RESEARCH ARTICLE | FEBRUARY 13 2024

Nonlinear Landau resonant interaction between whistler waves and electrons: Excitation of electron-acoustic waves

Donglai Ma ➡ ⑩ ; Xin An ⑩ ; Anton Artemyev ⑩ ; Jacob Bortnik ⑩ ; Vassilis Angelopoulos ⑩ ; Xiao-Jia Zhang ⑮



Phys. Plasmas 31, 022304 (2024) https://doi.org/10.1063/5.0171227









Nonlinear Landau resonant interaction between whistler waves and electrons: Excitation of electron-acoustic waves

Cite as: Phys. Plasmas **31**, 022304 (2024); doi: 10.1063/5.0171227 Submitted: 7 August 2023 · Accepted: 24 January 2024 · Published Online: 13 February 2024







Donglai Ma, $^{1,a)}$ (b) Xin An, 2 (b) Anton Artemyev, 2 (b) Jacob Bortnik, 1 (b) Vassilis Angelopoulos, 2 (b) and Xiao-Jia Zhang 2,3 (b)

AFFILIATIONS

- Department of Atmospheric and Oceanic Sciences, University of California, Los Angeles, California 90095, USA
- ²Department of Earth, Space and Planetary Sciences, University of California, Los Angeles, California 90095, USA

ABSTRACT

Electron-acoustic waves (EAWs) as well as electron-acoustic solitary structures play a crucial role in thermalization and acceleration of electron populations in Earth's magnetosphere. These waves are often observed in association with whistler-mode waves, but the detailed mechanism of EAW and whistler wave coupling is not yet revealed. We investigate the excitation mechanism of EAWs and their potential relation to whistler waves using particle-in-cell simulations. Whistler waves are first excited by electrons with a temperature anisotropy perpendicular to the background magnetic field. Electrons trapped by these whistler waves through nonlinear Landau resonance form localized field-aligned beams, which subsequently excite EAWs. By comparing the growth rate of EAWs and the phase mixing rate of trapped electron beams, we obtain the critical condition for EAW excitation, which is consistent with our simulation results across a wide region in parameter space. These results are expected to be useful in the interpretation of concurrent observations of whistler-mode waves and nonlinear solitary structures and may also have important implications for investigation of cross-scale energy transfer in the near-Earth space environment.

© 2024 Author(s). All article content, except where otherwise noted, is licensed under a Creative Commons Attribution-NonCommercial 4.0 International (CC BY-NC) license (http://creativecommons.org/licenses/by-nc/4.0/). https://doi.org/10.1063/5.0171227

I. INTRODUCTION

During geomagnetically active times, fast plasma flows in the Earth's plasma sheet transport energetic particles into the inner magnetosphere and form injection fronts (or dipolarization fronts; see reviews in Refs. 10, 11, and 21). In the leading edge of injection fronts, the magnetic B_z component (in the geocentric solar magnetospheric coordinate) typically has an abrupt enhancement, indicating a "dipolarization" of the geomagnetic field. 7,9,31,43 Injections consist of abruptly enhanced fluxes of high-energy ions and electrons in the energy range of 10 s-100 s of keV, which provide free energy to various plasma waves. 13,14,33 Indeed, a broad spectrum of electromagnetic emissions, extending from Doppler-shifted kinetic Alfvén waves of a few Hz to electron cyclotron harmonic waves of \sim 10 kHz, is embedded within dipolarization fronts and constitutes a significant fraction of the total energy transport in fast plasma flows. Among these emissions, whistler waves are excited by injected energetic electrons with a perpendicular ^{25,27,30,60} Electron-acoustic waves (EAWs), temperature anisotropy. 12 as well as electron-acoustic solitary structures, identified as broadband

electrostatic turbulence, are often observed in association with whistler waves. ^{2,15,17,29,34,38,41,56} Under typical conditions of plasma injections into the inner magnetosphere, EAWs and their related solitary structures effectively scatter electrons of 10 eV–1 keV in both pitch angle and energy, ^{8,37,45,53} whereas whistler waves provide electron scattering in the 1–100s keV energy range. ^{23,24,40,46,48,51} The concurrence of EAWs and whistler waves potentially results in a wide energy range for electron precipitation and acceleration. ³²

Linear wave theory, confirmed by spacecraft observations, indicates that even slightly oblique whistler waves (e.g., $\lesssim 30^\circ$ of wave normal angle with respect to the background magnetic field) have finite parallel electric fields, and trap electrons through the nonlinear Landau resonance. ^{1,4,29,61} It was demonstrated that such trapped electron populations can form beams that are unstable to the generation of EAWs or other electron-acoustic solitary structures with spatial scales on the order of tens of Debye lengths (e.g., double layers, phase space holes, also known as time domain structures; see Refs. 33, 34, and 54). The ratio of Landau resonant velocity to electron thermal velocity

³Department of Physics, University of Texas at Dallas, Richardson, Texas 75080, USA

^{a)}Author to whom correspondence should be addressed: dma96@atmos.ucla.edu

controls the type of nonlinear wave structures generated by electron populations trapped by whistler waves. The same mechanism of electron Landau trapping by lower frequency waves and further generation of higher frequency electrostatic structures has been confirmed to work for the excitation of electron-acoustic solitary structures through interactions between kinetic Alfvén waves and thermal electrons.3 It is worth mentioning that an alternative mechanism was proposed for the formation of electric field spikes, such as nonlinear fluid steepening of electron-acoustic modes.^{2,56} The link between whistler waves (or kinetic Alfvén waves) and EAWs (or solitary structures) provides a potentially important channel of energy transfer: Injected ions and electrons accelerated by the electromagnetic fields of dipolarization fronts at the macroscale (tens to hundreds of ion inertial length) first excite kinetic Alfvén and whistler waves at the intermediate scale (a few ion or electron inertial lengths), which subsequently generate EAWs and solitary structures at the microscale (tens of Debye lengths) and eventually deposit energy into thermal electrons. Such energy transfer from the macroscale to the microscale may contribute to electron thermalization and heating during one of the most energetic processes in the Earth's magnetosphere-the plasma injection and braking of fast plasma flows in the inner magnetosphere.

It is the aim of this study to explore the coupling from whistler waves to EAWs using particle-in-cell (PIC) simulations and to determine the favorable conditions for such coupling to occur. It is organized as follows: In Sec. II, we briefly describe the computational setup, in which whistler waves are naturally generated by energetic electrons with a perpendicular temperature anisotropy. In Sec. III, we investigate how trapped electron beams are formed through nonlinear Landau resonance between electrons and whistler waves, and how such electron beams excite EAWs and make EAWs survive. In Sec. IV, we derive the critical condition for EAW excitation and confirm its validity by comparing it with simulation results. We summarize our results in Sec. V.

