

THE EFFECT OF METRIC BEHAVIOR AT SPATIAL INFINITY ON POINTWISE WAVE DECAY IN THE ASYMPTOTICALLY FLAT STATIONARY SETTING.

KATRINA MORGAN

ABSTRACT. The current work considers solutions to the wave equation on asymptotically flat, stationary, Lorentzian spacetimes in $(1+3)$ dimensions. We investigate the relationship between the rate at which the geometry tends to flat and the pointwise decay rate of solutions. The case where the spacetime tends toward flat at a rate of $|x|^{-1}$ was studied in [33], where a t^{-3} pointwise decay rate was established. Here we extend the result to geometries tending toward flat at a rate of $|x|^{-\kappa}$ and establish a pointwise decay rate of $t^{-\kappa-2}$ for $\kappa \in \mathbb{N}$ with $\kappa \geq 2$. We assume a weak local energy decay estimate holds, which restricts the geodesic trapping allowed on the underlying geometry. We use the resolvent to connect the time Fourier Transform of a solution to the Cauchy data. Ultimately the rate of pointwise wave decay depends on the low frequency behavior of the resolvent, which is sensitive to the rate at which the background geometry tends to flat.

1. INTRODUCTION

This work examines the effect of the far away metric behavior on pointwise wave decay on asymptotically flat, stationary backgrounds in $(1+3)$ dimensions. A spacetime geometry is asymptotically flat if the metric coefficients tend toward the flat Minkowski metric ($\mathbf{m} = \text{diag}(-1, 1, 1, 1)$ in (t, x) coordinates) as $r \rightarrow \infty$. Here and throughout the paper we take $r := |x|$. A geometry is said to be stationary if the metric coefficients are time independent. The main result of this paper quantifies the relationship between the rate at which the background geometry tends to flat and the rate of pointwise wave decay.

The pointwise decay rates established in this work interpolate between two known cases: the flat Minkowski spacetime and asymptotically flat spacetimes tending toward flat at a rate of r^{-1} . Table 1 summarizes these results, which hold for compactly supported initial data. In the case of [33] and the current work, the assumptions on the initial data can be weakened.

Since we are working in three spatial dimensions, sharp Huygens' principle says that solutions to the wave equation on the flat Minkowski spacetime decay all the way to 0 in finite time at each point in space. In Table 1 we make a weaker statement that solutions to the wave equation have arbitrarily fast polynomial time decay.

The author was supported in part by the NSF under Grant No. 1440140 while the author was in residence at the Mathematical Sciences Research Institute in Berkeley, California. The author thanks Jason Metcalfe for his invaluable guidance on this project and an anonymous referee for their helpful feedback.

	METRIC BEHAVIOR	POINTWISE WAVE DECAY
[33] :	$\mathbf{g} = flat + \mathcal{O}(r^{-1-\delta})$	$ u(t, x) \lesssim_x t^{-3}$
Current Work:	$\mathbf{g} = flat + \mathcal{O}(r^{-\kappa-\delta})$	$ u(t, x) \lesssim_x t^{-\kappa-2}$
Sharp Huygens':	$\mathbf{g} = flat$	$ u(t, x) \lesssim_x t^{-\infty}$

TABLE 1. Summary of Decay Rates

The results in [\[33\]](#) and the current work rely on an assumption that a dispersive estimate called *weak local energy decay* holds. Heuristically, this estimate imposes sufficient restrictions on the behavior of the geometry within compact regions so only the long range metric behavior is left to be studied. A more detailed summary of local energy decay is provided later in the introduction.

Asymptotically flat spacetimes arise in general relativity, which has motivated a variety of mathematical questions about wave behavior in this setting. For example, the Schwarzschild metric describing the geometry of space in the presence of a single, non-rotating black hole and the Kerr metric describing spacetime in the presence of a single, axially symmetric, rotating black hole both tend toward flat at a rate of r^{-1} . A conjecture posited by physicist Richard Price in the 1970's in [\[27\]](#), known as Price's Law, predicted a t^{-3} decay rate for waves on the Schwarzschild metric. The question of proving Price's law was explored in [\[33\]](#) and also in [\[13\]](#), where they analyze the wave behavior via spherical modes using the spherical symmetry of the Schwarzschild metric. Pointwise decay rates for the Kerr spacetime were studied in [\[11\]](#) and [\[14\]](#). In [\[25\]](#) the authors proved Price's Law for non-stationary asymptotically flat spacetimes and established the t^{-3} decay rate for a class of perturbations of the Kerr spacetime. The techniques in [\[13\]](#), [\[33\]](#), and the current work involve taking the Fourier transform in time and therefore do not readily extend to non-stationary geometries.

Weak local energy decay on the Schwarzschild geometry was established in [\[3\]](#), [\[9\]](#), and [\[19\]](#). For the Kerr spacetime with low angular momentum, weak local energy decay estimates were proved in [\[2\]](#), [\[9\]](#), and [\[12\]](#). The assumptions in [\[33\]](#) therefore hold for Schwarzschild and Kerr with low angular momentum. A major challenge in obtaining local energy estimates for the Kerr and Schwarzschild geometries is the presence of trapping, in which a portion of the wave flow remains within a fixed set.

A natural question arising from Tataru's result in [\[33\]](#) is: What aspects of the Schwarzschild geometry dictate Price's Law? There are three locations that are a priori suspected to affect this decay rate: the event horizon, the photon sphere, and the behavior of the perturbation at spatial infinity. The current work shows that the metric behavior at spatial infinity dictates the pointwise decay rate of waves when the weak local energy estimate holds. On the Schwarzschild background, trapping occurs in two areas called the event horizon and the photon sphere. The trapping at the event horizon has been shown to be trivial due to what is known as the red-shift effect, which guarantees energy decay along the trapped rays ([\[11\]](#), [\[9\]](#)). The photon sphere corresponds to a fixed radius, and rays initially tangential to this surface remain there for all time. The behavior on Kerr backgrounds is more complicated. The trapping at the event horizon is similarly known to be trivial, but the other trapped set does not occur on a fixed radius and can only be described in phase space. In order to deal with trapping, a weak

local energy estimate with zero coefficients on the trapped set is often introduced. If this holds, then one obtains local energy estimates on the trapped set with a derivative loss. Our definition of the weak local energy decay estimate includes this derivative loss.

Questions similar to the aim of this paper were studied in [6] and [5] where the authors established local decay rates for waves on asymptotically flat, stationary spacetimes which tend toward flat at different rates. There are several key differences compared with the current work. First, we handle full Lorentzian perturbations of flat Minkowski space rather than restricting to perturbations of the Laplacian. This leads to the metrics considered in this paper containing $dt dx_i$ terms, which results in mixed space-time differential operators in our wave operator. Second, we allow for the possibility of unstable trapping on our background. In [6] and [5], a nontrapping assumption is used in order to obtain decay for the high frequency part of a solution to the wave equation (it is not needed for the low frequency part). Third, our result improves upon the established decay rates. Finally we note that [5] considers $(1+n)$ dimensional geometries for $n \geq 2$ and [6] considers n odd with $n \geq 3$. The current work only studies $(1+3)$ dimensional spacetimes.

1.1. The Wave Equation. The flat wave operator is given by

$$\square = -\partial_t^2 + \sum_{i=1}^3 \partial_{x_i}^2 = -\partial_t^2 + \Delta_x$$

where Δ_x is the spatial Laplacian. Throughout the paper we write $\Delta = \Delta_x$. Similarly we write $\nabla = \nabla_x$ for the spatial gradient. When both time and spatial derivatives are considered we use ∂ .

The wave operator associated to a Lorentzian metric $\mathbf{g} = \mathbf{g}_{\alpha\beta} d\alpha d\beta$ with signature $(3,1)$ is given by

$$\square_{\mathbf{g}} = \frac{1}{\sqrt{|\mathbf{g}|}} \partial_{\alpha} \sqrt{|\mathbf{g}|} \mathbf{g}^{\alpha\beta} \partial_{\beta} \quad (1.1)$$

where $|\mathbf{g}|$ is the determinant of the matrix associated to the metric, and α and β are summed over both time and space dimensions. We use Latin indices i, j to indicate only spatial dimensions are being considered and Greek indices to indicate both space and time dimensions are being considered. When we wish to specify Cartesian vs. spherical coordinates, we use α, β for Cartesian and γ, δ for spherical coordinates.

The flat metric (i.e. the Minkowski metric) is given in rectangular coordinates by

$$\mathbf{m} = -dt^2 + \sum_{i=1}^3 dx_i^2. \quad (1.2)$$

Taking $\mathbf{g} = \mathbf{m}$ in (1.1) thus yields $\square_{\mathbf{m}} = \square$, as one would expect.

1.2. Energy Estimates. We are interested in the Cauchy problem

$$(\square_{\mathbf{g}} + V)u = f, \quad u(0, x) = u_0, \quad \partial_t u(0, x) = u_1 \quad (1.3)$$

where V is a scalar potential. The assumptions placed on $\mathbf{g}$ and V are given in Section 1.5. The Cauchy data at time t is denoted $u[t] = (u(t, \cdot), \partial_t u(t, \cdot))$.

Definition 1.1. We say the evolution (1.3) satisfies the uniform energy bounds if:

$$\|u[t]\|_{\dot{H}^{k,1} \times H^k} \leq c_k(\|u[0]\|_{\dot{H}^{k,1} \times H^k} + \|f\|_{H^k L^1}), \quad t \geq 0, \quad k \geq 0. \quad (1.4)$$

Here H^k denotes the usual Sobolev space, and we say $\phi \in \dot{H}^{k,1}$ if $\nabla \phi \in H^k$.

1.2.1. Local Energy Decay. Local energy decay estimates originated in the work of Morawetz ([26]) where the author established a dispersive estimate for solutions to the flat wave equation. In [16] the authors presented a new approach for proving existence of solutions for nonlinear waves which relied on obtaining a Morawetz-type estimate. The use of local energy estimates has since become a standard tool for studying nonlinear wave equations (e.g. [4], [15], [20], [31], [21], [17], [18], [35], among many others).

The original Morawetz estimate considered a solution u to the homogeneous flat wave equation with initial data u_0, u_1 and states

$$\int_0^t \int_{\mathbb{R}^3} \frac{1}{|x|} |\nabla u|^2(t, x) \, dt dx \lesssim \|\nabla u_0\|_{L^2}^2 + \|u_1\|_{L^2}^2.$$

Restricting to compact regions in space, one is able to obtain similar bounds on u and its derivatives (see e.g. [16], [30], and [32]). Our definitions for the local energy norms will restrict to dyadic spatial regions. We use $\langle r \rangle$ to indicate a smooth function of r such that $\langle r \rangle \geq 2$ and $\langle r \rangle = r$ for $r \geq 3$, and we define $A_m := \{x : 2^m \leq \langle r \rangle \leq 2^{m+1}\}$. One benefit of using these dyadic regions is that $r \approx 2^m$ on the region of integration, so the weights in the local energy norm can roughly be treated as constant within the region of integration.

The local energy norm we use is defined by

$$\|u\|_{LE} = \sup_m \|\langle r \rangle^{-\frac{1}{2}} u\|_{L^2(\mathbb{R}_+ \times A_m)}.$$

Its H^1 analogue is given by

$$\|u\|_{LE^1} = \|\partial u\|_{LE} + \|\langle r \rangle^{-1} u\|_{LE},$$

and the dual norm is given by

$$\|f\|_{LE^*} = \sum_m \|\langle r \rangle^{\frac{1}{2}} f\|_{L^2(\mathbb{R}_+ \times A_m)}.$$

For functions with higher regularity we define the following norms

$$\|u\|_{LE^{1,N}} = \sum_{j \leq N} \|\partial^j u\|_{LE^1}, \quad \|f\|_{LE^{*,N}} = \sum_{j \leq N} \|\partial^j f\|_{LE^*}.$$

The spatial counterparts of the LE and LE^* space-time norms are

$$\|v\|_{\mathcal{LE}} = \sup_m \|\langle r \rangle^{-\frac{1}{2}} v\|_{L^2(A_m)}; \quad \|g\|_{\mathcal{LE}^*} = \sum_m \|\langle r \rangle^{\frac{1}{2}} g\|_{L^2(A_m)}$$

with the higher regularity norms defined by

$$\|v\|_{\mathcal{LE}^N} = \sum_{j \leq N} \|\nabla^j v\|_{\mathcal{LE}}, \quad \|g\|_{\mathcal{LE}^{*,N}} = \sum_{j \leq N} \|\nabla^j g\|_{\mathcal{LE}^*}.$$

Definition 1.2. We say the evolution (1.3) satisfies the local energy decay estimate if:

$$\|u\|_{LE^{1,N}} \leq c_N(\|u[0]\|_{H^{N,1} \times H^N} + \|f\|_{LE^{*,N}}), \quad N \geq 0. \quad (1.5)$$

Heuristically, the local energy decay estimate holds if the underlying geometry allows waves to spread out enough so that the energy within compact spatial regions decays sufficiently quickly to be integrable in time. Local energy decay has been used to establish other dispersive estimates such as Strichartz estimates (global, mixed norm estimates) in [24], [34], and [19] and pointwise estimates in [10], [25], [33] (among others).

The local energy decay estimate is known to hold in several nontrapping geometries. For sufficiently small perturbations of flat space without trapping, local energy decay was established in [1], [21], and [23]. The case of stationary product manifolds was considered in [8], [5], and [31]. The nontrapping case was studied more generally in [22]. If trapping occurs then the local energy decay estimate does not hold ([28], [29]).

1.2.2. Weak Local Energy Decay. Trapping on the background geometry may be stable or unstable. A spacetime with trapping where every trapped geodesic is unstable may still admit a weaker form of the local energy decay estimate. In the case of trapping, there is necessarily a loss of derivatives on the right hand side of the estimate (see e.g. [7]).

Definition 1.3. *We say the evolution (1.3) satisfies the weak local energy decay estimate if:*

$$\|u\|_{LE^{1,N}} \leq c_N(\|u[0]\|_{\dot{H}^{N+3,1} \times H^{N+3}} + \|f\|_{LE^{*,N+3}}), \quad N \geq 0. \quad (1.6)$$

1.3. Vector Fields and Weighted Sobolev Spaces. Our argument will use vector field methods. Specifically we are interested in the vector fields:

- Rotations: $\Omega = \{\Omega_{ab} \mid a, b = 1, 2, 3\}$ where $\Omega_{ab} = x_a \partial_b - x_b \partial_a$.
- Translations: $T = \{T_a \mid a = 1, 2, 3\}$ where $T_a = \nabla_a$.
- Scaling: $S = S_r - S_\tau$ where $S_r = r \partial_r$ and $S_\tau = \tau \partial_\tau$.

Note that the scaling vector field we use is taken in time frequency space and therefore differs from the scaling vector field in physical space which is given by $r \partial_r + t \partial_t$. This is because we use the vector field arguments only on the time Fourier transform side. We denote the collection of all such vector fields by $\Gamma = \{\Omega, T, S\}$. We write $\Gamma^{<n}$ to denote a linear combination of Γ^α for $|\alpha| < n$: $\Gamma^{<n} := \sum_{|\alpha| < n} c_\alpha \Gamma^\alpha$.

We use the vector fields to define a weighted Sobolev type norm. We will assume the initial data lies in such a space. The weighted Sobolev spaces $Z^{n,q}$ are defined by

$$\|\phi\|_{Z^{n,q}} = \sup_{i+j+k \leq n} \|\langle r \rangle^q T^i \Omega^j S_r^k \phi\|_{\mathcal{LE}^*}. \quad (1.7)$$

1.4. Symbol Classes. We will assume that the metric coefficients of the background geometry belong to certain symbol classes.

The symbol classes $S(r^q)$, $\ell^1 S(r^q)$, $S(\log r)$ are defined as follows:

$$\begin{aligned} \phi(x) \in S(r^q) &\Leftrightarrow \|\langle r \rangle^{j-q} \partial^j f(x)\|_{L^\infty(\mathbb{R}^3)} \lesssim_j 1 \quad j \in \{0, 1, 2, \dots\} \\ \phi(x) \in \ell^1 S(r^q) &\Leftrightarrow \sum_m 2^{m(j-q)} \|\partial^j f(x)\|_{L^\infty(A_m)} \lesssim_j 1 \quad j \in \{0, 1, 2, \dots\} \\ \phi(x) \in S(\log r) &\Leftrightarrow \|(\log \langle r \rangle)^{-1} f(x)\|_{L^\infty(\mathbb{R}^3)} \lesssim 1 \text{ and } \|\langle r \rangle^j \partial^j f(x)\|_{L^\infty(\mathbb{R}^3)} \lesssim_j 1, \\ &\quad j \in \{1, 2, 3, \dots\}. \end{aligned}$$

If $\phi \in S(r^q)$ is radial, we write $\phi \in S_{rad}(r^q)$. We indicate radial functions in the other symbol classes analogously.

In some of our calculations we use the notation ρ_ℓ^q to indicate a representative of the symbol class $\ell^1 S(r^q)$. Similarly, we use ρ^q to represent $S(r^q)$ and ρ_r^q to represent $S_{rad}(r^q)$. We allow ρ_ℓ^q , ρ_r^q , and ρ^q to stand for different functions at each appearance.

1.5. Statement of Main Theorem. We consider a Lorentzian metric $\mathbf{g}$ with the following properties:

Metric Assumptions

- (1) $\mathbf{g}$ is stationary (i.e. the metric coefficients are time independent).
- (2) The submanifolds $t = \text{constant}$ are space-like (i.e. the induced metric on the spatial submanifolds is positive definite).
- (3) Let $\kappa \in \mathbb{N}$ with $\kappa \geq 2$. The metric $\mathbf{g}$ is asymptotically flat in the sense that $\mathbf{g}$ can be written as

$$\mathbf{g} = \mathbf{m} + \mathbf{f} + \mathbf{h}$$

where

$$\mathbf{f} = f_{00}(x)dt^2 + f_{0i}(x)dt dx_i + f_{ij}(x)dx_i dx_j$$

with $f_{\alpha\beta} \in \ell^1 S(r^{-\kappa})$ for $\alpha, \beta \in \{0, 1, 2, 3\}$ and

$$\mathbf{h} = h_{tt}(r)dt^2 + h_{tr}(r)dt dr + h_{rr}(r)dr^2 + h_{\omega\omega}(r)r^2 d\omega^2$$

with $h_{\gamma\delta} \in S_{rad}(r^{-\kappa})$ for $\gamma, \delta \in \{t, r, \omega\}$. Here $d\omega^2 = d\theta^2 + \sin^2 \theta d\phi^2$.

We note these assumptions but with $\kappa = 1$ match those in [33], and thus the results apply here. We will appeal to Tataru's results for steps of the proof of the main theorem which are not sensitive to the rate at which the background geometry tends to flat.

Remark 1.4. In [33] the author also considers geometries exterior to an open ball where the lateral boundary $\mathbb{R} \times \partial B(0, R_0)$ is outgoing space-like. Our results are likely to also hold in this case. However, due to Birkhoff's theorem, the applications to black holes in this setting appear limited so we do not pursue the matter here.

Theorem 1.5. Let $\mathbf{g}$ be a $(1+3)$ -dimensional spacetime satisfying metric assumptions 1-3 above. Let V be a potential of the form

$$V(x) = V_\ell(x) + V_r(r), \quad V_\ell \in \ell^1 S(r^{-\kappa-2}), \quad V_r \in S_{rad}(r^{-\kappa-2}). \quad (1.8)$$

Assume the homogeneous Cauchy problem

$$(\square_g + V)u(t, x) = 0, \quad u(0, x) = u_0, \quad \partial_t u(0, x) = u_1 \quad (1.9)$$

satisfies the uniform energy bound (1.4) and the weak local energy decay assumption (1.6). If u solves (1.9) with $u_0 \in Z^{\nu+1, \kappa}$ and $u_1 \in Z^{\nu, \kappa+1}$ for $\nu \geq 31\kappa + 168$, then in normalized coordinates (see section 2) u satisfies the bounds

$$|u(t, x)| \lesssim \frac{1}{\langle t \rangle \langle t - r \rangle^{\kappa+1}} (\|u_0\|_{Z^{\nu+1, \kappa}} + \|u_1\|_{Z^{\nu, \kappa+1}}) \quad (1.10)$$

$$|\partial_t u(t, x)| \lesssim \frac{1}{\langle t \rangle \langle t - r \rangle^{\kappa+2}} (\|u_0\|_{Z^{\nu+1, \kappa}} + \|u_1\|_{Z^{\nu, \kappa+1}}). \quad (1.11)$$

1.6. Cutoff and Bump Functions. The function $\chi_{<1}(r)$ is defined to be a smooth function which is 1 for $r \leq 1$ and 0 for $r \geq 2$. We define $\chi_{>1}(r) := 1 - \chi_{<1}(r)$ so that $\chi_{>1}(r)$ is a smooth function which is 1 for $r \geq 2$ and 0 for $r \leq 1$. We define $\chi_{\approx 1}(r)$ to be a smooth function which is 1 for $1 \leq r \leq 2$ and 0 for $r < \frac{1}{2}$ and $r > 4$.

We define $\beta_{\approx m}(r)$ for $m \geq 0$ to be a smooth partition of unity which is subordinate to the dyadic intervals A_m .

When restricting r by dyadic regions, we use $\chi_{\approx m}$ to indicate $\chi_{\approx m}(r) = \chi_{\approx}(\frac{r}{2^m})$, $\chi_{>m}(r) = \chi_{>1}(\frac{r}{2^m})$, and $\chi_{<m}(r) = \chi_{<1}(\frac{r}{2^m})$. We note $\partial_r \chi_{<m}(r)$ and $\partial_r \chi_{>m}(r)$ are each supported on A_m and for some constants c_1 and c_2 we have

$$\|\partial_r \chi_{<m}(r)\|_{L^2(A_m)} = c_1 2^{\frac{3m}{2}} \quad \|\partial_r \chi_{>m}(r)\|_{L^2(A_m)} = c_2 2^{\frac{3m}{2}}.$$

In other contexts where we restrict to $r < R$ or $r > R$ for some constant R we write $\chi_{\approx R} = \chi_{\approx}(\frac{r}{R})$, etc..

1.7. Argument Summary. We first fix a coordinate system that allows us to write the operator $(\square_g + V)$ in the form

$$P = -\partial_t^2 + \Delta + \partial_t P^1 + P^2 \tag{1.12}$$

where P^1 and P^2 are spatial operators of order 1 and 2, respectively. The coefficients of the operators depend on the metric coefficients assumed in the main theorem. We then use the resolvent (denoted R_τ) to connect the time Fourier transform of a solution u to the Cauchy problem (1.9) with the initial data.

We define the resolvent to be the inverse of the image of P under the time Fourier transform, when the inverse exists. We will establish that if u solves (1.9) then

$$\hat{u}(\tau) = R_\tau(-i\tau u_0 + P^1 u_0 - u_1). \tag{1.13}$$

The final pointwise decay rate is then proved by analyzing the resolvent and inverting the Fourier transform.

