

MDPI

Article

# Laboratory Models of Planetary Core-Style Convective Turbulence

Emily K. Hawkins <sup>1,2,\*</sup>, Jonathan S. Cheng <sup>1,3</sup>, Jewel A. Abbate <sup>1</sup>, Timothy Pilegard <sup>1,4</sup>, Stephan Stellmach <sup>5</sup>, Keith Julien <sup>6</sup> and Jonathan M. Aurnou <sup>1</sup>

- Department of Earth, Planetary, and Space Sciences, University of California, Los Angeles, CA 90024, USA
- Department of Physics, Loyola Marymount University, Los Angeles, CA 90045, USA
- School of Physics and Astronomy, Rochester Institute of Technology, Rochester, NY 14623, USA
- Bedrock Ocean Exploration, Richmond, CA 94804, USA
- Institut für Geophysik, Westfälische Wilhelms-Universität Münster, 48149 Münster, Germany
- Department of Applied Mathematics, University of Colorado, Boulder, CO 80309, USA
- \* Correspondence: emily.hawkins@lmu.edu

Abstract: The connection between the heat transfer and characteristic flow velocities of planetary corestyle convection remains poorly understood. To address this, we present novel laboratory models of rotating Rayleigh-Bénard convection in which heat and momentum transfer are simultaneously measured. Using water (Prandtl number,  $Pr \simeq 6$ ) and cylindrical containers of diameter-to-height aspect ratios of  $\Gamma \simeq 3, 1.5, 0.75$ , the non-dimensional rotation period (Ekman number, E) is varied between  $10^{-7} \lesssim E \lesssim 3 \times 10^{-5}$  and the non-dimensional convective forcing (Rayleigh number, Ra) ranges from  $10^7 \lesssim Ra \lesssim 10^{12}$ . Our heat transfer data agree with those of previous studies and are largely controlled by boundary layer dynamics. We utilize laser Doppler velocimetry (LDV) to obtain experimental point measurements of bulk axial velocities, resulting in estimates of the nondimensional momentum transfer (Reynolds number, Re) with values between  $4 \times 10^2 \lesssim Re \lesssim 5 \times 10^4$ . Behavioral transitions in the velocity data do not exist where transitions in heat transfer behaviors occur, indicating that bulk dynamics are not controlled by the boundary layers of the system. Instead, the LDV data agree well with the diffusion-free Coriolis-Inertia-Archimedian (CIA) scaling over the range of Ra explored. Furthermore, the CIA scaling approximately co-scales with the Viscous-Archimedian-Coriolis (VAC) scaling over the parameter space studied. We explain this observation by demonstrating that the VAC and CIA relations will co-scale when the local Reynolds number in the fluid bulk is of order unity. We conclude that in our experiments and similar laboratory and numerical investigations with  $E \gtrsim 10^{-7}$ ,  $Ra \lesssim 10^{12}$ ,  $Pr \simeq 7$ , heat transfer is controlled by boundary layer physics while quasi-geostrophically turbulent dynamics relevant to core flows robustly exist in the fluid bulk.

Keywords: rotating convective turbulence; velocities; heat flow; planetary interiors; dynamos



Citation: Hawkins, E.K.; Cheng, J.S.; Abbate, J.A.; Pilegard, T.; Stellmach, S.; Julien, K.; Aurnou, J.M.
Laboratory Models of Planetary
Core-Style Convective Turbulence.
Fluids 2023, 8, 106. https://doi.org/10.3390/fluids8040106

Academic Editors: Ramesh Agarwal and D. Andrew S. Rees

Received: 22 August 2022 Revised: 12 March 2023 Accepted: 16 March 2023 Published: 23 March 2023



Copyright: © 2023 by the authors. Licensee MDPI, Basel, Switzerland. This article is an open access article distributed under the terms and conditions of the Creative Commons Attribution (CC BY) license (https://creativecommons.org/licenses/by/4.0/).

#### 1. Introduction

The rotating convective turbulence of electrically conducting liquid metal layers within planets is responsible for the generation of their global-scale magnetic fields [1,2]. The details of core-style turbulence, particularly the connection between heat and momentum transfer, remain elusive. These turbulent flows that exist at planetary conditions and are remote from direct measurement are difficult to directly model due to their complex and extreme nature. Laboratory experiments, direct numerical simulations (DNS), and theoretical models use various simplified approaches to understand the turbulent fluid physics of dynamo-generating mechanisms (e.g., [3–5]). The goal of these reduced models is to develop accurate scaling laws characterizing fundamental system behaviors that can be extrapolated to the extreme settings at which planetary dynamos exist.

One of the most common models used to study this problem is that of rotating Rayleigh-Bénard convection (RRBC), where a layer of fluid is heated from below while

Fluids 2023. 8, 106 2 of 35

simultaneously cooled from above and rotated about a vertical axis. The laboratory experiments in this RRBC study implement a right cylindrical geometry, thus removing any effects of spherical curvature. A number of recent studies have focused on characterizing the effects of cylindrical sidewalls on rotating heat transfer (e.g., [6–12]). These studies show that sidewall effects are most pronounced very close to the onset of convection. Overall, bulk cylindrical laboratory heat transfer data agree with those of unconfined periodic DNS. Furthermore, the recent study by Gastine and Aurnou [13] directly compares the heat transfer of the spherical shell DNS to that of the local Cartesian DNS and concludes that local, non-spherical models can be used to understand spherical systems (see also Wang et al. [14]). The results of these recent studies demonstrate that laboratory cylinders are a useful and relevant tool for studying spherical shell-core dynamics.

To illustrate our model system, Figure 1 displays a schematic of the northern hemisphere of Earth's core region. The cylinder located at a high latitude and aligned with the rotation axis represents one of the laboratory containers used in this work. Our experiments simulate a local, polar parcel of planetary core fluid convecting under the influence of axial rotation and buoyancy forcing. The interior of the cylinder is filled with an illustrative subset of the visualization of a geostrophically turbulent laboratory experiment. The bottom (top) boundaries of the cell are red (blue) to represent the warm (cold) fluid in parallel with the warmer/red inner core boundary (cooler/blue core—mantle boundary).

The outer core region in Figure 1 shows a meridional slice of the instantaneous radial velocity of a dynamo simulation adapted from Sheyko et al. [15]. From this meridional slice, we see that fluid motions are largely aligned with the rotation axis. Furthermore, these fluid motions, which are generated in the presence of a magnetic field and a spherical shell geometry, compare qualitatively to the visualization of the purely hydrodynamic cylindrical laboratory experiment shown in Figure 1. Thus, Figure 1 highlights the ability to compare a hydrodynamic laboratory experiment with a cylindrical geometry to a spherical shell DNS model that uses a magnetohydrodynamic description.

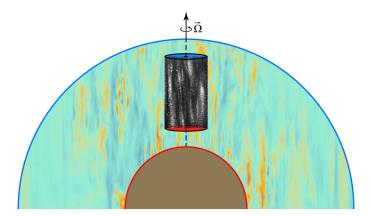


Figure 1. Schematic of the northern hemisphere of Earth's core region, with a solid inner core (tancolored) of radius  $r=0.35R_{\rm core}$ , rotating about the z-axis (dashed line) at rate  $\vec{\Omega}$ . The cylinder located at a high latitude and aligned with the rotation axis is representative of the laboratory containers used in this work. The interior of the cylinder is filled with an illustrative subset of the visualization of a geostrophically turbulent laboratory experiment ( $E\simeq 5\times 10^{-6}$ ,  $Ra\simeq 2\times 10^{10}$ ,  $Pr\simeq 6$ ). The bottom (top) edges of the experimental cell are red (blue) to represent the warm (cold) fluid under the presence of convection in parallel with the inner core boundary (core–mantle boundary). A meridional slice of the instantaneous radial velocity from a dynamo simulation ( $E=3\times 10^{-7}$ , E=1), E=10.05) adapted from [15] occupies the fluid outer core region. Orange (teal) indicates radially outward (inward) motions.

Numerous studies have found that the stabilizing effect of rotation on turbulent convection acts to vertically align motion in the fluid bulk along the rotation axis, regardless of geometry [16–26]. In addition, many dynamo studies have also found that the large-scale

Fluids 2023, 8, 106 3 of 35

dynamics of core-style flows are largely governed, at leading order, by the non-magnetic dynamics of rapidly rotating, so-called quasi-geostrophic turbulence [27–37]. These results motivate further exploration of rotating convective turbulence, which is carried out in this work. Thus, the use of a purely hydrodynamic fluid (water) in this study can be used to explore relevant non-magnetic quasi-geostrophic turbulent dynamics.

Heat transfer is the most common diagnostic used to characterize convective dynamics in studies of RRBC (e.g., [37–45]). RRBC studies typically assume that global heat transfer will agree with and correspond to bulk velocity dynamics. Furthermore, many studies have found that convective heat transport is controlled by the dynamics occurring in the boundary layers of the system (cf. [21,37,46,47]). Contrastingly, the theory for the dynamics of the quasi-geostrophic turbulence thought to exist in planetary dynamo systems is independent of boundary physics and system diffusivities [48–53]. Thus, it remains unclear as to whether or not flow regimes indicated by the study of heat transfer alone are indeed representative of the bulk dynamical behavior of convection systems. Therefore, simultaneous measurements of heat transfer and bulk system velocities are needed.

The experiments in this study display heat transfer that is controlled by the boundary layer physics, while bulk velocities follow a diffusion-free scaling related to a balance between the Coriolis force, inertia, and buoyancy in the system (cf. [25,53–55]). We show this by simultaneously measuring convective velocities and heat transfer and directly comparing these measurements to results from DNS. As the velocity data appear to robustly follow a diffusion-free scaling, our results indicate that quasi-geostrophic bulk turbulence is reached even when heat transfer still depends on diffusive processes occurring in the boundary layers.

In Section 2, we detail the system parameters and scaling behaviors relevant to this work. In Section 3, we describe our experimental and numerical methods. The non-rotating convection results are described in Section 4.1 and the rotating convection results are presented in Section 4.2. The co-scaling of the two velocity predictions in this study, the Viscous–Archimedian–Coriolis (VAC) and Coriolis–Inertia–Archimedian (CIA) scalings, is discussed in Section 4.2.1. An analysis of local flow estimates and their comparison to asymptotic theory is provided in Section 4.2.2. Lastly, a discussion of this work and its implications for planetary dynamo systems is given in Section 5.

# 2. System Parameters and Scaling Behaviors

2.1. Rayleigh-Bénard Convection (RBC)

# 2.1.1. System Parameters

The consideration of non-rotating Rayleigh–Bénard convection (RBC) is an important precursor for the study of analogous rotating systems. The control parameters for an RBC system are the dimensionless Rayleigh and Prandtl numbers. The Rayleigh number, *Ra*, describes the ratio of thermal buoyancy to viscous and thermal diffusion and is given as:

$$Ra = \frac{\text{thermal buoyancy}}{\text{thermal \& viscous diffusion}} = \frac{\alpha g \Delta T H^3}{\nu \kappa},$$
 (1)

where  $\alpha$  (1/K) is the thermal expansivity of the fluid, g (m/s<sup>2</sup>) is the gravitational acceleration,  $\Delta T$  (K) is the temperature difference across the fluid layer, H (m) is the height of the fluid layer,  $\nu$  (m<sup>2</sup>/s) is the kinematic viscosity of the fluid, and  $\kappa$  (m<sup>2</sup>/s) is the thermal diffusivity of the fluid. The Prandtl number, Pr, is the ratio between viscous and thermal diffusion and is given as:

$$Pr = \frac{\text{viscous diffusion}}{\text{thermal diffusion}} = \frac{\nu}{\kappa}.$$
 (2)

Lastly,  $\Gamma$  is the aspect ratio of the cylindrical container:

$$\Gamma = \frac{D}{H},\tag{3}$$

where D = 2R is the diameter of the fluid layer.

Fluids 2023, 8, 106 4 of 35

# 2.1.2. RBC Heat Transfer Scaling Behavior

In RBC, the non-dimensional heat transfer is described by the Nusselt number, Nu, which is the ratio of the total heat flux through the system relative to the conductive heat flux in the absence of convection:

$$Nu = \frac{\text{total heat flux}}{\text{conductive heat flux}} = \frac{qH}{k\Delta T},$$
 (4)

where q (W/m<sup>2</sup>) is the total heat flux through the system and k (W/mK) is the thermal conductivity of the fluid. The Nusselt number describes the efficiency of global heat transport, and trends in the Nusselt number are related to fundamental convective behaviors (e.g., [35,56–59]).

In  $Pr \sim \mathcal{O}(1)$  RBC systems that are far past convective onset, a scaling relationship between Nu and the input parameter Ra exists of the form:

$$Nu = cRa^{\alpha}, \tag{5}$$

where c is the pre-factor and  $\alpha$  is the scaling exponent. Over the accessible range of parameter space covered to date by  $Pr \sim \mathcal{O}(1)$  laboratory experiments and DNS (i.e.,  $10^3 \lesssim Ra \lesssim 10^{15}$ ), the scaling exponent,  $\alpha$ , in (5) is found to vary between  $\alpha \simeq 2/7$  and  $\alpha \simeq 1/3$  [38,40,60–66]. The  $\alpha \simeq 1/3$  scaling relation arises when the fluid bulk becomes approximately isothermal and the boundary layers do not interact due to vigorous bulk convective mixing [45,67–74]. We illustrate the RBC heat transfer behavior we test in this study in Figure 2a, where the partially faded fuchsia line,  $Nu_0^*$ , corresponds to the empirical best fit of the experimental data in Cheng et al. [45], with a pre-factor of c=0.11 and an exponent of  $\alpha=0.308$ .

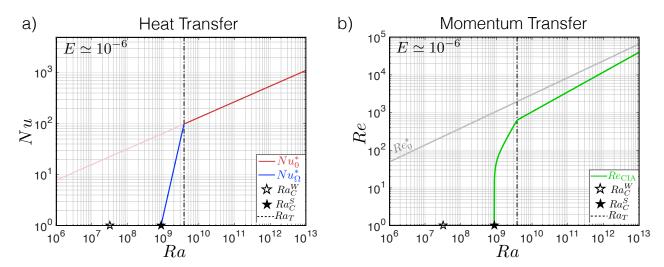


Figure 2. (a) Predicted Nu vs. Ra behavior when  $E \simeq 10^{-6}$ ,  $Pr \simeq 6$ . The partially faded fuchsia line,  $Nu_0^* = 0.11Ra^{0.308}$ , corresponds to RBC behavior [45]. For RRBC, the solid blue line,  $Nu_{\Omega}^* = (Ra/Ra_C^S)^{\beta}$ , where  $\beta = -0.45 - 0.58log_{10}(E)$ , is valid below the empirically defined transition Rayleigh number,  $Ra_T$  (dashed line) [3]. The open star denotes the onset of wall mode convection,  $Ra_C^W \simeq 31.8E^{-1}$ , and the solid star denotes the onset of steady convection,  $Ra_C^S \simeq 8.7E^{-4/3}$  [75,76]. When  $Ra > Ra_T$ , heat transfer is expected to return to the non-rotating scaling given by  $Nu_0^*$  (fuchsia). (b) Predicted Re vs. Ra behavior for  $E \simeq 10^{-6}$ ,  $Pr \simeq 6$ .  $Re_0^* = 0.102Ra^{0.447}$  is the non-rotating upper bound of the predicted RRBC behavior [77], where we expect Re to follow  $Re_{CIA} = (Ra(Nu-1)/Pr^2)^{2/5}E^{1/5}$  (green). The open and solid stars are the same as those in (a). Below  $Ra \leq Ra_T$ ,  $Nu_0^*$  is assumed in the calculation of Re, and above  $Ra > Ra_T$ ,  $Nu_0^*$  is also assumed.

Fluids 2023, 8, 106 5 of 35

# 2.1.3. RBC Momentum Transfer Scaling Behavior

The Reynolds number, *Re*, describes the momentum transfer in RBC systems and is the ratio of inertial advection to viscous diffusion, which is given as:

$$Re = \frac{\text{inertial advection}}{\text{viscous diffusion}} = \frac{uH}{v},$$
 (6)

where u (m/s) is a characteristic system velocity. Using a free-fall velocity,  $u_{\rm ff} \sim \sqrt{\alpha g \Delta T H}$ , to estimate the system scale velocity, u, in (6), a relationship between Re and Ra exists of the form:

$$Re = bRa^{\gamma} \sim Ra^{1/2} \tag{7}$$

when  $Pr \sim \mathcal{O}(1)$  [56,63,78–80]. A best fit exponent of  $\gamma \simeq 0.45$  has been found experimentally by Qiu and Tong [77]. The faded black line in Figure 2b shows Re as a function of Ra according to (7), with the best fitfit pre-factor, b=0.102, and exponent,  $\gamma=0.447$ , from [77]. This line serves as an upper-bounding estimate of the RRBC momentum transfer discussed in Section 2.2.3 below.

## 2.2. Rotating Convection (RRBC)

## 2.2.1. System Parameters

Two new non-dimensional control parameters are introduced to an RBC system with rotation: the Ekman number, *E*, and the Froude number, *Fr*. The Ekman number describes the ratio of viscous diffusion to the Coriolis force and is given as:

$$E = \frac{\text{viscous diffusion}}{\text{Coriolis}} = \frac{\nu}{2\Omega H^2},$$
 (8)

where  $\Omega$  (rad/s) is the angular rotation rate of the container. The Froude number characterizes the ratio of centrifugal buoyancy to gravitational buoyancy in the system and is given as:

$$Fr = \frac{\text{centrifugal buoyancy}}{\text{gravitational buoyancy}} = \frac{\Omega^2 R}{g},$$
 (9)

where R (m) is the radius of the cylindrical container.

Centrifugal buoyancy is argued to play a weak role in planetary core convection as estimates of the Froude number remain small, i.e.,  $Fr \lesssim 10^{-2}$ , in planetary dynamo systems. Thus, the DNS herein do not include centrifugal buoyancy effects. However, in our laboratory experiments, centrifugal buoyancy cannot be removed. According to Horn and Aurnou [81], heat transfer measurements are not significantly altered by centrifugal buoyancy until  $Fr \gtrsim \Gamma/2$ , i.e., when the Froude number exceeds half the aspect ratio of the container used. This constraint is only exceeded in our most extreme dataset, where  $E \simeq 1.27 \times 10^{-7}$  and  $\Gamma \simeq 0.74$ . The potential effect of the large centrifugal buoyancy force in this dataset is discussed in Section 4.2.

# 2.2.2. RRBC Heat Transfer Scaling Behavior

In a geometrically confined RRBC water system, convection will first onset in the form of wall modes [82–85]. The criterion for the onset of wall modes in low *E* RRBC under the assumption that the curvature of a confining cylinder can be neglected is given in Zhang and Liao [75] as:

$$Ra_C^W \simeq 31.8E^{-1}$$
. (10)

For Pr > 0.68, steady, bulk-filling convection is predicted to occur in a horizontally infinite plane layer in the limit of small E when: [76]

$$Ra_C^S \simeq 8.7E^{-4/3}$$
. (11)

Fluids 2023, 8, 106 6 of 35

After the onset of steady RRBC, different distinguishable scaling trends in heat transfer behavior have been observed and have been reviewed in Cheng et al. [3], Plumley and Julien [58], Kunnen [17], and Ecke and Shishkina [16].

Several recent studies have characterized the effects of lateral boundaries on rotating heat transfer [6–8,11,12,16,86–88]. In de Wit et al. [9] and Lu et al. [10], the 'sidewall circulation' in which enhanced heat transport occurs near container sidewalls is quantified. These studies also show that the bulk heat transfer in confined experiments is in good agreement with that of unconfined periodic DNS. The experiments herein do not have temperature sensors in locations that enable the quantification of heat transfer near container sidewalls. Nonetheless, in Section 4.2, we show that the measured heat transport data of the experiments in this work, which exist at least three times past steady onset, are in good agreement with those of laterally unconfined DNS.

Neglecting sidewall effects, a scaling relation for the heat transfer occurring in the regime past steady onset, where the entire fluid layer (i.e., the bulk and boundary layers) is thought to be rotationally constrained, can be expressed in terms of the criticality past steady onset as:

$$Nu_{\Omega} \sim (Ra/Ra_C^S)^{\beta}$$
 (12)

(e.g., [58]). As the Ekman number decreases,  $\beta$  steepens due to Ekman pumping effects, which boost the heat transfer [47,89,90]. Cheng et al. [3] estimate that the scaling exponent  $\beta$  in the rotationally constrained heat transfer regime depends on the Ekman number as: (cf. [10])

$$\beta = -0.45 - 0.59 \log_{10}(E). \tag{13}$$

The regime described by (12) is often referred to as the 'columnar' or 'convective Taylor column' regime, although we note that the direct connection between (12) and the existence of convective columns throughout this regime remains unclear (cf. [45,47,91]). Nonetheless, it is presently thought that convection in this regime occurs in the form of long, thin columns aligned with the axis of rotation when  $Ra/Ra_C^S \gtrsim 2$  (e.g., [18,82,92–96]). The cross-axial width of a steady RRBC column that develops at onset is: [76,97–102]

$$\ell_{\rm crit} \simeq 2.4 E^{1/3} H.$$
 (14)

Far past the onset of convection in RRBC, heat transfer behaves as if it is non-rotating according to (5) (e.g., [8,10,22,41,45,59,64,65,103]). The boundary layer physics involved in the transition of the RRBC heat transfer from being buoyancy dominated to rotationally constrained is complex, as elucidated in Julien et al. [89]. This boundary layer physics is not the focus of this study and, furthermore, we do not have temperature sensors in locations that enable us to resolve this physics. Nonetheless, we expect that the dominant boundary layer when RRBC heat transfer is buoyancy controlled can be estimated as the non-rotating thermal boundary layer:

$$\delta_T^{NR} \sim H/(2Nu),$$
 (15)

where  $\delta_T^{NR}$  (m) is the thickness of a single RBC thermal boundary layer [46,50,89].