II. COMPUTATIONAL SETUP

We use a fully relativistic, electromagnetic PIC code called OSIRIS 4.0. 19 Our simulations have two dimensions (2D) in configuration space and three dimensions in velocity space. The computation domain in the x-y plane consists of 625 \times 625 cells. Each cell contains 400 particles. We use periodic boundary conditions for both particles and fields. The background magnetic field \mathbf{B}_0 is in the x direction. The normalized strength of \mathbf{B}_0 is set as $\omega_{ce}/\omega_{pe}=0.25$, typical in the generation region of whistler waves in the inner magnetosphere. 20,49 Here, ω_{ce} is the electron gyrofrequency, and ω_{pe} is the electron plasma frequency. The plasma is initially uniform in space. Because our frequency range of interest is $\omega \gg \omega_{ci}$ (the ion gyrofrequency), ions are immobile as a charge-neutralizing background. Electrons are initialized with a single bi-Maxwellian distribution with a temperature anisotropy of $A = T_{\perp}/T_{\parallel} > 1$. The bi-Maxwellian model is a theoretical and typical construct in order to carry out the simulations. It is worth noting that the observed distributions sometimes may deviate significantly from Maxwellians.²⁰ The electron parallel beta is defined as

$$\beta_{||} = \frac{n_0 m_e v_{T_{||}}^2}{B_0^2 / 8\pi} = \frac{2v_{T_{||}}^2}{\left(c\omega_{ce}/\omega_{pe}\right)^2} = \frac{2v_{T_{||}}^2}{v_{Ae}^2},\tag{1}$$

where m_e is the electron mass, c is the speed of light, n_0 is the plasma density, v_{Ae} is the electron Alfvén speed, and $v_{T_{\parallel}}$ is the electron thermal

velocity in the parallel direction. β_{\parallel} describes the magnitude of the electron thermal velocity relative to the characteristic whistler phase velocity (0.5 v_{Ae} being the whistler phase velocity at 0.5 ω_{ce} and 0 degree wave normal angle). Given A and β_{\parallel} , we determine the initial T_{\perp} and T_{\parallel} and initialize the electron velocity distribution. The cell length Δx in both directions is set between λ_D and $2\lambda_D$, where the initial electron Debye length $\lambda_D = v_{T_{\parallel}}/\omega_{pe}$ (neglecting ions). For 2D simulations, the time step Δt is constrained by the Courant condition

$$\Delta t \lesssim \frac{\Delta x}{\sqrt{2}c}$$
. (2)

To understand the critical condition of EAW excitation, we scan the parameter space of β_{\parallel} and T_{\perp}/T_{\parallel} . In this scan, β_{\parallel} is varied from 0.005 to 0.3 with a logarithmic step, and T_{\perp}/T_{\parallel} is varied in the sequence (3.5, 4, 4.5, 5). The detailed parameters in each simulation are shown in Table I. Such choice is made to facilitate numerical work, and the observed anisotropies in space may deviate from the assumed values.⁵

III. EXCITATION OF EAWS BY WHISTLER WAVES THROUGH NONLINEAR ELECTRON TRAPPING

Whistler waves can be excited by the free energy provided by electron perpendicular temperature anisotropy. Linear kinetic theory and PIC simulations in previous works have shown that the dominant mode of maximum growth rate is in the parallel direction for $\beta_{\parallel} \gtrsim 0.025$ and shifts to the oblique direction for $\beta_{\parallel} \lesssim 0.025$. Here, we examine the evolution of the fields, as well as the electron Landau trapping in these fields, in the two different β_{\parallel} regimes and demonstrate that both regimes support electron-acoustic wave (EAW) excitation.

A. Small β_{\parallel} regime

Figure 1 illustrates the wave characteristics in the small β_{\parallel} regime of $\beta_{\parallel}=0.0091$ and $T_{\perp}/T_{\parallel}=5.0$ (case 8 in Table I). The whistler waves are first excited at $t\sim300\omega_{pe}^{-1}$ at the parallel wave number $k_x\sim3\omega_{pe}/c$. We denote the parallel whistler wave number as k'_{\parallel} and the EAW wave number as k_{\parallel} . The EAWs start to be excited at $t\sim1000\omega_{pe}^{-1}$ with the wave number k_{\parallel} ranging from $5\omega_{pe}/c$ to $20\omega_{pe}/c$ [Fig. 1(a)]. The upper band whistler-mode with maximum power, located at $\omega=0.18\omega_{pe}=0.72\omega_{ce}$ and $k'_{\parallel}=3.26\omega_{pe}/c$, has a parallel phase speed $v_{ph,\parallel}=\omega/k'_{\parallel}=0.055c$, which is about the same as (or slightly smaller than) that of EAWs [Fig. 1(b)]. Two representative time snapshots of wave fields (including δE_x and δB_y) at $t_A=800\omega_{pe}^{-1}$ and $t_B=1100\omega_{pe}^{-1}$, before and after the excitation of EAWs, respectively, are displayed in Figs. 1(c)–1(h). The

TABLE I. The detailed parameters for 32 simulations. The unit of cell length Δx is c/ω_{pe} , and the unit of time step Δt is ω_{pe}^{-1} . Each group has four cases corresponding to temperature anisotropies $T_{\perp}/T_{\parallel}=3.5,\,4,\,4.5,\,$ and 5.

Case No.	$oldsymbol{eta}_{\parallel}$	Δx	Δt	Case No.	$oldsymbol{eta}_{\parallel}$	Δx	Δt
1-4	0.005 00	0.0216	0.0145	17-20	0.052 68	0.040	0.0270
5-8	0.00901	0.0216	0.0145	21-24	0.09491	0.056	0.0370
9-12	0.01623	0.0216	0.0145	25-28	0.17100	0.072	0.0476
13–16	0.029 24	0.0300	0.0200	29-32	0.308 08	0.098	0.0670

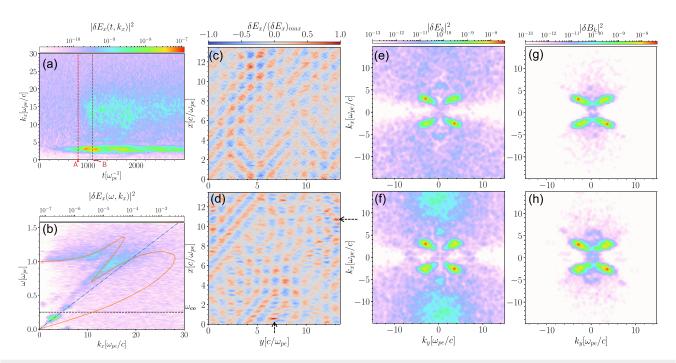


FIG. 1. Whistler anisotropy instability and the associated excitation of EAWs for $\beta_{\parallel}=0.0091$ and $T_{\perp}/T_{\parallel}=5.0$. (a) The wave number spectrum of the parallel electric field $\delta E_{\rm x}$ as a function of time. (b) The dispersion diagram of $\delta E_{\rm x}$ as a function of frequency and wave number. The black dashed line represents the electron cyclotron frequency $\omega_{\rm ce}=0.25\omega_{\rm pe}$. The blue dash-dotted line is the parallel phase velocity of the whistler wave $v_{\rm phase}$, $P=0.0546c=3.25v_{\rm T,\parallel}$. The orange line shows the EAW dispersion calculated from the electron distribution at $t=1100\omega_{\rm pe}^{-1}$. (c) and (d) The pattern of $\delta E_{\rm x}$ at two time snapshots, $t=800\omega_{\rm pe}^{-1}$ and $t=1100\omega_{\rm pe}^{-1}$, respectively. (e) and (f) The power density of $\delta E_{\rm x}$ in the wave number space at $t=800\omega_{\rm pe}^{-1}$ and $t=1100\omega_{\rm pe}^{-1}$, respectively.

electromagnetic, relatively long-wavelength ($\lambda \sim 1.7c/\omega_{pe}$) whistler waves propagate in the oblique direction with a wave normal angle (WNA) $\sim 45^{\circ}$, whereas the electrostatic, short-wavelength EAWs ($\lambda \sim 0.5c/\omega_{pe} \sim 30\lambda_{De}$) propagate in the parallel direction. Such electron-acoustic mode manifests as solitary structures in certain spatial domains [see such domains pointed by arrows in Fig. 1(d)] and, hence, has a broadband wave number spectrum.