Our argument will be different for high frequencies ($|\tau| \gtrsim 1$) and low frequencies ($|\tau| \lesssim 1$). Roughly speaking, the low frequency behavior is sensitive to the metric behavior at spatial infinity while the high frequency behavior is sensitive to trapping. We assume the weak local energy decay estimate holds so that some trapping may occur, but this estimate provides enough information to obtain decay for the high frequency part of our solution u . It is the low frequency behavior that depends on the metric perturbation at spatial infinity and dictates the pointwise decay rate. We obtain an expansion in powers of r^{-1} for the resolvent at zero frequency and use this to calculate the error when we estimate $R_\tau u_0$ by $(R_0 u_0)e^{-i\tau \langle r \rangle}$ for the resolvent at low frequencies. We then apply the inverse Fourier transform to the terms arising in this estimate. The behavior of these terms dictates the final pointwise decay rate.

This approach is due to [33]. A key difference in our analysis is that we need to go further down in the expansion of the zero resolvent in order to obtain improved decay rates. Changing the expansion then affects the error in the estimate for $R_\tau u_0$ when $|\tau|$ is small. The rate at which the background geometry tends toward flat (indicated by the parameter κ in the statement of the main theorem) ultimately determines how far down in the expansion of the zero resolvent we are able to go, which determines the error terms

in our low frequency resolvent estimate and in turn determines the result of inverting the Fourier transform.

Outline of the Paper

- Section 2: Replace $\square_g + V$ by P as in (1.12).
- Section 3: Define the resolvent and appeal to [33] to establish the desired mapping properties.
- Section 4: Analyze the resolvent at zero frequency. This analysis depends on the rate at which the background geometry tends toward flat.
- Section 5: Estimate the low frequency resolvent $R_\tau u_0$ by $(R_0 u_0)e^{-i\tau\langle r \rangle}$ and calculate the error using the results of section 4.
- Section 6: Establish pointwise bounds on derivatives of the resolvent that will be used when inverting the Fourier transform. We do not improve upon the bounds established in [33], but we do track the resulting regularity requirements more precisely and correct one proposition statement.
- Section 7: Invert the Fourier transform to prove theorem (1.5).

2. COORDINATE CHANGE

In this section we establish a normalized coordinate system in which the operator $\square_g + V$ in the statement of Theorem (1.5) can be replaced by an operator P of the form

$$P = -\partial_t^2 + \Delta + \partial_t P^1 + P^2 \quad (2.1)$$

where

$$P^1 = \partial_i p_1^i + p_1^i \partial_i, \quad p_1^i \in \ell^1 S(r^{-\kappa}) \quad (2.2)$$

and

$$\begin{aligned} P^2 &= \partial_i p_2^{ij} \partial_j + p_2^\omega \Delta_\omega + V_\ell + V_r, \\ p_2^{ij} &\in \ell^1 S(r^{-\kappa}); \quad V_\ell \in \ell^1 S(r^{-\kappa-2}); \quad p_2^\omega, V_r \in S_{rad}(r^{-\kappa-2}). \end{aligned} \quad (2.3)$$

The calculations in this section encode the geometric assumptions into the differential operator. Throughout the rest of the paper we will work in the normalized coordinates established here. The statement of the main theorem is given in these coordinates.

Metric Assumption 3 (given in the Introduction) can be restated in dual coefficients and using spherical coordinates as

$$\mathbf{g}^{\gamma\delta} = \mathbf{m}^{\gamma\delta} + \mathbf{f}^{\gamma\delta} + \mathbf{h}^{\gamma\delta} \quad (2.4)$$

where

$$\left[\mathbf{m}^{\gamma\delta} \right] = \begin{bmatrix} -1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & \frac{1}{r^2} & 0 \\ 0 & 0 & 0 & \frac{1}{r^2 \sin^2 \theta} \end{bmatrix}, \quad \left[\mathbf{f}^{\gamma\delta} \right] = \begin{bmatrix} \frac{\mathbf{f}^{tt}}{r \sin \theta} & \frac{\mathbf{f}^{tr}}{r \sin \theta} & \frac{\frac{\mathbf{f}^{t\theta}}{r}}{r^2 \sin \theta} & \frac{\frac{\mathbf{f}^{t\phi}}{r \sin \theta}}{r^2 \sin^2 \theta} \\ \frac{\mathbf{f}^{rt}}{r \sin \theta} & \frac{\mathbf{f}^{rr}}{r \sin \theta} & \frac{\frac{\mathbf{f}^{\theta r}}{r}}{r^2 \sin \theta} & \frac{\frac{\mathbf{f}^{r\phi}}{r \sin \theta}}{r^2 \sin^2 \theta} \\ \frac{\frac{\mathbf{f}^{\theta t}}{r}}{r^2 \sin \theta} & \frac{\frac{\mathbf{f}^{\theta r}}{r}}{r^2 \sin \theta} & \frac{\frac{\mathbf{f}^{\theta\theta}}{r^2}}{r^2 \sin \theta} & \frac{\frac{\mathbf{f}^{\theta\phi}}{r^2 \sin \theta}}{r^2 \sin^2 \theta} \\ \frac{\frac{\mathbf{f}^{\phi t}}{r \sin \theta}}{r^2 \sin \theta} & \frac{\frac{\mathbf{f}^{\phi r}}{r \sin \theta}}{r^2 \sin \theta} & \frac{\frac{\mathbf{f}^{\phi\theta}}{r^2 \sin \theta}}{r^2 \sin \theta} & \frac{\frac{\mathbf{f}^{\phi\phi}}{r^2 \sin^2 \theta}}{r^2 \sin^2 \theta} \end{bmatrix},$$

and

$$[\mathfrak{h}^{\gamma\delta}] = \begin{bmatrix} \mathfrak{h}^{tt} & \mathfrak{h}^{tr} & 0 & 0 \\ \mathfrak{h}^{rt} & \mathfrak{h}^{rr} & 0 & 0 \\ 0 & 0 & \frac{\mathfrak{h}^{\omega\omega}}{r^2} & 0 \\ 0 & 0 & 0 & \frac{\mathfrak{h}^{\omega\omega}}{r^2 \sin^2 \theta} \end{bmatrix}$$

with $\mathfrak{f}^{\gamma\delta} \in \ell^1 S(r^{-\kappa})$ and $\mathfrak{h}^{\gamma\delta} \in S_{rad}(r^{-\kappa})$. Furthermore $[\mathfrak{f}^{\gamma\delta}]$ and $[\mathfrak{h}^{\gamma\delta}]$ are symmetric (i.e. $\mathfrak{f}^{\gamma\delta} = \mathfrak{f}^{\delta\gamma}$ and $\mathfrak{h}^{\gamma\delta} = \mathfrak{h}^{\delta\gamma}$).

Lemma 2.1. *There exists a coordinate system so that $\mathfrak{g}$ satisfies Metric Assumptions 1-3 as well as the additional condition*

$$\mathfrak{h}^{rr} = -\mathfrak{h}^{tt}, \quad \mathfrak{h}^{tr} = 0.$$

Proof. In order to achieve $\mathfrak{h}^{tr} = 0$, we reset t via the coordinate change

$$dT = dt - \chi_{>R} \frac{\mathfrak{h}^{rt}}{1 + \mathfrak{h}^{rr}} dr$$

where R is a constant chosen to be sufficiently large so that $1 + \mathfrak{h}^{rr} \gtrsim 1$ for $r > R$. Note we have $\chi_{>R} \frac{\mathfrak{h}^{rt}}{1 + \mathfrak{h}^{rr}} \in S_{rad}(r^{-\kappa})$.

To see Assumption 1 still holds, we write the coordinate change as $T = t + Q(r)$ where $Q'(r) = -\chi_{>R} \frac{\mathfrak{h}^{rt}}{1 + \mathfrak{h}^{rr}}$. Thus $\frac{\partial}{\partial T} = \frac{\partial}{\partial t}$ so the metric coefficients remain independent of the time variable T .

To see Assumption 2 still holds, we calculate

$$\langle dT, dT \rangle = \langle dt, dt \rangle - 2\chi_{>R} \frac{\mathfrak{h}^{rt}}{1 + \mathfrak{h}^{rr}} \langle dt, dr \rangle + \left(\chi_{>R} \frac{\mathfrak{h}^{rt}}{1 + \mathfrak{h}^{rr}} \right)^2 \langle dr, dr \rangle.$$

Choosing R sufficiently large so that $\chi_{>R} \frac{\mathfrak{h}^{rt}}{1 + \mathfrak{h}^{rr}}$ is sufficiently small, the sign of $\mathfrak{g}^{TT}$ is the same as the sign of $\mathfrak{g}^{tt}$. The signature of the metric does not change under the change of coordinates, so the $t = \text{constant}$ submanifolds remain positive definite.

To establish Assumption 3 we need only calculate $\mathfrak{g}^{T\gamma}$ for $\gamma \in \{T, r, \theta, \phi\}$ since r, θ , and ϕ are unchanged. Direct calculation yields

$$\mathfrak{g}^{TT} = -1 + \mathfrak{f}^{tt} - \underbrace{2\chi_{>R} \frac{\mathfrak{h}^{rt}}{1 + \mathfrak{h}^{rr}} (\mathfrak{f}^{tr} + \mathfrak{h}^{tr}) + \left(\chi_{>R} \frac{\mathfrak{h}^{rt}}{1 + \mathfrak{h}^{rr}} \right)^2 (1 + \mathfrak{f}^{rr} + \mathfrak{h}^{rr})}_{\in \ell^1 S(r^{-\kappa})} + \mathfrak{h}^{tt}$$

$$\mathfrak{g}^{Tr} = \mathfrak{f}^{tr} - \chi_{>R} \frac{\mathfrak{h}^{rt}}{1 + \mathfrak{h}^{rr}} \mathfrak{f}^{rr} + \chi_{<R} \mathfrak{h}^{tr} \in \ell^1 S(r^{-\kappa})$$

$$r \mathfrak{g}^{T\theta} = \mathfrak{f}^{t\theta} - \chi_{>R} \frac{\mathfrak{h}^{rt}}{1 + \mathfrak{h}^{rr}} \mathfrak{f}^{r\theta} \in \ell^1 S(r^{-\kappa})$$

$$r \sin \theta \mathfrak{g}^{T\phi} = \mathfrak{f}^{t\phi} - \chi_{>R} \frac{\mathfrak{h}^{rt}}{1 + \mathfrak{h}^{rr}} \mathfrak{f}^{r\phi} \in \ell^1 S(r^{-\kappa}).$$

Thus after relabeling, $\mathfrak{g}$ can be written as in (2.4) with $\mathfrak{h}^{tr} = 0$, as desired.

Next we achieve $\mathfrak{h}^{rr} = -\mathfrak{h}^{tt}$ via the coordinate change

$$d\rho = \left(1 + \chi_{>R} \frac{-\mathfrak{h}^{tt} - \mathfrak{h}^{rr}}{1 + \mathfrak{h}^{rr}} \right)^{\frac{1}{2}} dr$$

where R is a constant chosen so $1 + \mathfrak{h}^{rr} \gtrsim 1$ for $r > R$. Note we have $\chi_{>R} \frac{-\mathfrak{h}^{tt} - \mathfrak{h}^{rr}}{1 + \mathfrak{h}^{rr}} \in S_{rad}(r^{-\kappa})$ and $\left(1 + \chi_{>R} \frac{-\mathfrak{h}^{tt} - \mathfrak{h}^{rr}}{1 + \mathfrak{h}^{rr}}\right)^{\frac{1}{2}} \in S_{rad}(1)$. The $t = \text{constant}$ subspaces are invariant under the change of coordinates and thus remain positive definite and the metric coefficients remain independent of t . It follows that Assumptions 1 and 2 still hold.

We now calculate $\mathfrak{g}^{\gamma\delta}$ in the new coordinate system. Since t, θ , and ϕ are unchanged, we need only calculate $\mathfrak{g}^{\rho\gamma}$ for $\gamma \in \{t, \rho, \theta, \phi\}$. Direct calculation yields

$$\begin{aligned} \mathfrak{g}^{\rho\rho} &= 1 + \mathfrak{f}^{rr} + \underbrace{\chi_{<R}(\mathfrak{h}^{rr} + \mathfrak{h}^{tt}) + \chi_{>R} \frac{-\mathfrak{h}^{tt} - \mathfrak{h}^{rr}}{1 + \mathfrak{h}^{rr}} \mathfrak{f}^{rr}}_{\in \ell^1 S(r^{-\kappa})} - \mathfrak{h}^{tt} \\ \mathfrak{g}^{\rho t} &= \left(1 + \chi_{>R} \frac{-\mathfrak{h}^{tt} - \mathfrak{h}^{rr}}{1 + \mathfrak{h}^{rr}}\right)^{\frac{1}{2}} \mathfrak{f}^{rt} \in \ell^1 S(r^{-\kappa}) \\ r \mathfrak{g}^{\rho\theta} &= \left(1 + \chi_{>R} \frac{-\mathfrak{h}^{tt} - \mathfrak{h}^{rr}}{1 + \mathfrak{h}^{rr}}\right)^{\frac{1}{2}} \mathfrak{f}^{r\theta} \in \ell^1 S(r^{-\kappa}) \\ r \sin \theta \mathfrak{g}^{\rho\phi} &= \left(1 + \chi_{>R} \frac{-\mathfrak{h}^{tt} - \mathfrak{h}^{rr}}{1 + \mathfrak{h}^{rr}}\right)^{\frac{1}{2}} \mathfrak{f}^{r\phi} \in \ell^1 S(r^{-\kappa}). \end{aligned}$$

After relabeling, $\mathfrak{g}$ can now be written as in (2.4) with the additional assumption $\mathfrak{h}^{rr} = -\mathfrak{h}^{tt}$, as desired. $\square$

Proposition 2.2. *In normalized coordinates (established in Lemma 2.1), the operator $\square_{\mathfrak{g}} + V$ can be replaced by an L^2 self-adjoint operator, which can be written as in (2.1) where (2.2) and (2.3) hold.*

Proof. The result of the proposition is obtained via direct calculation. We outline the key steps of the process. Converting from spherical to rectangular coordinates, the normalized metric $\mathfrak{g}$ can be written as

$$[\mathfrak{g}^{\alpha\beta}] = [\mathfrak{m}^{\alpha\beta}] + [\mathfrak{f}^{\alpha\beta}] + [\mathfrak{h}^{\alpha\beta}]$$

where

$$[\mathfrak{m}^{\alpha\beta}] = \begin{bmatrix} -1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{bmatrix}, \quad [\mathfrak{f}^{\alpha\beta}] = \begin{bmatrix} \mathfrak{f}^{tt} & \mathfrak{f}^{t1} & \mathfrak{f}^{t2} & \mathfrak{f}^{t3} \\ \mathfrak{f}^{t1} & \mathfrak{f}^{11} & \mathfrak{f}^{12} & \mathfrak{f}^{13} \\ \mathfrak{f}^{t2} & \mathfrak{f}^{12} & \mathfrak{f}^{22} & \mathfrak{f}^{23} \\ \mathfrak{f}^{t3} & \mathfrak{f}^{13} & \mathfrak{f}^{23} & \mathfrak{f}^{33} \end{bmatrix},$$

and

$$[\mathfrak{h}^{\alpha\beta}] = \begin{bmatrix} \mathfrak{h}^{tt} & 0 & 0 & 0 \\ 0 & \mathfrak{h}^{\omega\omega} & 0 & 0 \\ 0 & 0 & \mathfrak{h}^{\omega\omega} & 0 \\ 0 & 0 & 0 & \mathfrak{h}^{\omega\omega} \end{bmatrix} - \frac{\mathfrak{h}^{tt} + \mathfrak{h}^{\omega\omega}}{r^2} \begin{bmatrix} 0 & 0 & 0 & 0 \\ 0 & x_1^2 & x_1 x_2 & x_1 x_3 \\ 0 & x_1 x_2 & x_2^2 & x_2 x_3 \\ 0 & x_1 x_3 & x_2 x_3 & x_3^2 \end{bmatrix}$$

with $\mathfrak{f}^{\alpha\beta} \in \ell^1 S(r^{-\kappa})$ and $\mathfrak{h}^{\alpha\beta} \in S_{rad}(r^{-\kappa})$.

To make the operator self-adjoint, we conjugate by $|\mathfrak{g}|^{\frac{1}{4}}$ where $|\mathfrak{g}| = |\det(\mathfrak{g})|$:

$$\square_{\mathfrak{g}} + V \rightarrow |\mathfrak{g}|^{\frac{1}{4}} (\square_{\mathfrak{g}} + V) |\mathfrak{g}|^{-\frac{1}{4}}$$

The -1 coefficient of ∂_t^2 in (2.1) is achieved through multiplication by $(-\mathfrak{g}^{tt})^{-1}$, but we split this factor into multiplication by $(-\mathfrak{g}^{tt})^{-1/2}$ on the left and right so the operator remains self-adjoint. Thus we will replace $\square_{\mathfrak{g}} + V$ by

$$P = |\mathfrak{g}|^{1/4} (-\mathfrak{g}^{tt})^{-1/2} (\square_{\mathfrak{g}} + V) (-\mathfrak{g}^{tt})^{-1/2} |\mathfrak{g}|^{-1/4}.$$

Commuting and writing the operator in divergence form we find

$$P = \partial_{\alpha} A^2 B^{\alpha\beta} \partial_{\beta} + A(\partial_{\alpha} B^{\alpha\beta})(\partial_{\beta} A) + AB^{\alpha\beta}(\partial_{\alpha\beta} A) + (\mathfrak{g}^{tt})^{-1} V \quad (2.5)$$

where

$$A := (-\mathfrak{g}^{tt})^{-1/2} |\mathfrak{g}|^{-1/4} \quad \text{and} \quad B^{\alpha\beta} := |\mathfrak{g}|^{\frac{1}{2}} \mathfrak{g}^{\alpha\beta}.$$

Since $|\mathfrak{g}| - 1, \mathfrak{g}^{tt} + 1 \in \ell^1 S(r^{-\kappa}) + S_{rad}(r^{-\kappa})$, the scalar terms in (2.5) are of the form $V_{\ell} + V_r$ as in (2.3).

Consider the term $\partial_{\alpha} A^2 B^{\alpha\beta} \partial_{\beta} = \partial_{\alpha} (-\mathfrak{g}^{tt})^{-1} \mathfrak{g}^{\alpha\beta} \partial_{\beta}$ in (2.5). When $\alpha = \beta = t$ we find

$$\partial_t (-\mathfrak{g}^{tt})^{-1} \mathfrak{g}^{tt} \partial_t = -\partial_t^2. \quad (2.6)$$

When either $\alpha = t$ or $\beta = t$ (but not both), the desired form, $\partial_t(\partial_i p_1^i + p_1^i \partial_i)$ with $p_1^i \in \ell^1 S(r^{-\kappa})$, follows from the observation

$$(-\mathfrak{g}^{tt})^{-1} \mathfrak{g}^{ti} = \frac{\mathfrak{f}^{ti}}{1 - \mathfrak{f}^{tt} - \mathfrak{h}^{tt}} \in \ell^1 S(r^{-\kappa}), \quad i \in \{1, 2, 3\}.$$

When $\alpha, \beta \in \{1, 2, 3\}$ we use i, j instead of α, β since we are only considering spatial terms. If $i = j$ then

$$(-\mathfrak{g}^{tt})^{-1} \mathfrak{g}^{ii} = 1 + p_2^{ii} + (\mathfrak{h}^{tt} + \mathfrak{h}^{\omega\omega})(1 - x_i^2 r^{-2}) \quad (2.7)$$

where

$$p_2^{ii} := \frac{\mathfrak{f}^{ii} + \mathfrak{f}^{tt} + (\mathfrak{h}^{tt} + \mathfrak{h}^{\omega\omega})(1 - x_i^2 r^{-2})(\mathfrak{f}^{tt} + \mathfrak{h}^{tt})}{1 - \mathfrak{f}^{tt} - \mathfrak{h}^{tt}} \in \ell^1 S(r^{-\kappa}).$$

If $i \neq j$ then we find

$$(-\mathfrak{g}^{tt})^{-1} \mathfrak{g}^{ij} = p_2^{ij} - (\mathfrak{h}^{tt} + \mathfrak{h}^{\omega\omega}) x_i x_j r^{-2}, \quad i \neq j \quad (2.8)$$

where

$$p_2^{ij} := \frac{\mathfrak{f}^{ij} - (\mathfrak{f}^{tt} + \mathfrak{h}^{tt})(\mathfrak{h}^{tt} + \mathfrak{h}^{\omega\omega}) x_i x_j r^{-2}}{1 - \mathfrak{f}^{tt} - \mathfrak{h}^{tt}} \in \ell^1 S(r^{-\kappa}).$$

Combining (2.7) and (2.8) yields

$$(-\mathfrak{g}^{tt})^{-1} \mathfrak{g}^{ij} = \delta_{ij} + p_2^{ij} + (\mathfrak{h}^{tt} + \mathfrak{h}^{\omega\omega})(\delta_{ij} - x_i x_j r^{-2})$$

and we find

$$\partial_i (-\mathfrak{g}^{tt})^{-1} \mathfrak{g}^{ij} \partial_j = \Delta + \partial_i p_2^{ij} \partial_j + p_2^{\omega} \Delta_{\omega}$$

where $p_2^{\omega} = (\mathfrak{h}^{tt} + \mathfrak{h}^{\omega\omega}) r^{-2} \in S_{rad}(r^{-\kappa-2})$ and p_2^{ij} are as above. This concludes the proof of the proposition. $\square$

Throughout the rest of the paper we take P to be the operator established in Proposition 2.2

3. PRELIMINARY RESULTS ON THE RESOLVENT

The results in this section are not sensitive to the rate at which the background geometry tends toward flat. Instead, they rely on the energy assumptions in the statement of the main theorem. The results of section 2 mean we are interested in the Cauchy problem

$$Pu = f, \quad u(0, \cdot) = u_0, \quad \partial_t u(0, \cdot) = u_1. \quad (3.1)$$

The evolution satisfies the uniform energy bounds (1.4) and the weak local energy estimate (1.6) since these estimates are coordinate independent. We define the resolvent as follows:

Definition 3.1. *The operator P_τ associated to P is given by $\partial_t \mapsto i\tau$ so that*

$$P_\tau := \tau^2 + \Delta + i\tau P^1 + P^2.$$

Definition 3.2. *The resolvent associated to P , denoted R_τ , is defined by*

$$R_\tau := P_\tau^{-1}$$

when it exists.

The following proposition shows that the uniform energy assumption guarantees that R_τ exists when $\Im\tau < 0$ and links the time Fourier transform of a solution u to the initial data. It is this link that we will exploit in order to establish the final pointwise decay. The results of Proposition 3.3 are established in [33] section 3, but we include a proof since (3.2) is a fundamental piece of our argument.