RRBC heat transfer gradually becomes more and more rotationally constrained when lowering the convective supercriticality, and eventually, the viscous Ekman boundary layer becomes the dominant boundary layer, estimated here as: [64,97]

$$\delta_E \sim 3E^{1/2}H. \tag{16}$$

Thus, we predict that non-rotating-style heat transfer in RRBC will occur when the RBC thermal boundary layer becomes thinner than the Ekman boundary layer, i.e., when:

$$\delta_E/\delta_T^{\rm NR} > 1. \tag{17}$$

Numerous convection studies with a variety of applications have found that heat transport is controlled by the dominance of the relevant boundary layers. While the details of boundary

Fluids 2023, 8, 106 7 of 35

layer influence on heat transport vary, such applications include convection studies with non-uniform heating (e.g., [104–112]); studies with non-uniform boundaries [113–121]; and studies of rotating convection with libration, precession, or tidal deformation (e.g., [122–127]), among others.

Empirically, a 'transition' Rayleigh number,  $Ra_T$ , can be defined in an attempt to characterize the transition from non-rotating-style heat transfer to a rotationally constrained fluid layer. The empirical intersection of these two heat transfer scalings, (5) and (12), respectively, define  $Ra_T$  as: [65]

$$Ra_{T} = c^{1/(\beta - \alpha)} (Ra_{C}^{S})^{\beta/(\beta - \alpha)} = 8.7^{\beta/(\beta - \alpha)} c^{1/(\beta - \alpha)} E^{4\alpha/3(\alpha - \beta)}.$$
 (18)

Previous studies (e.g., [10,22,46,50,64,65,128,129]) find that the transition from non-rotating style heat transfer to the rotationally dominated heat transfer regime occurs over a gradual region that is not well described by (18). A number of different ideas exist regarding this gradual transition region (cf. [3,10,17]).

We choose to implement (18) in order to create a simple, testable framework for RRBC heat transfer behavior. Our framework does not incorporate recent findings surrounding wall mode heat transfer occurring near convective onset (i.e., [10]). Figure 2a is a 'strawman' that shows this framework by displaying Nu vs. Ra for an RRBC system with fixed  $E = 10^{-6}$ ,  $Pr \simeq 6$ . The onset of wall mode convection given by (10) is denoted by an open star and the onset of steady convection given by (11) is denoted by a solid star. After steady onset, the heat transfer efficiency, Nu, scales steeply with increases in Ra, as given by (12) (blue line). Past  $Ra_T$ , given by (18) (dashed line), the heat transfer behaves as though the system is not rotating according to (5) (fuchsia line).

#### 2.2.3. RRBC Momentum Transfer Scaling Behavior

The balance between buoyant energy production and dissipation in convection systems results in an exact relation between velocity and heat transfer. This holds for both RBC [62,130,131] and RRBC systems [25,55,132–135]. In RRBC, the use of this relation results in the generalized non-dimensional velocity scaling:

$$Re \sim \left[\frac{Ra(Nu-1)}{Pr^2}E\left(\frac{H}{\ell}\right)\right]^{\frac{1}{2}},$$
 (19)

where  $\ell$  (m) is the characteristic cross-axial length scale of convection. The term  $Ra(Nu-1)/Pr^2$  in (19) is the non-dimensional buoyancy flux (e.g., [134]). Thus, a non-dimensional velocity scaling in RRBC depends on: (1) the flux of buoyant energy; (2) the rotation period (via E); and (3) the anisotropy of the flow field,  $\ell/H$ .

Two theoretical length scales describing  $\ell$  exist in RRBC. The first is derived from linear theory [76] and describes the width of RRBC columns as given by (14):

$$\ell_{\rm crit}/H \sim E^{1/3}.\tag{20}$$

Substituting (20) into (19), we recover:

$$Re_{VAC} \sim \left(\frac{Ra(Nu-1)}{Pr^2}\right)^{\frac{1}{2}} E^{\frac{1}{3}}.$$
 (21)

Equation (21) is referred to as a 'VAC' scaling due to the underlying assumption that a triple balance exists between the Viscous–Archimedean (buoyancy)–Coriolis forces at steady onset [22,24,27,55,134,136].

The second length scale in RRBC provides a different scaling of *Re* than that of (21). Inertial theory predicts that: [54,137]

$$\ell_{\rm turb}/H \sim Ro^{1/2},\tag{22}$$

Fluids 2023, 8, 106 8 of 35

where Ro = ReE is the Rossby number. Ro describes the ratio of inertial advection to the Coriolis force:

 $Ro = \frac{u}{2\Omega H}. (23)$ 

The substitution of (22) into (19) yields:

$$Re_{\text{CIA}} \sim \left(\frac{Ra(Nu-1)}{Pr^2}\right)^{\frac{2}{5}} E^{\frac{1}{5}}$$
 (24)

and arises from a triple balance between the Coriolis–Inertia–Archimedean (buoyancy) forces, resulting in a 'CIA' scaling for *Re* [8,22,55,136,138].

Because  $Re \gg 1$  and  $Ro \ll 1$  in the fluid bulk of our experiments, viscous forces are subdominant to inertial forces, which, in turn, are subdominant to the Coriolis force. Thus, we expect diffusivity-free rotating turbulent flow to dominate bulk RRBC motions, as should be described by the CIA scaling, (24). For completeness, experimentally measured values of Re are compared to both theoretical scalings given by (21) and (24), and the results of this comparison are discussed in Section 4.2. Figure 2b is a 'strawman' that displays the predicted behavior of Re vs. Ra that we test in this study, where the green curve is given by the CIA scaling, (24) (with an assumed pre-factor of unity). Below  $Ra \leq Ra_T$ ,  $Nu_0^*$  is assumed in the calculation of Re shown by the green curve, and above  $Ra > Ra_T$ ,  $Nu_0^*$  is assumed.

#### 3. Methods

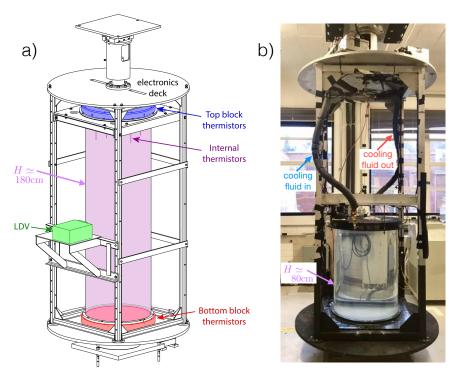
# 3.1. The NoMag Laboratory Device

RBC and RRBC experiments are conducted in axially aligned cylindrical containers filled with water. The 'NoMag' device used is comprised of a 1.27 cm thick optically clear cylindrical acrylic sidewall (thermal conductivity k=0.19 W/mK) of variable height bounded by two aluminum thermal blocks (k=167 W/mK) that allow for the exchange of heat (Figure 3). The Biot number, which characterizes the isothermality of the boundaries relative to that of the interior, is small in all experiments, i.e.,  $Bi \lesssim 0.05$ . Thus, we argue that thermal boundary gradients will only be a few percent of those within the fluid.

Convection is generated by passing a fixed heat flux through the bottom thermal block via a non-inductively wound electrical resistance heating pad. The fluid layer is cooled from above by a double-wound spiral heat exchanger in the top thermal block. Thermostated water is circulated through the heat exchanger to maintain a fixed temperature. The entire container and its aluminum structural frame are rotated about a vertical axis at a uniform rate varying from 0 to 55 revolutions per minute (rpm).

Figure 3a shows a schematic of the NoMag device with its tallest tank of height  $H\simeq 1.85$  m. Present diagnostic systems are also shown. Temperature measurements of the fluid layer are collected by 16 thermistors located within 1 cm of the fluid inside the top (blue) and bottom (red) boundaries, providing accurate measurement of the vertical temperature difference across the fluid layer. Thermistors can also be placed at various depths inside the fluid layer (purple) via ports in the top thermal block to sense the internal dynamics of the system. A custom calibration of all sensors is performed so that the precision of temperature measurements is within  $\pm 50$  mK. An MSE UltraLDV (green) is mounted in the rotating frame to conduct laser Doppler velocimetry (LDV) measurements and is placed at various heights and radial locations to collect point velocity measurements. Figure 3b displays an image of the device with the  $H\simeq 0.8$  m tank in place.

Fluids 2023, 8, 106 9 of 35



**Figure 3.** (a) Schematic of the NoMag laboratory device with the H=1.85 m acrylic container (purple). Diagnostic systems of the device include: (i) temperature measurements at the top (blue) and bottom (red) of the fluid boundaries; (ii) internal fluid temperature measurements (purple); (iii) point velocity measurements via laser Doppler velocimetry (LDV) (green). All electronics, including those used for data acquisition, are mounted onto the upper table structure in the rotating frame of the device and are labeled in black. (b) An image of the device with the H=0.802 m tank filled with water. The inlet (outlet) cooling fluid channel of the heat exchanger in the top thermal block is labeled in blue (red).

Three tank heights are used in this study to access a broad range of parameter space. These tanks have a fixed inner diameter of D=0.586 m and heights of H=0.202 m ( $\Gamma\simeq 3$ ), H=0.401 m ( $\Gamma\simeq 1.5$ ), and H=0.802 m ( $\Gamma\simeq 0.75$ ). In order to vary Ra, our control parameter is a flux-based Rayleigh number in which:

$$Ra_F = \frac{\text{buoyancy flux}}{\text{thermal \& viscous diffusion}} = \frac{\alpha g H^4 Q}{\rho_0 C_p \kappa^2 \nu A} = NuRa,$$
 (25)

where Q (W) is the heating power,  $C_p$  (J/kgK) is the specific heat capacity of the fluid, and A (m<sup>2</sup>) is the surface area of the thermal blocks in contact with the fluid. We varied Q from 30 to 1150 W in our experiments. Thus, our thermal control parameter, Q, which controls  $Ra_F$ , indirectly sets  $\Delta T$ , which controls Ra.

The control parameters for each experiment and all other experimental values can be found in Tables A1 and A2 in Appendix A. We also show the range of RRBC control parameters covered experimentally and numerically in this study in Table 1 below. For comparison, estimates of these parameters for Earth's outer core are also listed in Table 1 [139]. By examining the differences between the experimental and DNS parameters and those of the Earth's core, we see that the flux Rayleigh and Ekman numbers in our study differed by orders of magnitude from those of Earth's core, as is the case in any of the most extreme models of core dynamics to date (e.g., [34,45,140]). Nonetheless, rapidly rotating convective turbulence relevant to core dynamics is achieved in many of our extreme RRBC experiments (cf. Section 4.2). The Prandtl number of the DNS and experiments is moderate, whereas Pr for Earth's core is small (i.e.,  $Pr \lesssim 0.67$ ). Many similar fluid characteristics of non-magnetic rotating turbulence are found in previous studies with varied Pr [25,45,96,141]. In terms of

Fluids 2023, 8, 106 10 of 35

the Froude number, we carefully analyze our data for any possible effects due to centrifugation in our experiments with higher values of Fr and do not find any noticeable effects (cf. Section 4.2). Finally, we argue that our range of aspect ratios is relevant and interesting to study in relation to understanding core dynamics (cf. Section 1).

**Table 1.** RRBC control parameters for NoMag (this study), DNS (this study), and Earth's core [139].  $Ra_F$  is the flux Rayleigh number given by (25), E is the Ekman number given by (8), Pr is the Prandtl number given by (2), E is the Froude number given by (9), and E is the aspect ratio: for the NoMag and DNS experiments, E is defined as the diameter-to-height ratio of the cylinder used/simulated as given by (3), while for Earth's core, we provide the aspect ratio as the radius ratio of Earth's core relative to Earth's entire radius.

	$Ra_F$	E	Pr	Fr	Γ
NoMag	$5.0 \times 10^9$ – $6.1 \times 10^{13}$	$1.1 \times 10^{-7}$ – $2.9 \times 10^{-5}$	5.7-6.4	0.0018-1.0	0.73-2.9
DNS	$3.5 \times 10^8 - 1.4 \times 10^{11}$	$2.9 \times 10^{-6}$	5.9	0	0.74
Earth's Core	$4.0 \times 10^{29}$	$1.3 \times 10^{-15}$	0.10	0.0011	0.35

Each experiment is allowed to equilibrate thermally until the mean temperature on each thermal sensor does not change by more than approximately 1 percent over the course of 2 h. This process takes between 8 to 14 h on average for each experiment. Equilibrated data are then recorded for an additional 6–8 h, on average, at a rate of 10 samples per second. A minimum of  $\sim$ 1000 free-fall times exist for equilibrated data. Thus, each experiment is run for approximately 14–22 h total, on average. The vertical thermal diffusion time of the system,  $\tau_{\kappa} = H^2/\kappa$ , ranges between 3 and 50 days amongst all three containers used. Thus, while the total experimental time is small relative to the thermal diffusion time of our system, the time series of the portion of data analyzed always reaches a statistically steady state as described.

The relatively large  $\sim\!60$  cm tank diameter requires careful treatment of thermal losses through the  $\sim\!1$  cm thick acrylic sidewalls. The temperature of the room is set to be the same as the mean temperature of the fluid in order to minimize potential losses. The sidewalls are also wrapped in several layers of insulation: first, in a  $\sim\!0.3$  cm thick layer of foam insulation with a conductivity of 0.036 W/mK, next, in a layer of mylar to minimize radiative losses, and then, in a  $\sim\!2$  cm thick layer of foam. To further separate the fluid temperature from any potential influence of the ambient room, the outermost foam layer is surrounded by a  $\sim\!10$  cm thick layer of insulating styrofoam packing peanuts enclosed in plastic wrap. Horizontal conductive heat losses across the acrylic sidewall are calculated using recorded values of the room and mean fluid temperatures for each experiment. Thermal losses account for less than  $\sim\!5\%$  of the total input heating power for all cases.

# 3.2. Laser Doppler Velocimetry (LDV)

LDV enables the acquisition of point velocity measurements with small spatial and temporal resolution and with virtually no disturbance of the flow [142–145]. Titanium dioxide ( $TiO_2$ ) seeding particles are used as optical scattering sources ( $\sim$ 2 g/L concentration). A Measurement Science Enterprise (MSE) UltraLDV instrument is mounted in the rotating frame of the experiment and aligned in order to measure vertical convective flow velocities. A 130 mW split laser beam is passed through the acrylic cylindrical sidewall and into the water in the experimental tank. The split beams converge and intersect at a fixed distance inside the tank's sidewall, forming an interference pattern. In this intersection volume, passing particles reflect a signal back to the instrument, where the recorded frequency of the reflected particle light is then converted to a local velocity. For further details, see [146–148].

We verified our LDV data via spin up experiments on the NoMag device (cf. [7,148–150]). Several experimental parameters are varied, including the initial rotation rate, incremental change in the rotation rate, and tank height, in order to confirm

Fluids 2023, 8, 106 11 of 35

agreement with the well-characterized spin up theory. For all experiments, measured spin up times agree with respective theoretical spin-up calculations to within  $\pm 5\%$ . From these spin up experiments, the resolution of the velocity measurements using our device is determined to be  $\pm 2.0 \times 10^{-3}$  m/s [148]. Thus, we choose a resolution limit of  $\pm 2.5 \times 10^{-3}$  m/s for this work. Our analysis of the velocity measurements in Section 4 does not rely on any data with measured velocities smaller than this experimentally determined limit.

For all experiments in this study, the vertical root-mean-square (rms) velocity,  $u_{z,rms}$ , is calculated using the LDV velocity time series as:

$$u_{z,\text{rms}} = \sqrt{\frac{1}{N} \sum_{n=1}^{N} |u_{z,n}|^2},$$
 (26)

where  $u_{z,n}$  (m/s) is the vertical velocity time series containing N points. The measured Reynolds numbers are computed as:

$$Re = \frac{u_{z,\text{rms}}H}{v}.$$
 (27)

Note that weighting to account for any potential sampling biases in the measured velocities is determined to be unnecessary due to the normal/Gaussian statistical nature of the velocities in this work (cf. [148,151–153]). The Gaussian behavior of our experimental RRBC vertical velocities (all existing at  $Ra/Ra_C^S \gtrsim 60$ ) is consistent with the results of the recent numerical study by Aguirre-Guzman et al. [140]. For  $Pr \simeq 5$ , they find that above  $Ra/Ra_C^S \simeq 10$ , skewness and kurtosis statistics show that vertical velocities are isotropically turbulent (cf. [154,155]). They further show that while the vertical velocities are interestingly homogeneously turbulent past a supercriticality of  $\sim 10$ , the vertical vorticity, i.e., the vertical component of the curl of the velocity that displays the horizontal swirling nature of the flow, is anisotropic and non-Gaussian, indicating that the convective flow structures are multi-scale, as found in previous studies (e.g., [21,96,156]).

A fixed location in the fluid bulk for the point velocity data collection is chosen for each experimental container. The choice of location for the RRBC measurements can be anywhere within the bulk because the vertical flow is isotropically turbulent. In Section 4.1, we discuss how the location of the LDV point measurements may affect our non-rotating momentum transfer results. In all  $\Gamma \simeq 3$  experiments, data are collected at a non-dimensional location in the fluid bulk of  $h_{\rm LDV} \simeq 0.33H$ ,  $r_{\rm LDV} \simeq 0.84R$ , where R is the radius of the container, i.e., R = D/2. For all experiments in both the  $\Gamma \simeq 1.5$  and  $\Gamma \simeq 0.75$  tanks, the non-dimensional location of the velocity measurements is  $h_{\rm LDV} \simeq 0.57H$ ,  $r_{\rm LDV} \simeq 0.84R$ . The shorter  $\Gamma \simeq 3$  tank required a different non-dimensional height location in order to mount the LDV device. Once thermal equilibration is determined, a minimum of 2–4 h of data for rotating cases and a minimum of 4–6 h of data for non-rotating cases are recorded. The LDV data sampling rate is non-constant; on average, the sampling rates are a minimum of  $\sim$ 5 samples per second.

# 3.3. Direct Numerical Simulations (DNS)

Laboratory access to RRBC near steady onset is limited, requiring both a small input thermal forcing and a rapid rotation of the container, a setting in which it is difficult to control the thermal losses in experiments (cf. [10,45]). Thus, a set of fixed  $E \simeq 3 \times 10^{-6}$  DNS is conducted at low Ra values and  $Pr \simeq 6$  in order to examine the RRBC behavior near onset. We use the fundamental results of these DNS to further support and reinforce the results of the laboratory experiments herein, which are the focus of this study. Data from 10 novel DNS are provided in Table A3 in Appendix A. The DNS model a Cartesian fluid layer and do not include centrifugal buoyancy. The code used has been validated and employed in prior studies (see [47,64,65,157] for more details). The top and bottom boundaries are isothermal, rigid, and non-slip, with periodic horizontal boundary conditions. Fourier expansions are implemented in the horizontal directions and Chebyshev polynomials are used in the vertical direction.

Fluids **2023**, 8, 106 12 of 35

The vertical resolution is set in order to maintain at least 16 grid points within the Ekman boundary layer. The box aspect ratio is set as  $\Gamma=D/H\simeq 0.74$ , matching the aspect ratio of the  $H\simeq 0.8$  m experimental container. The onset of steady convection for these  $E\simeq 3\times 10^{-6}$  DNS is  $Ra_C^S\simeq 1.84\times 10^8$ . With no sidewall boundaries, we note that our DNS do not capture the heat transfer due to wall modes (cf. [6,9,10]). The reported Reynolds numbers of the DNS contain the time average of the rms of the vertical z-velocity at a non-dimensional height of z=0.57 to match the height location at which the majority of experimental LDV data are collected. The number of free-fall times per simulation ranges from  $2.53\times 10^3$  to  $2.05\times 10^5$ , and the number of thermal diffusion times captured ranges from  $2.53\times 10^{-2}$  to 5.67.

