (REEs) and slight
149 enrichment of Ba. Rejuvenated stage lavas are all rich in alkaline elements and can be classified
150 into foidite, basanite and alkali basalts ([Supplementary Fig. 1](#)), of which the three foidite samples
151 from Site O7 (Honolulu volcanics) have the highest alkaline (Na₂O+K₂O) and P₂O₅ and the
152 lowest SiO₂ contents ([Fig. 2](#)). Except for Site O7 foidite samples, the other rejuvenated stage
153 samples have SiO₂ between 37.7-45.3 wt%, MgO between 7.15-15.1 wt%, CaO between
154 9.4-13.3 wt%, Na₂O between 2.4-3.8 wt% and P₂O₅ between 0.33-0.92 wt%.

155 These rejuvenated stage lavas are strongly enriched in light rare earth elements (LREEs) and
156 large ion lithophile elements (LILEs) ([Fig. 3](#)) compared with shield stage lavas. These
157 rejuvenated stage lavas also have positive anomalies of Ba and negative anomalies of Zr-Hf ([Fig.](#)
158 [3](#)). Site O7 foidite lavas have the strongest enrichment of LILEs and Th-U and the strongest
159 negative anomalies of high field strength elements (HFSEs, e.g., Nb-Zr-Hf-Ti) ([Fig. 3](#)). These
160 rejuvenated stage lavas show systematically lower SiO₂ and Al₂O₃ and higher total iron and
161 CaO/Al₂O₃ compared to shield stage lavas ([Fig. 2](#), [Supplementary Fig. 2](#)). Rejuvenated stage
162 lavas also have systematically lower Ni than those of shield stage for a given MgO content ([Fig.](#)
163 [2b](#)). Moreover, rejuvenated stage lavas have overall higher CaO and MnO contents than those of
164 shield stage for a given MgO content ([Fig. 2 b&d](#)).

165 **3.3.2 *In situ* olivine and spinel chemistry**

166 Olivine phenocrysts and olivine-spinel pairs have been analyzed for high-precision major
167 and trace element compositions by EMPA, and the results are shown in [Supplementary Table](#)
168 [3&4](#), respectively. The volcanic rock samples of shield and rejuvenated stages in this study have
169 similar ranges in Fo# (molar Mg/(Mg+Fe)×100%) ([Fig. 4](#)). For shield stage volcanic rocks, the
170 highest Fo#s of olivines are 86.8, 89.1, 88.5 and 87.4 for Waimea, Wai'anae, Ko'olau and
171 Wailuku volcanics, respectively; while for rejuvenated stage volcanic rocks, the highest Fo#s
172 of olivines are 87.5, 86.9, 87.3 and 84.1 for Koloa, Honolulu, Hana and Kula, respectively. As
173 shown in [Fig. 4](#), at a given Fo, shield stage volcanic rocks have olivine Ni contents distinctly

174 higher and Mn-Ca contents lower than those of global MORBs. The olivines of rejuvenated stage
175 lavas also have Ni contents systematically higher than those of global MORBs but lower than
176 those of shield stage lavas for a given Fo#. Despite the distinct compositions of olivine Ni
177 between shield stage and rejuvenated stage lavas, their olivine Ca and Mn contents are
178 comparable for a given Fo# (Fig. 4). The olivines with the highest Fo# for both shield and
179 rejuvenated stages have similarly Mn and Ca contents, which are lower than those of MORBs.

180 **4. Discussion**

181 **4.1 Implication of whole-rock geochemistry on the role of CO₂**

182 It has been suggested that shield stage lavas characterized by depletion of CaO and
183 enrichment of SiO₂ cannot be explained by melting of peridotites (Hauri, 1996; Wagner and
184 Grove, 1998; Huang and Frey, 2005; Herzberg, 2006; Herzberg and Asimow, 2008; Dasgupta et
185 al., 2010), but can be explained by partial melting of mixed pyroxenite (recycled oceanic crust)
186 and peridotite (e.g., Sobolev et al., 2005&2007; Huang et al., 2007; Herzberg, 2011; Herzberg
187 and Asimow, 2008; Jackson et al., 2012; Mallik and Dasgupta, 2012). Rejuvenated stage lavas
188 are usually silica-undersaturated and enriched in CaO that are distinct from shield stage
189 volcanism (Fig. 2). The enrichment of LREEs and LILEs in rejuvenated lavas (Fig. 3) could be
190 explained by low degree of melting in the mantle source and a geochemically enriched mantle
191 source. However, the high CaO and SiO₂-poor nature of rejuvenated stage lavas cannot be
192 explained by melting of dry pyroxenite (e.g., Mallik and Dasgupta, 2013) or peridotite (e.g.,
193 Dasgupta et al., 2010) at any temperature and pressure. Results of high pressure/temperature
194 experiments indicate that, with the presence of CO₂, partial melts of either peridotite or
195 pyroxenite/eclogite have high CaO and low SiO₂ contents (e.g., Dasgupta et al., 2007; Mallik
196 and Dasgupta, 2013).

197 Plate subduction is an effective way to introduce sedimentary carbonate into deep mantle
198 (e.g., Zhang and Smith_Duque, 2014; Plank and Manning, 2019; but see also Thomson et al.,
199 2016). If recycled oceanic crust is involved in the Hawaiian mantle plume, sedimentary
200 carbonates may have played a role during mantle melting, especially at low degrees of melting.
201 Based on the results of melting experiment (e.g., Mallik and Dasgupta, 2013), melting of
202 carbonate-bearing MORB can only produce melts with MgO of <8 wt%, which cannot explain

203 Hawaiian rejuvenated stage lavas with MgO of >10 wt%. Partial melts of carbonated pyroxenite
204 and peridotite mixture and carbonated peridotite have both high MgO and CaO contents, as well
205 as low SiO₂ content (e.g., [Mallik and Dasgupta, 2013](#)). Carbonatitic melts typically have high
206 contents of CaO, MnO, Ba, and REEs, and are depleted in high field strength elements (HFSEs,
207 such as Nb-Ta-Zr-Hf-Ti) (e.g., [Hoernle et al., 2002](#); [Dixon et al., 2008](#); [Dalou et al., 2009](#); [Jones
208 et al., 2013](#)). We find that Site O7 samples from the Honolulu volcanics on Oahu Island, in
209 addition to their anomalously high CaO-MnO and extremely low SiO₂, have highly enriched
210 LREEs and relative depletion in Nb-Ta-Zr-Hf-Ti ([Fig. 3](#)). These observations could be explained
211 by melting a carbonated mantle source.

212 **4.2 Role of CO₂ indicated by olivine Ca-Mn**

213 Hawaiian rejuvenated stage volcanic rocks have higher CaO and MnO contents for a given
214 MgO content than shield stage lavas ([Fig. 2](#)), while their olivine Ca and Mn contents overlap at a
215 given olivine Fo# ([Fig. 4](#)), reflecting a difference in partitioning of Ca-Mn between olivine/melt
216 during the two volcanic stages. To better understand the partitioning of Ca-Mn between
217 olivine/melt in Hawaiian lavas, we have calculated the partition coefficients of Ca and Mn
218 between olivine/melt ([Supplementary Table 5](#)). Hawaiian lavas in this study, both shield and
219 rejuvenated stages, have variable amount of olivine accumulation, and the samples with
220 clinopyroxene accumulation are excluded. A mineral-melt pair in equilibrium is required before
221 calculation of partition coefficient. Although primary melt composition can be calculated by
222 assuming a Fo of ~91, the exact Ca content of the equilibrium primary olivine is not known.
223 Most of the rejuvenated stage lavas have variable degrees of olivine accumulations, thus, a way
224 to obtain olivine-melt pairs in equilibrium is to remove the accumulated olivines. Thus, in this
225 way, we can obtain the exact Ca composition of olivine in equilibrium. Ten to twenty olivine
226 grains with the highest Fo# and their average chemical compositions were used as the final
227 equilibrium olivine ([Supplementary Table 5](#)). These final equilibrium olivines have Fo# of 84-87
228 for rejuvenated stage lavas and 83-89 for shield stage lavas. We have removed this average
229 olivine composition step-by-step from the melt until the resulted melt was in equilibrium with
230 the average olivine based on PRIMelt3 program ([Herzberg and Asimow, 2015](#)), in which a melt
231 Fe²⁺/Fe_{tot} is specified. The resulted equilibrium melt compositions are shown in [Supplementary
232 Table 5](#). The equilibrium melts calculated for rejuvenated stages lavas have MgO of 8-12 wt%,

233 while the equilibrium melts of shield stage have MgO of 10-14 wt%. It is clear that the
234 calculated melts are not primary melts but evolved melts in equilibrium with the average olivines.
235 We have calculated the partition coefficients of Ca and Mn between olivine/melt based on these
236 olivine-melt pairs, which are shown in [Supplementary Table 6](#) and plotted in [Fig. 5](#). The
237 calculated olivine/melt partition coefficients of Ca and Mn for rejuvenated stage samples are
238 overall lower than those for shield stage for a given olivine Fo# and melt MgO content, despite
239 these evolved melts have degassed to certain degrees.

