

Reliable chaotic transition in incompressible fluid simulations

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Abstract

We consider a test problem for Navier–Stokes solvers based on the flow around a cylinder that exhibits chaotic behavior, to examine the performance of various numerical methods. We choose a range of Reynolds numbers for which the flow is time-dependent but can be characterized as essentially two-dimensional. The problem requires accurate resolution of chaotic dynamics over a long time interval. It also requires the use of a relatively large computational domain, part of which is curved. We review the performance of different finite element methods for the proposed range of Reynolds numbers. These tests indicate that some of the most established methods do not capture the correct behavior. The key requirements identified are pressure-robustness of the method, high resolution, and appropriate numerical dissipation when the smallest scales are under-resolved.

Keywords: Drag coefficient, Chaotic flow, Strouhal period, $H(\text{div})$ -conforming, numerical dissipation

MSC Codes: 35Q30, 65M60

1 Introduction

We consider a classic fluids problem, the chaotic flow around a cylinder, as a challenge for numerical schemes. As reviewed in [45], there is ample experimental data with which to compare simulations. To make the challenge more tractable, we focus on 2D simulations and a range of Reynolds numbers. Illustrating the difficulty posed by this problem, we review some published examples where erroneous results were obtained. We also show that some commonly used techniques fail to give acceptable results at Reynolds numbers of interest.

Flow around a cylinder is not only a fundamental problem in fluid dynamics but also a practical problem of interest in energy generation. For example, so-called bladeless turbines [3, 48, 15] have been proposed as a viable energy generation method. The design of such systems requires accurate and efficient simulations of the flow around cylinders. Another area of recent development in which such simulations are necessary is new design spaces for electric aircrafts, such as ground-effect transportation [7] and air taxis [36].

As computational simulation has emerged as a new form of experimentation in physics, it is essential to have some benchmarks that give guidance, as we present here. For many flow problems, there is no dispute regarding the Navier–Stokes equations as a suitable model, as we do here. However, this requires reliable methods to decide fundamental physics questions, such as whether the resulting flow is chaotic at a given Reynolds number.

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A critical factor in obtaining reliable results at high Reynolds numbers and avoiding erroneous chaos, is for a numerical method to have an appropriate mechanism to deal with the energy from under resolved scales [8]. This usually takes the form of numerical dissipation and can come from different stabilization terms in the discrete method. However, it is also critical for the method not to be overly dissipative, which would again yield erroneous results. In this work, we cover a number of schemes with different dissipative mechanisms providing significant data clarifying this.

Flow around a cylinder has been proposed as a test problem before [40, 25]. What is different here is that we focus on a range of Reynolds numbers for which the flow appears to be chaotic [45]. Thus, appropriate metrics must be used to compare different numerical methods. In particular, we focus on statistics predicting the periodicity of the flow, i.e. the Strouhal period [49], rather than just minimal and maximal drag and lift coefficient values. Moreover, the time interval for the simulation is quite long, so this simulation evaluates the ability of different discretizations with regard to accumulation of numerical error and challenges the implementations with regard to efficiency.

For a straightforward analysis of the resulting flow, we focus on simple metrics so that different methods can be compared easily. The Strouhal number or period of the time-dependent flow is such a quantity. This can be computed by analyzing the phase-diagram of the drag and lift coefficients, which are computed easily. It is vital to compute the period by considering the drag and lift simultaneously, as both have local maxima and minima making it difficult to determine the period in cases where the flow is not perfectly periodic. In [45], a simple approach was used to estimate the periodicity of chaotic flow. We present a slightly improved method compared to [45] for evaluating the Strouhal period that provides an indication of the transition to chaos. Additional metrics, such as the Lyapunov exponent and the fractal dimension of the attractor, provide a finer level of detail regarding the flow dynamics [45].

Evaluating computational methods for fluid flows is already viable in two spatial dimensions and sufficiently challenging to differentiate between them. We, therefore, focus on two dimensional flows. When the Reynolds number is more than 10^4 , the flow becomes three-dimensional [23]. Thus, we restrict ourselves to Reynolds numbers in the range $[10^2, 10^4]$ for simplicity. Three-dimensional effects of a different kind are seen [51] at lower Reynolds numbers due to the finite length of cylinders used in experiments. However, these are potentially like so-called blocking effects in two-dimensional flow [39], which become less significant at higher Reynolds numbers.

In this paper, we consider a variety of well-studied finite element methods with a solid theoretical foundation. These are low and high-order Taylor–Hood, with and without grad-div stabilization, Scott–Vogelius finite elements and a mass-conserving mixed stress (MSC) based formulation using $H(\text{div})$ -conforming velocities. The latter we consider both with an upwind and central flux in the convective term. It is rare for different numerical methods to be compared head-to-head on a challenging physical problem of interest. We hope that this paper will emphasize the value of doing so. Although we have studied several commonly used methods, and we compare results with other published simulation methods, there are many more methods that have been proposed for simulations of the type we do here. We hope our benchmark problem will attract others to determine to what extent their methods compare with the results presented here. Finally, we compare their performance to a method used in [45], and we find that the simulations in [45] were under-resolved for higher Reynolds numbers.

Our main conclusion is that high-order methods are necessary to obtain reliable results at higher Reynolds numbers. Surprisingly, the commonly used, lowest-order Taylor–Hood method is not accurate enough to be used even at the lowest Reynolds numbers considered here. A secondary conclusion is that it appears that methods that result in exactly divergence-free velocity solutions have a significant advantage at higher Reynolds numbers, even though one that we investigate is non-conforming. Furthermore, we also conclude that a viable method at high Reynolds numbers needs to contain the correct amount of numerical dissipation to capture the flow dynamics on coarser meshes. Finally, one of the main conclusions from [45] was that the simulation drag values differ substantially from experiments, as shown in Figure 1. We confirm this conclusion here, suggesting that cylinder vibration likely plays a significant role in practice [52].

The remainder of this paper is structured as follows. In Section 2, we discuss the equations and

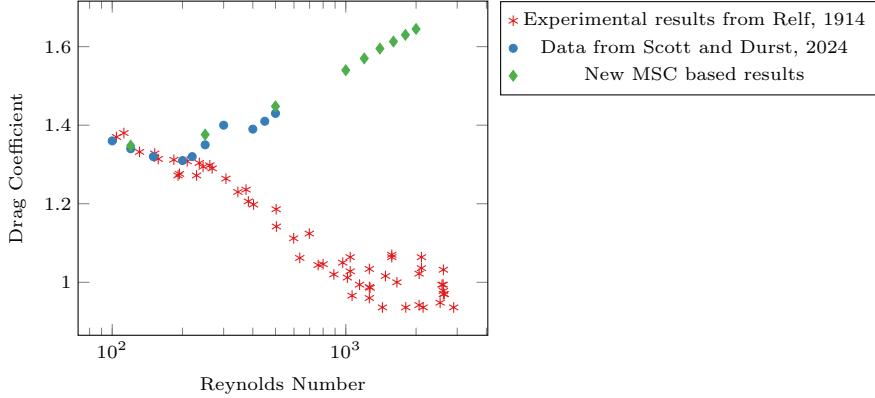


Figure 1: Relf experimental data [37] compared to the data computed in [45] with the Scot–Vogelius finite elements scheme and the new data presented here.

the quantities under consideration. In Section 3, we then present the details of the computational setup for our computations. Section 4 covers the details of the numerical methods under consideration. In Section 5, we present and discuss the numerical results achieved. In particular, we start at a low Reynolds number to illustrate how established methods struggle to achieve the correct dynamics. We then cover a range of Reynolds numbers up to $Re = 8000$. Finally, we give some concluding remarks in Section 6.