The perturbed magnetic field amplitude reaches $\sim 0.01B_0$ at $t_B = 1100\omega_{pe}^{-1}$. For large-amplitude whistler waves propagating in oblique directions, their parallel electric field can trap the electrons moving near the parallel phase speed $v_{ph,\parallel}$ (i.e., the Landau resonant velocity) in its potential well, which is the so-called nonlinear Landau resonance. 35 The response of trapped electrons to whistler waves is characterized by the formation of electron beams in the resonant islands around $\pm v_{ph,\parallel}$ as shown in Fig. 3(b). The resonant electrons are accelerated in the phase of $v_x \cdot \delta E_x < 0$, whereas they are decelerated in the phase of $v_x \cdot \delta E_x > 0$. This transport process in the velocity space gives rise to spatially modulated beams, which subsequently excite time domain structures (TDSs) as shown in Fig. 1(d). These TDSs are identified as the nonlinear electron-acoustic mode^{3,52} and survive Landau damping because of the plateau distribution created by the electron trapping [Fig. 4(a)]. It is worth noting that the beam velocity is slightly larger than $v_{ph,\parallel}$ [Fig. 3(b)], which makes the EAW phase velocity slightly larger than $v_{ph,\parallel}$ located at the center of the resonant island [Fig. 1(b)].

We further use the reduced, parallel velocity distribution obtained at $t_B=1100\omega_{pe}^{-1}$ (Fig. 4) to numerically solve for the dispersion relation of EAWs. The plateau at the wave phase velocity (i.e., $\partial f_0/\partial v|_{\omega/k}\simeq 0$) allows retaining only the principal part of the integral around the Landau contour, ⁵²

$$1 - \frac{1}{k^2} \int_{V} dv \frac{\partial f_0 / \partial v}{v - \omega / k} = 0, \tag{3}$$

with

$$\int_{L} dv \frac{\partial f_0 / \partial v}{v - \omega / k} = \text{p.v.} \int_{-\infty}^{+\infty} dv \frac{\partial f_0 / \partial v}{v - \omega / k},$$
(4)

where the subscript "L" denotes the Landau contour, and "p.v." represents the principal value integral. The roots of the dispersion relation yield the orange solid curve in Fig. 1(b). The "thumb dispersion," computed from a Maxwellian distribution assuming an infinitesimal resonant island, ⁵² is modified to branch at the beam velocity (i.e., the parallel phase velocity of whistler waves) due to the finite width of the resonant island in our simulation. This calculation clearly shows that the EAWs are excited by trapped beams approximately at the parallel phase velocity of whistler waves.

B. Large β_{\parallel} regime

Figure 2 shows the wave characteristics for $\beta_{\parallel}=0.05268$ and $T_{\perp}/T_{||}=4.5$ (case 19 in Table I). The whistler mode with the

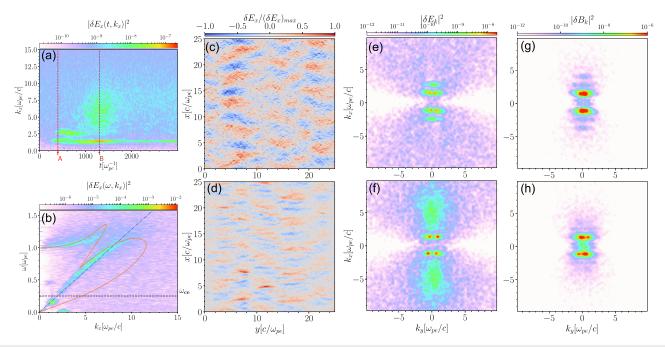


FIG. 2. The whistler anisotropy instability and associated excitation of EAWs for $\beta_{\parallel}=0.052$ 68 and $T_{\perp}/T_{\parallel}=4.5$. The format is the same as that of Fig. 1. The two selected snapshots are at $t_A=400\omega_{p_0}^{-1}$ and $t_B=1300\omega_{p_0}^{-1}$. The phase velocity of whistler waves is $v_{\text{phase},\parallel}=0.1283c=3.16v_{T,\parallel}$.

maximum linear growth rate in this case is in the parallel direction. The excited upper band whistler mode waves with maximum power are located at $\omega = 0.16\omega_{pe} = 0.64\omega_{ce}$, and the energy density of the wave magnetic field is maximized at the wave number $k_{x,max}$ $=\pm 1.26\omega_{pe}/c$ and $k_{\nu,\text{max}}=0$ [Figs. 2(g) and 2(h)]. It is interesting that there are low-frequency quasi-electrostatic modes at $\omega \sim 0$ and $k_x = 2k_{x,\text{max}}$ as shown in Figs. 2(a), 2(b), and 2(e). This standing electric field is generated because counter-propagating whistler waves are naturally excited by electron temperature anisotropy in a uniform magnetic field, similar to the spacecraft observations of chorus source region in the equatorial magnetosphere.⁵⁰ Specifically, the electron fluid should be force free parallel to the background magnetic field, $E_x + (\mathbf{v}_e \times \mathbf{B})_x = 0$ (neglecting density and temperature gradients for now). It can be shown that $(\mathbf{v}_e \times \mathbf{B})_x$ is finite by averaging over the fast wave periods for two counter-propagating whistler waves. 44 Thus, a standing electric field is generated.

The energy density of the slightly oblique whistler waves is still finite in the large β_{\parallel} regime [Figs. 2(c)–2(h)]. The parallel electric field of these oblique waves excites electron-acoustic modes via the same mechanism as the small β_{\parallel} regime. Figures 2(a) and 2(b) show that the electron-acoustic modes start to be excited at $t\sim 1000\omega_{pe}^{-1}$ and propagate at a phase velocity slightly larger than that of the whistler waves. These electron-acoustic modes appear as unipolar structures [Fig. 3(d)], rather than wave-like structures in case 8 [Fig. 3(a)]. Such a difference is likely due to the larger beam density of case 19 than that of case 8 [Figs. 3(b), 3(e), and 4], consistent with a previous study. 4

IV. CRITICAL CONDITION FOR EAW EXCITATION

The simulations demonstrate that EAWs can be excited in both low and high beta regimes. This leads to a question regarding

the essential criteria for EAW excitation. As β_{\parallel} changes, the parallel electric field amplitude and the phase velocity of the whistler waves change accordingly. These factors control the density and velocity of the trapped electron beams, which further control the growth rate of the beam instability. Furthermore, EAWs are subject to Landau damping, which depends on their phase velocities (or equivalently, the whistler wave phase velocities, we use the same $v_{ph,\parallel}$ in the following). Moreover, trapped electrons undergo phase mixing, 35 which smooths the beam distribution over time. The interplay between these processes determines the critical conditions for exciting EAWs driven by whistler waves. Our approach for solving this problem follows a similar method as outlined in Ref. 3, which explored the interaction between kinetic Alfvén waves and thermal electrons.

The concept of trapped beams involves electrons being confined within the wave potential well, oscillating at a frequency $\omega_{tr} = \sqrt{ek'_{\parallel}|\delta E_{\parallel}|/m_e}$ at the bottom of the potential well. ³⁹ Consequently, the half-width of the trapping island or the plateau created by Landau resonance can be estimated as

$$\frac{\Delta v_{tr}}{v_{T_{\parallel}}} = \frac{2\omega_{tr}}{k'_{\parallel}v_{T_{\parallel}}} = \frac{2}{k'_{\parallel}v_{T_{\parallel}}} \sqrt{\frac{ek'_{\parallel} \cdot k'_{\parallel}\delta\phi}{m_e}} = 2\sqrt{\frac{e\delta\phi}{T_{\parallel}}}.$$
 (5)

To compute the linear growth rate of the trapped electron beam, we model the parallel electron distribution by separating it into a background Maxwellian distribution $f_{0e} = \frac{1}{\sqrt{2\pi v_{T\parallel}}} \exp(-\frac{v_{\parallel}^2}{2v_{T\parallel}^2})$ [i.e., a reduced distribution normalized to a density of Eq. (1)] and a

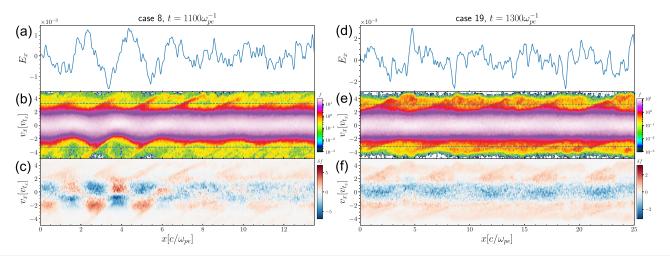


FIG. 3. Parallel electric fields and phase space densities in the two β_{\parallel} regimes. Left column: case 8 in the small β_{\parallel} regime at $t=1100\omega_{pe}^{-1}$. (a) The parallel electric field as a function of position x. (b) The phase space density of electrons in $v_x - x$ space. The blue dashed line shows the parallel phase speed of whistler waves. (c) The difference of electron phase space densities between $t=1100\omega_{pe}^{-1}$ and $t=0\omega_{pe}^{-1}$. Right column: case 19 in the large β_{\parallel} regime at $t=1300\omega_{pe}^{-1}$. (d)–(f) have the same format as (a)–(c), respectively.