#### 4. Results

# 4.1. RBC Heat and Momentum Transfer Measurements

Figure 4a shows our RBC heat transfer results, along with the data from [64,65]. The Nusselt number, Nu, is displayed on the y-axis and the Rayleigh number, Ra, is shown on the x-axis. Pink and cyan markers represent our data, where symbol shapes represent the different tanks used: circles ( $\Gamma \simeq 3$ ), squares ( $\Gamma \simeq 1.5$ ), and diamonds ( $\Gamma \simeq 0.75$ ). The  $\Gamma \simeq 3$  data are distinguished in pink from the data collected in the  $\Gamma \simeq 1.5$  and  $\Gamma \simeq 0.75$  tanks (cyan) due to the different location of velocity measurements in the  $\Gamma \simeq 3$  tank. The open gray triangles [64] and open upside-down gray triangles [65] are water data collected using UCLA's smaller-diameter RoMag device. The fuchsia line in Figure 4a,  $Nu_0 = 0.108Ra^{0.309}$ , represents the best fit to our data. The scaling exponent,  $\alpha = 0.309 \pm 0.003$ , and pre-factor,  $c = 0.108 \pm 0.009$ , are in good agreement with those found in Cheng et al. [45]:  $\alpha = 0.308$  and c = 0.11.

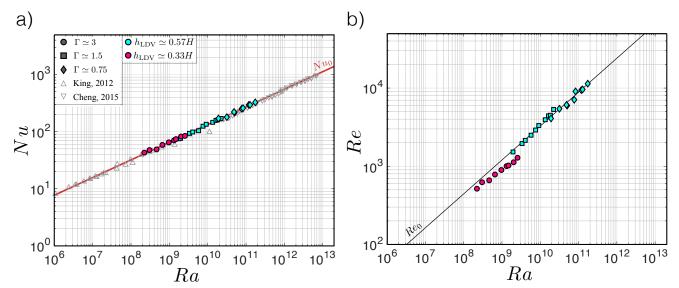


Figure 4. (a) Nu vs. Ra RBC results, where pink circles denote data with  $\Gamma \simeq 3$ , cyan squares denote data with  $\Gamma \simeq 1.5$ , and cyan diamonds denote data with  $\Gamma \simeq 0.75$ . Open gray triangles are the experimental data from [64], whereas upside-down gray triangles are the experimental data from [65]. The fuchsia line is the best fit to our data,  $Nu_0 = 0.108Ra^{0.309}$ , and is in good agreement with the empirical fit to the data from [45]. (b) Re vs. Ra RBC results, where the symbol shapes and colors are consistent with those in (a). The  $\Gamma \simeq 3$  data are collected at a location of  $h_{\rm LDV} \simeq 0.33H$ ,  $r_{\rm LDV} \simeq 0.84R$ , while the  $\Gamma \simeq 1.5$  and  $\Gamma \simeq 0.75$  data are collected at a location of  $h_{\rm LDV} \simeq 0.57H$ ,  $r_{\rm LDV} \simeq 0.84R$ . The solid black line,  $Re_0 = 0.151Ra^{0.434}$ , is the best fit to the data at the height of  $h \simeq 0.57H$ , and is in good agreement with the best fit to the water data in [77].

The RBC momentum transfer results of this study are shown in Figure 4b, with the Reynolds number, Re, displayed on the y-axis and the Rayleigh number, Ra, shown on the x-axis. The solid black line,  $Re_0 = (0.150 \pm 0.052)Ra^{(0.434 \pm 0.012)}$ , shows the best fit to

Fluids 2023, 8, 106 13 of 35

the  $\Gamma \simeq 1.5$  (squares) and  $\Gamma \simeq 0.75$  (diamonds) data collected at  $h_{\rm ldv} \simeq 0.57H$ ,  $r_{\rm ldv} \simeq 0.84R$ . This fit is in adequate agreement with the best fit of the experimental water data from Qiu and Tong [77]:  $Re_0^* = 0.102Ra^{0.447}$ .

The  $\Gamma \simeq 3$  pink data collected at  $h_{\rm ldv} \simeq 0.33H$ ,  $r_{\rm ldv} \simeq 0.84R$  does not follow the same scaling behavior as in our other two tanks. This is likely due to the difference in the location of the point velocity measurements and the regime of convection probed. In all of the RBC experiments, a large-scale overturning circulation (LSC) pattern should exist in the fluid bulk. The velocity field of such an LSC structure is non-uniform (aka position dependent) within the fluid bulk (e.g., [71,77,158–161]). It is possible that the  $\Gamma \simeq 3$  data at a non-dimensional height of  $h_{\rm ldv} \simeq 0.33H$  samples a corner of an LSC structure, whereas the rest of our data at a height of  $h_{\rm ldv} \simeq 0.57H$  (and larger Ra) measure velocities nearer to the center of the vertical upwellings and downwellings that comprise the LSC. Overall, our RBC heat and momentum transfer measurements agree well with the results of prior studies.

## 4.2. RRBC Heat and Momentum Transfer Measurements

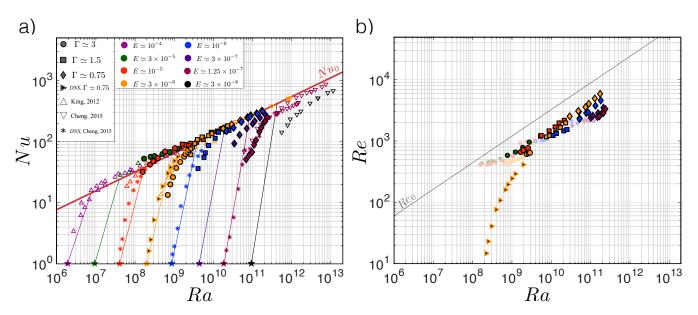
Figure 5a shows Nu versus Ra for our RRBC heat transfer measurements. Symbol color represents the non-dimensional rotation period, E. The different symbol shapes represent the data collected in the three different experimental containers used in this study and are consistent with Figure 4. The  $E \simeq 3 \times 10^{-6}$  DNS are displayed in Figure 5 as right pointing orange triangles with a black center. The open triangles represent data from [64], the open upside-down triangles represent data from [65], and the asterisks denote the direct numerical simulations from [65]. The onset values for steady bulk convection, given by (11), are shown as solid, colored stars on the x-axis. The colored, thin solid lines characterize the heat transfer regime occurring near the onset of convection given by (12) and (13). The numerical simulations reach closest to onset and provide the most data in this regime. The thick solid fuchsia line represents the best fit to the RBC data from Figure 4a.

Cheng et al. [65] note that it is possible that their two most extreme datasets at  $E \simeq 10^{-7}$  and  $E \simeq 3 \times 10^{-8}$  might be influenced by strong centrifugal acceleration, which scaled as  $\Omega^2 R$ . This may also be the case for our burgundy data at  $E \simeq 1.25 \times 10^{-7}$  ( $Fr \simeq 1 \geq (\Gamma/2 = 0.375)$ , [81]). However, our  $E \simeq 1.25 \times 10^{-7}$  data agree well with the  $E \simeq 10^{-7}$  numerics of [47] in which there is no centrifugal buoyancy. Thus, it is not obvious that the overall scaling behavior of this heat transfer data is altered by centrifugation (c.f. [9,45,162,163]). We conclude that our heat transfer data (both experimental and numerical) adequately agree with prior studies.

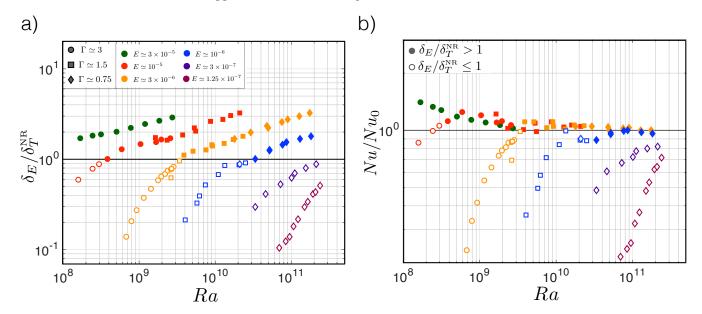
Figure 5b shows the *Re* versus *Ra* data for our RRBC experiments and DNS. The color and symbol shape are the same as those in Figure 5a, with one difference: faded symbols represent experimental data with measured velocities that are below the determined resolution limit of 2.5 mm/s for the LDV device used in this study (see Section 3.2). The solid, colored symbols thus represent the experimental velocity data above the 2.5 mm/s LDV device resolution limit that are used in the analysis of the RRBC momentum transfer data discussed below. The solid, faded black line is the best fit to the RBC momentum transfer data shown in Figure 4b, and serves as an upper bound for our rotating data. We note that the burgundy data ( $E \simeq 10^{-7}$ ,  $Er \simeq 1$ ) in Figure 5b do not show any observable influence of centrifugal buoyancy on vertical momentum transport.

Figure 6a displays the ratio of boundary layer thickness estimates,  $\delta_E/\delta_T^{\rm NR}$ , versus the Ra for the laboratory data. The RBC thermal boundary layer thickness,  $\delta_T^{\rm NR}$ , is estimated using  $\delta_T^{\rm NR} = H/(2Nu)$  given by (15) and the Ekman layer thickness,  $\delta_E$ , is estimated using  $\delta_E = 3E^{1/2}H$  given by (16) such that  $\delta_E/\delta_T^{\rm NR} = 6NuE^{1/2}$ . Previous studies have shown that whether RRBC heat transfer will behave as though the system is non-rotating versus being rotationally constrained depends on the overall dominance of an Ekman or a non-rotating thermal boundary layer in the system [21,46,47,64,89,164]. Thus, in Figure 6a we use hollow, colored symbols when  $\delta_E/\delta_T^{\rm NR} \leq 1$ , and symbols become solid when  $\delta_E/\delta_T^{\rm NR} > 1$ .

Fluids 2023, 8, 106 14 of 35



**Figure 5.** (a) Nu vs. Ra RRBC results, where circles denote laboratory data with  $\Gamma \simeq 3$ , squares denote laboratory data with  $\Gamma \simeq 1.5$ , and diamonds denote laboratory data with  $\Gamma \simeq 0.75$ . Color represents the non-dimensional rotation period, E. The right pointing orange triangles with black centers represent the fixed  $\Gamma \simeq 0.74$ ,  $E \simeq 3 \times 10^{-6}$  DNS of this study. Open, colored triangles represent the data from [64] and open, colored upside-down triangles represent the data from [65]. The asterisks denote the DNS from [65]. The onset of steady convection, given by (11), is shown as solid, colored stars. The colored, thin solid lines past onset characterize the regime given by (12) and (13). The thick fuchsia line shows the RBC empirical findings from Figure 4a. (b) Re vs. Ra RRBC results, where the colors and symbols are the same as those in (a) but with one difference: faded symbols represent the data with measured velocities that are below the determined resolution limit of 2.5 mm/s for the LDV device used in this study while solid symbols represent the data above this resolution limit. The solid, faded black line is the best fit to the RBC momentum transfer data shown in Figure 4b, serving as an upper bound for the rotating data.



**Figure 6.** (a) Estimated Ekman layer thickness,  $\delta_E=3E^{1/2}H$  [97], normalized by the estimated RBC thermal boundary layer thickness,  $\delta_T^{\rm NR}=H/(2Nu)$  [64], vs.  $\it Ra$ . Colored symbols are hollow when  $\delta_E/\delta_T^{\rm NR}\leq 1$  and solid when  $\delta_E/\delta_T^{\rm NR}>1$ . (b)  $\it Nu$  normalized by the non-rotating style scaling,  $\it Nu_0=0.108Ra^{0.309}$ , vs.  $\it Ra$ . Colored symbols are hollow or solid based on the criteria used in (a). We see that hollow symbols generally occur where  $\it Nu/Nu_0\lesssim 1$ .

Fluids 2023, 8, 106 15 of 35

Figure 6b displays the laboratory Nu data normalized by our empirical non-rotating style scaling,  $Nu_0 = 0.108Ra^{0.309}$ , plotted versus Ra. The same symbol significance used in Figure 6a is used here, that is, symbols are hollow if  $\delta_E/\delta_T^{\rm NR} \leq 1$  and solid if  $\delta_E/\delta_T^{\rm NR} > 1$ . We see that the symbols are hollow when  $Nu/Nu_0 \lesssim 1$ , whereas the symbols become solid roughly when  $Nu/Nu_0$  becomes unity. Thus, when  $\delta_E/\delta_T^{\rm NR} \leq 1$ , the Ekman layer remains thin, resulting in rotationally constrained heat transfer behavior. Contrastingly, when  $Nu/Nu_0$  roughly becomes unity, the thermal boundary layer becomes thinner than the Ekman layer and non-rotating-style heat transfer behavior is recovered. We conclude that the behavioral transitions in our RRBC heat transfer data correlate to the boundary layer physics of the system, consistent with prior studies (e.g., [22,47,59,128]).

Figure 7 shows the measured Re values from Figure 5b with the scaling predictions from Figure 2d separately for each fixed E dataset. Note that only the laboratory data that are above the LDV device resolution limit are included in each panel in Figure 7. Experimental data below the LDV device resolution limit (faded symbols in Figure 5b) are not included in any of the ensuing momentum transfer analysis. The inclusion of the orange  $E \simeq 3 \times 10^{-6}$  DNS in panel c in Figure 7 allows us to examine the Re behavior closer to the onset of convection than is possible with the laboratory experiments.

The empirical best fits of  $Nu_0$  and  $Nu_\Omega$  from Figure 5a are used in the computation of the green Re CIA theoretical curves shown in each panel in Figure 7 and given by (24). Specifically, below  $Ra \leq Ra_T$ ,  $Nu_\Omega$  is used in the calculation of these green curves, and above  $Ra > Ra_T$ ,  $Nu_0$  is used. A pre-factor of unity is assumed in the calculation of these curves. All of our measured laboratory Re values align well with the green CIA Re scaling predictions for each fixed E dataset in Figure 7. The DNS in panel c have decent agreement with the theoretical CIA prediction. We note that full volume average calculations of the DNS Re values more directly overlay onto the green CIA theory curve in panel c.

# 4.2.1. Co-Scaling of the VAC and CIA Predictions

Figure 8a shows measured Ro versus the  $Ro_{\text{CIA}}$ . A best fitfit line on this graph yields  $Ro = (0.87 \pm 0.03) Ro_{\text{CIA}}^{(0.98 \pm 0.01)}$ , indicating excellent agreement between experimental velocity measurements and associated predictions using the triple-balance assumption of the CIA scaling given by (24). Thus, our velocity measurements indicate that the Coriolis force, inertial advection, and buoyancy control the bulk momentum transport. Figure 8b displays measured Ro versus the  $Ro_{\text{VAC}}$ . Decent agreement also exists between LDV measurements and associated VAC predictions. Specifically, a best fitfit line on this graph yields  $Ro = (0.29 \pm 0.02) Ro_{\text{VAC}}^{(0.91 \pm 0.01)}$ . The VAC scaling prediction given by (21) assumes that viscous diffusion, instead of inertial advection, acts in balance with the Coriolis force and buoyancy to control the momentum transport.

Figure 8a,b indicate that both the VAC and CIA scalings of Re given by (21) and (24) adequately predict momentum transport behavior. To further highlight this, Figure 8c shows normalized Rossby numbers, Ro, versus Ra. Ro is defined as laboratory LDV values normalized by either CIA estimates,  $Ro = Ro/Ro_{CIA}$  (green), or by VAC estimates,  $Ro = Ro/Ro_{VAC}$  (magenta). The green  $Ro/Ro_{CIA}$  points are well-centered around 1, indicating that the CIA scaling estimates our measured convective velocities well, as expected from the excellent best fit shown in Figure 8a. The magenta  $Ro/Ro_{VAC}$  points show that the VAC scaling overestimates measured values of Ro by roughly a factor of 2. The magenta VAC normalized points contain an overall flat slope, similar to that of the green CIA normalized points, indicating that the VAC scaling also captures the velocity behavior of the system, even if offset by a small, constant factor.

Fluids 2023, 8, 106 16 of 35

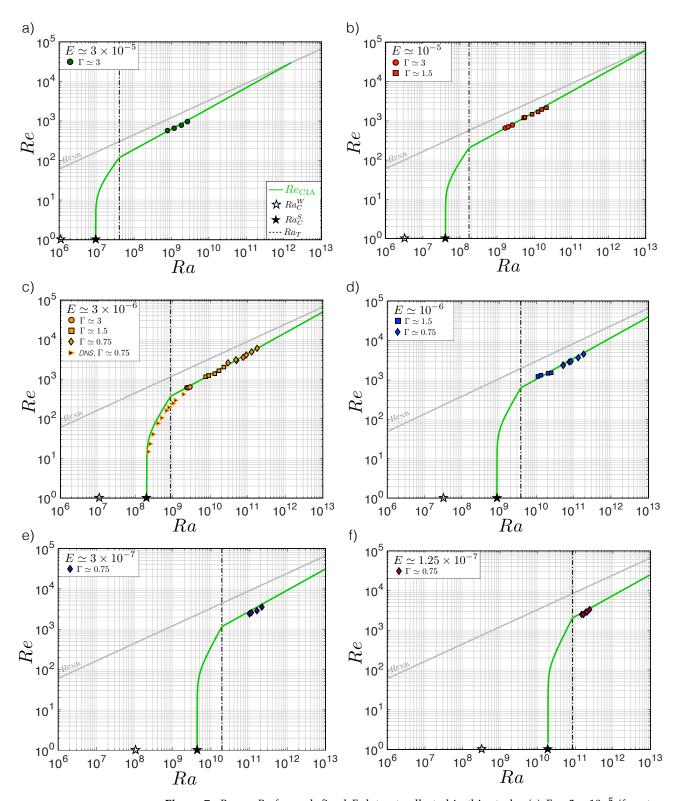
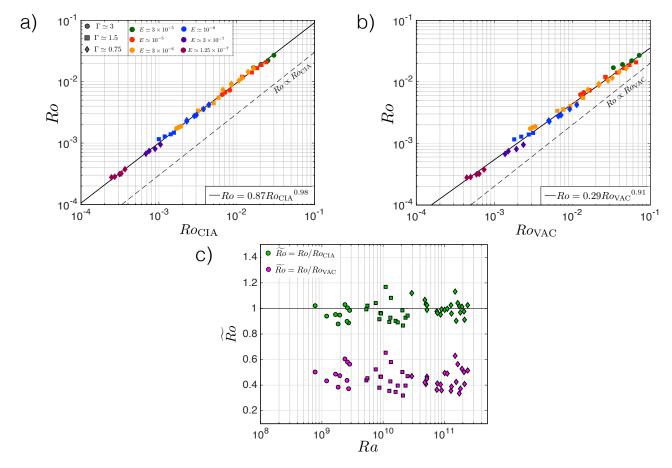


Figure 7. Re vs. Ra for each fixed E dataset collected in this study. (a)  $E \simeq 3 \times 10^{-5}$  (forest green), (b)  $E \simeq 10^{-5}$  (red), (c)  $E \simeq 3 \times 10^{-6}$  (orange), (d)  $E \simeq 10^{-6}$  (blue), (e)  $E \simeq 3 \times 10^{-7}$  (purple), and (f)  $E \simeq 1.25 \times 10^{-7}$  (burgundy). Symbol shape represents the experimental container used/simulated, consistent with Figure 5b. The theoretical scaling curves from Figure 2d at each respective fixed E value are plotted in each panel, where the green curves are given by  $Re_{\text{CIA}} = (Ra(Nu-1)/Pr^2)^{2/5}E^{1/5}$ . Below  $Ra \le Ra_T$ ,  $Nu_\Omega$  from Figure 5a is used in the calculation of these green curves, and above  $Ra > Ra_T$ ,  $Nu_\Omega$  is assumed. For each fixed E dataset, good agreement with CIA theory (green curve) is obtained.

Fluids 2023, 8, 106 17 of 35

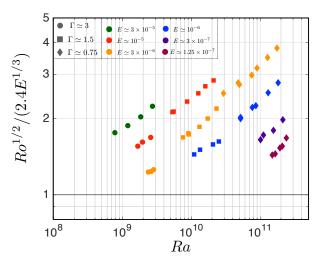


**Figure 8.** (a) Ro vs.  $Ro_{\text{CIA}}$  using laboratory data, where the best fitfit black line is given as  $Ro = 0.87Ro_{\text{CIA}}^{0.98}$ . The use of colors and symbol shapes is consistent with that of Figure 7. (b) Ro vs.  $Ro_{\text{VAC}}$ , where the best fitfit black line is given as  $Ro = 0.29Ro_{\text{VAC}}^{0.91}$ . (c) Normalized Rossby number, Ro, vs. Ra, where the green symbols contain a normalization on the y-axis of  $Ro = Ro/Ro_{\text{CIA}}$ , and the magenta symbols are a y-normalization of  $Ro = Ro/Ro_{\text{VAC}}$ . The green CIA normalized points are centered around 1, indicating that the CIA scaling estimates the measured values well, while the magenta VAC scaling overestimates the measured Ro values by roughly a factor of 2.

