

Hodge theory on ALG^* manifolds

By *Gao Chen* at Shanghai, *Jeff Viaclovsky* at Irvine and *Ruobing Zhang* at Princeton

Abstract. We develop a Fredholm theory for the Hodge Laplacian in weighted spaces on ALG^* manifolds in dimension four. We then give several applications of this theory. First, we show the existence of harmonic functions with prescribed asymptotics at infinity. A corollary of this is a non-existence result for ALG^* manifolds with non-negative Ricci curvature having group $\Gamma = \{e\}$ at infinity. Next, we prove a Hodge decomposition for the first de Rham cohomology group of an ALG^* manifold. A corollary of this is vanishing of the first Betti number for any ALG^* manifold with non-negative Ricci curvature. Another application of our analysis is to determine the optimal order of ALG^* gravitational instantons.

Contents

1. Introduction
 2. ALG^* model space
 3. Weighted analysis on the ALG^* model space
 4. Weighted analysis on ALG^* manifolds
 5. Applications
 - A. Computations on the ALG^* model space
- References

1. Introduction

In this paper, we are interested in complete Riemannian metrics in dimension four which are asymptotic to certain Ricci-flat model spaces at infinity. Many of these types of geometries have been previously studied, and are known as ALE, ALF, ALG, ALH, and two exceptional types known as ALG^* and ALH^* ; see for example [2, 4, 6, 13, 14] and the references therein. In this paper, we will concentrate on the first exceptional type, ALG^* , which we define next.

The corresponding author is Jeff Viaclovsky.

The first named author is supported by the Project of Stable Support for Youth Team in Basic Research Field, Chinese Academy of Sciences, YSBR-001. The second named author is supported by NSF Grants DMS-1811096 and DMS-2105478. The third named author is supported by NSF Grant DMS-1906265.

Let Nil_ν^3 be the Heisenberg nilmanifold of degree $\nu \in \mathbb{Z}_+$ with coordinates $(\theta_1, \theta_2, \theta_3)$ on the universal cover; see Section 2 for our conventions.

Definition 1.1 (ALG_ν^* model space). For $\nu \in \mathbb{Z}_+$, the ALG_ν^* model manifold is

$$\widehat{\mathfrak{M}}_\nu(R) \equiv (R, \infty) \times \text{Nil}_\nu^3.$$

Let $V = \kappa_0 + \frac{\nu}{2\pi} \log r$, where $\kappa_0 \in \mathbb{R}$, and assume that R satisfies $R > e^{\frac{2\pi}{\nu}(1-\kappa_0)}$. The model metric on $\widehat{\mathfrak{M}}_\nu(R)$ is given by

$$(1.1) \quad g_{\kappa_0}^{\widehat{\mathfrak{M}}} = V(dr^2 + r^2 d\theta_1^2 + d\theta_2^2) + V^{-1} \frac{\nu^2}{4\pi^2} (d\theta_3 - \theta_2 d\theta_1)^2.$$

Given $L > 0$, we let $g_{\kappa_0, L}^{\widehat{\mathfrak{M}}} = L^2 g_{\kappa_0}^{\widehat{\mathfrak{M}}}$.

The metric in (1.1) arises naturally from the Gibbons–Hawking ansatz; see Section 2. Note that the model metric restricts to a left-invariant metric on any cross-section $\{r_0\} \times \text{Nil}_\nu^3$. Next, let Γ be any finite group acting freely and isometrically on $\widehat{\mathfrak{M}}_\nu(R)$, and denote the quotient space by $\mathfrak{M}_\nu(R)$. We then define the following, which are the main objects of interest in this paper.

Definition 1.2 (ALG_ν^* - Γ manifold). A complete 4-manifold (X, g) is called an ALG_ν^* - Γ manifold of order $n > 0$ with parameters $\nu \in \mathbb{Z}_+$, $\kappa_0 \in \mathbb{R}$ and $L \in \mathbb{R}_+$ if there exist an ALG^* model space $(\widehat{\mathfrak{M}}_\nu(R), g_{\kappa_0, L}^{\widehat{\mathfrak{M}}})$ with $R > 0$, a compact subset $X_R \subset X$, and a diffeomorphism $\Phi: \widehat{\mathfrak{M}}_\nu(R)/\Gamma \rightarrow X \setminus X_R$ such that

$$(1.2) \quad |\nabla_g^k (\Phi^* g - g_{\kappa_0, L}^{\widehat{\mathfrak{M}}})|_g = O(\varepsilon^{-k-n}),$$

as $\varepsilon(x) \equiv r(x)V^{\frac{1}{2}}(x) \rightarrow \infty$, for any $x \in \widehat{\mathfrak{M}}_\nu(R)$, for any $k \in \mathbb{N}_0 \equiv \mathbb{Z}_+ \cup \{0\}$.

We note that allowing the quotient by Γ is analogous to the asymptotically flat (AF) case versus the asymptotically locally Euclidean (ALE) case. Fix $x_0 \in X_R$, and for $x \in X$, define $s(x) \equiv d_g(x_0, x)$. We also note that ALG_ν^* - Γ metrics satisfy the following properties:

- (1) $s(x) \sim \varepsilon(\Phi^{-1}(x))$ as $s(x) \rightarrow \infty$, see Remark 2.1,
- (2) $\text{Vol}_g(B_t(x_0)) \sim t^2$ as $t \rightarrow \infty$, and
- (3) $|\text{Rm}_g| = O(s^{-2}(\log s)^{-1})$ as $s(x) \rightarrow \infty$.

Fredholm theory of the Hodge Laplacian for the geometries listed above has been developed in many works; see for example [3, 5–7, 11, 14, 17, 19]. It is also worth adding a historical remark regarding the ALG^* model geometry. ALG^* hyper-Kähler manifolds appeared in the math literature before the special case of ALG manifolds. In fact, Cherkis–Kapustin first studied the ALG^* model geometry, ALG^* hyper-Kähler manifolds, as well as D_4 -ALG hyper-Kähler manifolds; see [9]. The ALG name was introduced in that paper (page 2, last paragraph), which included both the ALG and the ALG^* cases. Later, these cases were separated in order to distinguish the different model geometries.

In this paper, we will develop a Fredholm theory for the Hodge Laplacian on ALG_ν^* - Γ manifolds; see Proposition 4.5. Due to the peculiarities of the asymptotic geometry of ALG^* metrics, this is quite nontrivial; it is proved in Sections 3 and 4, which rely on some lengthy

formulas computed in the appendix. We will not describe this theory here in the introduction, but instead, we will turn our attention to a number of applications. Note that $h_0 \equiv 1$ is obviously a harmonic function. Our first application is to the existence of non-constant harmonic functions with prescribed asymptotics.

Theorem 1.3. *Let (X, g) be an ALG_v^* - Γ manifold of order $n > 0$, and let $k \in \mathbb{Z}_+$. If $r^k e^{\sqrt{-1}k\theta_1}$ is invariant under Γ , then for any $0 < \epsilon < \min\{k, n\}$, there exists a harmonic function $h_k: X \rightarrow \mathbb{C}$ such that*

$$(1.3) \quad h_k = r^k e^{\sqrt{-1}k\theta_1} + O(s^{k-\epsilon}),$$

as $s \rightarrow \infty$.

A corollary of this is the following non-existence result.

Corollary 1.4. *There do not exist any ALG_v^* - Γ manifolds of order $n > 0$ with $\Gamma = \{e\}$ and with non-negative Ricci curvature.*

To prove this, we will use Theorem 1.3 to find a certain harmonic 1-form which, using the Bochner formula, will lead to a contradiction. We note that there do exist complete ALG_v^* - Γ metrics with non-negative Ricci curvature for Γ nontrivial. A rough analogy with the AF setting: any AF metric with non-negative Ricci curvature must be Euclidean space. But there are many non-flat ALE metrics with Γ nontrivial and with non-negative Ricci curvature.

Our next application is a Hodge decomposition theorem for the first de Rham cohomology group. Define the following subspace of $\Omega^1(\hat{\mathcal{M}}_v(R)/\Gamma)$:

$$\mathcal{W}^1 \equiv \begin{cases} \mathbb{R} \cdot d\theta_2 & \text{if } \gamma^* d\theta_2 = d\theta_2 \text{ for all } \gamma \in \Gamma, \\ \{0\} & \text{otherwise.} \end{cases}$$

Theorem 1.5. *Let (X, g) be an ALG_v^* - Γ manifold. Then*

$$H_{dR}^1(X) \cong \{\omega \in \Omega^1(X) \mid d\omega = 0, \delta\omega = 0, \Phi^*\omega = \omega_0 + O(s^{-\epsilon}) \text{ as } s \rightarrow \infty, \omega_0 \in \mathcal{W}^1\}$$

for any ϵ satisfying $0 < \epsilon \ll 1$.

The proof of this is found in Section 5.2. There is an analogous result for the second de Rham cohomology group which can be proved using similar methods, but for simplicity, we do not state this in this paper. As a corollary of Theorem 1.5, we have the following vanishing theorem.

Corollary 1.6. *Let (X, g) be an ALG_v^* - Γ manifold with non-negative Ricci curvature. Then the first Betti number is $b^1(X) = 0$.*

We note that, since any such manifold has quadratic volume growth at infinity, this corollary can also be seen to follow from [1, Theorem 2.1 and Lemma 2.2]. Anderson's results are much more general, but his proof uses techniques from geometric measure theory. In contrast, our proof of Corollary 1.6 is more elementary, but is specialized to ALG^* geometry.

We next recall the following definition.

Definition 1.7. A hyper-Kähler 4-manifold (X, g, I, J, K) is a Riemannian 4-manifold (X, g) with a triple of Kähler structures $(g, I), (g, J), (g, K)$ such that $IJ = K$. Let $\omega_1, \omega_2, \omega_3$ denote the Kähler forms for I, J, K , respectively.

As we will see in Section 2.2, the model space has a hyper-Kähler structure, and we denote the triple of Kähler forms by $\omega_{i, \kappa_0, L}^{\mathfrak{M}}, i = 1, 2, 3$. In Section 2.3, we consider all possible isometric quotients of the model space which retain this hyper-Kähler structure. We show that, without loss of generality, we may assume that either Γ is trivial, or ν is even and Γ is a specific $\mathbb{Z}_2$ -action $\iota: \hat{\mathfrak{M}}_{2\nu}(R) \rightarrow \hat{\mathfrak{M}}_{2\nu}(R)$, up to hyper-Kähler rotation and scaling. In the latter case, denote the quotient space by $\mathfrak{M}_{2\nu}(R) = \hat{\mathfrak{M}}_{2\nu}(R)/\mathbb{Z}_2$. The model Kähler forms descend to the quotient, which we denote by $\omega_{i, \kappa_0, L}^{\mathfrak{M}}, i = 1, 2, 3$.

Next, we will consider the case of complete hyper-Kähler ALG_ν^* - Γ manifolds. Since any hyper-Kähler metric is Ricci-flat, by Corollary 1.4, Γ cannot be trivial. Consequently, by the remarks in the previous paragraph, we can make the following definition.

Definition 1.8. An ALG_ν^* gravitational instanton (X, g, I, J, K) of order $\mathfrak{n} > 0$ with parameters $\nu \in \mathbb{Z}_+, \kappa_0 \in \mathbb{R}, L \in \mathbb{R}_+$ is a hyper-Kähler 4-manifold which is also an $\text{ALG}_{2\nu}^*$ - $\mathbb{Z}_2$ manifold (X, g) with the $\mathbb{Z}_2$ -action given by ι such that, in addition to Definition 1.2, the hyper-Kähler forms satisfy

$$(1.4) \quad |\nabla_g^k (\Phi^* \omega_i - L^2 \omega_{i, \kappa_0, L}^{\mathfrak{M}})|_{g^{\mathfrak{M}}} = O(\varsigma^{-k-\mathfrak{n}})$$

for $i = 1, 2, 3$ as $\varsigma(x) \equiv r(x)V^{\frac{1}{2}}(x) \rightarrow \infty$, for any $k \in \mathbb{N}_0$.

Remark 1.9. Conditions (1.4) necessarily imply (1.2); see Section 5.3 below.

Our next main application is to determine the optimal order of ALG_ν^* gravitational instantons.

Theorem 1.10. *If (X, g, I, J, K) is an ALG_ν^* gravitational instanton of order $\mathfrak{n} > 0$ with respect to ALG_ν^* coordinates Φ , then there exists an ALG_ν^* coordinate system Φ' as in Definition 1.8 with order $\mathfrak{n}' = 2$.*

This is proved as an application of our Fredholm theory applied to a certain Dirac-type operator and for the Hodge Laplacian on 2-forms; see Section 5.3.

2. ALG^* model space

In this section, we explain some properties of ALG^* metrics in more detail.

2.1. The model metric. The 3-dimensional Heisenberg group is

$$H(1, \mathbb{R}) \equiv \left\{ \begin{bmatrix} 1 & \theta_2 & \theta_3 \\ 0 & 1 & \theta_1 \\ 0 & 0 & 1 \end{bmatrix} : \theta_1, \theta_2, \theta_3 \in \mathbb{R} \right\}.$$

For $\nu \in \mathbb{Z}_+$, the Heisenberg nilmanifold Nil_ν^3 of degree ν is the quotient of $H(1, \mathbb{R})$ by the left action of the subgroup

$$H(1, \mathbb{Z}) \equiv \left\{ \begin{bmatrix} 1 & 2\pi k & 4\pi^2 \nu^{-1} l \\ 0 & 1 & 2\pi m \\ 0 & 0 & 1 \end{bmatrix} : k, l, m \in \mathbb{Z} \right\}$$

generated by

$$\begin{aligned} \sigma_1(\theta_1, \theta_2, \theta_3) &\equiv (\theta_1 + 2\pi, \theta_2, \theta_3), \\ \sigma_2(\theta_1, \theta_2, \theta_3) &\equiv (\theta_1, \theta_2 + 2\pi, \theta_3 + 2\pi\theta_1), \\ \sigma_3(\theta_1, \theta_2, \theta_3) &\equiv (\theta_1, \theta_2, \theta_3 + 4\pi^2 \nu^{-1}). \end{aligned}$$

Note that the forms

$$d\theta_1, \quad d\theta_2, \quad \Theta \equiv \frac{\nu}{2\pi}(d\theta_3 - \theta_2 d\theta_1)$$

are a basis of left-invariant 1-forms. Also, it is clear that Nil_ν^3 is the total space of a degree ν circle fibration

$$S^1 \rightarrow \text{Nil}_\nu^3 \xrightarrow{\pi} T^2 \equiv \mathbb{R}_{\theta_1, \theta_2}^2 / \Lambda.$$

We next consider the Gibbons–Hawking ansatz

$$S^1 \rightarrow \widehat{\mathfrak{M}}_\nu(R) \equiv (R, \infty) \times \text{Nil}_\nu^3 \rightarrow \tilde{U} \equiv (\mathbb{R}^2 \setminus \overline{B_R(0)}) \times S^1,$$

with the radial harmonic function

$$V = \kappa_0 + \frac{\nu}{2\pi} \log r, \quad r \in (R, \infty), \quad \kappa_0 \in \mathbb{R}, \quad R > e^{\frac{2\pi}{\nu}(1-\kappa_0)};$$

for details of the Gibbons–Hawking ansatz construction, see [12, 16]. We use the coordinates

$$(x, y, \theta_2) = (r \cos(\theta_1), r \sin(\theta_1), \theta_2) \quad \text{on } (\mathbb{R}^2 \setminus \overline{B_R(0)}) \times S^1$$

and fix the orientation $r dr \wedge d\theta_1 \wedge d\theta_2$. Then we have

$$*\mathbb{R}^2 \times S^1 \odot d(V) = \frac{\nu}{2\pi} d\theta_1 \wedge d\theta_2,$$

and hence $\frac{1}{2\pi} [* \mathbb{R}^2 \times S^1 \odot d(V)] \in H^2(\tilde{U}; \mathbb{Z})$. Note that the form Θ is a connection form such that $\Omega = d\Theta = *dV$. The Gibbons–Hawking metric is

$$\begin{aligned} g_{\kappa_0}^{\widehat{\mathfrak{M}}} &= V(dx^2 + dy^2 + d\theta_2^2) + V^{-1}\Theta^2 \\ &= V(dr^2 + r^2 d\theta_1^2 + d\theta_2^2) + V^{-1} \frac{\nu^2}{4\pi^2} (d\theta_3 - \theta_2 d\theta_1)^2. \end{aligned}$$

If Γ is some specified finite group acting freely and isometrically on $\widehat{\mathfrak{M}}_\nu(R)$, then we will denote the quotient space by $\mathfrak{M}_\nu(R) = \widehat{\mathfrak{M}}_\nu(R)/\Gamma$.

Remark 2.1. Choose a point $p_0 \in \widehat{\mathfrak{M}}_\nu(R)$. By straightforward computations, one can see that there exists a constant $C > 0$ such that, for any $q \in \widehat{\mathfrak{M}}_\nu(R)$,

$$C^{-1} \cdot r(q) \cdot V(q)^{\frac{1}{2}} \leq d_{g_{\kappa_0}^{\widehat{\mathfrak{M}}}}(q, p_0) \leq C \cdot r(q) \cdot V(q)^{\frac{1}{2}}.$$

2.2. Hyper-Kähler structure. On $\widehat{\mathfrak{M}}_\nu(R)$, define an orthonormal basis

$$\{E^1, E^2, E^3, E^4\} = \{V^{\frac{1}{2}}dx, V^{\frac{1}{2}}dy, V^{\frac{1}{2}}d\theta_2, V^{-\frac{1}{2}}\Theta\}.$$

We define three almost complex structures $I, J, K \in C^\infty(\text{End}(T\widehat{\mathfrak{M}}))$ on $\widehat{\mathfrak{M}}_\nu(R)$ by requiring the dual linear maps $I^*, J^*, K^* \in C^\infty(\text{End}(T^*\widehat{\mathfrak{M}}))$ satisfying

$$\begin{aligned} I^*(E^1) &= -E^2, & I^*(E^3) &= -E^4, \\ J^*(E^1) &= -E^3, & J^*(E^2) &= E^4, \\ K^*(E^1) &= -E^4, & K^*(E^2) &= -E^3. \end{aligned}$$

It is clear that each complex structure is Hermitian with respect to g . Moreover, $K = IJ$ because $K^* = J^*I^*$. Using the convention that $\omega_J(X, Y) = g(JX, Y)$, the corresponding Kähler forms are

$$\begin{aligned} \omega_I &= \omega_{1, \kappa_0}^{\widehat{\mathfrak{M}}} = E^1 \wedge E^2 + E^3 \wedge E^4 = V dx \wedge dy + d\theta_2 \wedge \Theta, \\ (2.1) \quad \omega_J &= \omega_{2, \kappa_0}^{\widehat{\mathfrak{M}}} = E^1 \wedge E^3 - E^2 \wedge E^4 = V dx \wedge d\theta_2 - dy \wedge \Theta, \end{aligned}$$

$$(2.2) \quad \omega_K = \omega_{3, \kappa_0}^{\widehat{\mathfrak{M}}} = E^1 \wedge E^4 + E^2 \wedge E^3 = dx \wedge \Theta + V dy \wedge d\theta_2.$$

We notice that $d\omega_I = d\omega_J = d\omega_K = 0$. For example,

$$d\omega_I = dV \wedge *d\theta_2 - d\theta_2 \wedge d\Theta = dV \wedge *d\theta_2 - d\theta_2 \wedge *dV = 0,$$

where $*$ $\equiv *_{\mathbb{R}^2 \times S^1}$ is the Hodge star operator defined with respect to the flat metric on $\mathbb{R}^2 \times S^1$. The last equality holds since $*\alpha \wedge \beta = \alpha \wedge *\beta$ for 1-forms $\alpha, \beta \in \Omega^1(\mathbb{R}^2 \times S^1)$. The computations for $d\omega_J$ and $d\omega_K$ are similar.

By [15, Lemma 6.8], the triple

$$\{\omega_{1, \kappa_0}^{\widehat{\mathfrak{M}}}, \omega_{2, \kappa_0}^{\widehat{\mathfrak{M}}}, \omega_{3, \kappa_0}^{\widehat{\mathfrak{M}}}\} \equiv \{\omega_I, \omega_J, \omega_K\}$$

is a hyper-Kähler triple on $\widehat{\mathfrak{M}}_\nu(R)$. In particular, the complex structures I, J, K are integrable, and the metric is Ricci-flat Kähler with respect to all three of these complex structures.

2.3. Quotients. There are many possibilities for the group Γ ; see for example [10]. We will not analyze all the possibilities here, but will only address the question of which quotients retain the hyper-Kähler structure. To this end, we have the following proposition.

Proposition 2.2. *If Γ is a finite group acting freely and isometrically on $(\widehat{\mathfrak{M}}_\nu(R), g_{\kappa_0}^{\widehat{\mathfrak{M}}})$ and preserving I, J, K , then either Γ is generated by ξ_l and $\zeta_{k, l, m}$ for integers $k, l \in \mathbb{N}$, $m \in \mathbb{N}_0$ such that k divides ν and $0 \leq m \leq kl - 1$, or ν is even and Γ is generated by ξ_l , $\zeta_{k, l, m}$, $\iota_{n, t}$ for integers k, l, m satisfying the same conditions, $n \in \mathbb{N}_0$ satisfying $0 \leq n \leq \nu - 1$ with nl even, and $t \in \mathbb{R}$.*

Proof. Define

$$\begin{aligned} \xi_l(\theta_1, \theta_2, \theta_3) &\equiv (\theta_1, \theta_2, \theta_3 + 4\pi^2 l^{-1} \nu^{-1}), \\ \zeta_{k, l, m}(\theta_1, \theta_2, \theta_3) &\equiv (\theta_1, \theta_2 + 2\pi k^{-1}, \theta_3 + 2\pi k^{-1} \theta_1 + 4\pi^2 m l^{-1} k^{-1} \nu^{-1}), \\ \iota_{n, t}(\theta_1, \theta_2, \theta_3) &\equiv (\theta_1 + \pi, 2\pi n \nu^{-1} - \theta_2, 2\pi n \nu^{-1} \theta_1 + t - \theta_3). \end{aligned}$$

Then

$$\begin{aligned}\xi_l \sigma_1 \xi_l^{-1} &= \sigma_1, & \xi_l \sigma_2 \xi_l^{-1} &= \sigma_2, & \xi_l \sigma_3 \xi_l^{-1} &= \sigma_3, \\ \zeta_{k,l,m} \sigma_1 \zeta_{k,l,m}^{-1} &= \sigma_1 \sigma_3^{v/k}, & \zeta_{k,l,m} \sigma_2 \zeta_{k,l,m}^{-1} &= \sigma_2, & \zeta_{k,l,m} \sigma_3 \zeta_{k,l,m}^{-1} &= \sigma_3, \\ \iota_{n,t} \sigma_1 \iota_{n,t}^{-1} &= \sigma_1 \sigma_3^n, & \iota_{n,t} \sigma_2 \iota_{n,t}^{-1} &= \sigma_3^{v/2} \sigma_2^{-1}, & \iota_{n,t} \sigma_3 \iota_{n,t}^{-1} &= \sigma_3^{-1}.\end{aligned}$$

These imply that $\xi_l, \zeta_{k,l,m}, \iota_{n,t}$ descend to actions on $H(1, \mathbb{R})/H(1, \mathbb{Z})$. Moreover, it is easy to see that they induce actions on $\hat{\mathfrak{M}}$ which fix $(g_{\kappa_0}^{\hat{\mathfrak{M}}}, \omega_{1,\kappa_0}^{\hat{\mathfrak{M}}}, \omega_{2,\kappa_0}^{\hat{\mathfrak{M}}}, \omega_{3,\kappa_0}^{\hat{\mathfrak{M}}})$. Note that

$$\xi_l^l = \sigma_3, \quad \zeta_{k,l,m}^k = \xi_l^m \sigma_2, \quad \iota_{n,t}^2 = \xi_l^{nl/2} \sigma_1.$$

A routine calculation shows that they generate a finite subgroup.

