Probabilistic Seismic Source Inversion of the 1886 Charleston, South Carolina, Earthquake from Macroseismic Evidence: A Major Updating

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ABSTRACT

The source of the 1886 Charleston, South Carolina earthquake influences the computed seismic hazard of the Southeastern U.S. and thus impacts public policy and engineering practice. However, because the 1886 rupture predated seismic instruments, its source is highly uncertain. This study presents probabilistic seismic-source inversions of the Charleston earthquake from liquefaction evidence and historical intensity reports. Using the latest predictive models and a novel inversion approach, we seek to constrain the magnitude, location, and orientation of the 1886 rupture. Probability distributions of rupture magnitude are conditioned on both the "Woodstock Fault" – a commonly inferred source of the 1886 event – and on an unknown source, wherein the uncertainties of fault location and orientation are considered. These distributions are compared to the $M_w6.7$-$M_w7.5$ distribution adopted by the U.S. National Seismic Hazard Model Project (NSHMP). Collectively, the results do not provide strong support for the hypothesized Woodstock Fault. This is not to say the Woodstock Fault does not exist or could not have caused the 1886 rupture, but rather, that the position of the 1886 source model cannot be constrained by the data and models studied herein, given the large uncertainties inherent to each. While this is at odds with the underlying assumption of prior studies, the results nonetheless generally uphold the magnitude distribution adopted by the NSHMP. The largest uncertainties inherent to this distribution are identified and could be diminished in the future. Finally, we note that the inversion methodology used here is not specific to any region, or to certain types of evidence, but can be applied to any seismic zone and to any co-seismic response. This methodology allows for uncertainty to be accounted for in a more complete and transparent manner when inverting seismic source parameters from macroseismic data. Of course, any limitations, biases, or unmodeled uncertainties inherent to these data must be understood, and their implications acknowledged, as further discussed herein.

Keywords: seismic hazard, macroseismic intensity, soil liquefaction, inverse analysis

1. Introduction

Computed seismic hazards are especially uncertain in regions of infrequent seismicity, where the return periods of moderate-to-large earthquakes may exceed the historic observational period. To reduce this uncertainty, engineering geologists routinely perform forensic analyses of the macroseismic evidence (e.g.,

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32   liquefaction, landslides, intensity reports) produced by prehistoric and pre-instrumental earthquakes (e.g.,

33   among many, Obermeier, 1998; Schneider et al., 2001; Green et al., 2005; Kuhn, 2005; Obermeier et al.,

34   2005; Olson et al., 2005; Rodriguez-Marek and Ciani, 2008; Maurer et al., 2015a; Gheibi and Gassman,

35   2016; Yousuf et al., 2021; Chung et al., 2021; Rasanen et al., 2021; Bwambale et al., 2022). The goal of

36   these analyses, in effect, is to constrain the seismic-source parameters of paleoearthquakes, such that these

37   parameters may be input to probabilistic seismic hazard analyses. It follows that computed seismic hazards

38   are, in some regions, heavily influenced by analyses of macroseismic evidence. The South Carolina Coastal

39   Plain is one such region and is strongly influenced by interpretations of the 1886 Charleston earthquake.

40   The 1886 event induced widespread soil liquefaction across the Coastal Plain (Amick et al., 1990),

41   damaged structures in multiple U.S. states, including most structures in Charleston (Dutton, 1889; Wong et

42   al., 2005), produced perceptible shaking over 1500 km away in Canada (Bakun et al., 2002), and was larger

43   in magnitude than any earthquake to since occur in the Southeastern U.S. The source of the Charleston

44   event is thus a major seismic hazard for the region. A 2005 study, for example, predicted that a repeat of

45   the 1886 event would cause 900 deaths, 44,000 injuries, and economic losses of $20 billion in South

46   Carolina alone (Wong et al., 2005). In turn, the Charleston source controls the computed seismic hazard for

47   much of the Southeastern U.S., particularly for long-period structures (Petersen et al., 2020), and thus

48   impacts building codes, governing policies, and engineering practice. However, because the 1886 rupture

49   predated seismic instruments and did not manifest at the surface, its exact location and magnitude remain

50   uncertain, as do the regional amplitudes of resultant ground motions. To constrain these unknowns, and

51   thus prepare for a similar event, numerous researchers have studied macroseismic evidence, as summarized

52   in Table 1. Published confidence intervals (CIs) of the 1886 magnitude range from $M_w6.4$ to $M_w7.8$, as

53   interpreted from intensity reports (e.g., Bakun and Hopper, 2004; Cramer and Boyd, 2014) and liquefaction

54   evidence (e.g., Martin and Clough, 1994; Hayati and Andrus, 2008). Liquefaction features also suggest a

55   history of recurrent earthquakes in the region extending back 6,000 years (Talwani and Schaeffer, 2001)

56   with wide-ranging magnitude estimates of $M_w5.1$ to $M_w7.8$ (e.g., Hu et al., 2002; Gheibi et al., 2020).

57   Considering the existing literature, Petersen et al. (2014, 2020) assigned to the Charleston seismic zone a

58   magnitude probability distribution that ranged from $M_w6.7$ to $M_w7.5$ in the most recent U.S. National

59   Seismic Hazard Model Project (NSHMP) maps.

60   While much has been learned about the 1886 Charleston earthquake, prior analyses of the macroseismic

61   data (i.e., intensity reports, soil liquefaction) have several limitations. First, the analysis of this data has

62   multiple uncertainties, yet existing studies tend either to be deterministic or to account for uncertainties

63   informally. That is, they generally provide either a median estimate of the rupture magnitude or uncertainty

64   bounds that are nominal in nature. It is often unclear what the bounds are, exactly, and which uncertainties

65   are, and are not, accounted for. Second, most analyses assume that the 1886 event was caused by a particular

66   fault (i.e., the "Woodstock Fault") with known characteristics, even though the fault(s) responsible for the

67   event are debated and the characteristics of the Woodstock Fault are uncertain. The feasibility of the data

68   to constrain the source model has arguably not been fully explored, given that nearly all studies provide a

69   magnitude estimate conditioned on a single fault and do not investigate the uncertainty of this assumption.

70   Third, the inverse analysis of intensity and liquefaction data requires a series of models for predicting these

71   phenomena. Regionally distributed ground motions must be predicted, conditioned on a hypothetical

72   source, to include site-response effects at the locations of study. The probability of field observations (i.e.,

73   the observed intensity or liquefaction response) must then be computed, conditioned on the expected ground

74   motions. In this regard, major modeling advances have recently been made. The NGA-East project (Goulet

75   et al., 2018) resulted in the most advanced understanding of Eastern North America (ENA) ground motions

76   and site response (Harmon et al., 2019) to date. Models for correlating ground motions to macroseismic

77   intensities, including ENA-specific relationships, have been updated (e.g., Cramer, 2020). And models for

78   predicting the probability of liquefaction surface expression have been trained using all liquefaction case

79   histories globally compiled to date (Geyin and Maurer, 2020).

Table 1. Prior estimates of the 1886 Charleston earthquake magnitude; estimates are in moment magnitude (M_w) and ranges are at the 95% confidence level, unless noted otherwise (m_b = body wave magnitude; M_s = surface wave magnitude).

Study	Study Type	Magnitude
Bollinger (1977)	MMI	6.8-7.1 (m_b)*
Nuttli et al. (1986)	MMI	6.7 (m_b), 7.7 (M_s)[†]
Martin and Clough (1994)	Liquefaction	7.0-7.5[†]
Johnston (1996)	MMI	6.8-7.8
Bakun and Hopper (2004)	MMI	6.4-7.2
Heidari and Andrus (2010)	Liquefaction	6.8-7.0[†]
Cramer and Boyd (2014)	MMI	6.7-7.3

  * Upper bound magnitude estimate (range is not at the 95% confidence level)
  [†] Magnitude range is either not given or is not at the 95% confidence level

80

81   Accordingly, this study presents probabilistic seismic-source inversions of the 1886 Charleston

82   earthquake from historical intensity reports and liquefaction evidence. Each is studied using a novel

83   approach wherein the above shortcomings are addressed directly. With this approach, the likelihood of a

84   rupture with some location, geometry, and magnitude to produce a set of field observations (observed

85   intensities or liquefaction responses) is computed. Repeating for enumerable hypothetical faults results in

86   a regional scale understanding of the likely source parameters, to the degree the observational data permits.

87   Probability distributions of earthquake magnitude, conditioned on both an unknown source and on the

88   Woodstock Fault, are computed and compared to that used to develop the NSHMP maps (Petersen et al.,

89 2014, 2020). In the following, prior studies of the 1886 macroseismic data (see Table 1) are discussed

90 further. An overview of the analysis methodology is then presented, followed by implementation details.

91 Lastly, the macroseismic data are analyzed and a variety of results are presented and discussed.

92 **2. Prior analyses of 1886 macroseismic evidence**

93 *2.1 Analyses of MMI data*

94 Following the 1886 earthquake, Dutton (1889) compiled intensity reports throughout ENA and developed

95 isoseismal maps based on the Rossi-Forel intensity scale. Researchers have since reinterpreted these reports

96 to the Modified Mercalli Intensity (MMI) scale and analyzed them to infer seismic parameters. Bakun et

97 al. (2002), for example, compiled 1,034 MMI observations from these and other original reports. Prior

98 analyses of the 1886 MMI data have typically used intensity-prediction equations (IPEs), which predict

99 intensity as a function of rupture magnitude and site-to-source distance. Assuming some source location

100 and adopting an IPE, researchers have constrained the causative M_w which best fits the MMI data (e.g.,

101 Bollinger, 1977; Nuttli et al., 1986; Johnston, 1996; Bakun and Hopper, 2004) as summarized in Table 1.

102 Bakun and Hopper (2004), for example, developed an ENA-specific IPE, applied it to the Bakun et al.

103 (2002) MMI data, and reported a magnitude of $M_w6.9$ ($M_w6.4$-7.2 at the 95% confidence level). Using

104 different methods than prior researchers, Cramer and Boyd (2014) compared the mean MMI of the Bakun

105 et al. (2002) dataset against those from two reference events in similar tectonic settings ($M_w7.2$ 1929 Grand

106 Banks, Canada and $M_w7.6$ 2001 Bhuj, India) over a site-to-source distance of 600-1200 km. With this

107 approach, Cramer and Boyd (2014) estimated a median magnitude of $M_w7.0$ with uncertainty of $\pm 0.3M_w$.

108 Collectively, existing studies of the MMI data have reported estimates of $M_w6.4$ to $M_w7.8$. In producing

109 such estimates, these studies have generally assumed that the source was epicentrally located in the vicinity

110 of what is typically called the Woodstock Fault, an inferred N striking, W dipping fault ~25 km NW of

111 Charleston (e.g., Durá-Gómez and Talwani, 2009a,b; Chapman et al., 2016). The fault's surface projection,

112 as hypothesized by Durá-Gómez and Talwani (2009a,b), is mapped in Figure 1. While most studies based

113 on geophysical investigations or modern seismological data have supported this proposed alignment – at

114 least in a general sense (e.g., Pratt et al., 2022) – dramatically different hypotheses for the 1886 earthquake

115 have also been proposed. Marple and Hurd (2020), for example, recently suggested that the 40-km long

116 "Deer Park lineament," which is oriented roughly E-W, may have been responsible.
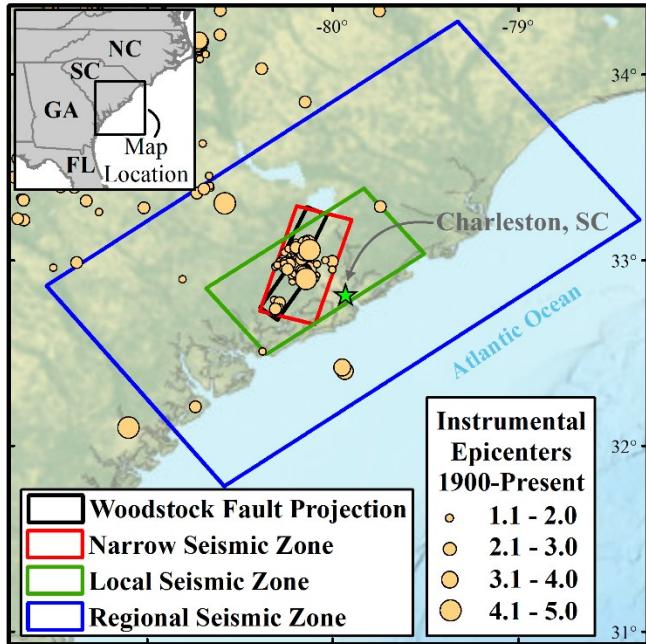
Figure 1. Woodstock Fault projection (Durá-Gómez and Talwani, 2009a,b) and the zonal weighting scheme assigned to the Charleston Seismic Zone in the most recent U.S. National Seismic Hazard Model Project (NSHMP) maps (Petersen et al., 2014, 2020).

117   Also shown in Figure 1 is the zonal weighting scheme assigned to the Charleston Seismic Zone by the
118   NSHMP (Petersen et al., 2014, 2020). In this scheme, a "Narrow" zone with weight of 0.3 delineates the
119   hypothesized Woodstock Fault while accounting for uncertainties in its position and branching structure.
120   The "Local" and "Regional" zones, with respective weights of 0.5 and 0.2, collectively extend offshore and
121   across the extents of the South Carolina Coastal Plain, thereby encompassing more distal faults and
122   liquefaction features that have not been tied to the 1886 event. Each zone is assigned the same $M_w6.7$ to
123    $M_w7.5$ probability distribution by Petersen et al. (2014, 2020). This weighting scheme was adopted from
124   the Central and Eastern United States Seismic Source Characterization for Nuclear Facilities Project
125   (Coppersmith et al., 2012), who concluded: "Neither the 1886 nor the prehistoric (i.e., pre-1886)
126   earthquakes in the Charleston area can be definitively attributed to any specific fault or fault zone at the
127   present time." In arriving at this conclusion, Coppersmith et al. (2012) noted: "the Charleston region is
128   associated with a pattern of observed seismicity that is not particularly remarkable for drawing attention to
129   the location of the 1886 earthquake." As summarized by Chapman et al. (2016), very different hypotheses
130   have arisen from analyses of microseismicity data. While it should be emphasized that the NSHMP scheme
131   does not describe the uncertainty of the 1886 rupture specifically, it does reflect the overall uncertainty of
132   moderate-to-large "1886-like" events in the region. Thus, while prior studies have generally assumed that
133   the 1886 source was in the "Narrow" zone, a scientific consensus has not been reached.