2 Setting the problem and model equations

Suppose that (\mathbf{u}, p) is a solution of the time-dependent Navier–Stokes equations in a domain $\Omega \subset \mathbb{R}^d$ containing an obstacle with boundary $\Gamma \subset \partial\Omega$. This fulfills the equations

$$\begin{aligned} \partial_t \mathbf{u} - \nu \Delta \mathbf{u} + \mathbf{u} \cdot \nabla \mathbf{u} + \nabla p &= \mathbf{0} & \text{in } \Omega, \\ \nabla \cdot \mathbf{u} &= 0 & \text{in } \Omega, \end{aligned} \tag{1}$$

with the kinematic viscosity ν , and together with boundary conditions

$$\mathbf{u} = \mathbf{g} \text{ on } \partial\Omega \setminus \Gamma, \quad \mathbf{u} = \mathbf{0} \text{ on } \Gamma. \tag{2}$$

For the well-posedness of these equations, see [17].

2.1 Weak formulation of the Navier–Stokes equations

While this test problem is not limited to finite-element simulations, we expect many will be done with finite elements. To this end, we require a variational formulation of (1). The Navier–Stokes equations can be written in a weak (or variational) form as follows:

Find $\mathbf{v} \in H_g^1(\Omega)$ and $p \in L_0^2(\Omega)$ such that

$$\begin{aligned} (\partial_t \mathbf{u}, \mathbf{w})_{L^2(\Omega)} + a(\mathbf{u}, \mathbf{w}) + b(\mathbf{w}, p) + c(\mathbf{u}, \mathbf{u}, \mathbf{w}) &= 0 & \forall \mathbf{w} \in (H_0^1(\Omega))^d, \\ b(\mathbf{u}, q) &= 0 & \forall q \in L_0^2(\Omega). \end{aligned} \tag{3}$$

The space $H_g^1(\Omega)$ is the space of vector-valued H^1 functions with trace g on the (outer) boundary, $H_0^1(\Omega)$ is the space of H^1 functions with trace 0 on the boundary and $L_0^2(\Omega)$ is the space of L^2 functions with mean zero. The bilinear forms $a(\cdot, \cdot)$ and $b(\cdot, \cdot)$ are defined as

$$a(\mathbf{v}, \mathbf{w}) := \int_{\Omega} \frac{\nu}{2} \mathcal{D}(\mathbf{v}) : \mathcal{D}(\mathbf{w}) \, d\mathbf{x}, \quad \text{and} \quad b(\mathbf{v}, q) := - \int_{\Omega} q (\nabla \cdot \mathbf{v}) \, d\mathbf{x},$$

respectively, where $\mathcal{D}(\mathbf{v}) = \nabla\mathbf{v} + \nabla\mathbf{v}^t$ and the colon ($:$) indicates the Frobenius inner-product of matrices. The convective (nonlinear) term takes various forms. By default, we take it to be

$$c(\mathbf{u}, \mathbf{v}, \mathbf{w}) = c_{conv}(\mathbf{u}, \mathbf{v}, \mathbf{w}) := \int_{\Omega} (\mathbf{u} \cdot \nabla \mathbf{v}) \cdot \mathbf{w} \, d\mathbf{x}.$$

For divergence-free H^1 functions, the various forms of the convective term are equivalent. Hence, we may take the above so-called convective form of the trilinear form for exactly divergence-free methods. However, for non-divergence-free methods, other forms are beneficial and for non- H^1 -conforming methods are necessary.

2.2 Drag and lift evaluation

The flow of fluid around an obstacle generates a force called drag, which is a fundamental concept in fluid dynamics [28]. It plays a critical role in determining the behavior of objects in flight and has been studied since the time of d'Alembert [18]. Drag is composed of two components: pressure drag β_p and viscous drag β_v , which can be calculated by evaluating the following functions using $\mathbf{v} = (1, 0)$:

$$\beta_v(\mathbf{v}) = \int_{\Gamma} ((\nu \mathcal{D}(\mathbf{u})) \mathbf{v}) \cdot \mathbf{n} \, ds, \quad \beta_p(\mathbf{v}) = \int_{\Gamma} -p \mathbf{v} \cdot \mathbf{n} \, ds.$$

The full drag β is defined by $\beta = \beta_v + \beta_p$. Similarly, lift is computed using $\mathbf{v} = (0, 1)$.

2.2.1 Alternate drag evaluation

Another way [25] to evaluate β is to test the weak formulation (3) with a non-conforming test function. This results in the functional

$$\omega(\mathbf{v}) = \int_{\Omega} \partial_t \mathbf{u} \cdot \mathbf{v} \, d\mathbf{x} + \int_{\Omega} \frac{\nu}{2} \mathcal{D}(\mathbf{u}) : \mathcal{D}(\mathbf{v}) + (\mathbf{u} \cdot \nabla \mathbf{u}) \cdot \mathbf{v} - p \nabla \cdot \mathbf{v} \, d\mathbf{x},$$

Then $\omega(\mathbf{v}) = \beta(\mathbf{v})$ for all $\mathbf{v} \in H^1(\Omega)^d$ [25, 19]. This approach may be traced back to [2]. It is also known that, for the case of strongly imposed Dirichlet boundary conditions, this approach of testing the residual with a non-conforming test-function doubles the rate of convergence for the drag and lift, see, for example [6] for the proof in the steady case.

Having two ways to compute drag and lift provides a valuable internal check of the accuracy of the simulation.

3 Problem Set-up

For our simulations, we consider the domain

$$\Omega = \{(x, y) : -30 < x < 300, |y| < 30, x^2 + y^2 > 1\}.$$

The boundary condition is set as $\mathbf{g} = (1, 0)^T$ and the cylinder diameter is the reference length, so the Reynolds number is given by $Re = 2/\nu$. We consider the time interval $[0, 500]$.

3.1 Challenge computation

The data in Figure 1 suggests a significant disagreement between computation and experiment. This could be for a variety of reasons. The most obvious is some flaw in the computational scheme. Thus, we propose this problem as a computational challenge to see what other methods predict.

Another reason for the disagreement could be that the model is wrong. Some people who have published experiments (see [45]) have raised the issue that the cylinders used could vibrate. In particular, the Relf data in Figure 1 is based on flows past thin wires, used in musical instruments for their

vibrational qualities. Thus, one response to the challenge is to allow the cylinder to vibrate back and forth in response to the oscillating lift and drag. The question would be to evaluate how oscillations of the cylinder change parameters describing the flow, such as drag, lift, and Strouhal number, representing the frequency of the generated vortices. One study with forced vibrations [33, Figure 6.9, page 179] suggests that the change could be substantial.

4 Discretization methods

The simulations in this paper are done using finite element methods (FEM). We, therefore, begin by presenting some details of the finite element approach.

To discretize the weak formulation (3) with finite element methods, we take a mesh of the domain Ω with characteristic length h , denoted as \mathcal{T}_h . Finite element spaces then discretize (3) by approximating the spaces $H^1(\Omega)$ and $L^2(\Omega)$ with spaces of piecewise polynomials on each element of the mesh, with varying degree of discontinuity of these polynomials over element edges.

We utilize many well-established simulation methods for which extensive numerical analysis is available, establishing not only convergence but also detailed simulation properties.

4.1 $H(\text{div})$ -conforming FEM

We consider an $H(\text{div})$ -conforming finite element method based on [20, 21]. This discretization is based on a mass-conserving mixed formulation with symmetric stresses, known as the MCS formulation. This approach discretizes the velocity in $H(\text{div})$ using Brezzi–Douglas–Marini (BDM) elements [47]. That is, the normal component of the velocity is continuous across element boundaries, while the tangential component is discontinuous. The pressure is discretized using discontinuous elements. This method is exactly mass conserving and pressure-robust. From a numerical point of view, pressure-robust methods have several advantages. First, the velocity remains unchanged by irrotational changes in any external forcing term, a property that most methods do not preserve on the discrete level [31]. More importantly, pressure-robust methods have the property that the velocity error is independent of the pressure error [26]. This contrasts classical Galerkin discretizations of the (Navier)–Stokes equations, where the velocity error depends on $\frac{1}{\nu}$ times the pressure error. From a physical point of view, exactly divergence-free methods have the advantage that, in addition to mass, kinetic energy, linear momentum, and angular momentum are conserved on the discrete level [10, 44].