Conversely, an isometry of the Gibbons–Hawking metric must map S^1 -fibers to S^1 -fibers, so there is a homomorphism h from Γ to the isometry group of $(\mathbb{R}^2 \setminus B_R(0^2)) \times S^1$. We first consider the kernel of h . Since Γ is a finite group, we see that the kernel of h is generated by ξ_l for some $l \in \mathbb{N}$. For the image of h , the isometries of the base are rotations and reflections in $\mathbb{R}^2$ and similarly on S^1 . Using (2.1) and (2.2), we see that ω_J and ω_K are invariant only if

$$(2.3) \quad (dx, dy, d\theta_2, \Theta) \mapsto (dx, dy, d\theta_2, \Theta), \text{ or}$$

$$(2.4) \quad (dx, dy, d\theta_2, \Theta) \mapsto (-dx, -dy, -d\theta_2, -\Theta).$$

Consider the case that every element in Γ satisfies (2.3); the only possibility is that $h(\Gamma)$ is generated by

$$h(\gamma): (r, \theta_1, \theta_2) \mapsto (r, \theta_1, \theta_2 + 2\pi a),$$

where $a \in \mathbb{R}$. Since Γ is a finite group, we can assume that $a = \frac{1}{k}$, where $k \in \mathbb{N}$. Using the condition that Θ is fixed under the action, we see that

$$\gamma(r, \theta_1, \theta_2, \theta_3) = (r, \theta_1, \theta_2 + 2\pi k^{-1}, \theta_3 + 2\pi k^{-1} \theta_1 + b),$$

where $b \in \mathbb{R}$. Since $\gamma \sigma_1 \gamma^{-1}$ is in $H(1, \mathbb{Z})$, we see that k divides v . Moreover, since γ^k is in the kernel of h , it must be $\xi_l^m \sigma_2$, where $m \in \mathbb{Z}$. This implies that $b = 4\pi^2 m l^{-1} k^{-1} v^{-1}$. By multiplying with σ_3 , we can assume that $0 \leq m \leq kl - 1$.

Next, consider that case that (2.4) happens for some element $\gamma \in \Gamma$. Then

$$\gamma(r, \theta_1, \theta_2, \theta_3) = (r, \theta_1 + \pi, c - \theta_2, c\theta_1 + t - \theta_3),$$

where $c, t \in \mathbb{R}$. Using $\gamma \sigma_1 \gamma^{-1} \in H(1, \mathbb{Z})$, we see that $c = 2\pi n v^{-1}$ for $n \in \mathbb{Z}$. By multiplying with σ_2 , we can assume that $0 \leq n \leq v - 1$. Using $\gamma \sigma_2 \gamma^{-1} \in H(1, \mathbb{Z})$, we see that v is even. Finally, since $\gamma^2 = \xi_l^{nl/2} \sigma_1$ is in the kernel of h , we see that nl is also even. $\square$

Remark 2.3. Note that, in the first case in Proposition 2.2, $h(\Gamma)$ is a cyclic group $\mathbb{Z}_k$ consisting of rotations of the S^1 factor. The kernel of h is also a cyclic group $\mathbb{Z}_l$. There is a short exact sequence

$$0 \rightarrow \mathbb{Z}_l \rightarrow \Gamma \xrightarrow{h} \mathbb{Z}_k \rightarrow 0,$$

so we must have $\Gamma = \mathbb{Z}_l \rtimes \mathbb{Z}_k$, and since Γ is abelian, we must have $\Gamma = \mathbb{Z}_l \times \mathbb{Z}_k$. In the second case, $h(\Gamma)$ is a dihedral group $D_k = \mathbb{Z}_k \rtimes \mathbb{Z}_2$, and similarly $\Gamma = \mathbb{Z}_l \rtimes D_k$. In fact, it is not hard to see from the above presentation that $\Gamma = (\mathbb{Z}_l \times \mathbb{Z}_k) \rtimes \mathbb{Z}_2$.

Remark 2.4. Recall that, in Definition 1.1, we defined the scaled metric $g_{\kappa_0, L}^{\widehat{\mathfrak{M}}} = L^2 g_{\kappa_0}^{\widehat{\mathfrak{M}}}$. The complex structures of course do not depend on any scaling, but we also define the rescaled Kähler forms $\omega_{i, \kappa_0, L}^{\widehat{\mathfrak{M}}} \equiv L^2 \omega_{i, \kappa_0}^{\widehat{\mathfrak{M}}}$ for $i = 1, 2, 3$.

The next proposition deals with the first case above.

Proposition 2.5. *In the first case in Proposition 2.2, the quotient space*

$$(2.5) \quad (\widehat{\mathfrak{M}}_v(R), g_{\kappa_0}^{\widehat{\mathfrak{M}}}, \omega_{1, \kappa_0}^{\widehat{\mathfrak{M}}}, \omega_{2, \kappa_0}^{\widehat{\mathfrak{M}}}, \omega_{3, \kappa_0}^{\widehat{\mathfrak{M}}}) / \Gamma$$

can be identified with $(\widehat{\mathfrak{M}}_{\widetilde{v}}(\widetilde{R}), g_{\widetilde{\kappa}_0, L}^{\widehat{\mathfrak{M}}}, \omega_{1, \widetilde{\kappa}_0, L}^{\widehat{\mathfrak{M}}}, \omega_{2, \widetilde{\kappa}_0, L}^{\widehat{\mathfrak{M}}}, \omega_{3, \widetilde{\kappa}_0, L}^{\widehat{\mathfrak{M}}})$ after a hyper-Kähler rotation, for some $\widetilde{v}$, $\widetilde{R}$, $\widetilde{\kappa}_0$, and L .

Proof. Letting $\widetilde{\theta}_1 = \theta_1 + 2\pi m l^{-1} v^{-1}$, $\widetilde{\theta}_2 = k\theta_2$ and $\widetilde{\theta}_3 = k\theta_3$, the Gibbons–Hawking metric becomes

$$\begin{aligned} g_{\kappa_0}^{\widehat{\mathfrak{M}}} &= V(dr^2 + r^2 d\widetilde{\theta}_1^2 + k^{-2} d\widetilde{\theta}_2^2) + V^{-1} \frac{v^2}{4\pi^2 k^2} (d\widetilde{\theta}_3 - \widetilde{\theta}_2 d\widetilde{\theta}_1)^2 \\ &= k^{-1} l^{-1} (V k^{-1} l) (k^2 (dr^2 + r^2 d\widetilde{\theta}_1^2) + d\widetilde{\theta}_2^2) \\ &\quad + k^{-1} l^{-1} (V k^{-1} l)^{-1} \frac{(v k^{-1} l)^2}{4\pi^2} (d\widetilde{\theta}_3 - \widetilde{\theta}_2 d\widetilde{\theta}_1)^2. \end{aligned}$$

Letting $\widetilde{V} = V k^{-1} l$, $\widetilde{r} = k r$, $\widetilde{v} = v k^{-1} l$, we have

$$g_{\kappa_0}^{\widehat{\mathfrak{M}}} = k^{-1} l^{-1} \left\{ \widetilde{V} (d\widetilde{r}^2 + \widetilde{r}^2 d\widetilde{\theta}_1^2 + d\widetilde{\theta}_2^2) + \widetilde{V}^{-1} \frac{\widetilde{v}^2}{4\pi^2} (d\widetilde{\theta}_3 - \widetilde{\theta}_2 d\widetilde{\theta}_1)^2 \right\}.$$

A similar calculation shows that

$$\omega_I = k^{-1} l^{-1} \left\{ \widetilde{V} \widetilde{r} d\widetilde{r} \wedge d\widetilde{\theta}_1 + \frac{\widetilde{v}}{2\pi} d\widetilde{\theta}_2 \wedge (d\widetilde{\theta}_3 - \widetilde{\theta}_2 d\widetilde{\theta}_1) \right\}.$$

We also see that

$$\begin{aligned} \omega_J &= k^{-1} l^{-1} (\cos(q) \widetilde{\omega}_J - \sin(q) \widetilde{\omega}_K), \\ \omega_K &= k^{-1} l^{-1} (\sin(q) \widetilde{\omega}_J + \cos(q) \widetilde{\omega}_K), \end{aligned}$$

for $q = -2\pi m l^{-1} v^{-1}$, where $\widetilde{\omega}_J$ and $\widetilde{\omega}_K$ are defined as in (2.1)–(2.2), but with respect to the $(\widetilde{r}, \widetilde{\theta}_1, \widetilde{\theta}_2, \widetilde{\theta}_3)$ -coordinates. This finishes the proof. $\square$

The next proposition deals with the second case above.

Proposition 2.6. *In the second case in Proposition 2.2, the quotient space (2.5) can be identified with*

$$(\widehat{\mathfrak{M}}_{\widetilde{v}}(\widetilde{R}), g_{\widetilde{\kappa}_0, L}^{\widehat{\mathfrak{M}}}, \omega_{1, \widetilde{\kappa}_0, L}^{\widehat{\mathfrak{M}}}, \omega_{2, \widetilde{\kappa}_0, L}^{\widehat{\mathfrak{M}}}, \omega_{3, \widetilde{\kappa}_0, L}^{\widehat{\mathfrak{M}}}) / \iota$$

after a hyper-Kähler rotation, for some $\widetilde{v}$, $\widetilde{R}$, $\widetilde{\kappa}_0$, L , where $\widetilde{v}$ is even and $\iota = \iota_{0,0}$.

Proof. In this case, from Remark 2.3, we have $\Gamma = (\mathbb{Z}_l \times \mathbb{Z}_k) \rtimes \mathbb{Z}_2$. Then $\mathbb{Z}_2$ acts on $\widehat{\mathfrak{M}}_v(R)/G$, where $G = \mathbb{Z}_l \times \mathbb{Z}_k$. Letting $\varphi(\theta_1, \theta_2, \theta_3) = (\theta_1 + 2\pi m l^{-1} v^{-1}, k\theta_2, k\theta_3)$, we see that $\varphi^{-1} \iota_{n, \iota} \varphi = \iota_{\widetilde{n}, \widetilde{\iota}}$ for some $\widetilde{n} \in \mathbb{Z}$ and $\widetilde{\iota} \in \mathbb{R}$. So we can assume that we are in the

second case of Proposition 2.5 with $\tilde{k} = \tilde{l} = 1$ and $\tilde{m} = 0$. Then we simply define

$$\hat{\theta}_2 \equiv \tilde{\theta}_2 - \pi \tilde{n} \tilde{v}^{-1}, \quad \hat{\theta}_3 \equiv \tilde{\theta}_3 - \pi \tilde{n} \tilde{v}^{-1} \tilde{\theta}_1 - \frac{\tilde{t}}{2}$$

to change $\tilde{n}$ and $\tilde{t}$ to 0. $\square$

Remark 2.7. The level sets $\{r = r_0\}$ on $\mathfrak{M}_{2\nu}(R) = \hat{\mathfrak{M}}_{2\nu}(R)/\iota$ are non-orientable line bundles over Klein bottles, and are *infrnilmanifolds*, which are double covered by nilmanifolds in $\hat{\mathfrak{M}}_{2\nu}(R)$.

3. Weighted analysis on the ALG* model space

In this section, we begin our analysis on the model space $(\hat{\mathfrak{M}}_\nu(R), g_{\kappa_0, L}^{\hat{\mathfrak{M}}})$. To simplify notation, we will abbreviate $(\hat{\mathfrak{M}}_\nu(R), g_{\kappa_0, L}^{\hat{\mathfrak{M}}})$ by just $(\hat{\mathfrak{M}}, g^{\hat{\mathfrak{M}}})$. Without loss of generality, by scaling, we can assume that the parameter L is equal to 1 in Definition 1.1. Let us introduce the following notation:

$$\begin{aligned} E &= 1, \quad E^1 = V^{\frac{1}{2}} dx, \quad E^2 = V^{\frac{1}{2}} dy, \quad E^3 = V^{\frac{1}{2}} d\theta_2, \quad E^4 = V^{-\frac{1}{2}} \Theta, \\ E^{12} &= E^1 \wedge E^2, \quad E^{13} = E^1 \wedge E^3, \quad \dots, \quad E^{1234} = E^1 \wedge E^2 \wedge E^3 \wedge E^4. \end{aligned}$$

We compute that

$$\begin{aligned} dE^1 &= \frac{1}{2} V^{-\frac{1}{2}} \frac{v}{2\pi r} dr \wedge dx = \frac{1}{2} V^{-\frac{1}{2}} \frac{v}{2\pi r^2} dy \wedge dx = \frac{1}{2} V^{-\frac{3}{2}} \frac{v}{2\pi r^2} E^2 \wedge E^1, \\ dE^2 &= \frac{1}{2} V^{-\frac{1}{2}} \frac{v}{2\pi r} dr \wedge dy = \frac{1}{2} V^{-\frac{1}{2}} \frac{v}{2\pi r^2} dx \wedge dy = \frac{1}{2} V^{-\frac{3}{2}} \frac{v}{2\pi r^2} E^1 \wedge E^2, \\ dE^3 &= \frac{1}{2} V^{-\frac{1}{2}} \frac{v}{2\pi r} dr \wedge d\theta_2 = \frac{1}{2} V^{-\frac{1}{2}} \frac{v}{2\pi} \frac{x dx + y dy}{r^2} \wedge d\theta_2 \\ &= \frac{1}{2} V^{-\frac{3}{2}} \frac{v}{2\pi} \frac{x E^1 + y E^2}{r^2} \wedge E^3, \\ dE^4 &= -\frac{1}{2} V^{-\frac{3}{2}} \frac{v}{2\pi} \frac{dr}{r} \wedge \Theta + V^{-\frac{1}{2}} \frac{v}{2\pi} d\theta_1 \wedge d\theta_2 \\ &= -\frac{1}{2} V^{-\frac{3}{2}} \frac{v}{2\pi} \frac{x dx + y dy}{r^2} \wedge \Theta + V^{-\frac{1}{2}} \frac{v}{2\pi} \frac{x dy - y dx}{r^2} \wedge d\theta_2 \\ &= -\frac{1}{2} V^{-\frac{3}{2}} \frac{v}{2\pi} \frac{x E^1 + y E^2}{r^2} \wedge E^4 + V^{-\frac{3}{2}} \frac{v}{2\pi} \frac{x E^2 - y E^1}{r^2} \wedge E^3. \end{aligned}$$

By Cartan's structural equations

$$dE^i = -E_j^i \wedge E^j, \quad E_j^i = -E_i^j,$$

we see that $|E_j^i| \leq C \cdot r^{-1} V^{-\frac{3}{2}}$. In other words, $|\nabla_{E_i} E_j| \leq C \cdot r^{-1} V^{-\frac{3}{2}}$, where E_i are the dual orthonormal basis of E^i . It follows that $|\nabla_{E_k} \nabla_{E_i} E_j| \leq C \cdot r^{-2} V^{-2}$, which implies that the curvature of the model space satisfies

$$(3.1) \quad |\text{Rm}| = O(r^{-2} V^{-2}) = O(r^{-2} (\log r)^{-2}),$$

as $r \rightarrow \infty$. For $j \in \mathbb{Z}$ fixed, let

$$(3.2) \quad \omega \equiv \sum_{I \subset \{1,2,3,4\}} e^{\sqrt{-1} \cdot j \cdot \theta_1} \omega_I(r) E^I,$$

where $\omega_I(r)$ are smooth functions in r and where the empty subset corresponds to $E = 1$. Then

$$\begin{aligned} & \left| \nabla^* \nabla \omega - \sum_{I \subset \{1,2,3,4\}} (\nabla^* \nabla (e^{\sqrt{-1} \cdot j \cdot \theta_1} \omega_I)) E^I \right| \\ & \leq C \sum_{I \subset \{1,2,3,4\}} (|\nabla(e^{\sqrt{-1} \cdot j \cdot \theta_1} \omega_I)| \cdot |\nabla E^I| + |\omega_I| \cdot |\nabla^* \nabla E^I|) \end{aligned}$$

for a constant C independent of j , where ∇^* is the L^2 -adjoint of ∇ . By the Weitzenböck formula,

$$|\Delta \omega - \nabla^* \nabla \omega| \leq C |\text{Rm}| |\omega|.$$

Our convention for the Hodge star operator is that $\alpha \wedge * \beta = g(\alpha, \beta) dV_g$, the divergence operator is $\delta = -*d*$, and the Laplacian is the Hodge Laplacian $\Delta = d\delta + \delta d$.

On the other hand, if we define the operator

$$L_{\mathbb{R}^2, j} \omega \equiv \sum_{I \subset \{1,2,3,4\}} e^{\sqrt{-1} \cdot j \cdot \theta_1} (\omega_I'' + r^{-1} \cdot \omega_I' - j^2 \cdot r^{-2} \cdot \omega_I) E^I,$$

then by (A.2),

$$\sum_{I \subset \{1,2,3,4\}} (\nabla^* \nabla (e^{\sqrt{-1} \cdot j \cdot \theta_1} \omega_I)) E^I = -V^{-1} L_{\mathbb{R}^2, j} \omega.$$

In conclusion, there exists a constant C , depending only upon the model space, such that

$$(3.3) \quad |\Delta \omega + V^{-1} L_{\mathbb{R}^2, j} \omega| \leq C \cdot r^{-2} \cdot V^{-2} |\omega| + C \cdot r^{-1} \cdot V^{-\frac{3}{2}} \cdot |\nabla \omega|$$

for any ω of the form (3.2). Estimate (3.3) will be used to carry out the weighted analysis on the ALG* model space $\hat{\mathfrak{M}}$. To start with, we define the weighted norms on the model space.

Definition 3.1 (Weighted Sobolev norms). For any $\mu \in \mathbb{R}$, we define the weight function

$$\varrho_\mu(\mathbf{x}) \equiv \varsigma(\mathbf{x})^{-\mu-1} \quad \text{for all } \mathbf{x} \in \hat{\mathfrak{M}},$$

where $\varsigma(\mathbf{x}) = r(\mathbf{x}) \cdot V(r(\mathbf{x}))^{\frac{1}{2}}$, $\mathbf{x} = \text{pr}_{\mathbb{R}^2}(\mathbf{x})$, $\text{pr}_{\mathbb{R}^2}: \hat{\mathfrak{M}} \rightarrow \mathbb{R}^2$ is the natural projection, and r is the radial distance to the cone vertex of $\mathbb{R}^2$. Then the Sobolev norms are defined as follows:

$$\|\omega\|_{L_\mu^2(\hat{\mathfrak{M}})} \equiv \left(\int_{\hat{\mathfrak{M}}} |\omega \cdot \varrho_\mu|^2 d\text{vol}_{\hat{\mathfrak{M}}} \right)^{\frac{1}{2}}, \quad \|\omega\|_{W_\mu^{k,2}(\hat{\mathfrak{M}})} \equiv \left(\sum_{m=0}^k \|\nabla^m \omega\|_{L_{\mu-m}^2(\hat{\mathfrak{M}})}^2 \right)^{\frac{1}{2}},$$

where ω is a tensor field on $\hat{\mathfrak{M}}$.

We will require the following weighted Sobolev estimates, which hold for tensor fields of any type.

Proposition 3.2. For any $\mu \in \mathbb{R}$ and $k \in \mathbb{N}_0$, there exists a constant

$$C = C(\kappa_0, \nu, \mu, k) > 0$$

such that, for any $\omega \in W_\mu^{k+2,2}(\hat{\mathfrak{M}}_\nu(R))$,

$$(3.4) \quad \|\omega\|_{W_\mu^{k+2,2}(\hat{\mathfrak{M}}_\nu(2R))} \leq C(\|\omega\|_{L_\mu^2(\hat{\mathfrak{M}}_\nu(R))} + \|\Delta \omega\|_{W_{\mu-2}^{k,2}(\hat{\mathfrak{M}}_\nu(R))}).$$

Proof. The argument is standard, so we will be brief; see for example [8, Proposition 6.16]. For $x \in \widehat{\mathfrak{M}}_\nu(2R)$, we consider the rescaled metric $\tilde{g} \equiv 100 \cdot d^{-2} \cdot g^{\widehat{\mathfrak{M}}}$, where d is the $g^{\widehat{\mathfrak{M}}}$ -distance between x and $\{r = R\}$. It is straightforward to check that $|\text{Rm}_{\tilde{g}}| \leq C_0$ on $B_2^{\tilde{g}}(x)$ for some constant $C_0 > 0$ independent of $x \in \widehat{\mathfrak{M}}_\nu(2R)$; see (3.1). The standard elliptic estimate is

$$\|\omega\|_{W^{k+2,2}(B_{1/2}^{\tilde{g}}(x))} \leq C \cdot \|\omega\|_{L^2(B_1^{\tilde{g}}(x))} + C \cdot \|\Delta\omega\|_{W^{k,2}(B_1^{\tilde{g}}(x))}.$$

Rescaling back to $g^{\widehat{\mathfrak{M}}}$ and using a simple covering argument, (3.4) follows. $\square$

Proposition 3.3. *For any $\mu \in \mathbb{R}$ and $k \in \mathbb{N}_0$, there exists a constant*

$$C = C(\kappa_0, \nu, \mu, k) > 0$$

such that, for any $\omega \in W_\mu^{k+3,2}(\widehat{\mathfrak{M}}_\nu(R))$,

$$\sum_{m=0}^k \sup_{x \in \widehat{\mathfrak{M}}_\nu(2R)} |(\varkappa(x))^{m-\mu} \nabla^m \omega(x)| \leq C \|\omega\|_{W_\mu^{k+3,2}(\widehat{\mathfrak{M}}_\nu(R))}.$$

Proof. The argument is similar to the proof of Proposition 3.2 and is omitted. $\square$

The following estimate is key to our weighted analysis.