134　　　　As a precursor to other analyses presented herein, the Bakun and Hopper (2004) IPE was first
135　　reimplemented on the Bakun et al. (2002) dataset and the same $M_w6.9$ estimate as Bakun and Hopper (2004)
136　　was obtained, indicating that the approach and dataset were correctly reproduced. Next, this approach was
137　　updated using the newest ENA IPE (Atkinson et al., 2014). Assuming the same source location and studying
138　　MMI data within 1000 km (the applicable distance of the Atkinson et al. (2014) IPE) and within all distances
139　　(to mirror Bakun and Hopper, 2004), median estimates of $M_w8.0$ and $M_w8.2$ were respectively obtained.
140　　The cause of the discrepant results obtained using Bakun and Hopper (2004) vs. Atkinson et al. (2014) can
141　　be seen in Figure 2, where both IPEs are plotted for three values of M_w. For a given M_w and epicentral
142　　distance, the Atkinson et al. (2014) model tends to predict a lesser MMI, indicating that a larger earthquake
143　　magnitude (i.e., $M_w8.0$ - $M_w8.2$) is needed to produce the same set of MMI observations.
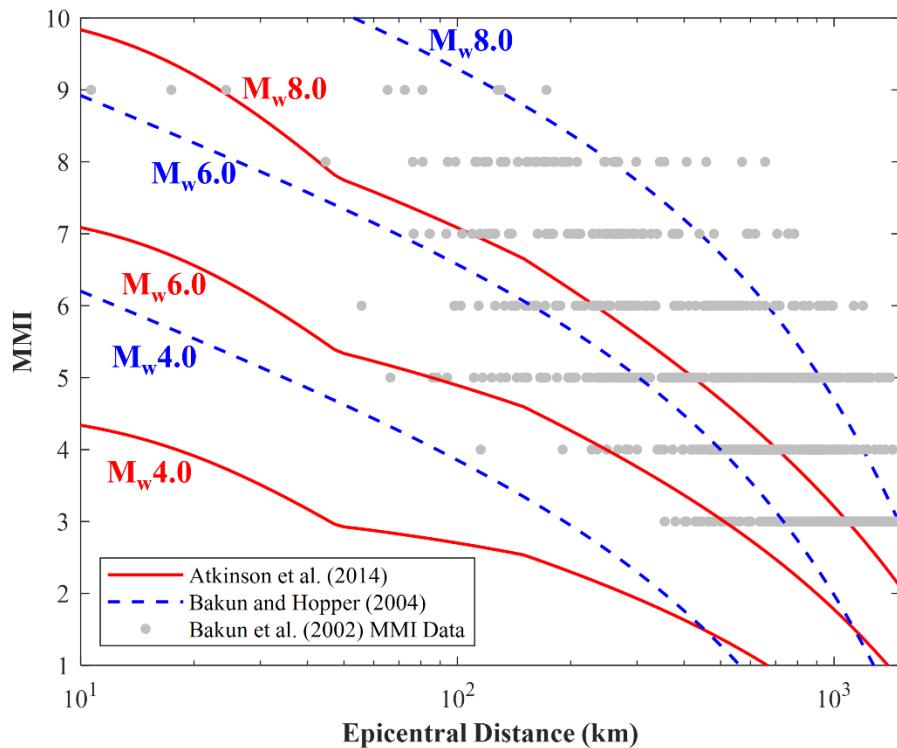
Figure 2. ENA-specific intensity prediction equations (IPEs) proposed by Bakun and Hopper (2004) and Atkinson et al. (2014) considering three values of M_w.

144　　　　While it may appear, per the latest ENA IPE, that the 1886 rupture was much larger than previously
145　　thought, there are limitations that give rise to the work that follows. Most notably, perhaps, is that existing
146　　ENA IPEs do not allow for consideration of site effects when predicting MMI. The Atkinson et al. (2014)
147　　IPE, for example, is intended for site class C, or stiff soil and soft rock profiles (i.e., sites where V_{S30}, the
148　　time averaged shear-wave velocity over the upper 30 m, is 360 m/s < V_{S30} < 760 m/s). Accordingly, if some
149　　of the 1886 MMI observations were on softer sites (e.g., in river valleys or along waterways, which seems

150   likely), then the estimate of $M_w8.0$-$M_w8.2$ obtained via the Atkinson et al. (2014) IPE could require

151   significant reduction. Moreover, the MMI studies in Table 1: (i) predate the latest knowledge of ENA

152   ground motions and site response (Goulet et al., 2018; Harmon et al., 2019); and (ii) do not rigorously

153   account for uncertainty. In this study, probabilistic site-adjusted ground motion intensity measures (IMs)

154   will be explicitly predicted using 17 ENA ground-motion models (GMMs) (whereas ground motions were

155   only implicitly predicted in prior studies). In turn, MMI values will be probabilistically predicted at study

156   sites using the latest IM-MMI models (e.g., Cramer, 2020). These predictions will be repeated for a

157   multitude of fault locations, orientations, and magnitudes to compute the likelihood that each source would

158   produce the 1886 MMI observations of Bakun et al. (2002). This approach, which will be subsequently

159   presented in detail, incorporates ground-motion IM uncertainty, IM-MMI uncertainty, and source location

160   uncertainty to produce a probability distribution of rupture magnitude. It must be noted that MMI data is

161   also subject to measurement uncertainty and reporting bias (e.g., Hough et al., 2000; Cramer and Boyd,

162   2014). While site-specific measurement uncertainties, correction factors, or weighting schemes could be

163   accommodated, a thorough reinterpretation of the more than 1000 original intensity reports would be

164   required, to include possible reassignment of MMI values in the Bakun et al. (2002) dataset and

165   development of observation-specific uncertainties. In the current effort, however, we study the existing data

166   directly, treat measurement uncertainty in a simple manner, and assign all observations equal weight. It

167   should be noted that all prior studies of the 1886 intensities have also used these data. Thus, while our

168   methodology has important advantages over prior efforts (e.g., the capacity to probabilistically constrain

169   the rupture location), a future study might further benefit from rigorous reinterpretation of the original

170   intensity reports.

171   *2.2 Analyses using liquefaction*

172   Paleoliquefaction evidence suggests that at least seven moderate-to-large earthquakes have impacted the

173   South Carolina Coastal Plain in the last 6,000 years (e.g., Gohn et al., 1984; Obermeier et al., 1985;

174   Obermeier et al., 1987; Talwani and Schaeffer, 2001). Three of these events are interpreted to have a source

175   in the vicinity of Charleston and a recurrence rate of ~500 years (Talwani and Schaeffer, 2001). The 1886

176   event was the most recent of these to generate liquefaction and the only instance in which liquefaction was

177   well documented as it occurred. Surface manifestations (e.g., ejecta, ground cracks) were mapped by Earle

178   Sloan, among others, and compiled by Dutton (1889). Additional liquefaction evidence was subsequently

179   discovered during trenching investigations (e.g., Obermeier et al., 1985; Talwani and Cox, 1985).

180   Collectively, this evidence has been analyzed to determine the magnitude of the earthquake that caused it.

181   Liquefaction models (e.g., Green et al., 2019; Maurer et al., 2015b) conventionally predict the future

182   triggering and surface manifestation of liquefaction, given in-situ geotechnical test data and some seismic

183 loading. In an inverse analysis, these models are used in reverse to constrain the seismic loading that would,

184 and would not, produce the observed response. By comparing this loading to that forward predicted by a

185 GMM for an assumed source, the magnitude of that source may be constrained. Studying primarily standard

186 penetration test (SPT) data from sites of interest, Martin and Clough (1994) carried out such an analysis

187 with the Seed et al. (1984) SPT-based liquefaction triggering model and the Ishihara (1985) liquefaction

188 manifestation model. Assuming a seismic source at the centroid of reported intensity (roughly consistent

189 with the hypothesized Woodstock Fault), adopting GMMs then available (e.g., Chapman et al., 1989), and

190 employing considerable judgement, Martin and Clough (1994) estimated that an $M_w7.0$-$M_w7.5$ event could

191 produce liquefaction consistent with that observed.

192   Studying cone penetration test (CPT) data, Hayati and Andrus (2008) used the Robertson and Wride

193 (1998) CPT-based triggering model and the Iwasaki et al. (1978) manifestation model to estimate a

194 magnitude of $M_w6.8$-$M_w7.3$. Because triggering models are trained almost exclusively with case-history

195 data from Holocene deposits, "aging correction" factors (or "deposit resistance" corrections) have been

196 proposed when applying triggering models in older soils. Specifically, it has been argued that aging effects,

197 or increases in the cyclic strength of soils over time, may be resolved into gains measurable by large strain

198 penetration tests and gains influenced by soil fabric phenomena undetectable at larger strain (e.g., Maurer

199 et al., 2014). Thus, penetration resistance may correlate to liquefaction resistance differently in Pleistocene

200 soils than in Holocene soils. Accordingly, Hayati and Andrus (2008) employed aging correction factors in

201 select geologic units. In producing their estimate of $M_w6.8$-$M_w7.3$, Hayati and Andrus (2008) assumed that

202 the Woodstock fault was the source and that a peak ground acceleration (*PGA*) of 0.3 g occurred throughout

203 Charleston, citing previous ground-motion predictions (e.g., Silva et al., 2003). Using this methodology,

204 they constrained the magnitude range for which predictions from liquefaction models matched observations

205 of response. Heidari and Andrus (2010) used a similar methodology as Hayati and Andrus (2008) but

206 applied the updated aging correction factors of Hayati and Andrus (2009) to obtain an estimate of $M_w6.8$-

207 $M_w7.0$. In addition to these studies, researchers have studied paleoliquefaction interpreted to be from older,

208 possibly similar events. Gheibi et al. (2020), for example, studied evidence induced by a "Charleston

209 Source" approximately ~550 and ~5,000 YBP. Assuming the source to be the Woodstock fault, they

210 computed respective minimum magnitudes of $M_w6.6$-$M_w7.2$ and $M_w6.2$-$M_w6.7$ for these two events, where

211 the uncertainty stems from which GMM is adopted to predict median ground motions.

212   Like prior studies of the MMI data, those of liquefaction evidence have greatly improved knowledge

213 of the regional seismic hazard, but also have limitations that motivate the present study. In brief, existing

214 studies: (i) predate both the NGA East Project (Goulet et al., 2018; Harmon et al., 2019) and the latest

215 liquefaction models trained on all globally available data (Geyin and Maurer, 2020); and (ii) do not

216 rigorously account for uncertainty. In this regard, prior studies do not account for source-model uncertainty

217   (i.e., they assume a single seismic source), do not account for the uncertainty of ground motions conditioned
218   on that source, and do consider the prediction of liquefaction in any probabilistic sense. In general,
219   published uncertainties of the 1886 rupture magnitude, whether derived from MMI or liquefaction data, are
220   arguably nominal in nature. In some studies, for example, only a single uncertainty is considered, such as
221   the epistemic uncertainty of which deterministic model is used (say, to compute ground motions). In such
222   cases, published uncertainty bounds (e.g., $M_w6.8$-$M_w7.0$) are ranges of the estimated *median* magnitude
223   considering one source of uncertainty. This may be distinctly different from the total uncertainty of the
224   1886 magnitude, which could be much greater. Consider, for example, that *instrumental* magnitudes have
225   95% CIs exceeding some of the ranges in Table 1 (e.g., Werner and Sornette, 2008).
226   In this study, and analogous to the analysis of MMI data, probabilistic site-adjusted ground motion IMs
227   will be predicted by 17 ENA GMMs. Conditioned on these IMs, the probability of liquefaction
228   manifestation will be computed by the fragility functions of Geyin and Maurer (2020), wherein multiple
229   models for soil aging effects will be ensembled. These predictions will be repeated for a multitude of
230   hypothetical sources to compute the likelihood that each would produce the observed regional liquefaction
231   response. Like the study of MMI data, this will result in a probability distribution of earthquake magnitude
232   and, arguably, a more complete and transparent understanding of the 1886 source model, at least insofar as
233   can be gained from the macroseismic evidence available for analysis.

234   **3. Macrosesimic data**
235   *3.1 MMI data*
236   Bakun et al. (2002) compiled 1,034 intensity reports from the 1886 event, including those of Dutton (1889)
237   and Bollinger and Stover (1976), and assigned MMI values per the: (i) MMI definitions of Wood and
238   Neumann (1931); and (ii) USGS National Earthquake Information Center practice for assigning intensity
239   (Stover and Coffman, 1993). In compiling these data, Bakun et al. (2002) either excluded MMI = 1 and
240   MMI = 2 reports (i.e., documented instances where shaking was not felt) or found no such reports to
241   compile. Like all previous studies of these data, we assume that MMI values were accurately assigned by
242   Bakun et al. (2002), with proper consideration of structural vulnerability to avoid biasing estimates. While
243   matters of MMI uncertainty and bias could be important and will be further discussed, they are largely
244   beyond the scope of the present analysis. It should be emphasized that our analysis is not an endorsement
245   of the Bakun et al. (2002) dataset. Rather, we study this data because it is the most recent compilation of
246   1886 intensity reports and because many prior publications studied either the Bakun et al. (2002)
247   compilation, or earlier collections of data that Bakun et al. (2002) built upon. As previously stated, a future
248   study could further benefit from a complete reinterpretation of the more than 1000 original intensity reports.
249   While the coordinates of one MMI report from Ottawa, Ontario were judged to be erroneous and corrected

250 by judgement, we otherwise adopt the Bakun et al. (2002) MMI data as presented therein. The locations of
251 these data are shown in Figure 3 and extend ~1500 km from Charleston, SC. A larger, alternative version
252 of Fig. 3 that provides an MMI raster surface is also provided in the electronic supplement.
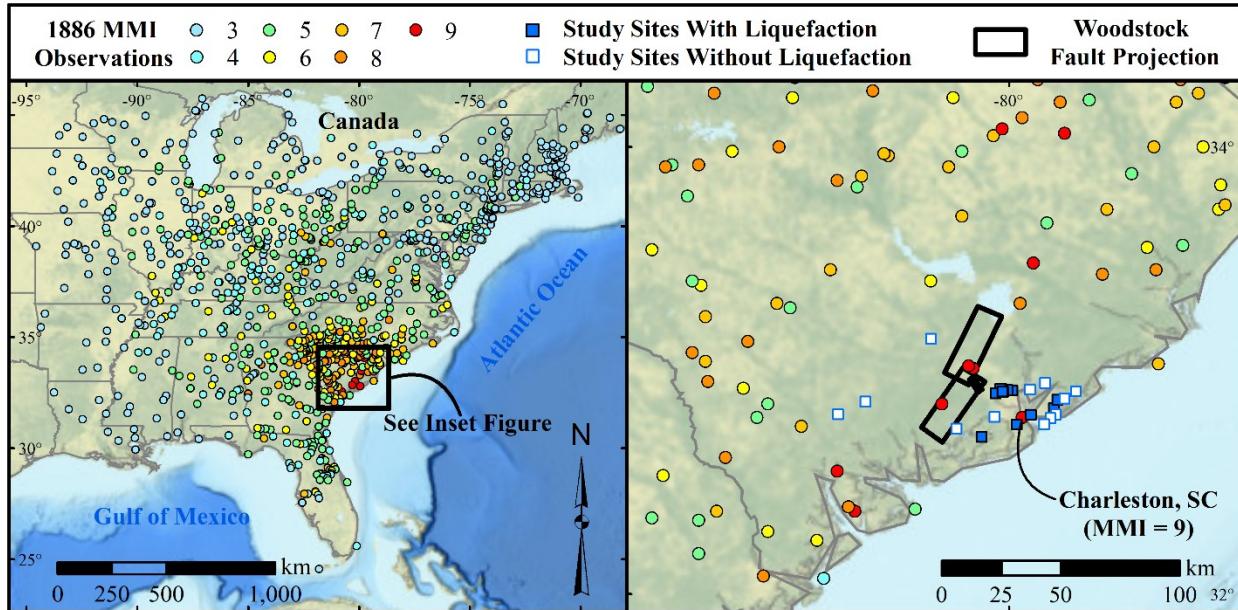
Figure 3. 1886 MMI observations (Bakun et al., 2002) and liquefaction evidence, as introduced subsequently. Also shown is the Woodstock Fault projection (Durá-Gómez and Talwani, 2009a,b).