This formulation introduces an additional variable $\sigma = \nu \nabla \mathbf{u}$ which is discretised in the function space $H(\text{curl div})$. Thus the energy form is modified to

$$a'(\sigma, \tau) := \frac{1}{\nu} \int_{\Omega} \sigma : \tau \, d\mathbf{x},$$

and a new form is defined by

$$b'(\tau, \mathbf{u}) := \sum_{T \in \mathcal{T}_h} \int_T \nabla \cdot (\tau) \cdot \mathbf{u} \, d\mathbf{x} + \sum_{F \in \mathcal{F}_h} \int_F [\![\tau_{nn}]\!] \mathbf{u} \cdot \mathbf{n} \, ds,$$

where \mathcal{T}_h is the triangular mesh, and \mathcal{F}_h is the set of element edges. The expression $[\![\tau_{nn}]\!]$ denotes the jump in the normal-normal component of the tensor τ across the edge e , and \mathbf{n} is the normal to the edge. Note that the orientation of the normal is not material. As the velocity is $H(\text{div})$, we have to modify the convective term to account for the discontinuity of the velocity in the tangential direction, and there are multiple choices available. For the most part, we consider the upwind flux for the convective term,

$$c_{upw}(\mathbf{u}_h, \mathbf{u}_h, \mathbf{w}_h) := \sum_{T \in \mathcal{T}_h} \int_T -(\mathbf{u}_h \cdot \nabla \mathbf{w}_h) \cdot \mathbf{u}_h \, d\mathbf{x} + \int_{\partial T} \mathbf{u}_h \cdot \mathbf{n} \hat{\mathbf{v}}_h \cdot \mathbf{w}_h \, ds, \quad (4)$$

where $\hat{\mathbf{v}}_h$ is the upwind flux, i.e., $\hat{\mathbf{v}}_h = \mathbf{u}_h$ if $\mathbf{u}_h(x) \cdot \mathbf{n}(x) \geq 0$ and otherwise $\hat{\mathbf{v}}_h = \mathbf{u}_h^{\text{oth}}$, where $\mathbf{u}_h^{\text{oth}}$ is the value of \mathbf{u}_h from the neighboring element. An alternative choice is a central flux. In this case the

convective term can be written as

$$c_{cf}(\mathbf{u}_h, \mathbf{u}_h, \mathbf{w}_h) := \sum_{T \in \mathcal{T}_h} \int_T (\mathbf{u}_h \cdot \nabla \mathbf{u}_h) \cdot \mathbf{w}_h \, d\mathbf{x} - \int_{\partial T} \mathbf{u}_h \cdot \mathbf{n} (\hat{\mathbf{u}}_h - \mathbf{u}_h) \cdot \mathbf{w}_h \, ds, \quad (5)$$

where $\hat{\mathbf{u}}_h = 0.5(\mathbf{u}_h - \mathbf{u}_h^{\text{oth}})$ is the centered flux. In general we shall use the upwind flux, and will therefore only specify if the centered flux is used. We note that the upwind choice is dissipative, whereas a central flux would not add any dissipation, i.e., is conservative.

Combining the above, the formulation is

$$\begin{aligned} a'(\sigma_h, \tau_h) + b'(\tau_h, \mathbf{u}_h) &= 0 \quad \forall \tau_h \in \Sigma_h, \\ (\partial_t \mathbf{u}_h, \mathbf{v}_h)_{L^2(\Omega)} + b'(\sigma_h, \mathbf{v}_h) + c_{upw}(\mathbf{u}_h, \mathbf{u}_h, \mathbf{v}_h) &+ b(\mathbf{v}_h, p_h) = 0 \quad \forall \mathbf{v}_h \in V_h, \\ b(\mathbf{u}_h, q_h) &= 0 \quad \forall q_h \in Q_h. \end{aligned}$$

Here, Σ_h is a space of discontinuous piecewise polynomials of degree k , whose values are trace-zero tensors satisfying continuity of the ‘normal-tangential component’ across edges and where the normal-tangential component is a polynomial of order $k-1$ on each edge, V_h denotes the BDMk finite-element space of order k , and Q_h consists of discontinuous piecewise polynomials of degree $k-1$.

As $H(\text{div})$ spaces are characterized by normal continuity across facets, the boundary condition for the normal components are imposed strongly. The implementation uses hybridization for the tangential component of the velocity. That is, a second finite element space is used to couple the tangential component across element facets and the tangential part of the boundary condition is enforced strongly on the facet space.

For time-stepping we use a second-order implicit-explicit (IMEX) scheme known as SBDF2 [1]. This scheme uses the BDF2 scheme to discretise the time-derivative, the Stokes term is fully implicit, and the non-linear convective term is treated explicitly with a second order extrapolation. This results in the fully discrete problem to find $(\sigma_h, \mathbf{u}_h, p_h) \in \Sigma_h \times V_h \times Q_h$, such that

$$\begin{aligned} a'(\sigma_h^n, \tau_h) + b'(\tau_h, \mathbf{u}_h^n) + b'(\sigma_h^n, \mathbf{v}_h) + \frac{1}{\Delta t} (\frac{3}{2} \mathbf{u}_h^n - 2 \mathbf{u}_h^{n-1} + \frac{1}{2} \mathbf{u}_h^{n-2}, \mathbf{v}_h)_{L^2(\Omega)} \\ + 2c_{upw}(\mathbf{u}_h^{n-1}, \mathbf{u}_h^{n-1}, \mathbf{v}_h) - c_{upw}(\mathbf{u}_h^{n-2}, \mathbf{u}_h^{n-2}, \mathbf{v}_h) + b(\mathbf{v}_h, p_h^n) + b(\mathbf{u}_h^n, q_h) = 0, \end{aligned}$$

holds for all $(\tau_h, \mathbf{v}_h, q_h) \in \Sigma_h \times V_h \times Q_h$. To regularize the saddle-point system and to fix the pressure constant in the case of pure Dirichlet boundary conditions, we may add a small perturbation to the lower right pressure block of the block-system resulting from the above equation. That is, we modify the equation for the divergence constraint to

$$-\int_{\Omega} q_h (\nabla \cdot \mathbf{u}_h^n) + \epsilon \int_{\Omega} p_h^n q_v \, d\mathbf{x} = -\int_{\Omega} q_h ((\nabla \cdot \mathbf{u}_h^n) - \epsilon p_h^n) \, d\mathbf{x} = 0, \quad (6)$$

with some $\epsilon > 0$. Since our discrete spaces V_h, Q_h satisfy the discrete de-Rahm complex [26], we have the property that $\nabla \cdot V_h = Q_h$. Consequently, we can replace q_h with $\nabla \cdot \mathbf{v}_h$ in (6). This gives us an equation for the pressure, which we may plug into the second equation in the above saddle-point system. This leads to the discrete system,

$$a'(\sigma_h^n, \tau_h) + b'(\tau_h, \mathbf{u}_h^n) + \frac{1}{\Delta t} (\frac{3}{2} \mathbf{u}_h^n, \mathbf{v}_h)_{L^2(\Omega)} + b'(\sigma_h^n, \mathbf{v}_h) - b(\mathbf{v}_h, \frac{1}{\epsilon} \nabla \cdot \mathbf{u}_h). \quad (7)$$

This is a smaller linear system to solve, and the system is no longer a saddle point problem, making it an easier system to solve numerically. We want to choose ϵ small to keep the perturbation from the divergence constraint small. However, this means that the $\frac{1}{\epsilon}$ -scaled term causes the linear system (7) to be very ill-conditioned. Consequently, we need to use direct linear solvers to handle these ill-conditioned linear systems, which in turn limits us with respect to the maximal size of the system we can solve and the parallel scalability. With the small perturbation, the divergence of \mathbf{u}_h is no longer exactly zero, but $\|\nabla \cdot \mathbf{u}_h^n\|_{L^2} \leq \epsilon \|p_h^n\|_{L^2}$. In the computations, we chose $\epsilon = 10^{-12}/\nu$.