Proposition 3.3. *Assume (3.1) satisfies the uniform energy bounds (1.4). If $\Im\tau < 0$, the operator $P_\tau : H^2 \rightarrow L^2$ is one-to-one and the range of P_τ is dense in L^2 . Furthermore, if u satisfies (3.1) then for $\Im\tau < 0$ we have*

$$\hat{u}(\tau, x) = R_\tau(\hat{f}(\tau) - i\tau u_0 + P^1 u_0 - u_1). \quad (3.2)$$

Proof. Let Q_τ be a family of τ dependent operators which are defined by $Q_\tau g = \hat{u}(\tau)$ where $u(t, x)$ solves the homogeneous Cauchy problem

$$Pu = 0, \quad u(0) = 0, \quad \partial_t u(0) = -g \in L^2.$$

We will show for $\Im\tau < 0$ that $Q_\tau P_\tau g = P_\tau Q_\tau g = g$ (i.e. P_τ is invertible and $R_\tau = Q_\tau$).

First we establish L^2 based bounds on $Q_\tau g$. By assumption, the evolution (3.1) satisfies the uniform energy bounds (1.4), which translate into L^2 based bounds for $Q_\tau g$. Using the notation $\|\phi\|_{\dot{H}^{N,1}} = \|\nabla \phi\|_{H^N}$ and setting $u(t, x) \equiv 0$ for $t < 0$ we use the Minkowski Integral Inequality and (1.4) to find for any $N \geq 0$

$$\|Q_\tau g\|_{\dot{H}^{N,1}} \leq \sum_{j \leq N} \int_0^\infty \left(\int_{\mathbb{R}^3} |e^{-it\tau} \nabla^{j+1} u(t, x)|^2 dx \right)^{1/2} dt \lesssim \frac{1}{|\Im\tau|} \|g\|_{H^N}.$$

Similarly we calculate

$$|\tau| \|Q_\tau g\|_{H^N} \leq \int_0^\infty e^{t\Im\tau} \|\partial_t u(t, \cdot)\|_{H^N(\mathbb{R}^3)} dt \lesssim \frac{1}{|\Im\tau|} \|g\|_{H^N}.$$

Therefore if $g \in H^N$ then $Q_\tau g \in H^{N+1}$ for $\Im\tau < 0$.

In general taking the time Fourier transform of Pu (again setting $u(t, x) \equiv 0$ for $t < 0$) and integrating by parts yields

$$0 = \int e^{-it\tau} Pu \, dt = P_\tau \hat{u}(\tau) + \partial_t u(0) + i\tau u(0) - P^1 u(0)$$

so that

$$P_\tau \hat{u}(\tau) = (Pu)^\wedge - i\tau u(0) + P^1 u(0) - \partial_t u(0). \quad (3.3)$$

Given $g \in H^1$, we have $Q_\tau g \in H^2$, so $Q_\tau g$ is in the domain of P_τ . Applying (3.3) to our definition of $Q_\tau g$, we find $P_\tau Q_\tau g = g$. Thus H^1 is contained in the range of $P_\tau : H^2 \rightarrow L^2$, so the range is dense in L^2 .

Next we aim to show $Q_\tau P_\tau g = g$. To this end, we claim that if $u(t, x)$ solves the nonhomogeneous Cauchy problem

$$Pu = f, \quad u(0, \cdot) = u_0, \quad \partial_t u(0, \cdot) = u_1$$

then

$$\hat{u}(\tau) = Q_\tau(\hat{f}(\tau) - i\tau u_0 + P^1 u_0 - u_1). \quad (3.4)$$

Once (3.4) is established, we can show $Q_\tau P_\tau g = g$. Indeed, assume (3.4) holds and let $g(x) \in H^2$ be given. We define $u(t, x) := g(x) \mathbf{1}_{t \geq 0}$, where $\mathbf{1}_{t \geq 0}$ is an indicator function that is 1 for $t \geq 0$ and 0 otherwise. Taking the time Fourier transform of $u(t, x)$ yields $\hat{u}(\tau) = \frac{1}{i\tau} g$. By (3.3) we have

$$\frac{1}{i\tau} P_\tau g = (Pu)^\wedge - i\tau u(0) + P^1 u(0) - \partial_t u(0). \quad (3.5)$$

Then applying Q_τ to (3.5) and using (3.4) gives

$$Q_\tau P_\tau g = i\tau Q_\tau \left((Pu)^\wedge - i\tau u(0) + P^1 u(0) - \partial_t u(0) \right) = g,$$

as desired.

It is left to show (3.4). To do this we use Duhamel's formula and find

$$u(t, x) = \int_0^t u_a(t-s, x; s) \, ds + \partial_t u_b + u_c + u_d, \quad (3.6)$$

where $u_1(t, x; s)$, $u_2(t, x)$, $u_3(t, x)$, and $u_4(t, x)$ solve the following:

$$\begin{aligned} Pu_a &= 0 & u_a(0, x; s) &= 0 & \partial_t u_a(0, x; s) &= -f(s, x) \\ Pu_b &= 0 & u_b(0, x) &= 0 & \partial_t u_b(0, x) &= u_0 \\ Pu_c &= 0 & u_c(0, x) &= 0 & \partial_t u_c(0, x) &= -P^1 u_0 \\ Pu_d &= 0 & u_d(0, x) &= 0 & \partial_t u_d(0, x) &= u_1. \end{aligned}$$

Note to find $\partial_t u(0, t) = u_1$, we use $Pu_b = 0$ to write $\partial_t^2 u_b = (\Delta + \partial_t P^1 + P^2)u_b$. We calculate $\hat{u}(\tau)$ by taking the time Fourier transform of each term in (3.6). For the first term we switch the order of integration, change variables by $t \mapsto t + s$, then switch back the order of integration to find

$$\int_0^\infty e^{-it\tau} \int_0^t u_a(t-s, x; s) \, ds dt = \int_0^\infty e^{-it\tau} \beta(t, x; \tau) dt$$

where $\beta(t, x; \tau) = \int_0^\infty e^{-is\tau} u_a(t, x; s) ds$. Since $\beta(t, x; \tau)$ satisfies

$$P\beta = 0, \quad \beta(0, x; \tau) = 0, \quad \partial_t \beta(0, x; \tau) = -\hat{f}(\tau, x)$$

we have $\hat{\beta}(\tau, x; \tau) = Q_\tau \hat{f}(\tau)$. Applying the time Fourier transform to the remaining terms in (3.6) yields

$$\hat{u}(\tau) = \hat{\beta}(\tau, x; \tau) + i\tau \hat{u}_b(\tau) + \hat{u}_c(\tau) + \hat{u}_d(\tau) = Q_\tau \left(\hat{f}(\tau) - i\tau u_0 + P^1 u_0 - u_1 \right)$$

as desired. This concludes the proof of (3.4) and thus the proof of the proposition. $\square$

If the weak local energy decay estimate (1.6) also holds, then we are able to obtain L^2 -based resolvent bounds which are stronger than those established in the proof of Proposition 3.3 and are uniform as $\Im\tau \rightarrow 0$. This makes it possible to extend the resolvent continuously to the real axis. The results proving the stronger L^2 -based bounds and the continuous extension of R_τ to $\tau \in \mathbb{R}$ are established in [33] and are not affected by our change to the assumed rate at which the background geometry tends toward flat. Therefore we state the key results from [33] without proof.

The $\mathcal{LE}_\tau$ norm, in which we measure the resolvent $v = R_\tau g$, is defined by

$$\|v\|_{\mathcal{LE}_\tau^N} = \|(|\tau| + \langle r \rangle^{-1})v\|_{\mathcal{LE}^N} + \|\nabla v\|_{\mathcal{LE}^N} + \|(|\tau| + \langle r \rangle^{-1})^{-1} \nabla^2 v\|_{\mathcal{LE}^N}. \quad (3.7)$$

Proposition 3.4 (see [33, Proposition 9 and Corollary 12]). *Assume (3.1) satisfies the uniform energy bounds (1.4) and the weak local energy estimate (1.6). If $\Im\tau \leq 0$ and $g \in \mathcal{LE}^{*,N+4}$ for fixed $N \in \mathbb{N}$, then $v = R_\tau g$ satisfies*

$$\|v\|_{\mathcal{LE}_\tau^N} \lesssim \|g\|_{\mathcal{LE}^{*,N+4}}. \quad (3.8)$$

Proposition 3.5 (see [33, Proposition 10 and Corollary 12]). *Assume the Cauchy problem (3.1) satisfies the uniform energy bounds (1.4) and the local energy estimate (1.6).*

If $\Im\tau \leq 0$, and $g \in \mathcal{LE}^$ satisfies*

$$\|T^i \Omega^j S^k g\|_{\mathcal{LE}^{*,4}} \lesssim 1, \quad i + 4j + 16k < M \quad (3.9)$$

for some positive integer M , then

$$\|T^i \Omega^j S^k (R_\tau g)\|_{\mathcal{LE}_\tau} \lesssim 1, \quad i + 4j + 16k < M - 4. \quad (3.10)$$

3.1. Strategy to Obtain Decay Rate. To prove Theorem 1.5, we must prove pointwise decay rates for solutions to the homogeneous Cauchy problem (3.1) with $f = 0$. Therefore by (3.2) we have

$$u(t, x) = \frac{1}{\sqrt{2\pi}} \int_{\Im\tau = -\varepsilon} R_\tau (-i\tau u_0 + P^1 u_0 - u_1) e^{it\tau} d\tau$$

for $\varepsilon > 0$. The continuous extension of R_τ to $\tau \in \mathbb{R}$ allows us to take the limit as $\varepsilon \rightarrow 0$ to obtain

$$u(t, x) = \frac{1}{\sqrt{2\pi}} \int_{\mathbb{R}} R_\tau (-i\tau u_0 + P^1 u_0 - u_1) e^{it\tau} d\tau. \quad (3.11)$$

The pointwise decay rates are obtained by writing $e^{it\tau} = \frac{d}{d\tau} \left(\frac{e^{it\tau}}{it} \right)$ and integrating by parts. Thus the final decay rate will depend on how many times we can integrate by parts, which depends on the regularity of the resolvent.

We separate the solution u into low and high frequency parts by defining

$$u_{>1}(t, x) := \frac{1}{\sqrt{2\pi}} \int_{\mathbb{R}} \chi_{>1}(|\tau|) R_{\tau}(-i\tau u_0 + P^1 u_0 - u_1) e^{it\tau} d\tau \quad (3.12)$$

and

$$u_{<1}(t, x) := \frac{1}{\sqrt{2\pi}} \int_{\mathbb{R}} \chi_{<1}(|\tau|) R_{\tau}(-i\tau u_0 + P^1 u_0 - u_1) e^{it\tau} d\tau. \quad (3.13)$$

Recall in Theorem 1.5 we assume the initial data satisfies $u_0 \in Z^{\nu+1, \kappa}$ and $u_1 \in Z^{\nu, \kappa+1}$. Here ν is a sufficiently large constant depending on κ , and κ indicates the rate at which the background geometry tends toward flat. The assumptions on the initial data mean we can write

$$-i\tau u_0 + P^1 u_0 - u_1 = \tau g_{\kappa}^{\nu+1} + g_{\kappa+1}^{\nu}$$

for some $g_{\kappa}^{\nu+1} \in Z^{\nu+1, \kappa}$ and some $g_{\kappa+1}^{\nu} \in Z^{\nu, \kappa+1}$. In sections 4 and 5 we analyze the resolvent $R_{\tau}g$ near 0 frequency for general g in an appropriately defined function space.

4. THE ZERO RESOLVENT

In this section we obtain an expansion of R_0g in powers of $\langle r \rangle^{-1}$. In section 5 we will approximate $R_{\tau}g$ by $R_0g e^{-i\tau r}$ for small τ and calculate the error using this expansion. Note for the arguments in this and the following section we harmlessly assume $r \geq 3$, since the results of these sections hold for $r \lesssim 1$ using weak local energy decay and Sobolev embeddings.

In our argument establishing an expansion of R_0g for large r we will find

$$(-\Delta)(\chi_{>R} R_0g) = h + \chi_{>R/2} P^2(\chi_{>R} R_0g)$$

where $\|h\|_{Z^{n, \lambda}} \lesssim \|g\|_{Z^{n+4, \lambda}}$. This motivates Lemma 4.1, where we will obtain an expansion of $(-\Delta)^{-1}g$ for $g \in Z^{n, \lambda}$.

Lemma 4.1. *Let $g \in Z^{n, \lambda}$ with $\lambda, n \in \mathbb{N}$. We have the following representation for $(-\Delta)^{-1}g$:*

$$(-\Delta)^{-1}g = \sum_{j=0}^{\lambda-2} (c_j \cdot \nabla^j \langle r \rangle^{-1} + e_j(r) \cdot (\nabla^j \langle r \rangle^{-1}) \langle r \rangle^{j-\lambda+1}) + d(r) \cdot \nabla^{\lambda-1} \langle r \rangle^{-1} + q(x) \quad (4.1)$$

where the coefficients satisfy

$$\sum_{j=0}^{\lambda-2} (|c_j| + \|e_j\|_{\ell^1 S(1)}) + \|d\|_{L^\infty} + \|S_r d\|_{\ell^1 S(1)} + \|q\|_{Z^{n+2, \lambda-2}} \lesssim \|g\|_{Z^{n, \lambda}}. \quad (4.2)$$

For $\lambda = 1$ we have only the last two terms in (4.1).

Proof. Set $v = (-\Delta)^{-1}g$. We begin by proving if $\langle r \rangle^\lambda g \in \mathcal{LE}^*$, then v can be expressed as in (4.1) where the following estimate holds

$$\sum_{j=0}^{\lambda-2} |c_j| + \|e_j\|_{\ell^1 S(1)} + \|d\|_{L^\infty} + \|S_r d\|_{\ell^1 S(1)} + \sum_{i \leq 2} \|\langle r \rangle^{-2+\lambda+i} \nabla^i q\|_{\mathcal{LE}^*} \lesssim \|\langle r \rangle^\lambda g\|_{\mathcal{LE}^*}. \quad (4.3)$$

We will then consider the case $g \in Z^{n, \lambda}$ and prove (4.2) using (4.3) and elliptic regularity arguments.

We have $v(x) = \frac{1}{4\pi} \int g(y) \frac{1}{|x-y|} dy$ using the kernel for the fundamental solution of the Laplacian in $\mathbb{R}^3$. We wish to bound the coefficients of the representation for v by the size of $\langle r \rangle^\lambda g$ measured in $\mathcal{LE}^*$, which is defined by the behavior of g on dyadic regions A_m . This motivates the following decomposition of g :

$$g_m := \beta_{\approx m}(r)g \quad \text{and} \quad v_m := (-\Delta)^{-1}g_m = \int g_m(y) \frac{1}{|x-y|} dy.$$

We further decompose each v_m into v_m^{low} and v_m^{high} as follows

$$v_m^{low} := \chi_{<m+2}(r)v_m = \chi_{<m+2}(|x|) \int \beta_{\approx m}(|y|)g(y) \frac{1}{|x-y|} dy \quad (4.4)$$

and

$$v_m^{high} := \chi_{>m+2}(r)v_m = \chi_{>m+2}(|x|) \int \beta_{\approx m}(|y|)g(y) \frac{1}{|x-y|} dy. \quad (4.5)$$

Thus we have $v = \frac{1}{4\pi} \sum_{m \geq 0} v_m^{low} + v_m^{high}$.

We claim

$$\sum_{i \leq 2} \left\| \langle r \rangle^{-2+\lambda+i} \nabla^i \sum_{m \geq 0} v_m^{low} \right\|_{\mathcal{LE}^*} \lesssim \left\| \langle r \rangle^\lambda g \right\|_{\mathcal{LE}^*} \quad (4.6)$$

so that $\sum_{m \geq 0} v_m^{low}$ can be included in the $q(x)$ term in (4.3) (we will see that $\sum_{m \geq 0} v_m^{high}$ also generates a term that will be included in $q(x)$). To prove (4.6) we begin by calculating

$$\begin{aligned} |(-\Delta)v_m^{low}| &\lesssim |g_m| + 2^{-m} |\chi'_{<m+2}(|x|)| \int \left| \beta_{\approx m}(|y|)g(y) \right| \frac{1}{|x-y|^2} dy \\ &\quad + 2^{-2m} |\chi''_{<m+2}(|x|)| \int \left| \beta_{\approx m}(|y|)g(y) \right| \frac{1}{|x-y|} dy \\ &\lesssim |g_m| + 2^{-\frac{3m}{2}} |\chi'_{<m+2}(|x|)| \|\beta_{\approx m}g\|_{L^2} + 2^{-\frac{3m}{2}} |\chi''_{<m+2}(|x|)| \|\beta_{\approx m}g\|_{L^2}. \end{aligned}$$

The second inequality is obtained using the Cauchy-Schwarz inequality and the fact that the cutoffs in x and y give $|x-y|^{-1} \lesssim 2^{-m}$. Integration by parts then yields $\|\nabla^2 v_m^{low}\|_{L^2} \lesssim \|g_m\|_{L^2}$ (there are no boundary terms since v_m^{low} is compactly supported). Next we use the Hardy inequality and the fact that v_m^{low} is supported on a bounded region to find

$$\|\nabla^i v_m^{low}\|_{L^2} \lesssim 2^{m(2-i)} \|g\|_{L^2(A_m)}, \quad i = 0, 1, 2. \quad (4.7)$$

Note if $|x| \approx 2^k$ and $|y| \approx 2^m$ with $k \leq m-1$, we have $|\nabla^i|x-y|^{-1}| \lesssim 2^{-m(1+i)}$ for $i = 0, 1, 2$ and $\|\nabla^i v_m^{low}\|_{L^2(A_k)} = \|\nabla^i v_m\|_{L^2(A_k)}$. Using the Cauchy-Schwarz inequality, we find

$$\|\nabla^i v_m\|_{L^2(A_k)} \lesssim 2^{\frac{3k}{2}} 2^{m(\frac{1}{2}-i)} \|g\|_{L^2(A_m)}, \quad i = 0, 1, 2 \quad (4.8)$$

when $k \leq m-1$. Now we use (4.7) and (4.8) to calculate

$$\begin{aligned} \|\langle r \rangle^{-2+\lambda+i} \nabla^i \sum_m v_m^{low} \|_{\mathcal{LE}^*} &\lesssim \sum_m \sum_{k < m+2} 2^{\frac{k}{2}} 2^{k(-2+\lambda+i)} \|\nabla^i v_m^{low}\|_{L^2(A_k)} \\ &\lesssim \sum_m \sum_{k < m} 2^{k(\lambda+i)} 2^{m(\frac{1}{2}-i)} \|g\|_{L^2(A_m)} \\ &\quad + \sum_m \sum_{k \approx m} 2^{m(-\frac{3}{2}+\lambda+i)} \|\nabla^i v_m^{low}\|_{L^2(A_{\approx m})} \\ &\lesssim \|\langle r \rangle^\lambda g\|_{\mathcal{LE}^*} \end{aligned}$$

for $i = 0, 1, 2$. This concludes the proof of (4.6).

Next we turn our attention to $\sum_m v_m^{high}$. For each m we integrate over $|y| \approx 2^m$ with $|x| \geq 2^{m+2}$, so $\frac{1}{|x-y|}$ is smooth in the region of integration and Taylor's theorem applies. We define y^j as follows. If the n^{th} component of ∇^j is $\partial_{i_1} \partial_{i_2} \cdots \partial_{i_j}$ then the n^{th} component of y^j is $y_{i_1} y_{i_2} \cdots y_{i_j}$. Note that $|y^j| \lesssim |y|^j$. In this notation, Taylor's theorem yields

$$\frac{1}{|x-y|} = \sum_{j=0}^{\lambda-1} \frac{\nabla^j |x|^{-1} \cdot (y)^j}{j!} + R_\lambda^x(y)$$

where

$$R_\lambda^x(y) = \lambda!^{-1} \nabla^\lambda (|x - ty|^{-1}) \cdot (y)^\lambda$$

for some $t \in (0, 1)$. Therefore

$$v_m^{high}(x) = \sum_{j=0}^{\lambda-1} j!^{-1} \chi_{>m+2}(|x|) \int g_m(y) (\nabla^j |x|^{-1}) \cdot y^j dy + \chi_{>m+2}(|x|) \int g_m(y) R_\lambda^x(y) dy. \quad (4.9)$$

We claim the last term in (4.9) can be included in q after summing over m . In other words, we wish to show

$$\sum_{i \leq 2} \|\langle r \rangle^{-2+\lambda+i} \nabla^i \tilde{q}\|_{\mathcal{LE}^*} \lesssim \|\langle r \rangle^\lambda g\|_{\mathcal{LE}^*}, \quad (4.10)$$

where $\tilde{q}$ is defined by

$$\tilde{q} := \sum_m \tilde{q}_m, \quad \text{and} \quad \tilde{q}_m := \chi_{>m+2}(r) \int g_m(y) R_\lambda^x(y) dy.$$

Note $|\nabla^\lambda(|x|^{-1})| \lesssim |x|^{-\lambda-1}$ and $|x - ty|^{-1} \lesssim |x|^{-1}$ since $|x| \geq 2|y|$ and $t \in (0, 1)$. It follows that

$$|R_\lambda^y(x)| \lesssim \frac{|y|^\lambda}{|x|^{\lambda+1}}, \quad (4.11)$$

which implies for $l > m+2$

$$\|\tilde{q}_m\|_{L^\infty(A_l)} \lesssim 2^{-l(\lambda+1)} \|r^\lambda g_m\|_{L^1} \lesssim 2^{(m-l)(\lambda+1)} \|\langle r \rangle^{\frac{1}{2}} g_m\|_{L^2}. \quad (4.12)$$

Straightforward calculation yields

$$\|\langle r \rangle^{-2+\lambda} \tilde{q}\|_{\mathcal{LE}^*} \leq \sum_m \sum_{l > m+2} \|\tilde{q}_m\|_{L^\infty(A_l)} 2^{l\lambda}. \quad (4.13)$$

Combining (4.12) and (4.13) then gives $\|\langle r \rangle^{-2+\lambda} \tilde{q}(x)\|_{\mathcal{LE}^*} \lesssim \|g\|_{\mathcal{LE}^*}$, as desired. The estimates for $\nabla^i \tilde{q}$ for $i = 1, 2$ follow analogously once we note $|\nabla^i R_\lambda^y(x)| \lesssim \frac{|y|^\lambda}{|x|^{\lambda+i+1}}$.