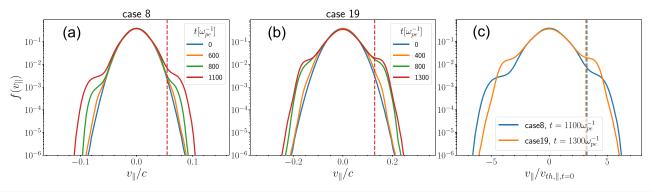


FIG. 4. The reduced electron distributions as a function of the parallel velocity for cases 8 and 19. (a) Case 8 at t=0,600,800, and $1100\omega_{pe}^{-1}$. (b) Case 19 at t=0,400,800, and $1300\omega_{pe}^{-1}$. (c) A comparison of phase space densities from both cases, for which velocities are normalized to the initial thermal velocity. The red dashed lines are the parallel phase speed of whistler waves.

perturbed distribution δf_{tr} around the parallel phase speed of the whistler wave $v_{vh.\parallel}$,

$$\delta f_{tr} = \Delta f \cdot \sin\left(\pi \frac{v_{||} - v_{ph,||}}{\Delta v_{tr}}\right). \tag{6}$$

The magnitude of the perturbed distribution can be estimated as

$$\Delta f = \frac{f_{0e}(v_{ph,\parallel} - \Delta v_{tr}/2) - f_{0e}(v_{ph,\parallel} + \Delta v_{tr}/2)}{2}$$

$$= \sum_{n=0}^{\infty} \frac{H_{2n+1}(\frac{v_{ph,\parallel}}{\sqrt{2}v_{T\parallel}})}{(2n+1)!} (\frac{\Delta v_{tr}}{2\sqrt{2}v_{T\parallel}})^{2n+1} f_{0e}(v_{ph,\parallel}), \qquad (7)$$

where we perform a Taylor expansion at the whistler phase velocity $v_{\parallel} = v_{ph,\parallel}$, and the Hermite Polynomials $H_{2n+1}(\cdot)$ are used for calculating derivatives of the Maxwellian distribution. The growth rate of the beam instability can finally be written as

$$\frac{\gamma_{tr}}{k_{\parallel}v_{T\parallel}} = \frac{\pi}{2} \frac{\omega}{k_{\parallel}v_{T\parallel}} \frac{\omega_{pe}^{2}}{k_{\parallel}^{2}} \left(\frac{\partial}{\partial v_{\parallel}} \delta f_{tr}\right)_{v_{\parallel}=\omega/k_{\parallel}}$$

$$\sim \frac{\pi}{2} \frac{\omega}{k_{\parallel}v_{T\parallel}} \frac{\omega_{pe}^{2}}{k_{\parallel}^{2}} \left(\frac{\partial}{\partial v_{\parallel}} \delta f_{tr}\right)_{v_{\parallel}=v_{ph,\parallel}}$$

$$= \frac{\pi}{2} \frac{\omega}{k_{\parallel}v_{T\parallel}} \frac{1}{k_{\parallel}^{2} \lambda_{D}^{2}} \frac{\pi}{2\sqrt{2}} v_{T\parallel} f_{0e}(v_{ph,\parallel})$$

$$\times \sum_{n=0}^{\infty} \frac{H_{2n+1} \left(\frac{v_{ph,\parallel}}{\sqrt{2}v_{T\parallel}}\right) \left(\frac{e\delta\phi}{2T_{\parallel}}\right)^{n}}{(2n+1)!}, \tag{8}$$

where ω/k_{\parallel} is the phase speed of EAWs, and the amplitude of electrostatic potential associated with the whistler wave is $\delta\phi=\delta E_x/k_x$. It is worth noting that the growth rate of trapped beam instabilities here is a function of $v_{ph,\parallel}/v_{T_{\parallel}}$ and $e\delta\phi/T_{\parallel}$ for a given $k_{\parallel}\lambda_D$. The zero-order

term (n=0) has no dependence on the wave amplitude, and it provides a positive growth rate due to the property $H_1\Big(v_{ph,\parallel}/\sqrt{2}v_{T\parallel}\Big)>0$ for $v_{ph,\parallel}>0$. The next higher order term (n=1) is linearly proportional to the wave amplitude $\delta\phi$. The coefficient of this term $H_3\Big(v_{ph,\parallel}/\sqrt{2}v_{T\parallel}\Big)=2\sqrt{2}(v_{ph,\parallel}/v_{T\parallel})((v_{ph,\parallel}/v_{T\parallel})^2-3)$ indicates that the wave growth rate increases with $\delta\phi$ for $v_{ph,\parallel}/v_{T\parallel}>\sqrt{3}$ and decreases with $\delta\phi$ for $v_{ph,\parallel}/v_{T\parallel}<\sqrt{3}$.

On the other hand, the Landau damping rate of EAWs by the background distribution f_{0e} can be written as

$$\begin{split} \frac{\gamma_{Le}}{k_{||}v_{T||}} &= \frac{\pi}{2} \frac{\omega}{k_{||}v_{T||}} \frac{\omega_{pe}^{2}}{k_{||}^{2}} f_{0e}^{\prime}(v_{ph,||}) \\ &= -\frac{\pi}{2} \frac{\omega}{k_{||}v_{T||}} \frac{1}{k_{||}^{2} \lambda_{D}^{2}} \frac{1}{\sqrt{2}} v_{T||} f_{0e}(v_{ph,||}) H_{1}\left(\frac{v_{ph,||}}{\sqrt{2}v_{T||}}\right). \end{split} \tag{9}$$

The overall amplification of beam instability and Landau damping on any initial perturbation wave field E_1 can be expressed as $E_1 e^{(\gamma_t,+\gamma_{Le})\Delta t}$. Taking the phase mixing effect into account, the beam distribution is smoothed within a few trapping periods. This timescale can be approximated as the inverse of phase mixing rate of trapped electrons $\Delta t \sim 1/\gamma_{\rm mixing}$, which can be estimated from the trapping frequency,

$$\frac{\gamma_{\text{mixing}}}{k_{||}v_{T||}} \sim \frac{\omega_{tr}}{k_{||}v_{T||}} = \frac{k'_{||}}{k_{||}} \cdot \frac{\omega_{tr}}{k'_{||}v_{T_{||}}} = \frac{1}{\kappa} \sqrt{\frac{e\delta\phi}{T_{||}}},\tag{10}$$

where the κ represents the wave number ratio of EAW and whistler: $k_{\parallel}/k'_{\parallel}$. Thus, the critical condition for EAW or nonlinear electrostatic structure excitation can be written as

$$\gamma_{tr} + \gamma_{Le} \ge N\gamma_{\text{mixing}},$$
 (11)