240 There are several observations indicating that CO₂ has played an essential role in the
241 activities of Hawaiian mantle plume. Sisson et al. (2009) argued that a carbonated garnet
242 lherzolite source is required to produce the pre-shield basanite-nephelinite lavas at Kilauea.
243 [Barsanti et al \(2009\)](#) showed that a suite of magmas of recent (1842-1844) Kilauea eruption
244 contain 2-6 wt% CO₂, and attributed the high CO₂ contents to enrichment by magma degassing
245 at low pressures. Based on the CO₂ emission rate and volcanic magma supply rate at Kilauea,
246 Anderson and Poland (2017 *Nature Geoscience* 10, 2017) estimated 1 wt% CO₂ in the parental
247 magma supplying Kilauea, which translates to 263 ppm C in their mantle source. This is
248 consistent with the result of a recent work using melt inclusions. [Tucker et al \(2019\)](#) showed that
249 parental magmas of Hawaii volcanos (Hualalai, Kilauea, Koolau, Loihi and Mauna Loa) have
250 0.39-1.0 wt% CO₂, and their mantle sources contain 380-480 ppm CO₂ that are at least a factor
251 of ~4 more than the MORB mantle. The negatively correlated Rb/Sr and ⁸⁷Sr/⁸⁶Sr in shield stage
252 Mahukona lavas (Huang et al., 2009) and the light Ca isotope compositions in most shield stage
253 Hawaiian lavas (Huang et al., 2011) may reflect a role of sedimentary carbonates in the Hawaiian
254 mantle plume. [Wirth and Rocholl \(2003\)](#) reported nanocrystalline diamond in pyroxenite
255 xenolith of Oahu Island, which may crystallize from rejuvenated stage magmas. Dixon et al.
256 (2008) argued for a role of carbonatite metasomatism in the petrogenesis of Hawaiian
257 rejuvenated stage lavas, requiring a CO₂ rich mantle source during the rejuvenated stage
258 volcanism.

259 As shown in [Fig. 2](#), partial melts of carbonated pyroxenites and carbonated peridotites have
260 higher contents of CaO and MnO compared to those from CO₂-free peridotites and pyroxenites.
261 Garnet, Clinopyroxene (Cpx) and orthopyroxene (Opx) are important hosts for Ca and Mn
262 during the mantle melting processes. The Cpx/melt partition coefficient of Mn (1.06-1.16) is

263 lower than that between garnet/melt (1-4.8), but higher than that between Opx/melt (0.66-1.05)
264 (Le Roux, 2011; Herzburg et al., 2013; Shea and Foley, 2019). The presence of CO₂ in the
265 mantle source would increase the stability of Opx relative to Cpx and garnet, thus, the
266 carbonated melts are enriched in Ca and Mn. This is supported by the observation that natural
267 carbonatites usually are enriched in Ca and Mn (e.g., Hoernle et al., 2002). To further investigate
268 the effect of CO₂ on the partitioning of Ca-Mn in olivine, we have selected the results of a suite
269 of high-quality experiments with resulted melts similar in composition to this study for
270 comparison.

271 Dasgupta et al (2007) conducted partial melting experiments on “peridotite+CO₂”, and
272 Mallik and Dasgupta (2013) conducted melting experiment on “eclogite+peridotite+CO₂”. These
273 two studies obtained carbonated silicate melts with SiO₂ of 30-48 wt% and CaO of 7-25 wt%,
274 which are comparable to the geochemistry of rejuvenated stage lavas. We also compared the Ca
275 partition coefficients for Hawaiian lavas with the results of experiments that produced olivines
276 equilibrium with dry silicate melts (Kogiso et al., 1998; Robinson et al., 1998; Mallik and
277 Dasgupta, 2012; Matzen et al., 2013 & 2017). These experiments either produced silicate melts
278 in equilibrium with olivine based on melting of peridotite (Robinson et al, 1998) and mixed
279 peridotite+eclogite (Kogiso et al, 1998; Mallik and Dasgupta, 2012), or modeled crystallization
280 of olivine from MORB-like melts (Matzen et al., 2013; Matzen et al., 2017). These experiments
281 all produced dry silicate melts with SiO₂ of 45-53 wt% and CaO of 6-11 wt% that are similar to
282 shield stage lavas and olivines with high precision Ca contents. We calculated the partition
283 coefficient of Ca between olivine/melt based on the above experiment studies, and the results are
284 plotted in Fig. 5.

285 To understand the effect of CO₂ on partitioning of Mn between olivine and melt, we have
286 compared our results with Mallik and Dasgupta (2013) (melting of eclogite+peridotite+CO₂),
287 Mallik and Dasgupta (2012) (melting of peridotite+eclogite), and Dasgupta et al (2007) (melting
288 of peridotite). The calculated partition coefficient of Mn for olivine-carbonated melt and
289 olivine-dry silicate melt are shown in Fig. 5. As shown in Fig. 5, the experimental partition
290 coefficients of Ca-Mn between olivine and carbonated silicate melt are overall lower than those
291 between olivine and dry silicate melt for a given Fo# and MgO content of melt. Rejuvenated
292 stage foidite samples have the lowest D_{Ca}^{Ol-melt} and D_{Mn}^{Ol-melt} for a given Fo# and magma MgO,

293 while the basanites and alkali basalts have $D_{Ca}^{Ol-melt}$ and $D_{Mn}^{Ol-melt}$ intermediate between foidites
294 and shield stage lavas (Fig. 5). $D_{Ca}^{Ol-melt}$ and $D_{Mn}^{Ol-melt}$ from the olivine-dry silicate melt pairs are
295 similar to those of shield stage lavas, while the partition coefficients of Ca-Mn calculated from
296 the olivine/carbonated silicate melt are comparable to those of rejuvenated stage (Fig. 5). We
297 suggest that the lower $D_{Ca}^{Ol-melt}$ and $D_{Mn}^{Ol-melt}$ of rejuvenated stage lavas relative to shield stage
298 lavas are consistent with the role of CO_2 during the mantle melting.

299 The relationships of $D_{Ca}^{Ol-melt}$ vs. magma MgO for the shield and rejuvenated lavas in this
300 study are compared with those resulted from experimental studies in Fig. 6a. As shown in Fig. 6a,
301 the $D_{Ca}^{Ol-melt}$ for the dry silicate melt system first slightly decreases and then increases with
302 decreasing MgO based on Gavrilenko et al (2016). The calculated $D_{Ca}^{Ol-melt}$ for shield stage lavas
303 basically follow the trend in $D_{Ca}^{Ol-melt}$ vs. MgO for the dry silicate melts (Fig. 6a). Similar to the
304 experimentally-derived $D_{Ca}^{Ol-melt}$ for carbonated silicate melt system, the $D_{Ca}^{Ol-melt}$ calculated for
305 rejuvenated stage lavas are well below the curve for the dry silicate melts. We have calculated
306 the difference in D_{Ca} ($\Delta D_{Ca}^{Ol-melt}$) between our calculated melts and experimental volatile-free
307 melts, and the results are shown in Supplementary Table 6. The foidites with the lowest SiO_2
308 tend to have the lowest $D_{Ca}^{Ol-melt}$ and $D_{Mn}^{Ol-melt}$ for a given olivine Fo and melt MgO content (Fig.
309 5), while the basanites and alkali basalts are intermediate between foidites and shield stage lavas
310 (Fig. 5). The overall high $\Delta D_{Ca}^{Ol-melt}$ for rejuvenated stage lavas relative to shield stage lavas (Fig.
311 6c) is consistent with the role of CO_2 in the mantle source, which significantly decreases SiO_2 in
312 melt. Another role of CO_2 is to cause negative anomalies of high field strength elements (HFSEs)
313 (e.g., decreases in Zr_N/Zr_N^* ; N, normalized to primitive mantle; Zr_N^* , calculated as $\sqrt{Nd_N \times Sm_N}$), because REEs are enriched relative to HFSEs in carbonated melts (Zhang et al.,
314 2017). As shown in Fig. 6d, the relatively low Zr_N/Zr_N^* (0.3 to 0.8) and elevated $\Delta D_{Ca}^{Ol-melt}$ for
315 rejuvenated stage lavas relative to shield stage further indicates the role of CO_2 in mantle melting.
316 It should be noted that the Honolulu foidites with the lowest SiO_2 , and $D_{Ca}^{Ol-melt}$ and $D_{Mn}^{Ol-melt}$
317 have the strongest negative anomalies of Nb and Zr (Fig. 3, Fig. 6d). This suggests that
318 rejuvenated stage foidites were subjected to the strongest influence of CO_2 during mantle
319 melting.