253 *3.2 Liquefaction data*

254 Twenty-four sites where liquefaction manifestations were or were not observed in 1886, and where CPT
255 testing was subsequently performed, will be studied. Liquefaction manifested (typically in the form of large
256 sand boils) at twelve of these sites, whereas no evidence of liquefaction was observed at the remainder.
257 These sites are summarized in Table 2, where citations are provided for the observed liquefaction response
258 and for the geotechnical tests, which are all available in the public domain. These study sites are also
259 mapped later in the paper and CPT data from each site is plotted in the electronic supplement.

Table 2. Summary of 1886 liquefaction data analyzed herein.

CPT ID	Longitude	Latitude	Manifestation	Geotechnical Reference	Liquefaction Reference
BKY07	-79.9061	32.9150	No	USGS (2021)	Dutton (1889)
BKY09	-79.8385	32.9443	No	USGS (2021)	Dutton (1889)
BKY23	-79.9855	32.9115	Yes	USGS (2021)	Dutton (1889)
BKY24	-80.0071	32.9118	Yes	USGS (2021)	Dutton (1889)
CHN01	-79.7900	32.8030	No	USGS (2021)	Dutton (1889)
CHN07	-79.8134	32.7874	No	USGS (2021)	Dutton (1889)
CHN12	-79.7989	32.8310	Yes	USGS (2021)	Amick et al. (1990)
CHN15	-79.6998	32.9073	No	USGS (2021)	Dutton (1889)

CHN28	-79.8428	32.7597	No	USGS (2021)	Dutton (1889)
CHN29	-79.7840	32.8682	Yes	USGS (2021)	Amick et al. (1990)
CHN31	-79.7520	32.8741	No	USGS (2021)	Dutton (1889)
CHN32	-80.0267	32.9043	Yes	USGS (2021)	Dutton (1889)
CHN33	-80.0323	32.9166	Yes	USGS (2021)	Dutton (1889)
CHN34	-80.0391	32.9170	Yes	USGS (2021)	Dutton (1889)
CHN50	-80.1235	32.7023	Yes	USGS (2021)	Dutton (1889)
CHN59	-79.9655	32.7575	Yes	USGS (2021)	Dutton (1889)
CHN64	-80.0604	32.8982	Yes	USGS (2021)	Dutton (1889)
CREC1	-80.0655	32.7921	No	Boller (2008)	Martin and Clough (1994); Boller (2008)
FHS3	-80.3507	33.1420	No	Hasek (2016)	Williamson and Gassman (2014); Hasek (2016)
HA74	-80.0300	32.9050	Yes	Heidari & Andrus (2012)	Dutton (1889)
HWD2	-80.2355	32.7394	No	Hasek (2016)	Talwani and Cox (1985); Hasek (2016)
S99634DS1	-79.9015	32.8017	Yes	Heidari & Andrus (2010)	Dutton (1889)
WLC2	-80.6444	32.8607	No	Geiger (2010)	Geiger (2010)
WRAP2	-80.7666	32.8040	No	Geiger (2010)	Geiger (2010)

260   In compiling these study sites, an interpreted, potential lateral spread at Fort Dorchester, SC (Talwani
261   et al., 2011) was omitted because the liquefaction response was ambiguous and because lateral spreading –
262   a distinct and complex manifestation of liquefaction – is not intended to be predicted by the liquefaction
263   model adopted herein. Whereas the MMI data are spread throughout ENA, the compiled liquefaction data
264   are within 100 km of Charleston. Because liquefaction is mechanistically implausible, irrespective of soil
265   properties, at *PGA*s less than ~0.09 g (de Magistris, 2013), the maximum site-to-source distance of
266   liquefaction observations is inherently limited. By corollary, an analysis of sites where liquefaction was not
267   observed, but which are very far from Charleston, would not provide meaningful constraint of the source
268   model, given that the computed probability of such an observation is 100% even for very large earthquakes.
269   In addition, while the MMI data from 1886 are unlikely to grow significantly, additional liquefaction data
270   could be compiled. That is, the liquefaction response was documented in 1886 at more than twenty-four
271   sites, but costly geotechnical testing must also be performed at each site. In this regard, it is known that
272   CPTs have been performed near additional sites of observation, yet these data are privately held and could
273   not be obtained for analysis. Nonetheless, a larger dataset could be studied in the future.

274   **4. Methodology**

275   The methodology that will be used to analyze macroseismic data was first introduced by Rasanen and
276   Maurer (2021, 2022), who collectively demonstrated and validated its use on landslide and liquefaction

277 evidence produced by eleven modern earthquakes with known source models. This is the first application
278 of the method to a prehistoric or pre-instrumental earthquake. The methodology will be covered in two
279 sections. The first provides a succinct conceptual overview. The second describes in detail the application
280 of the methodology to the 1886 Charleston earthquake.

281 *4.1 Conceptual Overview*

282 Our goal is to probabilistically constrain the 1886 source model. This is accomplished by computing the
283 likelihood that a rupture with some magnitude, location, and geometry, would produce a set of field
284 observations (reported MMI or liquefaction observations), wherein uncertainties inherent to these observed
285 outcomes are considered. In general, the likelihood of a parameter having some value, given a set of
286 observations, is the product of the probabilities of those observations, conditioned on the parameter value.
287 In other words, the likelihood of a rupture having some magnitude (M_w), location (L), and geometry (G),
288 given a set (x) of field observations at N different sites, can be computed as:

289
$$Likelihood\ (L, G, M_w | x) = P(X = x | L, G, M_w) = \prod_{i=1}^{N} P(X_i = x_i | L, G, M_w) \tag{1}$$

290 where $P(X_i = x_i | L, G, M_w)$ is the probability of what was observed at site i (i.e., the observed MMI or
291 liquefaction response) given an earthquake with parameters L, G, and M_w. By repeating for enumerable
292 possibilities, the actual rupture parameters may be probabilistically constrained by the likelihood function
293 (product of the probabilities of N observations), such that different source models will be found more and
294 less likely to produce the observed evidence.

295   If the evidence is MMI data, then the probability of any one field observation (obs) is:

296
$$P(\text{obs} | \text{EQK}: L, G, M_w) = \int_{IM} \int_{MMI} P(\text{obs} | MMI)\ f(MMI | IM)\ f(IM | L, G, M_w) \cdot dMMI \cdot dIM \tag{2}$$

297 where $f(IM | L, G, M_w)$ is the probability of an IM value conditioned on fault parameters L, G and M_w, and
298 site parameter V_{S30}, as computed by a GMM that considers site response; $f(MMI | IM)$ is the probability of
299 an MMI value conditioned on the IM value, and possibly on other parameters, as computed by an IM-MMI
300 model; and $P(\text{obs} | MMI)$ is the binomial probability that the predicted MMI is equal to the observed MMI.
301 In this work, we consider predictions and observations to agree if within ± 0.5 MMI (e.g., predicted MMIs
302 of 5.51 and 6.49 agree with an observed MMI of 6). Thus, we assign an implicit, uniform measurement
303 uncertainty of ± 0.5 MMI, but do not otherwise model the uncertainty of reported MMIs. In this regard, a
304 thorough reinterpretation of the original intensity data, to include assignment of site-specific uncertainty
305 distributions and corrections for bias, would be a valuable endeavor. In the current effort, however, we
306 adopt the Bakun et al. (2002) MMI dataset, which has been studied by other modern investigators of the
307 1886 earthquake.

308   If the evidence is observed liquefaction response rather than MMI, then the probability of field

309   observation is computed per Eq (3) if liquefaction manifestation was observed and per Eq (4) otherwise:

310    $$P(\text{Manifestation}|EQK{:}\,L, G, M_w) = \int_{IM} P(\text{Manifestation}|IM, M_w)\, f(IM|L, G, M_w) \cdot dIM \qquad (3)$$

311    $$P(\text{No Manifestation}|EQK{:}\,L, G, M_w) = 1 - \int_{IM} P(\text{Manifestation}|IM, M_w)\, f(IM|L, G, M_w) \cdot dIM \qquad (4)$$

312   In these equations, uncertainty is not assigned to the observed liquefaction response; $f(IM|L, G, M_w)$ has

313   the same meaning as above; and $P(\text{Manifestation}|IM, M_w)$ is computed by a probabilistic model that

314   predicts the triggering of liquefaction at depth and its subsequent manifestation at the ground surface using

315   subsurface geotechnical data. Ultimately, the uncertainties that are, and are not, accounted for in this work

316   will be explicitly discussed.

317   Application of this method is demonstrated conceptually in Figure 4 considering four MMI

318   observations and two hypothetical sources for the earthquake that produced the observations. In actual

319   analyses, a very large number of sources is considered. Figure 5 illustrates how the relative likelihoods of

320   these two sources are assessed. Shown in Figure 5a are the computed probabilities of individual

321   observations, given a rupture of source one, as computed by Eq. 2 for varying M_w. In Figure 5b, this is

322   repeated considering a rupture of source two. In Figure 5c, the likelihood of each source is computed as a

323   function of M_w by Eq. 1 (i.e., the product of the four probability distributions in Figure 5a or 5b). In this

324   simple example, source one has a far greater peak likelihood of producing the set of field observations,

325   whereas source two is very unlikely to do so, regardless of its M_w. By repeating this process for an array of

326   hypothetical sources and generating a likelihood distribution for each, the characteristics of the causative

327   rupture (e.g., location, orientation, magnitude) are probabilistically constrained, to the degree that evidence

328   permits. In this work, we compute M_w probability-distributions conditioned on both the hypothesized

329   Woodstock fault and on an unknown source. The latter is accomplished by aggregating probability

330   distributions from all hypothetical sources and thus includes the uncertainties of fault location and

331   geometry, whereas the former assumes the fault location and geometry are known.
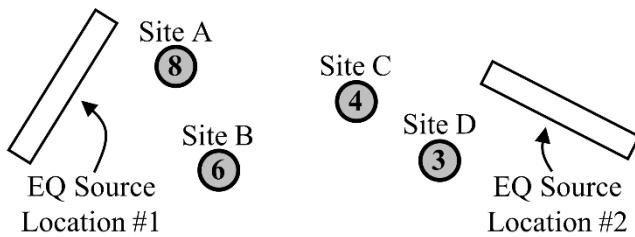
Figure 4. Hypothetical MMI inverse analysis consisting of four field sites, where MMI values were reported for each of the site locations (the value in the circle). In addition, two hypothetical sources for the earthquake that produced these observations are shown.

13
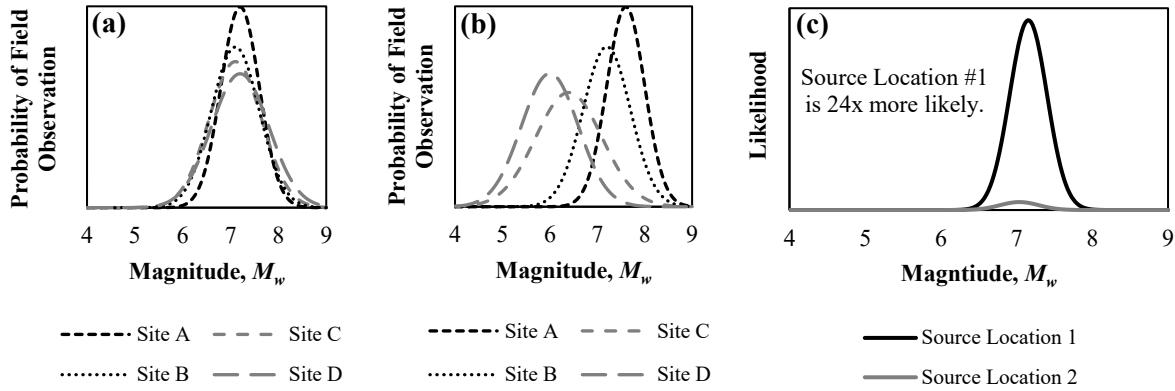
Figure 5. Approach for computing the likelihood of the sources depicted in Figure 4: (a) probabilities of individual observations, given an earthquake at location one of variable M_w; (b) probabilities of individual observations, given an earthquake at location two of variable M_w; (c) M_w-likelihood distributions for source locations one and two.

332 *4.2 Implementation Details*

333 The implementation of Eqs. 1-4 is next described in detail. This includes both general methods transferrable

334 to other regions and to other forms of evidence, as well as the specific models adopted for the 1886

335 earthquake. Because the analyses of MMI and liquefaction data are procedurally similar, we first fully

336 describe the former and then succinctly discuss differences specific to the latter. For completeness,

337 however, the liquefaction analysis procedure is fully replicated in the electronic supplement.

338   As detailed in the following, two approaches will be used to model hypothetical seismic sources. These

339 sources are first treated as earthquake epicenters on a grid pattern, which we term the "epicenter search".

340 To use modern GMMs, which are applicable only to faults, and because modeling large earthquakes as

341 points is inappropriate, each epicenter is converted to an amorphous fault realization using site-to-source

342 distance correlations, which implicitly assume some fault geometry as a function of magnitude. With this

343 approach, the source location may be investigated (i.e., probabilistically constrained) without consideration

344 of rupture geometry. While this reduces computational expense, it considers only a "median" geometry and

345 thus omits one source of uncertainty. Accordingly, and following constraint of the epicentral region, seismic

346 sources are next treated as faults having an array of locations, lengths, and orientations, which we term the

347 "fault search." Of these enumerable faults, one is the hypothesized Woodstock Fault, which we highlight

348 in the results for obvious reasons.

349 *4.2.1 MMI Analysis Procedure*

350 The analysis of MMI data is completed via the following 15 steps:

14

351   (1) Create an array of seismic sources. In this paper, a 62,500 km^2 grid of surficial points was centered on

352   Charleston. Within this grid, a finer point spacing increases spatial resolution while a coarser spacing

353   decreases computational demand (which can be significant, and tractable only with high performance

354   computing). Having found that 1 km and 10 km spacing produce nearly identical outcomes, we adopted

355   the latter spacing when studying MMI observations. These surficial points will be treated first as

356   earthquake epicenters (i.e., the "epicenter search") and second as surface projections of the centroids of

357   faults (i.e., the "fault search").