To handle the first term in (4.9) we define

$$c_j := \sum_m (j!)^{-1} \int g_m(y) y^j dy, \quad e_j(r) := \sum_m (j!)^{-1} \langle r \rangle^{\lambda-1-j} (-\chi_{<m+2}) \int g_m(y) y^j dy,$$

$$\text{and } d(r) := \sum_m ((\lambda-1)!)^{-1} \chi_{>m+2}(r) \int g_m(y) y^{\lambda-1} dy,$$

so

$$\begin{aligned} & \sum_m \sum_{j=0}^{\lambda-1} (j!)^{-1} \chi_{>m+2}(|x|) \int g_m(y) (\nabla^j |x|^{-1}) y^j dy \\ &= \sum_{j=0}^{\lambda-2} \left(c_j \cdot \nabla^j \langle r \rangle^{-1} + e_j(r) \cdot (\nabla^j \langle r \rangle^{-1}) \langle r \rangle^{j-\lambda+1} \right) + d(r) \cdot \nabla^{\lambda-1} \langle r \rangle^{-1}. \end{aligned}$$

The desired bounds (4.3) then follow directly using the inequality $\|\langle r \rangle^p g\|_{L^1(A_m)} \lesssim \|\langle r \rangle^{p+\frac{3}{2}} g\|_{L^2(A_m)}$, which is a straightforward application of Cauchy-Schwarz.

Now let $g \in Z^{n,\lambda}$. Since $\|\langle r \rangle^\lambda g\|_{\mathcal{LE}^*} \lesssim \|g\|_{Z^{n,\lambda}}$, the above argument shows that v admits a representation as in (4.1) such that (4.3) holds. To prove (4.2) we need to show the q term now satisfies

$$\|q\|_{Z^{n+2,\lambda-2}} \lesssim \|g\|_{Z^{n,\lambda}}. \quad (4.14)$$

Observe that for any ϕ with sufficient differentiability we have

$$|T^i \Omega^j S_r^k \phi| \lesssim \sum_{l=0}^{i+j+k} |\langle r \rangle^l \nabla^l \phi|. \quad (4.15)$$

By (4.3) and (4.15) we find

$$\|q\|_{Z^{2,\lambda-2}} \lesssim \|g\|_{\mathcal{LE}^*} = \|g\|_{Z^{0,\lambda}}. \quad (4.16)$$

Furthermore we can use the definition of $Z^{n,\lambda}$ and (4.15) to establish

$$\|q\|_{Z^{n+2,\lambda-2}} \lesssim \|q\|_{Z^{2,\lambda-2}} + \|\nabla^2 q\|_{Z^{n,\lambda}}. \quad (4.17)$$

Next we calculate

$$\begin{aligned}
\|\nabla^2 q\|_{Z^{n,\lambda}} &\lesssim \sup_{i+j+k \leq n} \sum_m 2^{\frac{m}{2}} 2^{m\lambda} \|T^i \Omega^j S_r^k \nabla^2(\chi_{\approx m} q)\|_{L^2(\mathbb{R}^3)} \\
&\approx \sup_{i+j+k \leq n} \sum_m 2^{\frac{m}{2}} 2^{m\lambda} \|T^i \Omega^j S_r^k \Delta(\chi_{\approx m} q)\|_{L^2(\mathbb{R}^3)} \\
&\lesssim \sup_{i+j+k \leq n} \sum_m 2^{m(\lambda+\frac{1}{2})} \left(\|T^i \Omega^j S_r^k \chi_{\approx m} \Delta q\|_{L^2(\mathbb{R}^3)} + \|T^i \Omega^j S_r^k (\Delta \chi_{\approx m}) q\|_{L^2(\mathbb{R}^3)} \right. \\
&\quad \left. + \|T^i \Omega^j S_r^k (\nabla \chi_{\approx m}) \cdot \nabla q\|_{L^2(\mathbb{R}^3)} \right) \\
&\lesssim \sup_{i+j+k \leq n} \sum_m 2^{m(\lambda+\frac{1}{2})} \left(\|T^i \Omega^j S_r^k \chi_{\approx m} \Delta q\|_{L^2(\mathbb{R}^3)} + \|2^{-2m} T^i \Omega^j S_r^k \chi'' \left(\frac{r}{2^m} \right) q\|_{L^2(\mathbb{R}^3)} \right. \\
&\quad \left. + 2^{-m} \|T^i \Omega^j S_r^k r^{-1} \chi' \left(\frac{r}{2^m} \right) q\|_{L^2(\mathbb{R}^3)} + 2^{-m} \|T^i \Omega^j S_r^k \chi' \left(\frac{r}{2^m} \right) \nabla q\|_{L^2(\mathbb{R}^3)} \right) \\
&\lesssim \|\Delta q\|_{Z^{n,\lambda}} + \|\nabla q\|_{Z^{n,\lambda-1}} + \|q\|_{Z^{n,\lambda-2}}
\end{aligned} \tag{4.18}$$

for all $n \geq 1$. The second line in (4.18) follows from

$$\|T^i \Omega^j S_r^k \nabla^2(\chi_{\approx m} q)\|_{L^2(\mathbb{R}^3)} \approx \|T^{\leq i} \Omega^{\leq j} S_r^{\leq k} \Delta(\chi_{\approx m} q)\|_{L^2(\mathbb{R}^3)},$$

which is obtained via integration by parts using the commutators

$$[\partial_i, \Omega], [\partial_i, S_r] \in \text{span}\{T\}, \quad [\Delta, \Omega] = [\Delta, T] = 0, \quad \text{and} \quad [\Delta, S_r] = 2\Delta.$$

We claim

$$\|\Delta q\|_{Z^{n,\lambda}} \lesssim \|g\|_{Z^{n,\lambda}}. \tag{4.19}$$

Once (4.19) is established, (4.14) follows by induction in n . Indeed, assume (4.19) holds. When $n = 1$, the RHS of (4.18) is controlled by $\|g\|_{Z^{1,\lambda}}$ using (4.19) to bound the first term and (4.16) to bound the last two terms (along with the fact $\|\nabla \phi\|_{Z^{n,\lambda}} \lesssim \|\phi\|_{Z^{n+1,\lambda-1}}$ for $\phi \in Z^{n+1,\lambda-1}$). Thus $\|\nabla^2 q\|_{Z^{1,\lambda}} \lesssim \|g\|_{Z^{1,\lambda}}$, so by (4.16) and (4.17) we obtain (4.14) in the case $n = 1$. The inductive step is proved in the same manner, with the inductive hypothesis being used to bound the last two terms on the RHS of (4.18).

To prove (4.19) we write

$$-\Delta q = g + \Delta(v - q). \tag{4.20}$$

By (4.1) we have

$$-\Delta(v - q) = -\Delta \left(\sum_{j=0}^{\lambda-2} \left(c_j \cdot \nabla^j \langle r \rangle^{-1} + e_j(r) \cdot (\nabla^j \langle r \rangle^{-1}) \langle r \rangle^{j-\lambda+1} \right) + d(r) \cdot \nabla^{\lambda-1} \langle r \rangle^{-1} \right).$$

Direct calculation using (4.3) yields

$$\begin{aligned}
\|\Delta(c_j \nabla^j \langle r \rangle^{-1})\|_{Z^{n,\lambda}} &\lesssim \|\langle r \rangle^\lambda g\|_{\mathcal{LE}^*} \\
\|\Delta(e_j(\nabla^j \langle r \rangle^{-1}) \langle r \rangle^{j-\lambda+1})\|_{Z^{n,\lambda}} &\lesssim \|\langle r \rangle^\lambda g\|_{\mathcal{LE}^*} \\
\|\Delta(d \nabla^{\lambda-1} \langle r \rangle^{-1})\|_{Z^{n,\lambda}} &\lesssim \|\langle r \rangle^\lambda g\|_{\mathcal{LE}^*},
\end{aligned}$$

so that $\|\Delta(v - q)\|_{Z^{n,\lambda}} \lesssim \|g\|_{Z^{0,\lambda}}$, which combined with (4.20) gives (4.19). This concludes the proof of (4.14) and thus the proof of the proposition. $\square$

In Propositions [4.3](#) and [5.6](#) we will need expressions for $(-\Delta)^{-1}g$ where $g \in S_{rad}(r^{-q})$ with $q \geq 2$, which are stated in the following lemma.

Lemma 4.2. *Let $g \in S_{rad}(r^{-q})$ and set $v = (-\Delta)^{-1}g$*

(1) *If $q \geq 4$ then there exist a constant c and an $e \in S_{rad}(1)$ such that*

$$v = c\langle r \rangle^{-1} + e(r)\langle r \rangle^{-(q-2)}.$$

If, furthermore, we have $\text{supp } g \subset \{r \geq \frac{R}{4}\}$ for some $R > 0$, then we also have

$$|c| + R^{-\frac{1}{2}}\|e\|_{S(1)} \lesssim R^{-\frac{1}{2}}\|g\|_{S_{rad}(r^{-q})}.$$

(2) *If $q = 3$ then there exists an $\varepsilon \in S_{rad}(\ln r)$ such that*

$$v = \varepsilon(r)\langle r \rangle^{-1}.$$

(3) *If $q = 2$, then $v \in S_{rad}(\ln r)$.*

Proof. 1 and 2. $q \geq 3$

Since $-\Delta v = g$, we can write $\partial_r^2(rv) = -rg$, and integrating from infinity yields

$$|-\partial_r(rv)| = \left| \int_r^\infty \partial_s^2(sv(s)) ds \right| = \left| \int_r^\infty sg(s) ds \right| \leq \int_r^\infty \langle s \rangle^{-q+1} ds = \langle r \rangle^{-q+2}.$$

So $\partial_r(rv) \in S_{rad}(r^{-q+2})$.

When $q \geq 4$ we have $-q+2 \leq -2$ and we find

$$rv = \int_0^r \partial_s(sv(s)) ds = \int_0^\infty \int_s^\infty -\rho g(\rho) d\rho ds + \int_r^\infty \int_s^\infty \rho g(\rho) d\rho ds = c + g_1$$

where $g_1 \in S_{rad}(r^{-q+3})$. Taking $e(r) := g_1(r)\langle r \rangle^{q-3}$, we find for large r

$$v = c\langle r \rangle^{-1} + e(r)\langle r \rangle^{-(q-2)}$$

for some constant c and $e \in S_{rad}(1)$, as desired.

When $q = 3$, we have $-q+2 = -1$ so that

$$rv = \int_0^r \partial_s(sv(s)) ds = \int_0^\infty \int_s^\infty -\rho g(\rho) d\rho ds + \int_r^\infty \int_s^\infty \rho g(\rho) d\rho ds = c + g_2$$

where $g_2 \in S_{rad}(\ln r)$. Note $c \in S_{rad}(1) \subseteq S_{rad}(\ln r)$. Thus taking $\varepsilon(r) = c + g_2$ we find for large r

$$v = \varepsilon(r)\langle r \rangle^{-1}$$

where $\varepsilon \in S_{rad}(\ln r)$, as desired.

Finally we consider the case in (1) where $\text{supp } g \subseteq \{r \geq \frac{R}{4}\}$. Using the representation we already obtained and the fact $\partial_r^2(rv) = rg$, we see $\|e\|_{S(1)} \lesssim \|g\|_{S_{rad}(r^{-q})}$. Furthermore,

$$\begin{aligned} |Rv(R)| &= \left| \int_0^R \int_s^\infty \partial_r^2(rv) \, dr \, ds \right| \\ &= \left| \int_0^R \int_s^\infty rg \, dr \, ds \right| \\ &\leq \int_0^R \int_{R/4}^\infty \langle r \rangle^{-q+1} \, dr \, ds \|g\|_{S_{rad}(r^{-q})} \\ &\lesssim R^{-q+3} \|g\|_{S_{rad}(r^{-q})}. \end{aligned}$$

Therefore we have

$$|c| = |Rv(R) - e(R)R^{-q+3}| \lesssim R^{-q+3} \|g\|_{S_{rad}(r^{-q})}.$$

It follows that $v = c\langle r \rangle^{-1} + e(r)\langle r \rangle^{-\kappa-1}$ satisfies the estimate

$$|c| + R^{-\frac{1}{2}} \|e\|_{S(1)} \lesssim R^{-\frac{1}{2}} \|h\|_{S_{rad}(r^{-q})}.$$

3. $q = 2$

We write the equation $v = (-\Delta)^{-1}g$ as

$$v(r_0) = - \int_0^{r_0} s^{-2} \int_0^s r^2 g(r) \, dr \, ds.$$

where $g \in S_{rad}(r^{-2})$ by assumption. We find

$$\left| \int_0^s r^2 g(r) \, dr \right| \lesssim \int_0^s 1 \, dr = s.$$

Thus for $g_1 \in S_{rad}(r)$ we have

$$|v(r_0)| = \left| \int_0^{r_0} s^{-2} g_1(s) \, ds \right| \lesssim |\ln r_0|,$$

and $v(r) = \varepsilon(r)$ with $\varepsilon(r) \in S_{rad}(\ln r)$, as desired. $\square$

We now prove an analogue of Proposition [4.1](#) with Δ^{-1} replaced by $R_0 = (\Delta + P^2)^{-1}$. Here we first see the role of long range metric behavior - the faster the metric tends to flat, the further we can go down in the expansion for $R_0 g$.

Proposition 4.3. *Let $g \in Z^{n+4,\lambda}$ with $\lambda, n \in \mathbb{N}$. Take $v = R_0 g$.*

If $1 \leq \lambda \leq \kappa$, then for large r , v can be written as in [\(4.1\)](#) where the following estimate holds

$$\sum_{j=0}^{\lambda-2} \left(|c_j| + \|e_j\|_{\ell^1 S(1)} \right) + \|d\|_{L^\infty} + \|S_r d\|_{\ell^1 S(1)} + \|q\|_{Z^{n+2,\lambda-2}} \lesssim \|g\|_{Z^{n+4,\lambda}}. \quad (4.21)$$

If $\lambda = \kappa + 1$, then for large r , v can be written as in (4.1) where the following estimate holds

$$\sum_{j=0}^{\lambda-2} |c_j| + \|e_0\|_{S(1)} + \sum_{j=1}^{\lambda-2} \|e_j\|_{\ell^1 S(1)} + \|d\|_{L^\infty} + \|S_r d\|_{\ell^1 S(1)} + \|q\|_{Z^{n+2, \lambda-2}} \lesssim \|g\|_{Z^{n+4, \lambda}}. \quad (4.22)$$

Before proving the proposition we provide a brief summary of the argument. Since we are concerned only with large r , we consider $\chi_{>R} v =: w$ and show

$$-\Delta w = h + \chi_{>R/2} P^2 w$$

where $\|h\|_{Z^{n, \lambda}} \lesssim \|g\|_{Z^{n+4, \lambda}}$. We then use Lemma 4.1 to obtain the desired form for w . For $\lambda \leq \kappa$, this works by showing that the above equation for w is perturbative with respect to (4.2). The $\lambda = \kappa + 1$ case is similar to the case $\lambda \leq \kappa$, but there is one term which fails to be perturbative. This non-perturbative term has the benefit of being radial and will be handled using Lemma 4.2

Proof. Set $w = \chi_{>R} v$ and write

$$P_0 w = \chi_{>R} g + [P_0, \chi_{>R}] v =: h. \quad (4.23)$$

Recall that $P_0 = \Delta + P^2$ with P^2 as in (2.3). Since $\chi_{>R} \equiv 0$ for $r < R$ and $\chi_{>R} \equiv 1$ for $r > 2R$, the commutator $[P_0, \chi_{>R}]$ is supported on $[R, 2R]$. Using the notation ρ_ℓ^q to indicate a representative of the symbol class $\ell^1 S(r^q)$ (and allowing ρ_ℓ^q to stand for different functions at each appearance), direct calculation yields

$$\begin{aligned} \|\langle r \rangle^\lambda [P_0, \chi_{>R}] v\|_{\mathcal{LE}^*} &\lesssim \sum_{m=\log R}^{\log 2R} \left[\|\langle r \rangle^{\frac{1}{2}+\lambda} (R^{-2} + R^{-1} \rho_\ell^{-\kappa-1} + R^{-2} \rho_\ell^{-\kappa}) v\|_{L^2(A_m)} \right. \\ &\quad \left. + \|\langle r \rangle^{\frac{1}{2}+\lambda} (R^{-1} + R^{-1} \rho_\ell^{-\kappa}) \nabla v\|_{L^2(A_m)} \right] \\ &\lesssim_R \|\langle r \rangle^{-1} v\|_{\mathcal{LE}} + \|\nabla v\|_{\mathcal{LE}} \\ &\lesssim \|v\|_{\mathcal{LE}_0} \end{aligned} \quad (4.24)$$

where

$$\|v\|_{\mathcal{LE}_0} = \|\langle r \rangle^{-1} v\|_{\mathcal{LE}} + \|\nabla v\|_{\mathcal{LE}} + \|\langle r \rangle \nabla^2 v\|_{\mathcal{LE}}$$

by the definition of the $\mathcal{LE}_\tau$ norm in (3.7). The constants in the inequalities in (4.24) depend on R , but this is not an issue since R is fixed. Now (4.23) and (4.24) yield

$$\|\langle r \rangle^\lambda h\|_{\mathcal{LE}^*} \lesssim \|\langle r \rangle^\lambda g\|_{\mathcal{LE}^*} + \|v\|_{\mathcal{LE}_0} \lesssim \|\langle r \rangle^\lambda g\|_{\mathcal{LE}^{*,4}}$$

since $\|v\|_{\mathcal{LE}_0} \lesssim \|\langle r \rangle g\|_{\mathcal{LE}^{*,4}}$ by Proposition 3.4

Similarly we find

$$\begin{aligned} \|h\|_{Z^{n, \lambda}} &\lesssim \|g\|_{Z^{n, \lambda}} + \sup_{i+j+k \leq n} \|\langle r \rangle^\lambda T^i \Omega^j S_r^k [P_0, \chi_{>R}] v\|_{\mathcal{LE}^*} \\ &\lesssim \|g\|_{Z^{n, \lambda}} + \sup_{i+j+k \leq n} \|\langle r \rangle^\lambda [P_0, \chi_{>R}] T^i \Omega^j S_r^k v\|_{\mathcal{LE}^*} + \|\langle r \rangle^\lambda [T^i \Omega^j S_r^k, \langle r \rangle^\lambda [P_0, \chi_{>R}]] v\|_{\mathcal{LE}^*} \\ &\lesssim_R \|g\|_{Z^{n, \lambda}} + \sup_{i \leq n} \|\langle r \rangle^\lambda T^i v\|_{\mathcal{LE}^*([R, 2R])} \\ &\lesssim_R \|g\|_{Z^{n, \lambda}} + \|v\|_{\mathcal{LE}_0^n} \end{aligned}$$

where $\mathcal{LE}^*([R, 2R])$ indicates the $\mathcal{LE}^*$ norm restricted to $R \leq r \leq 2R$. The last inequality follows due to the commutator $[P_0, \chi_{>R}]$ being supported on $[R, 2R]$ and (4.15). Thus by Proposition 3.4 we have

$$\|h\|_{Z^{n,\lambda}} \lesssim_R \|g\|_{Z^{n+4,\lambda}}. \quad (4.25)$$

Rewriting the equation $P_0 w = h$ using $P_0 = \Delta + P^2$, we obtain

$$-\Delta w = -h + \chi_{>R/2} P^2 w. \quad (4.26)$$

Note $\chi_{>R/2} P^2 w = P^2 w$ because the support of $P^2 w$ is contained in the region where $\chi_{>R/2} = 1$.

By Lemma 4.1, w can be written as in (4.1) where (4.2) holds so we have

$$\begin{aligned} & \sum_{j=0}^{\lambda-2} |c_j| + \|e_j\|_{\ell^1 S(1)} + \|d(r)\|_{L^\infty} + \|S_r d\|_{\ell^1 S(1)} + \|q\|_{Z^{n+2,\lambda-2}} \\ & \leq C \left\| \chi_{>\frac{R}{2}} P^2 \left(\sum_{j=0}^{\lambda-2} (c_j \cdot \nabla^j \langle r \rangle^{-1} + e_j \cdot (\nabla^j \langle r \rangle^{-1}) \langle r \rangle^{j-\lambda+1}) + d \cdot \nabla^{\lambda-1} \langle r \rangle^{-1} + q \right) \right\|_{Z^{n,\lambda}} \\ & \quad + C_R \|h\|_{Z^{n,\lambda}}. \end{aligned} \quad (4.27)$$

If we can show

$$\begin{aligned} & \left\| \chi_{>\frac{R}{2}} P^2 \left(\sum_{j=0}^{\lambda-2} (c_j \cdot \nabla^j \langle r \rangle^{-1} + e_j \cdot (\nabla^j \langle r \rangle^{-1}) \langle r \rangle^{j-\lambda+1}) + d \cdot \nabla^{\lambda-1} \langle r \rangle^{-1} + q \right) \right\|_{Z^{n,\lambda}} \\ & \lesssim R^{-1} \left(\sum_{j=0}^{\lambda-2} |c_j| + \|e_j\|_{\ell^1 S(1)} + \|d\|_{L^\infty} + \|S_r d\|_{\ell^1 S(1)} + \|q\|_{Z^{n+2,\lambda-2}} \right) \end{aligned} \quad (4.28)$$

then choosing R sufficiently large allows us to bootstrap the perturbative term in (4.28), and we can use (4.25) to obtain

$$\sum_{j=0}^{\lambda-2} (|c_j| + \|e_j\|_{\ell^1 S(1)}) + \|d(r)\|_{L^\infty} + \|S_r d\|_{\ell^1 S(1)} + \|q\|_{Z^{n+2,\lambda-2}} \leq (1 - cR^{-1})^{-1} C_R \|g\|_{Z^{n+4,\lambda}}$$

as desired. Thus we wish to show

$$\|\chi_{>\frac{R}{2}} P^2 c_j \cdot \nabla^j \langle r \rangle^{-1}\|_{Z^{n,\lambda}} \lesssim R^{-1} |c_j|, \quad 0 \leq j \leq \lambda - 2 \quad (4.29)$$

$$\|\chi_{>\frac{R}{2}} P^2 e_j \cdot (\nabla^j \langle r \rangle^{-1}) \langle r \rangle^{j-\lambda+1}\|_{Z^{n,\lambda}} \lesssim R^{-1} \|e_j\|_{\ell^1 S(1)}, \quad 1 \leq j \leq \lambda - 2 \quad (4.30)$$

$$\|\chi_{>\frac{R}{2}} P^2 d \nabla^{\lambda-1} \langle r \rangle^{-1}\|_{Z^{n,\lambda}} \lesssim R^{-1} (\|d\|_{L^\infty} + \|S_r d\|_{\ell^1 S(1)}) \quad (4.31)$$

$$\|\chi_{>\frac{R}{2}} P^2 q\|_{Z^{n,\lambda}} \lesssim R^{-1} \|q\|_{Z^{n+2,\lambda-2}} \quad (4.32)$$

where the implicit constants are independent of R . We will see (4.30), (4.31), and (4.32) hold for $\lambda \leq \kappa + 1$. And we will see (4.29) holds for $\lambda \leq \kappa + 1$ when $1 \leq j \leq \lambda - 2$. When $j = 0$, (4.29) holds only for $\lambda \leq \kappa$. Our argument handling $c_0 \langle r \rangle^{-1}$ when $\lambda = \kappa + 1$ will change the space we can assume e_0 is in, causing the difference in the result for $\lambda = \kappa + 1$.