Proposition 3.4. *Given $\kappa_0 \in \mathbb{R}$, $\nu \geq 1$ and $\mu \in \mathbb{R} \setminus \mathbb{Z}$, there exists a constant $R > 0$ depending only on κ_0 , ν and μ such that the following property holds. If ω is a smooth form compactly supported on a subset of $\{r > R\}$, then for any $k \in \mathbb{N}_0$, there exists a constant $C = C(\kappa_0, \nu, \mu, k) > 0$ such that*

$$(3.5) \quad \|\omega\|_{W_\mu^{k+2,2}(\widehat{\mathfrak{M}})} \leq C \cdot \|\Delta\omega\|_{W_{\mu-2}^{k,2}(\widehat{\mathfrak{M}})}.$$

Proof. We can decompose $\omega = \sum_{I \subset \{1,2,3,4\}} \omega_I E^I$ into two types. For the first type, ω_I depends only on r and θ_1 . For the second type,

$$\int_{T^2} \omega_I(r, \theta_1, \theta_2, \theta_3) d\theta_2 d\theta_3 = 0.$$

It is easy to see that the Hodge Laplacian Δ preserves this decomposition. So we only need to prove (3.5) in two steps.

Step 1. We will consider the case when ω_I depends only on r and θ_1 . In this case, we can write the coefficient function ω_I in terms of the Fourier expansion as follows:

$$(3.6) \quad \omega = \sum_{I \subset \{1,2,3,4\}} \sum_{j=-\infty}^{\infty} e^{\sqrt{-1} \cdot j \cdot \theta_1} \omega_{I,j}(r) E^I.$$

Let us start with the proof of the weighted L^2 -estimate

$$(3.7) \quad \|\omega\|_{L_\mu^2(\widehat{\mathfrak{M}})} \leq C \cdot \|\Delta\omega\|_{L_{\mu-2}^2(\widehat{\mathfrak{M}})}.$$

Since Δ preserves the decomposition in (3.6), and the subspaces for different j are L_μ^2 -orthogonal, we can assume that

$$(3.8) \quad \omega = \sum_{I \subset \{1,2,3,4\}} e^{\sqrt{-1} \cdot j \cdot \theta_1} \omega_I(r) E^I$$

for some fixed $j \in \mathbb{Z}$. Then we need to prove (3.7) for a constant C independent of j . The proof of (3.7) will be reduced to computations in $\mathbb{R}^2$. We will need the following two claims.

Claim 1. *Let $\alpha \neq -1$, $\beta \in \mathbb{R}$ and let $R > 0$ satisfy*

$$\left| \beta \cdot \frac{\nu}{2\pi} \cdot \left(\kappa_0 + \frac{\nu}{2\pi} \log R \right)^{-1} \right| \leq \frac{|\alpha + 1|}{2}.$$

Then, for any $f \in C_0^\infty(\{r \geq R\})$,

$$\int_R^\infty |f|^2 r^\alpha V^\beta dr \leq \frac{16}{(\alpha + 1)^2} \int_R^\infty |f'|^2 r^{\alpha+2} V^\beta dr.$$

Proof. It is straightforward that if R is chosen such that

$$\left| \beta \cdot \frac{\nu}{2\pi} \cdot \left(\kappa_0 + \frac{\nu}{2\pi} \log R \right)^{-1} \right| \leq \frac{|\alpha + 1|}{2},$$

then for any $r \geq R$,

$$\left| \frac{d}{dr} (r^{\alpha+1} V^\beta) \right| = r^\alpha V^\beta \cdot \left| \alpha + 1 + \beta \cdot \frac{\nu}{2\pi} \cdot V^{-1} \right| \geq \frac{|\alpha + 1|}{2} \cdot r^\alpha V^\beta.$$

So it follows that

$$\begin{aligned} \int_R^\infty |f|^2 r^\alpha V^\beta dr &\leq \frac{2}{|\alpha + 1|} \left| \int_R^\infty |f|^2 d(r^{\alpha+1} V^\beta) \right| \\ &= \frac{4}{|\alpha + 1|} \left| \int_R^\infty \operatorname{Re}(f \cdot \bar{f}') \cdot r^{\alpha+1} V^\beta dr \right| \\ &\leq \frac{4}{|\alpha + 1|} \left(\int_R^\infty |f|^2 r^\alpha V^\beta dr \right)^{\frac{1}{2}} \left(\int_R^\infty |f'|^2 r^{\alpha+2} V^\beta dr \right)^{\frac{1}{2}}. \end{aligned}$$

Then the desired inequality immediately follows. $\square$

Claim 2. *Let ω be defined in (3.8). Then, for $\mu \in \mathbb{R} \setminus \mathbb{Z}$ and $R > 0$ satisfying*

$$\left| -\mu \cdot \frac{\nu}{\pi} \cdot \left(\kappa_0 + \frac{\nu}{2\pi} \log R \right)^{-1} \right| \leq \min\{|j + \mu|, |j - \mu|\},$$

we have

$$\|\omega\|_{L_\mu^2(\widehat{\mathfrak{M}})}^2 \leq 16 \cdot (j + \mu)^{-2} \cdot (j - \mu)^{-2} \cdot \|r^2 \cdot L_{\mathbb{R}^2, j} \omega\|_{L_\mu^2(\widehat{\mathfrak{M}})}^2.$$

Proof. Recall that $\operatorname{dvol}_{\widehat{\mathfrak{M}}} = V \cdot r \cdot dr \wedge d\theta_1 \wedge d\theta_2 \wedge \Theta$, and thus

$$\|\omega\|_{L_\mu^2(\widehat{\mathfrak{M}})}^2 = 8\pi^3 \sum_{I \subset \{1,2,3,4\}} \int_R^\infty |\omega_I|^2 r^{-2\mu-1} V^{-\mu} dr.$$

Then let us estimate the weighted L^2 -integral of $r^2 \cdot L_{\mathbb{R}^2, j} \omega$. By definition,

$$\begin{aligned} \|r^2 L_{\mathbb{R}^2, j} \omega\|_{L_\mu^2(\hat{\mathfrak{M}})}^2 &= 8\pi^3 \sum_{I \subset \{1,2,3,4\}} \int_R^\infty |\omega_I'' + r^{-1} \cdot \omega_I' - j^2 \cdot r^{-2} \cdot \omega_I|^2 r^{-2\mu+3} V^{-\mu} dr \\ &= 8\pi^3 \sum_{I \subset \{1,2,3,4\}} \int_R^\infty |(r^{2j+1}(r^{-j} \omega_I'))'|^2 r^{-2\mu-2j+1} V^{-\mu} dr. \end{aligned}$$

Since ω has compact support in $\hat{\mathfrak{M}}$, applying Claim 1 to above integral,

$$\begin{aligned} &\|r^2 L_{\mathbb{R}^2, j} \omega\|_{L_\mu^2(\hat{\mathfrak{M}})}^2 \\ &\geq 8\pi^3 \cdot \frac{(-2\mu - 2j)^2}{16} \sum_{I \subset \{1,2,3,4\}} \int_R^\infty |r^{2j+1}(r^{-j} \omega_I')|^2 r^{-2\mu-2j-1} V^{-\mu} dr \\ &= 2\pi^3 (j + \mu)^2 \sum_{I \subset \{1,2,3,4\}} \int_R^\infty |(r^{-j} \omega_I')|^2 r^{-2\mu+2j+1} V^{-\mu} dr \\ &\geq 2\pi^3 (j + \mu)^2 \cdot \frac{(-2\mu + 2j)^2}{16} \sum_{I \subset \{1,2,3,4\}} \int_R^\infty |r^{-j} \omega_I|^2 r^{-2\mu+2j-1} V^{-\mu} dr \\ &= \frac{\pi^3}{2} (j + \mu)^2 (j - \mu)^2 \sum_{I \subset \{1,2,3,4\}} \int_R^\infty |\omega_I|^2 r^{-2\mu-1} V^{-\mu} dr \\ &= \frac{(j + \mu)^2 (j - \mu)^2}{16} \|\omega\|_{L_\mu^2(\hat{\mathfrak{M}})}^2. \end{aligned}$$

The proof of the claim is done. $\square$

Now we are ready to finish the proof of (3.7). Recall that, by (3.3),

$$\begin{aligned} |r^2 V \Delta \omega + r^2 L_{\mathbb{R}^2, j} \omega| &\leq C \cdot V^{-1} |\omega| + C \cdot r \cdot V^{-\frac{1}{2}} \cdot |\nabla \omega| \\ &\leq C \cdot V^{-1} (|\omega| + r \cdot V^{\frac{1}{2}} \cdot |\nabla \omega|) \end{aligned}$$

for $r \geq R$, where $C > 0$ is a constant independent of j . By Claim 2,

$$\|\omega\|_{L_\mu^2(\hat{\mathfrak{M}})} \leq C \cdot \|r^2 L_{\mathbb{R}^2, j} \omega\|_{L_\mu^2(\hat{\mathfrak{M}})}.$$

Consequently,

$$\begin{aligned} (3.9) \quad \|\omega\|_{L_\mu^2(\hat{\mathfrak{M}})} &\leq C \cdot \|\mathfrak{s}^2 \cdot \Delta \omega\|_{L_\mu^2(\hat{\mathfrak{M}})} + C \cdot V(R)^{-1} \|\omega\|_{W_\mu^{1,2}(\hat{\mathfrak{M}})} \\ &= C \cdot \|\Delta \omega\|_{L_{\mu-2}^2(\hat{\mathfrak{M}})} + C \cdot V(R)^{-1} \|\omega\|_{W_\mu^{1,2}(\hat{\mathfrak{M}})} \\ &\leq C \cdot \|\Delta \omega\|_{L_{\mu-2}^2(\hat{\mathfrak{M}})} + C \cdot V(R)^{-1} \|\omega\|_{W_\mu^{2,2}(\hat{\mathfrak{M}})}. \end{aligned}$$

If R is chosen sufficiently large, (3.7) follows by plugging (3.4) for $k = 0$ into (3.9).

Estimate (3.5) follows immediately from (3.7) and (3.4).

Step 2. We will consider the case when $\omega = \sum_{I \subset \{1,2,3,4\}} \omega_I(r, \theta_1, \theta_2, \theta_3) \cdot E^I$ satisfies

$$(3.10) \quad \int_{T^2} \omega_I(r, \theta_1, \theta_2, \theta_3) d\theta_2 d\theta_3 = 0.$$

We will prove that there exists $C > 0$ such that if ω satisfies (3.10), then

$$(3.11) \quad \|\omega\|_{W_{\mu}^{2,2}(\widehat{\mathfrak{M}})} \leq C \cdot \|\Delta\omega\|_{L_{\mu-2}^2(\widehat{\mathfrak{M}})}.$$

The higher-order estimate (3.7) then follows from Proposition 3.2.

First, we will show that there is a constant $C > 0$ such that if the coefficient function of ω satisfies (3.10), then for any sufficiently large R and for any $r \in [R, \infty)$,

$$(3.12) \quad \int_{\mathcal{A}_{r/2,2r}} |\omega|^2 d\text{vol}_{\widehat{\mathfrak{M}}} \leq C \cdot V(r) \cdot \int_{\mathcal{A}_{r/2,2r}} |\nabla\omega|^2 d\text{vol}_{\widehat{\mathfrak{M}}},$$

where $\mathcal{A}_{r/2,2r} \equiv \text{pr}_{\mathbb{R}^2}^{-1}(A_{r/2,2r}(0^2))$. There are two sub-cases to analyze:

- (1) ω_I satisfies $\int_{S^1} \omega_I(r, \theta_1, \theta_2, \theta_3) d\theta_3 = 0$ for all r, θ_1 , and θ_2 ,
- (2) $\omega_I = \omega_I(r, \theta_1, \theta_2)$ and it satisfies $\int_{S^1} \omega_I(r, \theta_1, \theta_2) d\theta_2 = 0$ for all r and θ_1 .

In the first case, applying the Poincaré inequality on the circle parametrized by θ_3 , we find that, for any $(r, \theta_1, \theta_2) \in [R, \infty) \times [0, \pi] \times [0, 2\pi]$,

$$\int_{S^1} |\omega_I|^2 d\theta_3 \leq C \int_{S^1} |\partial_{\theta_3} \omega_I|^2 d\theta_3 \leq \frac{C}{V(r)} \int_{S^1} |\nabla\omega|^2(r, \theta_1, \theta_2, \theta_3) d\theta_3.$$

In the second case, applying the Poincaré inequality on the circle parametrized by θ_2 , it follows that, for any $(r, \theta_1) \in [R, \infty) \times [0, \pi]$,

$$\int_{S^1} |\omega_I|^2 d\theta_2 \leq C \int_{S^1} |\partial_{\theta_2} \omega_I|^2 d\theta_2 \leq C \cdot V(r) \cdot \int_{S^1} |\nabla\omega|^2(r, \theta_1, \theta_2) d\theta_2.$$

Combining the above two cases and integrating over $\widehat{\mathfrak{M}}$, inequality (3.12) immediately follows from the condition $R \gg 1$.

Now we proceed to prove (3.11). Let $r_i = 2^{i+1} \cdot R$. Then $\{A_{r_i/2,2r_i}(0^2)\}_{i=0}^{\infty}$ is a covering of $\mathbb{R}^2 \setminus B_R(0^2)$ consisting of a sequence of annuli such that the number of overlaps at every point is bounded by 2. We denote $\mathcal{A}_i \equiv \text{pr}_{\mathbb{R}^2}^{-1}(A_{r_i/2,2r_i}(0^2))$ and $\mathfrak{s}_i \equiv \mathfrak{s}(r_i)$. Let $\{\chi_i\}_{i=0}^{\infty}$ be a partition of unity subordinate to the above covering such that $\text{Supp}(\chi_i) \subset A_{r_i/2,2r_i}(0^2)$ and

$$|\nabla_{\mathbb{R}^2} \chi_i|^2 + |\nabla_{\mathbb{R}^2}^2 \chi_i| \leq C \cdot r_i^{-2}$$

on $\mathbb{R}^2$. We still denote by χ_i the lifting of χ_i to $\widehat{\mathfrak{M}}$. In terms of the model metric $g^{\widehat{\mathfrak{M}}}$, the following estimate holds on $\widehat{\mathfrak{M}}$:

$$|\nabla \chi_i|^2 + |\nabla^2 \chi_i| \leq C \cdot \mathfrak{s}_i^{-2},$$

where $C > 0$ is independent of i . By (3.12),

$$\int_{\widehat{\mathfrak{M}}} |\chi_i \omega|^2 \leq C \cdot V_i \cdot \int_{\widehat{\mathfrak{M}}} |\nabla(\chi_i \omega)|^2,$$

where $V_i \equiv V(r_i)$. Since $\chi_i \omega$ has compact support in $\widehat{\mathfrak{M}}$, integrating by parts and applying the Cauchy–Schwarz inequality, we have that

$$\begin{aligned} \int_{\widehat{\mathfrak{M}}} |\chi_i \omega|^2 d\text{vol}_{\widehat{\mathfrak{M}}} &\leq C \cdot V_i \cdot \int_{\widehat{\mathfrak{M}}} (\chi_i \omega) \cdot \nabla^* \nabla (\chi_i \omega) d\text{vol}_{\widehat{\mathfrak{M}}} \\ &\leq C \cdot V_i \cdot \left(\int_{\widehat{\mathfrak{M}}} |\chi_i \omega|^2 d\text{vol}_{\widehat{\mathfrak{M}}} \right)^{\frac{1}{2}} \left(\int_{\widehat{\mathfrak{M}}} |\nabla^* \nabla (\chi_i \omega)|^2 d\text{vol}_{\widehat{\mathfrak{M}}} \right)^{\frac{1}{2}}. \end{aligned}$$

So it follows that

$$\int_{\hat{\mathfrak{M}}} |\chi_i \omega|^2 \, \text{dvol}_{\hat{\mathfrak{M}}} \leq C \cdot V_i^2 \cdot \int_{\hat{\mathfrak{M}}} |\nabla^* \nabla(\chi_i \omega)|^2 \, \text{dvol}_{\hat{\mathfrak{M}}}.$$

Using the Weitzenböck formula, we have that

$$\begin{aligned} |\nabla^* \nabla(\chi_i \omega)| &\leq |\Delta(\chi_i \omega)| + C \cdot |\text{Rm}| \cdot |\chi_i \cdot \omega| \\ &\leq |\Delta(\chi_i \omega)| + C \cdot \mathfrak{s}_i^{-2} \cdot V^{-1} \cdot |\omega| \\ &\leq |\Delta \omega| + C \cdot \mathfrak{s}_i^{-1} \cdot |\nabla \omega| + C \cdot \mathfrak{s}_i^{-2} \cdot |\omega|. \end{aligned}$$

Therefore,

$$\begin{aligned} \int_{\hat{\mathfrak{M}}} |\chi_i \omega|^2 \, \text{dvol}_{\hat{\mathfrak{M}}} &\leq C \cdot V_i^2 \int_{\hat{\mathfrak{M}}} |\nabla^* \nabla(\chi_i \omega)|^2 \, \text{dvol}_{\hat{\mathfrak{M}}} \\ &\leq \frac{C}{r_i^4} \int_{\mathcal{A}_i} (|\Delta \omega|^2 \cdot \mathfrak{s}_i^4 + |\nabla \omega|^2 \cdot \mathfrak{s}_i^2 + |\omega|^2) \, \text{dvol}_{\hat{\mathfrak{M}}}. \end{aligned}$$

Taking the weighted L^2 -norm, we find that

$$\begin{aligned} (3.13) \quad \|\omega\|_{L_\mu^2(\hat{\mathfrak{M}})}^2 &= \int_{\hat{\mathfrak{M}}} |\omega \cdot \varrho_\mu|^2 \, \text{dvol}_{\hat{\mathfrak{M}}} \\ &\leq C \cdot \sum_i \mathfrak{s}_i^{-2\mu-2} \cdot \int_{\hat{\mathfrak{M}}} |\chi_i \omega|^2 \, \text{dvol}_{\hat{\mathfrak{M}}} \\ &\leq C \cdot \sum_i \frac{1}{r_i^4} \int_{\mathcal{A}_i} (|(\Delta \omega) \cdot \varrho_{\mu-2}|^2 \\ &\quad + |(\nabla \omega) \cdot \varrho_{\mu-1}|^2 + |\omega \cdot \varrho_\mu|^2) \, \text{dvol}_{\hat{\mathfrak{M}}} \\ &\leq \frac{C}{R^4} \cdot (\|\Delta \omega\|_{L_{\mu-2}^2(\hat{\mathfrak{M}})}^2 + \|\nabla \omega\|_{L_{\mu-1}^2(\hat{\mathfrak{M}})}^2 + \|\omega\|_{L_\mu^2(\hat{\mathfrak{M}})}^2), \end{aligned}$$

where the last inequality follows since the number of overlaps is bounded by 2. It follows that

$$(3.14) \quad \|\omega\|_{L_\mu^2(\hat{\mathfrak{M}})} \leq C \cdot \|\Delta \omega\|_{L_{\mu-2}^2(\hat{\mathfrak{M}})} + \frac{C}{R^2} (\|\nabla \omega\|_{L_{\mu-1}^2(\hat{\mathfrak{M}})} + \|\omega\|_{L_\mu^2(\hat{\mathfrak{M}})}).$$

If R is chosen sufficiently large, then (3.11) follows from (3.14) and Proposition 3.2.

To finish the proof, we can write any form $\omega = \omega_1 + \omega_2$ as the sum of these two types of forms. We then have

$$\begin{aligned} (3.15) \quad \|\omega\|_{W_\mu^{2,2}(\hat{\mathfrak{M}})} &\leq \|\omega_1\|_{W_\mu^{2,2}(\hat{\mathfrak{M}})} + \|\omega_2\|_{W_\mu^{2,2}(\hat{\mathfrak{M}})} \\ &\leq C (\|\Delta \omega_1\|_{L_{\mu-2}^2(\hat{\mathfrak{M}})} + \|\Delta \omega_2\|_{L_{\mu-2}^2(\hat{\mathfrak{M}})}) \\ &= C \|\Delta \omega\|_{L_{\mu-2}^2(\hat{\mathfrak{M}})}, \end{aligned}$$

where the last equality follows since the two subspaces are orthogonal in $L_{\mu-2}^2(\hat{\mathfrak{M}})$. The higher-order estimates then follow from (3.15) and Proposition 3.2. $\square$

Corollary 3.5. *Given any $\mu \in \mathbb{R} \setminus \mathbb{Z}$, there exists a constant $\epsilon_0 > 0$ such that the following property holds. Let $\omega = \sum_{I \subset \{1,2,3,4\}} \omega_I E^I$ be a smooth form on $\hat{\mathfrak{M}}$ that satisfies $\Delta \omega = 0$ and $\|\omega\|_{W_\mu^{2,2}(\hat{\mathfrak{M}})} < \infty$. If*

$$\int_{T^2} \omega_I(r, \theta_1, \theta_2, \theta_3) \, d\theta_2 \, d\theta_3 = 0$$

for every $I \subset \{1, 2, 3, 4\}$, then $|\omega| = O(e^{-\epsilon_0 r})$ as $r \rightarrow \infty$.