358   (2) Select N study sites (i.e., the *MMI* observations compiled by Bakun et al. (2002)).

359   (3) Select an appropriate GMM. In this paper, the 17 GMMs developed by the NGA East Project (Goulet et

360   al., 2018) were coalesced in a logic tree using the weights proposed by Goulet et al. (2018). Using this

361   scheme, the least and greatest model weights were ~2% and 10%, respectively.

362   (4) For each seismic source created in (1):

363   (5) For each seismic-source M_w considered (a range of M$_w$4 to M$_w$8.2 – the applicable range of the adopted

364   GMMs – was used in this paper):

365   (6) For each of N study sites selected in (2), cycling from $i = 1$ to N:

366   (7) Compute the site-to-source distance(s) required by the GMM chosen in (3), as measured from study site

367   i to the seismic source selected in (4). For the GMMs adopted herein, the only such metric required is-the

368   closest distance to fault rupture (R_{RUP}). When the seismic sources in (1) are treated as epicenters, rather

369   than faults, the correlations of Scherbaum et al. (2004) were used to estimate a median R_{RUP} from

370   epicentral distance (R_{EPI}). In effect, these correlations, which are magnitude dependent, convert each

371   epicenter from (1) into a median realization of a multidimensional fault. Alternative approaches to

372   estimating R_{RUP} from point sources are provided by Bommer et al. (2016) and Thompson and Worden

373   (2018). Ultimately, the sources in (1) are explicitly modeled as faults via the "fault search" to determine

374   whether the field evidence can constrain the 1886 source more precisely. In doing so, R_{RUP} is directly

375   measured from fault planes and the uncertainty of fault orientation is considered.

376   (8) Using the GMM in (3), M_w from (5), and site-to-source distances from (7), compute the probability

377   density function (PDF) of expected *PGA* at site $i,$ modified for site effects. In general, this PDF is a

378   lognormal random variable described by a median and lognormal standard deviation, which are given by

379   a GMM. *PGA* predictions beyond ± 3 standard deviations of the median were truncated, as is typical, and

380   the PDF was scaled such that the area beneath it was one. In this study, the ENA weighted GMM predicts

381   *PGA* for reference rock conditions. Accordingly, the V_{S30}-dependent model of Harmon et al. (2019),

382   which is ENA-specific and developed as part of the NGA-East Project, was used to adjust *PGA*s for local

383   site effects. V_{S30} was estimated at the site of each field observation using the Heath et al. (2020) maps.

384   (9) For each possible *PGA* value at study site i, as computed in (8) for a given M_w and R_{RUP} pair:

385 (10) Select an IM-MMI model. In this paper, the Atkinson and Kaka (2007), Worden et al. (2012), Caprio et

386 al. (2015), and Cramer (2020) models were adopted. The inputs to these models vary but generally include

387 PGA, M_w, and R_{RUP}. The median MMI predicted by each model is plotted in Figure 6 as a function of

388 PGA considering an M_w6 event at $R_{RUP} = 100$ km.
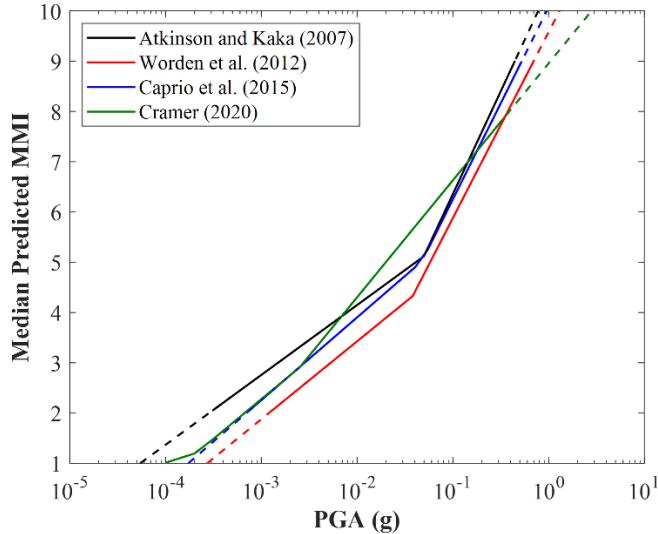
Figure 6. Predicted median MMI versus PGA according to four recent models (where applicable, $M_w = 6.0$; $R_{RUP} = 100$ km). Solid lines indicate the range of each model's proposed applicability (typically the range of training data); dashed lines indicate extrapolation beyond these respective bounds.

389 (11) Using the IM-MMI model selected in (10), the R_{RUP} from (7), and the PGA from (9), compute the PDF

390 of expected MMI at study site i. In this paper, MMI predictions beyond $1 \leq MMI \leq 10$ were truncated

391 because the definition of MMI gives a lower bound of 1 and because intensities greater than 10 are rarely

392 assigned in practice (Stover and Coffman, 1993). Following truncation (when applicable) the PDF was

393 scaled such that the area underneath it was one.

394 (12) For each possible MMI value at study site i, as computed in (11):

395 (13) Compute the probability of field observation as described in Eq. 2. Completing this equation (i.e., by

396 multiplying the probability of field observation by the probabilities of MMI and PGA, then summing over

397 all MMI and PGA values) gives the probability of the field observation at site i for a given seismic source

398 location, geometry, and M_w. Repeating steps 6-13 for all M_w results in a probability of field observation

399 curve for each source, examples of which are in Figures 5a and 5b.

400 (14) Compute the likelihood of a seismic source (as evidenced by MMI) as a function of M_w by multiplying

401 the probabilities of all field observations (i.e., multiply the curves in Figures 5a or 5b at each value of

402 M_w). The result, an example of which is shown in Figure 5c, is a likelihood distribution of M_w conditioned

16

403   on a single seismic source (i.e., assuming that a fault with given location and geometry, but unknown M_w,

404   is responsible for producing the field observations).

405 (15) Repeating steps 5-14 for all seismic sources created in (1) results in a likelihood distribution of M_w for

406   each. Collectively, this field of distributions describes the locations and magnitudes of earthquakes that

407   are, and are not, likely to produce the field evidence. To allow for relative likelihoods to be compared

408   visually, we normalize the likelihood of each source by the peak likelihood among all sources, such that

409   the most likely source has a normalized peak value of one. We then map contours of likelihood to identify

410   this location. The M_w distribution at this location (see Figure 5c) is the PDF of the inverted M_w conditioned

411   on the most likely source. Finally, by aggregating PDFs from all potential sources in (1), an overall PDF

412   of M_w, considering all possible sources, is produced. While a single source will always be "most likely,"

413   earthquakes at other locations typically also have potential to produce the evidence. This latter PDF,

414   conditioned on all possible sources, includes that uncertainty.

415   *4.2.2 Liquefaction Analysis Procedure*

416   The analysis of liquefaction is akin to that of MMI, differing only in steps (10) through (13) of the

417   preceding. Once site-adjusted PGAs are predicted at sites of observation (in this case, sites with liquefaction

418   observations), the factor of safety against liquefaction is computed throughout the soil profile using the

419   Boulanger and Idriss (2014) CPT-based triggering model, which is a function of PGA, M_w, and subsurface

420   geotechnical data. As part of this procedure, the effects of soil aging are separately accounted for with three

421   proposed models, in addition to a control analysis without any such accounting. Two of these models,

422   henceforth termed "A" and "B", are regional-scale corrections based on the Weems et al. (2014) geology

423   map of the Charleston region, whereas Model "C" uses site-specific measurements. Model A uses the

424   Hayati and Andrus (2009) aging-correction model based on the measured to estimated shear-wave velocity

425   ratio (MEVR). As MEVR increases, the shear stiffness of soil measured at small strain exceeds that inferred

426   from large strain measurements, which may be interpreted as an indicator of cementation. Median MEVR

427   values for each geologic unit were adopted from the regional sampling of Heidari and Andrus (2012). Model

428   B uses the Hayati and Andrus (2009) correction model based on depositional age, which we estimated for

429   each geologic unit from the Weems et al. (2014) map. As a deposit's age increases, an increasingly larger

430   correction is applied to the computed liquefaction resistance. Lastly, Model "C" again adopts the MEVR-

431   based approach of Hayati and Andrus (2009), but in this case MEVR is directly computed at each study

432   site using data from seismic CPTs (i.e., CPT resistances and shear-wave travel times) per the method of

433   Andrus et al. (2009). This approach produced corrections that tended to exceed those of Models A and B.

434   While corrections from Model C – being based on site-specific subsurface data – should be most efficient,

435   all such corrections are likely very uncertain and the best approach to account for soil aging is actively

436 debated. We thus argue that all three models, which scale upward the computed liquefaction triggering
437 resistance at depth, warrant consideration. However, given that the field observations are of surface
438 manifestations (i.e., liquefaction vents and dikes) rather than observations of liquefaction at discrete depths,
439 surface manifestations must be predicted for proper comparison to field observations. Accordingly, the
440 liquefaction potential index *(LPI)* of Iwasaki et al. *(1978)* was adopted, given its longstanding use:

441
$$LPI = \int_0^{20\,m} F\left(FS_{liq}\right) \times w(z)dz \qquad (5)$$

442 where $F\left(FS_{liq}\right)$ and $w(z)$ weight the respective influences of FS_{liq} and depth, z, on surface manifestation.
443 In brief, *LPI* assumes that manifestation depends on the FS_{liq} in each soil stratum, the thickness of all
444 liquefied strata, and the proximity of those strata to the surface. Given this definition, *LPI* can range from
445 0 to 100. Next, the probability of liquefaction manifestation was computed at each study site, considering
446 all values of M_w from (5) and all values of *PGA* from (9), using the fragility function of Geyin and Maurer
447 (2020), which relates the probability of liquefaction manifestation to *LPI*, and which was trained on all
448 globally available CPT-based liquefaction case histories. Lastly, and analogous to (13), the probability of
449 what was observed in the field was computed by Eq. 3 or 4, depending on whether manifestations were or
450 were not observed. Thus, the fact that the liquefaction hazard differs at different study sites is directly
451 measured and accounted for via site-specific CPT measurements and the latest CPT-based liquefaction
452 triggering and manifestation models. Using these data and tools, the probability of surficial liquefaction
453 manifestation is computed for a range of M_w and *PGA* at each study site. The liquefaction analysis
454 procedure is fully replicated in the electronic supplement.

455 **5. Results and discussion**

456 Using the above data and methodologies, the source of the 1886 Charleston rupture was rigorously
457 investigated. Results are separately presented for MMI and liquefaction evidence. For each, three types of
458 analyses were performed, namely the: (i) "epicenter search;" (ii) "fault search;" and (iii) "Woodstock
459 Fault," wherein the hypothesized Woodstock Fault is directly considered. The first provides preliminary
460 probability distributions of the source location and magnitude. The second provides final distributions that
461 include additional source-model uncertainties and conveys whether the orientation of the causative fault
462 can be constrained. The third provides results conditioned on the singular Woodstock Fault and should be
463 adopted if all other seismic sources, known and unknown, are dismissed as sources of the 1886 earthquake.

464 *5.1 Analysis of MMI Data: Epicenter Search*

465 Using the "epicenter search," the inversion methodology was applied with the Atkinson and Kaka (2007),
466 Worden et al. (2012), Caprio et al. (2015), and Cramer (2020) IM-MMI models. For brevity, these models

467   are henceforth titled AK07, WEA12, CEA15, and C20. For these analyses, MMI observations were studied

468   if within 600 km of Charleston. While the sensitivity of this decision will be analyzed further, it was made

469   based on a trait of the Bakun et al. (2002) MMI dataset that we view as problematic, but which prior

470   researchers have not discussed. Namely, Bakun et al. (2002) did not include "negative" observations (i.e.,

471   MMI values of 1 or 2) where ground motions were not perceived. This is notable, given that a *distribution*

472   of MMI values is naturally experienced at any given distance. Problematically, the omission of low MMI

473   values asymmetrically truncates this distribution at large distances (i.e., those where some observers report

474   MMI 1 or 2). The analysis of this dataset at large distances thus potentially has the effect of biasing inverted

475   magnitudes toward larger values, since those magnitudes need not adhere to small MMI values that were

476   experienced, but which were systematically undocumented. The analysis of data only at small distances is

477   equally undesirable, of course, given the inherent benefits of analyzing more data over a wider spectrum of

478   attenuation. Based on these competing interests and given that we begin to observe evidence of MMI

479   truncation at distances exceeding 600 km, this threshold was provisionally selected.

480   The resulting geospatial contours of seismic-source likelihood are shown in Figure 7 for the WEA12

481   model. Also shown is the surface projection of the hypothesized Woodstock fault (Durá-Gómez and

482   Talwani, 2009a,b), as well as instrumental epicenters ($M_w > 1.0$; 1900-present). These contours show the

483   source locations most and least likely to produce the MMI evidence. Using the WEA12 model, a zone of

484   high likelihood is computed just northwest of the Woodstock Fault projection, with the most likely source

485   located 2 km from the mapped projection. By contrast, source locations to the south and east are

486   comparatively very unlikely to produce the MMI observations. For brevity, results using the AK07, CEA15,

487   and C20 models are summarized in Figure 7 via stars, which denote the epicenters deemed most likely by

488   each model. Complete contour maps, analogous to those in Figure 7 for WEA12, are provided in the

489   electronic supplement. While CEA15 produces a result similar to WEA12, with the most likely source

490   inside the Woodstock Fault projection, AK07 and C20 suggest most likely sources ~10 km and ~35 km

491   east of the projection. Thus, all analyses point to a source north of Charleston. Some IM-MMI models

492   strongly support an earthquake source in the immediate vicinity of the hypothesized Woodstock Fault, while

493   others do not.
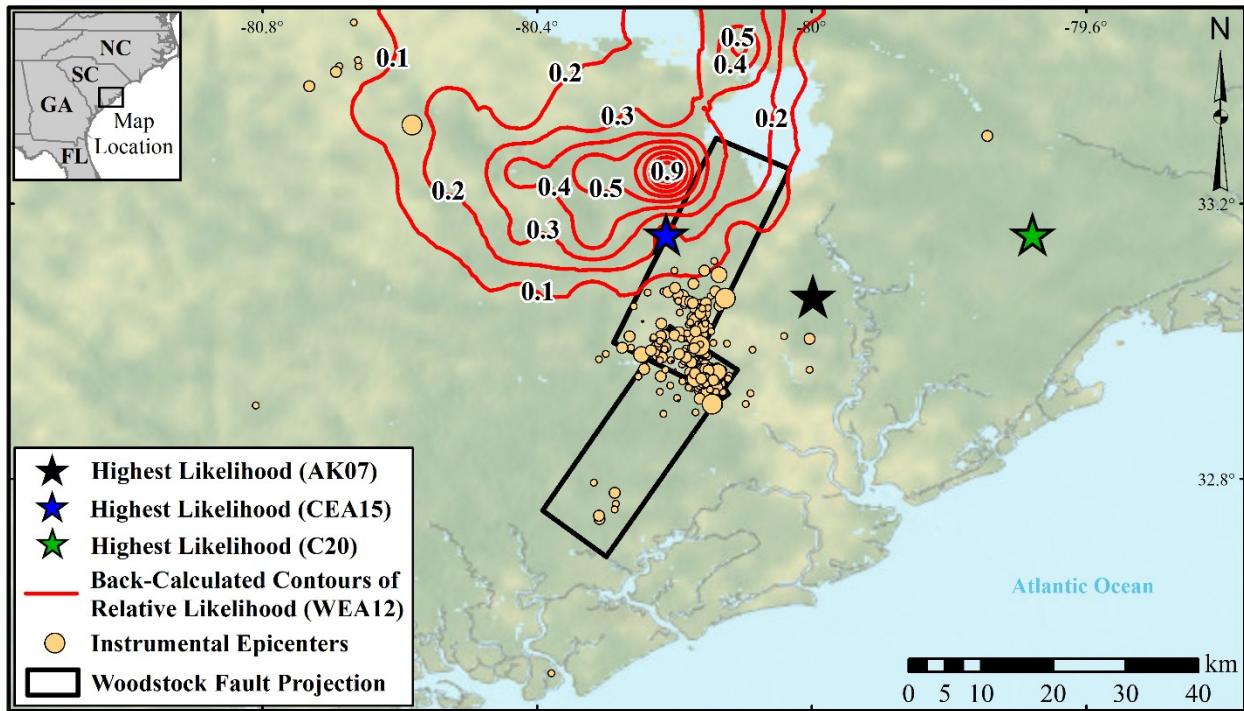
Figure 7. Likelihood contours produced using the WEA12 model which indicate source locations that are, and are not, likely of producing the MMI evidence. Stars = the most likely source locations per the AK07, CEA15, and C20 models; AK07 = Atkinson and Kaka (2007), WEA12 = Worden et al. (2012), CEA15 = Caprio et al. (2015), C20 = Cramer (2020); black rectangles = the Woodstock fault projection (Durá-Gómez and Talwani 2009a,b); tan circles = instrumental epicenters ($M_w > 1.0$; 1900-present) (USGS, 2022).