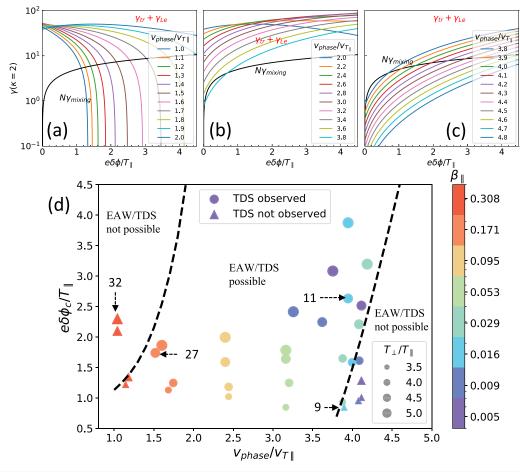


FIG. 5. The critical condition for nonlinear excitation of EAWs. (a)–(c) The comparison between the EAW growth rate and the phase mixing rate with different $v_{\rho h,\parallel}/v_{T\parallel}$. In evaluating the left-hand side of inequality (11), we use the typical parameter $k'_{\parallel}\lambda_D=0.05$ ($k'_{\parallel}\lambda_D$ being in the range 0.04–0.06 in the simulations), truncate the power series at n=4 (the relative error being $<5\times10^{-5}$), and parameterize the growth rate by $v_{\rho h,\parallel}/v_{T\parallel}$. In evaluating the phase mixing rate on the right-hand side of inequality (11), the number of e-foldings is chosen as N=5 (i.e., the signal-to-noise ratio is e^5). The result is shown as a black line. (d) The critical amplitude of whistler waves to drive EAWs or TDSs. The left boundary corresponds to the intersection of the mixing rate and the EAW/TDS growth rate in panel (a), and the right boundary corresponds to that in panel (c). The critical amplitudes are plotted for $\kappa=2$. The scattered dots show the PIC simulation results. The circles/triangles are results with/without EAW excitation. We use the average parallel electric field at saturation time in the simulations to estimate $\delta\phi=|\delta E_x|/k'$.

which means that the signal is observable after N e-foldings. Combining with Eqs. (8)–(11), we obtain the explicit critical condition as a function of $v_{ph,\parallel}/v_{T\parallel}$ and $e\delta\phi/T_{\parallel}$,

$$\frac{\sqrt{\pi}}{4} \frac{v_{ph,||}}{v_{T_{||}}} \exp\left(-\frac{v_{ph,||}}{2v_{T_{||}}}\right) \left[\frac{\pi}{2} \sum_{n=0}^{\infty} \frac{H_{2n+1}\left(\frac{v_{ph,||}}{\sqrt{2}v_{T_{||}}}\right)}{(2n+1)!} \left(\frac{e\delta\phi}{2T_{||}}\right)^{n} -H_{1}\left(\frac{v_{ph,||}}{\sqrt{2}v_{T_{||}}}\right)\right] \geq \frac{k_{||}^{2} \lambda_{D}^{2} N}{\kappa} \sqrt{\frac{e\delta\phi}{T_{||}}}$$

$$= k_{||}^{\prime 2} \lambda_{D}^{2} N \kappa \sqrt{\frac{e\delta\phi}{T_{||}}} = k_{||}^{\prime 2} d_{e}^{2} \frac{v_{T_{||}}^{2}}{c^{2}} N \kappa \sqrt{\frac{e\delta\phi}{T_{||}}}, \tag{12}$$

where the d_e is the electron inertial length.

We proceed to evaluate both sides of the inequality (11) and display them in three distinct regimes of $v_{ph,\parallel}/v_{T\parallel}$ [Figs. 5(a)–5(c)]. We choose typical $k'_{\parallel}\lambda_D=0.05$, $\kappa=2$, and N=5 in our simulation. In Fig. 5(a), for $1 \leq v_{ph,\parallel}/v_{T\parallel} \leq 1.9$, the EAW growth rate $(\gamma_{tr}+\gamma_{Le})$

drops below the phase mixing rate $(N\gamma_{\rm mixing})$ at a critical value $\delta\phi_c$ $(1.1 \le \frac{e\delta\phi_c}{T_\parallel} \le 4.5)$. Beyond this value, the excitation of EAWs is suppressed because the trapped electron beam is smoothed before EAWs grow to an observable level. In Fig. 5(b), for $2 < v_{ph,\parallel}/v_{T\parallel} < 3.6$, the EAW growth rate exceeds the phase mixing rate for any wave potential $\delta\phi$. In Fig. 5(c), for $3.8 \le v_{ph,\parallel}/v_{T\parallel} \le 4.8$, the EAW growth rate exceeds the phase mixing rate at a critical value, above which EAWs can be excited by the trapped electron beam. Thus, we plot the critical amplitude for a wide range of $v_{ph,\parallel}/v_{T\parallel}$. The boundaries given by the critical wave amplitudes divide the parameter space of $(v_{ph,\parallel},\delta\phi)$ into three regions [Fig. 5(d)]: an excitation band in the middle where the excitation of EAWs is allowed, and two stop bands at the two ends where the excitation of EAWs is prohibited.

We further compare the critical condition for EAW excitation with our PIC simulation results. As shown in Fig. 5(d), those simulations without EAW excitation are located in the stopband or near the boundaries between the stop and excitation bands, which supports our theoretical estimation. EAWs cannot be excited in cases where $\beta_\parallel>0.3$ and $\beta_\parallel<0.016,\,T_\perp/T_\parallel\leq4.$ We choose four representative cases [indicated by arrows and case numbers in Fig. 5(d)] to illustrate their dispersion diagrams in Fig. 6. Figures 6(a) and 6(c) show the two cases (no. 32 and 27) on two sides of the left boundary from the

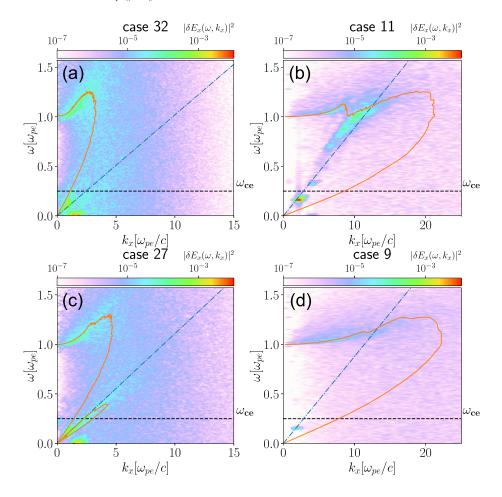


FIG. 6. The dispersion diagrams of Fourier-transformed $\delta E_{\rm X}$ as a function of frequency and wave number in four cases as marked in Fig. 5(d). The dispersion relations are calculated using the parallel electron distributions after wave saturation. (a) Case 32, (b) case 11, (c) case 27, and (d) case 9. The electron distribution functions used to calculate the dispersion relations are taken at $t=1200\omega_{pe}^{-1}$, $t=2000\omega_{pe}^{-1}$, and $t=4600\omega_{pe}^{-1}$, in the respective simulations in panels (a)–(d).

"small" β_{\parallel} regime. Although the wave amplitude $\delta\phi$ of case 32 is larger than that of case 27, EAWs are excited in the latter case, but not in the former case, consistent with the theoretical model. In comparison, Figs. 6(b) and 6(d) show the two cases (nos. 11 and 9) near the right boundary from the "large" β_{\parallel} regime. Case 11 having a wave amplitude larger than the critical amplitude can excite EAWs, whereas case 9 with a wave amplitude near the critical amplitude cannot excite EAWs. Interestingly, the "thumb" dispersion relation of EAWs/ Langmuir waves for a Maxwellian distribution⁵² is strongly bifurcated into new branches at the whistler/EAW phase velocities in the two cases with EAWs (cases 27 and 11), due to the finite plateau width in the electron distributions (created by whistler waves). The EAWs in cases 11 and 27 are mainly different in their phase velocities and range of wave numbers: case 11 with a high phase velocity is near the Langmuir branch, while case 27 with a low phase velocity is in the dispersionless acoustic branch. Such differences are governed by the ratio of whistler phase velocity to electron thermal velocity $v_{ph,\parallel}/v_{T_\parallel}$ and the finite resonant island in the electron distribution functions.