First we prove (4.32). Direct calculation yields

$$\begin{aligned}
\|\langle r \rangle^\lambda \chi_{>R/2} P^2 q\|_{\mathcal{LE}^*} &\lesssim \sum_{a \leq 2} \|\chi_{>R/2} \langle r \rangle^\lambda \rho_\ell^{-\kappa-2+a} \nabla^a q\|_{\mathcal{LE}^*} + \|\chi_{>R/2} \langle r \rangle^\lambda \rho_r^{-\kappa-2} \Delta_\omega q\|_{\mathcal{LE}^*} \\
&\quad + \|\chi_{>R/2} \langle r \rangle^\lambda \rho_r^{-\kappa-2} q\|_{\mathcal{LE}^*} \\
&\lesssim \sum_{a \leq 2} \|\chi_{>R/2} \langle r \rangle^{\lambda-\kappa-2+a} \nabla^a q\|_{\mathcal{LE}^*},
\end{aligned} \tag{4.33}$$

and we note (4.33) holds for any q with enough regularity. Now we have

$$\begin{aligned}
&\|\langle r \rangle^\lambda T^i \Omega^j S_r^k \chi_{>R/2} P^2 q\|_{\mathcal{LE}^*} \\
&\lesssim \sum_{a \leq 2} \|\langle r \rangle^{\lambda-\kappa-2+a} \nabla^a T^i \Omega^j S_r^k q\|_{\mathcal{LE}^*([\frac{R}{2}, \infty))} + \|\langle r \rangle^\lambda [T^i \Omega^j S_r^k, \chi_{>R/2} P^2] q\|_{\mathcal{LE}^*}.
\end{aligned} \tag{4.34}$$

For the commutator $[T^i \Omega^j S_r^k, \chi_{>R/2} P^2]$ we find

$$[T^i \Omega^j S_r^k, \chi_{>R/2} P^2] = [T^i, \chi_{>R/2}] \Omega^j S_r^k P^2 + T^i \Omega^j [S_r^k, \chi_{>R/2}] P^2 + \chi_{>R/2} [T^i \Omega^j S_r^k, P^2].$$

The commutators $[S_r^k, \chi_{>R/2}]$ and $[T^i, \chi_{>R/2}]$ are compactly supported and uniformly bounded in R . We note if Q^2 is an operator of the form

$$Q^2 = \partial_i h^{ij} \partial_j + h^\omega \Delta_\omega + h_r + h_\ell, \quad h^{ij} \in \ell^1 S(r^{-\kappa}), \quad h^\omega, h_r \in S_{rad}(r^{-\kappa-2}), \quad \text{and } h_\ell \in \ell^1 S(r^{-\kappa-2}),$$

then $[\Gamma, Q^2] = Q^2$ where we allow the precise form of Q^2 to change each time it appears. Since P^2 is an operator of the above form, it follows that $[\Gamma^n, P^2] = Q^2 \Gamma^{<n}$. Now we can use (4.33) to obtain

$$\sup_{i+j+k \leq n} \|\langle r \rangle^\lambda [T^i \Omega^j S_r^k, \chi_{>R/2} P^2] q\|_{\mathcal{LE}^*} \lesssim \sup_{i+j+k \leq n-1} \sum_{a \leq 2} \|\langle r \rangle^{\lambda-\kappa-2+a} \nabla^a T^i \Omega^j S_r^k q\|_{\mathcal{LE}^*([\frac{R}{2}, \infty))},$$

which combined with (4.34) yields

$$\|\chi_{>R/2} P^2 q\|_{Z^{n,\lambda}} \lesssim \|q\|_{Z^{n+2,\lambda-\kappa-2}([\frac{R}{2}, \infty))} \lesssim R^{-1} \|q\|_{Z^{n+2,\lambda-2}}, \tag{4.35}$$

using $|\nabla q| \leq r^{-1}(|S_r q| + |\Omega q|)$ (which holds for general q), as desired.

In the following calculations we will use the fact that if $\phi \in \ell^1 S(1)$, then

$$\|(\partial^\alpha \phi) \rho^q\|_{Z^{n,\lambda}} \lesssim \sum_{|B|=|\alpha|}^{n+|\alpha|} \|\langle r \rangle^{|B|} \partial^B \phi\|_{L^\infty(A_m)} \tag{4.36}$$

for $q \leq -\lambda - 2 + |\alpha|$, which is established by direct calculation.

To prove (4.31) we use (4.35) to find

$$\begin{aligned}
\|\chi_{>R/2} P^2 d(r) \nabla^{\lambda-1} \langle r \rangle^{-1}\|_{Z^{n,\lambda}} &\lesssim \|\chi_{>R/2} P^2 d(r) \rho^{-\lambda}\|_{Z^{n,\lambda}} \\
&= \|d(r) \rho^{-\lambda}\|_{Z^{n+2,\lambda-\kappa-2}([\frac{R}{2}, \infty))} \\
&\lesssim \|S_r(d\rho^{-\lambda})\|_{Z^{n+1,\lambda-\kappa-2}([\frac{R}{2}, \infty))} + \|\Omega(d\rho^{-\lambda})\|_{Z^{n+1,\lambda-\kappa-2}([\frac{R}{2}, \infty))} \\
&\quad + \|T(d\rho^{-\lambda})\|_{Z^{n+1,\lambda-\kappa-2}([\frac{R}{2}, \infty))} + \|d\rho^{-\lambda}\|_{Z^{0,\lambda-\kappa-2}([\frac{R}{2}, \infty))} \\
&\lesssim \|(S_r d) \rho^{-\lambda}\|_{Z^{n+1,\lambda-\kappa-2}([\frac{R}{2}, \infty))} + \|d\rho^{-\lambda}\|_{Z^{n+1,\lambda-\kappa-2}([\frac{R}{2}, \infty))} \\
&\lesssim \sum_{b=0}^{n+1} \|(S_r d) \rho^{-\lambda}\|_{Z^{b,\lambda-\kappa-2}([\frac{R}{2}, \infty))} + \|d\rho^{-\lambda}\|_{Z^{0,\lambda-\kappa-2}([\frac{R}{2}, \infty))} \\
&\lesssim R^{-1} \|S_r d\|_{\ell^1 S(1)} + R^{-1} \|d\|_{L^\infty}.
\end{aligned}$$

The last inequality follows from (4.36).

To prove (4.30) we use (4.32) and (4.36) to find

$$\|\chi_{>R/2} P^2 e_j (\nabla^j \langle r \rangle^{-1}) \langle r \rangle^{j-\lambda+1}\|_{Z^{n,\lambda}} \lesssim R^{-1} \|e_j (\nabla^j \langle r \rangle^{-1}) \langle r \rangle^{j-\lambda+1}\|_{Z^{n+2,\lambda-2}} \lesssim R^{-1} \|e_j\|_{\ell^1 S(1)}. \quad (4.37)$$

Finally we consider (4.29). Using (4.35) we find

$$\begin{aligned}
\|\chi_{>R/2} P^2 c_j \nabla^j \langle r \rangle^{-1}\|_{Z^{n,\lambda}} &\lesssim \|c_j \nabla^j \langle r \rangle^{-1}\|_{Z^{n+2,\lambda-\kappa-2}([R/2, \infty))} \\
&\lesssim |c_j| \|\langle r \rangle^{\lambda-\kappa-3-j}\|_{\mathcal{LE}^*([R/2, \infty))} \\
&\lesssim |c_j| \sum_{m > \log \frac{R}{2}} 2^{m(\lambda-\kappa-1-j)}.
\end{aligned}$$

If $\lambda < \kappa + 1 + j$, then this yields

$$\|\chi_{>R/2} P^2 c_j \nabla^j \langle r \rangle^{-1}\|_{Z^{n,\lambda}} \lesssim R^{-1} |c_j|$$

as desired. Thus the c_0 term fails to be perturbative when $\lambda = \kappa + 1$.

A straightforward calculation yields $|P^2(c_0 \langle r \rangle^{-1})| \lesssim c_0 \rho_r^{-\kappa-3} + c_0 \rho_\ell^{-\kappa-3}$. Note that $\rho_\ell^{-\kappa-3} \in Z^{\nu, \kappa+1}$ for all ν . We obtain decay as $R \rightarrow \infty$ so that $\|\chi_{>R/2} c_0 \rho_\ell^{-\kappa-3}\|_{Z^{n,\lambda}} \lesssim o_R(1) |c_0|$. Thus only the radial term $c_0 \rho_r^{-\kappa-3}$, which arises when the radial scalar term in P^2 lands on $c_0 r^{-1}$, fails to be perturbative.

To handle this piece we use Lemmas (4.1) and (4.2) to find for $\lambda = \kappa + 1$, if

$$-\Delta w = h_1 + h_2, \quad h_1 \in Z^{n,\lambda}, \quad h_2 \in S_{rad}(r^{-\lambda-2})$$

where $\text{supp}(h_2) \subseteq \{r \geq \frac{R}{4}\}$, then w can be written as in (4.1) with

$$\begin{aligned}
|c_0| + R^{-\frac{1}{2}} \|e_0\|_{S(1)} + \sum_{j=1}^{\lambda-2} |c_j| + \|e_j\|_{\ell^1 S(1)} + \|d(r)\|_{L^\infty} + \|S_r d\|_{\ell^1 S(1)} + \|q\|_{Z^{n+2,\lambda-2}} \\
\lesssim \|h_1\|_{Z^{n,\lambda}} + \|h_2\|_{R^{\frac{1}{2}} S_{rad}(r^{-\lambda-2})}.
\end{aligned}$$

Note that $\ell^1 S(1) \subset S(1)$ so that the change of space for e_0 as compared with Lemma (4.1) causes no problem. We now have (4.26) is perturbative with respect to this estimate since

$$R^{-\frac{1}{2}} \|\chi_{<R/2} c_0 \rho_r^{-\kappa-3}\|_{S_{rad}(r^{-\kappa-3})} \lesssim R^{-\frac{1}{2}} |c_0|.$$

Furthermore, the e_0 terms remain perturbative since our estimate (4.37) came with extra powers of R^{-1} and $\|\chi_{>\frac{R}{2}} e_0\|_{\ell^1 S(1)} \lesssim \log R \|e_0\|_{S(1)}$. This concludes the proof of the proposition. $\square$

5. LOW FREQUENCY ANALYSIS

In this section we use Proposition 4.3 to calculate the error when we estimate $R_\tau g$ by $(R_0 g)e^{-i\tau\langle r \rangle}$. Direct calculation yields for any function ϕ

$$\begin{aligned} P_\tau(\phi e^{-i\tau\langle r \rangle}) &= [(\Delta + P^2)\phi]e^{-i\tau\langle r \rangle} - 2i\tau[(\partial_r + \frac{1}{r})\phi]e^{-i\tau\langle r \rangle} \\ &\quad + \left[(\tau(\rho_\ell^{-\kappa-1} + \rho_\ell^{-\kappa}\nabla) + \tau^2\rho_\ell^{-\kappa})\phi \right]e^{-i\tau\langle r \rangle}. \end{aligned} \quad (5.1)$$

Since $P_0 = (\Delta + P^2)$, (5.1) gives

$$\begin{aligned} P_\tau\left(R_\tau g - (R_0 g)e^{-i\tau\langle r \rangle}\right) &= g(1 - e^{-i\tau\langle r \rangle}) - 2i\tau[(\partial_r + \frac{1}{r})R_0 g]e^{-i\tau\langle r \rangle} \\ &\quad + \left[(\tau(\rho_\ell^{-\kappa-1} + \rho_\ell^{-\kappa}\nabla) + \tau^2\rho_\ell^{-\kappa})R_0 g \right]e^{-i\tau\langle r \rangle}. \end{aligned} \quad (5.2)$$

Analyzing the right hand side of (5.2) with $R_0 g$ represented by the expansion found in Proposition 4.3 will yield a useful form of the error in the low frequency estimate. Lemmas 5.1 - 5.5 provide preliminary calculations that help us handle terms arising in the $-2i\tau[(\partial_r + \frac{1}{r})R_0 g]e^{-i\tau\langle r \rangle}$ piece in (5.2). In Proposition 5.6 we will use these lemmas to establish the error in the low frequency estimate.

Lemma 5.1. *Let n be a positive integer. Let φ_0 be given and assume φ_a satisfies*

$$-\frac{1}{2}\Delta\varphi_a = (\partial_r + \frac{1}{r})\varphi_{a-1} \quad (5.3)$$

and $\varphi_a \in S(r^{-1})$ for $1 \leq a \leq n$. Then

$$\begin{aligned} &-2i\tau[(\partial_r + \frac{1}{r})\varphi_0]e^{-i\tau\langle r \rangle} \\ &= \sum_{a=1}^n \left(-P_\tau((-i\tau)^a \varphi_a e^{-i\tau\langle r \rangle}) + (\tau^a \zeta_\kappa^\nu + \tau^{a+1} \zeta_\kappa^\nu + \tau^{a+2} \zeta_{\kappa-1}^\nu) e^{-i\tau\langle r \rangle} \right) \\ &\quad - 2(-i\tau)^{n+1}(\partial_r + \frac{1}{r})(\varphi_n) e^{-i\tau\langle r \rangle} \end{aligned} \quad (5.4)$$

Proof. By (5.1) and (5.3), we have

$$\begin{aligned}
-2i\tau\left[(\partial_r + \frac{1}{r})\varphi_0\right]e^{-i\tau\langle r\rangle} &= i\tau(\Delta\varphi_1)e^{-i\tau\langle r\rangle} \\
&= P_\tau(i\tau\varphi_1e^{-i\tau\langle r\rangle}) - 2\tau^2\left[(\partial_r + \frac{1}{r})\varphi_1\right]e^{-i\tau\langle r\rangle} - i\tau(P^2\varphi_1)e^{-i\tau\langle r\rangle} \\
&\quad + \tau\left[(\tau(\rho_\ell^{-\kappa-1} + \rho_\ell^{-\kappa}\nabla) + \tau^2\rho_\ell^{-\kappa})\varphi_1\right]e^{-i\tau\langle r\rangle} \\
&= i\tau P_\tau(\varphi_1e^{-i\tau\langle r\rangle}) - 2\tau^2\left[(\partial_r + \frac{1}{r})\varphi_1\right]e^{-i\tau\langle r\rangle} \\
&\quad + \left(\tau\zeta_\kappa^\nu + \tau^2\zeta_\kappa^\nu + \tau^3\zeta_{\kappa-1}^\nu\right)e^{-i\tau\langle r\rangle}.
\end{aligned} \tag{5.5}$$

The last inequality uses the fact that $\varphi_1 \in S(r^{-1})$ by assumption, and the term $\tau\zeta_\kappa^\nu$ is limited in its decay by the $\rho_r^{-\kappa-2}$ term in P^2 .

Repeating this argument, now for $2i\tau^2\left[(\partial_r + \frac{1}{r})i\varphi_1\right]e^{-i\tau\langle r\rangle}$ with φ_2 satisfying $-\frac{1}{2}\Delta\varphi_2 = (\partial_r + \frac{1}{r})\varphi_1$ and plugging the resulting expression into (5.5) we find

$$\begin{aligned}
-2i\tau(\partial_r + \frac{1}{r})(\varphi_0)e^{-i\tau\langle r\rangle} &= P_\tau\left(i\tau\varphi_1e^{-i\tau\langle r\rangle} + \tau^2\varphi_2e^{-i\tau\langle r\rangle}\right) + 2i\tau^3\left[(\partial_r + \frac{1}{r})\varphi_2\right]e^{-i\tau\langle r\rangle} \\
&\quad + \left(\tau\zeta_\kappa^\nu + \tau^2\zeta_\kappa^\nu + \tau^3(\zeta_\kappa^\nu + \zeta_{\kappa-1}^\nu) + \tau^4\zeta_\kappa^\nu\right)e^{-i\tau\langle r\rangle}
\end{aligned}$$

as long as $\varphi_1, \varphi_2 \in S(r^{-1})$.

Repeating this process yields (5.4) since $\varphi_a \in S(r^{-1})$ for $a \leq n$ by assumption. $\square$

Lemma 5.2. *If $r \geq 3$, then for $j \in \mathbb{N}_+$*

$$(\partial_r + \frac{1}{r})(\nabla^j\langle r\rangle^{-1}) = -\frac{1}{2}\Delta\left(\sum_{k=0}^{j-1}\nabla^{j-1-k}\frac{x}{r}\nabla^k\langle r\rangle^{-1}\right). \tag{5.6}$$

Proof. Direct calculation yields $[(\partial_r + \frac{1}{r}), \nabla] = -\frac{1}{2}[\Delta, \frac{x}{r}]$, and we use this identity to handle the $j = 1$ case:

$$(\partial_r + \frac{1}{r})\nabla\langle r\rangle^{-1} = \nabla(\partial_r + \frac{1}{r})r^{-1} + \frac{1}{2}\frac{x}{r}\Delta\langle r\rangle^{-1} - \frac{1}{2}\Delta\left(\frac{x}{r}r^{-1}\right) = -\frac{1}{2}\Delta\left(\frac{x}{r}\langle r\rangle^{-1}\right).$$

Now fix J and assume (5.6) holds for $j = J - 1$. We calculate

$$\begin{aligned}
(\partial_r + \frac{1}{r})\nabla^J\langle r\rangle^{-1} &= \nabla(\partial_r + \frac{1}{r})\nabla^{J-1}\langle r\rangle^{-1} + \frac{1}{2}\frac{x}{r}\Delta(\nabla^{J-1}\langle r\rangle^{-1}) - \frac{1}{2}\Delta\left(\frac{x}{r}\nabla^{J-1}\langle r\rangle^{-1}\right) \\
&= \nabla\left(-\frac{1}{2}\Delta\sum_{k=0}^{J-2}\nabla^{J-2-k}\frac{x}{r}\nabla^k\langle r\rangle^{-1}\right) + 0 - \frac{1}{2}\Delta\left(\frac{x}{r}\nabla^{J-1}\langle r\rangle^{-1}\right) \\
&= -\frac{1}{2}\Delta\sum_{k=0}^{J-1}\nabla^{J-1-k}\frac{x}{r}\nabla^k\langle r\rangle^{-1}
\end{aligned}$$

for $r \geq 3$, as desired. $\square$

Lemma 5.3. *Let $a \in \mathbb{N}$ be fixed and let φ^j be a family of functions indexed by j which satisfy*

$$(\partial_r + \frac{1}{r})\varphi^j = -\frac{1}{2}\Delta \sum_{k=a}^{j-1} \nabla^{j-1-k} \frac{x}{r} \varphi^k, \quad j \geq a+1 \quad (5.7)$$

and $(\partial_r + \frac{1}{r})\varphi^a = 0$. Then

$$(\partial_r + \frac{1}{r}) \sum_{k=a}^{j-1} \nabla^{j-1-k} \frac{x}{r} \varphi^k = -\frac{1}{2}\Delta \sum_{k=a+1}^{j-1} \nabla^{j-1-k} \frac{x}{r} \left(\sum_{\ell=a}^{k-1} \nabla^{k-1-\ell} \frac{x}{r} \varphi^\ell \right). \quad (5.8)$$

Note that if we define

$$\varphi_1^j := \sum_{k=a}^{j-1} \nabla^{j-1-k} \frac{x}{r} \varphi^k$$

then the lemma shows that φ_1^j is a family of functions indexed by j which satisfy the assumptions of the lemma with a replaced by $a+1$.

Proof. We calculate

$$\begin{aligned} (\partial_r + \frac{1}{r}) \sum_{k=a}^{j-1} \nabla^{j-1-k} \frac{x}{r} \varphi^k &= \sum_{k=a+1}^{j-1} \nabla^{j-1-k} \frac{x}{r} (\partial_r + \frac{1}{r}) \varphi^k + \sum_{k=a}^{j-2} [(\partial_r + \frac{1}{r}), \nabla^{j-1-k}] \frac{x}{r} \varphi^k \\ &= -\frac{1}{2} \sum_{k=a+1}^{j-1} \sum_{\ell=a}^{k-1} \nabla^{j-1-k} \left(\frac{x}{r} \Delta \right) \nabla^{k-1-\ell} \frac{x}{r} \varphi^\ell \\ &\quad - \frac{1}{2} \sum_{k=a}^{j-2} \sum_{\ell=1}^{j-1-k} \nabla^{j-1-k-\ell} [\Delta, \frac{x}{r}] \nabla^{\ell-1} \frac{x}{r} \varphi^k. \end{aligned} \quad (5.9)$$

The first equality uses the assumption $(\partial_r + \frac{1}{r})\varphi^a = 0$. The second equality uses (5.7) and the identity

$$[(\partial_r + \frac{1}{r}), \nabla^b] = \sum_{\ell=1}^b \nabla^{b-\ell} [(\partial_r + \frac{1}{r}), \nabla] \nabla^{\ell-1} = -\frac{1}{2} \sum_{\ell=1}^b \nabla^{b-\ell} [\Delta, \frac{x}{r}] \nabla^{\ell-1}. \quad (5.10)$$

Changing the order of summation in the first term on the right hand side of (5.9) and simply switching the indexing labels in the second term yields

$$\begin{aligned}
(\partial_r + \frac{1}{r}) \sum_{k=a}^{j-1} \nabla^{j-1-k} \frac{x}{r} \varphi^k &= -\frac{1}{2} \sum_{\ell=a}^{j-2} \sum_{k=\ell+1}^{j-1} \nabla^{j-1-k} \left(\frac{x}{r} \Delta \right) \nabla^{k-1-\ell} \frac{x}{r} \varphi^\ell \\
&\quad - \frac{1}{2} \sum_{\ell=a}^{j-2} \sum_{k=1}^{j-1-\ell} \nabla^{j-1-\ell-k} [\Delta, \frac{x}{r}] \nabla^{k-1} \frac{x}{r} \varphi^\ell \\
&= -\frac{1}{2} \sum_{\ell=a}^{j-2} \sum_{k=\ell+1}^{j-1} \nabla^{j-1-k} \left(\frac{x}{r} \Delta + [\Delta, \frac{x}{r}] \right) \nabla^{k-1-\ell} \frac{x}{r} \varphi^\ell \\
&= -\frac{1}{2} \Delta \sum_{k=a+1}^{j-1} \nabla^{j-1-k} \frac{x}{r} \sum_{\ell=a}^{k-1} \nabla^{k-1-\ell} \frac{x}{r} \varphi^\ell.
\end{aligned}$$

To obtain the second equality we redefine the k index by $k \mapsto k + \ell$. To obtain the third equality we switch the order of summation. This completes the proof of the lemma. $\square$

Lemma 5.4. *Let $j \geq 0$ and $r \geq 3$. Then*

$$\begin{aligned}
&-2i\tau(\partial_r + \frac{1}{r})(\nabla^j \langle r \rangle^{-1}) e^{-i\tau \langle r \rangle} \\
&= \sum_{a=1}^j \left(-P_\tau((-i\tau)^a F_a^j e^{-i\tau \langle r \rangle}) + (\tau^a \zeta_\kappa^\nu + \tau^{a+1} \zeta_\kappa^\nu + \tau^{a+2} \zeta_{\kappa-1}^\nu) e^{-i\tau \langle r \rangle} \right)
\end{aligned} \tag{5.11}$$

where $|F_a| \lesssim \langle r \rangle^{-1}$ for $1 \leq a \leq j$.