321 Gavrilenko et al (2016) gave an equation of CO_2 (wt.%) = $270 \times \Delta D_{Ca} - 3330 \times \Delta D_{Ca}^2 + 1.8$
322 $\times 10^6 \times \Delta D_{Ca}^3$ to calculate melt CO_2 content based on the correlation of CO_2 with the difference

323 (ΔD_{Ca}) in the D_{Ca} values of carbonated melts and volatile-free melts. Since both H_2O and CO_2
324 can lower $D_{Ca}^{Ol-melt}$ (Gavrilenko et al., 2016), the effect of H_2O should also be considered to
325 estimate magma CO_2 content based on $D_{Ca}^{Ol-melt}$. As shown in Dixon et al (1997), the alkali
326 basaltic to nephelinitic lavas from the North Arch Volcanic Field, Hawaii, can have H_2O up to
327 1.9 wt% and CO_2 up to 5.4 wt% based on analyses of basalt glass and vesicles. Dixon et al (2008)
328 further estimated 350 ppm H_2O in the Hawaii plume mantle, which is ~5 factors higher than the
329 depleted MORB mantle. Dixon et al (2008) estimated a bulk partition coefficient of ~0.01 for
330 H_2O in the Hawaiian mantle. Hawaiian rejuvenated lavas are usually considered to have partial
331 melting extents of <5%, thus, we have calculated rejuvenated magma H_2O contents based on a
332 bulk partition coefficient of 0.01 for H_2O and batch melting degrees of 2% and 4%, and our
333 calculation results in H_2O contents of 1.17 wt% and 0.71 wt%, respectively. Shield stage magma
334 H_2O is calculated by assuming a batch melting degree of 20% in Hawaiian mantle source, and a
335 magma H_2O content of 0.17 wt% is obtained. For comparison, Hauri (2002 Chem Geol 183,
336 115-141) reported 0.03 to 0.84 wt % H_2O in melt inclusions from five shield stage lavas from
337 Hawaii. We estimated the influence of H_2O on $D_{Ca}^{Ol-melt}$ (ΔD_{Ca} at a given H_2O content) based on
338 Gavrilenko et al (2016). Then, we calculated the difference between $\Delta D_{Ca}^{Ol-melt}$ and ΔD_{Ca} at a
339 given H_2O content, which is used to calculate the content of CO_2 based on the equation of
340 Gavrilenko et al (2016). The calculated results of CO_2 are shown in Supplementary Table 6.

341 The mantle-derived primary magmas usually crystallize olivines with $Fo\#$ up to ~91, which
342 are higher than the olivine phenocrysts in this study (Fig. 4). Thus, the result using the above
343 method only represents the CO_2 concentration of evolved magmas, rather than primary magmas.
344 The primary magma composition can be obtained by addition of equilibrium olivines to melt,
345 however, the Ca contents of high $Fo\#$ (>90) olivines in equilibrium with primary magmas are not
346 known. Therefore, we cannot obtain the primary magma CO_2 content directly based on the
347 equation of Gavrilenko et al (2016). To obtain primary magma CO_2 content, we assume that
348 magmas have not significantly degassed during fractionation. Then, the equilibrium olivines
349 were added step by step to melts until the melts are in equilibrium with olivine with $Fo\#$ of 91
350 based on PREMELT3 MEGA.XLSM of Herzberg and Asimow (2015), the resulted primary
351 magma compositions and fraction of olivine added are shown in Supplementary Table 6. The
352 primary magma CO_2 contents were obtained after correction to fraction of olivine added to melts.

353 The calculation procedure and results are shown in [Supplementary Table 6](#). The CO₂ contents of
354 the calculated shield stage primary magmas are close to zero. The calculated primary magma
355 CO₂ for rejuvenated stage are up to 5.6 wt% and 8.3 wt% for batch melting degrees of 2% and
356 4%, respectively ([Supplementary Table 6](#)).

357 We have compared our primary magma CO₂ contents with those calculated based on a
358 constant CO₂/Ba ratio. Because CO₂ is similar to Ba in incompatibility during mantle melting
359 (e.g., [Anderson and Poland, 2017](#); [Miller et al., 2019](#)), the primary magma CO₂ content can be
360 estimated if the mantle source CO₂/Ba ratio can be determined. As suggested by previous studies
361 the Hawaiian mantle has a CO₂/Ba ratio of ~86 ([Anderson and Poland, 2017](#)). The calculated
362 primary magma CO₂ contents through this method is also shown in [Supplementary Table 6](#). The
363 calculated primary magma CO₂ contents are 0.45 to 3.15 wt% for the shield stage, while the
364 primary magma CO₂ contents are 2.8 to 10 wt% for rejuvenated stage. We have plotted the
365 primary magma CO₂ contents for anhydrous melting and a melting degree of 4% in this study
366 with those based on the constant CO₂/Ba of 86 in [Fig. 7](#). As shown in [Fig. 7](#), the foidite samples
367 tend to have the highest primary magma CO₂ contents, while the basanites and alkali basalts are
368 similarly low relative to the foidite samples. The results of anhydrous melting are overall close to
369 those based on assuming a constant CO₂/Ba ratio, however, the results of hydrous melting,
370 melting degree of 4%, for rejuvenated stage are generally lower than those based on assuming a
371 constant CO₂/Ba ratio. Since the latter method is independent of uncertainties in magma
372 degassing and melting degrees, the lower primary magma CO₂ contents could have been caused
373 by these uncertainties. Despite discrepancy between the two methods, they are overall consistent
374 in resulting in primary magma CO₂ contents up to ~10 wt% for rejuvenated stage lavas and low
375 CO₂ contents for shield stage lavas, suggesting rejuvenated stage lavas have a similar source CO₂
376 to the plume mantle. However, as a constant CO₂/Ba ratio of Hawaiian mantle source is not
377 verified for shield and rejuvenated stage lavas, our work provides an independent constraint on
378 CO₂ enrichment in Hawaiian rejuvenated lavas.

379 **4.3 Olivine Ni: effects of source vs. temperature**

380 There are different views on the origin of Hawaiian rejuvenated stage volcanism, either
381 from the shallow metasomatized mantle lithosphere ([Chen and Frey, 1983](#); [Yang et al., 2003](#);
382 [Bianco, 2005](#); [Garcia et al., 2010](#)), the deep mantle plume component ([Frey et al., 2005](#); [Dixon et](#)

383 al., 2008; Sisson et al., 2009; Hofmann and Farnetani, 2013; DeFelice et al., 2019; Harrison et al.,
384 2020), or both (Beguelin et al., 2019). The deep plume origin for rejuvenated stage melts is
385 supported by their distinct isotopic compositions from the Pacific MORBs (Supplementary Fig.
386 3). Despite enrichment of LILEs and LREEs in rejuvenated stage lavas, they generally have
387 more depleted isotopic compositions compared to shield stage lavas (Yang et al., 2003; Hofmann
388 and Farnetani, 2013; Beguelin et al., 2019; DeFelice et al., 2019; Harrison et al., 2020). As
389 shown in Supplementary Fig. 3, such a depleted signature cannot be explained by involvement of
390 Pacific-type depleted upper mantle, but likely to have been derived from an intrinsic mantle
391 plume component. However, it remains unclear how the low-degree melts of rejuvenated stage
392 preferentially sample the isotopically depleted component in the mantle plume (Ribe and
393 Christensen, 1999; Bianco and Ito, 2008; Beguelin et al. 2019; DeFelice et al. 2019). For
394 example, it was suggested that rejuvenated stage lavas originated from a deep isotopically
395 depleted zone separated from the primary melting zone that formed shield stage lavas (Ribe and
396 Christensen, 1999), or from a depleted periphery of a zoned mantle plume (Bianco and Ito,
397 2008).