We conclude from Figure 8 that the CIA scaling predictions fit our data extremely well, in good agreement with the findings from planetary dynamo models (cf. [4,33,34,165,166]). No pre-factors are used in the calculations of the CIA and VAC scaling predictions given by (24) and (21), respectively. Importantly, the VAC scaling also fits our velocity data reasonably well over the full range of *Ra* investigated. Thus, the CIA and VAC scalings appeared to co-scale, with both adequately predicting the measured data.

The VAC and CIA momentum transfer scalings only differed in the assumed cross-axial length scale of the flow. This length scale is not directly measured in this study. We estimate the local onset and turbulent scales using (20) and (22), respectively. Figure 9 shows the ratio of these two length scales,  $Ro^{1/2}/(2.4E^{1/3})$ , plotted versus Ra for our RRBC experimental data. The ratio of the length scales in Figure 9 only differs by a factor of 4. Thus, the onset scale,  $\ell_{\rm crit}$ , associated with the VAC scaling and the turbulent scale,  $\ell_{\rm turb}$ , associated with the CIA scaling are not formally scaled separately in the range of the parameter space explored experimentally (cf. [28,167–169]). This is also the case for the DNS in this study.

Fluids **2023**, 8, 106 18 of 35



**Figure 9.** The turbulent length scale,  $\ell_{\rm turb} = Ro^{1/2}$ , normalized by the onset scale,  $\ell_{\rm crit} = 2.4E^{1/3}$ , for the experimental data of this study. The symbol shapes and colors are consistent with those in Figure 8a,b. The ratio  $\ell_{\rm turb}/\ell_{\rm crit}$  is of order unity for all experiments shown.

We conclude from Figure 9 that the VAC and CIA scalings both explained our velocity data because their underlying length scales are closely comparable in value. Thus, Figures 8 and 9 lead us to postulate that the two scalings of  $\ell$  discussed, the onset scale,  $\ell_{\rm crit}$ , and the inertial scale,  $\ell_{\rm turb}$ , must be comparable in this study. We now demonstrate that when local bulk Reynolds numbers,  $Re_{\ell}$ , are moderate (e.g., mathematically order unity) and local bulk Rossby numbers remain small (i.e.,  $Ro_{\ell} \lesssim \mathcal{O}(1)$ ), the two length scales that define the VAC and CIA scalings, respectively,  $\ell_{\rm crit}$  and  $\ell_{\rm turb}$ , will co-scale. In Section 4.2.2, we demonstrate that these two conditions are satisfied, i.e.,  $Re_{\ell} \sim \mathcal{O}(10^1-10^2)$  and  $Ro_{\ell} \lesssim \mathcal{O}(1)$ , for our RRBC laboratory data. Thus, the following analysis holds true for our experiments.

Using  $\ell_{\rm crit} \sim E^{1/3} H$ , the local Reynolds number can be defined as:

$$Re_{\ell} = \frac{u\ell_{\text{crit}}}{v} \sim ReE^{1/3}.$$
 (28)

Assuming that the local Reynolds number is mathematically of order unity, i.e.:

$$Re_{\ell} \sim ReE^{1/3} \sim \mathcal{O}(1),$$
 (29)

and since Ro = ReE, then,

$$Ro \sim (Re_{\ell}E^{-1/3})E \sim E^{2/3}.$$
 (30)

Using (30), we see that the inertial scale,  $\ell_{\text{turb}} \sim Ro^{1/2}H$ , scales equivalently to the onset scale:

$$\ell \sim Ro^{1/2}H \sim (E^{2/3})^{1/2}H \sim E^{1/3}H.$$
 (31)

Thus, we crucially note that when  $Re_{\ell} \sim \mathcal{O}(1)$  in RRBC systems, the onset and turbulent cross-axial scales are comparable, i.e.: (cf. [28,31,32,34])

$$\ell \sim E^{1/3} H \sim Ro^{1/2} H.$$
 (32)

The VAC scaling of (21) and the CIA scaling of (24) are thus comparable:

$$Re_{VAC} \sim Re_{CIA}$$
. (33)

From the above analysis and Figures 8 and 9, we conclude that the bulk dynamics are robustly governed by a Coriolis–Inertia–Archimedian (buoyancy) balance in the range of Ra explored in this study (cf. [4,34,53,59,170]), while  $\ell_{\rm turb}$  never departs significantly from  $\ell_{\rm crit}$ . We purport that viscous forces act secondarily on bulk dynamics under the CIA

Fluids **2023**, 8, 106 19 of 35

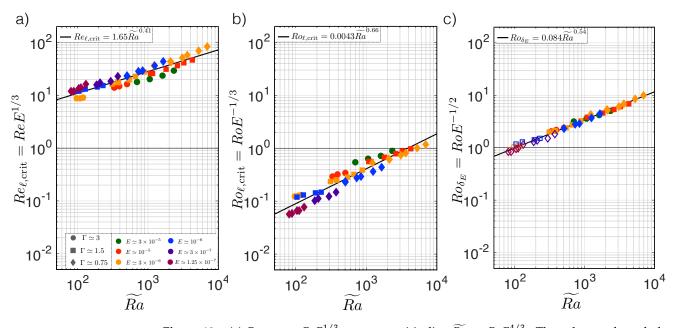
balance (cf. [89]), and either length scale estimate can be used to accurately predict the cross-axial scale of flows.

# 4.2.2. Local Flow Estimates and Comparison to Asymptotic Theory

Figures 8 and 9 motivate further exploration of local estimates within the bulk and boundary layers of our experiments. Using the onset scale given by (20) for  $\ell$ , a local Reynolds number for the fluid bulk can be defined as:

$$Re_{\ell,\text{crit}} = \frac{u\ell_{\text{crit}}}{v} = ReE^{1/3},$$
 (34)

where Re is the global Reynolds number computed by (27) for our experiments. The definition of  $Re_{\ell,\text{crit}}$  in (34) does not use any empirical pre-factors for simplicity. Figure 10a displays  $Re_{\ell,\text{crit}}$  versus  $\widetilde{Ra} = RaE^{4/3} \simeq Ra/Ra_C^S$ , a measure of the convective supercriticality of experiments. Color and symbol shape are consistent with that of Figure 9, and the best fit black line is given as  $Re_{\ell,\text{crit}} = (1.65 \pm 0.25)\widetilde{Ra}^{(0.41 \pm 0.02)}$ . Bulk local Re estimates are  $\mathcal{O}(10^1 - 10^2)$ . Note that the turbulent length scale given by (22) can instead be used for  $\ell$  in (34) and approximately identical results are obtained.



**Figure 10.** (a)  $Re_{\ell, crit} = ReE^{1/3}$  vs. supercriticality,  $\widetilde{Ra} = RaE^{4/3}$ . The colors and symbols are consistent with those in Figure 9. Bulk local Re estimates are  $\mathcal{O}(10^1-10^2)$ . (b)  $Ro_{\ell, crit} = RoE^{-1/3}$  vs.  $\widetilde{Ra}$ . Bulk local Re estimates remain small, i.e.,  $Ro_{\ell, crit} \lesssim 1$ . Panels (a,b) together further confirm that the bulk dynamics in all experiments are quasi-geostrophically turbulent. (c)  $Ro_{\delta_E} = RoE^{-1/2}$  vs.  $\widetilde{Ra}$ . For comparison with Figure 6, the hollow symbols indicate cases in which  $\delta_E/\delta_T^{\rm NR} \leq 1$ , whereas the solid symbols indicate cases in which  $\delta_E/\delta_T^{\rm NR} > 1$ . The estimated local Re0 values within the boundary layers using the Ekman boundary layer thickness are greater than unity (i.e., solid) for the majority of the data shown, in agreement with Figure 6 that shows that the majority of the data plotted in (c) are not strongly controlled by rotation.

We define a local bulk Rossby number as:

$$Ro_{\ell,\text{crit}} = \frac{u}{2\Omega\ell_{\text{crit}}} = RoE^{-1/3}.$$
 (35)

Figure 10b displays  $Ro_{\ell,\mathrm{crit}}$  versus  $\widetilde{Ra}$ , where the best fit black line is given as  $Ro_{\ell,\mathrm{crit}} = (0.0043 \pm 0.0009)\widetilde{Ra}^{(0.66 \pm 0.03)}$ . Bulk local Ro estimates remain small, i.e.,  $Ro_{\ell,\mathrm{crit}} \lesssim 1$ . Quasigeostrophic turbulence exists in the fluid bulk when  $Re_{\ell} \gtrsim \mathcal{O}(10)$  and  $Ro_{\ell} \lesssim 1$  (e.g.,

Fluids 2023, 8, 106 20 of 35

[5,52,167,171–174]). Thus, collectively, Figure 10a,b indicate, via local arguments of the fluid bulk, that the experiments plotted possess quasi-geostrophically turbulent bulk dynamics, consistent with the global analysis highlighted in Figure 8.

Local boundary layer dynamics are explored in Figure 10c, which displays  $Ro_{\delta_E}$  versus  $\widetilde{Ra}$ .  $Ro_{\delta_E}$  is a local Rossby number in the fluid boundary layers that we define as:

$$Ro_{\delta_E} = \frac{u}{2\Omega\delta_E} = RoE^{-1/2},\tag{36}$$

where  $\delta_E$  is the Ekman boundary layer thickness given by (16). No pre-factor is used in our definition of (36) above. The hollow symbols in Figure 10c indicate cases in which the rotational effects dominate the boundary layers according to the criteria in Figure 6, i.e.,  $\delta_E/\delta_T^{\rm NR} \leq 1$ . Comparing the number of hollow data points in Figure 6 versus those in Figure 10c, we see that there are fewer hollow points in Figure 10c (but the same number of solid points in the two figures). The reason for this is that the velocity data points that are below the LDV device resolution limit are not displayed in Figure 10c. Many of the velocity data points below this resolution limit exist in the heat transfer regime that is controlled by rotation. Figure 10c shows that estimated local Ro values within the Ekman boundary layers are greater than unity for the majority of the data shown, confirming that boundary layer dynamics are not strongly controlled by rotation. We note that plotting  $Ro_{\delta_E}$  versus  $\widetilde{Ra}$ , as in Figure 10c, results in well-collapsed data, with a best fit black line given as  $Ro_{\delta_E} = (0.084 \pm 0.005)\widetilde{Ra}^{(0.54 \pm 0.01)}$ .

According to asymptotic theory, diffusion-free heat transfer should behave as: [5,18, 50,58,98]

$$Nu \sim Ra^{3/2}E^2Pr^{-1/2}$$
. (37)

This balance is not achieved in our experiments and DNS because the boundary layer physics controls heat transfer behavior in the parameter space explored. Following the flux-based scaling arguments of Aurnou et al. [53], the diffusion-free CIA estimate of the global Reynolds number can be estimated using (37) as:

$$Re_{CIA} \sim \left(\frac{RaNu}{Pr^2}\right)^{\frac{2}{5}} E^{\frac{1}{5}} \sim \left(\frac{Ra}{Pr^2} \frac{Ra^{3/2}E^2}{Pr^{1/2}}\right)^{\frac{2}{5}} E^{\frac{1}{5}} \sim \frac{RaE}{Pr}.$$
 (38)

Taking  $Re_{CIA}$  as an estimate of the global Reynolds number, the scaling predictions for the local estimate parameters displayed in Figure 10 are:

$$Re_{\ell,\text{crit}} = Re_{\text{CIA}}E^{1/3} \simeq \frac{RaE}{Pr}E^{1/3} = \frac{\widetilde{Ra}}{Pr},$$
 (39a)

$$Ro_{\ell,\text{crit}} = (Re_{\text{CIA}}E)E^{-1/3} \simeq \frac{RaE}{Pr}E^{2/3} = \frac{\widetilde{Ra}E^{1/3}}{Pr},$$
 (39b)

$$Ro_{\delta_{\rm E}} = (Re_{\rm CIA}E)E^{-1/2} \simeq \frac{RaE}{Pr}E^{1/2} = \frac{RaE^{1/6}}{Pr}.$$
 (39c)

Thus, in the asymptotic diffusion-free limit, local estimates all scale with  $\widetilde{Ra}^1$ . This scaling is not seen in panels a–c of Figure 10. However, we do not expect these asymptotic scalings to hold in our data since the diffusion-free heat transfer regime given by (37) is not attained.

If we re-do the above analysis using the experimental heat transfer relation found herein,  $Nu \sim Ra^{1/3}$ , rather than (37), the CIA scaling of Re now becomes:

$$Re_{\text{CIA}} \sim \left(\frac{RaNu}{Pr^2}\right)^{\frac{2}{5}} E^{\frac{1}{5}} \sim \left(\frac{RaRa^{1/3}}{Pr^2}\right)^{\frac{2}{5}} E^{\frac{1}{5}} \sim \frac{\widetilde{Ra}^{8/15}}{Pr^{4/5}} E^{-23/45}.$$
 (40)

Fluids 2023, 8, 106 21 of 35

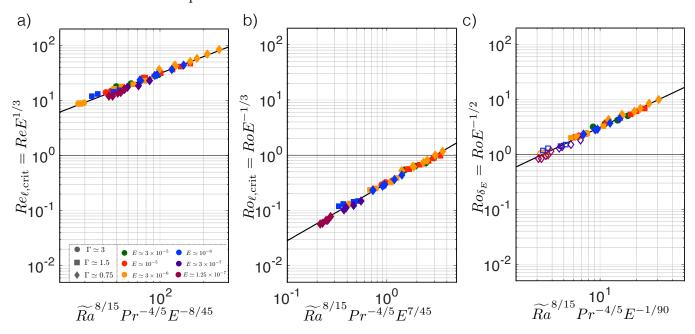
Thus, we predict that  $Re_{\ell,\text{crit}}$ ,  $Ro_{\ell,\text{crit}}$ , and  $Ro_{\delta_E}$  will experimentally scale with  $\widetilde{Ra}$  as:

$$Re_{\ell,\text{crit}} = Re_{\text{CIA}}E^{1/3} \simeq \widetilde{Ra}^{8/15}Pr^{-4/5}E^{-8/45},$$
 (41a)

$$Ro_{\ell,\text{crit}} = (Re_{\text{CIA}}E)E^{-1/3} \simeq \widetilde{Ra}^{8/15}Pr^{-4/5}E^{7/45},$$
 (41b)

$$Ro_{\delta_E} = (Re_{CIA}E)E^{-1/2} \simeq \widetilde{Ra}^{8/15}Pr^{-4/5}E^{-1/90}.$$
 (41c)

Figure 11a shows the local Reynolds number,  $Re_{\ell,\rm crit}$ , plotted against Equation (41a). Similarly, Figure 11b displays the local Rossby number,  $Ro_{\ell,\rm crit}$ , plotted versus Equation (41b), and Figure 11c shows the local boundary layer Rossby number,  $Ro_{\delta_E}$ , versus Equation (41c). All data in each panel collapse in y, indicating that each prediction of its respective local estimate is accurate. Further, the best fit black lines all yield a fit of  $y \simeq 0.3x^{1.0}$ . Thus, boundary layer controlled heat transfer and inviscid, quasi-geostrophic bulk dynamics provide accurate predictions of local parameters in both the bulk and boundary layers of experiments.



**Figure 11.** (a) Local Reynolds number,  $Re_{\ell,\text{crit}}$ , versus its scaling prediction given by (41a),  $\widetilde{Ra}^{8/15}Pr^{-4/5}E^{-8/45}$ . (b) Local Rossby number,  $Ro_{\ell,\text{crit}}$ , versus Equation (41b),  $\widetilde{Ra}^{8/15}Pr^{-4/5}E^{7/45}$ . (c) Local boundary layer Rossby number,  $Ro_{\delta_E}$ , versus Equation (41c),  $\widetilde{Ra}^{8/15}Pr^{-4/5}E^{-1/90}$ . The colors and symbol shapes are consistent with those in Figure 10. All best fit black lines yield a fit of  $y \simeq 0.3x^{1.0}$ , indicating that the predicted x-parameter in each panel does indeed collapse the local estimate plotted on the y-axis.

# 5. Discussion

The results of a laboratory-numerical survey of rotating Rayleigh–Bénard convection in which the heat and momentum transfer are simultaneously measured are presented. Thermal measurements are collected to characterize convective heat transfer in conjunction with measurements of fluid bulk convective-point velocities via laser Doppler velocimetry. For non-rotating experiments, data agree well with those of past studies, thus validating the convective measurements in this work. Specifically, non-rotating heat transfer data scale as  $Nu \sim Ra^{1/3}$  while the momentum transfer measurements exhibit a  $Re \sim Ra^{1/2}$  scaling.

Rotating convection DNS and laboratory heat transfer data show that transitions in heat transfer behavior are determined by boundary layer dynamics. Bulk dynamics are robustly governed by the inviscid CIA balance. Simultaneous measurements of heat and momentum transfer show that the CIA *Re* scaling holds from near convective onset to

Fluids 2023, 8, 106 22 of 35

nearly  $\sim$ 800 times critical (at low E, i.e.,  $E \simeq 3 \times 10^{-6}$ ). Importantly, this velocity behavior is not correlated with boundary layer dynamics, unlike heat transfer behavior. The onset and turbulent length scale estimates co-scale such that the VAC and CIA scalings of *Re* co-scale when the local Reynolds number in the fluid bulk is of order unity. We argue that this explains the results shown in Figures 8 and 9. An open question remains as to if, and under what conditions, the theoretical viscous and inertial length scales of RRBC underlying the VAC and CIA velocity scalings become clearly scale separable (e.g., [24,175–177]).

An analysis of the local flow estimates of our experiments also shows that bulk flow dynamics display properties of quasi-geostrophic turbulence. Local bulk Reynolds numbers are large while local bulk Rossby numbers remain small. Estimates of the local Reynolds numbers in the boundary layers are of order unity and demonstrate that heat transfer behavior is intimately tied to boundary layer dynamics. Thus, an analysis of local flow dynamics conveys the same findings as a global analysis. Further, the relationship between local estimates and convective supercriticality is properly accounted for by considering both boundary layer dominated heat transfer and inviscid bulk dynamics.

Quasi-geostrophic turbulence, relevant to dynamo flows, develops in the fluid bulk of experiments, even relatively near to onset. This finding suggests that we can access diffusion-free bulk dynamics in models that are far from planetary core conditions. Cheng and Aurnou [139] asserted that dynamo models cannot be extrapolated to core conditions because the heat transfer in these models is not diffusion-free. Meanwhile, many dynamo models exhibit diffusion-free CIA-balanced flow dynamics (e.g., [4,34,165,166]). Our findings elucidate that although the heat transfer in these models remains controlled by boundary layer dynamics, diffusion-free bulk flow dynamics are indeed captured.

Over the nearly 100 years in which laboratory convection experiments have been conducted, a long-standing assumption has been that heat transfer behavior is indicative of the full system dynamics (cf. [38,41,88,96]). This study clarifies that laboratory studies and comparable numerical work contain boundary layer controlled heat transfer that prevents bulk dynamics from becoming fully diffusion-free. The CIA *Re* scaling is dependent on heat transfer via *Nu*. Thus, bulk flows in present-day laboratory-numerical models contain inviscid, CIA-balanced flow dynamics that are limited by boundary layer controlled heat transfer [36].

Another interesting outstanding question remains as to the conditions under which present-day DNS and laboratory models might be able to achieve diffusion-free heat transfer. A recent laboratory study by Bouillaut et al. [163] claims to find the diffusion-free heat transfer scaling,  $Nu \sim Ra^{3/2}E^2Pr^{-1/2}$ , using radiatively driven convection. Overall, the work herein illuminates that boundary layer processes throttle the dynamics in current RRBC laboratory experiments. This boundary layer throttling prohibits the observation of the asymptotic RRBC flow state described by (38), which likely dominates in planetary cores [53].

**Author Contributions:** Conceptualization, E.K.H. and J.M.A.; Methodology, E.K.H., J.A.A., T.P., S.S. and J.M.A.; Software, S.S.; Formal analysis, E.K.H., K.J., S.S. and J.M.A.; Investigation, E.K.H. and J.M.A.; Resources, S.S. and J.M.A.; Data curation, E.K.H. and S.S.; Writing—original draft, E.K.H.; Writing—review & editing, J.S.C., J.A.A., T.P., S.S., K.J. and J.M.A.; Visualization, J.S.C.; Supervision, J.M.A.; Project administration, J.M.A.; Funding acquisition, S.S. and J.M.A. All authors have read and agreed to the published version of the manuscript.

**Funding:** This research was funded by the NSF geophysics program (grant numbers #1547269 and #1853196), and by the Gauss Centre for Supercomputing e.V. (www.gauss-centre.eu, accessed on 5 September 2019) in which computing time was provided through the John von Neumann Institute for Computing on the GCS Supercomputer JUWELS at the Jülich Supercomputing Centre.