Proof. We define $\varphi_0^j := \nabla^j \langle r \rangle^{-1}$. By Lemma 5.2 φ_0^j satisfies (5.7) with $a = 0$. Furthermore $(\partial_r + \frac{1}{r})\varphi_0^j = 0$, so the assumptions of Lemma 5.3 are satisfied.

Then defining $\varphi_1^j := \sum_{k=0}^{j-1} \nabla^{j-1-k} \frac{x}{r} \varphi_0^k$, we have

$$(\partial_r + \frac{1}{r})\varphi_1^j = -\frac{1}{2} \Delta \sum_{k=1}^{j-1} \nabla^{j-1-k} \frac{x}{r} \varphi_1^k =: -\frac{1}{2} \Delta \varphi_2^j$$

by Lemma 5.3. Now φ_1^j satisfies (5.7) with $a = 1$ and

$$(\partial_r + \frac{1}{r})\varphi_1^j = (\partial_r + \frac{1}{r}) \frac{x}{r} \langle r \rangle^{-1} = 0$$

so we can iterate the process again. We define

$$\varphi_\ell^j := \sum_{k=\ell-1}^{j-1} \nabla^{j-1-k} \frac{x}{r} \varphi_{\ell-1}^k \tag{5.12}$$

and see $(\partial_r + \frac{1}{r})\varphi_\ell^j = (\partial_r + \frac{1}{r}) \left(\frac{x}{r} \right)^j \langle r \rangle^{-1} = 0$ so the assumptions of Lemma 5.3 are satisfied at each iteration and we find for $0 \leq n \leq j-1$

$$(\partial_r + \frac{1}{r})\varphi_n^j = -\frac{1}{2} \Delta \sum_{k=n}^{j-1} \nabla^{j-1-k} \frac{x}{r} \varphi_n^k =: -\frac{1}{2} \Delta \varphi_{n+1}^j.$$

We note $\varphi_n^j \in S(r^{-j-1+n})$ so $\varphi_n^j \in S(r^{-1})$ for $n \leq j$. Now (5.11) follows by (5.4) since $(\partial_r + \frac{1}{r})\varphi_j^j = 0$. $\square$

We define $a \wedge b$ to be a smooth function such that $a \wedge b = \min(a, b)$ when $a \gg b$ or $b \ll a$.

Lemma 5.5. *Let $\varphi_0 \in S_{rad}(r^{-\kappa-1})$. Then*

$$\begin{aligned} & 2\tau[(\partial_r + \frac{1}{r})\varphi_0]e^{-i\tau\langle r \rangle} \\ &= \sum_{a=1}^{\kappa-1} \left(P_\tau(\tau^a F_a e^{-i\tau\langle r \rangle}) + (\tau^a \zeta_\kappa^\nu + \tau^{a+1} \zeta_\kappa^\nu + \tau^{a+1} \zeta_{\kappa-1}^\nu) e^{-i\tau\langle r \rangle} \right) + (\tau^\kappa \zeta_1^\nu + \tau^{\kappa+1} \zeta_0^\nu) e^{-i\tau\langle r \rangle} \\ & \quad - \tau^\kappa P_\tau \left((\langle r \rangle^{-1} \varepsilon_1(r \wedge |\tau|^{-1}) + \tau(\varepsilon_2(r \wedge |\tau|^{-1}) - \varepsilon_2(|\tau|^{-1})) e^{-i\tau\langle r \rangle} \right). \end{aligned} \quad (5.13)$$

where $|F_a| \lesssim \langle r \rangle^{-1}$ and $\varepsilon_1, \varepsilon_2 \in S(\log r)$.

Proof. Since $\varphi_0 \in S_{rad}(r^{-\kappa-1})$, we see $(\partial_r + \frac{1}{r})\varphi_0 \in S_{rad}(r^{-\kappa-2})$. Lemma 4.2 implies

$$(\partial_r + \frac{1}{r})\varphi_n = -\frac{1}{2}\Delta\varphi_{n+1}, \quad \varphi_n \in S_{rad}(r^{-\kappa-1+n})$$

for $0 \leq n \leq \kappa - 2$. Then by Lemma 5.1 we have

$$\begin{aligned} & -2i\tau[(\partial_r + \frac{1}{r})\varphi_0]e^{-i\tau\langle r \rangle} \\ &= \sum_{a=1}^{\kappa-1} \left(-P_\tau((-i\tau)^a F_a e^{-i\tau\langle r \rangle}) + (\tau^a \zeta_\kappa^\nu + \tau^{a+1} \zeta_\kappa^\nu + \tau^{a+1} \zeta_{\kappa-1}^\nu) e^{-i\tau\langle r \rangle} \right) \\ & \quad - 2(-i\tau)^\kappa (\partial_r + \frac{1}{r})(\varphi_{\kappa-1}) e^{-i\tau\langle r \rangle}. \end{aligned} \quad (5.14)$$

Since $\varphi_{\kappa-1} \in S_{rad}(r^{-2})$, we see $(\partial_r + \frac{1}{r})\varphi_{\kappa-1} \in S_{rad}(r^{-3})$. By Lemma 4.2 there exists an $\varepsilon_1 \in S_{rad}(\log r)$ such that

$$-\frac{1}{2}\Delta\langle r \rangle^{-1}\varepsilon_1(r) = (\partial_r + \frac{1}{r})\varphi_{\kappa-1}.$$

We wish to avoid the logarithmic growth in r , so we use the modified function

$$\tilde{\varphi}_\kappa := \langle r \rangle^{-1} \varepsilon_1(r \wedge |\tau|^{-1}).$$

Now we have

$$-\frac{1}{2}\Delta\tilde{\varphi}_\kappa = \chi_{<|\tau|^{-1}}(\partial_r + \frac{1}{r})\varphi_{\kappa-1} = (\partial_r + \frac{1}{r})\varphi_{\kappa-1} - \chi_{>|\tau|^{-1}}(\partial_r + \frac{1}{r})\varphi_{\kappa-1}.$$

Using (5.1) we find

$$\begin{aligned}
& 2\tau^\kappa \left(\partial_r + \frac{1}{r} \right) \varphi_{\kappa-1} e^{-i\tau \langle r \rangle} \\
&= P_\tau \left(\tau^\kappa (-\tilde{\varphi}_\kappa) e^{-i\tau \langle r \rangle} \right) + 2i\tau^{\kappa+1} \left(\partial_r + \frac{1}{r} \right) (\tilde{\varphi}_\kappa) e^{-i\tau \langle r \rangle} \\
&\quad + \tau^\kappa \left(\tau (\rho_\ell^{-\kappa-1} + \rho_\ell^{-\kappa} \nabla) + \tau^2 \rho_\ell^{-\kappa} \right) (\tilde{\varphi}_\kappa) e^{-i\tau \langle r \rangle} + \tau^\kappa P^2 (\tilde{\varphi}_\kappa) e^{-i\tau \langle r \rangle} \\
&\quad + 2\tau^\kappa \chi_{>|\tau|^{-1}} \left[\left(\partial_r + \frac{1}{r} \right) \varphi_{\kappa-1} \right] e^{-i\tau \langle r \rangle} \\
&= P_\tau \left(\tau^\kappa (-\tilde{\varphi}_\kappa) e^{-i\tau \langle r \rangle} \right) + 2\tau^{\kappa+1} \chi_{<|\tau|^{-1}} \langle r \rangle^{-1} (\partial_r \varepsilon_1(r)) e^{-i\tau \langle r \rangle} \\
&\quad + \left(\tau^\kappa \zeta_\kappa^\nu + \tau^{\kappa+1} (\zeta_{\kappa-1}^\nu + \zeta_0^\nu) + \tau^{\kappa+2} \zeta_{\kappa-2}^\nu \right) e^{-i\tau \langle r \rangle}.
\end{aligned} \tag{5.15}$$

Here we used the fact $|\tau|^{-1} \chi_{>|\tau|^{-1}} (\partial_r + \frac{1}{r}) \varphi_{\kappa-1} \in \ell^1 S(r^{-2})$ because $(\partial_r + \frac{1}{r}) \varphi_{\kappa-1} \in S_{rad}(r^{-3})$, and the cutoff function allows us to pull out a τ factor when summing in the $\ell^1 S(r^{-2})$ norm.

We still need to handle the term $2\tau^{\kappa+1} \chi_{<|\tau|^{-1}} \langle r \rangle^{-1} (\partial_r \varepsilon_1(r)) e^{-i\tau \langle r \rangle}$. By Lemma 4.2 there exists an $\varepsilon_2 \in S_{rad}(\log r)$ such that

$$-\frac{1}{2} \Delta \varepsilon_2 = \langle r \rangle^{-1} (\partial_r \varepsilon_1(r)) \in S_{rad}(r^{-2}).$$

To remove the logarithmic growth in r we use the modified function

$$\tilde{\varphi}_{\kappa+1} := \varepsilon_2(r \wedge |\tau|^{-1}) - \varepsilon_2(|\tau|^{-1})$$

and find

$$-\frac{1}{2} \Delta \tilde{\varphi}_{\kappa+1} = \chi_{<|\tau|^{-1}} \langle r \rangle^{-1} (\partial_r \varepsilon_1(r)).$$

Then by (5.1) we have

$$\begin{aligned}
& 2\tau^{\kappa+1} \chi_{<|\tau|^{-1}} \langle r \rangle^{-1} (\partial_r \varepsilon_1(r)) e^{-i\tau \langle r \rangle} \\
&= P_\tau \left(\tau^{\kappa+1} (-\tilde{\varphi}_{\kappa+1}) e^{-i\tau \langle r \rangle} \right) + 2\tau^{\kappa+2} \left(\partial_r + \frac{1}{r} \right) (\tilde{\varphi}_{\kappa+1}) e^{-i\tau \langle r \rangle} \\
&\quad + \tau^{\kappa+1} \left(\tau (\rho_\ell^{-\kappa-1} + \rho_\ell^{-\kappa} \nabla) + \tau^2 \rho_\ell^{-\kappa} \right) (\tilde{\varphi}_{\kappa+1}) e^{-i\tau \langle r \rangle} + (P^2 \tilde{\varphi}_{\kappa+1}) e^{-i\tau \langle r \rangle} \\
&= P_\tau \left(\tau^{\kappa+1} (-\tilde{\varphi}_{\kappa+1}) e^{-i\tau \langle r \rangle} \right) + \tau^{\kappa+1} \zeta_0^\nu e^{-i\tau \langle r \rangle} + \left(\tau^\kappa \zeta_{\kappa-1}^\nu + \tau^{\kappa+1} \zeta_{\kappa-1}^\nu \right) e^{-i\tau \langle r \rangle}.
\end{aligned} \tag{5.16}$$

Combining (5.14), (5.15), and (5.16) then yields (5.13), as desired. $\square$

We are now ready to calculate the error in the approximation of $R_\tau g$ by $(R_0 g) e^{-i\tau \langle r \rangle}$, which is established in Proposition 5.6. The precise form of the error depends on the regularity and decay assumed for g .

Proposition 5.6. *Let $g_\lambda^\nu \in Z^{\nu, \lambda}$ with $1 \leq \lambda \leq \kappa + 1$ and $\nu > 3\lambda$. Assume $|\tau| \lesssim 1$ and $\Im \tau \leq 0$.*

(1) If $1 \leq \lambda \leq \kappa$ then

$$R_\tau g_\lambda^\nu = (R_0 g_\lambda^\nu) e^{-i\tau \langle r \rangle} + R_\tau (\chi_{>|\tau|^{-1}} g_\lambda^\nu) + \sum_{m=1}^{\lambda-1} \tau^m (F_m + R_0 \zeta_{\lambda-m}^{\nu-3\lambda}) e^{-i\tau \langle r \rangle} + \tau^\lambda (R_\tau h_{\nu-3\lambda})$$

where $|F_m| \lesssim \langle r \rangle^{-1}$, $\zeta_{\lambda-m}^{\nu-3\lambda} \in Z^{\nu-3\lambda, \lambda-m}$, and $h_{\nu-3\lambda}$ satisfies

$$\sup_{i+j+k+q \leq \nu-3\lambda} \|\langle r \rangle^q (\partial_r + i\tau)^q T^i \Omega^j S^k h_{\nu-3\lambda}\|_{\mathcal{L}\mathcal{E}^*} \lesssim 1. \quad (5.17)$$

(2) If $\lambda = \kappa + 1$ then

$$R_\tau g_{\kappa+1}^\nu = (R_0 g_{\kappa+1}^\nu) e^{-i\tau \langle r \rangle} + R_\tau (\chi_{>|\tau|^{-1}} g_{\kappa+1}^\nu) + \sum_{m=1}^{\kappa} \tau^m (F_m + R_0 \zeta_{\kappa+1-m}^{\nu-3m}) e^{-i\tau r} \\ + \tau^\kappa \varepsilon(r, \tau) e^{-i\tau r} + \tau^{\kappa+1} (R_\tau h_{\nu-3\kappa-3})$$

where $|F_m| \lesssim \langle r \rangle^{-1}$, $\zeta_{\kappa+1-m}^{\nu-3m} \in Z^{\nu-3m, \kappa+1-m}$, $\varepsilon(r, \tau)$ is of the form

$$\varepsilon(r, \tau) = \langle r \rangle^{-1} \varepsilon_1(r \wedge |\tau|^{-1}) + \tau \left(\varepsilon_2(r \wedge |\tau|^{-1}) - \varepsilon_2(|\tau|^{-1}) \right)$$

with $\varepsilon_1, \varepsilon_2 \in S(\log r)$, and $h_{\nu-3\lambda}$ satisfies (5.17) with $\lambda = \kappa + 1$.

The term $\varepsilon(r, \tau)$ in the statement of Proposition 5.6 for $\lambda = \kappa + 1$ arises due to the $e_0(r)$ term in Proposition 4.3 in the $\lambda = \kappa + 1$ case. The final decay rate is ultimately determined by this $\varepsilon(r, \tau)$ term.

We use ζ_λ^ν to represent a function in $Z^{\nu, \lambda}$ and allow ζ_λ^ν to change from line to line. The purpose of this notation is to keep track of what function spaces each term in our calculations is in while reserving g_λ^ν to indicate the arbitrary but fixed function in $Z^{\nu, \lambda}$ given in the statement of the proposition. In general $Z^{N_1, L_1} \subset Z^{N_2, L_2}$ for $N_2 \leq N_1$ and $L_2 \leq L_1$. The monotonicity of $Z^{\nu, \lambda}$ allows us to collect terms with different regularity and decay and write them as one term:

$$\zeta_{\lambda_1}^{\nu_1} + \zeta_{\lambda_2}^{\nu_2} = \zeta_{\min(\lambda_1, \lambda_2)}^{\min(\nu_1, \nu_2)}. \quad (5.18)$$

Proof. Define E_λ^ν to be the error associated to $R_\tau g_\lambda^\nu$: $E_\lambda^\nu := R_\tau g_\lambda^\nu - (R_0 g_\lambda^\nu) e^{-i\tau \langle r \rangle}$. We begin by establishing a useful expression for $P_\tau(E_\lambda^\nu)$ using (5.2). Our expansion in Proposition 4.3 shows $R_0 g_\lambda^\nu \in S(r^{-1}) + Z^{\nu-2, \lambda-2}$. It follows that

$$\rho_\ell^{-\kappa-1} R_0 g_\lambda^\nu \in Z^{\nu-2, \kappa}, \quad \rho_\ell^{-\kappa} \nabla R_0 g_\lambda^\nu \in Z^{\nu-3, \kappa}, \quad \text{and} \quad \rho_\ell^{-\kappa} R_0 g_\lambda^\nu \in Z^{\nu-2, \kappa-1}. \quad (5.19)$$

Furthermore we have

$$g_\lambda^\nu - g_\lambda^\nu e^{-i\tau \langle r \rangle} = \chi_{>|\tau|^{-1}}(r) g_\lambda^\nu + \tau \chi_{<|\tau|^{-1}}(r) \langle r \rangle g_\lambda^\nu \frac{1 - e^{-i\tau \langle r \rangle}}{\tau \langle r \rangle} - \tau^\lambda \chi_{>|\tau|^{-1}}(r) \tau^{-\lambda} g_\lambda^\nu e^{-i\tau \langle r \rangle}. \quad (5.20)$$

We use (5.2), (5.19), and (5.20) to find

$$\begin{aligned}
 P_\tau(E_\lambda^\nu) = & \underbrace{\chi_{>|\tau|^{-1}}(r)g_\lambda^\nu}_{(I)} + \tau \underbrace{\left(\chi_{<|\tau|^{-1}}(r)\langle r \rangle g_\lambda^\nu \frac{1 - e^{-i\tau\langle r \rangle}}{\tau r} \right)}_{(II)} - \tau^\lambda \underbrace{\left(\chi_{>|\tau|^{-1}}(r)\tau^{-\lambda}g_\lambda^\nu e^{-i\tau\langle r \rangle} \right)}_{(II)} \\
 & - \underbrace{2i\tau(\partial_r + \frac{1}{r})(R_0g_\lambda^\nu)e^{-i\tau\langle r \rangle}}_{(III)} + \underbrace{(\tau\zeta_\kappa^{\nu-3} + \tau^2\zeta_{\kappa-1}^{\nu-2})e^{-i\tau\langle r \rangle}}_{(IV)}.
 \end{aligned} \tag{5.21}$$

Next we claim that

$$\mathfrak{h}_{\nu-3\lambda} = \zeta_0^{\nu-3\lambda}e^{-i\tau\langle r \rangle} \tag{5.22}$$

satisfies (5.17). Indeed, when $r \geq 3$, we find for any $\phi(x)$

$$\begin{aligned}
 S^k(\phi e^{-i\tau\langle r \rangle}) &= (S_r^k\phi)e^{-i\tau\langle r \rangle}, \quad \Omega^j(\phi e^{-i\tau\langle r \rangle}) = (\Omega^j\phi)e^{-i\tau\langle r \rangle} \\
 |T^i(\phi e^{-i\tau\langle r \rangle})| &\lesssim \sum_{a=0}^i |(T^a\phi)e^{-i\tau\langle r \rangle}|, \quad (\partial_r + i\tau)^q(\phi e^{-i\tau\langle r \rangle}) = (\partial_r^q\phi)e^{-i\tau\langle r \rangle}.
 \end{aligned}$$

Thus when $i + j + k + l \leq \nu - 3\lambda$ we have

$$\|r^l(\partial_r + i\tau)^l T^i \Omega^j S_r^k (\zeta_0^{\nu-3\lambda}e^{-i\tau\langle r \rangle})\|_{\mathcal{LE}^*} \lesssim \|T^i \Omega^j S_r^{k+l} \zeta_0^{\nu-3\lambda}\|_{\mathcal{LE}^*} \lesssim 1.$$

Our formula for E_λ^ν will be recursive in λ , so we begin by calculating E_1^ν directly.

Case 1: $\lambda = 1$

We calculate E_1^ν using (5.21) and the expansion for $R_0g_1^\nu$ given by Proposition 4.3:

$$R_0g_1^\nu = d(r)\langle r \rangle^{-1} + q \tag{5.23}$$

where $d \in L^\infty$, $S_r d \in \ell^1 S(1)$, and $q \in Z^{\nu-2, \lambda-2}$. Terms II and IV in (5.21) are readily seen to be of the form $\tau h_{\nu-3}$ using (5.22). Term I in (5.21) can also be included in $\tau h_{\nu-3}$. The cutoff function restricts I to the region where $r|\tau| \lesssim 1$, so (5.17) reduces to $h_{\nu-3} \in Z^{\nu-3, 0}$. Since $rg_1^\nu \in Z^{\nu, 0}$,

$$\partial_r \left(\frac{1 - e^{-i\tau\langle r \rangle}}{\tau r^n} \right) = \frac{ie^{-i\tau\langle r \rangle}}{r^n} - n \frac{1 - e^{-i\tau\langle r \rangle}}{\tau r^{n+1}}, \quad \partial_r \left(\frac{ie^{-i\tau\langle r \rangle}}{r^n} \right) = -n \frac{ie^{-i\tau\langle r \rangle}}{r^{n+1}} + \frac{\tau e^{-i\tau\langle r \rangle}}{r^n},$$

and

$$S \left(\frac{1 - e^{-i\tau\langle r \rangle}}{\tau r} \right) = 0,$$

we have

$$\langle r \rangle g_1^\nu \frac{1 - e^{-i\tau\langle r \rangle}}{\tau \langle r \rangle} \in Z^{\nu, 0} \subset Z^{\nu-3, 0},$$

as desired. Here and throughout we harmlessly assume $r \geq 3$. We note the above calculations and (5.20) imply

$$R_\tau(g_1^\nu(1 - e^{-i\tau\langle r \rangle})) = R_\tau(\chi_{>|\tau|^{-1}}g_1^\nu) + \tau(R_\tau h_\nu). \tag{5.24}$$

This equation will help us handle terms that will arise when $\lambda \geq 2$ by providing a base case for an inductive argument.

For term III in (5.21), we use (5.23) and write $\partial_r = r^{-1}S_r$ to find $(\partial_r + \frac{1}{r})R_0g_1^\nu \in Z^{\nu-3, 0}$, so this term also satisfies (5.17).

Combining our calculations for terms I, II, III , and IV and applying R_τ to both sides of (5.21) yields

$$E_1^\nu = R_\tau(\chi_{>|\tau|^{-1}} g_1^\nu) + \tau R_\tau h_{\nu-3} \quad (5.25)$$

as desired.