We will refer to the approach above as MCS_k , where k is the order of the velocity finite element space. When we use the central flux in the convective term, we refer to the method as MCS_k^{cf} . These approaches are implemented using Netgen/NGSolve [42, 41] and based on a template available on github*.

Remark 4.1 (Other $H(\text{div})$ -conforming methods). We note there are other $H(\text{div})$ -conforming mass conserving finite element methods that do not reformulate the problem into a first-order system. These include [11, 43] and hybridized approaches, e.g., [30, 29].

4.2 Taylor–Hood

To compare with more established methods, we also consider Taylor–Hood elements consisting of H^1 -conforming finite elements of order $k \geq 2$ for the velocity and order $k - 1$ for the pressure.

To stabilize this scheme, we optionally add grad-div stabilization to improve the conservation of mass [34, 16]. That is, we add the term

$$j_h(\mathbf{u}_h, \mathbf{v}_h) := \gamma_{gd} \int_{\Omega} \nabla \cdot \mathbf{u}_h \nabla \cdot \mathbf{v}_h \, d\mathbf{x}, \quad (8)$$

to our bilinear form. In our computations, we take $\gamma_{gd} = 10^3$, and we refer to this method as gdTH_k , where k is the order of the velocity space. In case no grad-div stabilization is used, we refer to the method as TH_k . Furthermore, we note that it is known that for $\gamma_{gd} \rightarrow \infty$, the grad-div stabilized Taylor–Hood solution converges to the Scott–Vogelius solution [9, 32].

For the convective term, there are multiple choices possible. We take the divergence form [24]

$$c_{div}(\mathbf{u}, \mathbf{v}, \mathbf{w}) = c_{conv}(\mathbf{u}, \mathbf{v}, \mathbf{w}) + \frac{1}{2} \int_{\Omega} (\nabla \cdot \mathbf{u}) \mathbf{v} \cdot \mathbf{w} \, d\mathbf{x}, \quad (9)$$

which is equivalent to the standard convective form on the continuous level, but vanishes under less restrictive conditions for $\mathbf{v} = \mathbf{w}$ on the discrete level.

Thus the variational formulation takes the form

$$\begin{aligned} & (\partial_t \mathbf{u}_h, \mathbf{w}_h)_{L^2(\Omega)} + a(\mathbf{u}_h, \mathbf{w}_h) + b(\mathbf{w}_h, p_h) + j_h(\mathbf{u}_h, \mathbf{w}_h) \\ & \quad + c_{div}(\mathbf{u}_h, \mathbf{u}_h, \mathbf{w}_h) = 0 \quad \forall \mathbf{w}_h \in V_h, \\ & \quad b(\mathbf{u}_h, q_h) = 0 \quad \forall q_h \in Q_h, \end{aligned} \quad (10)$$

where V_h is the standard Lagrange space of continuous, piecewise polynomials of degree $k \geq 2$, and Q_h is the standard Lagrange space of continuous, piecewise polynomials of degree $k - 1$.

To discretize in time we use the fully implicit Crank–Nicolson scheme, that is we solve for $(\mathbf{u}_h, p_h) \in V_h \times Q_h$, such that

$$\begin{aligned} & \frac{1}{\Delta t} (\mathbf{u}_h^n - \mathbf{u}_h^{n-1}, \mathbf{w}_h)_{L^2(\Omega)} + \frac{1}{2} (a(\mathbf{u}_h^n, \mathbf{w}_h) + c_{div}(\mathbf{u}_h^n, \mathbf{u}_h^n, \mathbf{w}_h)) + b(\mathbf{w}_h, p_h^n) \\ & \quad + j_h(\mathbf{u}_h^n, \mathbf{w}_h) + \frac{1}{2} (a(\mathbf{u}_h^{n-1}, \mathbf{w}_h) + c_{div}(\mathbf{u}_h^{n-1}, \mathbf{u}_h^{n-1}, \mathbf{w}_h)) + b(\mathbf{u}_h^n, q_h) = 0, \end{aligned} \quad (11)$$

for all $(\mathbf{v}_h, q_h) \in V_h \times Q_h$. Note that we do not take into account the pressure from the previous time step, nor do we add the grad-div stabilization term from the last time step. This is because the method needs to correctly reflect that the pressure is the Lagrange–multiplier to weakly enforce the divergence constraint [38, Section 4.1.4]. Similarly, the role of the grad-div form is to penalize deviations from the divergence constraint.

We use a quasi-Newton method with line search to solve the nonlinear systems arising in every time step. More precisely, we only recompute the factorization of the inverse Jacobian if the residual decreases by less than a factor of 0.3. The line search consists of a check that the Newton update has decreased the residual. Otherwise, we reduce the size of the update by half recursively until the residual has decreased compared to the previous iteration. We start each quasi-Newton iteration with (u_h, p_h)

*<https://github.com/NGSolve/modeltemplates>

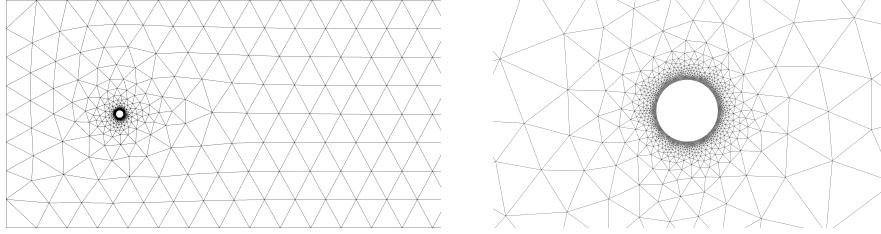


Figure 2: Left: Left section of mesh for the TH, SV, and MCS methods with $h = 8$. Right: Zoom in to the cylinder of the same mesh.

from the previous time step. As the scheme is fully implicit, we can use time steps independent of the mesh size, and as the Jacobian rarely has to be refactorized, the method is essentially as efficient as IMEX schemes.

The Taylor–Hood methods are also implemented using Netgen/NGSolve. The meshes used are constructed as those for the $H(\text{div})$ -conforming method in Section 4.1.

4.3 Scott–Vogelius

As a final method, we also consider the Scott–Vogelius finite element pair. This finite element pair is very similar to the Taylor–Hood pair, using H^1 -conforming finite elements for the velocity but discontinuous elements of order $k - 1$ for the pressure. Consequently, we use the same variational formulation time-stepping scheme used for the unstabilized Taylor–Hood method. For $k \geq 4$ in two dimensions, this method is inf-sup stable on meshes without nearly singular vertices [50, 46] and for $k \geq d = 2$ on barycentric refined meshes [35, 53]. This method is also exactly divergence-free and pressure-robust.

There are different ways that the Scott–Vogelius method [47] can be implemented. In [45], the iterated penalty method was used to avoid explicit knowledge of the pressure space. However, this approach was plagued by nearly singular vertices [19] in the meshes used. Here, we utilize the Netgen mesh generator, which, for the cylinder problem, does not have any singular or nearly singular vertices. Thus, the mixed-method implementation of Scott–Vogelius can be used, and in this case, there is only one change from the (unstabilized) Taylor–Hood method, namely, the pressure elements are discontinuous instead of continuous. One drawback of this approach is the linear systems are larger, and it appears that for Reynolds number 8000 (Table 7), it is no longer possible to keep the divergence very small. Upon closer examination, it was found that the Newton step was not converging at a few time steps, and the divergence also spiked there. In general, the divergence was very small except at these few times. Furthermore, it is worth mentioning that, unlike the unstabilized Taylor–Hood method, the lack of convergence of Newton’s method did not lead to a blow-up of the solution.