**Data Availability Statement:** The data presented in this study are available in Appendix A herein.

**Acknowledgments:** We thank the three anonymous reviewers for their suggestions to improve this manuscript. The NoMag Fabrication Team is thanked: Jonathan Bridgeman, Austin Chadwick, Christina Daniel, Alexander Grannan, Chirag Jariwala, Mike Lavell, Drew Levy, Ryan Marakas,

Fluids **2023**, *8*, 106 23 of 35

Samuel May, Tong Mu, David Ortega, Steffen Tai, Steven Tomlinson, and Anthony Vong. Ron Aurnou, Ron Mears, and Joel Neal are also thanked for their design and engineering support and expertise. We also thank Anthony Willis and Ben Harding with Bigger Bang Communications Ltd. for their collaboration on the experimental visualization displayed in Figure 1, and Taylor Lonner for her help with Figure 1.

**Conflicts of Interest:** The authors declare no conflict of interest.

# Appendix A. Data Tables

**Table A1.** Dimensional experimental values: H is the fluid container height (container diameter is a fixed D=58.6 cm),  $\Omega$  is the rotation rate of the container, Q is the mean input heating power,  $\overline{T_{\rm fluid}}$  is the mean temperature of the fluid,  $\Delta T$  is the temperature difference across the fluid layer, and  $u_{\rm Z,\,rms}$  is the z-rms velocity in the fluid bulk. The dashed horizontal lines separate the different aspect ratio containers used.  $u_{\rm Z,\,rms}$  values that are gray indicate measurements below the determined resolution limit of 2.5 mm/s for the LDV device used in this study. A description of the errors in the parameters listed can be found in Appendix B.

H (cm)	Ω (rpm)	Q (W)	T <sub>fluid</sub> (C)	Δ <i>T</i> (C)	u <sub>z,rms</sub> (mm/s)
20.2	0	49.89	24.36	1.44	2.28
20.2	0	80.80	24.77	1.91	2.74
20.2	0	124.9	24.00	2.98	2.96
20.2	0	200.0	25.34	4.04	3.50
20.2	0	315.9	25.03	5.93	3.93
20.2	0	499.6	24.16	8.65	4.44
20.2	0	598.6	23.87	9.84	4.55
20.2	0	801.1	25.80	12.05	4.84
20.2	0	1102	24.92	16.03	5.57
40.1		49.75	24.33	1.56	3.42
40.1	0	102.8	24.44	2.66	4.34
40.1	0	135.0	24.07	3.36	4.77
40.1	0	199.0	24.43	4.68	5.50
40.1	0	297.9	24.86	5.88	6.39
40.1	0	403.6	25.16	7.29	7.23
40.1	0	599.4	25.24	10.21	8.65
40.1	0	797.1	25.73	12.33	9.60
40.1	0	902.5	26.02	13.67	9.64
40.1	0	1149	25.73	16.66	11.4
80.2		66.64		1.93	4.55
80.2	0	125.6	23.92	3.29	6.09
80.2	0	249.5	22.81	5.57	6.96
80.2	0	250.8	22.73	5.47	7.01
80.2	0	402.1	24.01	7.91	8.01
80.2	0	502.2	22.61	9.41	10.6
80.2	0	697.3	25.14	11.65	10.3
80.2	0	800.7	23.68	13.36	10.9
80.2	0	1107	25.42	16.77	12.4
20.2	3.6	49.98	25.07	1.02	1.91
20.2	3.6	75.22	24.98	1.52	2.01
20.2	3.6	99.14	24.83	1.97	2.04
20.2	3.6	163.4	24.83	3.12	2.28
20.2	3.6	272.9	24.83	4.82	2.54
20.2	3.6	447.9	25.44	7.20	2.88
20.2	3.6	731.7	26.17	10.66	3.33
20.2	3.6	1145	26.38	15.43	4.04

Fluids **2023**, 8, 106 24 of 35

Table A1. Cont.

H (cm)	Ω (rpm)	Q (W)	T <sub>fluid</sub> (C)	Δ <i>T</i> (C)	u <sub>z,rms</sub> (mm/s)
20.2	10.6	31.43	24.98	0.98	1.79
20.2	10.6	49.63	25.41	1.45	1.78
20.2	10.6	75.45	25.18	1.81	1.89
20.2	10.6	103.2	25.21	2.33	1.98
20.2	10.6	202.1	25.21	3.59	2.07
20.2	10.6	399.7	25.59	6.30	2.47
20.2	10.6	649.1	26.61	9.39	2.79
20.2	10.6	800.2	26.44	11.12	3.01
20.2	10.6	1099	26.78	14.21	3.27
40.1		50.10	24.40	<del>1</del> . <del>3</del> 2	2.02
40.1	2.7	67.28	24.45	1.88	2.18
40.1	2.7	102.6	24.46	2.55	2.29
40.1	2.7	201.2	24.81	4.16	2.68
40.1	2.7	200.4	24.41	4.49	2.72
40.1	2.7	391.6	25.10	6.79	3.24
40.1	2.7	599.1	25.09	9.85	3.75
40.1	2.7	797.3	25.64	11.96	4.27
40.1	2.7	1148	25.60	15.93	4.74
20.2	35.0	49.78	25.67	4.02	1.86
20.2	35.0	81.23	26.05	4.61	1.90
20.2	35.0	125.7	25.40	5.60	1.97
20.2	35.0	200.9	26.60	6.46	2.03
20.2	35.0	315.7	26.30	8.20	2.13
20.2	35.0	498.7	25.37	10.64	2.32
20.2	35.0	598.3	25.31	11.74	2.40
20.2	35.0	700.3	26.47	12.56	2.45
20.2	35.0	803.0	26.73	13.37	2.58
20.2	35.0	945.8	25.95	15.25	2.66
20.2	35.0	1104	25.64	16.93	2.76
40.1	9.0	51.45	25.40	1.99	2.00
40.1	9.0	67.30	24.68	2.12	1.98
40.1	9.0	103.0	24.53	2.69	2.04
40.1	9.0	134.9	25.20	3.00	2.17
40.1	9.0	201.3	24.10	4.16	2.30
40.1	9.0	296.7	25.18	5.79	2.53
40.1	9.0	398.9	25.62	6.83	2.70
40.1	9.0	402.7	24.87	7.24	2.68
40.1	9.0	599.3	25.13	10.03	3.06
40.1	9.0	801.1	27.69	11.62	3.42
40.1	9.0	1144	27.78	15.36	4.11
80.2	$\frac{5.0}{2.3}$	$\frac{1144}{66.45}$	$\frac{27.76}{24.35}$	$\frac{15.56}{1.83}$	<del>2.35</del>
80.2	2.3	126.3	23.80	3.10	2.89
80.2	2.3	251.0	22.82	5.52	3.46
80.2	2.3	250.8	22.72	5.42	3.58
80.2	2.3	402.1	23.93	7.78	4.05
80.2	2.3	402.1 499.7	23.93 24.69	7.78 8.99	4.54
80.2 80.2	2.3 2.3	799.5	23.69	13.42	5.59 6.45
00.4	2.3	1107	25.47	16.76	6.45

Fluids **2023**, 8, 106 25 of 35

Table A1. Cont.

H (cm)	Ω (rpm)	Q (W)	T <sub>fluid</sub> (C)	Δ <i>T</i> (C)	u <sub>z,rms</sub> (mm/s)
40.1	26.7	49.12	25.37	3.10	2.15
40.1	26.7	101.5	25.63	4.33	2.18
40.1	26.7	134.7	24.99	4.81	2.22
40.1	26.7	219.3	24.39	6.04	2.39
40.1	26.7	389.9	25.92	8.14	2.88
40.1	26.7	600.9	25.16	10.22	2.53
40.1	26.7	902.4	26.50	14.79	3.13
40.1	26.7	1148	26.38	17.97	3.30
80.2	7.0	66.44	24.58	2.13	2.10
80.2	7.0	125.5	24.06	3.47	2.40
80.2	7.0	250.1	22.88	5.72	2.75
80.2	7.0	250.3	22.93	5.72	2.67
80.2	7.0	401.9	24.04	7.97	3.22
80.2	7.0	503.5	22.74	9.64	3.40
80.2	7.0	799.8	23.85	13.67	4.22
80.2	7.0	1108	25.61	17.18	4.95
80.2	23.2	66.43	25.14	3.28	1.96
80.2	23.2	126.5	24.62	4.65	2.17
80.2	23.2	251.8	24.40	7.47	2.41
80.2	23.2	402.8	25.01	9.84	2.63
80.2	23.2	502.9	24.61	11.21	2.88
80.2	23.2	799.1	24.78	15.54	3.15
80.2	23.2	1109	26.61	19.07	3.71
80.2	55.7	66.52	26.48	6.21	2.28
80.2	55.7	95.12	26.55	7.61	2.30
80.2	55.7	126.3	26.51	8.59	2.30
80.2	55.7	168.6	26.34	9.77	2.32
80.2	55.7	252.1	26.25	11.24	2.33
80.2	55.7	400.4	26.92	13.18	2.60
80.2	55.7	498.5	27.48	14.22	2.62
80.2	55.7	698.2	27.67	16.42	2.93
80.2	55.7	798.4	27.64	17.78	3.00
80.2	55.7	1097	26.87	21.08	3.53

**Table A2.** Non-dimensional experimental values: Γ is the aspect ratio of the cylindrical container used, E is the Ekman number, Ra is the Rayleigh number, Pr is the Prandtl number, Nu is the Nusselt number, Re is the Reynolds number, Ro is the Rossby Number,  $Ro_c$  is the convective Rossby number given as  $Ro_c = \sqrt{RaE^2/Pr}$ , and Fr is the Froude number. The dashed horizontal lines separate the different aspect ratio containers used. Re and Ro values that are gray coincide with velocity measurements below the determined resolution limit of 2.5 mm/s for the LDV device.

Γ	E	Ra	Pr	Nu	Re	Ro	Roc	Fr
2.91	∞	$2.25 \times 10^{8}$	6.16	42.6	513.6	∞	∞	0
2.91	$\infty$	$3.07 \times 10^{8}$	6.10	48.0	$623.2 \pm 119.5$	$\infty$	$\infty$	0
2.91	$\infty$	$4.58 \times 10^{8}$	6.22	49.1	$661.1 \pm 108.7$	$\infty$	$\infty$	0
2.91	$\infty$	$6.70 \times 10^{8}$	6.01	59.2	$780.3 \pm 91.8$	$\infty$	$\infty$	0
2.91	$\infty$	$9.67 \times 10^{8}$	6.06	64.4	$899.7 \pm 83.9$	$\infty$	$\infty$	0
2.91	$\infty$	$1.34 \times 10^{9}$	6.20	70.7	$996.5 \pm 73.0$	$\infty$	$\infty$	0
2.91	$\infty$	$1.50 \times 10^{9}$	6.24	74.7	$1014 \pm 70.7$	$\infty$	$\infty$	0
2.91	$\infty$	$2.05 \times 10^{9}$	5.94	81.4	$1126 \pm 69.4$	$\infty$	$\infty$	0
2.91	$\infty$	$2.60 \times 10^{9}$	6.07	84.5	$1272\pm59.4$	$\infty$	$\infty$	0

Fluids **2023**, 8, 106 26 of 35

Table A2. Cont.

Γ	E	Ra	Pr	Nu	Re	Ro	Roc	Fr
1.46	$\infty$	$1.93 \times 10^{9}$	6.17	76.4	$1535 \pm 189.0$	$\infty$	$\infty$	0
1.46	$\infty$	$3.31 \times 10^{9}$	6.15	94.1	$1953 \pm 149.4$	$\infty$	$\infty$	0
1.46	$\infty$	$4.08 \times 10^{9}$	6.21	98.2	$2127 \pm 134.7$	$\infty$	$\infty$	0
1.46	$\infty$	$5.82 \times 10^{9}$	6.15	103	$2471 \pm 117.8$	$\infty$	$\infty$	0
1.46	$\infty$	$7.63 \times 10^{9}$	6.04	124	$2918 \pm 103.2$	$\infty$	$\infty$	0
1.46	$\infty$	$9.29 \times 10^{9}$	6.08	136	$3340 \pm 92.4$	$\infty$	$\infty$	0
1.46	$\infty$	$1.33 \times 10^{10}$	6.02	144	$3961 \pm 76.9$	$\infty$	$\infty$	0
1.46	$\infty$	$1.65 \times 10^{10}$	5.95	158	$4439 \pm 70.2$	$\infty$	$\infty$	0
1.46	$\infty$	$1.86\times10^{10}$	5.90	162	$4492 \pm 70.5$	$\infty$	$\infty$	0
1.46	$\infty$	$2.23 \times 10^{10}$	5.95	169	$5296 \pm 60.1$	$\infty$	$\infty$	0
0.73		$-1.91 \times 10^{10}$	6.15	169	$-4089 \pm 284.5$			
0.73	$\infty$	$3.16 \times 10^{10}$	6.24	188	$5403 \pm 209.9$	$\infty$	$\infty$	0
0.73	$\infty$	$4.98 \times 10^{10}$	6.42	221	$6014 \pm 178.9$	$\infty$	$\infty$	0
0.73	$\infty$	$4.87 \times 10^{10}$	6.44	226	$6047 \pm 177.4$	$\infty$	$\infty$	0
0.73	$\infty$	$7.62 \times 10^{10}$	6.22	250	$7120 \pm 160.0$	$\infty$	$\infty$	0
0.73	$\infty$	$8.32 \times 10^{10}$	6.46	263	$9128 \pm 117.5$	$\infty$	$\infty$	0
0.73	$\infty$	$1.20 \times 10^{11}$	6.04	294	$9350 \pm 127.5$	$\infty$	$\infty$	0
0.73	$\infty$	$1.26 \times 10^{11}$	6.28	295	$9574 \pm 116.7$	$\infty$	$\infty$	0
0.73	$\infty$	$1.76 \times 10^{11}$	6.00	323	$11,400 \pm 107.7$	$\infty$	$\infty$	0
2.91	$2.90 \times 10^{-5}$	$1.67 \times 10^{8}$	6.05	52.8	437.1	$1.27 \times 10^{-2}$	$1.53 \times 10^{-1}$	0.0042
2.91	$2.91 \times 10^{-5}$	$2.46 \times 10^{8}$	6.06	56.4	458.9	$1.34 \times 10^{-2}$	$1.85 \times 10^{-1}$	0.0042
2.91	$2.92 \times 10^{-5}$	$3.17 \times 10^{8}$	6.09	58.3	465.1	$1.36 \times 10^{-2}$	$2.11 \times 10^{-1}$	0.0042
2.91	$2.92 \times 10^{-5}$	$5.02 \times 10^{8}$	6.09	62.4	518.5	$1.52 \times 10^{-2}$	$2.65 \times 10^{-1}$	0.0042
2.91	$2.92 \times 10^{-5}$	$7.76 \times 10^{8}$	6.09	68.4	$579.0 \pm 129.3$	$1.69 \times 10^{-2}$	$3.30 \times 10^{-1}$	0.0042
2.91	$2.88 \times 10^{-5}$	$1.20 \times 10^{9}$	6.00	75.9	$665.8 \pm 115.7$	$1.92 \times 10^{-2}$	$4.07 \times 10^{-1}$	0.0042
2.91	$2.83 \times 10^{-5}$	$1.85 \times 10^{9}$	5.88	83.8	$781.1 \pm 101.5$	$2.21 \times 10^{-2}$	$5.02 \times 10^{-1}$	0.0042
2.91	$2.82 \times 10^{-5}$	$2.71 \times 10^9$	5.85	90.9	$954.0 \pm 84.3$	$2.69 \times 10^{-2}$	$6.07 \times 10^{-1}$	0.0042
2.91	$9.77 \times 10^{-6}$	$1.59 \times 10^{8}$	6.06	31.9	410.2	$4.00 \times 10^{-3}$	$5.01 \times 10^{-2}$	0.037
2.91	$9.67 \times 10^{-6}$	$2.42 \times 10^{8}$	6.00	41.8	411.7	$3.98 \times 10^{-3}$	$6.14 \times 10^{-2}$	0.037
2.91	$9.73 \times 10^{-6}$	$2.97 \times 10^{8}$	6.03	47.3	432.2	$4.22 \times 10^{-3}$	$6.83 \times 10^{-2}$	0.037
2.91	$9.72 \times 10^{-6}$	$3.84 \times 10^{8}$	6.03	54.1	455.6	$4.42 \times 10^{-3}$	$7.76 \times 10^{-2}$	0.037
2.91	$9.72 \times 10^{-6}$ $9.72 \times 10^{-6}$	$5.90 \times 10^{8}$	6.03	69.1	475.4	$4.62 \times 10^{-3}$	$9.61 \times 10^{-2}$	0.037
2.91	$9.64 \times 10^{-6}$	$1.04 \times 10^{9}$	5.97	79.3	572.5	$5.52 \times 10^{-3}$	$1.27 \times 10^{-1}$	0.037
2.91	$9.42 \times 10^{-6}$	$1.67 \times 10^9$	5.82	84.2	$660.1 \pm 122.2$	$6.23 \times 10^{-3}$	$1.59 \times 10^{-1}$	0.037
2.91	$9.38 \times 10^{-6}$	$1.96 \times 10^9$	5.84	89.5	$715.9 \pm 113.8$	$6.72 \times 10^{-3}$	$1.72 \times 10^{-1}$	0.037
2.91	$9.45 \times 10^{-6}$	$2.55 \times 10^9$	5.79	93.0	$772.0 \pm 104.0$	$7.30 \times 10^{-3}$	$1.98 \times 10^{-1}$	0.037
$-\frac{2.71}{1.46}$	$-\frac{9.13 \times 10}{9.84 \times 10^{-6}}$	$-\frac{2.63 \times 10}{1.63 \times 10^{9}}$	- 6.18 -	92.2	906.0	$-\frac{7.80 \times 10^{-3}}{8.89 \times 10^{-3}}$	$-\frac{1.50 \times 10}{1.60 \times 10^{-1}}$	- 0.0024
1.46	$9.79 \times 10^{-6}$	$2.34 \times 10^9$	6.15	86.7	980.3	$9.60 \times 10^{-3}$	$1.91 \times 10^{-1}$	0.0024
1.46	$9.79 \times 10^{-6}$	$3.71 \times 10^9$	6.15	98.4	1030	$1.01 \times 10^{-2}$	$2.41 \times 10^{-1}$	0.0024
1.46	$9.71 \times 10^{-6}$	$5.29 \times 10^9$	6.09	119	$1213 \pm 243.3$	$1.18 \times 10^{-2}$	$2.86 \times 10^{-1}$	0.0024
1.46	$9.78 \times 10^{-6}$	$5.58 \times 10^9$	6.16	109	$1215 \pm 245.5$ $1225 \pm 238.5$	$1.10 \times 10^{-2}$ $1.20 \times 10^{-2}$	$2.94 \times 10^{-1}$	0.0024
1.46	$9.64 \times 10^{-6}$	$8.77 \times 10^9$	6.05	142	$1478 \pm 202.9$	$1.43 \times 10^{-2}$	$3.67 \times 10^{-1}$	0.0024
1.46	$9.64 \times 10$ $9.65 \times 10^{-6}$	$1.27 \times 10^{10}$	6.05	142	$1478 \pm 202.9$ $1713 \pm 175.6$	$1.43 \times 10$ $1.65 \times 10^{-2}$	$4.42 \times 10^{-1}$	0.0024
1.46	$9.53 \times 10^{-6}$ $9.53 \times 10^{-6}$	$1.59 \times 10^{10}$	5.96	163	$1713 \pm 175.0$ $1974 \pm 156.2$	$1.88 \times 10^{-2}$	$4.42 \times 10$ $4.92 \times 10^{-1}$	0.0024
1.46	$9.54 \times 10^{-6}$	$2.12 \times 10^{10}$	5.96	176	$1974 \pm 136.2$ $2190 \pm 140.8$	$2.09 \times 10^{-2}$	$5.68 \times 10^{-1}$	0.0024
1.40	2.04 × 10 °	Z.1Z × 10	5.97	1/0	∠170 ± 1 <del>4</del> 0.0	2.09 X 10 2	3.00 × 10 1	0.0024

Fluids **2023**, *8*, 106 27 of 35

Table A2. Cont.