494   Shown in Figure 8 are the cumulative distribution functions (CDFs) of M_w inverted from MMI data, as

495   computed with each of the four IM-MMI models. These results include the uncertainty of an unknown

496   source (epicenter) location, the uncertainties of expected shaking intensities conditioned on a given source,

497   considering also the epistemic uncertainty of GMM selection, and the uncertainties of the IM-MMI models.

498   For comparison, the CDF assigned to the Charleston source in the NSHMP (Petersen et al., 2014, 2020) is

499   also shown and ranges from $M_w6.7$ to $M_w7.5$ with a median of $M_w7.1$. AK07 and C20 produce nearly

500   identical CDFs with a median of $M_w7.1$. WEA12 and CEA15 produce similarly identical CDFs, but with a

501   median of $M_w7.3$. The grouping of these outcomes can be traced to Figure 6. AK07 and C20 tend to predict

502   larger MMIs for a given PGA. As a result, smaller magnitude ruptures result from the use of AK07 and

503   C20 within the inversion methodology. Each of the individual CDFs has relatively low uncertainty. This

504   can be attributed to: (i) the large quantity of MMI observations; (ii) the fact that not all uncertainties are

505   considered (e.g., those of site conditions or rupture geometry, among others); and (iii) the relatively large

506   variance of MMI observations over any given site-to-source distance. Regarding the last, extreme outliers

507   from the mean MMI trend have the effect of yielding very low likelihoods for small and large magnitude

508   events. That is, only a narrow range of magnitudes are likely to simultaneously produce MMI observations

509   of, say, 3, 5, and 7 at the same site-to-source distance and on similar site profiles. Conversely, a wider range

510   of magnitudes could simultaneously produce MMI observations of 4, 5, and 6. While geostatistical analyses

511   could conceivably identify and delete observations that are extreme outliers, the justification would be

512   largely speculative without an intensive reinvestigation. In other words, we are unaware of any objective

513   basis for deleting some MMI reports but not others in the absence of a complete reinterpretation of the more

514   than 1000 original intensity reports. Lastly, to include the epistemic uncertainty of IM-MMI model

515   selection, each CDF was weighted to produce an ensemble CDF, as shown in Figure 8. This selection of

516   weights is heuristic, as no quantitative justification could be identified (e.g., based on model residuals in

517   ENA). The weights and justifications are as follows. C20 (0.4) is the latest ENA-specific model and was

518   trained using the largest quantity of ENA data. AK07 (0.3) is also ENA-specific, but also found no need

519   for region-specific models within North America. For that reason, and because differences in methodology

520   (e.g., data selection, intensity scales, regression techniques) can result in very different models with

521   apparently similar performance (i.e., on the respective training sets), we hesitate to omit other reputable

522   models even if nonspecific to ENA. Accordingly, WEA12 (0.15), which was trained only on California

523   data, and CEA15 (0.15), which was trained on global data including some from ENA, are also weighted in

524   the ensemble. While CEA15 might otherwise warrant greater weighting, its training set was truncated at a

525   site-to-source distance of 200 km, which calls into question its suitability in ENA, where ground motions

526   are commonly felt at greater distances (as in the 1886 event).
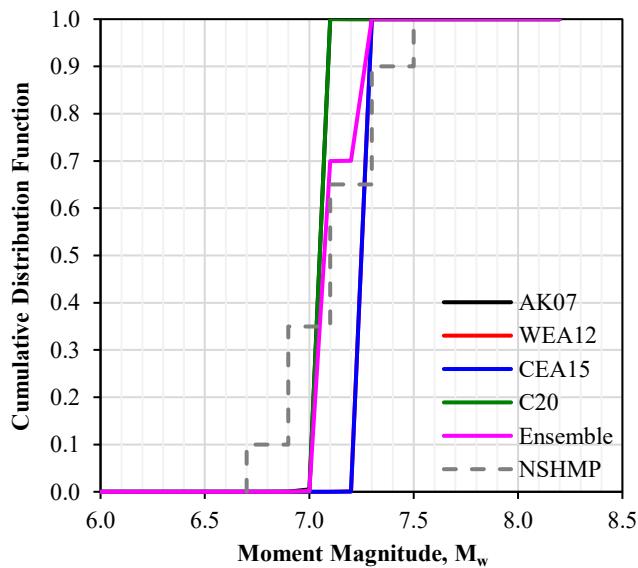
Figure 8. CDFs of rupture magnitude, as inverted from MMI observations using four IM-MMI models, wherein seismic sources are modelled as epicenters. AK07 = Atkinson and Kaka (2007), WEA12 = Worden et al. (2012), CEA15 = Caprio et al. (2015), C20 = Cramer (2020), and NSHMP = National Seismic Hazard Model Project (Petersen et al., 2014, 2020).

527 *5.2 Analysis of MMI Data: Fault Search (Seismic Sources Modelled as Faults)*

528 To determine whether MMI observations can constrain the seismic source beyond a general area, the

529 gridded sources were next treated as surface projections of the centroids of faults. For simplicity, the faults

530 were initially assumed to be strike-slip with a dip of 90 degrees. The orientations of the faults were

531 discretized into strike azimuths of 30° increments (i.e., 0°, 30°, 60°…). The lengths of the faults were

532 computed via the magnitude-dependent correlation of Wells and Coppersmith (1994). The depth to the top

533 of rupture was estimated using the correlation of Kaklamanos et al. (2011), which produces estimates

534 consistent with the inferred depths of modern ruptures in the region (Chapman et al., 2016). Thus, while

535 not every aspect of rupture geometry was assigned uncertainty, the analyses can nonetheless determine

536 whether some fault alignments (e.g., the hypothesized Woodstock Fault) are more likely than others to

537 produce the observed evidence. With this approach, the inversion methodology was again applied to the

538 Bakun et al. (2002) dataset using each of the four IM-MMI models.

539 Results are shown in Figure 9 for the WEA12 model in the style of a heat map. Faults more likely to

540 produce the MMI observations have thicker lineweight and are darker in color. Faults very unlikely to

541 produce the evidence have thin lineweight and light color, and thus blend with the map's background. Also

542 shown is the surface projection of the hypothesized Woodstock Fault (Durá-Gómez and Talwani 2009a,b)

543 as well as instrumental epicenters. While most modelled faults are relatively unlikely to produce the

544 evidence, the faults deemed most likely align with the hypothesized Woodstock Fault, albeit their

545 magnitudes are greater than most prior estimates. The singular fault most likely to produce the MMI

546 evidence, for example, has a median M_w of 7.60. Faults striking E-W and N-S have lesser, but nontrivial,

547 likelihoods of producing the MMI evidence.

548 For brevity, results using the AK07, CEA15, and C20 models are each summarized in Figure 9 by

549 single lines, which denote the singular faults deemed most likely by each model. Complete heat maps,

550 analogous to those in Figure 9 for WEA12, are provided in the electronic supplement. These three models

551 produce results that are similar to one another and different from WEA12, with the most likely faults

552 oriented perpendicular to the Woodstock Fault and located partly offshore. These most-likely faults range

553 in median magnitude from M_w7.20 to M_w7.90. A close inspection of these results (electronic supplement)

554 reveals that each of these models finds faults consistent with the Woodstock Fault to have low relative

555 likelihoods of producing the MMI evidence, as compared to all other hypothetical faults. It is worth noting

556 that the two ENA-specific IM-MMI models (AK07 and C20) point to faults inconsistent with the

557 Woodstock Fault. While these results do not necessarily indicate that the Woodstock Fault was not

558 responsible for the 1886 earthquake, they do indicate that the causative fault cannot currently be constrained

559 by the available MMI evidence. That is, faults striking SE and partly offshore are just as likely to produce

560 this evidence as faults striking NE and onshore. While different assumptions about fault depth, length, and

561 dip invariably change the results, a parametric analysis indicates that these changes are minor, upholding

562 the overall conclusion that the MMI data cannot constrain the 1886 fault rupture. If anything, the data point

563 to a fault striking SE and partly offshore. This conclusion might be different if: (i) more MMI observations

564 were available in the near field; and (ii) the affected area was not on a coastline. Analogous to the inversion

565 of an epicenter from wave-arrival times, which relies on distributed instruments, this inversion relies on

566 distributed observers to "record" the amplitudes of ground motions. Thus, in events such as this, where the

567 near-field is not more densely populated with observations than the far field, and where observations are

568 geographically asymmetric, it may be more difficult to detect directions and rates of ground-motion

569 attenuation, and thus more difficult to constrain the causative fault from which motions propagated.
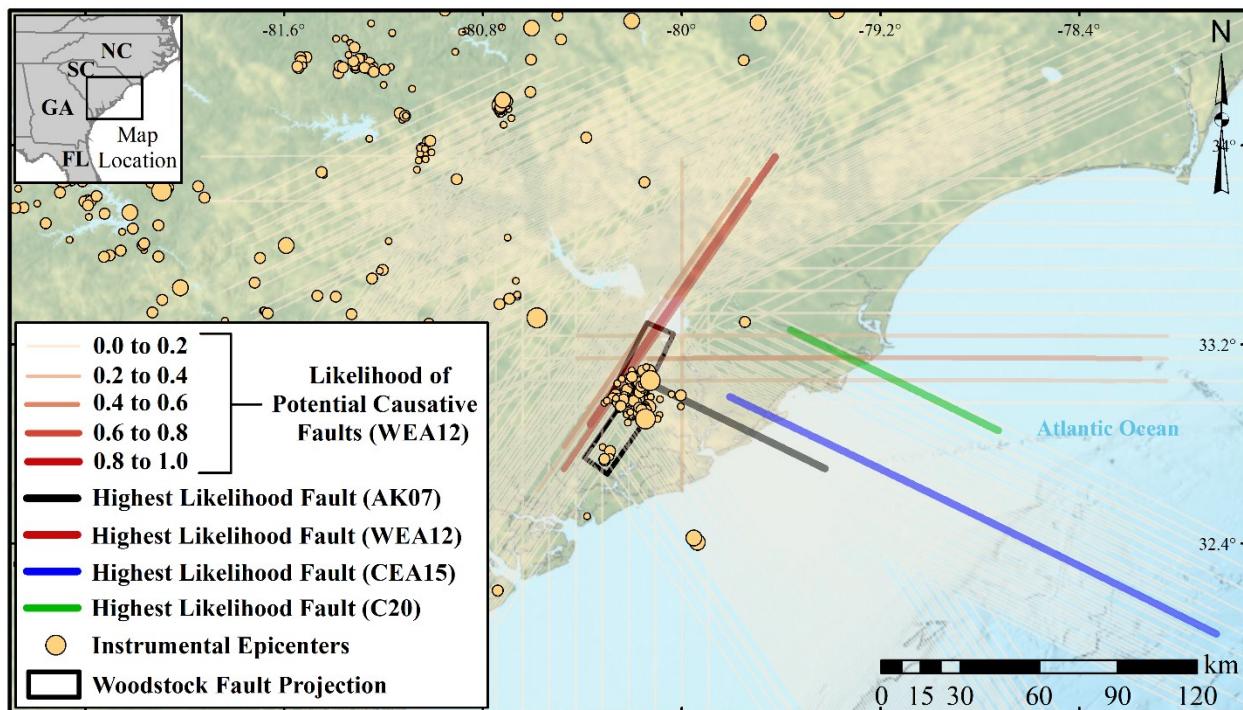
Figure 9. Likelihoods of hypothetical faults with differing orientations and lengths to produce the MMI evidence, as obtained using the WEA12 model. Also shown are the singular faults deemed most likely using the AK07, CEA15, and C20 models; AK07 = Atkinson and Kaka (2007), WEA12 = Worden et al. (2012), CEA15 = Caprio et al. (2015), C20 = Cramer (2020); black rectangles = the Woodstock projection (Durá-Gómez and Talwani 2009a,b); tan circles = instrumental epicenters ($M_w > 1.0$; 1900-present) (USGS, 2022).

570 Shown in Figure 10 are the CDFs of M_w inverted from MMI data via the "fault search", as computed

571 with each of the four IM-MMI models. These results include the uncertainty of an unknown source location,

572 the uncertainty of unknown fault orientation, the uncertainties of expected shaking intensities conditioned

573 on a given source and considering the epistemic uncertainty of GMM selection, and the uncertainties of the

574   IM-MMI models. It should be emphasized that these results do not include every source of uncertainty.

575   Omitted, for example, are the uncertainties of: (i) fault depth and length; (ii) fault dip; and (iii) site V_{S30}. For

576   each of these inputs only a median prediction was used. For comparison, the CDF assigned to the Charleston

577   source in the NSHMP (Petersen et al., 2014, 2020) is also shown. Like the "epicenter search" results in

578   Figure 8, the AK07 and C20 models produce similar CDFs with a lesser median of ~M_w7.2, while WEA12

579   and CEA15 produce similar CDFs with a greater median of ~M_w7.7. It is readily apparent that the epistemic

580   uncertainty of IM-MMI model selection is considerable. To account for this uncertainty, each CDF was

581   weighted per the prior scheme to produce an ensemble CDF, as shown in Figure 10, which has a median of

582   ~M_w7.25 and 95% CI of M_w7.05 to M_w7.85. In this regard, efforts to better quantify the suitability of various

583   IM-MMI models to ENA, and in turn to refine the weighting scheme used herein, could have significant

584   influence on the overall conclusions of this study (and presumably also on studies of other seismic sources).