Proof. Let R_0 be the constant R in Proposition 3.4 depending on κ_0 , ν and μ . If η is a smooth form on $\widehat{\mathfrak{M}}$ with compact support on a subset of $\{r > R_0\}$, (3.11) and (3.13) imply that

$$(3.16) \quad \|\eta\|_{L^2_\mu(\widehat{\mathfrak{M}})}^2 \leq \frac{C}{R^4} \cdot \|\Delta\eta\|_{L^2_{\mu-2}(\widehat{\mathfrak{M}})}^2.$$

For any $R > R_0$, let φ_R be a cut-off function on $\mathbb{R}^2$ such that

$$\varphi_R = \begin{cases} 1, & r \geq R+2, \\ 0, & r \leq R+1, \end{cases}$$

and $|\nabla\varphi_R|_{\mathbb{R}^2} + |\nabla^2\varphi_R|_{\mathbb{R}^2} \leq C$. We still denote by φ_R the lifting of φ_R to the model space $\widehat{\mathfrak{M}}$. Then

$$|\nabla\varphi_R|_{\widehat{\mathfrak{M}}}^2 + |\nabla^2\varphi_R|_{\widehat{\mathfrak{M}}} \leq C \cdot V^{-1}.$$

Since $\varphi_R\omega$ can be approximated by compactly supported smooth forms in the $W_\mu^{k,2}$ -norm, (3.16) also holds for $\varphi_R\omega$. Using (3.16), we estimate

$$\begin{aligned} \int_{r \geq R+3} |\omega \cdot \varrho_\mu|^2 \, \text{dvol}_{\widehat{\mathfrak{M}}} &\leq \frac{C}{R^4} \cdot \int_{r \geq R+1} |\Delta(\varphi_R\omega) \cdot \varrho_{\mu-2}|^2 \, \text{dvol}_{\widehat{\mathfrak{M}}} \\ &\leq CR^{-2\mu-2} V^{-\mu+1} \cdot \int_{R+1 \leq r \leq R+2} |\Delta(\varphi_R\omega)|^2 \, \text{dvol}_{\widehat{\mathfrak{M}}} \\ &\leq CR^{-2\mu-2} V^{-\mu-1} \cdot \int_{R+1 \leq r \leq R+2} (|\omega|^2 + |\nabla\omega|^2 \cdot V) \, \text{dvol}_{\widehat{\mathfrak{M}}} \\ &\leq CR^{-2\mu-2} V^{-\mu-1} \cdot \int_{R \leq r \leq R+3} |\omega|^2 \, \text{dvol}_{\widehat{\mathfrak{M}}} \\ &\leq C_0 \cdot \int_{R \leq r < R+3} |\omega|^2 \cdot \varrho_\mu^2 \, \text{dvol}_{\widehat{\mathfrak{M}}}. \end{aligned}$$

Notice that the fourth inequality follows from Proposition 3.2. Therefore,

$$\int_{r \geq R+3} |\omega \cdot \varrho_\mu|^2 \, \text{dvol}_{\widehat{\mathfrak{M}}} \leq \frac{C_0}{C_0 + 1} \cdot \int_{r \geq R} |\omega \cdot \varrho_\mu|^2 \, \text{dvol}_{\widehat{\mathfrak{M}}},$$

which implies that $\int_{r \geq R} |\omega \cdot \varrho_\mu|^2 \, \text{dvol}_{\widehat{\mathfrak{M}}} = O(e^{-\epsilon_0 \cdot R})$ for some $\epsilon_0 > 0$ as $R \rightarrow \infty$. Therefore, ω also decays exponentially by Proposition 3.2 and Proposition 3.3. $\square$

4. Weighted analysis on ALG^* manifolds

We next transfer the previous estimates on the model space to any ALG^* manifold. Without loss of generality, by scaling, we can assume that the parameter L is equal to 1 in Definition 1.2.

Definition 4.1 (Weighted Sobolev norms). Let (X, g) be an ALG_ν^* - Γ manifold for $\nu \in \mathbb{Z}_+$ together with a diffeomorphism $\Phi: \widehat{\mathfrak{M}}_\nu(R)/\Gamma \rightarrow X \setminus X_R$. For any fixed parameters $\kappa_0 \in \mathbb{R}$ and $\mu \in \mathbb{R}$, we define the weight function $\widehat{\varrho}_\mu$ on X ,

$$\widehat{\varrho}_\mu(\mathbf{x}) \equiv \begin{cases} 1, & r \leq 2R, \\ \varrho_\mu(\Phi^{-1}(\mathbf{x})), & r \geq 3R, \end{cases}$$

where R is the radius as in Definition 1.2, and ϱ_μ is the weight function on $\mathfrak{M} = \widehat{\mathfrak{M}}/\Gamma$ as in Definition 3.1. Then the weighted Sobolev norms are defined as follows:

$$\|\omega\|_{L^2_\mu(X)} \equiv \left(\int_X |\omega \cdot \widehat{\varrho}_\mu|^2 \, \text{dvol}_X \right)^{\frac{1}{2}}, \quad \|\omega\|_{W^{k,2}_\mu(X)} \equiv \left(\sum_{m=0}^k \|\nabla^m \omega\|_{L^2_{\mu-m}(X)}^2 \right)^{\frac{1}{2}}.$$

Remark 4.2. Notice that Proposition 3.2 and Proposition 3.3 are stated on the ALG* model space $\widehat{\mathfrak{M}}$. Since the weighted norms on an ALG* gravitational instanton X and its model space $\widehat{\mathfrak{M}}$ differ by a uniform multiplicative constant, in our applications, we may also quote these two propositions for ALG* manifolds.

We need the following notation on an ALG* manifold (X, g) and its asymptotic model $(\mathfrak{M}, g^\mathfrak{M})$.

Definition 4.3. Let $\mathcal{H}^p_\mu(X)$ be the space of all smooth p -forms ω on (X, g) that satisfy $\Delta_X \omega = 0$ and $\|\omega\|_{L^2_\mu(X)} < \infty$, where Δ_X is the Hodge Laplacian on (X, g) .

Definition 4.4. Given $p \in \{0, 1, 2\}$ and $\mathfrak{q} \in \mathbb{Z}$, let $\mathcal{Z}^p_{\mathfrak{q}}(\widehat{\mathfrak{M}})$ be the linear space of p -forms with a basis $\{u_{\mathfrak{q},i} \cdot e^{\sqrt{-1} \cdot m_{\mathfrak{q},i} \cdot \theta_1} : 1 \leq i \leq \dim \mathcal{Z}^p_{\mathfrak{q}}(\widehat{\mathfrak{M}})\}$, where for each i ,

$$u_{\mathfrak{q},i} = \sum_I \omega_{I,i}(r) E^I,$$

$\omega_{I,i}(r)$ is a radial function, and $I \subset \{1, 2, 3, 4\}$ satisfies $|I| = p$. For any $p \in \{0, 1, 2\}$ and $\mathfrak{q} \in \mathbb{Z}$, the linear space $\mathcal{Z}^p_{\mathfrak{q}}(\widehat{\mathfrak{M}})$ is characterized as follows.

- When $p = 0$, the basis is defined in Lemma A.1
- When $p = 1$, the basis of $\mathcal{Z}^1_{\mathfrak{q}}(\widehat{\mathfrak{M}}) \equiv \mathcal{Z}^{1,I}_{\mathfrak{q}}(\widehat{\mathfrak{M}}) \oplus \mathcal{Z}^{1,\Pi}_{\mathfrak{q}}(\widehat{\mathfrak{M}})$ is defined in Lemma A.2 and Lemma A.4
- When $p = 2$, the basis of $\mathcal{Z}^2_{\mathfrak{q}}(\widehat{\mathfrak{M}}) \equiv \mathcal{Z}^{2,+}_{\mathfrak{q}}(\widehat{\mathfrak{M}}) \oplus \mathcal{Z}^{2,-}_{\mathfrak{q}}(\widehat{\mathfrak{M}})$ is defined in (A.8) and Lemma A.8.

Our main result in this section is the following Fredholm package for the Hodge Laplacian on ALG* manifolds with respect to the weighted Sobolev norms.

Proposition 4.5. *Let (X, g) be an $\text{ALG}^*_v\text{-}\Gamma$ manifold of order $\mathfrak{n} > 0$. Then the following properties hold.*

- (1) *For any $\mu \in \mathbb{R} \setminus \mathbb{Z}$ and $k \in \mathbb{N}_0$, there exist constants $R(X, \mu) > 0$ and $C(X, \mu, k) > 0$ such that, for any p -form $\omega \in W^{k+2,2}_\mu(X)$,*

$$(4.1) \quad \|\omega\|_{W^{k+2,2}_\mu(X)} \leq C \cdot (\|\Delta_X \omega\|_{W^{k,2}_{\mu-2}(X)} + \|\omega\|_{L^2(\{r \leq 3R\} \subset X)}).$$

- (2) *For any $\mu \in \mathbb{R} \setminus \mathbb{Z}$ and $k \in \mathbb{N}_0$, the operator*

$$\Delta_X: W^{k+2,2}_\mu(X) \rightarrow W^{k,2}_{\mu-2}(X)$$

is a Fredholm operator. Then, for any p -form $\omega \in W_{\mu-2}^{k,2}(X)$, $\Delta_X \xi = \omega$ has a solution $\xi \in W_{\mu}^{k+2,2}(X)$ if and only if, for all $\eta \in \mathcal{H}_{-\mu}^p(X)$,

$$\int_X (\omega, \eta) \, \text{dvol}_g = 0.$$

- (3) For any $\mu \in \mathbb{R} \setminus \mathbb{Z}$, $k \in \mathbb{N}_0$, and p -form $\omega \in W_{\mu}^{k,2}(X)$, there exists some $\xi \in W_{\mu+2}^{k+2,2}(X)$ such that $\Delta_X \xi = \omega$ when $r \geq 2R$.
- (4) Let $\mu \in \mathbb{R} \setminus \mathbb{Z}$, $\delta \in (0, \min\{1, \mathfrak{n}\})$. Consider any form $\omega \in W_{\mu}^{k,2}(X)$ such that $\Delta_X \omega = 0$ when $r \geq 2R$. If $\mathbb{Z} \cap [\mu - \delta, \mu] = \emptyset$, then we have $\omega \in W_{\mu-\delta}^{k,2}(X)$. If there is some $\mathfrak{q} \in \mathbb{Z} \cap (\mu - \delta, \mu)$, then the pull-back of ω to $\widehat{\mathfrak{M}}$ can be written as the sum of an element in $\mathcal{Z}_{\mathfrak{q}}^p(\widehat{\mathfrak{M}})$ and an element in $W_{\mu-\delta}^{k,2}(\widehat{\mathfrak{M}})$.

Proof. Recall that, by definition, the ALG_v^* - Γ manifold (X, g) is equipped with a diffeomorphism $\Phi: \widehat{\mathfrak{M}}_v(R)/\Gamma \rightarrow X \setminus X_R$ such that $|\Phi^*g - g^{\widehat{\mathfrak{M}}}| = O(\mathfrak{s}^{-\mathfrak{n}})$. For the simplicity of notation, we do not distinguish the function $\mathfrak{s}$ defined on $\widehat{\mathfrak{M}}_v(R)$ with the function $\mathfrak{s} \circ \Phi^{-1}$ defined on $X \setminus X_R$.

For (1), for any p -form ω with compact support in $\{r > R\} \subset \mathfrak{M}$, we have

$$(1 + O(\mathfrak{s}(R)^{-\mathfrak{n}}))^{-1} \|\omega\|_{W_{\mu}^{k+2,2}(\mathfrak{M})} \leq \|\omega\|_{W_{\mu}^{k+2,2}(X)} \leq (1 + O(\mathfrak{s}(R)^{-\mathfrak{n}})) \|\omega\|_{W_{\mu}^{k+2,2}(\mathfrak{M})}.$$

Then we will not distinguish $W_{\mu}^{k+2,2}(\mathfrak{M})$ with $W_{\mu}^{k+2,2}(X)$. Moreover,

$$\|\Delta_X \omega\|_{W_{\mu-2}^{k,2}(X)} - \|\Delta_{\mathfrak{M}} \omega\|_{W_{\mu-2}^{k,2}(X)} \leq C \cdot \mathfrak{s}(R)^{-\mathfrak{n}} \cdot \|\omega\|_{W_{\mu}^{k+2,2}(X)}.$$

We also identify a form on $\mathfrak{M}$ with a Γ -invariant form on $\widehat{\mathfrak{M}}$. By (3.5), for R sufficiently large, we have

$$\|\omega\|_{W_{\mu}^{k+2,2}} \leq C \|\Delta_X \omega\|_{W_{\mu-2}^{k,2}}.$$

Let χ be a cut-off function on $\mathbb{R}^2$ with $\text{Supp}(\chi) \subset B_{2R}$ and $\chi \equiv 1$ in B_R . If $\omega \in W_{\mu}^{k+2,2}(X)$, then $(1 - \chi)\omega$ can be approximated by smooth forms with compact support in $\{r > R\}$. Therefore,

$$\begin{aligned} \|\omega\|_{W_{\mu}^{k+2,2}(X)} &\leq \|\chi\omega\|_{W_{\mu}^{k+2,2}(X)} + \|(1 - \chi)\omega\|_{W_{\mu}^{k+2,2}(X)} \\ &\leq C(\|\omega\|_{W_{\mu}^{k+2,2}(\{r \leq 2R\} \subset X)} + \|\Delta_X((1 - \chi)\omega)\|_{W_{\mu-2}^{k,2}(X)}) \\ &\leq C(\|\omega\|_{L^2(\{r \leq 3R\} \subset X)} + \|\Delta_X \omega\|_{W_{\mu-2}^{k,2}(\{r \leq 3R\} \subset X)} + \|\Delta_X \omega\|_{W_{\mu-2}^{k,2}(X)}) \\ &\leq C(\|\omega\|_{L^2(\{r \leq 3R\} \subset X)} + \|\Delta_X \omega\|_{W_{\mu-2}^{k,2}(X)}). \end{aligned}$$

For (2), it is straightforward to check that, for any $k \in \mathbb{N}_0$ and p -form $\omega \in W_{\mu}^{k+2,2}(X)$, we have

$$\|\Delta_X \omega\|_{W_{\mu-2}^{k,2}(X)} \leq C \cdot \|\omega\|_{W_{\mu}^{k+2,2}(X)}.$$

First, to prove $\dim(\ker(\Delta_X)) < \infty$, let us take any sequence $\omega_j \in \ker(\Delta_X)$ with

$$\|\omega_j\|_{W_{\mu}^{k+2,2}(X)} = 1.$$

Rellich's theorem implies that the inclusion

$$W_{\mu}^{k+2,2}(X) \hookrightarrow L^2(\{r \leq 3R\} \subset X)$$

is compact. Combining this and (4.1), a subsequence of ω_j converges to ω_∞ in the $W_\mu^{k+2,2}$ -norm. Then the unit sphere of $\ker(\Delta_X)$ is compact, and hence $\dim(\ker(\Delta_X)) < \infty$. Moreover, standard elliptic estimate implies $\ker(\Delta_X) = \mathcal{H}_\mu^p(X)$.

Next, we will show $\text{Image}(\Delta_X)$ is closed in the $W_{\mu-2}^{k,2}$ -norm. This follows from the claim that, for any p -form $\omega \in (\mathcal{H}_\mu^p(X))^\perp \subset W_\mu^{k+2,2}(X)$ in terms of the $W_\mu^{k+2,2}$ -inner product, we have

$$\|\omega\|_{W_\mu^{k+2,2}(X)} \leq C \|\Delta_X \omega\|_{W_{\mu-2}^{k,2}(X)}.$$

Suppose not; then there exists a sequence of $\omega_j \in (\mathcal{H}_\mu^p(X))^\perp$ such that

$$\|\omega_j\|_{W_\mu^{k+2,2}(X)} = 1 \quad \text{and} \quad \|\Delta_X \omega_j\|_{W_{\mu-2}^{k,2}(X)} \rightarrow 0.$$

Combining Rellich's theorem and (4.1), passing to a subsequence, we have a limit

$$\omega_\infty \in \mathcal{H}_\mu^p(X) \subset W_\mu^{k+2,2}(X).$$

Since $\omega_\infty \in (\mathcal{H}_\mu^p(X))^\perp$, we have $\omega_\infty = 0$.

Now we show the solvability criterion. This implies that $\dim(\text{coker}(\Delta_X)) < \infty$. When $k = 0$, we can characterize $\text{Image}(\Delta_X)$ in $L_{\mu-2}^2(X)$ by finding its orthogonal complement in terms of the $L_{\mu-2}^2(X)$ -inner product. Notice that $\tilde{\eta} \perp \text{Image}(\Delta_X)$ if and only if

$$\int_X (\Delta_X \xi, \tilde{\eta}) \cdot (\hat{\varrho}_{\mu-2})^2 \, d\text{vol}_g = 0 \quad \text{for all } \xi \in W_\mu^{2,2}(X).$$

This is identical to $\Delta_X((\hat{\varrho}_{\mu-2})^2 \cdot \tilde{\eta}) = 0$ in the distributional sense. By standard elliptic regularity, this also coincides with the condition

$$\eta \equiv (\hat{\varrho}_{\mu-2})^2 \cdot \tilde{\eta} \in \mathcal{H}_{-\mu}^p(X).$$

Then the above implies that $\omega \in L_{\mu-2}^2(X)$ satisfies $\Delta_X \xi = \omega$ for some $\xi \in W_\mu^{2,2}(X)$ if and only if, for any $\eta \in \mathcal{H}_{-\mu}^p(X)$,

$$\int_X (\omega, \eta) \, d\text{vol}_g = 0.$$

When $k \in \mathbb{Z}_+$, the proof follows from the standard elliptic regularity theory.

The proof of (3) is similar to the proof of [5, Theorem 4.4], so is omitted. To prove (4), for any $\omega \in \mathcal{H}_\mu^p(X)$, by the asymptotics

$$|\Phi^* g - g^{\mathfrak{M}}| = O(\mathfrak{s}^{-\delta}),$$

we find that $\Delta_{\mathfrak{M}} \omega \in W_{\mu-\delta-2}^{k,2}(X)$ for all $k \in \mathbb{N}_0$. By (3), there is a solution $\eta \in W_{\mu-\delta}^{k+2,2}(X)$ such that $\Delta_{\mathfrak{M}} \eta = \Delta_{\mathfrak{M}} \omega$ when $r \geq 2R$. Then $\Delta_{\mathfrak{M}}(\eta - \omega) = 0$ when $r \geq 2R$. Let us pull back $\eta - \omega$ to $\hat{\mathfrak{M}}$ and write it as $\zeta' + \zeta''$, where each coefficient function of ζ' has zero $\mathbb{T}^2$ -average and ζ'' is $\mathbb{T}^2$ -invariant, and hence its coefficient functions depend only on (r, θ_1) . By Corollary 3.5, ζ' decays exponentially.

The next step is to analyze the $\mathbb{T}^2$ -invariant form ζ'' . Using separation by variables in the coordinates (r, θ_1) from Section 2, we observe that the Fourier expansion of ζ'' can be written as

$$\zeta'' = \sum_{j=-\infty}^{\infty} \sum_{i=1}^{\dim \mathbb{Z}_j^p(\hat{\mathfrak{M}})} c_{j,i} \cdot \omega_{j,i},$$

where $\{\omega_{j,i}\}_{1 \leq i \leq \dim \mathbb{Z}_j^p(\hat{\mathfrak{M}})}$ is a basis of $\mathbb{Z}_j^p(\hat{\mathfrak{M}})$.

First, since $\zeta'' \in L_\mu^2(\widehat{\mathfrak{M}})$, the components $e^{\sqrt{-1} \cdot j \cdot \theta_1}$ and $e^{-\sqrt{-1} \cdot j \cdot \theta_1}$ of ζ'' are also in $L_\mu^2(\widehat{\mathfrak{M}})$. This implies that $c_{j,i} = 0$ for all $j > \mu$. Next, we define the integer $\mathfrak{q} \equiv \lceil (\mu - \delta) \rceil$ and the form

$$\widehat{\zeta} \equiv \sum_{j=-\infty}^{\mathfrak{q}-1} \sum_{i=1}^{\dim \mathbb{Z}_j^p(\widehat{\mathfrak{M}})} c_{j,i} \cdot u_{j,i} \cdot e^{\sqrt{-1} \cdot m_{j,i} \cdot \theta_1}.$$

If $\mathfrak{q} < \mu$, then we write $\zeta'' = \zeta_{\mathfrak{q}} + \widehat{\zeta}$, where

$$\zeta_{\mathfrak{q}} \equiv \sum_{i=1}^{\dim \mathbb{Z}_{\mathfrak{q}}^p(\widehat{\mathfrak{M}})} c_{\mathfrak{q},i} \cdot u_{\mathfrak{q},i} \cdot e^{\sqrt{-1} \cdot m_{\mathfrak{q},i} \cdot \theta_1} \in \mathbb{Z}_{\mathfrak{q}}^p(\widehat{\mathfrak{M}}).$$

If $\mathfrak{q} > \mu$, then $\widehat{\zeta} = \zeta''$.

In the following, we will prove $\widehat{\zeta} \in W_{\mu-\delta}^{k,2}(\widehat{\mathfrak{M}})$. A technical issue is that, for example, when $p = 0$,

$$\int_{r=t} (r^k e^{\sqrt{-1} \cdot k \cdot \theta_1}, r^{-k} e^{\sqrt{-1} \cdot k \cdot \theta_1}) d\theta_1 \wedge d\theta_2 \wedge \Theta \neq 0.$$

To solve this issue, we define

$$\widehat{\zeta}_1 \equiv \sum_{j=-\infty}^{-|\mathfrak{q}|-10} \sum_{i=1}^{\dim \mathbb{Z}_j^p(\widehat{\mathfrak{M}})} c_{j,i} \cdot u_{j,i} \cdot e^{\sqrt{-1} \cdot m_{j,i} \cdot \theta_1}.$$

Then $\widehat{\zeta} = \widehat{\zeta}_1 + \widehat{\zeta}_2$, where

$$\widehat{\zeta}_2 = \sum_{j=-|\mathfrak{q}|-9}^{\mathfrak{q}-1} \sum_{i=1}^{\dim \mathbb{Z}_j^p(\widehat{\mathfrak{M}})} c_{j,i} \cdot u_{j,i} \cdot e^{\sqrt{-1} \cdot m_{j,i} \cdot \theta_1} \in W_{\mu-\delta}^{k,2}(\widehat{\mathfrak{M}}).$$

The main improvement is the components of $\widehat{\zeta}_1$ are orthogonal to each other. Therefore,

$$\int_{r=t} (\widehat{\zeta}_1, \widehat{\zeta}_1) d\theta_1 \wedge d\theta_2 \wedge \Theta = 8\pi^3 \cdot \sum_{j=-\infty}^{-|\mathfrak{q}|-10} \sum_{i=1}^{\dim \mathbb{Z}_j^p(\widehat{\mathfrak{M}})} |c_{j,i}|^2 \cdot |u_{j,i}|^2.$$

We only need to show that $\widehat{\zeta}_1 \in W_{\mu-\delta}^{k,2}(\widehat{\mathfrak{M}})$ using the information that $\widehat{\zeta}_1 \in W_\mu^{k,2}(\widehat{\mathfrak{M}})$. To this end, let us fix a large number $r_0 \gg 1$. By Definition 4.4, there are a constant

$$C = C(r_0, \kappa_0, \nu, \mu) > 0$$

and constants $-2 \leq \mathfrak{b}_{j,i} \leq 2$ such that, for any $j \leq -|\mathfrak{q}| - 10$ and $t > r_0$, we have

$$C^{-1} \cdot V^{\mathfrak{b}_{j,i}}(t) \cdot t^j \leq |u_{j,i}|(t) \leq C \cdot V^{\mathfrak{b}_{j,i}}(t) \cdot t^j.$$

Therefore, for any $j \leq -|\mathfrak{q}| - 10$, $t_1 > 4r_0$ and $t_2 \in (r_0, 2r_0)$,

$$\begin{aligned} \frac{|u_{j,i}|(t_1)}{|u_{j,i}|(t_2)} &\leq C \cdot \left(\frac{V(t_1)}{V(t_2)} \right)^{\mathfrak{b}_{j,i}} \cdot \left(\frac{t_1}{t_2} \right)^j \leq C \cdot \left(\frac{V(t_1)}{V(t_2)} \right)^2 \cdot \left(\frac{t_1}{t_2} \right)^{-|\mathfrak{q}|-10} \\ &\leq C \cdot V^2(t_1) \cdot t_1^{-|\mathfrak{q}|-10}. \end{aligned}$$

Note that the constants $C = C(r_0, \kappa_0, \nu, \mu) > 0$ are allowed to change line by line. So

$$\begin{aligned} \int_{r=t_1} (\hat{\xi}_1, \hat{\xi}_1) d\theta_1 \wedge d\theta_2 \wedge \Theta \\ \leq C \cdot V^4(t_1) \cdot t_1^{-2|\mathfrak{q}|-20} \int_{r_0}^{2r_0} \left(\int_{r=t_2} (\hat{\xi}_1, \hat{\xi}_1) d\theta_1 \wedge d\theta_2 \wedge \Theta \right) dt_2 \\ \leq C \cdot V^4(t_1) \cdot t_1^{-2|\mathfrak{q}|-20} \|\hat{\xi}_1\|_{L^2_\mu(\hat{\mathfrak{M}})}^2. \end{aligned}$$

Then we have $\hat{\xi}_1 \in L^2_{\mu-\delta}(\hat{\mathfrak{M}})$, and the higher-order estimate follows from the standard elliptic regularity. $\square$

5. Applications

In this section, we will apply the above results to prove Theorem 1.3, Corollary 1.4, Theorem 1.5, Corollary 1.6 and Theorem 1.10 from the introduction.