Case 2: $2 \leq \lambda \leq \kappa + 1$ We proceed as in the $\lambda = 1$ case. For term I in (5.21), we see $\langle r \rangle g_\lambda^\nu \frac{1-e^{-i\tau\langle r \rangle}}{\tau\langle r \rangle} \in Z^{\nu, \lambda-1}$ uniformly in τ as $\tau \rightarrow 0$. So we can write

$$I = \tau \chi_{<|\tau|^{-1}}(r) \langle r \rangle g_\lambda^\nu \frac{1-e^{-i\tau\langle r \rangle}}{\tau\langle r \rangle} = \tau \chi_{<|\tau|^{-1}}(r) \zeta_{\lambda-1}^\nu. \quad (5.26)$$

For term II in (5.21), we note $\chi_{>|\tau|^{-1}}(r) \tau^{-\lambda} g_\lambda^\nu \in Z^{\nu, 0}$ uniformly in τ as $\tau \rightarrow 0$. So we can write

$$II = \tau^\lambda \chi_{>|\tau|^{-1}} \tau^{-\lambda} g_\lambda^\nu e^{-i\tau\langle r \rangle} = \tau^\lambda \zeta_0^\nu e^{-i\tau\langle r \rangle}. \quad (5.27)$$

Substituting (5.26) and (5.27) into (5.21) we have

$$\begin{aligned} P_\tau(E_\lambda^\nu) &= \chi_{>|\tau|^{-1}} g_\lambda^\nu + \tau \chi_{<|\tau|^{-1}} \zeta_{\lambda-1}^\nu + \left(\tau \zeta_{\kappa-3}^{\nu-3} + \tau^2 \zeta_{\kappa-1}^{\nu-2} + \tau^\lambda \zeta_0^\nu \right) e^{-i\tau\langle r \rangle} \\ &\quad - 2i\tau \left[\left(\partial_r + \frac{1}{r} \right) R_0 g_\lambda^\nu \right] e^{-i\tau\langle r \rangle}. \end{aligned} \quad (5.28)$$

We claim that terms of the right hand side of (5.28) of the form

$$\tau^m \zeta_{\lambda-m}^\nu e^{-i\tau\langle r \rangle}, \quad 1 \leq m \leq \lambda - 1 \quad (5.29)$$

produce error terms which can be handled inductively. First consider the case $m = \lambda - 1$. By (5.24) we see that after applying R_τ to both sides of (5.28), terms on the right hand side that are of the form (5.29) with $m = \lambda - 1$ become

$$\tau^{\lambda-1} R_\tau(\zeta_1^\nu e^{-i\tau\langle r \rangle}) = \tau^{\lambda-1} R_\tau(\chi_{<|\tau|^{-1}} \zeta_1^\nu) + \tau^\lambda (R_\tau h_\nu), \quad (5.30)$$

and we can appeal to case 1 of the proposition to handle the first term, while the second term is expected to appear in E_λ^ν (In fact this term has more regularity than the h term in the statement of the proposition. We will see the last term on the right hand side of (5.28) limits the amount of regularity we can get for $h_{\nu-3\lambda}$.)

Now consider (5.29) for $1 \leq m \leq \lambda - 2$. Substituting (5.26) and (5.27) into (5.20) gives

$$R_\tau(g_\lambda^\nu e^{-i\tau\langle r \rangle}) = R_\tau(\chi_{<|\tau|^{-1}} g_\lambda^\nu) + \tau R_\tau(\chi_{<|\tau|^{-1}} \zeta_{\lambda-1}^\nu) + \tau^\lambda (R_\tau h_\nu).$$

Therefore after applying R_τ to both sides of (5.28), we see terms on the right hand side that are of the form (5.29) with $1 \leq m \leq \lambda - 2$ become

$$\tau^m R_\tau(\zeta_{\lambda-m}^\nu e^{-i\tau\langle r \rangle}) = \tau^m R_\tau(\chi_{<|\tau|^{-1}} \zeta_{\lambda-m}^\nu) + \tau^{m+1} R_\tau(\chi_{<|\tau|^{-1}} \zeta_{\lambda-m-1}^\nu) + \tau^\lambda (R_\tau h_\nu), \quad (5.31)$$

and we can proceed inductively for the first 2 terms, while the last term is expected to appear in E_λ^ν .

Next we consider the term $-2i\tau[(\partial_r + \frac{1}{r})R_0 g_\lambda^\nu]e^{-i\tau\langle r \rangle}$ in (5.28) using Proposition 4.3. We note $(\partial_r + \frac{1}{r})\langle r \rangle^{-1} = 0$ so the term $2i\tau(\partial_r + \frac{1}{r})c_0\langle r \rangle^{-1}$ vanishes. It is left to consider the terms

- (A) $-2i\tau[(\partial_r + \frac{1}{r})c_j \nabla^j \langle r \rangle^{-1}]e^{-i\tau\langle r \rangle}$ for $1 \leq j \leq \lambda - 2$
- (B) $-2i\tau[(\partial_r + \frac{1}{r})e_j(r)(\nabla^j \langle r \rangle^{-1})\langle r \rangle^{j-\lambda+1}]e^{-i\tau\langle r \rangle}$ for $1 \leq j \leq \lambda - 2$

- (C) $-2i\tau[(\partial_r + \frac{1}{r})e_0\langle r \rangle^{-\lambda}]e^{-i\tau\langle r \rangle}$ (Here the cases $\lambda \leq \kappa$ and $\lambda = \kappa + 1$ must be considered separately since $e_0 \in \ell^1 S(1)$ for $\lambda \leq \kappa$ and $e_0 \in S(1)$ for $\lambda = \kappa + 1$ by Proposition 4.3.)
- (D) $-2i\tau[(\partial_r + \frac{1}{r})d(r)\nabla^{\lambda-1}\langle r \rangle^{-1}]e^{-i\tau\langle r \rangle}$
- (E) $-2i\tau[(\partial_r + \frac{1}{r})q(x)]e^{-i\tau\langle r \rangle}$.

For term E we have $q \in Z^{\nu-2, \lambda-2}$, so writing $\partial_r = r^{-1}S_r$ we see $(\partial_r + \frac{1}{r})q \in Z^{\nu-3, \lambda-1}$. Thus we can write $E = \tau\zeta_{\lambda-1}^{\nu-3}e^{-i\tau\langle r \rangle}$.

For term B we have $e_j \in \ell^1 S(1)$ so $(\partial_r + \frac{1}{r})e_j(\nabla^j\langle r \rangle^{-1})\langle r \rangle^{j-\lambda+1} \in \ell^1 S(r^{-\lambda-1})$. Since $\ell^1 S(r^{-2}) \subseteq Z^{N,0}$ for any N , we can write $B = \tau\zeta_{\lambda-1}^{\nu-3}e^{-i\tau\langle r \rangle}$.

To handle term A we multiply both sides of (5.11) by the constant c_j to find

$$A = \sum_{a=1}^j \left(P_\tau(\tau^a F_a e^{-i\tau\langle r \rangle}) + (\tau^a \zeta_\kappa^\nu + \tau^{a+1} \zeta_\kappa^\nu + \tau^{a+2} \zeta_{\kappa-1}^\nu) e^{-i\tau\langle r \rangle} \right)$$

where $|F_a| \lesssim \langle r \rangle^{-1}$. Here we absorbed the constant into the functions ζ_κ^ν and $\zeta_{\kappa-1}^\nu$.

Next we consider term D. Using Lemma 5.4 and the fact that $S_r d \in \ell^1 S(1)$, we calculate

$$\begin{aligned} D &= -2i\tau(\partial_r d)(\nabla^{\lambda-1}\langle r \rangle^{-1})e^{-i\tau\langle r \rangle} - d2i\tau(\partial_r + \frac{1}{r})(\nabla^{\lambda-1}\langle r \rangle^{-1}) \\ &= \tau\zeta_{\lambda-1}^\nu e^{-i\tau\langle r \rangle} + \sum_{a=1}^{\lambda-1} dP_\tau((-i\tau)^a F_a e^{-i\tau\langle r \rangle}) + (\tau^a \zeta_\kappa^\nu + \tau^{a+1} \zeta_\kappa^\nu + \tau^{a+2} \zeta_{\kappa-1}^\nu) e^{-i\tau\langle r \rangle} \\ &= \tau\zeta_{\lambda-1}^\nu e^{-i\tau\langle r \rangle} + \sum_{a=1}^{\lambda-1} P_\tau(d(-i\tau)^a F_a e^{-i\tau\langle r \rangle}) - [P_\tau, d](-i\tau)^a F_a e^{-i\tau\langle r \rangle} \\ &\quad + \left(\tau^a \zeta_\kappa^\nu + \tau^{a+1} \zeta_\kappa^\nu + \tau^{a+2} \zeta_{\kappa-1}^\nu \right) e^{-i\tau\langle r \rangle}. \end{aligned}$$

We find by direct calculation $[P_\tau, d]F_a e^{-i\tau\langle r \rangle} = \left(\rho_\ell^{-\lambda+a-2} + \tau\rho_\ell^{-\lambda+a-1} \right) e^{-i\tau\langle r \rangle}$. so that

$$D = \sum_{a=1}^{\lambda-1} P_\tau \left(\tau^a F_a e^{-i\tau\langle r \rangle} \right) + (\tau^a (\zeta_\kappa^\nu + \zeta_{\lambda-a}^\nu) + \tau^{a+1} (\zeta_\kappa^\nu + \zeta_{\lambda-a-1}^\nu) + \tau^{a+2} \zeta_{\kappa-1}^\nu) e^{-i\tau\langle r \rangle}$$

since $\ell^1 S(r^{-2}) \subset Z^{N,0}$ for any N .

Substituting our expressions for terms A, B, D, and E into (5.28) and simplifying yields

$$\begin{aligned} P_\tau(E_\lambda^\nu) &= \chi_{>|\tau|^{-1}} g_\lambda^\nu + \sum_{a=1}^{\lambda-1} \left(P_\tau(\tau^a F_a e^{-i\tau\langle r \rangle}) \right) + \tau\chi_{<|\tau|^{-1}} \zeta_{\lambda-1}^\nu + \left(\tau\zeta_{\lambda-1}^{\nu-3} + \tau^2\zeta_{\lambda-2}^{\nu-2} \right) e^{-i\tau\langle r \rangle} \\ &\quad + \sum_{m=3}^{\lambda} \tau^m \zeta_{\lambda-m}^\nu e^{-i\tau\langle r \rangle} + \underbrace{2\tau(\partial_r + \frac{1}{r})(e_0\langle r \rangle^{-\lambda})e^{-i\tau\langle r \rangle}}_{(C)} \end{aligned} \tag{5.32}$$

for $1 \leq \lambda \leq \kappa + 1$.

For term C we consider the cases $2 \leq \lambda \leq \kappa$ and $\lambda = \kappa + 1$ separately.

Case 2(a): $2 \leq \lambda \leq \kappa$

Here we have $e_0 \in \ell^1 S(1)$ so $e_0 \langle r \rangle^{-\lambda} \in \ell^1 S(r^{-\lambda})$ and we find

$$C = 2\tau(\partial_r + \frac{1}{r})(e_0(r)\langle r \rangle^{-\lambda})e^{-i\tau\langle r \rangle} = \tau\zeta_{\lambda-1}^\nu e^{-i\tau\langle r \rangle}. \quad (5.33)$$

We can absorb C into the term $\tau\zeta_{\lambda-1}^{\nu-3}e^{-i\tau\langle r \rangle}$ in (5.32). Then applying R_τ to both sides of (5.32) and using (5.30) and (5.31) yields

$$\begin{aligned} E_\lambda^\nu &= R_\tau(\chi_{>|\tau|^{-1}}g_\lambda^\nu) + \sum_{a=1}^{\lambda-1}(\tau^a F_a e^{-i\tau\langle r \rangle}) + \tau R_\tau(\chi_{<|\tau|^{-1}}\zeta_{\lambda-1}^{\nu-3}) + \tau^2 R_\tau(\chi_{<|\tau|^{-1}}\zeta_{\lambda-2}^{\nu-3}) \\ &\quad + \sum_{m=3}^{\lambda-1} \tau^m R_\tau(\chi_{<|\tau|^{-1}}\zeta_{\lambda-m}^\nu) + \tau^\lambda(R_\tau h_{\nu-3}). \end{aligned} \quad (5.34)$$

Part 1 of the proposition then follows by induction in λ and the established base case for $\lambda = 1$. We note the term $\tau R_\tau(\chi_{<|\tau|^{-1}}\zeta_{\lambda-1}^{\nu-3})$ leads to the loss of regularity for h .

Case 2(b): $\lambda = \kappa + 1$

Since $e_0 \langle r \rangle^{-\kappa-1} \in S_{rad}(r^{-\kappa-1})$, we can use Lemma 5.5. Combining (5.13) and (5.32) then applying R_τ yields

$$\begin{aligned} E_{\kappa+1}^\nu &= R_\tau(\chi_{>|\tau|^{-1}}g_{\kappa+1}^\nu) + \sum_{a=1}^{\kappa}(\tau^a F_a e^{-i\tau\langle r \rangle}) + \tau R_\tau(\chi_{<|\tau|^{-1}}\zeta_\kappa^\nu + \zeta_\kappa^{\nu-3}e^{-i\tau\langle r \rangle}) \\ &\quad + \tau^2 R_\tau(\zeta_{\kappa-1}^{\nu-2}e^{-i\tau\langle r \rangle}) + \sum_{m=3}^{\kappa} \tau^m R_\tau(\zeta_{\kappa+1-m}^\nu e^{-i\tau\langle r \rangle}) + \tau^\kappa \varepsilon(r, \tau)e^{-i\tau\langle r \rangle} + \tau^{\kappa+1}R_\tau h_\nu. \end{aligned} \quad (5.35)$$

Part 1 of the proposition, (5.30), and (5.31) then give

$$\begin{aligned} E_{\kappa+1}^\nu &= R_\tau(\chi_{>|\tau|^{-1}}g_{\kappa+1}^\nu) + \sum_{m=1}^{\kappa} \left(\tau^m (F_m + R_0 \zeta_{\lambda-m}^{\nu-3m}) e^{-i\tau\langle r \rangle} \right) + \tau^\kappa \varepsilon(r, \tau) e^{-i\tau\langle r \rangle} \\ &\quad + \tau^{\kappa+1}(R_\tau h_{\nu-3\kappa-3}) \end{aligned} \quad (5.36)$$

as desired. This concludes the proof of the proposition. $\square$

6. POINTWISE RESOLVENT BOUNDS

In this section we establish the pointwise resolvent bounds that will be used in the proof of the main theorem (see Propositions 6.5 and 6.8). The results of this section do not improve on the results in [33], but we do track the required regularity more precisely.

Our argument uses the Sobolev embedding

$$\|\phi\|_{L^\infty(\mathbb{S}^2)} \lesssim \|\phi\|_{L^2(\mathbb{S}^2)} + \|\Omega^2 \phi\|_{L^2(\mathbb{S}^2)} \quad (6.1)$$

to obtain useful $L_r^2 L_\omega^\infty(A_m)$ bounds on g and $R_\tau g$. For reference we begin with two preliminary lemmas resulting from a straightforward application of (6.1).

Lemma 6.1. *If ϕ satisfies*

$$\|\langle r \rangle^p \phi\|_{L^2(A_m)} + \sum_{|\alpha|=2} \|\langle r \rangle^p \Omega^\alpha \phi\|_{L^2(A_m)} \lesssim 1$$

then $2^{m(1+p)} \|\phi\|_{L_r^2 L_\omega^\infty(A_m)} \lesssim 1$.

Proof. A change of coordinates yields $\|\langle r \rangle^p \phi\|_{L^2(A_m)}^2 \approx 2^{2m(1+p)} \|\phi\|_{L_r^2 L_\omega^2(A_m)}^2$. Then from the Sobolev embedding (6.1) we obtain

$$2^{m(1+p)} \|\phi\|_{L_r^2 L_\omega^\infty(A_m)} \lesssim \|\langle r \rangle^p \phi\|_{L^2(A_m)} + \|\langle r \rangle^p \Omega^2 \phi\|_{L^2(A_m)} \lesssim 1. \quad (6.2)$$

□

In the following proofs we use the notation Q_ℓ to denote any operator of the form

$$\tau(\partial_i h^i + h^i \partial_i) + \partial_i h^{ij} \partial_j + h_\ell, \quad h^i, h^{ij} \in \ell^1 S(r^{-\kappa}), \quad h_\ell \in \ell^1 S(r^{-\kappa-2}) \quad (6.3)$$

and use Q_r to denote an operator of the form

$$h^\omega \Delta_\omega + h_r, \quad h^\omega, h_r \in S_{rad}(r^{-\kappa-2}). \quad (6.4)$$

Lemma 6.2. *If $\phi, S_r \phi, \Omega^2 \phi, \Omega^2 S_r \phi \in \mathcal{LE}^*$, then $|\phi| \lesssim \langle r \rangle^{-2}$. Furthermore, if $\phi \in Z^{n,q}$, then $|\partial_r^p \phi| \lesssim \langle r \rangle^{-2-p-q}$ for $p \leq n-3$.*

Proof. Fix m . By assumption, $\langle r \rangle^{\frac{1}{2}} \phi, \langle r \rangle^{\frac{1}{2}} \Omega^2 \phi, \langle r \rangle^{\frac{3}{2}} \partial_r \phi, \langle r \rangle^{\frac{3}{2}} \Omega^2 \partial_r \phi \in L^2(A_m)$, so by (6.1), we have $2^{\frac{3m}{2}} \|\phi\|_{L_r^2 L_\omega^\infty(A_m)} \lesssim 1$ and $2^{\frac{5m}{2}} \|\partial_r \phi\|_{L_r^2 L_\omega^\infty(A_m)} \lesssim 1$. Using the Fundamental Theorem of Calculus and Cauchy-Schwarz, we find pointwise bounds on ϕ :

$$\|\phi\|_{L^\infty(A_m)} \lesssim 2^{-\frac{m}{2}} \|\phi\|_{L_r^2 L_\omega^\infty(A_m)} + 2^{\frac{m}{2}} \|\partial_r \phi\|_{L_r^2 L_\omega^\infty(A_m)} \lesssim 2^{-2m}.$$

Thus $|\phi| \lesssim \langle r \rangle^{-2}$ since m was arbitrary.

Now take $\phi \in Z^{n,q}$. We can write $\langle r \rangle^p \partial_r^p S_r^p$ as a linear combination of S_r^k for $k \leq p$ so $|\langle r \rangle^{q+p} \partial_r^p \phi| \lesssim \sum_{k \leq p} |\langle r \rangle^q S_r^k \phi|$. Since $[\Omega, r] = 0$ and $[S_r, r] = r$, it follows by the definition of $Z^{n,q}$ that $\Omega^j S_r^k \langle r \rangle^{q+p} \partial_r^p \phi \in \mathcal{LE}^*$ when $j+k \leq 3$ if $p \leq n-3$. Therefore by the first part of the proposition we have $|\langle r \rangle^{q+p} \partial_r^p \phi| \lesssim \langle r \rangle^{-2}$. □

The following calculation will be useful for the remaining lemmas and propositions in this section. Writing

$$(\partial_r^2 + \tau^2) = P_\tau - (2r^{-1} \partial_r + r^{-2} \Delta_\omega + Q_\ell + Q_r),$$

where Q_ℓ, Q_r are as in (6.3) and (6.4), we obtain

$$\begin{aligned} (\partial_r^2 + \tau^2) r v_{ijk} &= r(\partial_r^2 + \tau^2) v_{ijk} + 2\partial_r v_{ijk} \\ &= r P_\tau v_{ijk} - r^{-1} \Delta_\omega v_{ijk} - r(Q_\ell + Q_r) v_{ijk}. \end{aligned} \quad (6.5)$$

Commuting P_τ with $T^i \Omega^j S^k$ yields

$$P_\tau v_{ijk} = g_{ij \leq k} + Q_\ell(v_{<ijk} + v_{\leq i < jk} + v_{\leq i \leq j < k}) + Q_r(v_{<ijk} + v_{\leq ij < k}). \quad (6.6)$$

Then we rewrite the first term of (6.5) using (6.6) to find

$$(\partial_r^2 + \tau^2) r v_{ijk} = -r^{-1} \Delta_\omega v_{ijk} + r(Q_\ell + Q_r) v_{\leq i \leq j \leq k} + r g_{\leq i \leq j \leq k}. \quad (6.7)$$

In Proposition 6.3 we state the same pointwise bounds and outgoing radiation condition established in [33, Proposition 16]. However, we obtain different numerology for the number of vector fields that can be applied to $v = R_\tau g$ so that the results hold. Thus we offer a concrete justification for the change in the vector field numerology but provide only a brief outline of the argument. We will use notation as in Proposition 3.5 so that M indicates the regularity assumed for g . We take $v_{ijk} = T^i \Omega^j S^k v$ and $g_{ijk} = T^i \Omega^j S^k g$. Similarly, we write $v_{<i<j<k} = T^{<i} \Omega^{<j} S^{<k} v$ and use analogous notation for g .

Proposition 6.3. *Assume $\Im \tau \leq 0$. Let $g \in \mathcal{LE}^*$ satisfy (3.9) and possibly depend on τ . Set $v = R_\tau g$.*

(i) *If $|\tau| \gtrsim 1$, then*

$$|T^i \Omega^j S^k v(\tau)| \lesssim (|\tau| \langle r \rangle)^{-1}, \quad i + 4j + 16k \leq M - 20. \quad (6.8)$$

(ii) *If $|\tau| \lesssim 1$, then*

$$|T^i \Omega^j S^k v(\tau)| \lesssim \begin{cases} \min\{1, (|\tau| \langle r \rangle)^{-1}\} & i = 0 \\ \langle r \rangle^{-1} & i \geq 1 \end{cases} \quad i + 4j + 16k \leq M - 20. \quad (6.9)$$

(iii) *If $\tau \in \mathbb{R} \setminus \{0\}$, then we have the outgoing radiation condition:*

$$\lim_{|x| \rightarrow \infty} r(\partial_r + i\tau) T^i \Omega^j S^k v(\tau) = 0, \quad i + 4j + 16k \leq M - 20. \quad (6.10)$$

proof summary. (i) and (ii) The estimate

$$\sum_m 2^{\frac{m}{2}} \|(\partial_r^2 + \tau^2)(rv_{ijk})\|_{L_r^2 L_\infty(A_m)} \lesssim 1 \quad (6.11)$$

implies

$$|v_{ijk}| \lesssim (\langle r \rangle |\tau|)^{-1} \quad \text{and} \quad |\partial_r v_{ijk}| \lesssim \langle r \rangle^{-1} \quad (6.12)$$

using the fundamental solution for $(\partial_r^2 + \tau^2)$, which is given by $\tau^{-1} e^{-i\tau|s|}$.

The pointwise bounds for $|\tau| \gtrsim 1$ and for the case $\langle r \rangle \gtrsim |\tau|^{-1}$ when $|\tau| \lesssim 1$ are obtained using (6.11). To show (6.11) holds, we bound each term on the right hand side of (6.7) by applying Lemma 6.1 to the assumption $\|g_{ijk}\|_{\mathcal{LE}^*} \lesssim 1$ for $i + 4j + 16k \leq M$ and the resulting fact $\|v_{ijk}\|_{\mathcal{LE}_\tau} \lesssim 1$ for $i + 4j + 16 \leq M - 4$ (which holds by Proposition 3.5).

To see where the vector field loss occurs, consider the term $r^{-1} \Delta_\omega v_{ijk}$ on the right hand side of (6.7). We wish to show this term satisfies (6.11). Using (6.2) in Lemma 6.1, it suffices to show

$$\sum_m \|\langle r \rangle^{-\frac{3}{2}} \Delta_\omega v_{ijk}\|_{L^2(A_m)} + \|\langle r \rangle^{-\frac{3}{2}} \Omega^2(\Delta_\omega v_{ijk})\|_{L^2(A_m)} \lesssim 1. \quad (6.13)$$

When $|\tau| \gtrsim 1$, we have

$$\|\langle r \rangle^{-\frac{1}{2}} v_{ijk}\|_{L^2(A_m)} \lesssim |\tau| \|\langle r \rangle^{-\frac{1}{2}} v_{ijk}\|_{L^2(A_m)} \lesssim \|v_{ijk}\|_{\mathcal{LE}_\tau} \lesssim 1 \quad i + 4j + 16k < M - 4. \quad (6.14)$$

Replacing Δ_ω by $\sum_{|\alpha|=2} \Omega^\alpha$ and using (6.14) yields

$$\begin{aligned} & \sum_m \|\langle r \rangle^{-\frac{3}{2}} \Delta_\omega v_{ijk}\|_{L^2(A_m)} + \|\langle r \rangle^{-\frac{3}{2}} \Omega^2(\Delta_\omega v_{ijk})\|_{L^2(A_m)} \\ &= \sum_m \|\langle r \rangle^{-\frac{3}{2}} v_{i(j+2)k}\|_{L^2(A_m)} + \|\langle r \rangle^{-\frac{3}{2}} v_{i(j+4)k}\|_{L^2(A_m)} \\ &\lesssim 1, \quad i + 4j + 16k \leq M - 20. \end{aligned}$$

This shows the $M - 20$ vector field loss (as compared to the erroneously stated $M - 12$ loss in [33]).