4.4 Meshes

We used several meshes for our experiments. For the methods described in Section 4, the meshes are unstructured simplicial meshes constructed using Netgen [42] using a bulk mesh parameter h_{\max} and a local mesh size $h_{\min} = h_{\max}/250$ on the boundary of the cylinder. The part of the mesh containing the cylinder with $h_{\max} = 8$ is visualized in Figure 2. This is the coarsest mesh considered.

For $h_{\max} = 8, 4, 2$, the Taylor–Hood finite element space with $k = 4$ has a total of 69947, 163107, and 437439 degrees of freedom (dofs), respectively, for the saddle point system. The Scott–Vogelius saddle point systems have 87858, 205176, and 551434 dofs, respectively, due to the larger pressure space. On the same meshes, the MSC method with $k = 4$ has 95856, 224487, and 604047 dofs for the stress and velocity space in which we solve the problem. However, we emphasize that a direct comparison of dofs is not useful to compare methods, as different types of systems must be solved for each method, and the methods have significant differences regarding their approximation power.

4.5 Computing the period

Following [45], we estimate the period of the drag-lift trajectory as follows. For selected points along the trajectory (about 100 per period), we consider points from roughly one-half period away to three-halves away and determine the point of closest approach. We start with an estimated average period, and for each initial point we, obtain an estimated period. Histograms of such estimates are given in Figure 6. Taking the average period from this data gives an estimate of the Strouhal period that can be used to refine the process, providing a new estimated average period. To allow for estimates of the period at a finer grain than the individual time steps, we used a quadratic interpolation of estimated periods to allow for prediction of the period to second order. That is, we found the trajectory error at the optimal point and then considered the error at the time steps on either side of this. Fitting this data with a quadratic and finding the minimum of the quadratic provided an estimated period to within a fraction of Δt .

5 Computational experiments

We consider Reynolds numbers between 120 and 8000. The results for Reynolds numbers 120, 250, 500, 1000, 2000, 4000, and 8000 are shown in Tables 1 to 7. Furthermore, results for Reynolds numbers between 1100 and 1800 are shown in Table 8. Finally, the vorticity of the resulting flow computed on the finest mesh with the MSC method is shown for Reynolds numbers 500, 1000, and 2000 in Figures 3 to 5.

5.1 Reynolds 120–500

We began with a small Reynolds number (120) to provide a low bar for diverse methods. Nevertheless, Table 1 shows that unstabilized, low-order ($k = 2$) Taylor–Hood, which is commonly thought of as the workhorse for fluid flow, struggled to get a viable result. By contrast, all of the high-order methods agree to several digits in the prediction of the Strouhal period and drag. The stabilized Taylor–Hood for $k = 2$ is between the high-order predictions and the unstabilized $k = 2$ results. Note that at this Reynolds number, stabilization of high-order ($k = 4$) Taylor–Hood does not make a difference, except on the size of the divergence error.

Table 2 tells a similar story with Reynolds number 250. The high-order methods are in agreement and the low-order methods are substandard, in particular suggesting erroneous chaotic behavior. This chaotic behavior was also observed in [45], leading to the suggestion that the flow should become chaotic at lower Reynolds numbers compared to some of the literature. We now understand this to be due to under-resolved simulations in [45], where high-frequency errors get amplified by the non-linear term.

For Reynolds numbers 500 and higher, we decided to drop the lowest-order Taylor–Hood simulation data. This is due to the method’s poor performance at the lower Reynolds numbers. In some cases, the unstabilized method did not even continue to time 500, with the Newton iteration failing and the solution blowing up. For Reynolds number 500, c.f. Table 3, even the stabilized lowest-order Taylor–Hood simulation data is not viable. For higher Reynolds numbers, we no longer report results using this method. Finally, the vorticity of the velocity solution from the MCS method on the finest mesh is shown in Figure 3. Here, we see the periodic flow in the wake of the cylinder for about 70 spatial units, after which the flow transitions away into more chaotic behavior. However, since our mesh is very coarse in this part of the domain compared to the resolution around the cylinder, c.f. Figure 2, we do not claim that this is the fully resolved behavior so far downstream in the channel.

5.2 Reynolds 1000 and higher

For Reynolds numbers 1000 and higher, only the high-order grad-div stabilized Taylor–Hood, Scott–Vogelius and the MCS method remain viable. To better understand the behavior of the methods regarding their numerical dissipation, we also consider the MCS method with the centered flux for the convective term. This remains an exactly divergence-free but does not include any numeric dissipation,

Table 1: Results for Reynolds number 120. The consensus values for the average drag is 1.348 and for the Strouhal period is 11.34.

Method	h_{\max}/h_{\min}	Δt	t_{start}	t_{end}	Drag	Period	$\ \nabla \cdot \mathbf{u}_h\ _{\ell^\infty(L^2)}$
MCS ₄	8 / 0.032	0.005	280	480	1.3486	11.339 ± 0.00073	NA
MCS ₄	4 / 0.016	0.0025	280	480	1.3485	11.337 ± 0.00085	NA
MCS ₄	2 / 0.008	0.00125	280	480	1.3482	11.337 ± 0.00098	NA
SV ₄	8 / 0.032	0.01	280	480	1.3542	11.303 ± 0.00354	2.83×10^{-7}
gdTH ₄	8 / 0.032	0.01	280	480	1.3541	11.304 ± 0.00350	3.44×10^{-3}
gdTH ₄	4 / 0.016	0.01	280	480	1.3492	11.339 ± 0.00005	1.52×10^{-3}
gdTH ₄	2 / 0.008	0.01	280	480	1.3476	11.338 ± 0.00010	5.72×10^{-4}
gdTH ₂	8 / 0.032	0.01	280	480	1.4633	11.139 ± 0.28907	1.35×10^{-2}
gdTH ₂	4 / 0.016	0.01	280	480	1.3222	11.814 ± 0.05813	1.10×10^{-2}
TH ₄	8 / 0.032	0.01	280	480	1.3443	11.379 ± 0.00049	3.74×10^0
TH ₄	4 / 0.016	0.01	280	480	1.3492	11.339 ± 0.00005	2.57×10^0
TH ₂	8 / 0.032	0.01	280	380	1.1136	9.518 ± 3.01488	8.61×10^1
TH ₂	4 / 0.016	0.01	280	480	1.0662	7.898 ± 2.89044	5.42×10^0
TH ₂	2 / 0.008	0.01	280	480	1.2397	9.368 ± 2.66038	2.35×10^0

Table 2: Results for Reynolds number 250. The consensus values for the average drag is 1.376 and for the Strouhal period is 9.680.

Method	h_{\max}/h_{\min}	Δt	t_{start}	t_{end}	Drag	Period	$\ \nabla \cdot \mathbf{u}_h\ _{\ell^\infty(L^2)}$
MCS ₄	8 / 0.032	0.005	280	480	1.3750	9.681 ± 0.00011	NA
MCS ₄	4 / 0.016	0.0025	280	480	1.3769	9.673 ± 0.00013	NA
MCS ₄	2 / 0.008	0.00125	280	480	1.3760	9.680 ± 0.00023	NA
gdTH ₄	8 / 0.032	0.01	280	480	1.3867	9.658 ± 0.01316	5.77×10^{-3}
gdTH ₄	4 / 0.016	0.01	280	480	1.3749	9.679 ± 0.00007	3.84×10^{-3}
gdTH ₄	2 / 0.008	0.01	280	480	1.3765	9.684 ± 0.00009	1.21×10^{-3}
gdTH ₂	8 / 0.032	0.01	280	480	1.3531	9.460 ± 1.79078	1.67×10^{-2}
gdTH ₂	4 / 0.016	0.01	280	480	1.3529	9.732 ± 0.35677	1.47×10^{-2}
TH ₂	2 / 0.008	0.01	280	480	1.0144	9.488 ± 2.64040	1.22×10^2

Table 3: Results for Reynolds number 500. The consensus values for the average drag is 1.448 and for the Strouhal period is 8.82.