Γ	Е	Ra	Pr	Nu	Re	Ro	Roc	Fr
2.91	$2.91 \times 10^{-6}$	$6.79 \times 10^{8}$	5.96	13.5	432.7	$1.26 \times 10^{-3}$	$3.11 \times 10^{-2}$	0.40
2.91	$2.89 \times 10^{-6}$	$7.95 \times 10^{8}$	5.90	20.2	440.7	$1.29 \times 10^{-3}$	$3.35 \times 10^{-2}$	0.40
2.91	$2.93 \times 10^{-6}$	$9.32 \times 10^{8}$	6.00	26.5	454.9	$1.33 \times 10^{-3}$	$3.65 \times 10^{-2}$	0.40
2.91	$2.85 \times 10^{-6}$	$1.15 \times 10^{9}$	5.82	37.1	482.2	$1.37 \times 10^{-3}$	$4.01 \times 10^{-2}$	0.40
2.91	$2.87 \times 10^{-6}$	$1.44 \times 10^{9}$	5.86	46.6	501.3	$1.44 \times 10^{-3}$	$4.50 \times 10^{-2}$	0.40
2.91	$2.93 \times 10^{-6}$	$1.77 \times 10^{9}$	6.00	57.3	535.1	$1.57 \times 10^{-3}$	$5.04 \times 10^{-2}$	0.40
2.91	$2.94 \times 10^{-6}$	$1.94 \times 10^{9}$	6.01	62.4	551.9	$1.62 \times 10^{-3}$	$5.28 \times 10^{-2}$	0.40
2.91	$2.86 \times 10^{-6}$	$2.22 \times 10^{9}$	5.84	68.1	591.6	$1.66 \times 10^{-3}$	$5.58\times10^{-2}$	0.40
2.91	$2.84 \times 10^{-6}$	$2.40 \times 10^{9}$	5.80	73.4	$614.5 \pm 133.0$	$1.75\times10^{-3}$	$5.79 \times 10^{-2}$	0.40
2.91	$2.89 \times 10^{-6}$	$2.62 \times 10^{9}$	5.92	76.0	$621.3 \pm 126.5$	$1.80 \times 10^{-3}$	$6.09 \times 10^{-2}$	0.40
2.91	$2.91 \times 10^{-6}$	$2.86 \times 10^{9}$	5.96	80.1	$641.3 \pm 121.3$	$1.87 \times 10^{-3}$	$6.39 \times 10^{-2}$	0.40
$\bar{1}.\bar{4}6^{-}$	$-\bar{2}.\bar{87} \times 10^{-6}$	$-2.62 \times 10^{9}$	5.99	61.0	918.4	$2.64 \times 10^{-3}$	$-6.01 \times 10^{-2}$	0.027
1.46	$2.92 \times 10^{-6}$	$2.67 \times 10^{9}$	6.11	77.1	894.5	$2.61 \times 10^{-3}$	$6.11 \times 10^{-2}$	0.027
1.46	$2.93 \times 10^{-6}$	$3.37 \times 10^{9}$	6.14	93.7	917.9	$2.69 \times 10^{-3}$	$6.87 \times 10^{-2}$	0.027
1.46	$2.89 \times 10^{-6}$	$3.91 \times 10^{9}$	6.03	109	990.6	$2.86 \times 10^{-3}$	$7.35 \times 10^{-2}$	0.027
1.46	$2.96 \times 10^{-6}$	$5.07 \times 10^{9}$	6.03	119	1024	$3.04 \times 10^{-3}$	$8.58 \times 10^{-2}$	0.027
1.46	$2.89 \times 10^{-6}$	$7.52 \times 10^{9}$	6.21	125	$1158 \pm 260.6$	$3.34 \times 10^{-3}$	$1.01 \times 10^{-1}$	0.027
1.46	$2.86 \times 10^{-6}$	$9.10 \times 10^{9}$	6.00	143	$1249 \pm 246.8$	$3.57 \times 10^{-3}$	$1.11 \times 10^{-1}$	0.027
1.46	$2.91 \times 10^{-6}$	$9.24 \times 10^{9}$	6.08	136	$1218 \pm 244.3$	$3.54 \times 10^{-3}$	$1.13 \times 10^{-1}$	0.027
1.46	$2.89 \times 10^{-6}$	$1.30 \times 10^{10}$	6.04	146	$1396\pm214.8$	$4.03 \times 10^{-3}$	$1.34  imes 10^{-1}$	0.027
1.46	$2.73 \times 10^{-6}$	$1.73 \times 10^{10}$	5.66	168	$1654 \pm 203.8$	$4.51 \times 10^{-3}$	$1.51 \times 10^{-1}$	0.027
1.46	$2.72 \times 10^{-6}$	$2.30 \times 10^{10}$	5.65	181	$1993 \pm 170.2$	$5.43 \times 10^{-3}$	$1.74\times10^{-1}$	0.027
0.73	$-2.87 \times 10^{-6}$	$1.80 \times 10^{10}$	6.17	167	2100	$6.03 \times 10^{-3}$	$1.55 \times 10^{-1}$	0.0018
0.73	$2.90 \times 10^{-6}$	$2.95 \times 10^{10}$	6.26	194	$2552 \pm 440.0$	$7.42 \times 10^{-3}$	$1.99 \times 10^{-1}$	0.0018
0.73	$2.97 \times 10^{-6}$	$4.95 \times 10^{10}$	6.42	217	$2987 \pm 359.3$	$8.88 \times 10^{-3}$	$2.61\times10^{-1}$	0.0018
0.73	$2.98 \times 10^{-6}$	$4.82 \times 10^{10}$	6.44	224	$3082 \pm 346.3$	$9.19 \times 10^{-3}$	$2.58 \times 10^{-1}$	0.0018
0.73	$2.90 \times 10^{-6}$	$7.46 \times 10^{10}$	6.23	252	$3588 \pm 315.0$	$1.04 \times 10^{-2}$	$3.17 \times 10^{-1}$	0.0018
0.73	$2.84 \times 10^{-6}$	$9.02 \times 10^{10}$	6.11	270	$4097 \pm 286.3$	$1.17 \times 10^{-2}$	$3.46 \times 10^{-1}$	0.0018
0.73	$2.91 \times 10^{-6}$	$1.27 \times 10^{11}$	6.27	292	$4933 \pm 227.9$	$1.44 \times 10^{-2}$	$4.14 \times 10^{-1}$	0.0018
0.73	$2.79 \times 10^{-6}$	$1.76 \times 10^{11}$	5.99	323	$5962 \pm 206.5$	$1.66 \times 10^{-2}$	$4.79 \times 10^{-1}$	0.0018
1.46	$9.69 \times 10^{-7}$	$4.06 \times 10^{9}$	6.00	36.1	985.3	$9.55 \times 10^{-4}$	$2.52 \times 10^{-2}$	0.23
1.46	$9.64 \times 10^{-7}$	$5.77 \times 10^{9}$	5.96	55.3	1005	$9.69 \times 10^{-4}$	$3.00 \times 10^{-2}$	0.23
1.46	$9.78 \times 10^{-7}$	$6.18 \times 10^{9}$	6.06	66.3	1021	$9.99 \times 10^{-4}$	$3.12 \times 10^{-2}$	0.23
1.46	$9.91 \times 10^{-7}$	$7.49 \times 10^9$	6.16	86.9	1071	$1.06 \times 10^{-3}$	$3.46 \times 10^{-2}$	0.23
1.46	$9.57 \times 10^{-7}$	$1.10 \times 10^{10}$	5.92	115	$1217 \pm 211.3$	$1.17 \times 10^{-3}$	$4.13 \times 10^{-2}$	0.23
1.46	$9.74 \times 10^{-7}$	$1.33 \times 10^{10}$	6.04	144	$1317 \pm 296.3$	$1.28 \times 10^{-3}$	$4.57 \times 10^{-2}$	0.23
1.46	$9.45 \times 10^{-7}$	$2.07 \times 10^{10}$	5.83	148	$1473 \pm 216.5$	$1.39 \times 10^{-3}$	$5.63 \times 10^{-2}$	0.23
1.46	$9.48 \times 10^{-7}$	$2.49 \times 10^{10}$	_ 5.85	155	$_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$ $_{-}$	$1.47 \times 10^{-3}$	$6.18 \times 10^{-2}$	0.23
0.73	$-9.45 \times 10^{-7}$	$-2.12 \times 10^{10}$	6.13	151	1886	$1.79 \times 10^{-3}$	$5.55 \times 10^{-2}$	0.016
0.73	$9.57 \times 10^{-7}$	$3.35 \times 10^{10}$	6.21	172	2130	$2.04 \times 10^{-3}$	$7.03 \times 10^{-2}$	0.016
0.73	$9.83 \times 10^{-7}$	$5.14 \times 10^{10}$	6.41	210	$2376 \pm 452.4$	$2.34 \times 10^{-3}$	$8.80 \times 10^{-2}$	0.016
0.73	$9.82 \times 10^{-7}$	$5.16 \times 10^{10}$	6.40	213	$2314 \pm 467.4$	$2.27 \times 10^{-3}$	$8.82 \times 10^{-2}$	0.016
0.73	$9.57 \times 10^{-7}$	$7.69 \times 10^{10}$	6.21	247	$2859 \pm 397.1$	$2.74 \times 10^{-3}$	$1.07 \times 10^{-1}$	0.016
0.73	$9.87 \times 10^{-7}$	$8.59 \times 10^{10}$	6.43	258	$2934 \pm 365.5$	$2.90 \times 10^{-3}$	$1.14 \times 10^{-1}$	0.016
0.73	$9.61 \times 10^{-7}$	$1.31 \times 10^{11}$	6.25	287	$3733 \pm 301.9$	$3.59 \times 10^{-3}$	$1.39 \times 10^{-1}$	0.016
0.73	$9.23 \times 10^{-7}$	$1.82 \times 10^{11}$	5.97	313	$4563 \pm 268.3$	$4.21 \times 10^{-3}$	$1.61 \times 10^{-1}$	0.016
0.73	$2.82 \times 10^{-7}$	$3.38 \times 10^{10}$	6.04	93.8	1784	$5.04 \times 10^{-4}$	$2.11 \times 10^{-2}$	0.18
0.73	$2.85 \times 10^{-7}$	$4.65 \times 10^{10}$	6.12	129	1955	$5.58 \times 10^{-4}$	$2.49 \times 10^{-2}$	0.18
0.73	$2.87 \times 10^{-7}$	$7.26 \times 10^{10}$	6.16	164	2161	$6.19 \times 10^{-4}$	$3.11 \times 10^{-2}$	0.18
0.73	$2.83 \times 10^{-7}$	$1.01 \times 10^{11}$	6.06	197	$2387 \pm 496.9$	$6.76 \times 10^{-4}$	$3.65 \times 10^{-2}$	0.18
0.73	$2.85 \times 10^{-7}$	$1.12 \times 10^{11}$	6.12	218	$2596 \pm 450.7$	$7.40 \times 10^{-4}$	$3.86 \times 10^{-2}$	0.18
0.73	$2.84 \times 10^{-7}$	$1.57 \times 10^{10}$	6.10	249	$2848 \pm 413.3$	$8.10 \times 10^{-4}$	$4.56 \times 10^{-2}$	0.18
0.73	$2.72 \times 10^{-7}$	$2.13 \times 10^{11}$	5.82	280	$3497 \pm 365.9$	$9.53 \times 10^{-4}$	$5.21 \times 10^{-2}$	0.18

Fluids 2023, 8, 106 28 of 35

T 1	1	A2.		
เวเ	NIA	Δ')	10	u+

Γ	E	Ra	Pr	Nu	Re	Ro	Roc	Fr
0.73	$1.14\times10^{-7}$	$6.90 \times 10^{10}$	5.84	52.0	2145	$2.44 \times 10^{-4}$	$1.24\times10^{-2}$	1.02
0.73	$1.14 \times 10^{-7}$	$8.48 \times 10^{10}$	5.82	60.8	2168	$2.46 \times 10^{-4}$	$1.37 \times 10^{-2}$	1.02
0.73	$1.14  imes 10^{-7}$	$9.55 \times 10^{10}$	5.83	68.2	2160	$2.46 \times 10^{-4}$	$1.46  imes 10^{-2}$	1.02
0.73	$1.14 \times 10^{-7}$	$1.08 \times 10^{11}$	5.86	89.1	2170	$2.48 \times 10^{-4}$	$1.55 \times 10^{-2}$	1.02
0.73	$1.14 \times 10^{-7}$	$1.23 \times 10^{11}$	5.87	106	2182	$2.49 \times 10^{-4}$	$1.66 \times 10^{-2}$	1.02
0.73	$1.13 \times 10^{-7}$	$1.50 \times 10^{11}$	5.77	147	$2467 \pm 525.5$	$2.78 \times 10^{-4}$	$1.82 \times 10^{-2}$	1.02
0.73	$1.13 \times 10^{-7}$	$1.66 \times 10^{11}$	5.69	168	$2520 \pm 528.6$	$2.80 \times 10^{-4}$	$1.90 \times 10^{-2}$	1.02
0.73	$1.11 \times 10^{-7}$	$1.94 \times 10^{11}$	5.66	206	$2829 \pm 474.5$	$3.13 \times 10^{-4}$	$2.05 \times 10^{-2}$	1.02
0.73	$1.11 \times 10^{-7}$	$2.10 \times 10^{11}$	5.67	219	$2883 \pm 461.3$	$3.21 \times 10^{-4}$	$2.13 \times 10^{-2}$	1.02
0.73	$1.13\times10^{-7}$	$2.39\times10^{11}$	5.78	254	$3347 \pm 386.8$	$3.77\times10^{-4}$	$2.29 \times 10^{-2}$	1.02

**Table A3.** Non-dimensional values of DNS in this study: Γ is the aspect ratio, E is the Ekman number, Ra is the Rayleigh number, Pr is the Prandtl number, Nu is the Nusselt number, Re is the Reynolds number, Ro is the Rossby Number, Ro is the convective Rossby number given as  $Ro_c = \sqrt{RaE^2/Pr}$ , Fr is the Froude number, and  $N_x$ ,  $N_y$ , and  $N_z$  are the number of grid points in the x-, y-, and z-directions, respectively.

Γ	E	Ra	Pr	Nu	Re	Ro	$Ro_{ m ff}$	Fr	$N_x$	$N_y$	$N_z$
0.74	$2.9 \times 10^{-6}$	$2.20 \times 10^{8}$	5.92	1.61	14.57	$6.40 \times 10^{-5}$	$1.77 \times 10^{-2}$	0	288	288	240
0.74	$2.9 \times 10^{-6}$	$2.50 \times 10^8$	5.92	2.50	23.03	$1.07  imes 10^{-4}$	$1.88  imes 10^{-2}$	0	288	288	240
0.74	$2.9 \times 10^{-6}$	$3.00 \times 10^{8}$	5.92	5.26	40.91	$1.95 \times 10^{-4}$	$2.06 \times 10^{-2}$	0	288	288	240
0.74	$2.9 \times 10^{-6}$	$4.00 \times 10^{8}$	5.92	11.7	75.61	$3.54 \times 10^{-4}$	$2.38 \times 10^{-2}$	0	288	288	240
0.74	$2.9 \times 10^{-6}$	$5.00 \times 10^{8}$	5.92	18.5	107.6	$5.06 \times 10^{-4}$	$2.67 \times 10^{-2}$	0	288	288	240
0.74	$2.9 \times 10^{-6}$	$6.80 \times 10^{8}$	5.92	30.8	160.4	$7.71 \times 10^{-4}$	$3.11 \times 10^{-2}$	0	288	288	240
0.74	$2.9 \times 10^{-6}$	$8.00 \times 10^8$	5.92	38.3	193.8	$9.35 \times 10^{-4}$	$3.37 \times 10^{-2}$	0	288	288	240
0.74	$2.9 \times 10^{-6}$	$1.00 \times 10^{9}$	5.92	48.8	244.3	$1.17 \times 10^{-4}$	$3.77 \times 10^{-2}$	0	384	384	240
0.74	$2.9 \times 10^{-6}$	$1.20 \times 10^{9}$	5.92	56.6	288.5	$1.37 \times 10^{-3}$	$4.13 \times 10^{-2}$	0	384	384	240
0.74	$2.9 \times 10^{-6}$	$1.95\times10^{9}$	5.92	73.7	414.5	$1.92\times10^{-3}$	$5.26\times10^{-2}$	0	384	384	384

# Appendix B. Experimental Methods: Additional Details

The rotation of the fluid chamber and the supporting aluminum structural frame about the vertical axis is controlled by a Danaher gearhead motor (DT90-010) with a Yaskawa speed reducer (SGMPH-04AAE41D). The measurement of the rotation rate for a given experiment is carefully computed by hand. A Sorenson DCS150-8E programmable DC power supply is connected to the heating pad described in Section 3.1 to provide constant heating power to the bottom fluid boundary. A Thermo NESLAB HX-150 aircooled recirculating chiller with a CP-55 pump is used to extract the heat provided by the heat pad to the fluid layer in order to maintain convection with fixed temperatures at the boundaries. The input heating power is accurately calculated by measuring the current and voltage provided by the power supply to the heat pad. To read the current, a shunt resistor with a fixed resistance is connected directly in line with the power supply and heat pad. The voltage across the shunt is measured and Ohm's law is used to calculate the current passing through the shunt resistor and thus the heat pad. The top fluid temperature that is maintained by the recirculating chiller is read using the thermistors described in Section 3.1. If interested, further details regarding the operation of the NoMag device and associated measurement techniques can be found in Hawkins [148].

Temperature data and voltage measurements for the input heating power calculation described are collected using a National Instruments (NI) PXI 1042Q chassis with an NI PXI 8105 embedded controller and an NI PXI 6221 communication module. The communication module is connected to a digitizing module chassis, an NI SCXI 1000. This chassis houses two modules (two NI SCXI 1102C modules) that contain 32 analog voltage input channels (in which the temperature sensors and voltage signals are connected via NI SCXI 1303

Fluids 2023, 8, 106 29 of 35

isothermal terminal blocks). A third module (NI SCXI 1581 32-channel analog output current excitation module) exists in the SCXI 1000 chassis to provide a small excitation current to the thermistors so that they can be properly read. The modules in the SCXI 1000 convert the analog voltage signals into digital readings. The digital readings are then transmitted via an SCXI data cable to the main PXI chassis/embedded controller, where all data are saved at a rate of 10 Hz. Separately, a desktop Dell PC equipped with Windows 10 is connected to the MSE LDV described in Section 3.2 to collect and save velocity data at rates  $\gtrsim$ 5 Hz. Custom data acquisition software from MSE is used to collect the velocity data from the LDV. Temperature and velocity data are matched via recorded timestamps in both datasets.

In terms of propagating errors for experimental data we now describe the methods we use to determine the errors in all the dimensional parameters. We accurately measure the height, H, (and radius, R) of the cylindrical containers used via a combination of measuring tape and calipers, and we determine the error in these measurements to be  $\pm 1$  mm. The rotation rate of the container for each experiment is calculated by hand to precision within  $\pm 0.1$  rpm. The error in the temperature of each thermistor used is  $\pm 50$  mK, as mentioned in Section 3.1. The input heating power (in Watts) is calculated using the measurements of the current and voltage across the heat pad provided by the DC power supply, as described (i.e.,  $Q_{\rm input} = I_{\rm shunt} V_{\rm heat \, pad}$ ). The heating power, Q, recorded for each experiment in Table A1 of Appendix A is calculated as  $Q = Q_{\rm input} - Q_{\rm lost}$ , where  $Q_{\rm lost}$  is the amount of heating power lost out of the sidewalls of the experiment. As discussed in Section 3.1, the heat loss through the top and bottom boundaries of the cylindrical chamber is negligible as our boundaries are well insulated and isothermal (i.e., the Biot numbers are small with  $Bi \lesssim 0.05$  in all cases).

We calculate the heat lost out of the cylindrical sidewall using a set of temperature sensors on the outer acrylic wall that are placed azimuthally around the mid-plane as: [178]

$$Q_{\text{lost}} = \frac{|\overline{T_{\text{fluid}}} - \overline{T_{\text{SW}}}| 2\pi k_{\text{acrylic}} H}{\log(R_o/R_i)},$$
(A1)

where  $\overline{T_{\text{fluid}}}$  is the mean fluid temperature (i.e., the fluid temperature at the mid-plane) in K (shown in degrees Celsius in Table A1 of Appendix A),  $\overline{T_{\text{SW}}}$  is the mean temperature of the sensors on the outer cylindrical sidewall at the mid-plane in K,  $R_0$  is the outermost cylinder radius (accounting for the wall thickness of the container) in m,  $R_i$  is the fluid radius (i.e., the inner radius of the container) in m,  $k_{\text{acrylic}}$  is the thermal conductivity of the acrylic sidewall in W/m·K, and H is the container height in m. These sidewall sensors are within several layers of thermal insulation, as described in Section 3.1.

Thus, the error in Q is determined as: [179]

$$\delta Q = \delta Q_{\text{input}} + \delta Q_{\text{lost}} = \left(0.05 + Q_{\text{lost}} \sqrt{\left(\frac{2(0.05K)}{|\overline{T}_{\text{fluid}} - \overline{T}_{\text{SW}}|}\right)^2}\right) \text{ Watts,}$$
 (A2)

where the errors in the other components of (A1) are determined to be negligible in comparison to the first term included under the square root in (A2) and thus are not included in (A2). The error in Q is found to be most significant for cases with the lowest input heating power in each tank, which makes sense because these cases are most susceptible to being influenced by variations in the surrounding room temperature (despite our best attempts to insulate the experiment and isolate it from the effects of the surrounding room). Lastly, the error in the rms z-velocity values provided in Table A1 of Appendix A is determined to be  $\pm 0.002$  m/s (see Section 3.2 [148] for more details).