585   Despite this ambiguity, the ensemble CDF in Figure 10 suggests that ruptures larger than M_w7.5 – the

586   largest value considered in the NSHMP – could produce the observed evidence.

587   Of course, it should also be emphasized that the numerous hypothetical faults considered in our

588   methodology and aggregated to form the results in Figure 10 may not exist. That is, some of these gridded

589   faults are more likely to produce the 1886 MMI data than the hypothesized Woodstock Fault, but these

590   various faults are not necessarily present. Conversely, there is evidence (e.g., seismological data,

591   geophysical surveys) supporting the Woodstock Fault's existence (Durá-Gómez and Talwani 2009a,b;

592   Chapman et al., 2016). Accordingly, the preceding analyses are next repeated, conditioned solely on the

593   Woodstock Fault, thereby removing all source uncertainties aside from magnitude.
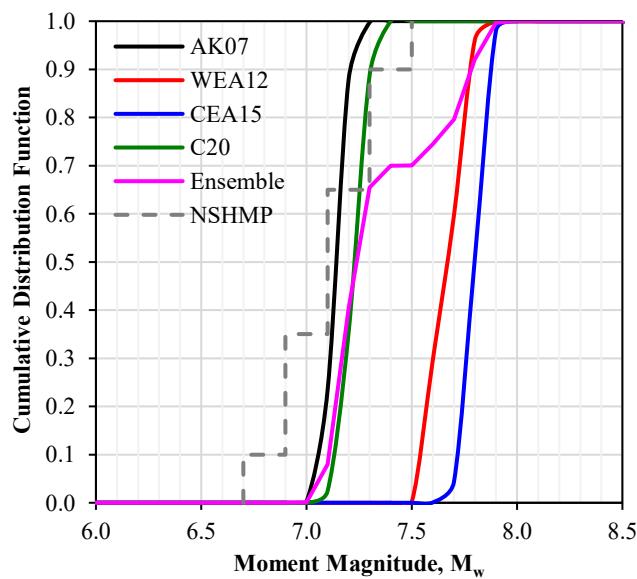
Figure 10. CDFs of rupture magnitude, as inverted from MMI observations using four IM-MMI models, wherein seismic sources are modelled as faults. AK07 = Atkinson and Kaka (2007), WEA12 = Worden et

al. (2012), CEA15 = Caprio et al. (2015), C20 = Cramer (2020), and NSHMP = National Seismic Hazard Model Project (Petersen et al., 2014, 2020).

594   *5.3 Analysis of MMI Data: Woodstock Fault*

595   While the parameters of the Woodstock Fault are uncertain, we adopt the source model of Durá-Gómez and

596   Talwani (2009a,b) exactly as proposed therein, the surface projection of which appears in Figures 1, 7, and

597   9. While Chapman et al. (2016) do not identify the Woodstock Fault by name, they infer and describe from

598   more recent seismological records a source that is similar: "We interpret …that the modern seismicity is

599   the lingering aftershock sequence of the 1886 shock and that the mainshock occurred on a south-striking,

600   west-dipping fault plane ..." To evaluate the prior site-to-source distance threshold of 600 km (i.e., that

601   which was used to exclude MMI observations), these analyses were also repeated using MMI data at

602   different site-to-source distances. Shown in Figure 11, for example, are the CDFs of M_w conditioned on the

603   Woodstock Fault and inverted from MMI data in two distance bins: (a) < 600 km; and (b) < 400 km. Several

604   observations are made from Figure 11 as follows.

605   First, it is observed that the ensemble CDF in Figure 11a is only marginally less uncertain than the

606   ensemble CDF in Figure 10, meaning that conditioning the analysis on the Woodstock Fault does not

607   significantly alter the results. In other words, the uncertainties of fault location and geometry are relatively

608   minor, given the data available for analysis. As previously discussed in Figure 9, these data are unable to

609   constrain the causative fault. That is, faults with diverse locations and orientations have similar likelihoods

610   of producing the MMI evidence, which may be due to a paucity of near-field observations. However,

611   because similar rupture magnitudes are inferred for these various faults irrespective of their positions, the

612   CDFs conditioned on the Woodstock Fault and the CDFs conditioned on an unknown fault are similar.

613   Second, a site-to-source distance bias is observed from Figures 11a and 11b. While the CDFs become

614   more vertical as more observations are included (reflecting a decrease in uncertainty) the CDFs also tend

615   to increase in magnitude. Observations at distances up to 600 km suggest larger rupture magnitudes than

616   those at distances up to 400 km. The median magnitude, for example, increases by ~0.25M_w according to

617   CEA15, by ~0.1M_w according AK07 and C20, and by ~0.01M_w according to WEA12. It could be shown

618   that the prior "epicenter search" and "fault search" results have a similar degree of sensitivity. This distance

619   bias could be present in either: (i) one or more of the component models utilized (e.g., the GMMs, site-

620   response model, or IM-MMI models); or (ii) in the MMI data itself. With respect to the adopted models, all

621   were shown to be unbiased during their respective trainings and cannot be further tested in the absence of

622   additional data. It is worth noting again, however, that the CEA15 model, which exhibits the greatest site-

623   to-source distance sensitivity in Figure 11, is herein applied to data much more distant than it was trained

624   on. It was for this reason that CEA15 was given low weighting despite having a large global training set.

625   With respect to the MMI data, we reiterate that analyzing more data over a wider spectrum of attenuation

25

626   is beneficial but remain adamant that the Bakun et al. (2002) dataset is apt to introduce bias at large

627   distances, given that small MMI values were systematically undocumented. To further probe the issue of

628   IM-MMI model applicability, the analyses were performed using only the data to which each model is

629   applicable (as stated by the original authors, or otherwise interpreted by the present authors). Specifically,

630   the WEA12, AK07, CEA15, and C20 models were respectively applied to observations within site-to-

631   source distances of 400 km, 800 km, 200 km, and 1500 km. The results are shown in Figure 11c. The most

632   salient changes are: (i) the CEA15 CDF is more uncertain (due to fewer field observations) and has a

633   reduced median of ~M_w7.2; and (ii) the C20 CDF is less uncertain (due to more field observations) and has

634   an increased median of ~M_w7.35. In summary, we are unsure why the results show evidence of site-to-

635   source distance bias at distances less than 600 km. However, we argue that the bias observed at larger

636   distances (e.g., the C20 result in Figure 11c) is at least partly due to the aforementioned bias of the Bakun

637   et al. (2002) dataset. As such, we prefer not to glean new conclusions from Figure 11c or champion its

638   results over those in Figure 11a. Our preferred CDF conditioned on the Woodstock Fault thus has a median

639   of M_w7.10 and 95% CI of M_w7.0 to M_w7.75. Given an unknown source, our preferred CDF (Figure 10) has

640   median of ~M_w7.25 and 95% CI of M_w7.05 to M_w7.85. For either result, the epistemic uncertainty of IM-

641   MMI model selection has significant influence.
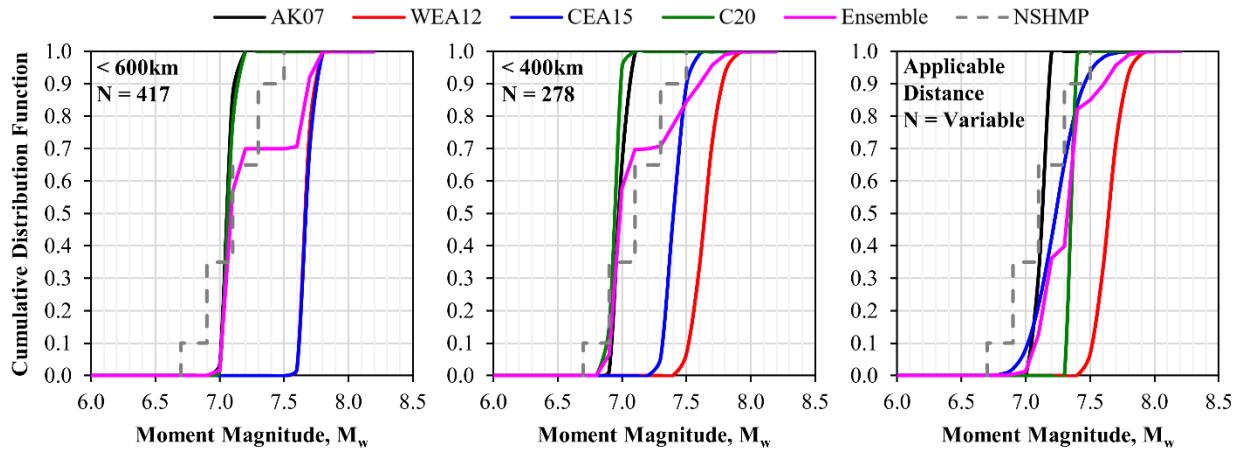
Figure 11. CDFs of rupture magnitude, as inverted from MMI observations using four IM-MMI models and assuming the Woodstock Fault (Durá-Gómez and Talwani, 2009a,b) to be the earthquake source. AK07 = Atkinson and Kaka (2007), WEA12 = Worden et al. (2012), CEA15 = Caprio et al. (2015), C20 = Cramer (2020); NSHMP = National Seismic Hazard Model Project (Petersen et al., 2014, 2020).

642   *5.4 Analysis of Liquefaction Data: Epicenter Search*

643   The inversion methodology was next applied to the 24 study sites summarized in Table 2 where the 1886

644   liquefaction response was observed, or has since been investigated, and where CPT testing has been

645   performed. Analogous to the analysis of MMI data, the seismic source was first treated as an epicenter with

unknown location. The resulting contours of seismic-source likelihood, which depict the epicenters most and least likely to produce the observed liquefaction response, are mapped in Figure 12, as are the 24 study sites. Also shown are instrumental epicenters and the hypothesized Woodstock Fault (Durá-Gómez and Talwani, 2009a,b). These initial results include no correction to account for the effects of soil aging on liquefaction. As shown in Figure 12, the analysis identified a source area east of the Woodstock fault. While the single most likely epicenter is 9 km from the proposed fault projection, epicenters within the projection have up to 70% relative likelihood. Notably, these geolocation results are largely insensitive to the treatment of soil aging. Because aging correction models tend to adjust the computed liquefaction resistances by similar amounts, the inverted magnitude tends to be significantly affected, whereas the inverted, most likely source location does not. For this reason, geolocation results are not shown for each aging-correction model. All give similar results to those in Figure 12, indicating that an epicenter more than ~20 km in any direction from the northern Charleston Peninsula (e.g., the Charleston Airport) is relatively unlikely to produce the observed evidence.
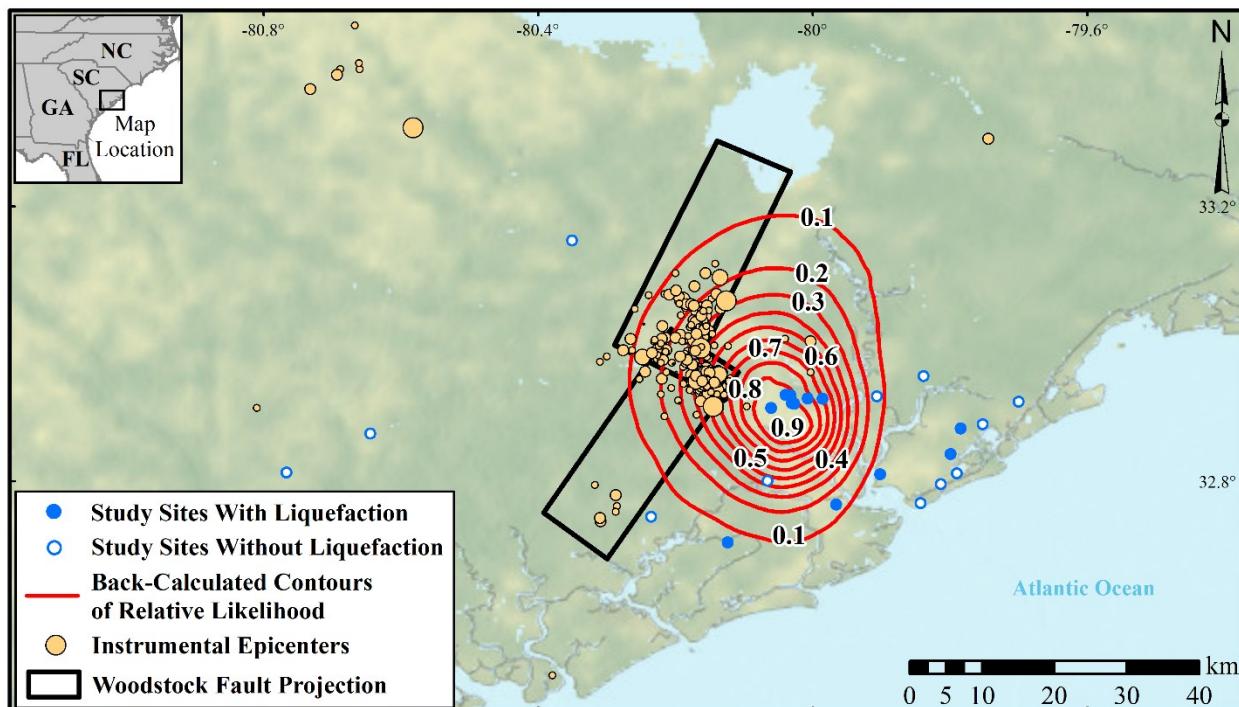
Figure 12. Likelihood contours indicating source locations that are, and are not, likely of producing the observed liquefaction response. Black rectangles = the Woodstock fault projection (Durá-Gómez and Talwani 2009a,b); tan circles = instrumental epicenters (M_w>1.0; 1900-present) (USGS, 2022).

Shown in Figure 13 are the CDFs of M_w inverted from liquefaction evidence, as computed using: (i) no aging correction; and (ii) each of the three correction methods discussed in the *Liquefaction Analysis*

661   *Procedure* (i.e., Models "A, B, and C"). It is apparent from Figure 13 that the uncertainties of *whether* and

662   *how* to correct for the effects of soil aging have significant influence, with the inverted median magnitudes

663   ranging from ~M_w6.3 (no aging correction) to ~M_w7.0 (Model C). Irrespective of soil aging, it is also

664   apparent that a magnitude inverted from the available liquefaction evidence is: (i) more uncertain than a

665   magnitude inverted from the available MMI evidence; and (ii) more uncertain than past studies have

666   reported. The former is attributable to there being far more MMI data than liquefaction data. The latter, as

667   previously discussed, is attributable to published uncertainty bounds (e.g., "M_w6.8-M_w7.0") being ranges

668   of the *median* magnitude considering one source of uncertainty, which is distinctly different from a CDF

669   of magnitude. Consequently, these results suggest more uncertainty than is adopted in the NSHMP.