Method	h_{\max}/h_{\min}	Δt	t_{start}	t_{end}	Drag	Period	$\ \nabla \cdot \mathbf{u}_h\ _{\ell^\infty(L^2)}$
MCS ₄	8 / 0.032	0.002	280	480	1.4464	8.836 ± 0.00004	NA
MCS ₄	4 / 0.016	0.001	280	480	1.4499	8.811 ± 0.00006	NA
MCS ₄	2 / 0.008	0.0005	280	480	1.4477	8.822 ± 0.00016	NA
gdTH ₄	8 / 0.032	0.01	280	450	1.4537	8.823 ± 0.43717	7.82×10^{-3}
gdTH ₄	4 / 0.016	0.01	280	480	1.4488	8.877 ± 0.00796	4.77×10^{-3}
gdTH ₄	2 / 0.008	0.01	280	480	1.4483	8.811 ± 0.00007	2.85×10^{-3}
gdTH ₂	8 / 0.032	0.01	280	480	1.3838	8.966 ± 1.90009	2.09×10^{-2}
gdTH ₂	4 / 0.016	0.01	280	480	1.3753	8.716 ± 1.20772	1.82×10^{-2}

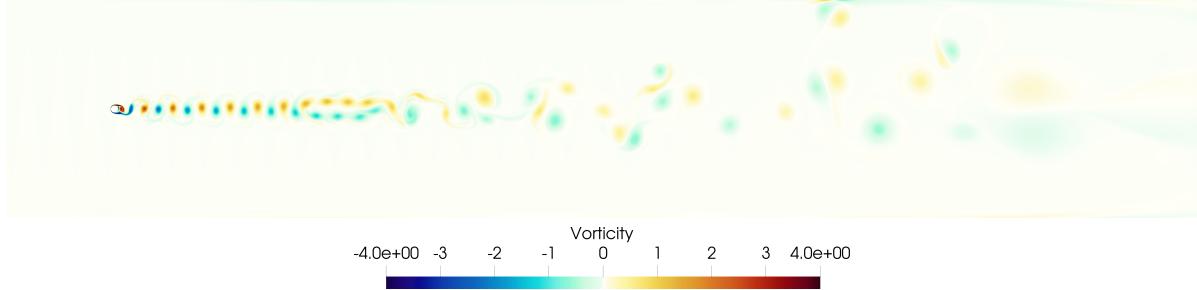


Figure 3: Vorticity of the flow with Reynolds number 500 at $t = 300$ computed using the MCS_4 method on the mesh with $h_{\max} = 2$.

similar to Scott–Vogelius with Crank–Nicolson time-stepping. This will allow us to judge the effect of the numerical dissipation from the upwind convective term in the MCS method used so far, and better judge the effect of the pressure-robust methods. The results are presented in Table 4 for Reynolds number 1000. Additionally, the vorticity of the MSC method on the finest mesh at time $t = 300$ is shown in Figure 4.

At Reynolds number 1000, the drag-lift dynamics are very close to periodic but have a very small standard deviation in the computed period. In particular, we see in Figure 4 that the solution remains periodic for about 40 units downstream of the cylinder before the solution transitions into chaos. However, we again cannot claim high accuracy of the solution in this part of the domain, compared to close to the cylinder. Furthermore, the MCS_4^{cf} , gdTH_4 and SV_4 drag-lift dynamics are only periodic on the finest mesh considered.

In particular, it is interesting to see in Table 4 that the MCS method with the upwind convective term results in a periodic solution on all three meshes. This is in stark contrast to the MCS method with the centered flux in the convective term, Scott–Vogelius, and grad-div stabilized Taylor–Hood results. In these methods, the standard deviation of the period is very large on the coarsest mesh, and the periodic solution is obtained only using the finest mesh. We note that both the MCS with a centered flux and the Scott–Vogelius method do not contain any numerical dissipation and that the grad-div stabilized Taylor–Hood has extra numerical dissipation through the stabilization term. Interestingly, on the second mesh, the MCS method with centered flux has a smaller standard deviation than the Scott–Vogelius method, and the grad-div stabilized Taylor–Hood has the largest standard deviation on the second mesh. Furthermore, we see that the time-step choice is not the critical factor here, as the gdTH_4 results are comparable with those using a smaller time step.

Overall, it is harder to get a close agreement for the average drag and even harder for the Strouhal period at this and higher Reynolds numbers. This illustrates the need for very good methods at higher Reynolds numbers.

In Table 5, we see for Reynolds number 2000 that the drag-lift dynamics is no longer periodic, as also observed in [12]. Similar to [12, Fig. 7], we observe an oscillation of the wake of the flow behind the cylinder in Figure 5.

As the drag-lift dynamics is no longer periodic at Reynolds number 2000, we computed the flow for a sequence of Reynolds number values between 1000 and 2000 to see how the chaos emerges, as shown in Table 8. We obtain a plausibly periodic solution at Reynolds number 1125, but with an ever-increasing standard deviation of the Strouhal period as the Reynolds number increases, even on the finest meshes.

For Reynolds numbers 4000 and higher, we no longer attempt to provide consensus values for the Strouhal period or drag; see Table 6 and Table 7. Furthermore, Figure 6 illustrates how the spread in the computed Strouhal period increases with increasing Reynolds number.

Table 4: Results for Reynolds number 1000. The consensus values for the average drag is 1.54 and for the Strouhal period is 8.36.

Method	h_{\max}/h_{\min}	Δt	t_{start}	t_{end}	Drag	Period	$\ \nabla \cdot \mathbf{u}_h\ _{\ell^\infty(L^2)}$
MCS ₄	8 / 0.032	0.001	280	480	1.5420	8.351 ± 0.00005	NA
MCS ₄	4 / 0.016	0.0005	280	480	1.5504	8.322 ± 0.00036	NA
MCS ₄	2 / 0.008	0.00025	280	480	1.5384	8.361 ± 0.00318	NA
MCS ₄ ^{cf}	8 / 0.032	0.0005	280	480	1.5238	8.516 ± 1.29090	NA
MCS ₄ ^{cf}	4 / 0.016	0.0005	280	480	1.5615	8.295 ± 0.05556	NA
MCS ₄ ^{cf}	2 / 0.008	0.00025	280	480	1.5320	8.393 ± 0.00879	NA
gdTH ₄	8 / 0.032	0.01	280	480	1.5343	8.481 ± 1.30527	1.07×10^{-2}
gdTH ₄	4 / 0.016	0.01	280	480	1.5162	8.297 ± 0.42448	6.79×10^{-3}
gdTH ₄	4 / 0.016	0.005	280	480	1.5163	8.302 ± 0.41216	6.78×10^{-3}
gdTH ₄	2 / 0.008	0.01	280	480	1.5444	8.365 ± 0.00169	5.13×10^{-3}
SV ₄	8 / 0.032	0.01	280	480	1.5400	8.550 ± 1.31361	9.78×10^{-6}
SV ₄	4 / 0.016	0.01	280	480	1.5168	8.278 ± 0.24277	4.40×10^{-7}
SV ₄	2 / 0.008	0.01	280	480	1.5444	8.365 ± 0.00157	1.18×10^{-7}