398 As shown in Fig. 4a, the olivine Ni contents of rejuvenated stage lavas are overall lower
399 than those of shield stage, but higher than the global MORBs for a given olivine Fo#. In contrast,
400 rejuvenated stage lavas have whole-rock Ni lower than shield stage lavas for a given MgO
401 content (Fig. 2c). Increasing magma total alkaline (Na₂O+K₂O) increases the partition coefficient
402 of Ni in olivine when total alkaline > 8 wt% (Förster et al., 2018). Most of our studied rocks
403 have K₂O+Na₂O contents significantly lower than 8 wt%, hence, it is unexpected to have notable
404 difference in $D_{Ni}^{Ol-melt}$ between shield and rejuvenated stage lavas. As we discussed above, CO₂
405 played a key role in the origin of rejuvenated stage magmatism. However, Girnis et al (2013)
406 showed that CO₂ have negligible influence on the partitioning of Ni between olivine/melt. As
407 such, we suggest the relatively lower olivine Ni contents of the rejuvenated stage compared with
408 the shield stage reflects a lower Ni content in their parental magmas, as indicated by whole-rock
409 Ni contents (Fig. 2c).

410 Melting temperature/pressure, source lithology, and source Ni content could influence
411 primary magma Ni content (Sobolev et al., 2005&2007; Putirka et al, 2011; Matzen et al., 2013).
412 Partial melts of eclogite/pyroxenite react with peridotite in the Hawaii mantle plume to convert

413 olivines to Opx and to form stage-2 pyroxenite with a lower proportion of (or no) olivine in the
414 source (Sobolev et al., 2005&2007). Thus, the mantle source with a lowered olivine proportion
415 would generate primary melts with elevated Ni (Sobolev et al., 2005). As indicated by Matzen et
416 al (2013, 2017), $D_{Ni}^{Ol-melt}$ increases with decreasing temperature and pressure, thus, elevated
417 source temperatures and pressures would result in an increase of olivine Ni in Hawaiian shield
418 lavas. In addition to the stage-2 pyroxenite in the mantle source, the higher olivine Ni of
419 Hawaiian shield stage lavas may at least partly have been caused by higher melting temperatures
420 and pressures (e.g., Matzen et al., 2013). Thus, evaluation of difference in primary magma
421 temperatures among the shield stage lavas, rejuvenated stage lavas and normal MORBs would
422 help understand the origin of their Fo-Ni systematics (Fig. 4a).

423 In this study, we have calculated the magma temperatures of shield and rejuvenated stages
424 based on the olivine-spinel aluminum exchange thermometry. This thermometry is advantageous
425 over the olivine-liquid thermometry since it is independent of equilibrium pressure and melt
426 compositions. Details on the use of the olivine-spinel aluminum exchange thermometry are
427 referred to Wan et al (2008) and Coogan et al (2014). We selected the volcanic rock samples
428 from the rejuvenated and shield stages that contain the highest olivine Fo#. We used the equation
429 given by Coogan et al (2014) ($T(K)=10000/(0.575+0.884Cr\#-0.897\ln(Al_2O_3^{Ol}/Al_2O_3^{Sp}))$), which
430 has extended the use of this thermometry to a relatively oxidizing environment. The data of
431 olivine-spinel are filtered based on the requirements of Coogan et al (2014), i.e., Cr# between
432 0-69%, and $Fe^{3+}/Total\ FeO < 35\%$. The calculated results are shown in Supplementary Table 4
433 and plotted in Fig. 8a. As shown in Fig. 8a, shield stage olivines with the highest Fo# (~89) tend
434 to have the highest crystallization temperatures. Rejuvenated stage olivines have statistically
435 lower crystallization temperatures and Fo# than those of shield stage as shown in Fig. 8 c-d. The
436 slightly lower olivine crystallization temperatures of rejuvenated stage lavas could also be
437 indicated by their slightly higher olivine Al_2O_3 (Supplementary Fig. 4) (Coogan et al., 2014).
438 However, the lavas of both shield and rejuvenated stages approximately follow the trend of melt
439 temperature vs. equilibrium olivine Fo# of the Icelandic primary magmas as shown in Herzberg
440 and Asimow (2015). Thus, for the given highest olivine Fo# (e.g., ~91), broadly similar primary
441 magma temperatures for the shield and rejuvenated stage lavas are expected (Fig. 8). As shown
442 in Fig. 8a, at the highest Fo# of up to 91, the MORB olivines tend to have lower temperatures

443 compared with those of Hawaiian olivines. This indicates that the Hawaiian primary magmas,
444 both shield and rejuvenated stages, crystallized high Fo# olivines at higher temperatures than
445 those of normal MORBs.

446 Since Hawaiian shield and rejuvenated stage lavas show different Sr-Nd-Hf isotopic
447 compositions (Supplementary Fig. 3; Yang et al., 2003; Hofmann and Farnetani, 2013; DeFelice
448 et al., 2019), the overall lower olivine Ni contents of rejuvenated stage lavas relative to those of
449 shield stage may reflect a mantle source effect. As shown in Fig. 8, high Fo olivines from both
450 shield and rejuvenated stage lavas have similar crystallization temperatures. The effect of
451 melting-crystallization temperature difference (Matzen et al., 2013) may have not contributed
452 significantly to the observed olivine Ni difference.

453 The mantle source of Hawaiian shield stage lavas has been considered to contain eclogites
454 that were most likely derived from ancient recycled oceanic crust. Subducted altered oceanic
455 crust, as an important carrier of secondary carbonates (Zhang and Smith-Duque, 2014), may
456 constitute a source of CO₂ in the Hawaiian mantle plume. Thus, it is possible that CO₂-bearing
457 mixed eclogite/peridotite exists in the Hawaii mantle plume. However, because the plume mantle
458 source has melted to high degrees to form shield stage lavas, most of CO₂ would be extracted
459 from the source mantle after high degrees of melting. As CO₂ is highly incompatible during
460 mantle melting (e.g., Dixon et al., 2008), the resulted refractory residual mantle after extraction
461 of shield stage lavas would be difficult to produce carbonated melts. Because Hawaiian
462 rejuvenated stage lavas were formed 0.5-2 Myr after shield stage and ~100 km from the plume
463 axis, it is possible that the rejuvenated lavas were sourced from the edge of the mantle plume,
464 where mantle upwelling and decompression rate are lower than the plume axis (Fig. 9).

465 As indicated by this study, the rejuvenated stage lavas were most likely sourced from
466 carbonated peridotites that are intrinsic to the Hawaiian mantle plume. Thus, the mantle plume
467 peridotites may have experienced metasomatism of low degree carbonated melts. This is
468 consistent with the enriched trace element patterns and negative anomalies of HFSEs in
469 rejuvenated lavas (Fig. 3). However, the distinct Sr-Nd-Hf isotopic compositions between
470 rejuvenated and shield stage lavas indicate that such carbonated melts were unlikely derived
471 from a CO₂-bearing eclogite/pyroxenite (carbonate-bearing recycled oceanic crust) feeding
472 shield stage lavas. A possible source for the carbonated melts in the mantle plume is subducted

473 lithospheric mantle (e.g., [Kelemen and Manning, 2015](#)). Although carbonate precipitation is
474 generally considered to occur mainly in the oceanic crust during seafloor alteration, it would also
475 occur in shallow lithospheric mantle of the bending plate before subduction ([Kelemen and](#)
476 [Manning, 2015](#)), and in the exposed abyssal peridotites as represented by slow- to ultra-slow
477 spreading ridge settings (e.g., [Dick et al., 2000](#); [Schroeder et al., 2002](#)). These carbonate-bearing
478 subducted peridotites may constitute the source required by Hawaiian rejuvenated stage lavas. As
479 shown in [Fig. 9](#), the low degree carbonated melts of such carbonate-bearing peridotites near the
480 plume edge can metasomatize the mantle plume peridotites. Thus, along with the upwelling of
481 mantle plume, such carbonated peridotites would decompress and melt, which may explain the
482 CO₂-rich rejuvenated stage lavas.