V. SUMMARY

In this study, using a series of 2D PIC simulations, we demonstrate the excitation of EAWs and nonlinear electrostatic structures through the nonlinear Landau resonant interaction between whistler waves and electrons. We further derive a critical condition for such excitation of EAWs and nonlinear electrostatic structures in the parameter space of whistler wave amplitude and phase velocity, which shows good agreement with the simulation results. The main results are as follows.

- 1. In all our PIC simulations, whistler waves are naturally generated through the temperature anisotropy instability. In the small β_{\parallel} regime of $\beta_{\parallel} \lesssim 0.025$, oblique, quasi-electrostatic whistler waves are excited, whereas in the large β_{\parallel} regime of $\beta_{\parallel} \gtrsim 0.025$, quasi-parallel, electromagnetic whistler waves are excited. In both regimes, whistler waves can have strong enough parallel electric fields to accelerate/form electron beams in their potential wells, i.e., nonlinear Landau resonance.
- 2. Electron beams trapped by whistler waves subsequently excite EAWs, which may further evolve into TDSs nonlinear electrostatic structures (time domain structures, TDS). The EAW phase velocities are approximately equal to the beam velocities, or equivalently the whistler phase velocities. We obtain the dispersion relation of EAWs using the electron distributions from PIC simulations and show that the finite plateau distribution (created by the beam) allows the survival of EAWs even when their phase velocities are close to the electron thermal velocity.
- 3. We derive the critical condition for EAW excitation by comparing the EAW growth rate and the phase mixing rate of a trapped electron beam. This critical condition is constructed in the parameter space of normalized whistler wave amplitude $e\delta\phi/T_{\parallel}$ and normalized whistler phase velocity $v_{ph,\parallel}/v_{T\parallel}$: At $v_{ph,\parallel}/v_{T\parallel} \lesssim 2$, there exists an upper bound of $e\delta\phi/T_{\parallel}$ for EAW excitation; at $v_{ph,\parallel}/v_{T\parallel} \gtrsim 3.6$, there exists a lower bound of $e\delta\phi/T_{\parallel}$ for the EAW excitation; at $2 \lesssim v_{ph,\parallel}/v_{T\parallel} \lesssim 3.6$, EAWs are unconditionally excited. These theoretical predictions are consistent with the PIC simulation results.

The modulation of high-frequency electrostatic waves (either Langmuir or electron-acoustic waves) by whistler waves has been widely observed in Earth's inner magnetosphere,²⁹ magnetotail,¹⁵ magnetopause reconnection region, 28,57 and planetary magnetospheres. 42 Our results provide a clear, quantitative explanation for these observations. Particularly, the critical condition for EAW excitation can be tested against in situ spacecraft observations. Moreover, such coupling from whistler to high-frequency electrostatic waves indicates a channel of energy transfer across different spatial scales. Taking the fast plasma injections from Earth's magnetotail to the inner magnetosphere as an example, whistler waves, which are generated by mesoscale electron injections, transfer their energy to the high-frequency, Debye-scale electrostatic waves. The coupling process involving nonlinear Landau resonance serves as a cross-scale energy channel for the dissipation of injected energy in the form of electron heating.3

ACKNOWLEDGMENTS

This work was supported by NASA Award Nos. 80NSSC20K0917, 80NSSC22K1634, and 80NSSC23K0413, NSF Award No. 2108582, and NASA Contract No. NAS5-02099. We would like to acknowledge high-performance computing support from Cheyenne (doi:10.5065/D6RX99HX) provided by NCAR's Computational and Information Systems Laboratory, sponsored by the National Science Foundation. We would also like to thank E. Paulo Alves for insightful discussions, and the OSIRIS Consortium, consisting of UCLA and IST, (Lisbon, Portugal) for the use of OSIRIS and for providing access to the OSIRIS 4.0 framework.

AUTHOR DECLARATIONS Conflict of Interest

The authors have no conflicts to disclose.

Author Contributions

Donglai Ma: Conceptualization (equal); Data curation (lead); Formal analysis (lead); Methodology (equal); Software (lead); Validation (lead); Visualization (lead); Writing – original draft (lead). Xin An: Conceptualization (equal); Funding acquisition (lead); Methodology (lead); Project administration (lead); Supervision (lead); Writing – review & editing (equal). Anton Artemyev: Supervision (supporting); Writing – review & editing (equal). Jacob Bortnik: Writing – review & editing (equal). Vassilis Angelopoulos: Supervision (supporting); Writing – review & editing (equal). Xiao-Jia Zhang: Writing – review & editing (supporting).

DATA AVAILABILITY

The data that support the findings of this study are available from the corresponding author upon reasonable request.

REFERENCES

¹O. Agapitov, A. V. Artemyev, D. Mourenas, F. Mozer, and V. Krasnoselskikh, "Nonlinear local parallel acceleration of electrons through Landau trapping by oblique whistler mode waves in the outer radiation belt," Geophys. Res. Lett. 42, 10140, https://doi.org/10.1002/2015GL066887 (2015).