5.1. Existence of harmonic functions.

Proof of Theorem 1.3. From Section A.2, we see that the function $r^k e^{\sqrt{-1}k\theta_1}$ is a harmonic function on the model space $\hat{\mathfrak{M}}$. Moreover, it descends to $\mathfrak{M}$ by our assumption. Let ψ be a cut-off function on X such that

$$\psi = \begin{cases} 1, & s \geq 2R_0, \\ 0, & s \leq R_0. \end{cases}$$

Since

$$|\nabla_{g^{\hat{\mathfrak{M}}}}^l (\Phi^* g - g^{\hat{\mathfrak{M}}})| = O(\mathfrak{s}^{-l-\mathfrak{n}}) \quad \text{as } \mathfrak{s} \rightarrow \infty$$

for all $l \in \mathbb{N}_0$, and $\mathfrak{s}(x) \equiv r(x) \cdot V^{\frac{1}{2}}(x)$, we have

$$\Delta_g(\psi \cdot r^k e^{\sqrt{-1}k\theta_1}) \in W_{\mu-2}^{l,2}(X)$$

for any $l \in \mathbb{N}_0$ and $\mu > k - \mathfrak{n}$. We can choose μ such that, in addition,

$$\max\{0, k - \mathfrak{n}\} < \mu < k - \epsilon$$

and μ is not an integer. Using the maximum principle, we have $\mathcal{H}_{-\mu}(X) = \{0\}$. Applying item (2) of Proposition 4.5, there exists some $u \in W_{\mu}^{l+2,2}(X)$ which solves the equation

$$\Delta_g u = -\Delta_g(\psi \cdot r^k e^{\sqrt{-1}k\theta_1}).$$

Then

$$h_k \equiv \psi \cdot r^k e^{\sqrt{-1}k\theta_1} + u$$

is harmonic with respect to g . Proposition 3.3 then implies (1.3). $\square$

Proof of Corollary 1.4. Recall that

$$|d(re^{\sqrt{-1}\theta_1})|_{g^{\mathfrak{M}}} = \sqrt{2} \cdot V^{-\frac{1}{2}} \rightarrow 0 \quad \text{as } \mathfrak{s} \rightarrow \infty.$$

This implies that $|dh_1|_g = o(1)$ as $s \rightarrow \infty$, where $s(x) \equiv d_g(x, x_0)$ and x_0 is a fixed point in some compact subset $X_R \subset X$. Since $\text{Ric}_g \geq 0$ and dh_1 is harmonic, by Bochner's formula, we have $\Delta_g(|dh_1|_g^2) \leq 0$. Since dh_1 is not identically zero, the maximum of $|dh_1|_g^2$ would be achieved in the interior. The strong maximum principle then leads to a contradiction. $\square$

5.2. Hodge theory on ALG* manifolds.

Proof of Theorem 1.5. As in [18, Lemma 6.11], we define a smooth function f such that $f(r) = r$ when $r \leq R$ and $f(r) = 2R$ when $r \geq 2R$. Therefore, the map $F: X \rightarrow X$ defined by

$$F(x) = \begin{cases} \Phi((f(r), \theta_1, \theta_2, \theta_3)) & \text{if } x = \Phi(r, \theta_1, \theta_2, \theta_3), \\ x & \text{if } x \in X_R, \end{cases}$$

is well-defined and homotopic to the identity. Then any smooth closed 1-form η_1 on X is cohomologous to $\eta_2 \equiv F^*\eta_1$. Clearly,

$$\Phi^*F^*\eta_1 = a_1 d\theta_1 + a_2 d\theta_2 + a_3 d\theta_3$$

on the set $r \geq 2R$, where a_1, a_2, a_3 are constants. So then $\eta_2 \in W_\mu^{k,2}(X)$ for all $k \in \mathbb{N}_0$ and $0 < \mu \ll 1$. The mapping $v \mapsto (\eta_2, v)_{L^2(X)}$ is a bounded linear functional on $\mathcal{H}_{-2-\mu}^1(X)$. By the Riesz representation theorem, there exists $\eta_3 \in \mathcal{H}_{-2-\mu}^1(X)$ such that

$$(\eta_2, v)_{L^2(X)} = (\eta_3, v)_{L^2(X)} \quad \text{for all } v \in \mathcal{H}_{-2-\mu}^1(X).$$

By Proposition 4.5, there exists a 1-form $\eta_4 \in W_{\mu+2}^{k+2,2}(X)$ such that

$$\eta_2 - \eta_3 = d\delta\eta_4 + \delta d\eta_4.$$

The boundary term in the integral

$$\int_{r \leq R} \left((\eta_3, d\delta\eta_3 + \delta d\eta_3) - (d\eta_3, d\eta_3) - (\delta\eta_3, \delta\eta_3) \right) d\text{vol}_g$$

goes to 0 when $R \rightarrow \infty$. So we see that $d\eta_3 = \delta\eta_3 = 0$. So η_1 is also cohomologous to the closed and co-closed 1-form $\eta_5 \equiv \delta d\eta_4 \in W_\mu^{k,2}(X)$.

So then η_5 is a harmonic 1-form in $W_\mu^{k,2}(X)$ for any μ satisfying $0 < \mu \ll 1$. By Proposition 4.5, η_5 admits a harmonic expansion. By Definition 4.4, the leading term η_6 of the pull-back of η_5 to $\hat{\mathfrak{M}}$ is given by

$$(5.1) \quad \begin{aligned} \eta_6 = & (A_0 + B_0 V^2) dz + (A'_0 + B'_0 V^2) d\bar{z} \\ & + (C_0 + D_0 V^2) d\theta_2 + (C'_0 V^{-1} + D'_0 V) \Theta, \end{aligned}$$

where $A_0, A'_0, B_0, B'_0, C_0, C'_0, D_0, D'_0$ are constants and $z \equiv r e^{\sqrt{-1}\theta_1}$. Since η_5 is both closed and co-closed, we have

$$(5.2) \quad |d\eta_6|_{g^{\hat{\mathfrak{M}}}} + |\delta_{\hat{\mathfrak{M}}} \eta_6|_{g^{\hat{\mathfrak{M}}}} = O(\varkappa^{-1-\epsilon})$$

for a constant $\epsilon > 0$, as $\varkappa \rightarrow \infty$. Then (5.1) and (5.2) imply that

$$\eta_6 = A_0 dz + A'_0 d\bar{z} + C_0 d\theta_2.$$

By Theorem 1.3, there exists a harmonic function $h_1: X \rightarrow \mathbb{C}$ such that $h_1 = z + O(s^{-\epsilon})$ for some $\epsilon > 0$ as $s \rightarrow \infty$. Then the closed and co-closed 1-form

$$\eta_7 = \eta_5 - A_0 dh_1 - A'_0 d\bar{h}_1$$

is cohomologous to η_1 . This shows that the natural mapping

$$\{\omega \in \Omega^1(X) \mid d\omega = 0, \delta\omega = 0, \Phi^*\omega = \omega_0 + O(\varkappa^{-\epsilon}) \text{ as } \varkappa \rightarrow \infty, \omega_0 \in \mathcal{W}^1\} \rightarrow H_{dR}^1(X)$$

is surjective. To show the injectivity, assume that

$$\omega \in \{\omega \in \Omega^1(X) \mid d\omega = 0, \delta\omega = 0, \Phi^*\omega = \omega_0 + O(\varkappa^{-\epsilon}) \text{ as } \varkappa \rightarrow \infty, \omega_0 \in \mathcal{W}^1\}$$

satisfies $\omega = du$ for a function $u: X \rightarrow \mathbb{R}$. On the compact subset $\{r \leq R\} \subset X$, u is bounded. For each point $\Phi(r, \theta_1, \theta_2, \theta_3)$ on $\{r \geq R\} \subset X$, the path

$$\gamma_{r, \theta_1, \theta_2, \theta_3}(t) \equiv \Phi(t, \theta_1, \theta_2, \theta_3), t \in [R, r],$$

connects $\Phi(R, \theta_1, \theta_2, \theta_3)$ and $\Phi(r, \theta_1, \theta_2, \theta_3)$. Using

$$u(\Phi(r, \theta_1, \theta_2, \theta_3)) = u(\Phi(R, \theta_1, \theta_2, \theta_3)) + \int_{\gamma_{r, \theta_1, \theta_2, \theta_3}} \omega,$$

we have that $u = O(s^{1-\epsilon})$ as $s \rightarrow \infty$ because the integral of $(\Phi^{-1})^*d\theta_2$ on $\gamma_{r, \theta_1, \theta_2, \theta_3}$ is 0. Since ω is co-closed, from Proposition 4.5 and Definition 4.4, u admits a harmonic expansion $u = A_0 + B_0 V + O(s^{-\epsilon})$ as $s \rightarrow \infty$. Integrating by parts,

$$0 = \int_{r < R} ((\delta du, 1) - (du, d1)) \, d\text{vol}_g = \int_{r=R} \partial_n u \, d\sigma_g,$$

where ∂_n is the derivative with respect to the unit normal, and $d\sigma_g$ is the induced area element. Clearly,

$$\begin{aligned} \Phi^* \partial_n u &= V^{-\frac{1}{2}} \partial_r (\Phi^* u) + O(\varkappa^{-1-\epsilon}), \\ \Phi^* d\sigma_g &= r V^{\frac{1}{2}} d\theta_1 \wedge d\theta_2 \wedge \Theta + O(\varkappa^{-\epsilon}) \end{aligned}$$

as $\varkappa \rightarrow \infty$. This implies that $B_0 = 0$ by taking $R \rightarrow \infty$. So $u = A_0$ by the maximum principle. This implies that $\omega = 0$, which completes the proof of the injectivity. $\square$

Proof of Corollary 1.6. Let

$$\omega \in \{\omega \in \Omega^1(X) \mid d\omega = 0, \delta\omega = 0, \Phi^*\omega = \omega_0 + O(\varkappa^{-\epsilon}) \text{ as } \varkappa \rightarrow \infty, \omega_0 \in \mathcal{W}^1\}.$$

Recall that

$$|d\theta_2|_{g^{\mathfrak{M}}} = V^{-\frac{1}{2}} = o(1) \quad \text{as } \varkappa \rightarrow \infty.$$

If X has non-negative Ricci curvature, then the same argument as in the proof of Corollary 1.4 above would imply that $\omega \equiv 0$. Therefore, $b^1(X) = 0$ by Theorem 1.5. $\square$

5.3. Asymptotics of ALG* gravitational instantons. We begin with a few remarks about hyper-Kähler structures. Let (M, g, I, J, K) be a hyper-Kähler 4-manifold. Recall that we denote the triple of 2-forms as $\omega = (\omega_1, \omega_2, \omega_3)$, where $\omega_1, \omega_2, \omega_3$ are the Kähler forms associated to I, J, K , respectively. These 2-forms satisfy

$$(5.3) \quad \begin{aligned} \omega_1 \wedge \omega_2 &= 0, & \omega_1 \wedge \omega_3 &= 0, & \omega_2 \wedge \omega_3 &= 0, \\ \omega_1 \wedge \omega_1 - \omega_2 \wedge \omega_2 &= 0, & \omega_1 \wedge \omega_1 - \omega_3 \wedge \omega_3 &= 0. \end{aligned}$$

Conversely, any triple of symplectic forms ω_i satisfying (5.3) determines a hyper-Kähler structure if we replace ω_3 by $-\omega_3$ if necessary. To see this, using the algebraic isomorphism of homogeneous spaces

$$\text{SO}_0(3, 3)/(\text{SO}(3) \times \text{SO}(3)) \cong \text{SL}(4, \mathbb{R})/\text{SO}(4)$$

(see for example [20, Chapter 7]), a triple ω uniquely determines a Riemannian metric g_ω such that each ω_j is self-dual with respect to g_ω and $\text{dvol}_{g_\omega} = \frac{1}{2}\omega_1 \wedge \omega_1$.

Next, define $\mathcal{Q}: \Lambda^2 \oplus \Lambda^2 \oplus \Lambda^2 \rightarrow \Lambda^4 \oplus \Lambda^4 \oplus \Lambda^4 \oplus \Lambda^4 \oplus \Lambda^4$ by

$$(5.4) \quad \begin{aligned} \mathcal{Q}(\omega_1, \omega_2, \omega_3) = & (\omega_1 \wedge \omega_2, \omega_1 \wedge \omega_3, \omega_2 \wedge \omega_3, \\ & \omega_1 \wedge \omega_1 - \omega_2 \wedge \omega_2, \omega_1 \wedge \omega_1 - \omega_3 \wedge \omega_3). \end{aligned}$$

The kernel of the linearized operator of $\mathcal{Q}$ at a hyper-Kähler triple will be

$$\Lambda_-^2 \oplus \Lambda_-^2 \oplus \Lambda_-^2 \oplus E_M,$$

where E_M is the rank-4 bundle given by

$$\begin{aligned} E_M \equiv \{ \theta \in \Lambda_+^2 \oplus \Lambda_+^2 \oplus \Lambda_+^2 \mid & \omega_i \wedge \theta_j + \omega_j \wedge \theta_i = 0, i < j \\ & \text{and } \omega_1 \wedge \theta_1 = \omega_2 \wedge \theta_2 = \omega_3 \wedge \theta_3 \}. \end{aligned}$$

Recall that there is a Dirac-type operator $\mathcal{D}: \Gamma(TM) \rightarrow \Gamma(E_M)$, defined as follows: for every $Y \in \Gamma(TM)$, $\mathcal{D}(Y)$ is the projection of the Lie derivative $\mathcal{L}_Y(\omega)$ to E_M . The operator $-\mathcal{D}\mathcal{D}^*$ can be identified with the Laplacian on functions since E_M admits a basis of parallel sections; see [5, Section 3].

Proof of Theorem 1.10. Using the Fredholm theory developed above, the argument is very similar to the proof of [5, Theorem A]. Decompose the difference

$$(5.5) \quad \Phi^* \omega^X - \omega^{\mathfrak{M}} = \eta_+ + \eta_-,$$

where η_+ is a self-dual triple and η_- is an anti-self-dual triple with respect to $g^{\mathfrak{M}}$ and the volume form $\frac{1}{2}(\omega_1^{\mathfrak{M}} \wedge \omega_1^{\mathfrak{M}})$. We choose an irrational number $\epsilon \in (0, 1)$ and view the right-hand side of (5.5) as an element of $W_{-\epsilon}^{k,2}(\mathfrak{M}_{2\nu}(R))$ for any $k \in \mathbb{Z}_+$. For sufficiently large R , expanding (5.4) yields

$$\mathcal{Q}(\Phi^* \omega^X) = \mathcal{Q}(\omega^{\mathfrak{M}}) + \mathcal{L}_{\omega^{\mathfrak{M}}}(\Phi^* \omega^X - \omega^{\mathfrak{M}}) + \mathcal{N}_{\omega^{\mathfrak{M}}}(\Phi^* \omega^X - \omega^{\mathfrak{M}}),$$

where $\mathcal{L}_{\omega^{\mathfrak{M}}}$ is the linearized operator at $\omega^{\mathfrak{M}}$, and $\mathcal{N}_{\omega^{\mathfrak{M}}}$ are the nonlinear terms. We then have

$$(5.6) \quad \mathcal{L}_{\omega^{\mathfrak{M}}}(\eta_+) = \mathcal{L}_{\omega^{\mathfrak{M}}}(\eta_+ + \eta_-) = -\mathcal{N}_{\omega^{\mathfrak{M}}}(\eta_+ + \eta_-).$$

From the structure of (5.4), the right-hand side of (5.6) must be in $W_{-2\epsilon}^{k,2}(\mathfrak{M}_{2\nu}(R))$. Therefore, the projection θ of η_+ to $E_{\mathfrak{M}_{2\nu}(R)}$ satisfies

$$\theta \in W_{-\epsilon}^{k,2}(\mathfrak{M}_{2\nu}(R)) \quad \text{and} \quad \eta_+ - \theta \in W_{-2\epsilon}^{k,2}(\mathfrak{M}_{2\nu}(R)).$$

By Proposition 4.5 (3), there exists a bounded linear operator

$$G_{g^{\mathfrak{M}}}: W_{-\epsilon}^{k,2}(\Gamma(E_{\mathfrak{M}_{2\nu}(R)})) \rightarrow W_{-\epsilon+2}^{k+2,2}(\Gamma(E_{\mathfrak{M}_{2\nu}(R)}))$$

such that

$$-\mathcal{D}_{g^{\mathfrak{M}}} \mathcal{D}_{g^{\mathfrak{M}}}^* G_{g^{\mathfrak{M}}} \theta = \theta.$$

Consider the vector field $Y = \mathcal{D}_{g^{\mathfrak{M}}}^* G_{g^{\mathfrak{M}}} \theta$ which satisfies $Y \in W_{1-\epsilon}^{k+1,2}(\mathfrak{M}_{2\nu}(R))$. By elliptic regularity, Y is smooth, and by Proposition 3.3, $|\nabla^m Y| = O(\mathfrak{s}^{1-m-\epsilon})$ for any $m \in \mathbb{N}_0$ as $\mathfrak{s} \rightarrow \infty$. It is important to point out that Y does not depend on the choice of k . Define $\Phi_t: \mathfrak{M}_{2\nu}(2R) \rightarrow \mathfrak{M}_{2\nu}(R)$ by $\Phi_t(x) = \exp_{g^{\mathfrak{M}},x}(tY_x)$. If $0 \leq t \leq 1$, since $1 - \epsilon < 1$, the vector tY_x has norm much smaller than the conjugate radius at x (which is comparable to $\mathfrak{s}(x)$), so Φ_t is a diffeomorphism onto its image for R sufficiently large. Then we have

$$\Phi_1^* \omega^{\mathfrak{M}} - \omega^{\mathfrak{M}} - \mathcal{L}_Y \omega^{\mathfrak{M}} \in W_{-2\epsilon}^{k,2}(\mathfrak{M}_{2\nu}(R)).$$

Let $\Phi' = \Phi \circ \Phi_1$. From (5.5), we then have

$$\begin{aligned} (\Phi')^* \omega^X - \omega^{\mathfrak{M}} - \eta_+ - \eta_- - \mathcal{L}_Y \omega^{\mathfrak{M}} &= \Phi_1^* \Phi^* \omega^X - \Phi^* \omega^X - \mathcal{L}_Y \omega^{\mathfrak{M}} \\ &= (\Phi_1^* - \text{Id})(\Phi^* \omega^X - \omega^{\mathfrak{M}}) + \Phi_1^* \omega^{\mathfrak{M}} \\ &\quad - \omega^{\mathfrak{M}} - \mathcal{L}_Y \omega^{\mathfrak{M}} \in W_{-2\epsilon}^{k,2}(\mathfrak{M}_{2\nu}(R)). \end{aligned}$$

Similar to (5.6) above, we have the expansion

$$\begin{aligned} (5.7) \quad \mathcal{L}_{\omega^{\mathfrak{M}}}(\eta_+ + (\mathcal{L}_Y \omega^{\mathfrak{M}})^+) &= \mathcal{L}_{\omega^{\mathfrak{M}}}(\eta_+ + \eta_- + \mathcal{L}_Y \omega^{\mathfrak{M}}) \\ &= -\mathcal{N}_{\omega^{\mathfrak{M}}}(\eta_+ + \eta_- + \mathcal{L}_Y \omega^{\mathfrak{M}}). \end{aligned}$$

From the structure of (5.4), the right-hand side of (5.7) must be in $W_{-2\epsilon}^{k,2}(\mathfrak{M}_{2\nu}(R))$. Since the projection of $\eta_+ + (\mathcal{L}_Y \omega^{\mathfrak{M}})^+$ to $E\mathfrak{M}_{2\nu}(R)$ is $\theta + \mathcal{D}_{g^{\mathfrak{M}}} Y = 0$, we see that

$$\eta_+ + (\mathcal{L}_Y \omega^{\mathfrak{M}})^+ \in W_{-2\epsilon}^{k,2}(\mathfrak{M}_{2\nu}(R)).$$

In other words, redefining $\eta_{\pm}$, we may assume that

$$(\Phi')^* \omega^X - \omega^{\mathfrak{M}} = \eta_+ + \eta_-$$

with $\eta_- \in W_{-\epsilon}^{k,2}(\mathfrak{M}_{2\nu}(R))$ and $\eta_+ \in W_{-2\epsilon}^{k,2}(\mathfrak{M}_{2\nu}(R))$. Since η_- is anti-self-dual with respect to $\omega^{\mathfrak{M}}$,

$$\begin{aligned} \Delta_{g^{\mathfrak{M}}} \eta_- &= (-*_{g^{\mathfrak{M}}} d *_{g^{\mathfrak{M}}} d - d *_{g^{\mathfrak{M}}} d *_{g^{\mathfrak{M}}}) \eta_- \\ &= -*_{g^{\mathfrak{M}}} d *_{g^{\mathfrak{M}}} d \eta_- + d *_{g^{\mathfrak{M}}} d \eta_-. \end{aligned}$$