483 5 Conclusions

484 The origin of Hawaiian rejuvenated stage magmatism remains a subject of debate. In this
485 study, we have analyzed the whole-rock major and trace elements, *in situ* high-precision
486 geochemistry of olivines and olivine-spinel pairs of Hawaiian rejuvenated and shield stage lavas.
487 Unlike shield stage, rejuvenated stage lavas with anomalously high CaO and low SiO₂ and
488 negative anomalies of HFSEs, which could be explained by the influence of CO₂ in the source. It
489 is notable that the olivine Ni contents of rejuvenated stage lavas are systematically lower than
490 those of shield stage, but higher than normal MORBs for a given olivine Fo#. While rejuvenated
491 stage lavas have whole-rock CaO and MnO higher than shield stage, their olivine Ca and Mn
492 contents are similar to each other for a given Fo#. Our study results in that D_{Ca}^{Ol-melt} and
493 D_{Mn}^{Ol-melt} for rejuvenated stage lavas are systematically lower than those of shield stage, which
494 we suggest is caused by the influence of CO₂. Our calculation based on the reduced D_{Ca}^{Ol-melt}
495 relative to dry basaltic melts indicates that rejuvenated stage primary melts were rich in CO₂,
496 while shield stage melts were relatively dry (low volatile contents). The temperatures of primary
497 magmas of rejuvenated stage are similar to the shield stage, but systematically higher than
498 normal primary MORB magmas. The relatively low olivine Ni of rejuvenated lavas compared to
499 shield stage can be attributed to the mantle source lithology dominated by peridotites, rather than
500 stage-2 pyroxenite in shield stage mantle source. It is suggested that rejuvenated stage lavas were
501 originated from the melting of carbonated peridotites in the Hawaiian mantle plume.

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509 **Data Availability Statement**

510 Data archiving is underway, and will be deposited in Mendeley Data after the manuscript is
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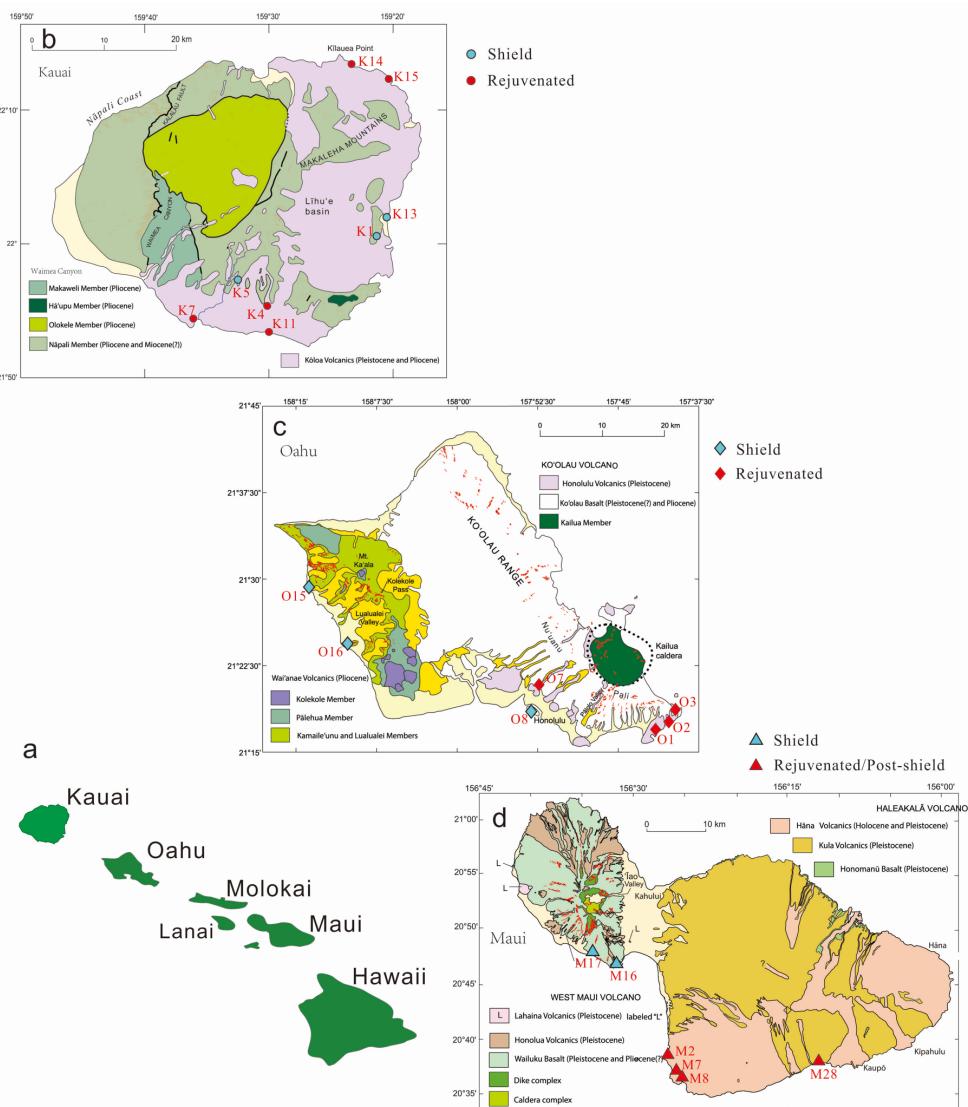
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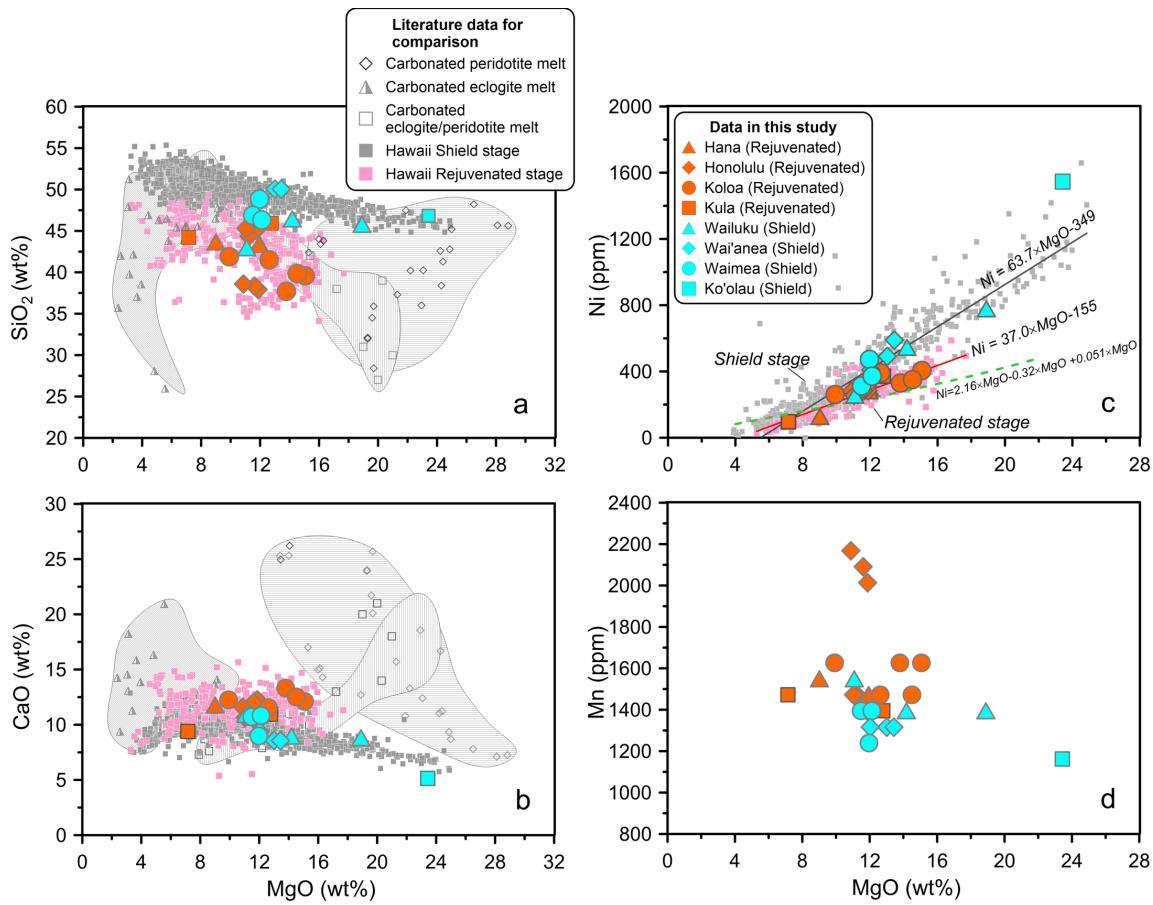
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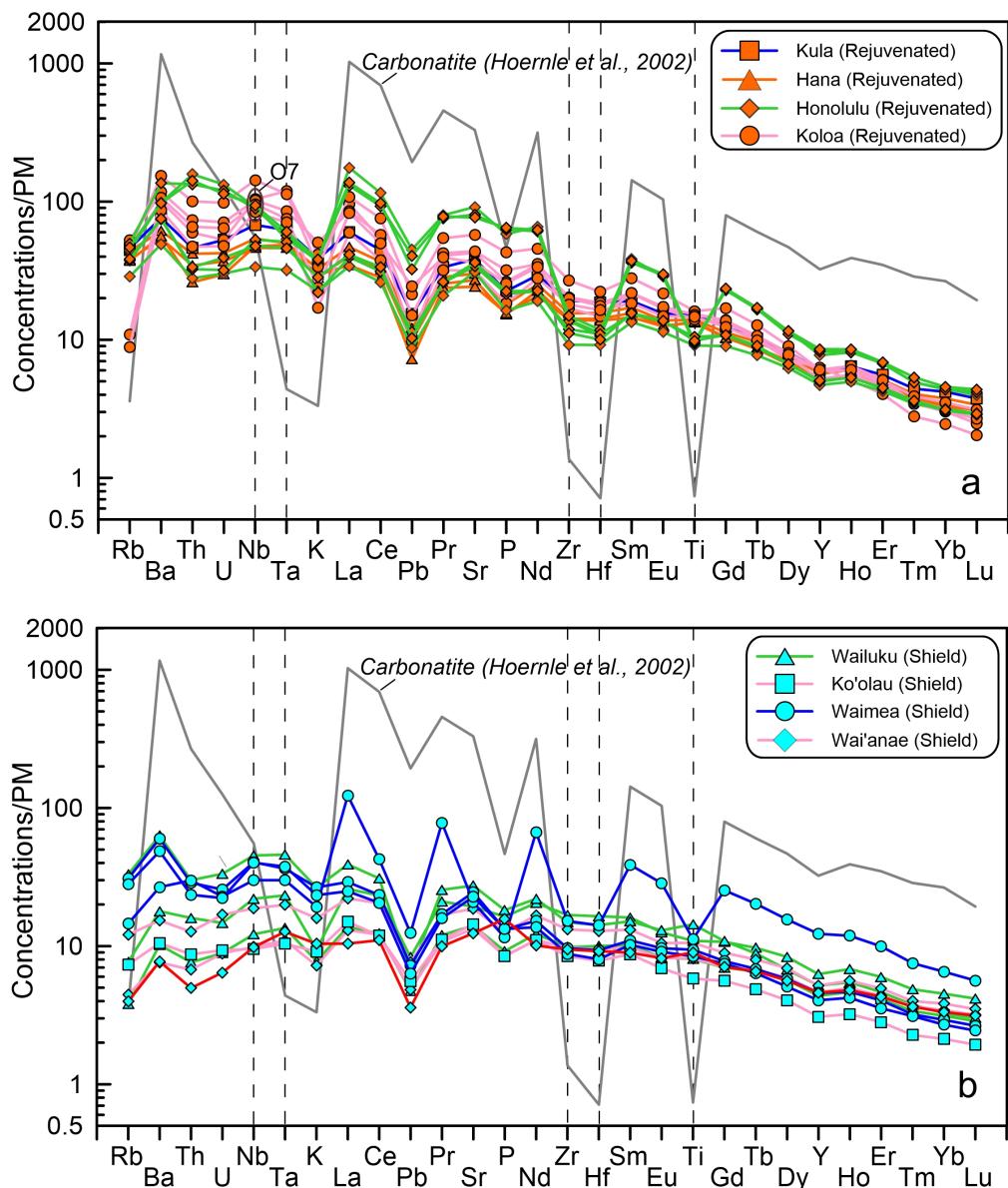
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720 **Figure 1.** Geological setting and sampling locations of this study. The geologic maps are
721 modified according to [Sherrod et al \(2007\)](#).
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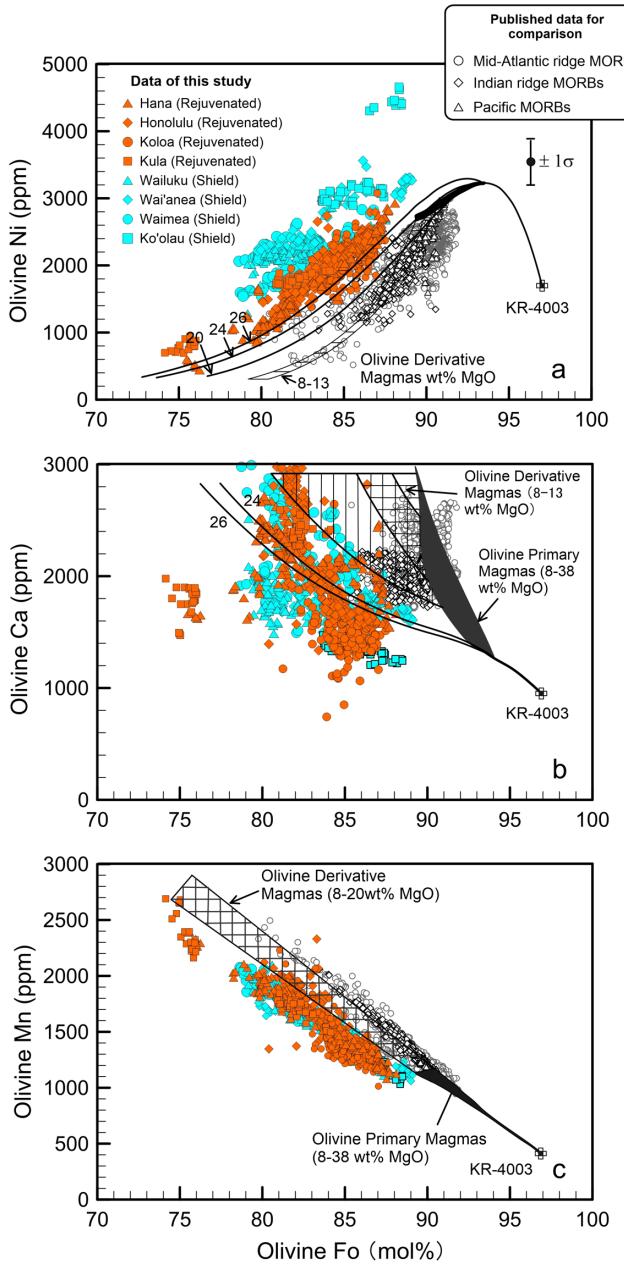
724 **Figure 2. Plots of bulk-rock MgO vs. (a) SiO₂, (b) CaO, (c) Ni and (d) Mn for volcanic rocks**
 725 **of Hawaiian islands.** Green dashed line in (c), primary accumulated fractional melts of fertile
 726 mantle peridotite (KR-4003) modeled by calculated by Herzburg (2011). Source of literature
 727 data for comparison: Carbonated peridotite melt (Hirose, 1997; Dasgupta et al., 2007),
 728 carbonated eclogite melt (Kiseeva et al., 2012; Kiseeva et al., 2013; Hammouda et al., 2010;
 729 Gerbode and Dasgupta, 2010), carbonated eclogite+peridotite melt (Mallik and Dasgupta,
 730 2013), data of Hawaiian shield and rejuvenated stage in (a)-(c) for comparison are from
 731 <http://georoc.mpch-mainz.gwdg.de/georoc/Entry.html>.