- ²O. Agapitov, J. Drake, I. Vasko, F. Mozer, A. Artemyev, V. Krasnoselskikh, V. Angelopoulos, J. Wygant, and G. D. Reeves, "Nonlinear electrostatic steepening of whistler waves: The guiding factors and dynamics in inhomogeneous systems," Geophys. Res. Lett. 45, 2168–2176, https://doi.org/10.1002/2017GL076957 (2018).
- ³X. An, J. Bortnik, and X. J. Zhang, "Nonlinear Landau resonant interaction between kinetic Alfvén waves and thermal electrons: Excitation of time domain structures," J. Geophys. Res. **126**, e2020JA028643, https://doi.org/10.1029/2020JA028643 (2021).
- ⁴X. An, J. Li, J. Bortnik, V. Decyk, C. Kletzing, and G. Hospodarsky, "Unified view of nonlinear wave structures associated with whistler-mode chorus," Phys. Rev. Lett. **122**, 045101 (2019).
- ⁵X. An, C. Yue, J. Bortnik, V. Decyk, W. Li, and R. M. Thorne, "On the parameter dependence of the whistler anisotropy instability," J. Geophys. Res. 122, 2001–2009, https://doi.org/10.1002/2017JA023895 (2017).
- ⁶V. Angelopoulos, J. Chapman, F. Mozer, J. Scudder, C. Russell, K. Tsuruda, T. Mukai, T. Hughes, and K. Yumoto, "Plasma sheet electromagnetic power generation and its dissipation along auroral field lines," J. Geophys. Res. 107, SMP 14-1, https://doi.org/10.1029/2001JA900136 (2002).
- ⁷V. Angelopoulos, J. P. McFadden, D. Larson, C. W. Carlson, S. B. Mende, H. Frey, T. Phan, D. G. Sibeck, K. H. Glassmeier, U. Auster *et al.*, "Tail reconnection triggering substorm onset," Science 321, 931–935 (2008).
- ⁸A. Artemyev, O. Agapitov, F. Mozer, and V. Krasnoselskikh, "Thermal electron acceleration by localized bursts of electric field in the radiation belts," Geophys. Res. Lett. 41, 5734–5739, https://doi.org/10.1002/2014GL061248 (2014).
- ⁹D. N. Baker, T. Pulkkinen, V. Angelopoulos, W. Baumjohann, and R. McPherron, "Neutral line model of substorms: Past results and present view," J. Geophys. Res. **101**, 12975–13010, https://doi.org/10.1029/95JA03753 (1996).
- ¹⁰J. Birn, A. Artemyev, D. Baker, M. Echim, M. Hoshino, and L. Zelenyi, "Particle acceleration in the magnetotail and aurora," Space Sci. Rev. 173, 49–102 (2012).
- ¹¹J. Birn, A. Runov, and Y. Khotyaintsev, "Magnetotail processes," Magnetospheres in the Solar System (AGU, 2021), pp. 243–275.
- ¹²H. Breuillard, O. Le Contel, A. Retino, A. Chasapis, T. Chust, L. Mirioni, D. B. Graham, F. Wilder, I. Cohen, A. Vaivads et al., "Multispacecraft analysis of dipolarization fronts and associated whistler wave emissions using mms data," Geophys. Res. Lett. 43, 7279–7286, https://doi.org/10.1002/2016GL069188 (2016).
- ¹³C. Chaston, J. Bonnell, L. Clausen, and V. Angelopoulos, "Energy transport by kinetic-scale electromagnetic waves in fast plasma sheet flows," J. Geophys. Res. 117, A09202, https://doi.org/10.1029/2012JA017863 (2012).
- ¹⁴C. C. C. Chaston, J. W. Bonnell, J. R. Wygant, F. Mozer, S. D. Bale, K. Kersten, A. W. Breneman, C. A. Kletzing, W. S. Kurth, G. B. Hospodarsky, C. W. Smith, and E. A. Macdonald, "Observations of kinetic scale field line resonances," Geophys. Res. Lett. 41, 209–215, https://doi.org/10.1002/2013GL058507 (2014).
- ¹⁵Z. Chen, J. Yu, J. Wang, Z. He, N. Liu, J. Cui, and J. Cao, "High-frequency electrostatic waves modulated by whistler waves behind dipolarization front," J. Geophys. Res. 127, e2022JA030935, https://doi.org/10.1029/2022JA030935 (2022).
- 16 Computational and Information Systems Laboratory, Cheyenne: HPE/SGI ICE XA System (University Community Computing) (National Center for Atmospheric Research, Boulder, CO, 2019).
- ¹⁷C. Dillard, I. Vasko, F. Mozer, O. Agapitov, and J. Bonnell, "Electron-acoustic solitary waves in the Earth's inner magnetosphere," Phys. Plasmas 25, 022905 (2018).
- ¹⁸R. Ergun, K. Goodrich, J. Stawarz, L. Andersson, and V. Angelopoulos, "Large-amplitude electric fields associated with bursty bulk flow braking in the Earth's plasma sheet," J. Geophys. Res. 120, 1832–1844, https://doi.org/10.1002/2014JA020165 (2015).
- ¹⁹R. A. Fonseca, L. O. Silva, F. S. Tsung, V. K. Decyk, W. Lu, C. Ren, W. B. Mori, S. Deng, S. Lee, T. Katsouleas et al., "Osiris: A three-dimensional, fully relativistic particle in cell code for modeling plasma based accelerators," in *International Conference on Computational Science* (Springer, 2002), pp. 342–351.
- ²⁰X. Fu, M. M. Cowee, R. H. Friedel, H. O. Funsten, S. P. Gary, G. B. Hospodarsky, C. Kletzing, W. Kurth, B. A. Larsen, K. Liu *et al.*, "Whistler anisotropy instabilities as the source of banded chorus: Van Allen Probes

- observations and particle-in-cell simulations," J. Geophys. Res. 119, 8288–8298, https://doi.org/10.1002/2014JA020364 (2014).
- ²¹C. Gabrielse, M. Gkioulidou, S. Merkin, D. Malaspina, D. L. Turner, M. W. Chen, S. Ohtani, Y. Nishimura, J. Liu, J. Birn *et al.*, "Mesoscale phenomena and their contribution to the global response: A focus on the magnetotail transition region and magnetosphere-ionosphere coupling," Front. Astron. Space Sci. 10, 1151339 (2023).
- ²²S. P. Gary, K. Liu, and D. Winske, "Whistler anisotropy instability at low electron β: Particle-in-cell simulations," Phys. Plasmas 18, 082902 (2011).
- ²³R. Horne and R. Thorne, "Relativistic electron acceleration and precipitation during resonant interactions with whistler-mode chorus," Geophys. Res. Lett. 30, 1527, https://doi.org/10.1029/2003GL016973 (2003).
- ²⁴R. B. Horne and R. M. Thorne, "Potential waves for relativistic electron scattering and stochastic acceleration during magnetic storms," Geophys. Res. Lett. 25, 3011–3014, https://doi.org/10.1029/98GL01002 (1998).
- 25J. A. Hwang, D. Y. Lee, L. Lyons, A. Smith, S. Zou, K. Min, K. H. Kim, Y. J. Moon, and Y. Park, "Statistical significance of association between whistler-mode chorus enhancements and enhanced convection periods during high-speed streams," J. Geophys. Res. 112, A09213, https://doi.org/10.1029/2007JA012388 (2007).
- ²⁶C. Kennel, "Low-frequency whistler mode," Phys. Fluids **9**, 2190–2202 (1966).
- ²⁷O. Le Contel, A. Roux, C. Jacquey, P. Robert, M. Berthomier, T. Chust, B. Grison, V. Angelopoulos, D. Sibeck, C. Chaston *et al.*, "Quasi-parallel whistler mode waves observed by THEMIS during near-earth dipolarizations," Ann. Geophys. 27, 2259–2275 (2009).
- ²⁸J. Li, J. Bortnik, X. An, W. Li, C. T. Russell, M. Zhou, J. Berchem, C. Zhao, S. Wang, R. B. Torbert *et al.*, "Local excitation of whistler mode waves and associated Langmuir waves at dayside reconnection regions," Geophys. Res. Lett. 45, 8793–8802, https://doi.org/10.1029/2018GL078287 (2018).
- ²⁹J. Li, J. Bortnik, X. An, W. Li, R. M. Thorne, M. Zhou, W. S. Kurth, G. B. Hospodarsky, H. O. Funsten, and H. E. Spence, "Chorus wave modulation of Langmuir waves in the radiation belts," Geophys. Res. Lett. 44, 11713, https://doi.org/10.1002/2017GL075877 (2017).
- 30 W. Li, R. Thorne, V. Angelopoulos, J. Bonnell, J. McFadden, C. Carlson, O. LeContel, A. Roux, K. Glassmeier, and H. Auster, "Evaluation of whistler-mode chorus intensification on the nightside during an injection event observed on the THEMIS spacecraft," J. Geophys. Res. 114, A00C14, https://doi.org/10.1029/2008JA013554 (2009).
- ³¹J. Liu, V. Angelopoulos, X. J. Zhang, D. L. Turner, C. Gabrielse, A. Runov, J. Li, H. O. Funsten, and H. Spence, "Dipolarizing flux bundles in the cisgeosynchronous magnetosphere: Relationship between electric fields and energetic particle injections," J. Geophys. Res. 121, 1362–1376, https://doi.org/10.1002/2015JA021691 (2016).
- ³²Q. Ma, D. Mourenas, A. Artemyev, W. Li, R. M. Thorne, and J. Bortnik, "Strong enhancement of 10–100 keV electron fluxes by combined effects of chorus waves and time domain structures," Geophys. Res. Lett. 43, 4683–4690, https://doi.org/10.1002/2016GL069125 (2016).
- ³³D. M. Malaspina, A. Ukhorskiy, X. Chu, and J. Wygant, "A census of plasma waves and structures associated with an injection front in the inner magnetosphere," J. Geophys. Res. 123, 2566–2587, https://doi.org/10.1002/2017JA025005 (2018).
- ³⁴F. Mozer, O. Agapitov, A. Artemyev, J. Drake, V. Krasnoselskikh, S. Lejosne, and I. Vasko, "Time domain structures: What and where they are, what they do, and how they are made," Geophys. Res. Lett. 42, 3627–3638, https://doi.org/10.1002/2015GL063946 (2015).
- 35T. O'neil, "Collisionless damping of nonlinear plasma oscillations," Phys. Fluids 8, 2255–2262 (1965).
- ³⁶T. O'neil and J. Malmberg, "Transition of the dispersion roots from beam-type to Landau-type solutions," Phys. Fluids 11, 1754–1760 (1968).
- ³⁷A. Osmane and T. I. Pulkkinen, "On the threshold energization of radiation belt electrons by double layers," J. Geophys. Res. 119, 8243–8248, https://doi. org/10.1002/2014JA020236 (2014).
- 38A. Osmane, D. L. Turner, L. B. Wilson, A. P. Dimmock, and T. I. Pulkkinen, "Subcritical growth of electron phase-space holes in planetary radiation belts," Astrophys. J. 846, 83 (2017).
- ³⁹P. Palmadesso, "Resonance, particle trapping, and Landau damping in finite amplitude obliquely propagating waves," Phys. Fluids 15, 2006–2013 (1972).