But since both triples are closed, $d\eta_- = -d\eta_+$, and therefore, $\Delta_{g^{\mathfrak{M}}} \eta_- \in W_{-2\epsilon-2}^{k-2,2}(\mathfrak{M}_{2\nu}(R))$. By (3) and (4) in Proposition 4.5, if $\epsilon < \frac{1}{2}$, then $\eta_- \in W_{-2\epsilon}^{k,2}(\mathfrak{M}_{2\nu}(2R))$. We can replace (Φ, ϵ) with $(\Phi', 2\epsilon)$ and repeat this process until $\epsilon > \frac{1}{2}$. If $\frac{1}{2} < \epsilon < 1$, then η_- is the sum of a triple in $\mathcal{Z}_{-1}^{2,-} \otimes \mathbb{R}^3$ and a triple in $W_{-2\epsilon}^{k,2}(\mathfrak{M}_{2\nu}(2R))$. By Lemma A.8, all the non-zero forms in $\mathcal{Z}_{-1}^{2,-}$ are not invariant under the $\mathbb{Z}_2$ -action $\iota: \mathfrak{M} \rightarrow \mathfrak{M}$. So we see that we can replace (Φ, ϵ) with $(\Phi', 2\epsilon)$ and repeat this process again. If $1 < \epsilon < 2$, then η_- is the sum of $\eta_{\mathcal{Z}} \in \mathcal{Z}_{-2}^{2,-} \otimes \mathbb{R}^3$ and a triple in $W_{-\epsilon-1}^{k,2}(\mathfrak{M}_{2\nu}(2R))$. Moreover, the decay rate of $d\eta_{\mathcal{Z}}$ is strictly larger than 3. So by the explicit definition of $\mathcal{Z}_{-2}^{2,-}$ in Lemma A.8, $|\nabla^k \eta_{\mathcal{Z}}| = O(\mathfrak{s}^{-2-k})$ as $\mathfrak{s} \rightarrow \infty$, for any $k \in \mathbb{N}_0$. The above is a finite iteration, so $\eta_- - \eta_{\mathcal{Z}} \in W_{-\epsilon-1}^{k,2}(\mathfrak{M}_{2\nu}(2R))$ for any $k \in \mathbb{N}_0$. Then, by Proposition 3.3, $|\nabla^k(\eta_- - \eta_{\mathcal{Z}})| = O(\mathfrak{s}^{-\epsilon-1-k})$ as $\mathfrak{s} \rightarrow \infty$, for any $k \in \mathbb{N}_0$. Therefore, by replacing Φ with Φ' , (1.4) is satisfied for $n = 2$. $\square$

A. Computations on the ALG* model space

In this section, we derive some expansions of harmonic forms on the model space

$$(\hat{\mathfrak{M}}_\nu(R), g_{\kappa_0, L}^{\hat{\mathfrak{M}}}).$$

By scaling, we may assume that $L = 1$. For simplicity, we will denote $\hat{\mathfrak{M}} \equiv \hat{\mathfrak{M}}_\nu(R)$. Define an orthonormal basis of 1-forms by

$$\{e^0, e^1, e^2, e^3\} = \{V^{\frac{1}{2}}dr, V^{\frac{1}{2}}r d\theta_1, V^{\frac{1}{2}}d\theta_2, V^{-\frac{1}{2}}\Theta\},$$

and use the orientation $e^0 \wedge e^1 \wedge e^2 \wedge e^3$. The following formulas are straightforward to verify and will be used throughout the appendix. We have

$$\begin{aligned} de^0 &= 0, \quad de^1 = \partial_r(V^{\frac{1}{2}}r)V^{-1}r^{-1}e^0 \wedge e^1, \quad de^2 = \partial_r(V^{\frac{1}{2}})V^{-1}e^0 \wedge e^2, \\ de^3 &= \partial_r(V^{-\frac{1}{2}})e^0 \wedge e^3 + \frac{\nu}{2\pi}V^{-\frac{3}{2}}r^{-1}e^1 \wedge e^2, \end{aligned}$$

and

$$\begin{aligned} d(e^0 \wedge e^1) &= d(e^0 \wedge e^2) = d(e^2 \wedge e^3) = 0, \\ d(e^1 \wedge e^3) &= V^{-\frac{1}{2}}r^{-1}e^0 \wedge e^1 \wedge e^3, \\ d(e^1 \wedge e^2) &= (V^{-\frac{3}{2}}\partial_r(V) + V^{-\frac{1}{2}}r^{-1})e^0 \wedge e^1 \wedge e^2, \\ d(e^0 \wedge e^3) &= -\frac{\nu}{2\pi}V^{-\frac{3}{2}}r^{-1}e^0 \wedge e^1 \wedge e^2. \end{aligned}$$

A.1. Laplacian on functions. First, we need to characterize the harmonic functions of the form

$$(A.1) \quad h = f(r)e^{\sqrt{-1}k\theta_1}, \quad k \in \mathbb{Z},$$

with prescribed growth order. The following lemma identifies the space $\mathcal{Z}_{\mathfrak{q}}^0(\hat{\mathfrak{M}})$.

Lemma A.1. *Given an integer $\mathfrak{q} \in \mathbb{Z}$, let $\mathcal{Z}_{\mathfrak{q}}^0(\hat{\mathfrak{M}})$ be the linear space with a basis*

$$\begin{aligned} \mathcal{B}_0^0 &\equiv \{1, V\}, \\ \mathcal{B}_{\mathfrak{q}}^0 &\equiv \{r^{\mathfrak{q}}e^{\sqrt{-1}\cdot\mathfrak{q}\cdot\theta_1}, r^{\mathfrak{q}}e^{-\sqrt{-1}\cdot\mathfrak{q}\cdot\theta_1}\}, \quad \mathfrak{q} \in \mathbb{Z} \setminus \{0\}. \end{aligned}$$

If a function h satisfies (A.1) for some $k \in \mathbb{Z}$ and solves $\Delta h = 0$ on $\hat{\mathfrak{M}}$, then

$$h \in \begin{cases} \mathcal{Z}_0^0(\hat{\mathfrak{M}}) & \text{if } k = 0, \\ \mathcal{Z}_k^0(\hat{\mathfrak{M}}) \oplus \mathcal{Z}_{-k}^0(\hat{\mathfrak{M}}) & \text{if } k \in \mathbb{Z} \setminus \{0\}. \end{cases}$$

Proof. Taking the differential dh of h , we have that

$$dh = e^{\sqrt{-1}k\theta_1}(f'dr + \sqrt{-1}kf d\theta_1) = e^{\sqrt{-1}k\theta_1}(f'V^{-\frac{1}{2}}e^0 + \sqrt{-1}kfV^{-\frac{1}{2}}r^{-1}e^1).$$

Then

$$\begin{aligned} *dh &= e^{\sqrt{-1}k\theta_1}(f'V^{-\frac{1}{2}}e^1 \wedge e^2 \wedge e^3 - \sqrt{-1}kfV^{-\frac{1}{2}}r^{-1}e^0 \wedge e^2 \wedge e^3) \\ &= e^{\sqrt{-1}k\theta_1}(f'r d\theta_1 \wedge d\theta_2 \wedge \Theta - \sqrt{-1}kfr^{-1}dr \wedge d\theta_2 \wedge \Theta). \end{aligned}$$

Applying d ,

$$\begin{aligned} d * dh &= e^{\sqrt{-1}k\theta_1} \sqrt{-1}k d\theta_1 \wedge (-\sqrt{-1}kfr^{-1}dr \wedge d\theta_2 \wedge \Theta) \\ &\quad + e^{\sqrt{-1}k\theta_1} \partial_r(f'r) dr \wedge d\theta_1 \wedge d\theta_2 \wedge \Theta \\ &= e^{\sqrt{-1}k\theta_1} (\partial_r(f'r) - k^2 fr^{-1}) V^{-1} r^{-1} e^0 \wedge e^1 \wedge e^2 \wedge e^3. \end{aligned}$$

Finally, we have

$$(A.2) \quad \Delta h = - * d * dh = -e^{\sqrt{-1}k\theta_1} (f'' + r^{-1}f' - k^2 r^{-2}f) V^{-1}.$$

The ODE for a harmonic function is therefore $f'' + r^{-1}f' - k^2 r^{-2}f = 0$, and the solutions are given by

$$f(r) = \begin{cases} C_1 + C_2 V & k = 0 \\ C_1 r^k + C_2 r^{-k} & k \neq 0. \end{cases} \quad \square$$

A.2. Forms of Type I on the model space. A 1-form ω is said to be of *Type I* if ω satisfies

$$(A.3) \quad \omega = e^{\sqrt{-1}k\theta_1} (f(r)e^0 + a(r)e^1).$$

This subsection studies the Type I solutions of $\Delta\omega = 0$.

Let us denote $dz = dr + \sqrt{-1}r d\theta_1$. The following lemma characterizes the Type I solutions of $\Delta\omega = 0$.

Lemma A.2. *Given an integer $\mathfrak{q} \in \mathbb{Z}$, let $\mathcal{Z}_{\mathfrak{q}}^{1,I}(\widehat{\mathfrak{M}})$ be the linear space with a basis*

$$\begin{aligned} \mathcal{B}_0^{1,I} &\equiv \{dz, V^2 dz, d\bar{z}, V^2 d\bar{z}\}, \\ \mathcal{B}_{\mathfrak{q}}^{1,I} &\equiv \{e^{\sqrt{-1}(\mathfrak{q}+1)\theta_1} r^{\mathfrak{q}} \cdot dz, e^{-\sqrt{-1}(\mathfrak{q}+1)\theta_1} (v - 4\mathfrak{q}\pi V) r^{\mathfrak{q}} \cdot dz, \\ &\quad e^{-\sqrt{-1}(\mathfrak{q}+1)\theta_1} r^{\mathfrak{q}} \cdot d\bar{z}, e^{\sqrt{-1}(\mathfrak{q}+1)\theta_1} (v - 4\mathfrak{q}\pi V) r^{\mathfrak{q}} \cdot d\bar{z}\}, \quad \mathfrak{q} \neq 0. \end{aligned}$$

If a complex-valued 1-form ω of Type I satisfies (A.3) for some $k \in \mathbb{Z}$ and solves $\Delta\omega = 0$ on $\widehat{\mathfrak{M}}$, then

$$\omega \in \begin{cases} \mathcal{Z}_{-1}^{1,I}(\widehat{\mathfrak{M}}) & \text{if } k = 1, \\ \mathcal{Z}_{k-2}^{1,I}(\widehat{\mathfrak{M}}) \oplus \mathcal{Z}_{-k}^{1,I}(\widehat{\mathfrak{M}}) & \text{if } k \in \mathbb{Z} \setminus \{1\}. \end{cases}$$

Proof. In the proof, instead of directly analyzing Type I, we will first reduce (A.3) by analyzing the following ansatz:

$$(A.4) \quad \omega = e^{\sqrt{-1}k\theta_1} (f(r)e^0 + a(r)e^1) = e^{\sqrt{-1}k\theta_1} u(r) (e^{-\sqrt{-1}\theta_1} V^{\frac{1}{2}} dz).$$

Then

$$\omega = e^{\sqrt{-1}k\theta_1} u(r) (e^0 + \sqrt{-1}e^1),$$

that is, $a(r) = \sqrt{-1}f(r)$. Note that, for the complex structure I defined as above, we have $I^*(e^0) = -e^1$. Then

$$e^0 - \sqrt{-1}I^*e^0 = e^0 + \sqrt{-1}e^1 \in \Lambda_I^{1,0}.$$

Since the metric is Kähler, the Hodge Laplacian preserves the type of forms, so (A.4) is a natural ansatz.

To begin with, let us compute the Laplacian of ω . First,

$$*\omega = e^{\sqrt{-1}k\theta_1}u(r)(e^1 \wedge e^2 \wedge e^3 - \sqrt{-1}e^0 \wedge e^2 \wedge e^3).$$

This is

$$*\omega = e^{\sqrt{-1}k\theta_1}u(r)V^{\frac{1}{2}}(r d\theta_1 \wedge d\theta_2 \wedge \Theta - \sqrt{-1}dr \wedge d\theta_2 \wedge \Theta).$$

Then

$$\begin{aligned} d*\omega &= e^{\sqrt{-1}k\theta_1}\sqrt{-1}kuV^{\frac{1}{2}}d\theta_1 \wedge (-\sqrt{-1}dr \wedge d\theta_2 \wedge \Theta) \\ &\quad + e^{\sqrt{-1}k\theta_1}\partial_r(rV^{\frac{1}{2}}u)dr \wedge d\theta_1 \wedge d\theta_2 \wedge \Theta \\ &= e^{\sqrt{-1}k\theta_1}(\partial_r(rV^{\frac{1}{2}}u) - kuV^{\frac{1}{2}})V^{-1}r^{-1}e^0 \wedge e^1 \wedge e^2 \wedge e^3, \end{aligned}$$

which implies that

$$\delta\omega = -*d*\omega = -e^{\sqrt{-1}k\theta_1}(\partial_r(ruV^{\frac{1}{2}}) - kuV^{\frac{1}{2}})V^{-1}r^{-1}.$$

We define $p(r) = (\partial_r(rV^{\frac{1}{2}}u) - kuV^{\frac{1}{2}})V^{-1}r^{-1}$. Next, we have

$$d\delta\omega = -(e^{\sqrt{-1}k\theta_1}\sqrt{-1}kp d\theta_1 + e^{\sqrt{-1}k\theta_1}p'dr),$$

or

$$(A.5) \quad d\delta\omega = -e^{\sqrt{-1}k\theta_1}(p'V^{-\frac{1}{2}}e^0 + \sqrt{-1}kpV^{-\frac{1}{2}}r^{-1}e^1).$$

Next, we compute

$$\begin{aligned} d\omega &= e^{\sqrt{-1}k\theta_1}\{\sqrt{-1}k d\theta_1 \wedge (uV^{\frac{1}{2}}(dr + \sqrt{-1}r d\theta_1)) + \partial_r(rV^{\frac{1}{2}}u)dr \wedge (\sqrt{-1}d\theta_1)\} \\ &= \sqrt{-1}e^{\sqrt{-1}k\theta_1}(\partial_r(rV^{\frac{1}{2}}u) - kuV^{\frac{1}{2}})V^{-1}r^{-1}e^0 \wedge e^1 \equiv \sqrt{-1}e^{\sqrt{-1}k\theta_1}pe^0 \wedge e^1. \end{aligned}$$

Noting that $*e^0 \wedge e^1 = e^2 \wedge e^3$, applying Hodge star, we have

$$*d\omega = \sqrt{-1}e^{\sqrt{-1}k\theta_1}pe^2 \wedge e^3,$$

which implies that

$$\begin{aligned} d*d\omega &= \sqrt{-1}e^{\sqrt{-1}k\theta_1}\sqrt{-1}kp d\theta_1 \wedge e^2 \wedge e^3 + \sqrt{-1}e^{\sqrt{-1}k\theta_1}p'dr \wedge e^2 \wedge e^3 \\ &= e^{\sqrt{-1}k\theta_1}\{-kpV^{-\frac{1}{2}}r^{-1}e^1 \wedge e^2 \wedge e^3 + \sqrt{-1}p'V^{-\frac{1}{2}}e^0 \wedge e^2 \wedge e^3\}. \end{aligned}$$

Applying Hodge star,

$$*d*d\omega = e^{\sqrt{-1}k\theta_1}(kpV^{-\frac{1}{2}}r^{-1}e^0 + \sqrt{-1}p'V^{-\frac{1}{2}}e^1).$$

Combining with (A.5) from above, we see that

$$\begin{aligned} \Delta\omega &= -e^{\sqrt{-1}k\theta_1}\{p'V^{-\frac{1}{2}}e^0 + \sqrt{-1}kpV^{-\frac{1}{2}}r^{-1}e^1\} \\ &\quad - e^{\sqrt{-1}k\theta_1}(kpV^{-\frac{1}{2}}r^{-1}e^0 + \sqrt{-1}p'V^{-\frac{1}{2}}e^1) \\ &= -e^{\sqrt{-1}k\theta_1}(p' + kr^{-1}p)V^{-\frac{1}{2}}(e^0 + \sqrt{-1}e^1). \end{aligned}$$

Assume that ω is a harmonic 1-form. Then we have the homogeneous first-order system

$$p' + kr^{-1}p = 0.$$

The general solution is $p = C_1 r^{-k}$. Recalling that

$$p(r) = (\partial_r(rV^{\frac{1}{2}}u) - kuV^{\frac{1}{2}})V^{-1}r^{-1},$$

we have the ODE

$$\partial_r(ruV^{\frac{1}{2}}) - kuV^{\frac{1}{2}} = C_1 Vr^{-k+1}.$$

Let us define $\tilde{u} \equiv V^{\frac{1}{2}}u$; then this ODE can be written as

$$(A.6) \quad \partial_r(r\tilde{u}) - k\tilde{u} = C_1 Vr^{-k+1}.$$

The general solution of the associated homogeneous equation $\partial_r(r\tilde{u}) - k\tilde{u} = 0$ is $\tilde{u} = C_2 r^{k-1}$.

If $k = 1$, then ODE (A.6) becomes $r\tilde{u}' = C_1 V$, which has a particular solution

$$\tilde{u} = C_1 \left(\kappa_0 \log(r) + \frac{\nu}{4\pi} (\log(r))^2 \right).$$

Therefore, for $k = 1$, we have the general solution

$$\begin{aligned} u(r) &= V^{-\frac{1}{2}} \left(C_2 + C_1 \left(\kappa_0 \log(r) + \frac{\nu}{4\pi} (\log(r))^2 \right) \right) \\ &= V^{-\frac{1}{2}} \left(\left(C_2 - \kappa_0^2 \cdot \frac{\pi}{\nu} \cdot C_1 \right) + \frac{\pi}{\nu} \cdot C_1 \cdot V^2 \right). \end{aligned}$$

Letting

$$\tilde{C}_1 \equiv \frac{\pi}{\nu} \cdot C_1 \quad \text{and} \quad \tilde{C}_2 \equiv C_2 - \kappa_0^2 \cdot \frac{\pi}{\nu} \cdot C_1,$$

we can then write the general solution for $k = 1$ as $u(r) = \tilde{C}_2 V^{-\frac{1}{2}} + \tilde{C}_1 V^{\frac{3}{2}}$.

Next, we assume that $k \neq 1$. Then we need to solve

$$\partial_r(r\tilde{u}) - k\tilde{u} = C_1 Vr^{-k+1}.$$

Using the integrating factor r^{1-k} , we see that the general solution is

$$\begin{aligned} \tilde{u} &= C_2 r^{k-1} - \frac{C_1 r^{1-k}}{4\pi(k-1)^2} \left(\frac{\nu}{2} + 2(k-1)\pi\kappa_0 + \nu(k-1)\log(r) \right) \\ &= C_2 r^{k-1} - \frac{C_1 r^{1-k}}{4\pi(k-1)^2} \left(\frac{\nu}{2} + 2(k-1)\pi V \right). \end{aligned}$$

Letting

$$\tilde{C}_1 \equiv \frac{C_1 r^{1-k}}{4\pi(k-1)^2},$$

we therefore have the following general solution for $k \neq 1$:

$$u = C_2 V^{-\frac{1}{2}} r^{k-1} + \tilde{C}_1 r^{1-k} \left(\nu V^{-\frac{1}{2}} + 4(k-1)\pi V^{\frac{1}{2}} \right).$$

Notice that the Laplacian is a real operator, so we also have conjugate solutions. This completes the proof. $\square$

Remark A.3. We note that the solutions with $\tilde{C}_1 = 0$ are exactly the ones with $p = 0$, so by the above computations, they are exactly the solutions which satisfy $d\omega = 0$ and $\delta\omega = 0$. The solutions with $\tilde{C}_1 \neq 0$ are neither closed nor co-closed.

A.3. One-forms of Type II on the model space. A 1-form ω is said to be of *Type II* if

$$(A.7) \quad \omega = e^{\sqrt{-1}k\theta_1} (b(r)e^2 + c(r)e^3).$$

Then we have the following lemma.

Lemma A.4. *Given an integer $\mathfrak{q} \in \mathbb{Z}$, let $\mathcal{Z}_{\mathfrak{q}}^{1,\Pi}(\widehat{\mathfrak{M}})$ be the linear space with a basis*

$$\begin{aligned} \mathcal{B}_0^{1,\Pi} &\equiv \{-d\theta_2 + \sqrt{-1}V^{-1}\Theta, -V^2d\theta_2 + \sqrt{-1}V\Theta, \\ &\quad -d\theta_2 - \sqrt{-1}V^{-1}\Theta, -V^2d\theta_2 - \sqrt{-1}V\Theta\}, \\ \mathcal{B}_{\mathfrak{q}}^{1,\Pi} &\equiv \{e^{\sqrt{-1}\mathfrak{q}\theta_1}r^{\mathfrak{q}} \cdot (-d\theta_2 + \sqrt{-1}V^{-1}\Theta), e^{-\sqrt{-1}\mathfrak{q}\theta_1}r^{\mathfrak{q}} \cdot (-d\theta_2 - \sqrt{-1}V^{-1}\Theta), \\ &\quad e^{-\sqrt{-1}\mathfrak{q}\theta_1}(vV^{-\frac{1}{2}} - 4\mathfrak{q}\pi V^{\frac{1}{2}})r^{\mathfrak{q}} \cdot (-V^{\frac{1}{2}}d\theta_2 + \sqrt{-1}V^{-\frac{1}{2}}\Theta), \\ &\quad e^{\sqrt{-1}\mathfrak{q}\theta_1}(vV^{-\frac{1}{2}} - 4\mathfrak{q}\pi V^{\frac{1}{2}})r^{\mathfrak{q}} \cdot (-V^{\frac{1}{2}}d\theta_2 - \sqrt{-1}V^{-\frac{1}{2}}\Theta)\}, \quad \mathfrak{q} \in \mathbb{Z} \setminus \{0\}. \end{aligned}$$

If a complex-valued 1-form ω of Type II satisfies (A.7) for some $k \in \mathbb{Z}$ and solves $\Delta\omega = 0$ on $\widehat{\mathfrak{M}}$, then

$$\omega \in \begin{cases} \mathcal{Z}_0^{1,\Pi}(\widehat{\mathfrak{M}}) & \text{if } k = 0, \\ \mathcal{Z}_k^{1,\Pi}(\widehat{\mathfrak{M}}) \oplus \mathcal{Z}_{-k}^{1,\Pi}(\widehat{\mathfrak{M}}) & \text{if } k \in \mathbb{Z} \setminus \{0\}. \end{cases}$$

The proof relies on the following relationship between Type II forms and Type I forms.