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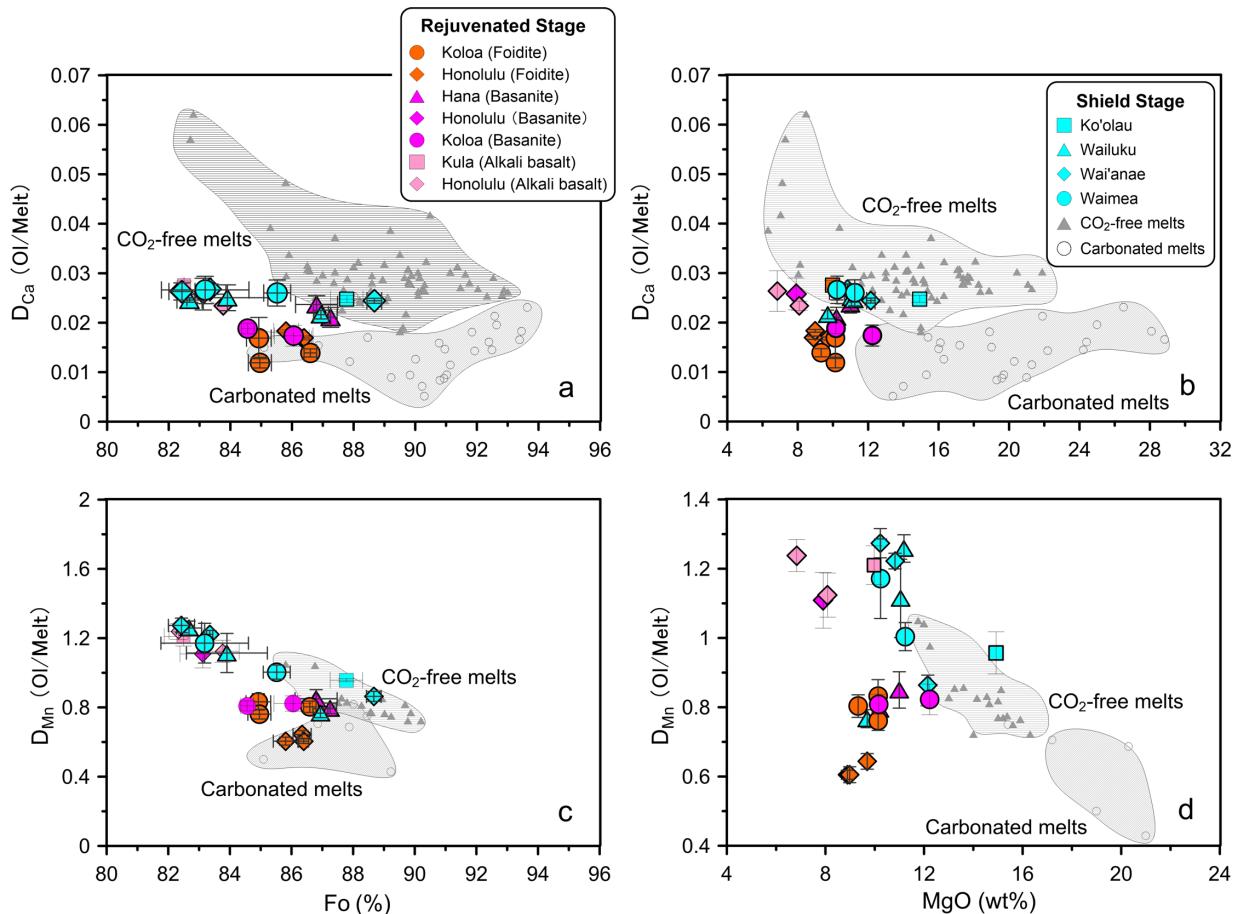
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734 **Figure 3. Trace element patterns for the volcanic rock samples from Hawaiian Islands.** Data
 735 are normalized to the primitive mantle data of [McDonough and Sun \(1995\)](#). The typical
 736 carbonatite (Sample #: 85LB25, the carbonatite lava) for comparison from [Hoernle et al](#)
 737 ([2002](#)).
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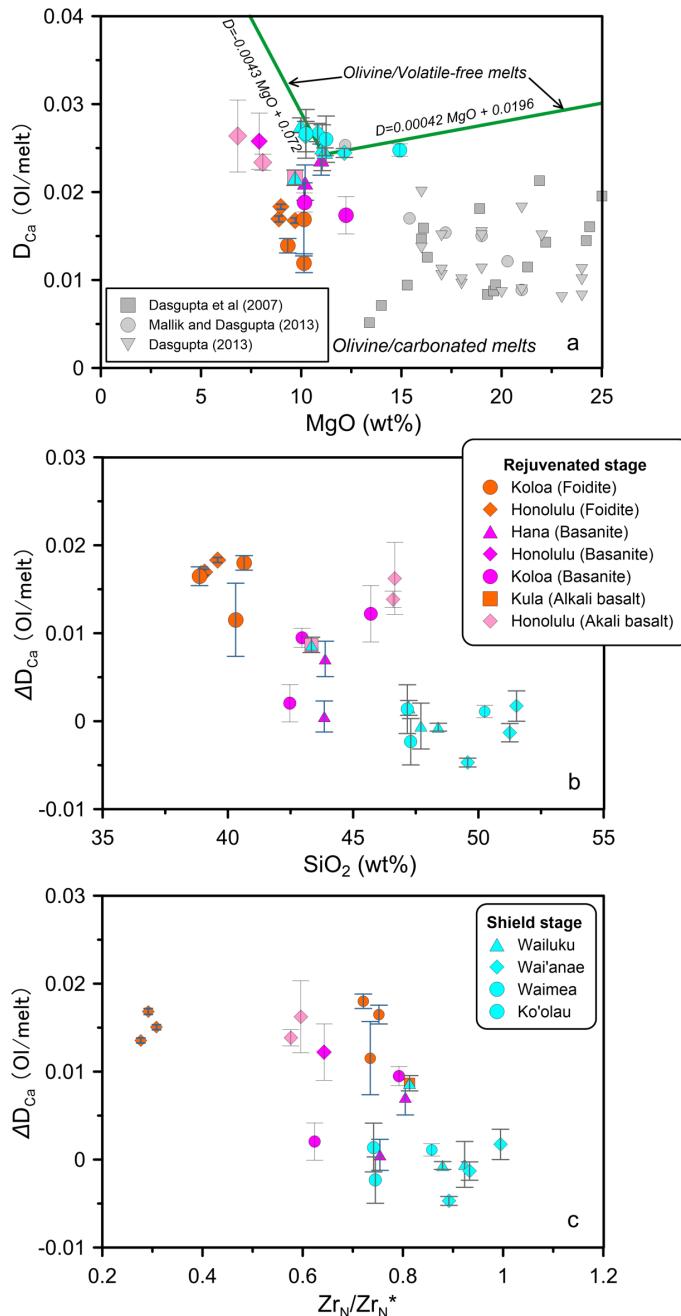
740 **Figure 4. Plots of olivine Mg number (Fo) vs. (a) Ni, (b) Ca and (c) Mn.** The calculated
 741 olivine compositions are based on [Herzberg \(2011\)](#). Black area shows the olivines of primary
 742 magmas (MgO of 8–38 wt%) from fertile peridotite KR-4003 with 1964 ppm Ni, 3.45 wt%
 743 CaO, 1007 ppm Mn, and 8.02 wt% FeO. Numbered lines in (a) and (b) are calculated olivines
 744 of olivine-fractionated derivative magmas, and the numbers indicate the MgO contents of
 745 olivine derivative magmas. Short lines with end-bars in (a) and (d) are $\pm 1\sigma$ Ni and Fe/Mn
 746 variations of the primary magmas from which olivines crystallize. Data of olivine for global
 747 MORBs (Indian ridge, Mid-Atlantic Ridge, and East Pacific Rise) and Hawaiian OIBs are
 748 from [Sobolev et al. \(2007\)](#).