- ⁴⁰G. Reeves, H. E. Spence, M. Henderson, S. Morley, R. Friedel, H. Funsten, D. Baker, S. Kanekal, J. Blake, J. Fennell *et al.*, "Electron acceleration in the heart of the Van Allen radiation belts," Science 341, 991–994 (2013).
- ⁴¹L. A. Reinleitner, D. A. Gurnett, and D. L. Gallagher, "Chorus-related electro-static bursts in the Earth's outer magnetosphere," Nature 295, 46–48 (1982).
- ⁴²L. A. Reinleitner, W. Kurth, and D. A. Gurnett, "Chorus-related electrostatic bursts at Jupiter and Saturn," J. Geophys. Res. 89, 75–83, https://doi.org/ 10.1029/JA089iA01p00075 (1984).
- ⁴³A. Runov, V. Angelopoulos, M. Sitnov, V. Sergeev, J. Bonnell, J. McFadden, D. Larson, K. H. Glassmeier, and U. Auster, "Themis observations of an earthward-propagating dipolarization front," Geophys. Res. Lett. 36, L14106, https://doi.org/10.1029/2009GL038980 (2009).
- 44 T. Sano, M. Hata, D. Kawahito, K. Mima, and Y. Sentoku, "Ultrafast wave-particle energy transfer in the collapse of standing whistler waves," Phys. Rev. E 100, 053205 (2019).
- 45 Y. Shen, A. Artemyev, X. J. Zhang, I. Y. Vasko, A. Runov, V. Angelopoulos, and D. Knudsen, "Potential evidence of low-energy electron scattering and ion-ospheric precipitation by time domain structures," Geophys. Res. Lett. 47, e2020GL089138, https://doi.org/10.1029/2020GL089138 (2020).
- ⁴⁶Y. Y. Shprits, D. A. Subbotin, N. P. Meredith, and S. R. Elkington, "Review of modeling of losses and sources of relativistic electrons in the outer radiation belt II: Local acceleration and loss," J. Atmos. Sol.-Terr. Phys. 70, 1694–1713 (2008).
- ⁴⁷J. Stawarz, R. Ergun, and K. Goodrich, "Generation of high-frequency electric field activity by turbulence in the Earth's magnetotail," J. Geophys. Res. 120, 1845–1866, https://doi.org/10.1002/2014JA020166 (2015).
- ⁴⁸D. Summers, R. M. Thorne, and F. Xiao, "Relativistic theory of wave-particle resonant diffusion with application to electron acceleration in the magnetosphere," J. Geophys. Res. 103, 20487–20500, https://doi.org/10.1029/98JA01740 (1998)
- 49X. Tao, R. Thorne, W. Li, B. Ni, N. Meredith, and R. Horne, "Evolution of electron pitch angle distributions following injection from the plasma sheet," J. Geophys. Res. 116, A04229, https://doi.org/10.1029/2010JA016245 (2011).
- 50 U. Taubenschuss, O. Santolík, D. B. Graham, H. Fu, Y. V. Khotyaintsev, and O. Le Contel, "Different types of whistler mode chorus in the equatorial source region," Geophys. Res. Lett. 42, 8271–8279, https://doi.org/10.1002/2015GL066004 (2015).

- ⁵¹R. Thorne, W. Li, B. Ni, Q. Ma, J. Bortnik, L. Chen, D. Baker, H. E. Spence, G. Reeves, M. Henderson *et al.*, "Rapid local acceleration of relativistic radiation-belt electrons by magnetospheric chorus," Nature 504, 411–414 (2013).
- ⁵²F. Valentini, T. M. O'Neil, and D. H. Dubin, "Excitation of nonlinear electron acoustic waves," Phys. Plasmas 13, 052303 (2006).
- ⁵³I. Vasko, O. Agapitov, F. Mozer, A. Artemyev, V. Krasnoselskikh, and J. Bonnell, "Diffusive scattering of electrons by electron holes around injection fronts," J. Geophys. Res. 122, 3163–3182, https://doi.org/10.1002/2016JA023337 (2017).
- 54I. Vasko, O. Agapitov, F. Mozer, J. Bonnell, A. Artemyev, V. Krasnoselskikh, G. Reeves, and G. Hospodarsky, "Electron-acoustic solitons and double layers in the inner magnetosphere," Geophys. Res. Lett. 44, 4575–4583, https://doi.org/10.1002/2017GL074026 (2017b).
- 55I. Vasko, O. V. Agapitov, F. Mozer, and A. Artemyev, "Thermal electron acceleration by electric field spikes in the outer radiation belt: Generation of field-aligned pitch angle distributions," J. Geophys. Res. 120, 8616–8632, https://doi.org/10.1002/2015JA021644 (2015).
- ⁵⁶I. Y. Vasko, O. V. Agapitov, F. S. Mozer, J. W. Bonnell, A. V. Artemyev, V. V. Krasnoselskikh, and Y. Tong, "Electrostatic steepening of whistler waves," Phys. Rev. Lett. **120**, 195101 (2018).
- 57S. Wang, D. B. Graham, X. An, L. Li, Q. G. Zong, X. Z. Zhou, W. Li, and Z. Liu, "Electrostatic waves around a magnetopause reconnection diffusion region and their associations with whistler and lower-hybrid waves," J. Geophys. Res. 50(18), e2023GL104905, https://doi.org/10.1029/2023GL104905 (2023).
- 58H. J. Weber and G. B. Arfken, Essential Mathematical Methods for Physicists (ISE-Elsevier, 2003).
- 59C. Yue, X. An, J. Bortnik, Q. Ma, W. Li, R. M. Thorne, G. D. Reeves, M. Gkioulidou, D. G. Mitchell, and C. A. Kletzing, "The relationship between the macroscopic state of electrons and the properties of chorus waves observed by the Van Allen Probes," Geophys. Res. Lett. 43, 7804–7812, https://doi.org/10.1002/2016GL070084 (2016).
- ⁶⁰X. Zhang, V. Angelopoulos, A. Artemyev, and J. Liu, "Energy transport by whistler waves around dipolarizing flux bundles," Geophys. Res. Lett. 46, 11718–11727, https://doi.org/10.1029/2019GL084226 (2019).
- ⁶¹X. J. Zhang, A. Artemyev, V. Angelopoulos, E. Tsai, C. Wilkins, S. Kasahara, D. Mourenas, S. Yokota, K. Keika, T. Hori *et al.*, "Superfast precipitation of energetic electrons in the radiation belts of the earth," Nat. Commun. 13, 1611 (2022).