Lemma A.5. *Any form ω_2 of Type II may be written as $J^*(\omega_1)$, where ω_1 is a form of Type I for the hyper-Kähler complex structure J and the dual linear operator J^* . So the decay rates of harmonic Type II forms are the same as Type I forms.*

Proof. We use the complex structure J given by

$$J^*(dx) = -d\theta_2, \quad J^*(V^{\frac{1}{2}}dy) = V^{-\frac{1}{2}}\Theta.$$

Recall that

$$e^0 + \sqrt{-1}e^1 = V^{\frac{1}{2}}(dr + \sqrt{-1}r d\theta_1) = V^{-\frac{1}{2}}e^{-\sqrt{-1}\theta_1}(dx + \sqrt{-1}dy).$$

We then have

$$\begin{aligned} J^*(e^0 + \sqrt{-1}e^1) &= V^{\frac{1}{2}}e^{-\sqrt{-1}\theta_1}(J dx + \sqrt{-1}J dy) \\ &= V^{\frac{1}{2}}e^{-\sqrt{-1}\theta_1}(-d\theta_2 + \sqrt{-1}V^{-1}\Theta) \\ &= -e^{-\sqrt{-1}\theta_1}(e^2 - \sqrt{-1}e^3). \end{aligned}$$

So if ω_1 is a 1-form of Type I and satisfies $\omega_1 = e^{\sqrt{-1}k\theta_1}u(r)(e^0 + \sqrt{-1}e^1)$, then

$$\begin{aligned} J^*(\omega_1) &= -e^{\sqrt{-1}(k-1)\theta_1}u(r)(e^2 - \sqrt{-1}e^3), \\ J^*(\overline{\omega_1}) &= -e^{\sqrt{-1}(1-k)\theta_1}\overline{u(r)}(e^2 + \sqrt{-1}e^3). \end{aligned}$$

On 1-forms, $\Delta = \nabla^*\nabla$, and since J is parallel, ω_2 is harmonic if and only if ω_1 is harmonic. These generate all Type II harmonic 1-forms. $\square$

A.4. 2-forms on the model space. Let us define the 2-forms

$$\begin{aligned}\omega_{\pm}^1 &= e^0 \wedge e^1 \pm e^2 \wedge e^3, \\ \omega_{\pm}^2 &= e^0 \wedge e^2 \mp e^1 \wedge e^3, \\ \omega_{\pm}^3 &= e^0 \wedge e^3 \pm e^1 \wedge e^2.\end{aligned}$$

Then we have $*\omega_{\pm}^i = \pm\omega_{\pm}^i$, $i = 1, 2, 3$. Note that

$$\begin{aligned}\omega_{\pm}^1 &= Vr dr \wedge d\theta_1 \pm d\theta_2 \wedge \Theta, \\ \omega_{\pm}^2 &= V dr \wedge d\theta_2 \mp r d\theta_1 \wedge \Theta, \\ \omega_{\pm}^3 &= dr \wedge \Theta \pm Vr d\theta_1 \wedge d\theta_2.\end{aligned}$$

We compute the exterior derivatives

$$\begin{aligned}d\omega_{\pm}^1 &= 0, \\ d\omega_{\pm}^2 &= \mp dr \wedge d\theta_1 \wedge \Theta = \mp V^{-\frac{1}{2}} r^{-1} e^0 \wedge e^1 \wedge e^3, \\ d\omega_{\pm}^3 &= \left(-\frac{v}{2\pi} \pm (Vr)'\right) dr \wedge d\theta_1 \wedge d\theta_2 \\ &= \left(-\frac{v}{2\pi} \pm (Vr)'\right) V^{-\frac{3}{2}} r^{-1} e^0 \wedge e^1 \wedge e^2.\end{aligned}$$

If we expand further, we have

$$d\omega_{\pm}^3 = -\left(\frac{v}{\pi} + V\right) dr \wedge d\theta_1 \wedge d\theta_2 = -\left(\frac{v}{\pi} + V\right) V^{-\frac{3}{2}} r^{-1} e^0 \wedge e^1 \wedge e^2.$$

Remark A.6. Recall that $\omega_{\pm}^i$ are not the hyper-Kähler forms. For this, we should define $E^0 = V^{-\frac{1}{2}} dx$, $E^1 = V^{-\frac{1}{2}} dy$. But for the following calculations, we can freely choose an orthonormal basis.

Remark A.7. If ω is self-dual, then we may also write $\omega = a\omega_I + b\omega_J + c\omega_K$, and hence

$$\Delta\omega = (\Delta a)\omega_I + (\Delta b)\omega_J + (\Delta c)\omega_K.$$

Since we already know the result about harmonic functions, we can define

$$(A.8) \quad \mathcal{Z}_{\mathfrak{q}}^{2,+}(\hat{\mathfrak{M}}) \equiv \mathcal{Z}_{\mathfrak{q}}^0(\hat{\mathfrak{M}}) \otimes (\mathbb{R}\omega_I \oplus \mathbb{R}\omega_J \oplus \mathbb{R}\omega_K).$$

In the following, we only need to consider ASD 2-forms.

We want ω to be invariant under the S^1 -action. Let us consider the general ansatz

$$(A.9) \quad \omega = e^{\sqrt{-1}k\theta_1} (a(r)\omega_{-}^1 + b(r)\omega_{-}^2 + c(r)\omega_{-}^3).$$

Then we prove the following lemma.

Lemma A.8. Given an integer $\mathfrak{q} \in \mathbb{Z}$, let $\mathcal{Z}_{\mathfrak{q}}^{2,-}(\hat{\mathfrak{M}})$ be the linear space with the following bases.

(1) For every $\mathfrak{q} \in \mathbb{Z} \setminus \{0, 1, -1\}$,

$$\begin{aligned}
\mathcal{B}_{\mathfrak{q}}^{2,-} \equiv & \left\{ r^{\mathfrak{q}} e^{\sqrt{-1}\mathfrak{q}\theta_1} \omega_-^1, r^{\mathfrak{q}} e^{-\sqrt{-1}\mathfrak{q}\theta_1} \omega_-^1, \right. \\
& \frac{\sqrt{-1}}{\mathfrak{q}-1} r^{\mathfrak{q}} \left(-\frac{(\mathfrak{q}-1)V}{2\mathfrak{q}} + \frac{2\mathfrak{q}-1}{2\mathfrak{q}^2} \frac{\nu}{2\pi} + \frac{(\mathfrak{q}-1)^2}{4\mathfrak{q}^3} \frac{\nu^2}{4\pi^2} V^{-1} \right) e^{\sqrt{-1}(1-\mathfrak{q})\theta_1} \omega_-^2 \\
& + r^{\mathfrak{q}} \left(-\frac{V}{2\mathfrak{q}} + \frac{1}{2\mathfrak{q}^2} \frac{\nu}{2\pi} - \frac{\mathfrak{q}+1}{4\mathfrak{q}^3} \frac{\nu^2}{4\pi^2} V^{-1} + \frac{1}{4\mathfrak{q}^3} \frac{\nu^3}{8\pi^3} V^{-2} \right) e^{\sqrt{-1}(1-\mathfrak{q})\theta_1} \omega_-^3, \\
& - \frac{\sqrt{-1}}{\mathfrak{q}-1} r^{\mathfrak{q}} \left(-\frac{(\mathfrak{q}-1)V}{2\mathfrak{q}} + \frac{2\mathfrak{q}-1}{2\mathfrak{q}^2} \frac{\nu}{2\pi} + \frac{(\mathfrak{q}-1)^2}{4\mathfrak{q}^3} \frac{\nu^2}{4\pi^2} V^{-1} \right) e^{-\sqrt{-1}(1-\mathfrak{q})\theta_1} \omega_-^2 \\
& + r^{\mathfrak{q}} \left(-\frac{V}{2\mathfrak{q}} + \frac{1}{2\mathfrak{q}^2} \frac{\nu}{2\pi} - \frac{\mathfrak{q}+1}{4\mathfrak{q}^3} \frac{\nu^2}{4\pi^2} V^{-1} + \frac{1}{4\mathfrak{q}^3} \frac{\nu^3}{8\pi^3} V^{-2} \right) e^{-\sqrt{-1}(1-\mathfrak{q})\theta_1} \omega_-^3, \\
& V^{-1} r^{\mathfrak{q}} e^{\sqrt{-1}(\mathfrak{q}+1)\theta_1} \omega_-^2 \\
& + \frac{1}{\sqrt{-1}(\mathfrak{q}+1)} V^{-2} r^{\mathfrak{q}} \left((\mathfrak{q}+1)V - \frac{\nu}{2\pi} \right) e^{\sqrt{-1}(\mathfrak{q}+1)\theta_1} \omega_-^3, \\
& V^{-1} r^{\mathfrak{q}} e^{-\sqrt{-1}(\mathfrak{q}+1)\theta_1} \omega_-^2 \\
& \left. - \frac{1}{\sqrt{-1}(\mathfrak{q}+1)} V^{-2} r^{\mathfrak{q}} \left((\mathfrak{q}+1)V - \frac{\nu}{2\pi} \right) e^{-\sqrt{-1}(\mathfrak{q}+1)\theta_1} \omega_-^3 \right\}.
\end{aligned}$$

(2) For $\mathfrak{q} = -1$,

$$\begin{aligned}
\mathcal{B}_{-1}^{2,-} \equiv & \left\{ r^{-1} e^{-\sqrt{-1}\theta_1} \omega_-^1, r^{-1} e^{\sqrt{-1}\theta_1} \omega_-^1, r^{-1} \omega_-^2, V^{-2} r^{-1} \omega_-^3, \right. \\
& \frac{\sqrt{-1}}{2} r^{-1} \left(V + \frac{3\nu}{4\pi} + \frac{\nu^2}{4\pi^2} V^{-1} \right) e^{2\sqrt{-1}\theta_1} \omega_-^2 \\
& + r^{-1} \left(\frac{V}{2} + \frac{\nu}{4\pi} - \frac{\nu^3}{32\pi^3} V^{-2} \right) e^{2\sqrt{-1}\theta_1} \omega_-^3, \\
& - \frac{\sqrt{-1}}{2} r^{-1} \left(V + \frac{3\nu}{4\pi} + \frac{\nu^2}{4\pi^2} V^{-1} \right) e^{-2\sqrt{-1}\theta_1} \omega_-^2 \\
& \left. + r^{-1} \left(\frac{V}{2} + \frac{\nu}{4\pi} - \frac{\nu^3}{32\pi^3} V^{-2} \right) e^{-2\sqrt{-1}\theta_1} \omega_-^3 \right\}.
\end{aligned}$$

(3) For $\mathfrak{q} = 0$,

$$\begin{aligned}
\mathcal{B}_0^{2,-} \equiv & \left\{ \omega_-^1, V e^{\sqrt{-1}\theta_1} \omega_-^1, V e^{-\sqrt{-1}\theta_1} \omega_-^1, \right. \\
& \sqrt{-1} \left(\frac{1}{3} \frac{2\pi}{\nu} V^2 + \frac{1}{2} V + \frac{1}{2} \frac{\nu}{2\pi} \right) e^{\sqrt{-1}\theta_1} \omega_-^2 \\
& + \left(\frac{1}{3} \frac{2\pi}{\nu} V^2 + \frac{1}{6} V \right) e^{\sqrt{-1}\theta_1} \omega_-^3, \\
& - \sqrt{-1} \left(\frac{1}{3} \frac{2\pi}{\nu} V^2 + \frac{1}{2} V + \frac{1}{2} \frac{\nu}{2\pi} \right) e^{-\sqrt{-1}\theta_1} \omega_-^2 \\
& + \left(\frac{1}{3} \frac{2\pi}{\nu} V^2 + \frac{1}{6} V \right) e^{-\sqrt{-1}\theta_1} \omega_-^3, \\
& V^{-1} e^{\sqrt{-1}\theta_1} \omega_-^2 + \frac{1}{\sqrt{-1}} V^{-2} \left(V - \frac{\nu}{2\pi} \right) e^{\sqrt{-1}\theta_1} \omega_-^3, \\
& \left. V^{-1} e^{-\sqrt{-1}\theta_1} \omega_-^2 - \frac{1}{\sqrt{-1}} V^{-2} \left(V - \frac{\nu}{2\pi} \right) e^{-\sqrt{-1}\theta_1} \omega_-^3 \right\}.
\end{aligned}$$

(4) For $\mathfrak{q} = 1$,

$$\begin{aligned} \mathcal{B}_1^{2,-} \equiv & \left\{ r e^{\sqrt{-1}\theta_1} \omega_-^1, r e^{-\sqrt{-1}\theta_1} \omega_-^1, r \omega_-^2, \right. \\ & r \left(\frac{1}{2} V^2 - \frac{\nu}{2\pi} V + \frac{3}{2} \left(\frac{\nu}{2\pi} \right)^2 - \frac{3}{2} \left(\frac{\nu}{2\pi} \right)^3 V^{-1} + \frac{3}{4} \left(\frac{\nu}{2\pi} \right)^4 V^{-2} \right) \omega_-^3, \\ & V^{-1} r e^{2\sqrt{-1}\theta_1} \omega_-^2 + \frac{1}{2\sqrt{-1}} V^{-2} r \left(2V - \frac{\nu}{2\pi} \right) e^{2\sqrt{-1}\theta_1} \omega_-^3, \\ & \left. V^{-1} r e^{-2\sqrt{-1}\theta_1} \omega_-^2 - \frac{1}{2\sqrt{-1}} V^{-2} r \left(2V - \frac{\nu}{2\pi} \right) e^{-2\sqrt{-1}\theta_1} \omega_-^3 \right\}. \end{aligned}$$

If a complex-valued 2-form ω satisfies (A.9) for some $k \in \mathbb{Z}$ and solves $\Delta\omega = 0$ on $\hat{\mathfrak{M}}$, then

$$\omega \in \begin{cases} \bigoplus_{\mathfrak{q}=-2}^2 \mathcal{Z}_{\mathfrak{q}}^{2,-}(\hat{\mathfrak{M}}), & k \in \{-1, 1\}, \\ \bigoplus_{\mathfrak{q}=-1}^1 \mathcal{Z}_{\mathfrak{q}}^{2,-}(\hat{\mathfrak{M}}), & k = 0, \\ \left(\bigoplus_{m=-1}^1 \mathcal{Z}_{k+m}^{2,-}(\hat{\mathfrak{M}}) \right) \oplus \left(\bigoplus_{m=-1}^1 \mathcal{Z}_{-k+m}^{2,-}(\hat{\mathfrak{M}}) \right), & k \in \mathbb{Z} \setminus \{-1, 0, 1\}. \end{cases}$$

Remark A.9. We have that $*\omega = -\omega$ for any ASD 2-form ω , so

$$\Delta\omega = d\delta\omega + \delta d\omega = -d * d * \omega - * d * d\omega = d * d\omega - * d * d\omega = (\text{Id} - *)d * d\omega.$$

Therefore, we only need to compute $d * d\omega$ and then take twice the projection onto the space of ASD 2-forms.

Proof. First, we compute

$$\begin{aligned} d\omega = & e^{\sqrt{-1}k\theta_1} \sqrt{-1}k d\theta_1 \wedge (a(r)\omega_-^1 + b(r)\omega_-^2 + c(r)\omega_-^3) \\ & + e^{\sqrt{-1}k\theta_1} (a'dr \wedge \omega_-^1 + b'dr \wedge \omega_-^2 + c'dr \wedge \omega_-^3) \\ & + e^{\sqrt{-1}k\theta_1} (ad\omega_-^1 + bd\omega_-^2 + cd\omega_-^3). \end{aligned}$$

Using the above, this is

$$\begin{aligned} d\omega = & e^{\sqrt{-1}k\theta_1} \left\{ \sqrt{-1}kaV^{-\frac{1}{2}}r^{-1}e^1 \wedge \omega_-^1 + \sqrt{-1}kbV^{-\frac{1}{2}}r^{-1}e^1 \wedge \omega_-^2 \right. \\ & + \sqrt{-1}kcV^{-\frac{1}{2}}r^{-1}e^1 \wedge \omega_-^3 + a'V^{-\frac{1}{2}}e^0 \wedge \omega_-^1 + b'V^{-\frac{1}{2}}e^0 \wedge \omega_-^2 \\ & + c'V^{-\frac{1}{2}}e^0 \wedge \omega_-^3 + bV^{-\frac{1}{2}}r^{-1}e^0 \wedge e^1 \wedge e^3 \\ & \left. - c\left(\frac{\nu}{\pi} + V\right)V^{-\frac{3}{2}}r^{-1}e^0 \wedge e^1 \wedge e^2 \right\}. \end{aligned}$$

Then we have

$$\begin{aligned} d\omega = & e^{\sqrt{-1}k\theta_1} \left\{ -\sqrt{-1}kaV^{-\frac{1}{2}}r^{-1}e^1 \wedge e^2 \wedge e^3 - \sqrt{-1}kbV^{-\frac{1}{2}}r^{-1}e^0 \wedge e^1 \wedge e^2 \right. \\ & - \sqrt{-1}kcV^{-\frac{1}{2}}r^{-1}e^0 \wedge e^1 \wedge e^3 - a'V^{-\frac{1}{2}}e^0 \wedge e^2 \wedge e^3 \\ & + b'V^{-\frac{1}{2}}e^0 \wedge e^1 \wedge e^3 - c'V^{-\frac{1}{2}}e^0 \wedge e^1 \wedge e^2 \\ & \left. + bV^{-\frac{1}{2}}r^{-1}e^0 \wedge e^1 \wedge e^3 - c\left(\frac{\nu}{\pi} + V\right)V^{-\frac{3}{2}}r^{-1}e^0 \wedge e^1 \wedge e^2 \right\}. \end{aligned}$$

Collecting terms, we have

$$\begin{aligned} d\omega = e^{\sqrt{-1}k\theta_1} & \left\{ -\sqrt{-1}kaV^{-\frac{1}{2}}r^{-1}e^1 \wedge e^2 \wedge e^3 - a'V^{-\frac{1}{2}}e^0 \wedge e^2 \wedge e^3 \right. \\ & + [-\sqrt{-1}kcV^{-\frac{1}{2}}r^{-1} + b'V^{-\frac{1}{2}} + bV^{-\frac{1}{2}}r^{-1}]e^0 \wedge e^1 \wedge e^3 \\ & \left. + \left[-\sqrt{-1}kbV^{-\frac{1}{2}}r^{-1} - c'V^{-\frac{1}{2}} - c\left(\frac{\nu}{\pi} + V\right)V^{-\frac{3}{2}}r^{-1} \right]e^0 \wedge e^1 \wedge e^2 \right\}. \end{aligned}$$

Now let us define

$$\begin{aligned} p(r) &= -\sqrt{-1}kcV^{-\frac{1}{2}}r^{-1} + b'V^{-\frac{1}{2}} + bV^{-\frac{1}{2}}r^{-1}, \\ q(r) &= -\sqrt{-1}kbV^{-\frac{1}{2}}r^{-1} - c'V^{-\frac{1}{2}} - c\left(\frac{\nu}{\pi} + V\right)V^{-\frac{3}{2}}r^{-1} \end{aligned}$$

so that

$$\begin{aligned} d\omega = e^{\sqrt{-1}k\theta_1} & \left\{ -\sqrt{-1}kaV^{-\frac{1}{2}}r^{-1}e^1 \wedge e^2 \wedge e^3 - a'V^{-\frac{1}{2}}e^0 \wedge e^2 \wedge e^3 \right. \\ & \left. + pe^0 \wedge e^1 \wedge e^3 + qe^0 \wedge e^1 \wedge e^2 \right\}. \end{aligned}$$

Then we have

$$*d\omega = e^{\sqrt{-1}k\theta_1} \left\{ \sqrt{-1}kaV^{-\frac{1}{2}}r^{-1}e^0 - a'V^{-\frac{1}{2}}e^1 - pe^2 + qe^3 \right\}.$$

So we have

$$\begin{aligned} d * d\omega = e^{\sqrt{-1}k\theta_1} & \left\{ \sqrt{-1}k d\theta_1 \wedge \left\{ \sqrt{-1}kaV^{-\frac{1}{2}}r^{-1}e^0 - a'V^{-\frac{1}{2}}e^1 - pe^2 + qe^3 \right\} \right. \\ & + e^{\sqrt{-1}k\theta_1} \left\{ -\partial_r(a'V^{-\frac{1}{2}})dr \wedge e^1 - a'V^{-\frac{1}{2}}\partial_r(V^{\frac{1}{2}}r)V^{-1}r^{-1}e^0 \wedge e^1 \right. \\ & \quad - p'dr \wedge e^2 - p\partial_r(V^{\frac{1}{2}})V^{-1}e^0 \wedge e^2 + q'dr \wedge e^3 \\ & \quad \left. \left. + q\left(\partial_r(V^{-\frac{1}{2}})e^0 \wedge e^3 + \frac{\nu}{2\pi}V^{-\frac{3}{2}}r^{-1}e^1 \wedge e^2\right) \right\} \right\}. \end{aligned}$$

Simplifying some,

$$\begin{aligned} d * d\omega = e^{\sqrt{-1}k\theta_1} & \left\{ k^2aV^{-1}r^{-2}e^0 \wedge e^1 - \sqrt{-1}kpV^{-\frac{1}{2}}r^{-1}e^1 \wedge e^2 \right. \\ & + \sqrt{-1}kqV^{-\frac{1}{2}}r^{-1}e^1 \wedge e^3 - \partial_r(a'V^{-\frac{1}{2}})V^{-\frac{1}{2}}e^0 \wedge e^1 \\ & - a'V^{-\frac{1}{2}}\partial_r(V^{\frac{1}{2}}r)V^{-1}r^{-1}e^0 \wedge e^1 - p'V^{-\frac{1}{2}}e^0 \wedge e^2 \\ & - p\partial_r(V^{\frac{1}{2}})V^{-1}e^0 \wedge e^2 + q'V^{-\frac{1}{2}}e^0 \wedge e^3 \\ & \left. + q\left(\partial_r(V^{-\frac{1}{2}})e^0 \wedge e^3 + \frac{\nu}{2\pi}V^{-\frac{3}{2}}r^{-1}e^1 \wedge e^2\right) \right\}. \end{aligned}$$