749

750 **Figure 5. Plots of D_{Ca} (Ol/melt) vs. (a) Fo and (b) melt MgO , and D_{Mn} (Ol/melt) vs. (c)**
 751 **olivine Mg number (Fo) and (d) melt MgO .** Source of literature data for comparison: data
 752 for D_{Mn} (Mallik and Dasgupta, 2012 & 2013; Matzen et al., 2017), data for D_{Ca} (Mallik and
 753 Dasgupta, 2012 & 2013; Matzen et al., 2017; Dasgupta et al., 2007; Robinson et al., 1998;
 754 Kogiso et al., 1998).

755



756

757 **Figure 6. Plots of (a) $D_{Ca}^{Ol/melt}$ vs. equilibrium melt MgO , $\Delta D_{Ca}^{Ol/melt}$ vs. (b) SiO_2 and (c)**
758 **Zr_N/Zr_N^* .** The results of high-pressure/temperature experiments from Dasgupta et al. (2007,
759 2013) and Mallik and Dasgupta (2013) are plotted in (a) for comparison. The green solid line
760 in (a) indicates the results of such experiments with volatile-free melts from Gavrilenko et al.
761 (2016). $\Delta D_{Ca}^{Ol/melt}$ indicates difference between the olivine D_{Ca} values of calculated melts and
762 volatile-free melts. Zr_N , normalized to primitive mantle data of McDonough and Sun (1995).
763 Zr_N^* , calculated as square root of $Nd_N \times Sm_N$.