We now provide the details of our error propagation for the two output parameters of significance in our study, the Nusselt and Reynolds numbers. The errors in all other

Fluids 2023, 8, 106 30 of 35

non-dimensional parameters provided in Table A2 of Appendix A are calculated using the errors in the dimensional values described. We calculate the error in Nu as: [179]

$$\delta Nu = Nu \sqrt{\left(\frac{\delta Q}{Q}\right)^2 + \left(\frac{\delta H}{H}\right)^2 + \left(\frac{\delta \Delta T}{\Delta T}\right)^2 + 2\left(\frac{\delta R}{R}\right)^2},$$
 (A3)

where we do not include the error in the thermal conductivity of the fluid, k, of (4) because it is negligible compared to the other terms listed under the square root in (A3). The relative errors in the Nusselt numbers of all experiments do not exceed more than  $\pm 4\%$ . When we plot vertical error bars in Figure 5a, they are not visible. Thus, we deduce that the errors in the laboratory Nu values always remain small and we, therefore, do not list them in Table A2 of Appendix A.

The error in the experimental *Re* values is given as: [179]

$$\delta Re = Re\sqrt{\left(\frac{(0.002)^2}{(u_{z, \text{rms}})^2}\right)^2 + \left(\frac{\delta H}{H}\right)^2},\tag{A4}$$

where once again we do not include the negligible error in the fluid viscosity,  $\nu$ , in (6). The error in  $u_{z,\,rms}$  is squared relative to the square of the value itself because this is the variance of our root-mean-square value of the velocity. The error for each experimental Re value (that is not considered low resolution, as indicated by the gray color) is presented in Table A2 of Appendix A. The relative errors in Re do not exceed  $\pm 22\%$ . Further details about the methods we used to analyze the experimental errors are available upon request from the corresponding author.

#### References

- 1. Elsassar, W. On the Origin of the Earth's Magnetic Field. Phys. Rev. 1939, 55, 489–498. [CrossRef]
- 2. Roberts, P.H.; King, E.M. On the genesis of the Earth's magnetism. Rep. Prog. Phys. 2013, 76, 096801. [CrossRef] [PubMed]
- 3. Cheng, J.S.; Aurnou, J.M.; Julien, K.; Kunnen, R. A heuristic framework for next-generation models of geostrophic convective turbulence. *Geophys. Astrophys. Fluid Dyn.* **2018**, *112*, 277–300. [CrossRef]
- 4. Schwaiger, T.; Gastine, T.; Aubert, J. Relating force balances and flow length scales in geodynamo simulations. *Geophys. J. Int.* **2021**, 224, 1890–1904. [CrossRef]
- 5. Maffei, S.; Krouss, M.; Julien, K.; Calkins, M. On the inverse cascade and flow speed scaling behaviour in rapidly rotating Rayleigh-Bénard convection. *J. Fluid Mech.* **2021**, *913*, A18. [CrossRef]
- 6. Horn, S.; Schmid, P.J. Prograde, Retrograde, and Oscillatory modes in rotating Rayleigh-Bénard Convection. *J. Fluid Mech.* **2017**, 831, 182–211. [CrossRef]
- 7. Aurnou, J.; Bertin, V.; Grannan, A.; Horn, S.; Vogt, T. Rotating thermal convection in liquid gallium: Multi-modal flow, absent steady columns. *J. Fluid Mech.* **2018**, *846*, 846–876. [CrossRef]
- 8. Aujogue, K.; Pothérat, A.; Sreenivasan, B.; Debray, F. Experimental study of the convection in arotating tangent cylinder. *J. Fluid Mech.* **2018**, *843*, 355–381. [CrossRef]
- 9. de Wit, X.; Guzman, A.A.; Madonia, M.; Cheng, J.S.; Clercx, H.; Kunnen, R. Turbulent rotating convection confined in a slender cylinder: The sidewall circulation. *Phys. Fluid Dyn.* **2020**, *5*, 023502. [CrossRef]
- 10. Lu, H.; Ding, G.; Shi, J.; Xia, K.; Zhong, J. Heat-transport scaling and transition in geostrophic rotating convection with varying aspect ratio. *Phys. Rev. Fluids* **2021**, *6*, L071501. [CrossRef]
- 11. Ecke, R.E.; Zhang, X.; Shishkina, O. Connecting wall modes and boundary zonal flows in rotating Rayleigh-Bénard convection. *Phys. Rev. Fluids* **2022**, *7*, L011501. [CrossRef]
- 12. Wedi, M.; Moturi, V.; Funfschilling, D.; Weiss, S. Experimental evidence for the boundary zonal flow in rotating Rayleigh-Bénard convection. *J. Fluid Mech.* **2022**, 939, A14. [CrossRef]
- 13. Gastine, T.; Aurnou, J.M. Latitudinal regionalization of rotating spherical shell convection. J. Fluid Mech. 2022, 954, R1. [CrossRef]
- 14. Wang, G.; Santelli, L.; Lohse, D.; Verzicco, R.; Stevens, R. Diffusion-free scaling in rotating spherical Rayleigh Bénard Convection. *Geophys. Res. Lett.* **2021**, 48, e2021GL095017. [CrossRef]
- 15. Sheyko, A.; Finlay, C.; Favre, J.; Jackson, A. Scale separated low viscosity dynamos and dissipation within the Earth's core. *Sci. Rep.* **2018**, *8*, 12566. [CrossRef] [PubMed]
- 16. Ecke, R.E.; Shishkina, O. Turbulent Rotating Rayleigh-Bénard Convection. Ann. Rev. Fluid Mech. 2023, 55, 603–638. [CrossRef]
- 17. Kunnen, R. The geostrophic regime of rapidly rotating turbulent convection. J. Turbul. 2021, 22, 267–296. [CrossRef]
- 18. Sprague, M.; Julien, K.; Knobloch, E.; Werne, J. Numerical simulation of an asymptotically reduced system for rotationally constrained convection. *J. Fluid Mech.* **2006**, *551*, 141–174. [CrossRef]

Fluids 2023, 8, 106 31 of 35

19. Kunnen, R.; Geurts, B.; Clercx, H. Experimental and numerical investigation of turbulent convection in a rotating cylinder. *J. Fluid Mech.* **2010**, *642*, 445–476. [CrossRef]

- 20. King, E.; Aurnou, J.M. Thermal evidence for Taylor columns in turbulent rotating Rayleigh-Bénard convection. *Phys. Rev. E* **2012**, 85, 016313. [CrossRef]
- 21. Julien, K.; Rubio, A.M.; Grooms, I.; Knobloch, E. Statistical and physical balances in low-Rossby-number Rayleigh-Bénard convection. *Geophys. Astrophys. Fluid Dyn.* **2012**, *106*, 392–428. [CrossRef]
- 22. Gastine, T.; Wicht, J.; Aubert, J. Scaling regimes in spherical shell rotating convection. J. Fluid Mech. 2016, 808, 690–732. [CrossRef]
- 23. Chong, K.; Yang, Y.; Haung, S.; Zhong, J.; Stevens, R.; Verzicco, R.; Lohse, D.; Xia, K. Confined Rayleigh–B'enard, rotating Rayleigh–B'enard, and double diffusive convection: A unifying view on turbulent transport enhancement through coherent structure manipulation. *Phys. Rev. Lett.* **2017**, *119*, 064501. [CrossRef] [PubMed]
- 24. Rajaei, H.; Alards, K.; Kunnen, R.; Clercx, H. Velocity and acceleration statistics in rapidly rotating Rayleigh–Bénard convection. *J. Fluid Mech.* **2018**, *857*, 374–397. [CrossRef]
- 25. Vogt, T.; Horn, S.; Aurnou, J. Oscillatory thermal-inertial flows in liquid metal rotating convection. *J. Fluid Mech.* **2021**, *911*, A5. [CrossRef]
- 26. Long, R.S. Regimes and Scaling Laws for Convection with and without Rotation. Ph.D. Thesis, The University of Leeds, Leeds, UK. 2021.
- 27. Soderlund, K.; King, E.; Aurnou, J.M. The influence of magnetic fields in planetary dynamo models. *Earth Planet. Sci. Lett.* **2012**, 333, 9–20. [CrossRef]
- 28. Calkins, M.; Julien, K.; Tobias, S.; Aurnou, J. A multi-scale dynamo model driven by quasigeostrophic convection. *J. Fluid Mech.* **2015**, *780*, 143–166. [CrossRef]
- 29. Yadav, R.; Gastine, T.; Christensen, U.R.; Wolk, S.; Poppenhaeger, K. Approaching a realistic force balance in geodynamo simulations. *Proc. Natl. Acad. Sci. USA* **2016**, *113*, 12065–12070. [CrossRef]
- 30. Aurnou, J.; King, E. The cross-over to magnetostrophic convection in planetary dynamo systems. *Proc. R. Soc. A* **2017**, 473, 20160731. [CrossRef]
- 31. Calkins, M. Quasi-geostrophic dynamo theory. Phys. Earth Planet. Inter. 2018, 276, 182–189. [CrossRef]
- 32. Aubert, J. Approaching Earth's core conditions in high-resolution geodynamo simulations. *Geophys. J. Int.* **2019**, 219, 137–151. [CrossRef]
- 33. Schwaiger, T.; Gastine, T.; Aubert, J. Force balance in numerical geodynamo simulations: A systematic study. *Geophys. J. Int.* **2019**, 219, 101–114. [CrossRef]
- 34. Calkins, M.; Orvedahl, R.J.; Featherstone, N.A. Large-scale balances and asymptotic behaviour in spherical dynamos. *Geophys. J. Int.* **2021**, 227, 1228–1245. [CrossRef]
- 35. Yan, M.; Tobias, S.; Calkins, M. Scaling behaviour of small-scale dynamos driven by Rayleigh-Bénard convection. *J. Fluid Mech.* **2021**, *915*, A15. [CrossRef]
- 36. Orvedahl, R.; Featherstone, N.A.; Calkins, M. Large-scale magnetic field saturation and the Elsasser number in rotating spherical dynamo models. *Mon. Not. R. Astron. Soc. Lett.* **2021**, *507*, *67*–71. [CrossRef]
- 37. Kolhey, P.; Stellmach, S.; Heyner, D. Influence of boundary conditions on rapidly rotating convection and its dynamo action in a plane fluid layer. *Phys. Rev. Fluids* **2022**, *7*, 043502. [CrossRef]
- 38. Rossby, H.T. A study of Bénard convection with and without rotation. J. Fluid Mech. 1969, 36, 309–335. [CrossRef]
- 39. Kerr, R.M.; Herring, J.R. Prandtl number dependence of Nusselt number in direct numerical simulations. *J. Fluid Mech.* **2000**, 419, 325–344. [CrossRef]
- 40. Funfschilling, D.; Brown, E.; Nikolaenko, A.; Ahlers, G. Heat transport by turbulent Rayleigh-Bénard convection in cylindrical samples. *J. Fluid Mech.* **2005**, *536*, 145–154. [CrossRef]
- 41. Weiss, S.; Stevens, R.; Zhong, J.; Clercx, H.; Lohse, D.; Ahlers, G. Finite-Size Effects Lead to Supercritical Bifurcations in Turbulent Rotating Rayleigh-Bénard Convection. *Phys. Rev. Lett.* **2010**, *105*, 224501. [CrossRef] [PubMed]
- 42. Choblet, G. On the scaling of heat transfer for mixed heating convection in a spherical shell. *Phys. Earth Planet. Inter.* **2012**, 206, 31–42. [CrossRef]
- 43. Horn, S.; Shishkina, O. Rotating non-Oberbeck-Boussinesq Rayleigh-Bénard convection in water. *Phys. Fluids* **2014**, 26, 055111. [CrossRef]
- 44. Ecke, R.E. Scaling of heat transport near onset in rapidly rotating convection. *Phys. Lett. A* 2015, 379, 2221–2223. [CrossRef]
- 45. Cheng, J.S.; Madonia, M.; Guzman, A.A.; Kunnen, R. Laboratory exploration of heat transfer regimes in rapidly rotating turbulent convection. *Phys. Rev. Fluids* **2020**, *5*, 113501. [CrossRef]
- 46. King, E.; Stellmach, S.; Noir, J.; Hansen, U.; Aurnou, J.M. Boundary layer control of rotating convection systems. *Nature* **2009**, 457, 301–304. [CrossRef]
- 47. Stellmach, S.; Lischper, M.; Julien, K.; Vasil, G.; Cheng, J.S.; Ribeiro, A.; King, E.M.; Aurnou, J.M. Approaching the Asymptotic Regime of Rapidly Rotating Convection: Boundary Layers versus Interior Dynamics. *Phys. Rev. Lett.* **2014**, *113*, 254501. [CrossRef]
- 48. Stevenson, D. Turbulent thermal convection in the presence of rotation and a magnetic field: A hueristic theory. *Geophys. Astrophys. Fluid Dyn.* **1979**, *12*, 139–169. [CrossRef]
- 49. Julien, K.; Knobloch, E. A new class of equations for rotationally constrained flows. *Theor. Comput. Fluid Dyn.* **1998**, *11*, 251–261. [CrossRef]

Fluids 2023, 8, 106 32 of 35

50. Julien, K.; Knobloch, E.; Rubio, A.M.; Vasil, G.M. Heat transport in Low-Rossby-Number Rayleigh-Bénard convection. *Phys. Rev. Lett.* **2012**, *109*, 254503. [CrossRef]

- 51. Barker, A.; Dempsey, A.; Lithwick, Y. Theory and simulations of rotating convection. Astrophys. J. 2014, 791, 13. [CrossRef]
- 52. Plumley, M.; Julien, K.; Marti, P.; Stellmach, S. The effects of Ekman pumping on quasi-geostrophic Rayleigh-Bénard convection. *J. Fluid Mech.* **2016**, *803*, 51–71. [CrossRef]
- 53. Aurnou, J.M.; Horn, S.; Julien, J. Connections between nonrotating, slowly rotating, and rapidly rotating turbulent convection transport scalings. *Phys. Rev. Reas.* **2020**, *2*, 043115. [CrossRef]
- 54. Ingersoll, A.; Pollard, D. Motion in the interiors and atmospheres of Jupiter and Saturn: Scale analysis, anelastic equations, barotropic stability criterion. *Icarus* **1982**, *52*, *62*–80. [CrossRef]
- 55. Aubert, J.; Brito, D.; Nataf, H.; Cardin, P.; Masson, J. A systematic experimental study of rapidly rotating spherical convection in water and liquid gallium. *Phys. Earth Planet. Inter.* **2001**, *128*, 51–74. [CrossRef]
- 56. Kraichnan, R.H. Turbulent thermal convection at arbitrary Prandtl number. *Phys. Fluids* **1962**, *5*, 1374–1389. [CrossRef]
- 57. Spiegel, E.A. Convection in stars: I. Basic Boussinesq convection. Ann. Rev. Astro. Astrophys. 1971, 9, 323–352. [CrossRef]
- 58. Plumley, M.; Julien, K. Scaling laws in Rayleigh-Bénard convection. Earth Space Sci. 2019, 6, 1580–1592. [CrossRef]
- 59. Long, R.S.; Mound, J.E.; Davies, C.J.; Tobias, S.M. Scaling behaviour in spherical shell rotating convection with fixed-flux thermal boundary conditions. *J. Fluid Mech.* **2020**, *889*, A7. [CrossRef]
- 60. Shraiman, B.; Siggia, E. Heat transport in high-Rayleigh-number convection. Phys. Rev. A 1990, 42, 3650–3653. [CrossRef]
- 61. Chilla, F.; Ciliberto, S.; Innocenti, C.; Pampaloni, E. Boundary layer and scaling properties in turbulent thermal convection. *Il Nuovo Cimento D* **1993**, *15*, 1229–1249. [CrossRef]
- 62. Grossmann, S.; Lohse, D. Scaling in thermal convection: A unifying theory. J. Fluid Mech. 2000, 407, 27–56. [CrossRef]
- 63. Grossman, S.; Lohse, D. Prandtl and Rayleigh number dependence of the Reynolds number in turbulent thermal convection. *Phys. Rev. E* **2002**, *66*, 016305. [CrossRef] [PubMed]
- 64. King, E.; Stellmach, S.; Aurnou, J.M. Heat transfer by rapidly rotating Rayleigh-Bénard convection. *J. Fluid Mech.* **2012**, 691, 568–582. [CrossRef]
- 65. Cheng, J.S.; Stellmach, S.; Ribeiro, A.; Grannan, A.; King, E.; Aurnou, J.M. Laboratory-numerical models of rapidly rotating convection in planetary cores. *Geophys. J. Int.* **2015**, 201, 1–17. [CrossRef]
- 66. Yang, Y.; Zhu, X.; Wang, B.; Liu, Y.; Zhou, Q. Experimental investigation of turbulent Rayleigh-Bénard convection of water in a cylindrical cell: The Prandtl number effects for Pr > 1. Phys. Fluids **2020**, 32, 015101. [CrossRef]
- 67. Priestly, C. Buoyant motion in a turbulent environment. Aust. J. Phys 1953, 6, 279–290. [CrossRef]
- 68. Malkus, W.V.R. The heat transport and spectrum of thermal turbulence. *Proc. R. Soc. A* 1954, 225, 196–212.
- 69. Brown, E.; Nikolaenko, A.; Funfschilling, D.; Ahlers, G. Heat transport in turbulent Rayleigh-Bénard convection: Effect of finite top-and bottom-plate conductivities. *Phys. Fluids* **2005**, *17*, 075108. [CrossRef]
- 70. Sun, C.; Ren, L.; Song, H.; Xia, K. Heat transport by turbulent Rayleigh-Bénard convection in 1 m diameter cylindrical cells of widely varying aspect ratio. *J. Fluid Mech.* **2005**, 542, 165–174. [CrossRef]
- 71. Ahlers, G.; Grossmann, S.; Lohse, D. Heat transfer and large scale dynamics in turbulent Rayleigh-Bénard convection. *Rev. Mod. Phys.* **2009**, *81*, 503. [CrossRef]
- 72. Chilla, F.; Schumacher, J. New perspectives in turbulent Rayleigh-Bénard convection. Eur. Phys. J. E 2012, 35, 1–25. [CrossRef]
- 73. Doering, C.; Toppaladoddi, S.; Wettlaufer, J. Absence of evidence for the ultimate regime in two-dimensional Rayleigh-bénard convection. *Phys. Rev. Lett.* **2019**, 123, 259401. [CrossRef]
- 74. Iyer, K.; Scheel, J.; Schumacher, J.; Sreenivasan, K. Classical 1/3 scaling of convection holds up to  $Ra = 10^{15}$ . *Proc. Natl. Acad. Sci. USA* **2020**, 117, 7594–7598. [CrossRef] [PubMed]
- 75. Zhang, K.; Liao, X. The onset of convection in rotating circular cylinders with experimental boundary conditions. *J. Fluid Mech.* **2009**, *6*22, 63–73. [CrossRef]
- 76. Chandrasekhar, S. Hydrodynamic and Hydromagnetic Stability; Dover Publications: Mineola, NY, USA, 1961.
- 77. Qiu, X.L.; Tong, P. Onset of coherent oscillations in turbulent Rayleigh Bénard convection. *Phys. Rev. Lett.* **2001**, *87*, 094501. [CrossRef]
- 78. Qiu, X.L.; Tong, P. Large-scale velocity structures in turbulent thermal convection. *Phys. Rev. E* **2001**, *64*, 036304. [CrossRef] [PubMed]
- 79. Xi, H.; Zhou, Q.; Xia, K. Azimuthal motion of the mean wind in turbulent thermal convection. *Phys. Rev. E* **2006**, 73, 056312. [CrossRef] [PubMed]
- 80. Brown, E.; Funfschilling, D.; Ahlers, G. Anomalous Reynolds-number scaling in turbulent Rayleigh-Bénard convection. *J. Stat. Mech.* **2007**, *10*, P10005. [CrossRef]
- 81. Horn, S.; Aurnou, J. Regimes of Coriolis-Centrifugal Convection. Phys. Rev. Lett. 2018, 120, 204502. [CrossRef]
- 82. Zhong, F.; Ecke, R.E.; Steinberg, V. Rotating Rayleigh Bénard convection: Asymmetric modes and vortex states. *J. Fluid Mech.* **1993**, 249, 135–159. [CrossRef]
- 83. Herrmann, J.; Busse, F. Asymptotic theory of wall-attached convection in a rotating fluid layer. *J. Fluid Mech.* **1993**, 255, 183–194. [CrossRef]
- 84. Goldstein, H.; Knobloch, E.; Mercader, I.; Net, M. Convection in a rotating cylinder. Part 1: Linear theory for moderate Prandtl numbers. *J. Fluid Mech.* **1993**, 248, 583–604. [CrossRef]

Fluids 2023, 8, 106 33 of 35

85. Boubnov, B.; Golitsyn, G. *Convection in Rotating Fluids*; Fluid Mechanics and Its Applications; Springer Science and Business Media: Berlin/Heidelberg, Germany, 1995; Volume 29.