Collecting terms,

$$\begin{aligned} d * d\omega = e^{\sqrt{-1}k\theta_1} & \left\{ [k^2aV^{-1}r^{-2} - \partial_r(a'V^{-\frac{1}{2}})V^{-\frac{1}{2}} \right. \\ & \quad \left. - a'V^{-\frac{1}{2}}\partial_r(V^{\frac{1}{2}}r)V^{-1}r^{-1}]e^0 \wedge e^1 \right. \\ & + \{-p'V^{-\frac{1}{2}} - p\partial_r(V^{\frac{1}{2}})V^{-1}\}e^0 \wedge e^2 \\ & + \sqrt{-1}kqV^{-\frac{1}{2}}r^{-1}e^1 \wedge e^3 + \{q'V^{-\frac{1}{2}} + q\partial_r(V^{-\frac{1}{2}})\}e^0 \wedge e^3 \\ & \left. + \left\{ -\sqrt{-1}kpV^{-\frac{1}{2}}r^{-1} + q\frac{\nu}{2\pi}V^{-\frac{3}{2}}r^{-1} \right\}e^1 \wedge e^2 \right\}. \end{aligned}$$

Recall that we need to take the projection onto the ASD part of this. So we have the following equations. The ω_-^1 coefficient is

$$k^2 a V^{-1} r^{-2} - \partial_r(a' V^{-\frac{1}{2}}) V^{-\frac{1}{2}} - a' V^{-\frac{1}{2}} \partial_r(V^{\frac{1}{2}} r) V^{-1} r^{-1},$$

which simplifies to $-V^{-1}(a'' + r^{-1}a' - k^2 r^{-2}a)$. So we have the formula

$$\Delta(e^{\sqrt{-1}k\theta_1} a(r) \omega_-^1) = -e^{\sqrt{-1}k\theta_1} (a'' + r^{-1}a' - k^2 r^{-2}a) V^{-1} \omega_-^1.$$

Recall that solutions of the ODE $a'' + r^{-1}a' - k^2 r^{-2}a = 0$ are given by

$$a(r) = \begin{cases} C_1 + C_2 V, & k = 0, \\ C_1 r^k + C_2 r^{-k}, & k \neq 0. \end{cases}$$

These prove the ω_-^1 part in Lemma A.8. Next, the ω_-^2 component is

$$\begin{aligned} (\text{Id} - *)e^{\sqrt{-1}k\theta_1} \{(-p' V^{-\frac{1}{2}} - p \partial_r(V^{\frac{1}{2}}) V^{-1}) e^0 \wedge e^2 + \sqrt{-1}kq V^{-\frac{1}{2}} r^{-1} e^1 \wedge e^3\} \\ = e^{\sqrt{-1}k\theta_1} \{(-p' V^{-\frac{1}{2}} - p \partial_r(V^{\frac{1}{2}}) V^{-1}) \omega_-^2 + \sqrt{-1}kq V^{-\frac{1}{2}} r^{-1} \omega_-^2\} \\ = e^{\sqrt{-1}k\theta_1} (-p' V^{-\frac{1}{2}} - p \partial_r(V^{\frac{1}{2}}) V^{-1} + \sqrt{-1}kq V^{-\frac{1}{2}} r^{-1}) \omega_-^2 \\ = -e^{\sqrt{-1}k\theta_1} \left(p' + \frac{1}{2}(\log(V))' p - \sqrt{-1}kqr^{-1} \right) V^{-\frac{1}{2}} \omega_-^2. \end{aligned}$$

For the ω_-^3 component, we have

$$\begin{aligned} (\text{Id} - *)e^{\sqrt{-1}k\theta_1} \left\{ (q' V^{-\frac{1}{2}} + q \partial_r(V^{-\frac{1}{2}})) e^0 \wedge e^3 \right. \\ \left. + \left(-\sqrt{-1}kp V^{-\frac{1}{2}} r^{-1} + q \frac{\nu}{2\pi} V^{-\frac{3}{2}} r^{-1} \right) e^1 \wedge e^2 \right\} \\ = e^{\sqrt{-1}k\theta_1} \left\{ (q' V^{-\frac{1}{2}} + q \partial_r(V^{-\frac{1}{2}})) \omega_-^3 \right. \\ \left. - \left(-\sqrt{-1}kp V^{-\frac{1}{2}} r^{-1} + q \frac{\nu}{2\pi} V^{-\frac{3}{2}} r^{-1} \right) \omega_-^3 \right\} \\ = e^{\sqrt{-1}k\theta_1} \left\{ q' V^{-\frac{1}{2}} + q \partial_r(V^{-\frac{1}{2}}) + \sqrt{-1}kp V^{-\frac{1}{2}} r^{-1} - q \frac{\nu}{2\pi} V^{-\frac{3}{2}} r^{-1} \right\} \omega_-^3 \\ = e^{\sqrt{-1}k\theta_1} \left\{ q' - \frac{3}{2} \frac{\nu}{2\pi} V^{-1} r^{-1} q + \sqrt{-1}kpr^{-1} \right\} V^{-\frac{1}{2}} \omega_-^3 \\ = e^{\sqrt{-1}k\theta_1} \left\{ q' - \frac{3}{2}(\log(V))' q + \sqrt{-1}kpr^{-1} \right\} V^{-\frac{1}{2}} \omega_-^3. \end{aligned}$$

Consequently, we have the formula

$$\begin{aligned} \Delta(e^{\sqrt{-1}k\theta_1} (b(r) \omega_-^2 + c(r) \omega_-^3)) \\ = e^{\sqrt{-1}k\theta_1} \left\{ -\left(p' + \frac{1}{2}(\log(V))' p - \sqrt{-1}kqr^{-1} \right) V^{-\frac{1}{2}} \omega_-^2 \right. \\ \left. + \left(q' - \frac{3}{2}(\log(V))' q + \sqrt{-1}kpr^{-1} \right) V^{-\frac{1}{2}} \omega_-^3 \right\}. \end{aligned}$$

The harmonic condition becomes the following first-order system for p and q :

$$\begin{aligned} \partial_r p + \frac{1}{2}(\log(V))' p - \sqrt{-1}kqr^{-1} &= 0, \\ \partial_r q - \frac{3}{2}(\log(V))' q + \sqrt{-1}kpr^{-1} &= 0. \end{aligned}$$

Define $\tilde{p} = V^{\frac{1}{2}} \cdot p$, $\tilde{q} = V^{-\frac{3}{2}} \cdot q$. Then the equations become

$$\begin{aligned}\tilde{p}' - \sqrt{-1}kr^{-1}V^2\tilde{q} &= 0, \\ \tilde{q}' + \sqrt{-1}kr^{-1}V^{-2}\tilde{p} &= 0.\end{aligned}$$

If $k = 0$, then we have $(\tilde{p}, \tilde{q}) = (C_1, C_2)$, so $(p, q) = (C_1V^{-\frac{1}{2}}, C_2V^{\frac{3}{2}})$. So next, we assume that $k \neq 0$. To solve this, let us differentiate the second equation to get

$$\tilde{q}'' + \sqrt{-1}k(r^{-1}V^{-2})'\tilde{p} + \sqrt{-1}kr^{-1}V^{-2}\tilde{p}' = 0.$$

Using the system again, this is

$$\begin{aligned}0 &= \tilde{q}'' + \sqrt{-1}k(r^{-1}V^{-2})'\frac{1}{\sqrt{-1}k}rV^2(-\tilde{q})' + \sqrt{-1}kr^{-1}V^{-2}\sqrt{-1}kr^{-1}V^2\tilde{q} \\ &= \tilde{q}'' - (r^{-1}V^{-2})'rV^2\tilde{q}' - k^2r^{-2}\tilde{q}.\end{aligned}$$

Some simplification yields

$$\tilde{q}'' + \left(\frac{1}{r} + 2(\log(V))'\right)\tilde{q}' - k^2r^{-2}\tilde{q} = 0.$$

For $k \neq 0$, the solutions are $\tilde{q} = C_1r^{-k}V^{-1} + C_2r^kV^{-1}$. Then we can use the second equation

$$\tilde{p} = -\frac{1}{\sqrt{-1}k}rV^2\tilde{q}'$$

to solve for $\tilde{p}$. This yields

$$\begin{aligned}\tilde{p} &= -\frac{1}{\sqrt{-1}k}rV^2\left((-kC_1r^{-k-1} + kC_2r^{k-1})V^{-1} - (C_1r^{-k} + C_2r^k)V^{-2}\frac{v}{2\pi}r^{-1}\right) \\ &= -\frac{1}{\sqrt{-1}k}\left((-kC_1r^{-k} + kC_2r^k)V - \frac{v}{2\pi}(C_1r^{-k} + C_2r^k)\right) \\ &= -\frac{1}{\sqrt{-1}k}\left\{C_1\left(-kV - \frac{v}{2\pi}\right)r^{-k} + C_2\left(kV - \frac{v}{2\pi}\right)r^k\right\}.\end{aligned}$$

So to summarize, we have $(\tilde{p}, \tilde{q}) = (C_1, C_2)$ if $k = 0$, and

$$(\tilde{p}, \tilde{q}) = C_1\left(\frac{\sqrt{-1}}{k}\left(-kV - \frac{v}{2\pi}\right)r^{-k}, r^{-k}V^{-1}\right) + C_2\left(\frac{\sqrt{-1}}{k}\left(kV - \frac{v}{2\pi}\right)r^k, r^kV^{-1}\right)$$

if $k \neq 0$.

Next, recall the main system is

$$\begin{aligned}-\sqrt{-1}kcV^{-\frac{1}{2}}r^{-1} + b'V^{-\frac{1}{2}} + bV^{-\frac{1}{2}}r^{-1} &= p, \\ -\sqrt{-1}kbV^{-\frac{1}{2}}r^{-1} - c'V^{-\frac{1}{2}} - c\left(\frac{v}{\pi} + V\right)V^{-\frac{3}{2}}r^{-1} &= q.\end{aligned}$$

The system for b and c then becomes

$$\begin{aligned}\text{(A.10)} \quad b' + r^{-1}b - \sqrt{-1}kr^{-1}c &= V^{\frac{1}{2}}p, \\ c' + c\left(\frac{v}{\pi} + V\right)V^{-1}r^{-1} + \sqrt{-1}kr^{-1}b &= -V^{\frac{1}{2}}q.\end{aligned}$$

We let $\tilde{b} = rb$, $\tilde{c} = V^2rc$. Then the system becomes

$$\begin{aligned}\tilde{b}' - \sqrt{-1}kV^{-2}r^{-1}\tilde{c} &= V^{\frac{1}{2}}rp, \\ \tilde{c}' + \sqrt{-1}kV^2r^{-1}\tilde{b} &= -V^{\frac{5}{2}}rq.\end{aligned}$$

A.4.1. $k = 0$. In this case, we know from above that $(p, q) = (C_1 V^{-\frac{1}{2}}, C_2 V^{\frac{3}{2}})$, so the system becomes $\tilde{b}' = C_1 r$, $\tilde{c}' = -C_2 V^4 r$. The solution is

$$\begin{aligned}\tilde{b} &= A_1 + \frac{C_1}{2} r^2, \\ \tilde{c} &= A_2 - C_2 r^2 \left(\frac{1}{2} V^4 - \frac{v}{2\pi} V^3 + \frac{3}{2} \left(\frac{v}{2\pi} \right)^2 V^2 - \frac{3}{2} \left(\frac{v}{2\pi} \right)^3 V + \frac{3}{4} \left(\frac{v}{2\pi} \right)^4 \right).\end{aligned}$$

So then

$$\begin{aligned}b &= A_1 r^{-1} + \frac{C_1}{2} r, \\ c &= A_2 V^{-2} r^{-1} - C_2 r \left(\frac{1}{2} V^2 - \frac{v}{2\pi} V + \frac{3}{2} \left(\frac{v}{2\pi} \right)^2 - \frac{3}{2} \left(\frac{v}{2\pi} \right)^3 V^{-1} + \frac{3}{4} \left(\frac{v}{2\pi} \right)^4 V^{-2} \right).\end{aligned}$$

A.4.2. $k \neq 0$. If $k \neq 0$, then we differentiate the first equation of

$$\begin{aligned}\tilde{b}' - \sqrt{-1} k V^{-2} r^{-1} \tilde{c} &= V^{\frac{1}{2}} r p, \\ \tilde{c}' + \sqrt{-1} k V^2 r^{-1} \tilde{b} &= -V^{\frac{5}{2}} r q\end{aligned}$$

and use the second equation and the system for (p, q) to get (after some computation)

$$\tilde{b}'' + V^{-1} r^{-1} \left(\frac{v}{\pi} + V \right) \tilde{b}' - k^2 r^{-2} \tilde{b} = V^{-\frac{1}{2}} \left(\frac{v}{\pi} + 2V \right) p.$$

Let us make the substitution $\hat{b} = V \tilde{b}$ to get the equation

$$\hat{b}'' + r^{-1} \hat{b}' - k^2 r^{-2} \hat{b} = V^{\frac{1}{2}} \left(\frac{v}{\pi} + 2V \right) p.$$

The homogeneous equation is

$$\hat{b}'' + r^{-1} \hat{b}' - k^2 r^{-2} \hat{b} = 0,$$

which has solutions $\hat{b} = A_1 r^k + A_2 r^{-k}$, equivalently, $b = V^{-1} (A_1 r^{k-1} + A_2 r^{-k-1})$. We can then solve for c using equation (A.10) to get

$$c = \frac{1}{\sqrt{-1} k} \left(A_1 V^{-2} r^{k-1} \left(kV - \frac{v}{2\pi} \right) + A_2 V^{-2} r^{-k-1} \left(-kV - \frac{v}{2\pi} \right) \right).$$

Next, we find a particular solution of

$$\hat{b}'' + r^{-1} \hat{b}' - k^2 r^{-2} \hat{b} = \left(\frac{v}{\pi} + 2V \right) \tilde{p},$$

where

$$\tilde{p} = C_1 \frac{\sqrt{-1}}{k} \left(-kV - \frac{v}{2\pi} \right) r^{-k} + C_2 \frac{\sqrt{-1}}{k} \left(kV - \frac{v}{2\pi} \right) r^k.$$

We define $h = r^{-k} \hat{b}$. Then, after some computation, the equation becomes

$$h'' + (2k + 1) r^{-1} h' = \left(\frac{v}{\pi} + 2V \right) r^{-k} \tilde{p}.$$

Multiplying by r^{2k+1} , we may write this as

$$(r^{2k+1} h')' = \left(\frac{v}{\pi} + 2V \right) r^{k+1} \tilde{p}.$$

This can be integrated to yield

$$h' = r^{-2k-1} \int \left(\frac{v}{\pi} + 2V \right) r^{k+1} \tilde{p} dr.$$

We can integrate again to obtain

$$h = \int r^{-2k-1} \int \left(\frac{v}{\pi} + 2V \right) r^{k+1} \tilde{p} dr.$$

Converting back to b , we have

$$b = V^{-1} r^{k-1} h = V^{-1} r^{k-1} \int r^{-2k-1} \int \left(\frac{v}{\pi} + 2V \right) r^{k+1} \tilde{p} dr.$$

When $k \neq \pm 1$, we find that

$$\begin{aligned} b = & \frac{C_1 \sqrt{-1}}{k} r^{-k+1} \left(\frac{kV}{2k-2} + \frac{2k-1}{2(k-1)^2} \frac{v}{2\pi} + \frac{k^2}{4(k-1)^3} \frac{v^2}{4\pi^2} V^{-1} \right) \\ & + \frac{C_2 \sqrt{-1}}{k} r^{k+1} \left(\frac{kV}{2k+2} + \frac{-2k-1}{2(k+1)^2} \frac{v}{2\pi} - \frac{k^2}{4(k+1)^3} \frac{v^2}{4\pi^2} V^{-1} \right). \end{aligned}$$

By (A.10),

$$\begin{aligned} c = & C_1 r^{-k+1} \left(\frac{V}{2k-2} + \frac{1}{2(k-1)^2} \frac{v}{2\pi} - \frac{k-2}{4(k-1)^3} \frac{v^2}{4\pi^2} V^{-1} - \frac{1}{4(k-1)^3} \frac{v^3}{8\pi^3} V^{-2} \right) \\ & + C_2 r^{k+1} \left(\frac{V}{-2k-2} + \frac{1}{2(k+1)^2} \frac{v}{2\pi} \right. \\ & \quad \left. - \frac{k+2}{4(k+1)^3} \frac{v^2}{4\pi^2} V^{-1} + \frac{1}{4(k+1)^3} \frac{v^3}{8\pi^3} V^{-2} \right). \end{aligned}$$

When $k = 1$,

$$\begin{aligned} b = & -C_1 \sqrt{-1} \left(\frac{1}{3} \frac{2\pi}{v} V^2 + \frac{1}{2} V + \frac{1}{2} \frac{v}{2\pi} \right) + C_2 \sqrt{-1} r^2 \left(\frac{V}{4} - \frac{3}{8} \frac{v}{2\pi} - \frac{1}{32} \frac{v^2}{4\pi^2} V^{-1} \right), \\ c = & -C_1 \left(\frac{1}{3} \frac{2\pi}{v} V^2 + \frac{1}{6} V \right) + C_2 r^2 \left(-\frac{V}{4} + \frac{1}{8} \frac{v}{2\pi} - \frac{3}{32} \frac{v^2}{4\pi^2} V^{-1} + \frac{1}{32} \frac{v^3}{8\pi^3} V^{-2} \right). \end{aligned}$$

When $k = -1$,

$$\begin{aligned} b = & -C_1 \sqrt{-1} r^2 \left(\frac{V}{4} - \frac{3}{8} \frac{v}{2\pi} - \frac{1}{32} \frac{v^2}{4\pi^2} V^{-1} \right) + C_2 \sqrt{-1} \left(\frac{1}{3} \frac{2\pi}{v} V^2 + \frac{1}{2} V + \frac{1}{2} \frac{v}{2\pi} \right), \\ c = & C_1 r^2 \left(-\frac{V}{4} + \frac{1}{8} \frac{v}{2\pi} - \frac{3}{32} \frac{v^2}{4\pi^2} V^{-1} + \frac{1}{32} \frac{v^3}{8\pi^3} V^{-2} \right) - C_2 \left(\frac{1}{3} \frac{2\pi}{v} V^2 + \frac{1}{6} V \right). \end{aligned}$$

Notice that the Laplacian is a real operator, so we also have conjugate solutions. This completes the proof. $\square$

Acknowledgement. The authors would like to thank Hans-Joachim Hein, Rafe Mazzeo, and Song Sun for numerous helpful discussions. We also would like to thank the anonymous referee for their helpful questions and insightful remarks which greatly improved the exposition of the paper.

References

- [1] *M. T. Anderson*, On the topology of complete manifolds of nonnegative Ricci curvature, *Topology* **29** (1990), no. 1, 41–55.
- [2] *S. Bando*, *A. Kasue* and *H. Nakajima*, On a construction of coordinates at infinity on manifolds with fast curvature decay and maximal volume growth, *Invent. Math.* **97** (1989), no. 2, 313–349.
- [3] *R. Bartnik*, The mass of an asymptotically flat manifold, *Comm. Pure Appl. Math.* **39** (1986), no. 5, 661–693.
- [4] *O. Biquard* and *V. Minerbe*, A Kummer construction for gravitational instantons, *Comm. Math. Phys.* **308** (2011), no. 3, 773–794.
- [5] *G. Chen* and *X. Chen*, Gravitational instantons with faster than quadratic curvature decay (II), *J. reine angew. Math.* **756** (2019), 259–284.
- [6] *G. Chen* and *X. Chen*, Gravitational instantons with faster than quadratic curvature decay. I, *Acta Math.* **227** (2021), no. 2, 263–307.
- [7] *G. Chen* and *X. Chen*, Gravitational instantons with faster than quadratic curvature decay (III), *Math. Ann.* **380** (2021), no. 1–2, 687–717.
- [8] *G. Chen*, *J. Viaclovsky* and *R. Zhang*, Collapsing Ricci-flat metrics on elliptic K3 surfaces, *Comm. Anal. Geom.* **28** (2020), no. 8, 2019–2133.
- [9] *S. A. Cherkis* and *A. Kapustin*, Hyper-Kähler metrics from periodic monopoles, *Phys. Rev. D* (3) **65** (2002), no. 8, Article ID 084015.
- [10] *H. Y. Chu* and *J. Shin*, Free actions of finite groups on the 3-dimensional nilmanifold, *Topology Appl.* **144** (2004), no. 1–3, 255–270.
- [11] *T. Hausel*, *E. Hunsicker* and *R. Mazzeo*, Hodge cohomology of gravitational instantons, *Duke Math. J.* **122** (2004), no. 3, 485–548.
- [12] *S. W. Hawking*, Gravitational instantons, *Phys. Lett. A* **60** (1977), no. 2, 81–83.
- [13] *H.-J. Hein*, Gravitational instantons from rational elliptic surfaces, *J. Amer. Math. Soc.* **25** (2012), no. 2, 355–393.
- [14] *H.-J. Hein*, *S. Sun*, *J. Viaclovsky* and *R. Zhang*, Nilpotent structures and collapsing Ricci-flat metrics on the K3 surface, *J. Amer. Math. Soc.* **35** (2022), no. 1, 123–209.
- [15] *N. J. Hitchin*, The self-duality equations on a Riemann surface, *Proc. Lond. Math. Soc.* (3) **55** (1987), no. 1, 59–126.
- [16] *N. J. Hitchin*, *A. Karlhede*, *U. Lindström* and *M. Roček*, Hyper-Kähler metrics and supersymmetry, *Comm. Math. Phys.* **108** (1987), no. 4, 535–589.
- [17] *R. Mazzeo*, Elliptic theory of differential edge operators. I, *Comm. Partial Differential Equations* **16** (1991), no. 10, 1615–1664.
- [18] *R. B. Melrose*, The Atiyah–Patodi–Singer index theorem, *Res. Notes Math.* **4**, A K Peters, Wellesley 1993.
- [19] *V. Minerbe*, A mass for ALF manifolds, *Comm. Math. Phys.* **289** (2009), no. 3, 925–955.
- [20] *S. Salamon*, Riemannian geometry and holonomy groups, *Pitman Res. Notes in Math.s Ser.* **201**, Longman Scientific & Technical, Harlow 1989.

Gao Chen, Institute of Geometry and Physics, University of Science and Technology of China,
Shanghai 201315, P. R. China
e-mail: chengao1@ustc.edu.cn

Jeff Viaclovsky, Department of Mathematics, University of California, Irvine, CA 92697, USA
e-mail: jviaclov@uci.edu

Ruobing Zhang, Department of Mathematics, Princeton University, Princeton, NJ 08544, USA
e-mail: ruobingz@princeton.edu

Eingegangen 13. Oktober 2021, in revidierter Fassung 9. Januar 2023