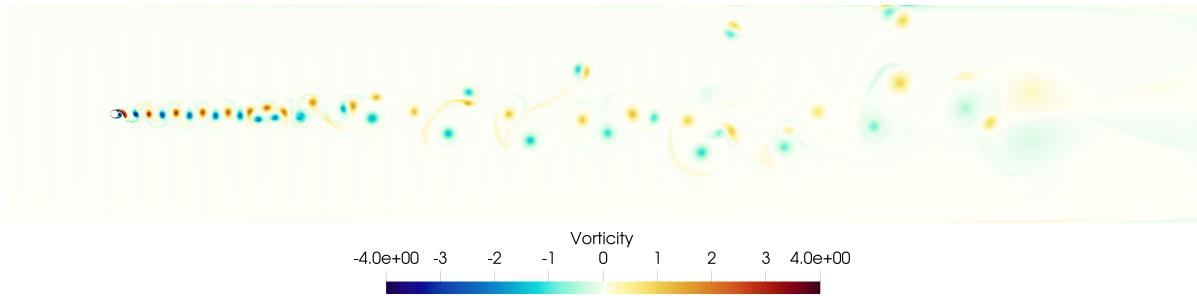


Figure 4: Vorticity of the flow with Reynolds number 1000 at $t = 300$ computed using the MCS₄ method on the mesh with $h_{\max} = 2$.

Table 5: Results for Reynolds number 2000. The consensus values for the average drag is 1.65 and for the Strouhal period is 8.4.

Method	h_{\max}/h_{\min}	Δt	t_{start}	t_{end}	Drag	Period	$\ \nabla \cdot \mathbf{u}_h\ _{\ell^\infty(L^2)}$
MCS ₄	8 / 0.032	0.001	280	480	1.6447	8.441 ± 0.99875	NA
MCS ₄	4 / 0.016	0.0005	280	480	1.6470	8.417 ± 0.95675	NA
MCS ₄	2 / 0.008	0.00025	280	480	1.6411	8.407 ± 0.94202	NA
SV ₄	2 / 0.008	0.01	280	480	1.6603	8.066 ± 0.68569	1.75×10^{-7}
gdTH ₄	4 / 0.016	0.01	280	480	1.5798	8.902 ± 1.27449	9.55×10^{-3}
gdTH ₄	2 / 0.008	0.01	280	480	1.6524	8.365 ± 0.96296	5.61×10^{-3}

Table 6: Results for Reynolds number 4000.

Method	h_{\max}/h_{\min}	Δt	t_{start}	t_{end}	Drag	Period	$\ \nabla \cdot \mathbf{u}_h\ _{\ell^\infty(L^2)}$
MCS ₄	8 / 0.032	0.001	280	480	1.7778	7.867 ± 1.32713	NA
MCS ₄	4 / 0.016	0.0005	280	480	1.5922	9.030 ± 2.33259	NA
MCS ₄	2 / 0.008	0.00025	280	480	1.7717	7.792 ± 1.32658	NA
SV ₄	2 / 0.008	0.01	280	480	1.4719	8.904 ± 2.68945	1.28×10^{-6}
gdTH ₄	4 / 0.016	0.01	280	480	1.6283	8.941 ± 2.01980	1.25×10^{-2}
gdTH ₄	2 / 0.008	0.01	280	480	1.5620	8.997 ± 2.61238	9.48×10^{-3}

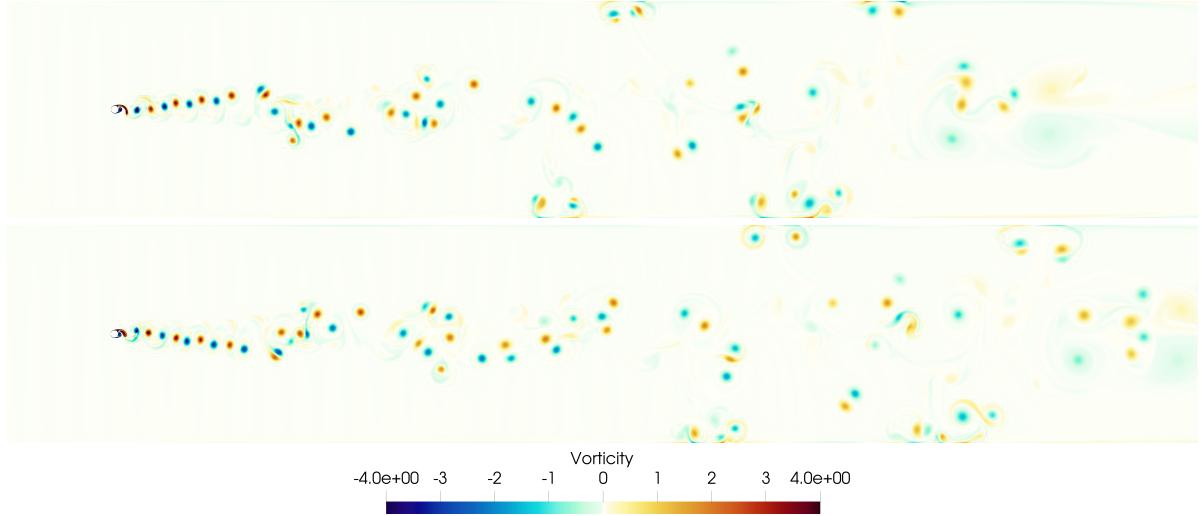


Figure 5: Vorticity of the flow with Reynolds number 2000 computed using the MCS_4 method on the mesh with $h_{\max} = 2$. Top: $t = 310$, Bottom: $t = 350.4$.

Table 7: Results for Reynolds number 8000

Method	h_{\max}/h_{\min}	Δt	t_{start}	t_{end}	Drag	Period	$\ \nabla \cdot \mathbf{u}_h\ _{\ell^\infty(L^2)}$
MCS_4	4 / 0.016	0.0004	280	480	1.6900	8.924 ± 2.46617	NA
MCS_4	2 / 0.008	0.0002	280	480	1.7262	9.252 ± 2.32805	NA
SV_4	2 / 0.008	0.01	280	480	1.6481	9.353 ± 2.23443	3.52×10^{-3}
gdTH_4	4 / 0.016	0.01	280	480	1.5174	8.883 ± 2.27784	1.29×10^{-2}
gdTH_4	2 / 0.008	0.01	280	480	1.6522	9.135 ± 2.38565	1.17×10^{-2}

Table 8: Investigating at what Reynolds number the flow becomes chaotic.

Re	Method	h_{\max}/h_{\min}	Δt	t_{start}	t_{end}	Drag	Period
1100	MCS_4	8 / 0.032	0.001	280	480	1.5581	8.279 ± 0.00258
1125	MCS_4	8 / 0.032	0.001	280	480	1.5628	8.262 ± 0.00822
1150	MCS_4	8 / 0.032	0.001	280	480	1.5660	8.245 ± 0.02200
1175	MCS_4	8 / 0.032	0.001	280	480	1.5718	8.230 ± 0.07870
1200	MCS_4	8 / 0.032	0.001	280	480	1.5743	8.206 ± 0.33194
1200	MCS_4	4 / 0.016	0.0005	280	480	1.5811	8.245 ± 0.91742
1200	MCS_4	2 / 0.008	0.00025	280	480	1.5728	8.227 ± 0.48112
1400	MCS_4	8 / 0.032	0.001	280	480	1.5961	8.534 ± 1.07792
1400	MCS_4	4 / 0.016	0.0005	280	480	1.5947	8.576 ± 1.17649
1400	MCS_4	2 / 0.008	0.00025	280	480	1.5959	8.536 ± 1.11533
1600	MCS_4	8 / 0.032	0.001	280	480	1.6103	8.520 ± 1.15913
1600	MCS_4	4 / 0.016	0.005	280	480	1.6137	8.504 ± 1.16444
1600	MCS_4	2 / 0.008	0.0025	280	480	1.6136	8.507 ± 1.05176
1800	MCS_4	8 / 0.032	0.001	280	480	1.6312	8.462 ± 1.00298
1800	MCS_4	4 / 0.016	0.005	280	480	1.6305	8.475 ± 1.06059
1800	MCS_4	2 / 0.008	0.0025	280	480	1.6327	8.481 ± 0.98248