670   Magnitudes below M_w6.7 and above M_w7.5, for example (i.e., the limits of the NSHMP weighting), have

671   non-trivial probabilities of producing the field evidence. It should also be noted that these results do not

672   include uncertainty within the aging-correction models (which are certainly uncertain). That is, the models

673   provide a median correction factor. To include the epistemic uncertainty of model selection, each CDF was

674   weighted to produce an ensemble. As with the MMI results, our selection of weights is based more on

675   judgement than on quantitative evidence. While we agree with the developers of aging-correction models

676   that corrections are likely warranted, these corrections are actively debated and likely have large

677   uncertainty. The weights selected and their justifications are as follows. Model C (0.5) is the only site-

678   specific method. It uses detailed subsurface geotechnical measurements from each study site to compute

679   site-specific corrections and thus arguably warrants the greatest weighting. In contrast, Models A (0.25)

680   and B (0.25) both rely on a sampling of the median characteristics of the geologic unit in which each study

681   site resides (e.g., the age of the unit). These models, which do not use site-specific information, are

682   therefore better suited for regional scale analyses (i.e., where subsurface data is unavailable). Nonetheless,

683   we argue these models warrant weighting given the overall uncertainty and debate surrounding aging

684   corrections. The ensemble CDF has a median of ~M_w6.90 and 95% CI of M_w6.05 to M_w7.85.
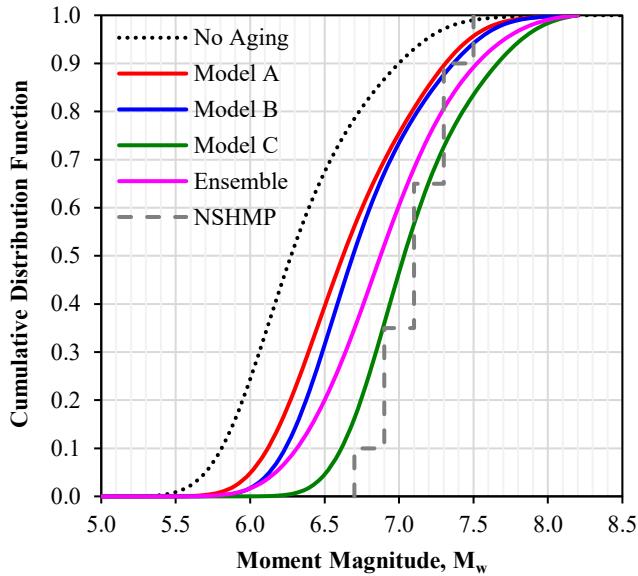
Figure 13. CDFs of rupture magnitude, as inverted from liquefaction observations using three age correction models, wherein seismic sources are modelled as epicenters. NSHMP = National Seismic Hazard Model Project (Petersen et al., 2014, 2020).

*5.5 Analysis of Liquefaction Data: Fault Search (Seismic Sources Modelled as Faults)*

To determine whether the liquefaction data can constrain the seismic source beyond a point location, the epicenters were next converted to faults. The discretization and treatment of hypothetical fault sources was identical to the MMI analyses. The results without correction for soil aging are shown in Figure 14. It can be seen that the data are incapable of constraining the seismic source to the Woodstock Fault, or to any other hypothetical alignment. The faults deemed most likely to produce the field evidence strike SE (like the predominant result from MMI analysis), but faults with nearly any orientation also have high relative likelihoods. These faults all strike through the general area previously identified via the epicenter search (i.e., through the northern Charleston Peninsula). While faults striking NE do have some likelihood of producing the evidence, these faults deviate from the proposed position of the Woodstock Fault. Results using each of the aging correction models are very similar to those in Figure 14 and are therefore not presented. The use of these models increases the inverted magnitudes (and thus increases the fault lengths in Figure 14), but otherwise has little effect on the inferred fault location or orientation. As with the MMI analyses, this should not be interpreted to mean that the Woodstock Fault or any similar alignment was not the source of the 1886 earthquake. Rather, this should be interpreted to mean that many aspects of the causative fault cannot be constrained with the available macroseismic evidence. This evidence is only marginally supportive of the Woodstock Fault's existence, which is not to say it doesn't exist. This conclusion might be different if more liquefaction study sites were available to the analysis, or if the component prediction models were less uncertain.
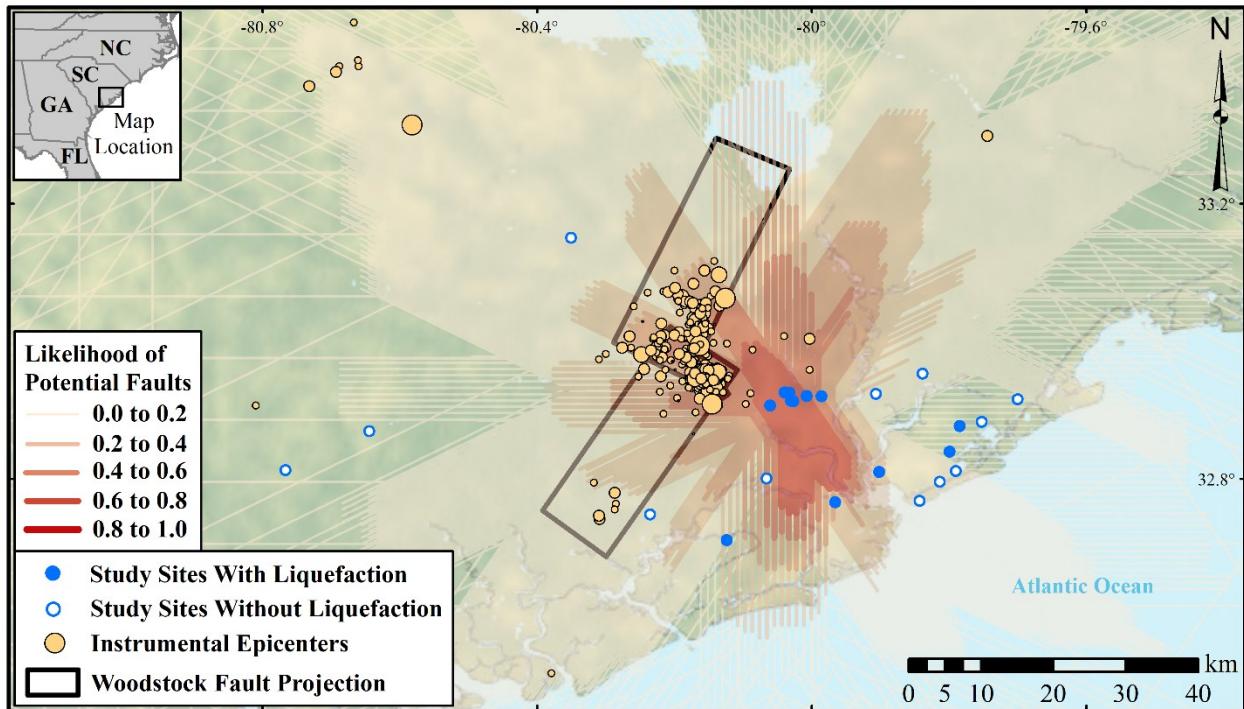
Figure 14. Likelihoods of hypothetical faults with differing orientations and lengths to produce the liquefaction evidence (no aging correction). Black rectangles = Woodstock fault projection (Durá-Gómez and Talwani 2009a,b); tan circles = instrumental epicenters (M_w>1.0; 1900-present) (USGS, 2022).

704   Shown in Figure 15 are the CDFs of M_w inverted from liquefaction data via the "fault search." The

705   large uncertainty of a magnitude inferred from the liquefaction data is again apparent, as is the influence of

706   soil aging and its correction. It is again emphasized that these results do not include every source of

707   uncertainty. Omitted, for example, are the uncertainties of: (i) fault depth and length; (ii) fault dip; (iii) site

708   V_{S30}; and (iv) the aging-correction factor. For each of these inputs only a median prediction was used. To

709   account for the uncertainty of selecting aging-correction models, each CDF was weighted per the prior

710   scheme to produce an ensemble CDF, as shown in Figure 15. This ensemble has a median of M_w6.95 and

711   95% CI of M_w6.15 to M_w7.85, suggesting a similar median magnitude as the NSHMP (Petersen et al., 2014,

712   2020) but much greater uncertainty.
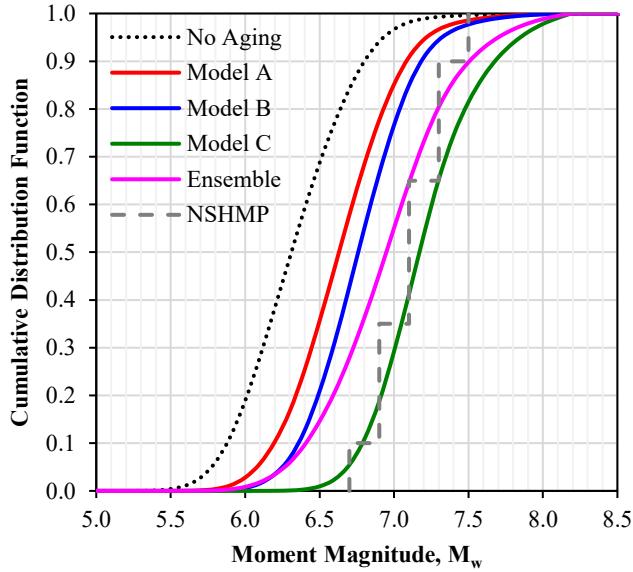
Figure 15. CDFs of rupture magnitude, as inverted from liquefaction observations using three age correction models, wherein seismic sources are modelled as faults. NSHMP = National Seismic Hazard Model Project (Petersen et al., 2014, 2020).

713 *5.6 Analysis of Liquefaction Data: Woodstock Fault*

714 Finally, the preceding analyses were repeated, but conditioned solely on the Woodstock Fault as

715 hypothesized by Durá-Gómez and Talwani (2009a,b). The computed CDFs of M_w, including the weighted

716 ensemble, are shown in Figure 16. Our preferred CDF (i.e., the ensemble) has a median of $M_w7.20$ and 95%

717 CI of $M_w6.30$ to $M_w8.10$. Relative to the CDF conditioned on an unknown source in Figure 15, conditioning

718 on the Woodstock Fault increases the inverted magnitude by ~$0.2M_w$ and increases the M_w uncertainty.

719 This counterintuitive increase in uncertainty is attributable to the epistemic uncertainty of selecting aging-

720 correction models, which is observed to increase with increasing rupture magnitude. Because conditioning

721 on the Woodstock Fault increases the inverted rupture magnitude, differences between the aging-correction

722 models become more apparent. Nonetheless, and similar to the analysis of MMI data, the uncertainties of

723 fault location and geometry appear relatively minor compared to other uncertainties, given the data

724 available for analysis. As shown in Figure 14, faults with diverse orientations were found to have similar

725 likelihoods of producing the liquefaction evidence. In other words, the data are sufficient to constrain the

726 source to a general area but are insufficient to constrain the source to a specific fault alignment. However,

727 because similar magnitudes are inferred for these faults irrespective of their positions, the CDFs conditioned

728 on the Woodstock Fault and the CDFs conditioned on an unknown source are similar.
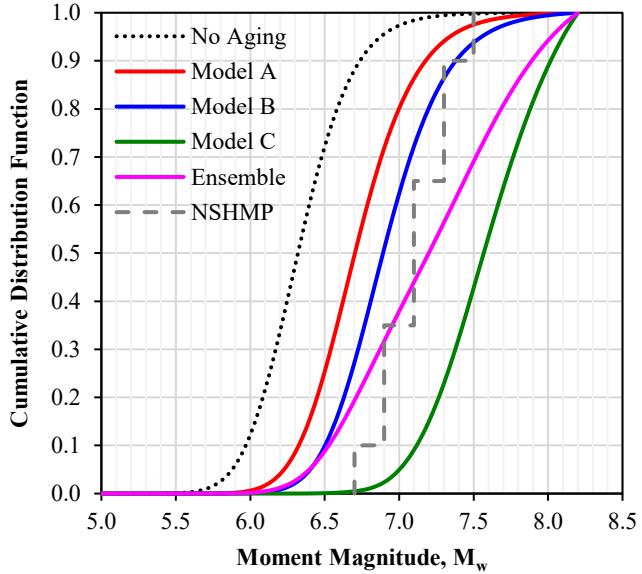
Figure 16. CDFs of rupture magnitude, as inverted from liquefaction observations using three age correction models and assuming the Woodstock Fault (Durá-Gómez and Talwani, 2009a,b) to be the earthquake source. NSHMP = National Seismic Hazard Model Project (Petersen et al., 2014, 2020).

729   *5.7 Forward Predictions of an 1886-Like Rupture*

730   Based on the preceding, which indicated that the intensity and liquefaction data alone cannot well constrain

731   the source of the 1886 event beyond a general region, we forward predict the median PGAs expected in an

732   "1886-like" rupture. As part of this effort, two fault locations and three magnitudes are considered. The

733   purpose of these predictions is to: (i) further illustrate why constraint of the source model is difficult, given

734   data limitations; and (ii) provide predictions for this scenario event using the latest predictive models, which

735   may be useful for regional-scale consequence modelling or science communication. Shown in Figure 17,

736   for example, are predictions considering an $M_w 7.1$ rupture and two strike-slip fault scenarios: (i) fault 1,

737   which is similar to the Woodstock fault; and (ii) fault 2, which is further east, perpendicular to fault 1, and

738   similar to hypothetical faults that were shown to have high likelihoods of producing the MMI and

739   liquefaction evidence (see Figures 9 and 14). Among previously hypothesized sources of the 1886

740   earthquake, fault 2 is most similar to the Deer Park lineament proposed by Marple and Hurd (2020),

741   although this fault is not as well supported by modern seismicity and is very different from most

742   hypothesized sources. Although the predicted PGAs do differ in the nearfield (e.g., within 50 km of the

743   source), the overall expectation at regional scale is obviously similar for two perpendicular faults.