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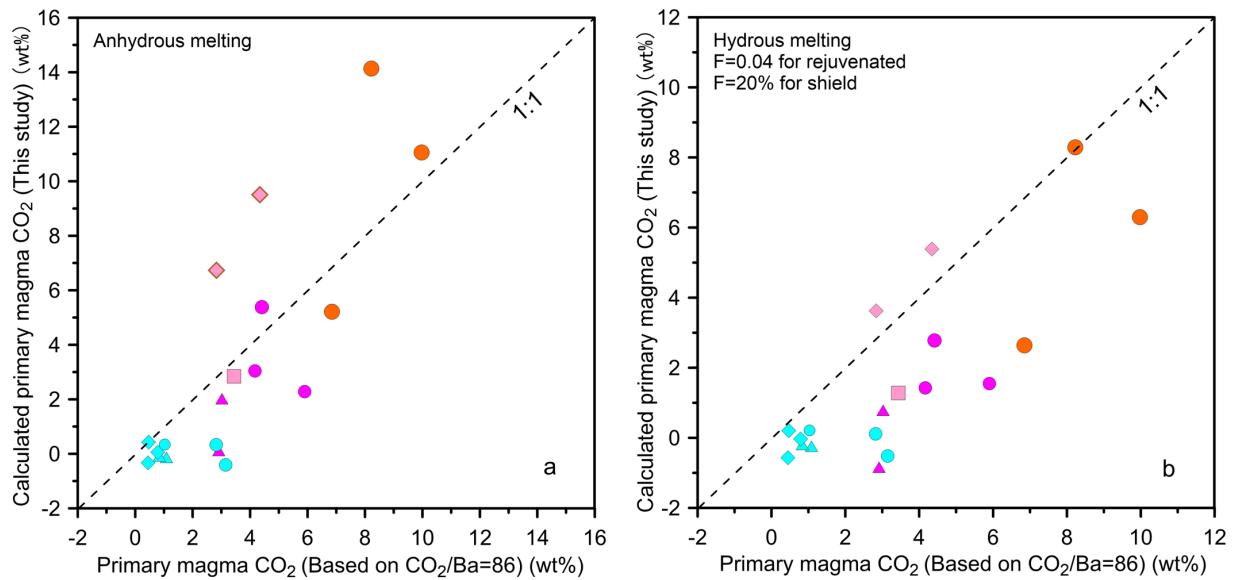


Figure 7. Plots of calculated primary magma CO_2 based on $\text{CO}_2/\text{Ba}=86$ vs. calculated primary magma CO_2 on basis of (a) anhydrous and (b) hydrous melting. Anhydrous melting assumes no influence of water on partition of Ca in olivine. F, melting degree of mantle based on which water content is calculated. Primary magma CO_2 content is calculated based on $\text{CO}_2/\text{Ba}=86$ and is corrected based on fraction of olivine added during primary magma calculation.

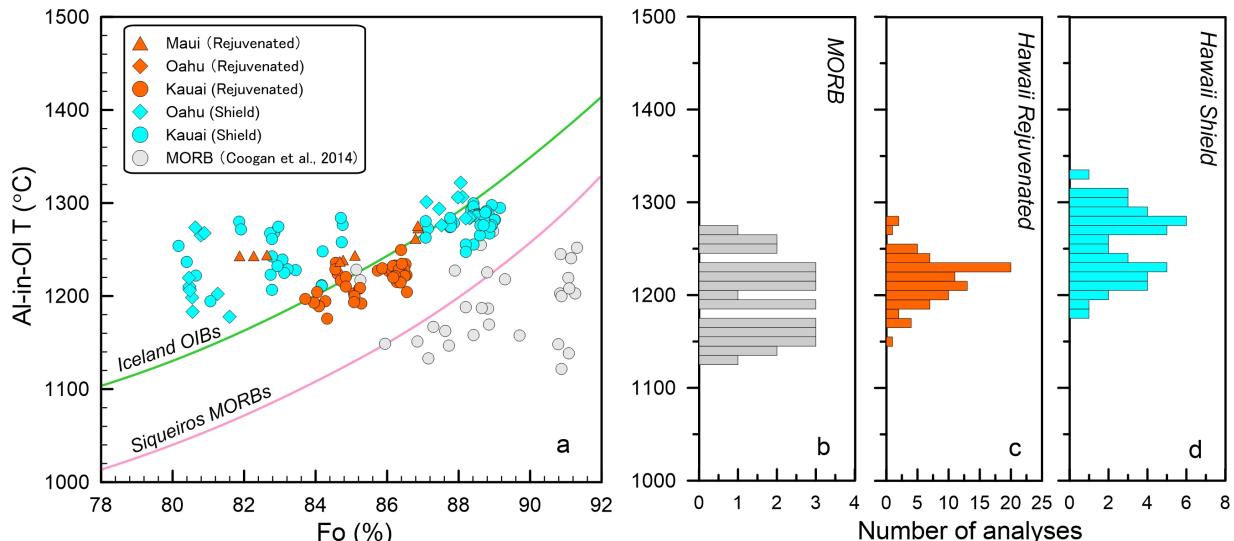
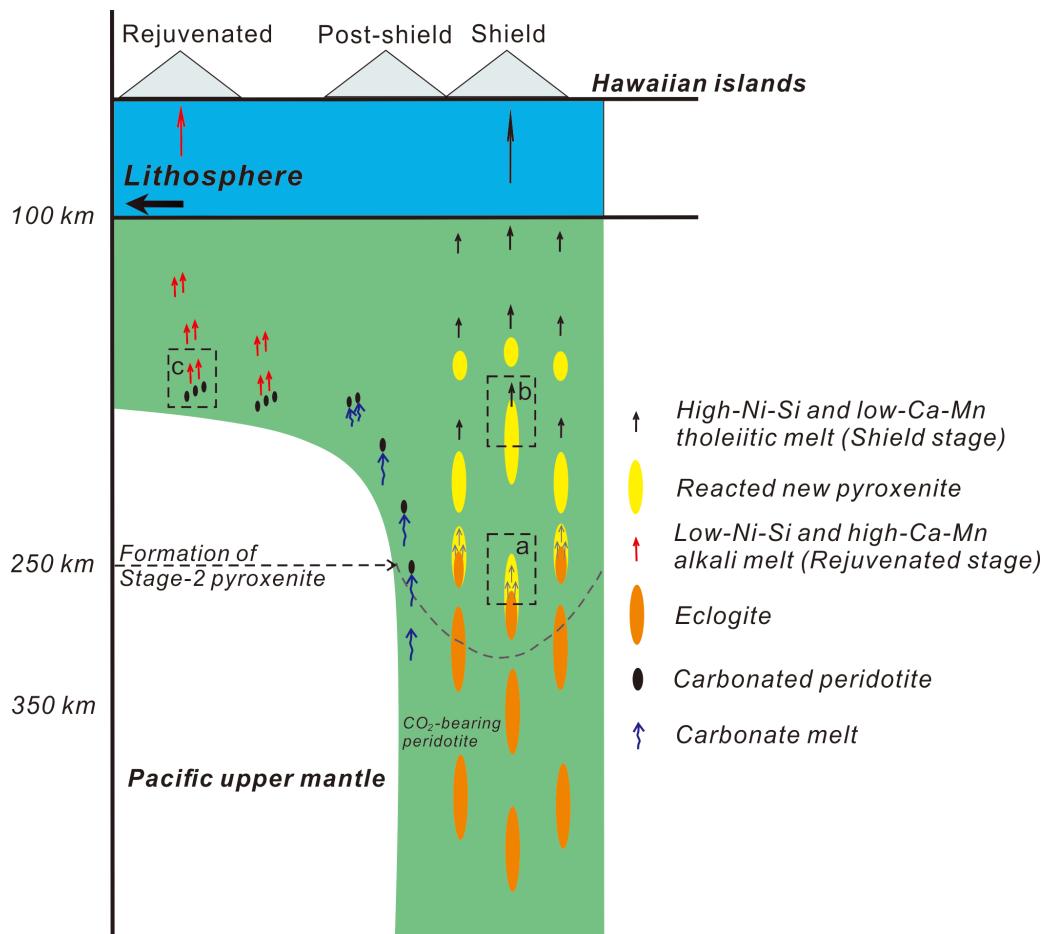


Figure 8. Al-in-olivine crystallization temperature as a function of olivine Fo (a) and histogram for comparing Hawaii rejuvenated and shield stage crystallization temperature (b-c). Data of MORB temperature for comparison are from Coogan et al (2014). The solid lines for Iceland OIB and Siqueiros MORB are calculated by PREMELT3 MEGA.XLSM based on Herzberg and Asimow (2015).



781 **Figure 9. Model showing genesis of shield stage and rejuvenated stage volcanism.** *a.*
782 Reaction of eclogite melt with peridotite to form Stage-2 pyroxenite (Sobolev et al., 2005); *b.*
783 High-degree melt of stage-2 pyroxenite to produce Hawaii Shield stage volcanism; *c.* Melting
784 of carbonated peridotite to form CO₂-rich rejuvenated stage melts.