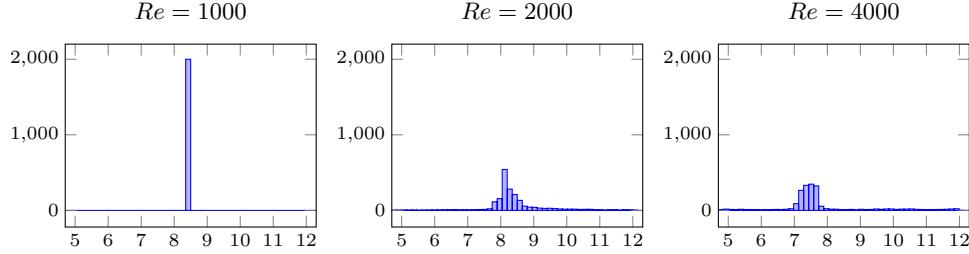


Figure 6: Histogram of the computed periods of the flow at Reynolds numbers 1000, 2000, 4000 using the MCS_4 method with $h_{\max} = 2$.

5.2.1 Just enough numerical dissipation

As far as we know, there are no direct comparisons between conforming and non-conforming pressure-robust methods, so our results give some data on this issue. A direct comparison between fully discontinuous Galerkin and an $H(\text{div})$ -conforming method was performed in [14]. Our results indicate that the $H(\text{div})$ -conforming MCS method with an upwind stabilized convective term is the most accurate scheme considered here, in particular at higher-Reynolds numbers. This is in line with the fact that at high Reynolds numbers, some artificial dissipation is needed unless all scales are resolved [8]. This is due to the non-linear coupling in the Navier–Stokes equations pushing energy (enstrophy in our present, two-dimensional case) to ever higher frequencies. Consequently, for a computational method for high Reynolds number flows to be globally energy consistent with respect to the resolved scales and give optimal resolution where the solution is smooth, it is argued in [8] that the method should have optimal convergence properties for the numerical or artificial dissipation when the solution is smooth and respect the power spectrum where it is rough.

That this is the case for an $H(\text{div})$ -conforming method with upwind stabilization is indicated by the fact that in the inviscid limit, an $H(\text{div})$ -conforming finite element approximation with an upwind flux leads to an optimal error estimate of order $\mathcal{O}(h^{k+0.5})$ [4], something which is not obtained with the conservative central flux [22]. Furthermore, in the divergence-free and conforming case, extra stabilization is needed to obtain the optimal order error estimate in the inviscid limit [5].

Furthermore, we note that in [14], it was also concluded that for an $H(\text{div})$ method, the extra dissipation from the upwind flux in the convective term is of benefit, for three-dimensional under resolved Navier–Stokes flows. Finally, it is argued in [13] that numerical schemes should contain some numerical dissipation in the convective term to overcome problems caused by discretely energy-conserving schemes, such as the inability to represent physical dissipation in the inviscid limit and the resulting accumulation of energy in small scales.

5.2.2 Computational effort

Finally, to give insight into the computational effort required by the studied methods, we present the timings for computation required to reach $t = 10$ on the same computer hardware for the MCS_4 , gdTH_4 and SV_4 methods at $Re = 1000$ in Table 9. Here, we see that the matrix factorization of the resulting sparse linear(ized) systems and the application of each time step are the cheapest for the MCS method and most expensive for the stabilized Taylor–Hood method. However, since the time step must be chosen smaller in the IMEX time-stepping scheme in the MCS method, the time-to-solution is faster for the fully implicit time-stepping schemes on each considered mesh. Nevertheless, we note that because our quasi-Newton scheme reuses the factorization of the Jacobian matrix, as long as this reduces the residual by a factor of 0.1, only one matrix factorization was necessary during the first 1000 time steps where the vortex street has not yet developed.

As seen in Table 4, the coarsest MSC discretization provided the most accurate estimates of the drag and Strouhal period, while both stabilized Taylor–Hood and Scott–Vogelius required finer meshes to resolve all the necessary scales. To estimate the computational time for an extended solution, we

Table 9: Timings for computation until $t = 10$ at $Re = 1000$ on the three considered meshes for the MCS_4 , $gdTH_4$ and SV_4 methods. Matrix factorizations computed using the sparse direct solver PARDISO through the Intel Math Kernel Library. Computed using four shared memory threads on an Intel Core i7-13700F processor.

Method	h_{\max} / h_{\min}	Δt	Total Time (s)	Time Factorizing (s)	# Matrix Factorizations
MCS_4	8 / 0.032	0.001	258.1	0.8	2
MCS_4	4 / 0.016	0.0005	1125.0	2.0	2
MCS_4	2 / 0.008	0.00025	6781.4	5.0	2
$gdTH_4$	8 / 0.032	0.01	90.8	1.9	1
$gdTH_4$	4 / 0.016	0.01	208.4	5.1	1
$gdTH_4$	2 / 0.008	0.01	613.9	56.2	1
SV_4	8 / 0.032	0.01	65.6	1.6	1
SV_4	4 / 0.016	0.01	151.4	3.7	1
SV_4	2 / 0.008	0.01	407.8	10.8	1

extrapolate the data in Table 9 to $t = 500^{\dagger}$. The computational wall time for the MSC method on the coarsest mesh is approximately 4h. The grad-div stabilized Taylor–Hood computation on the finest mesh required 135 updates of the Jacobean. This gives an approximate computational time of 10h. The Scott–Vogelius method required 425 updates of the Jacobean. This gives an approximate computational time of 7h. Consequently, the MSC method (with upwinding) provided an effective tool for exploration on coarser meshes, as indicated in the first four entries in Table 8. On the other hand, to resolve the drag and period to high precision required more resolution in some cases.

5.3 Other methods

We survey briefly other methods that have been used for simulating flow around a cylinder.

5.3.1 Vortex blob

The publication [12, Figure 5] suggests that drag and lift for flow around a cylinder is nearly periodic at $Re = 1000$, in agreement with our findings, with chaotic flow for $Re \geq 2000$.

5.3.2 OpenFoam

OpenFoam may not be the ideal discretization approach for low Reynolds numbers. In [27], Strouhal numbers and frequencies computed with OpenFoam were found to be in error by greater than a factor of two for $Re \in [55, 161]$.

6 Conclusions

We considered several finite element methods for solving flow past a cylinder. All of these methods are well studied both computationally and theoretically. We found that high-order and pressure-robust methods with the correct amount of numerical dissipation are necessary to obtain reliable results at higher Reynolds numbers. The widely-used, lowest-order Taylor–Hood method was not accurate enough to give appropriate results, even at relatively low Reynolds numbers. We were able to confirm some earlier results by other researchers. We found that exact incompressibility and appropriate numerical dissipation are important at higher Reynolds numbers. This is particularly the case when not all scales are resolved, which is often the case at high Reynolds numbers. In particular, we found that both H^1 -conforming and $H(\text{div})$ -conforming pressure-robust methods without any numerical dissipation from the convective term did not lead to accurate results unless very fine meshes were considered. Finally, we studied in detail how chaos appears to emerge for Reynolds numbers just above 1000 using the

[†]The actual runs were performed on different nodes of the cluster “oscar” at Brown University with different computational hardware and without control over other computations effecting computational run time.

pressure-robust MSC method with upwinding which appears to have the correct amount of numerical dissipation for reliable results.

Data Availability Statement

The code used to realize the presented results, the raw data generated by this code and videos of key simulations are available on zenodo at <https://doi.org/10.5281/zenodo.11490084>.

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