744   Importantly, and as previously discussed, the Bakun et al. (2002) dataset does not contain an abundance of

745   1886 MMI observations in the vicinity of Charleston, or in the near field more generally. Only 16

746 observations, for example, are within 100 km of Charleston. As a result, the predicted PGAs, and by
747 corollary MMIs, are relatively similar at most observation sites.
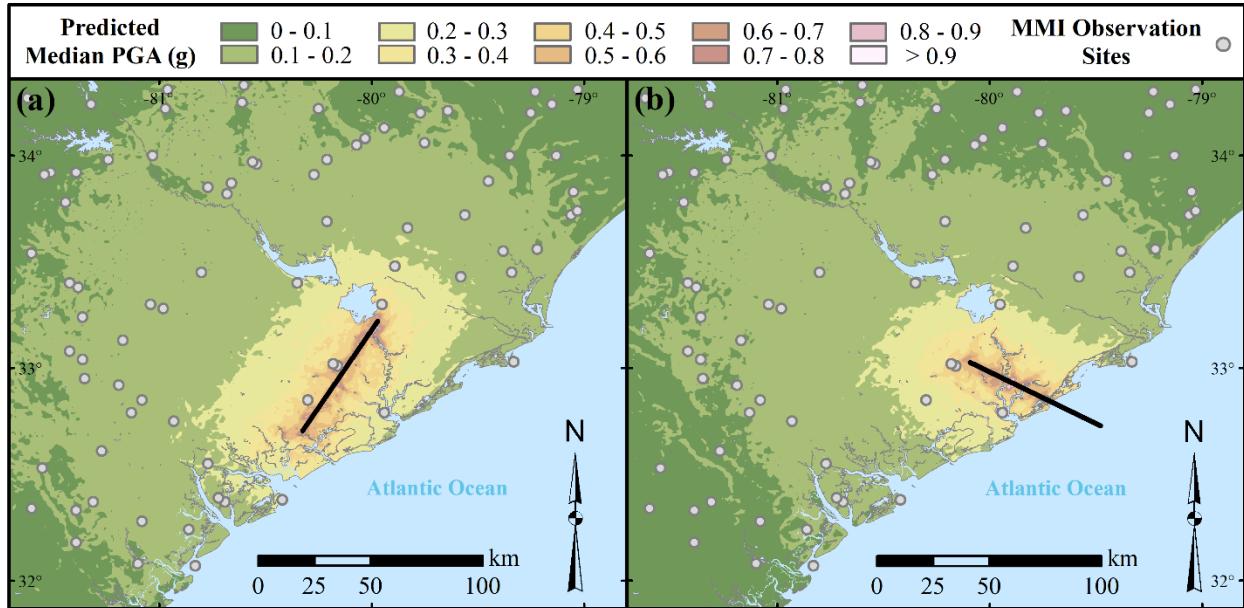
Figure 17. Predictions of median PGA in the Charleston region, as computed using 17 NGA-East GMMs (Goulet et al., 2018) and the Harmon et al. (2019) site-response model, and considering an $M_w7.1$ rupture on: (a) fault scenario 1; and (b) fault scenario 2, as described in the text.

748 Plotted in Figure 18a, considering these two fault scenarios, are the predicted MMIs at each of the
749 intensity report locations in the Bakun et al. (2002) dataset. For this example, the Atkinson and Kaka (2007)
750 IM-MMI model was adopted. It can be seen that large differences between the predicted MMIs (say, > 0.5
751 MMI) are rare. In turn, the difficulty of source-model constraint is apparent, given that the two scenarios
752 lead to perceptible differences at relatively few observation sites. This ambiguity is analogous to that which
753 arises when inverting faults from seismic displacements, wherein the classic "double couple" model
754 requires near-field or supplemental information to solve. Plotted in Figure 18b are the MMI values predicted
755 for fault scenario 1 versus the MMI values observed in 1886; an analogous plot is shown in Figure 18c for
756 fault scenario 2. While scenario 2 fits the data marginally better, as measured by mean absolute error, the
757 fits are essentially the same, as expected. Building off prior results, it cannot be concluded that scenario 1,
758 which approximates the Woodstock Fault, fits the MMI data better than alternative scenarios.
759 While a version of Figure 18a at sites of liquefaction evidence would show larger differences between
760 expected MMIs (given that all such sites are in the near field), the large finite-sample uncertainty of this
761 smaller dataset leads to a similar outcome, with a variety of faults having relatively high likelihoods of
762 producing the evidence (see Figure 14). It follows that additional near-field evidence (whether MMI or
763 liquefaction) could be especially influential to future studies of the 1886 earthquake. Following the same

33

764 approach, predictions were made for $M_w6.6$ and $M_w7.6$ ruptures. These results are shown in the electronic

765 supplement, where each of the forward predictions is also provided as a GIS map package.
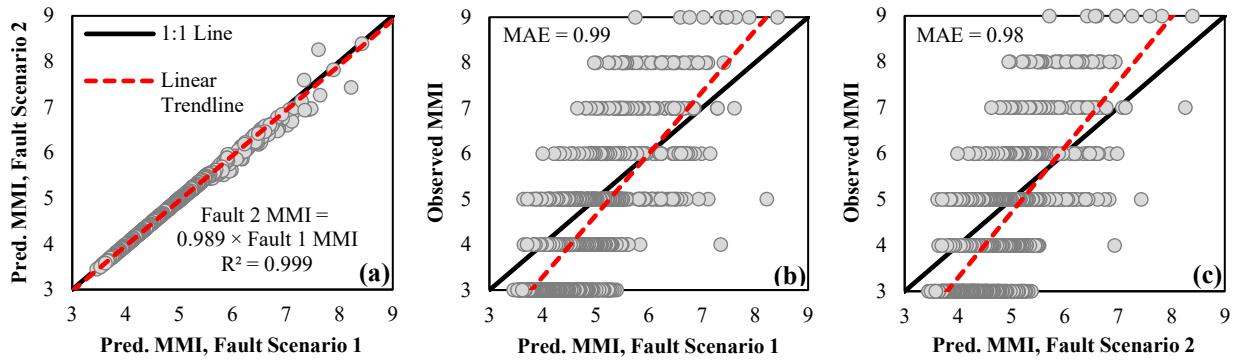
Figure 18. MMI values predicted by the Atkinson and Kaka (2007) model at sites of MMI observations: (a) fault scenario 1 predictions vs. fault scenario 2 predictions; (b) predictions vs. observations for fault scenario 1; (c) predictions vs. observations for fault scenario 2. MAE = mean absolute error.

766 *5.8 Limitations and Uncertainties*

767 The findings presented herein are inherently tied to the field evidence currently available and adopted for

768 analysis. Reinterpretation or augmentation of these observations, which we assume to be independent

769 events, would potentially change these findings, as would the adoption of new component models (e.g., to

770 predict ground motions, site response, MMI, liquefaction triggering, or liquefaction manifestation).

771 Undoubtedly, the modeling of these phenomena will continually advance, warranting future analyses of the

772 1886 Charleston earthquake. ENA ground-motion modeling has advanced, for example, yet GMMs remain

773 especially uncertain and untested at large magnitudes and will continue to evolve. And, as a supplement to

774 empirical GMMs, physics-based ground motion simulations could provide new insights into the 1886

775 rupture via more explicit modeling of influential site, path, and source effects (e.g., stress drop), as has been

776 shown for other historical earthquakes (e.g., Lozos, 2016). While empirical GMMs are ubiquitous in

777 earthquake science/engineering and implicitly account for many salient effects, a more explicit accounting

778 could help to reduce uncertainty. Moreover, the goal of this study was to assess the degree to which

779 macroseismic evidence can constrain the 1886 source model (i.e., rupture magnitude and position), rather

780 than to perform a broad investigation of all geophysical and modern seismological evidence, which would

781 fall under the purview of other investigators with different expertise. As one example, our inversions of

782 rupture magnitude conditioned on an unknown source allowed for randomly oriented fault sources.

783 However, with support of the interpreted regional stress regime, it may be reasonable to limit the range of

784 orientations considered, or alternatively, to subjectively assign weights to each orientation, thereby

increasing or decreasing the computed likelihoods, all else being equal. These possibilities could be investigated in the future. In addition, it should be noted that many, but not all, sources of uncertainty were accounted for in the analyses. Neglected, for example, were the uncertainties of site characterization (i.e., V_S30) and site observations (i.e., MMI and liquefaction responses). The inclusion of these and other uncertainties could potentially broaden the M_w CDFs computed herein. It must similarly be emphasized that a thorough reinterpretation of the more than 1000 original intensity reports, to include the assignment of site-specific uncertainties, bias corrections, and/or weighting schemes, could potentially change the results we present. And, as previously emphasized, our ensemble M_w CDFs utilize judgement-based weighting schemes. While justifications were provided for these weights, readers might argue for other weights, and thus draw other conclusions from our suites of M_w CDFs. Ultimately, future analyses will confirm or revise the conclusions reached in this study and summarized below.

6. Conclusions

Prior studies of MMI and liquefaction data resulting from the 1886 Charleston, SC, earthquake have several limitations. Namely, these studies tend to: (i) either be deterministic or account for uncertainties in an informal manner (e.g., it is often unclear what published uncertainty bounds represent and which uncertainties are, and are not, included); (ii) assume that the 1886 event was caused by a particular fault (i.e., the Woodstock Fault) without investigating the uncertainty of this assumption or the ability of the field data to constrain source traits beyond magnitude (i.e., fault location, geometry); and (iii) rely on models for predicting various phenomena (e.g., ground motions, site response, liquefaction response, MMI) that have since been superseded or augmented (e.g., by the NGA East project's 17 GMMs). Accordingly, this study presented probabilistic seismic-source inversions of the 1886 earthquake using the latest predictive models and a novel inversion methodology. With this approach, the likelihood of a rupture with some magnitude, location, and geometry to produce a set of field observations is computed. Repeating for enumerable hypothetical sources results in a regional scale constraint of the likely source traits, to the extent that observational data permits. With this approach, magnitude probability distributions conditioned on both an unknown source and on the hypothesized Woodstock Fault (Durá-Gómez and Talwani, 2009a,b) were computed and compared to that used in the U.S. NSHMP (Petersen et al., 2014, 2020). The most salient findings of this study, subject to the limitations and uncertainties summarized in section 5.8, are:

- Neither the location nor orientation of the 1886 fault rupture could be confidently constrained by the macroseismic evidence and models utilized herein, given the large uncertainties inherent to each. Hypothetical faults in a range of locations and with various alignments were deemed to have high relative likelihoods of producing this evidence. Considering both types of evidence and all analyses, faults striking SE and partly offshore were predominantly identified as having the greatest likelihood.

818   Yet these faults ranged in location and other faults, with very different strikes, were also often found
819   to be relatively likely.

820 • Collectively, the evidence does not provide strong support for the hypothesized Woodstock Fault.
821   One analysis (that of MMI data using the WEA12 IM-MMI model) found a NE-striking "Woodstock
822   like" fault to be the most likely source of the field evidence. Yet most analyses – in so far as
823   supporting the Woodstock Fault – point to the likelihood of a seismic source somewhere north of
824   Charleston but deem the Woodstock Fault to be relatively unlikely. This is not to say the fault does
825   not exist or could not have been the source of the 1886 rupture (geophysical investigations and
826   modern seismological data clearly suggest that active faults exist in the area), but rather, that many
827   aspects of the 1886 source model cannot be well constrained with the available macroseismic
828   evidence and models. This result might change if: (i) more MMI and liquefaction evidence were
829   available – particularly in the near field; (ii) the MMI evidence was reinterpreted to remove or correct
830   outliers, and to assign observation-specific uncertainties; or (iii) if the various required component
831   models were less uncertain.

832 • In the absence of these developments, certain aspects of the 1886 fault rupture can only be constrained
833   with other seismologic, geologic, and/or coseismic data and interpretation.

834 • When conditioned on the Woodstock Fault proposed by Durá-Gómez and Talwani (2009a,b), and
835   generally supported by others, our preferred CDF of M_w inverted from MMI data has a median of
836   M_w7.10 and 95% CI of M_w7.0 to M_w7.75. Of all uncertainties considered, the epistemic uncertainty
837   of IM-MMI model selection was larger than any other, since different models may give significantly
838   different MMI predictions for a given IM. In this regard, efforts to better quantify the suitability of
839   IM-MMI models to the study region, and in turn, to refine the weighting scheme used herein, could
840   have a significant influence on our overall conclusions.

841 • The results from MMI analysis show some site-to-source distance bias, with magnitudes inverted
842   from more distal MMI observations tending to be larger. Possible reasons for this bias were discussed
843   but could be further investigated in the future.

844 • When conditioned on the Woodstock Fault, our preferred CDF of M_w inverted from liquefaction data
845   has a median of M_w7.20 and 95% CI of M_w6.30 to M_w8.10. The greater M_w uncertainty, relative to
846   that from MMI analysis, is attributable to there being fewer liquefaction study sites, the compilation
847   of which requires both an observation of liquefaction response *and* subsurface geotechnical testing.
848   In addition, the uncertainties of whether and how to correct for soil-aging effects considerably
849   augmented the uncertainty. As such, efforts to quantify the suitability and uncertainties of aging-
850   correction models could have significant influence on our overall conclusions.

- When conditioned on an unknown seismic source, the CDFs of M_w inverted from MMI and liquefaction data did not differ greatly from the CDFs conditioned on the Woodstock Fault. Our interpretation is that while faults with a range of locations and alignments were found similarly likely to produce the field evidence, these faults were inferred to have mostly similar magnitude distributions. Thus, the results of this study pertaining to the magnitude of the 1886 rupture would not necessarily change if the rupture's exact position was known. This should not be interpreted to mean that the uncertainties of source location and geometry are inconsequential to inverse-analyses of macroseismic data. As demonstrated via simulated inversions of modern earthquakes (Rasanen and Maurer, 2021, 2022), these uncertainties are often considerable. This is especially the case, for example, when the effects of a distant, large M_w rupture cannot be distinguished from the effects of a nearby, small M_w rupture. In such cases, constraint of the source location can dramatically reduce the overall uncertainty of the inverted magnitude.

- Collectively, the results largely support the M_w distribution adopted by the NSHMP, which ranges from $M_w 6.7$ to $M_w 7.5$ with a median of $M_w 7.1$. While analyses indicate that $M_w < 6.7$ ruptures have potential to produce the observed liquefaction response (particularly when the uncertainty of selecting an aging-correction model is considered), the MMI evidence suggests a near-zero likelihood of such magnitudes. Conversely, both types of evidence suggest that $M_w > 7.5$ events have potential to produce the field evidence. While this conclusion hinges on which component models are adopted to predict soil-aging effects and MMI, we fail to find conclusive evidence for outright rejecting some models in favor of others. In the absence of such evidence, the possibility of $M_w > 7.5$ ruptures would merit consideration.

- Ultimately, the flexible inversion methodology employed herein is not specific to ENA, or to certain types of macroseismic evidence, but rather is applicable to any seismic zone and to any co-seismic response for which probabilistic prediction models exist. This methodology allows for uncertainty to be accounted for in a more complete and transparent manner when inverting seismic parameters from macroseismic evidence.

Declaration of competing interests

The authors declare no competing interests.

Silvia Mazzoni of the John Garrick Institute for Risk Sciences at UCLA for developing NGA East ground

motion characterization tools based on the results of Goulet et al. (2018), which were utilized in this study.

Appendix A. Supplementary data

All data and models utilized in this study are publicly available, as cited herein. The supplemental material

includes 13 additional figures that largely pertain to the analysis of MMI data using the "epicenter search"

and "fault search" methods. The results presented in these figures are summarized and discussed in the

main text. Also included are maps (figures and GIS files) of the predicted median PGAs in an "1886-like"

event, considering three magnitudes and two fault alignments, as well as measurements from 24 CPTs.

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