- 86. Favier, B.; Knobloch, E. Robust wall states in rapidly rotating Rayleigh Bénard convection. J. Fluid Mech. 2020, 895, R1. [CrossRef]
- 87. Zhang, X.; Ecke, R.; Shishkina, O. Boundary zonal flows in rapidly rotating turbulent thermal convection. *J. Fluid Mech.* **2021**, 915, A62. [CrossRef]
- 88. Grannan, A.; Cheng, J.; Aggarwal, A.; Hawkins, E.; Xu, Y.; Horn, S.; Sánchez-Álvarez, J.; Aurnou, J. Experimental pub crawl from Rayleigh-Bénard to magnetostrophic convection. *J. Fluid Mech.* **2022**, *939*, R1. [CrossRef]
- 89. Julien, K.; Aurnou, J.; Calkins, M.; Knobloch, E.; Marti, P.; Stellmach, S.; Vasil, G. A nonlinear model for rotationally constrained convection with Ekman pumping. *J. Fluid Mech.* **2016**, 798, 50–87. [CrossRef]
- 90. Plumley, M.; Julien, K.; Marti, P.; Stellmach, S. Sensitivity of rapidly rotating Rayleigh-Bénard convection to Ekman pumping. *Phys. Rev. Fluids* **2017**, *2*, 094801. [CrossRef]
- 91. Nieves, D.; Rubio, A.; Julien, K. Statistical classification of flow morphology in rapidly rotating Rayleigh-Bénard convection. *Phys. Fluids* **2014**, *26*, 086602. [CrossRef]
- 92. Boubnov, B.; Golitsyn, G. Experimental study of convective structures in rotating fluids. *J. Fluid Mech.* **1986**, 167, 503–531. [CrossRef]
- 93. Boubnov, B.; van Heijst, G. Experiments on convection from a horizontal plate with and without background rotation. *Exp. Fluids* **1994**, *16*, 155–164. [CrossRef]
- 94. Sakai, S. The horizontal scale of rotating convection in the geostrophic regime. J. Fluid Mech. 1997, 333, 85–95. [CrossRef]
- 95. Grooms, I.; Julien, K.; Weiss, J.; Knobloch, E. Model of convective Taylor columns in rotating Rayleigh-Bénard convection. *Phys. Rev. Lett.* **2010**, *104*. [CrossRef] [PubMed]
- 96. Aurnou, J.; Calkins, M.; Julien, K.; Nieves, D.; Soderlund, K.; Stellmach, S. Rotating convective turbulence in Earth and planetary cores. *Earth Planet. Sci. Lett.* **2015**, 246, 52–71. [CrossRef]
- 97. Greenspan, H.P. The Theory of Rotating Fluids; Cambridge University Press: Cambridge, UK, 1969.
- 98. Julien, K.; Knobloch, E. Strongly nonlinear convection cells in a rapidly rotating fluid layer: The tilted f-plane. *J. Fluid Mech.* **1998**, 360, 141–178. [CrossRef]
- 99. Jones, C.A.; Soward, A.M.; Mussa, A.I. The onest of thermal convection in a rapidly rotating sphere. *J. Fluid Mech.* **2000**, 405, 157–179. [CrossRef]
- 100. Zhang, K.; Schubert, G. Magnetohydrodynamics in rapidly rotating spherical systems. *Ann. Rev. Fluid Mech.* **2000**, 32, 409–443. [CrossRef]
- 101. Stellmach, S.; Hansen, U. Cartesian convection driven dynamos at low Ekman number. Phys. Rev. E 2004, 70, 056312. [CrossRef]
- 102. Aubert, J. Steady zonal flows in spherical shell dynamos. J. Fluid Mech. 2005, 542, 53-67. [CrossRef]
- 103. Boubnov, B.; Golitsyn, G. Temperature and velocity field regimes of convective motions in a rotating plane fuid layer. *J. Fluid Mech.* **1990**, 219, 215–239. [CrossRef]
- 104. Boubnov, B.; Golitsyn, G. Convection from local sources. In *Convection in Rotating Fluids*; Springer Science and Business Media: Berlin/Heidelberg, Germany, 1995; Volume 29.
- 105. Shang, X.D.; Qiu, X.; Tong, P.; Xia, K. Measurements of the local convective heat flux in turbulent Rayleigh-Bénard convection. *Phys. Rev. E* **2004**, *70*, 026308. [CrossRef]
- 106. Hignett, P.; Ibbetson, A.; Killworth, P. On rotating thermal convection driven by non-uniform heating from below. *J. Fluid Mech.* **2006**, *109*, 161–187. [CrossRef]
- 107. Ripsei, P.; Biferale, L.; Sbragaglia, M.; Wirth, A. Natural convection with mixed insulating and conducting boundary conditions: Low- and high-Rayleigh-number regimes. *J. Fluid Mech.* **2014**, 742, 636–663. [CrossRef]
- 108. Wang, F.; Haung, S.; Xia, K. Thermal convection with mixed thermal boundary conditions: Effects of insulating lids at the top. *J. Fluid Mech.* **2017**, *817*, R1. [CrossRef]
- 109. Bakhuis, D.; Ostilla-Monico, R.; Poel, E.V.D.; Verzicco, R.; Lohse, D. Mixed insulating and conducting thermal boundary conditions in Rayleigh-Bénard convection. *J. Fluid Mech.* **2018**, *835*, 491–511. [CrossRef]
- 110. Qu, M.; Yang, D.; Liang, Z.; Liu, D. Experimental and numerical investigation on heat transfer of ultra-supercritical water in vertical upward tube under uniform and non-uniform heating. *Int. J. Heat Mass Transfer* **2018**, 127, 769–783. [CrossRef]
- 111. Evgrafova, A.; Sukhanovskii, A. Specifics of heat flux from localized heater in cylindrical layer. *Int. J. Heat Mass Transfer* **2019**, 135, 761–768. [CrossRef]
- 112. Sukhanovskii, A.; Vasiliev, A. Physical mechanism of the convective heat flux increasing in the case of mixed boundary conditions in Rayleigh-Bénard convection. *Int. J. Heat Mass Transfer* **2022**, *185*, 122411. [CrossRef]
- 113. Ciliberto, S.; Laroche, C. Random roughness of boundary increases the turbuent convection scaling exponent. *Phys. Rev. Lett.* **1999**, *82*, 3998. [CrossRef]
- 114. Qiu, X.L.; Xia, K.; Tong, P. Experimental study of velocity boundary layer near a rough conducting surface in turbulent natural convection. *J. Turbulence* **2005**, *6*, N30. [CrossRef]
- 115. Shishkina, O.; Wagner, C. Modelling the influence of wall roughness on heat transfer in thermal convection. *J. Fluid Mech.* **2011**, 686, 568–582. [CrossRef]
- 116. Toppaladoddi, S.; Succi, S.; Wettlaufer, J. Roughness as a route to the ultimate regime of thermal convection. *Phys. Rev. Lett.* **2017**, 118, 074503. [CrossRef] [PubMed]

Fluids 2023, 8, 106 34 of 35

117. Zhu, X.; Stevens, R.; Verzicco, R.; Lohse, D. Roughness facilitated local 1/2 scaling does not imply the onset of the ultimate regime of thermal convection. *Phys. Rev. Lett.* **2017**, *119*, 154501. [CrossRef] [PubMed]

- 118. Xie, Y.; Xia, K. Turbulent thermal convection over rough plates with varying roughness geometries. *J. Fluid Mech.* **2017**, *8*25, 573–599. [CrossRef]
- 119. Zhang, Y.Z.; Sun, C.; Bao, Y.; Zhou, Q. How surface roughness reduces heat transport for small roughness heights in turbulent Rayleigh-Bénard convection. *J. Fluid Mech.* **2017**, *826*, R2. [CrossRef]
- 120. Zhu, X.; Stevens, R.; Shishkina, O.; Verzicco, R.; Lohse, D.  $Nu \sim Ra^{1/2}$  scaling enabled by multiscale wall roughness in Rayleigh Bénard convection. *J. Fluid Mech.* **2019**, *869*, R4. [CrossRef]
- 121. Dong, D.L.; Wang, B.F.; Dong, Y.H.; Huang, Y.X.; Jiang, N.; Liu, Y.L.; Lu, Z.M.; Qiu, X.; Tang, Z.Q.; Zhou, Q. Influence of spatial arrangements of roughness elements on turbulent Rayleigh-Bénard convection. *Phys. Fluids* **2020**, *32*.
- 122. Lavorel, G.; Le Bars, M. Experimental study of the interaction between convective and elliptical instabilities. *Phys. Fluids* **2010**, 22, 045114. [CrossRef]
- 123. Ogilvie, G.; Lesur. On the interaction between tides and convection. Month. Not. Royal Astro. Soc. 2012, 422, 1975–1987. [CrossRef]
- 124. Wei, X. The combined effect of precession and convection on the dynamo action. Astrophys. J. 2016, 827, 123. [CrossRef]
- 125. Vormann, J.; Hansen, U. Characteristics of a precessing flow under the influence of a convecting temperature field in a spheroidal shell. *J. Fluid Mech.* **2020**, *891*, A15. [CrossRef]
- 126. Vidal, J.; Barker, A. Efficiency of tidal dissipation in slowly rotating fully convective stars or planets. *Month. Not. Royal Astro. Soc.* **2020**, 497, 4472–4485. [CrossRef]
- 127. Barker, A.; Astoul, A. On the interaction between fast tides and convection. *Month. Not. Royal Astro. Soc.* **2021**, *506*, L69–L73. [CrossRef]
- 128. Ecke, R.; Niemela, J. Heat transport in the geostrophic regime of rotating Rayleigh-Bénard convection. *Phys. Rev. Lett.* **2014**, 113, 114301. [CrossRef] [PubMed]
- 129. Yang, Y.; Verzicco, R.; Lohse, D.; Stevens, R. What rotation rate maximizes heat transport in rotating Rayleigh-B'enard convection with Prandtl number larger than one? *Phys. Rev. Fluids* **2020**, *5*, 053501. [CrossRef]
- 130. Siggia, E. High Rayleigh number convection. J. Fluid Mech. 1994, 26, 137–168. [CrossRef]
- 131. Gastine, T.; Wicht, J.; Aurnou, J. Turbulent Rayleigh–Bénard convection in spherical shells. *J. Fluid Mech.* **2015**, 778, 721–764. [CrossRef]
- 132. Christensen, U.R. Dynamo Scaling Laws and Applications to the Planets. Space Sci. Rev. 2010, 152, 565-590. [CrossRef]
- 133. Jones, C.A. Planetary Magnetic Fields and Fluid Dynamos. Ann. Rev. Fluid Mech. 2011, 43, 583-614. [CrossRef]
- 134. King, E.; Buffett, B. Flow speeds and length scales in geodynamo models: The role of viscosity. *Earth Planet. Sci. Lett.* **2013**, 371, 156–162. [CrossRef]
- 135. Jones, C.A. Treatise on Geophysics; Elsevier: Amsterdam, The Netherlands, 2015; pp. 115–159.
- 136. Gillet, N.; Jones, C.A. The quasi-geostrophic model for rapidly rotating spherical convection outside the tangent cylinder. *J. Fluid Mech.* **2006**, 554, 343–369. [CrossRef]
- 137. Hide, R. Jupiter and Saturn. Proc. R. Soc. Lond. A. 1974, 336, 63-84.
- 138. Cardin, P.; Olson, P. Chaotic thermal convection in a rapidly rotating spherical shell: Consequences for flow in the outer core. *Phys. Earth Planet. Inter.* **1994**, *82*, 235–259. [CrossRef]
- 139. Cheng, J.; Aurnou, J. Tests of diffusion-free scaling behaivors in numerical dynamo datasets. *Earth Planet. Sci. Lett.* **2016**, 436. [CrossRef]
- 140. Aguirre-Guzman, A.; Madonia, M.; Cheng, J.; Ostilla-Monico, R.; Clercx, H.; Kunnen, R. Flow- and temperature-based statistics characterizing the regimes in rapidly rotating turbulent convection in simulations employing no-slip boundary conditions. *Phys. Rev. Fluids* **2022**, *7*, 013501. [CrossRef]
- 141. Calkins, M.; Long, L.; Nieves, D.; Julien, K.; Tobias, S. Convection-driven kinematic dynamos at low Rossby and magnetic Prandtl numbers. *Phys. Rev. Fluids* **2016**, *1*, 083701. [CrossRef]
- 142. Yeh, Y.; Cummins, H. Localized fluid flow measurements with an He–Ne laser spectrometer. *Appl. Phys. Lett.* **1964**, *4*, 176–178. [CrossRef]
- 143. Warn-Varnas, A.; Fowlis, W.; Piacsek, S.; Lee, S. Numerical solutions and laser-Doppler measurements of spin-up. *J. Fluid Mech.* **1978**, *85*, 609–639. [CrossRef]
- 144. Drain, L. The Laser Doppler Technique; John Wiley and Sons: Hoboken, NJ, USA, 1980.
- 145. Noir, J.; Hemmerlin, F.; Wicht, J.; Baca, S.; Aurnou, J.M. An experimental and numerical study of librationally driven flow in planetary cores and subsurface oceans. *Phys. Earth Planet. Inter.* **2009**, *173*, 141–152. [CrossRef]
- 146. Noir, J.; Calkins, M.; Lasbleis, M.; Cantwell, J.; Aurnou, J.M. Experimental study of libration-driven zonal flows in a straight cylinder. *Phys. Earth Planet. Inter.* **2010**, *182*, 98–1106. [CrossRef]
- 147. Damaschke, N.; Kuhn, V.; Nobach, H. A fair review of non-parametric bias-free autocorrelation and spectral methods for randomly sampled data in laser Doppler velocimetry. *Digit. Signal Process.* **2018**, *76*, 22–33. [CrossRef]
- 148. Hawkins, E.K. *Experimental Investigations of Rapidly Rotating Convective Turbulence in Planetary Cores*; ProQuest Dissertations Publishing: Ann Arbor, MI, USA, 2020.
- 149. Brito, D.; Aurnou, J.M.; Cardin, P. Turbulent viscosity measurements relevant to planetary core-mantle dynamics. *Phys. Earth Planet. Inter.* **2004**, 141, 3–8. [CrossRef]

Fluids 2023, 8, 106 35 of 35

- 150. Burmann, F.; Noir, J. Effects of bottom topography on the spin-up in a cylinder. Phys. Fluids 2018, 30, 106601. [CrossRef]
- 151. Johnson, D.; Modarress, D.; Owen, F. An experimental verification of Laser-Velocimeter sampling bias and its correction. *J. Fluid Eng.* **1984**, *106*, 5–12. [CrossRef]
- 152. Edwards, R. Report of the special panel on statistical particle bias problems in Laser Anemometry. *J. Fluid Eng.* **1987**, *109*, 89–93. [CrossRef]
- 153. Efstathiou, C.; Luhar, M. Mean turbulence statistics in boundary layers over high-porosity foams. *J. Fluid Mech.* **2018**, *841*, 351–379. [CrossRef]
- 154. Batchelor, G. The Theory of Homogeneous Turbulence; Cambrdige University Press: Cambrdige, UK, 1953.
- 155. Wilczek, M.; Daitche, A.; Freidrich, R. On the velocity distribution in homogeneous isotropic turbulence: Correlations and deviations from Gaussianity. *J. Fluid Mech.* **2011**, *676*, 191–217. [CrossRef]
- 156. Schaeffer, N.; Pais, M. On symmetry and anisotropy of Earth-core flows. Geophys. Res. Lett. 2011, 38, L10309. [CrossRef]
- 157. Stellmach, S.; Hansen, U. An efficient spectral method for the simulation of dynamos in Cartesian geometry and its implementation on massively parallel computers. *Geochem. Geophys. Geosyst.* **2008**, *9*, Q05003. [CrossRef]
- 158. Shang, X.; Qiu, X.; Tong, P.; Xia, K. Measured local heat transport in turbulent Rayleigh-Bénard convection. *Phys. Rev. Lett.* **2003**, 90, 074501. [CrossRef]
- 159. Sun, C.; Xia, K.; Tong, P. Three-dimensional flow structures and dynamics of turbulent thermal convection in a cylindrical cell. *Phys. Rev. E* **2005**, *72*, 026302. [CrossRef] [PubMed]
- 160. Zurner, T.; Schindler, F.; Vogt, T.; Eckert, S.; Shumacher, J. Combined measurement of velocity and temperature in liquid metal convection. *J. Fluid Mech.* **2019**, *876*, 1108–1128. [CrossRef]
- 161. Xu, Y.; Horn, S.; Aurnou, J. Thermoelectric precession in turbulent magnetoconvection. J. Fluid Mech. 2022, 930, A8. [CrossRef]
- 162. Horn, S.; Aurnou, J.M. Rotating convection with centrifugal buoyancy: Numerical predictions for laboratory experiments. *Phys. Rev. Fluids* **2019**, *4*, 073501. [CrossRef]
- 163. Bouillaut, V.; Miquel, B.; Julien, K.; Aumaitre, S.; Gallet, B. Experimental observation of the geostrophic turbulence regime of rapidly rotating convection. *Proc. Natl. Acad. Sci. USA* **2021**, *118*, e2105015118. [CrossRef] [PubMed]
- 164. Stevens, R.; Clercx, H.; Lohse, D. Heat transport and flow structure in rotating Rayleigh-Bénard convection. *Eur. J. Mech Fluids* **2013**, *40*, 41–49. [CrossRef]
- 165. Christensen, U.R. Zonal flow driven by strongly supercritical convection in rotating spherical shells. *J. Fluid Mech.* **2002**, *470*, 115–133. [CrossRef]
- 166. Christensen, U.R.; Aubert, J. Scaling properties of convection-driven dynamos in rotating spherical shells and applications to planetary mangetic fields. *Geophys. J. Int.* **2006**, *166*, 97–114. [CrossRef]
- 167. Favier, B.; Silvers, L.J.; Proctor, M.R.E. Inverse cascade and symmetry breaking in rapidly rotating Boussinesq convection. *Phys. Fluids* **2014**, *9*, 096605. [CrossRef]
- 168. Rubio, A.M.; Julien, K.; Knobloch, E.; Weiss, J.B. Upscale Energy Transfer in Three-Dimensional Rapidly Rotating Turbulent Convection. *Phys. Rev. Lett.* **2014**, *112*, 144501. [CrossRef]
- 169. Moffatt, K.; Dormy, E. Self-Exciting Fluid Dynamos; Cambrdige University Press: Cambrdige, UK, 2019.
- 170. Aguirre-Guzman, A.; Madonia, M.; Cheng, J.S.; Ostilla-Monico, R.; Clercx, H.; Kunnen, R. Force balance in rapidly rotating Rayleigh B'enard convection. *J. Fluid Mech.* **2021**, *928*, A16. [CrossRef]
- 171. Mininni, P.; Alexakis, A.; Pouquet, A. Scale interactions and scaling laws in rotating flows at moderate Rossby numbers and large Reynolds numbers. *Phys. Fluids* **2009**, *21*, 015108. [CrossRef]
- 172. Guervilly, C.; Hughes, D.W.; Jones, C.A. Large-scale vortices in rapidly rotating Rayleigh-Bénard convection. *J. Fluid Mech.* **2014**, 758, 407–435. [CrossRef]
- 173. Kunnen, R.; Ostilla-Monico, R.; Poel, E.V.D.; Verzicco, R.; Lohse, D. Transition to geostrophic convection: The role of boundary conditions. *J. Fluid Mech.* **2016**, 799, 413–432. [CrossRef]
- 174. Aguirre-Guzmán, A.; Madonia, M.; Cheng, J.; Ostilla-Monico, R.; Clercx, H.; Kunnen, R. Competition between Ekman Plumes and Vortex Condensates in Rapidly Rotating Thermal Convection. *Phys. Rev. Lett.* **2020**, *125*, 214501. [CrossRef]
- 175. Rajaei, H.; Kunnen, R.; Clercx, H. Exploring the geostrophic regime of rapidly rotating convection with experiments. *Phys. Fluids* **2017**, *29*, 214501. [CrossRef]
- 176. Guervilly, C.; Cardin, P.; Schaeffer, N. Turbulent convective length scales in planetary cores. Nature 2019, 570, 368–371. [CrossRef]
- 177. Madonia, M.; Guzman, A.A.; Clercx, H.; Kunnen, R. Velocimetry in rapidly rotating convection: Spatial correlations, flow structures, and length scales. *Europhys. Lett.* **2021**, *135*, 54002. [CrossRef]
- 178. John, H.; Leinhard, I.V.; Leinhard, V. A Heat Transfer Textbook; Phlogiston Press: Cambridge, UK, 2008.
- 179. Taylor, J. Introduction to Error Analysis, the Study of Uncertainties in Physical Measurements; University Science Books: Perth, Australia, 1997.

**Disclaimer/Publisher's Note:** The statements, opinions and data contained in all publications are solely those of the individual author(s) and contributor(s) and not of MDPI and/or the editor(s). MDPI and/or the editor(s) disclaim responsibility for any injury to people or property resulting from any ideas, methods, instructions or products referred to in the content.