When $|\tau| \lesssim 1$ and $|\tau|^{-1} \gtrsim \langle r \rangle$, then (6.12) is insufficient since $(|\tau| \langle r \rangle)^{-1}$ is unbounded. The advantage in this case is that the $(\langle r \rangle^{-1} + |\tau|)^{-1}$ weight in the second order term of $\|v_{ijk}\|_{\mathcal{LE}_\tau}$ is bounded below by $\langle r \rangle$ when $\langle r \rangle \lesssim |\tau|^{-1}$. $\square$

We will use Proposition 6.3 to establish pointwise bounds on $(\tau \partial_\tau)^p (v e^{i\tau r})$. Note $\tau \partial_\tau (v e^{i\tau r}) = [(-S + r(\partial_r + i\tau))v] e^{i\tau r}$. This motivates the following lemma, which will be used to prove the subsequent proposition stating the pointwise bounds on $(\tau \partial_\tau)^p (v e^{i\tau r})$ for $|\tau| \gtrsim 1$. We remark that while the above calculation shows we are primarily concerned with $(\partial_r + i\tau)^p v_{00k}$, our methods will generate T and Ω vector fields as we induct in k , so we handle $(\partial_r + i\tau)^p v_{ijk}$.

Lemma 6.4. *Let $g \in Z^{n,q}$. If $\tau \in \mathbb{R}$ and $|\tau| \gtrsim 1$, then $v = R_\tau g$ satisfies the pointwise bounds*

$$|(\partial_r + i\tau)^p v_{ijk}| \lesssim |\tau|^{p-1} \langle r \rangle^{-p-1}, \quad p \leq q, \quad p \leq n-3, \quad \text{and} \quad i + 4j + 16k \leq n - 20 - 8p. \quad (6.15)$$

Proof. Note $i + j + k < i + 4j + 16k < n$ so $g \in Z^{n,q}$ implies g satisfies (3.9) with $M = n$, and the results of Proposition 6.3 apply with $M = n$.

When $p = 0$, (6.15) follows from (6.8).

Now let $p = 1$. All but the last term on the right hand side of (6.7) are pointwise bounded by $\langle r \rangle^{-2}$ using (6.8) and the fact that $\kappa \geq 2$. We replace $\Delta_\omega v_{ijk}$ by $v_{i(j+2)k}$ so the bounds hold when $i + 4j + 16k \leq n - 28$. For the final term, by Lemma 6.2 we have $|\partial_r^k g| \lesssim \langle r \rangle^{-2-k-q}$ when $k \leq n - 3$, so $|rg| \lesssim \langle r \rangle^{-2}$ since $1 = p \leq q$. Thus we have $|(\partial_r^2 + \tau^2)(rv_{ijk})| \lesssim \langle r \rangle^{-2}$.

We rewrite $\partial_r^2 + \tau^2 = (\partial_r - i\tau)(\partial_r + i\tau)$ and use an integrating factor to write

$$\partial_r [((\partial_r + i\tau)rv_{ijk})e^{-i\tau r}] = ((\partial_r^2 + \tau^2)rv_{ijk})e^{-i\tau r}$$

By (6.8) and (6.10), we have $\lim_{r \rightarrow \infty} (\partial_r + i\tau)(rv_{ijk}) = 0$, so we can integrate from infinity to find

$$|(\partial_r + i\tau)(rv_{ijk})| \lesssim \int_r^\infty |\langle s \rangle^{-2}| ds = \langle r \rangle^{-1}.$$

It follows that $|(\partial_r + i\tau)v_{ijk}| \lesssim \langle r \rangle^{-2}$, as desired.

We proceed by induction. Fix p and assume $|(\partial_r + i\tau)^a v_{ijk}| \lesssim |\tau|^{a-1} \langle r \rangle^{-a-1}$ for $a < p$ when $i + 4j + 16 \leq n - 20 - 8a$. Applying $(\partial_r + i\tau)^{p-1}$ to (6.7) we find

$$\begin{aligned} & (\partial_r - i\tau)(\partial_r + i\tau)^p (rv_{ijk}) \\ &= \sum_{m=0}^{p-1} \left((-1)^{p-m+1} c_m r^{-(p-m)} (\partial_r + i\tau)^m \Delta_\omega v_{ijk} \right) + r(\partial_r + i\tau)^{p-1} (Q_\ell + Q_r) v_{\leq i \leq j \leq k} \\ & \quad + C(\partial_r + i\tau)^{p-2} (Q_\ell + Q_r) v_{\leq i \leq j \leq k} + (r(\partial_r + i\tau)^{p-1} + C(\partial_r + i\tau)^{p-2}) g_{\leq i \leq j \leq k}. \end{aligned} \quad (6.16)$$

Each term on the right hand side of (6.16) is bounded in magnitude by $|\tau|^{p-1} \langle r \rangle^{-p-1}$. The first term is bounded, by the inductive hypothesis, when $i + 4j + 16k \leq n - 20 - 8p$. For the Q_ℓ and Q_r terms, we commute $(\partial_r + i\tau)$ with the coefficients of the operators and view the derivatives as vector fields. The bounds then follow using our assumption $\kappa \geq 2$ once we note

$$[(\partial_r + i\tau)^{p-1}, \rho_\bullet^{-\kappa}] = \sum_{m=1}^{p-1} c_m \rho_\bullet^{-\kappa-m} (\partial_r + i\tau)^{p-1-m} \quad (6.17)$$

for $\bullet \in \{\ell, r\}$. For the last g_{ijk} terms we use Lemma 6.2 (which requires our assumption $p \leq n - 3$) to find

$$\begin{aligned} & |r(\partial_r + i\tau)^{p-1} g_{\leq i \leq j \leq k}| + |(\partial_r + i\tau)^{p-2} g_{\leq i \leq j \leq k}| \\ &= \left| r \sum_{m=0}^{p-1} c_m (i\tau)^m \partial_r^{p-1-m} g_{\leq i \leq j \leq k} \right| + \left| \sum_{m=0}^{p-2} (i\tau)^m c_m \partial_r^{p-2-m} g_{\leq i \leq j \leq k} \right| \\ &\lesssim |\tau|^{p-1} \langle r \rangle^{-1-p}. \end{aligned}$$

The last inequality holds since we assume $p \leq q$.

Now we have

$$|(\partial_r - i\tau)(\partial_r + i\tau)^p (rv_{ijk})| \lesssim |\tau|^{p-1} \langle r \rangle^{-p-1}$$

Integrating as before then yields (6.15). $\square$

Proposition 6.5 establishes the pointwise bounds we will use in the proof of the main theorem. This corresponds to Proposition 17 in [33]. Note we use Tataru's method of proof and correct an error in the proposition statement.

Proposition 6.5. *Let $g \in Z^{n,q}$. If $\tau \in \mathbb{R}$ and $|\tau| \gtrsim 1$, then $v = R_\tau g$ satisfies the pointwise bounds*

$$|(\tau \partial_\tau)^p (v e^{i\tau \langle r \rangle})| \lesssim |\tau|^{p-1} \langle r \rangle^{-1}, \quad p \leq q \quad \text{and} \quad 16p \leq n - 20. \quad (6.18)$$

Proof. If $p = 0$, then (6.18) follows from (6.8).

To handle $p = 1$, we write $(\tau \partial_\tau) (v e^{i\tau \langle r \rangle}) = (-Sv + r(\partial_r + i\tau)v) e^{i\tau \langle r \rangle}$. Then $|Sv| \lesssim |\tau|^{-1} \langle r \rangle^{-1}$ using (6.8), and $|r(\partial_r + i\tau)v| \lesssim \langle r \rangle^{-1}$ using (6.15). Both (6.8) and (6.15) hold under our assumptions $p \leq 16p \leq n - 20$, which implies $p \leq n - 3$ since p is nonnegative).

For general p , we write

$$(\tau \partial_\tau)^p \left(v e^{i\langle r \rangle \tau} \right) = \left(\sum_{j=0}^p \sum_{\ell=0}^j c_{j\ell} r^\ell (\partial_r + i\tau)^\ell (-S)^{p-j} v \right) e^{i\langle r \rangle \tau}.$$

Each term on the right hand side is bounded by $|\tau|^{\ell-1} \langle r \rangle^{-1} \lesssim |\tau|^{p-1} \langle r \rangle^{-1}$ using (6.15) and our assumption $16p \leq n - 20$. $\square$

Next we find pointwise resolvent bounds for $|\tau| \lesssim 1$. In this case we are interested in the term $R_\tau h_{\nu-3\kappa-3}$ in our expression for $R_\tau g$ in Proposition 5.6. The terms included in h in the proof of Proposition 5.6 depend on τ , so we consider a τ dependent function g .

Lemma 6.6. *Let $g \in \mathcal{LE}^*$, possibly depending on τ , satisfy*

$$\|\langle r \rangle^q (\partial_r + i\tau)^q T^i \Omega^j S^k g\|_{\mathcal{LE}^*} \lesssim 1, \quad q + i + 4j + 16k \leq n. \quad (6.19)$$

If $\tau \in \mathbb{R}$, $|\tau| \lesssim 1$, and $p < n - 3$ then

$$|\partial_r^p (g e^{i\tau r})| \lesssim r^{-p-2}, \quad (6.20)$$

$$|(\partial_r + i\tau)^p g| \lesssim r^{-p-2}, \quad (6.21)$$

and

$$|\partial_r^p g| \lesssim r^{-2}. \quad (6.22)$$

Furthermore, if $|\tau r| \lesssim 1$ then

$$|\partial_r^p g| \lesssim r^{-p-2}. \quad (6.23)$$

Proof. To prove (6.20) we calculate $S_r(g e^{i\tau r}) = (r(\partial_r + i\tau)g) e^{i\tau r}$ so that

$$S_r^k(g e^{i\tau r}) = \sum_{m=1}^k c_m (r^m (\partial_r + i\tau)^m g) e^{i\tau r}. \quad (6.24)$$

We also have $\Omega^j g e^{i\tau r} = (\Omega^j g) e^{i\tau r}$. Finally we calculate

$$|T^i g e^{i\tau r}| = \left| \sum_{m=0}^i c_m (T^{i-m} g) T^m e^{i\tau r} \right| \lesssim \sum_{m=0}^i c_m |T^m g|$$

since we assumed $|\tau| \lesssim 1$. Therefore (6.19) implies $g e^{i\tau r} \in Z^{n,0}$ and (6.20) follows by Lemma 6.2. Then (6.21) follows from (6.20) since $\partial_r^p (g e^{i\tau r}) = ((\partial_r + i\tau)^p g) e^{i\tau r}$.

To prove (6.22), note the case $p = 0$ follows from (6.21). Now assume (6.22) holds for $a < p$. We calculate

$$(\partial_r + i\tau)^p = \sum_{m=0}^p c_m \partial_r^m (i\tau)^{p-m} = \partial_r^p + \sum_{m=0}^{p-1} c_m \partial_r^m (i\tau)^{p-m} \quad (6.25)$$

so that

$$|\partial_r^p g| \lesssim |(\partial_r + i\tau)^p g| + \sum_{m=0}^{p-1} |\partial_r^m g|$$

since we assume $|\tau| \lesssim 1$. Then (6.22) follows from (6.21) and the inductive hypothesis.

Finally, to prove (6.23), we see the case $p = 0$ follows from (6.21). Now assume (6.23) holds for $a < p$. By (6.25) we have

$$|r^p \partial_r^p g| \lesssim |r^p (\partial_r + i\tau)^p g| + \sum_{m=0}^{p-1} r^m |\partial_r^m (i\tau r)^{p-m}| \lesssim \langle r \rangle^{-2}$$

where the last inequality follows from (6.21), the assumption $|\tau r| \lesssim 1$, and the inductive hypothesis. $\square$

Lemma 6.7. *Let $g \in Z^{n,0}$, possibly depending on τ , satisfy*

$$\|\langle r \rangle^q (\partial_r + i\tau)^q T^i \Omega^j S^k g\|_{\mathcal{LE}^*} \lesssim 1, \quad q + i + 4j + 16k \leq n. \quad (6.26)$$

If $\tau \in \mathbb{R}$ and $|\tau| \lesssim 1$ then $v = R_\tau g$ satisfies

$$|(\partial_r + i\tau)^p v_{ijk}| \lesssim |\tau|^{-1} \langle r \rangle^{-p-1}, \quad p \leq n-3 \quad \text{and} \quad i + 4j + 16k \leq n - 20 - 8p. \quad (6.27)$$

Proof. If $p = 0$, (6.27) follows from (6.9).

To handle $p = 1$, we again use (6.7). All but the last term are bounded by $|\tau|^{-1} \langle r \rangle^{-2}$ using (6.9). We note the assumption $g \in Z^{n,0}$ does not allow us to use Lemma 6.2. g may depend on τ , so $Sg \neq S_\tau g$. Take $\psi = (\partial_r + i\tau)(rv_{ijk})$. Then the radiation condition (6.10) allows us to integrate from infinity as before to find

$$\begin{aligned} |\psi(r_0)e^{-i\tau r_0}| &= \left| \int_{r_0}^{\infty} (r^{-1} \Delta_\omega v_{ijk} + r(Q_\ell + Q_r)v_{\leq i \leq j \leq k}) e^{-i\tau r} dr + \int_{r_0}^{\infty} r g_{\leq i \leq j \leq k} e^{-i\tau r} dr \right| \\ &\lesssim |\tau|^{-1} \langle r_0 \rangle^{-1} + \left| \int_{r_0}^{\infty} r g_{\leq i \leq j \leq k} e^{-i\tau r} dr \right|. \end{aligned}$$

For the last term we integrate by parts and use Lemma 6.6 to calculate

$$\begin{aligned} &\left| \int_{r_0}^{\infty} r g_{\leq i \leq j \leq k} e^{i\tau r} \partial_r \left(\frac{e^{-2i\tau r}}{-2i\tau} \right) dr \right| \\ &= \left| \frac{-i}{2\tau} r_0 g_{\leq i \leq j \leq k}(r_0) e^{-i\tau r_0} + \frac{1}{2i\tau} \int_{r_0}^{\infty} e^{-2i\tau r} (g e^{i\tau r} + r \partial_r (g e^{i\tau r})) dr \right| \\ &\lesssim |\tau|^{-1} \langle r_0 \rangle^{-1}. \end{aligned}$$

Thus we have

$$|r(\partial_r + i\tau)v_{ijk}| = |(\partial_r + i\tau)(rv_{ijk}) - v_{ijk}| \lesssim |\tau| \langle r \rangle^{-1}$$

so that $|(\partial_r + i\tau)v_{ijk}| \lesssim |\tau|^{-1} \langle r \rangle^{-2}$, as desired.

We proceed by induction. Fix p and assume $|(\partial_r + i\tau)^a v_{ijk}| \lesssim |\tau|^{-1} \langle r \rangle^{-a-1}$ for $a < p$. We again use (6.16). All but the g_{ijk} terms are bounded by $|\tau|^{-1} \langle r \rangle^{-p-1}$ using the inductive hypothesis and (6.17). For the g_{ijk} terms we integrate by parts as in the $p = 1$ case to find

$$\left| \int_{r_0}^{\infty} (r(\partial_r + i\tau)^{p-1} + C(\partial_r + i\tau)^{p-2}) g_{\leq i \leq j \leq k} e^{-i\tau r} dr \right| \lesssim |\tau|^{-1} \langle r_0 \rangle^{-p}.$$

Then (6.27) follows. $\square$

In Proposition [6.8](#) we establish the pointwise resolvent bounds which will be used in the proof of the main theorem for small τ . The result is the same as that in Proposition 18 in [\[33\]](#) with a more precise statement on the regularity requirements.

Proposition 6.8. *Let $g \in Z^{n,0}$, possibly depending on τ , satisfy*

$$\|\langle r \rangle^q (\partial_r + i\tau)^q T^i \Omega^j S^k g\|_{\mathcal{L}\mathcal{E}^*} \lesssim 1, \quad q + i + 4j + 16k \leq n. \quad (6.28)$$

If $\tau \in \mathbb{R}$ and $|\tau| \lesssim 1$ then $v = R_\tau g$ satisfies the following pointwise bounds:

(1) *If $\langle r \rangle \lesssim |\tau|^{-1}$, then*

$$|(\tau \partial_\tau)^p v| \lesssim 1, \quad 16p \leq n - 20. \quad (6.29)$$

(2) *If $\langle r \rangle \gtrsim |\tau|^{-1}$, then*

$$|(\tau \partial_\tau)^p (v e^{i\tau \langle r \rangle})| \lesssim (|\tau| \langle r \rangle)^{-1}, \quad 16p \leq n - 20. \quad (6.30)$$

Proof. **1. Small r :** $\langle r \rangle \lesssim |\tau|^{-1}$

We write $\tau \partial_\tau = -S + r \partial_r$ and find

$$(\tau \partial_\tau)^p v = \sum_{m=0}^p c_m (r \partial_r)^{p-m} (-S)^m v.$$

Since $(r \partial_r)^p = \sum_{j=0}^p c_j r^j \partial_r^j$, it is sufficient to show $|r^p \partial_r^p v_{00k}| \lesssim 1$ for $16k \leq n - 20 - 16p$. As before, we will use [\(6.7\)](#), which introduces Ω and T vector fields, so we will instead bound $|r^p \partial_r^p v_{ijk}|$ then set $i, j = 0$.

When $p = 0$, we have $|v_{ijk}| \lesssim 1$ by [\(6.9\)](#) for $i + 4j + 16k \leq n - 20$. When $p = 1$, we have $|\partial_r v_{ijk}| \lesssim |\nabla v_{ijk}| \lesssim \langle r \rangle^{-1}$ when $i + 4j + 16k \leq n - 20$ by [\(6.9\)](#).

Fix p and assume $|r^a \partial_r^a v_{ijk}| \lesssim 1$ for $a < p$ when $i + 4j + 16k \leq n - 20 - 16a$. Applying $r^{p-2} \partial_r^{p-2}$ to $r^2 \partial_r^2 v_{ijk}$ and commuting yields

$$r^p \partial_r^p v_{ijk} = r^{p-2} \partial_r^{p-2} (r^2 \partial_r^2 v_{ijk}) - c_1 r^{p-2} \partial_r^{p-2} v_{ijk} - c_2 r^{p-1} \partial_r^{p-1} v_{ijk}. \quad (6.31)$$

The last two terms in [\(6.31\)](#) are bounded by the inductive hypothesis. To handle the first term in [\(6.31\)](#) we use [\(6.7\)](#) and obtain

$$r^2 \partial_r^2 v_{ijk} = -\Delta_\omega v_{ijk} + r^2 (Q_\ell + Q_r) v_{\leq i \leq j \leq k} - r^2 \tau^2 v_{ijk} + r^2 g_{\leq i \leq j \leq k} - 2r \partial_r v_{ijk}.$$

Now we calculate

$$\begin{aligned} & |r^{p-2} \partial_r^{p-2} (-\Delta_\omega v_{ijk} + r^2 (Q_\ell + Q_r) v_{\leq i \leq j \leq k} - r^2 \tau^2 v_{ijk} + r^2 g_{\leq i \leq j \leq k} - 2r \partial_r v_{ijk})| \\ & \lesssim \left| (r^{p-2} \partial_r^{p-4} + c_1 r^{p-1} \partial_r^{p-3} + c_2 r^p \partial_r^{p-2}) [(Q_\ell + Q_r) v_{\leq i \leq j \leq k} + g_{\leq i \leq j \leq k} - \tau^2 v_{ijk}] \right| \\ & \quad + |r^{p-2} \partial_r^{p-2} v_{(i+j+2)k}| + |(r^{p-2} \partial_r^{p-2} + r^{p-1} \partial_r^{p-1}) v_{ijk}|. \end{aligned}$$

The Q_ℓ and Q_r terms are handled in a manner analogous to the argument using [\(6.17\)](#). Each term on the right hand side is then bounded by the inductive hypothesis, the assumption $|\tau r| \lesssim 1$, and Lemma [6.6](#).

2. Large r : $\langle r \rangle \gtrsim |\tau|^{-1}$

As in Proposition [6.5](#) it suffices to prove the proposition for $(\tau \partial_\tau)^k (v e^{i\tau r})$.

If $p = 0$, the desired bound follows by [\(6.9\)](#).

Fix p and assume $|(\tau\partial_\tau)^a(v e^{i\tau r})| \lesssim (|\tau|\langle r \rangle)^{-1}$ when $a < p$. We write

$$(\tau\partial_\tau)^p(v e^{i\tau r}) = \left(\sum_{j=0}^p \sum_{\ell=0}^j c_j \ell^\ell (\partial_r + i\tau)^\ell (-S)^{p-j} v \right) e^{i\tau r}.$$

By Lemma [6.7](#), each term on the right hand side is bounded by $|\tau|^{-1}\langle r \rangle^{-1}$, as desired. $\square$

7. PROOF OF MAIN THEOREM

We are now ready to prove Theorem [1.5](#). By assumption the initial data satisfies $u_0 \in Z^{\nu+1, \kappa}$ and $u_1 \in Z^{\nu, \kappa+1}$. Since $P^1 : Z^{n, q} \rightarrow Z^{n-1, q+\kappa}$, we can write

$$R_\tau(-i\tau u_0 + P^1 u_0 - u_1) = R_\tau(\tau g_\kappa^{\nu+1} + g_{\kappa+1}^\nu)$$

for some $g_\kappa^{\nu+1} \in Z^{\nu+1, \kappa}$ and some $g_{\kappa+1}^\nu \in Z^{\nu, \kappa+1}$. Therefore [\(3.11\)](#) becomes

$$u(t, x) = \frac{1}{\sqrt{2\pi}} \int_{\mathbb{R}} R_\tau(\tau g_\kappa^{\nu+1} + g_{\kappa+1}^\nu) e^{i\tau t} d\tau. \quad (7.1)$$

We will use cutoff functions to break u into high and low frequency components as in [\(3.12\)](#) and [\(3.13\)](#).

7.1. High Frequency Case. ($|\tau| \gtrsim 1$) We will decompose the expression $R_\tau(\tau g_\kappa^{\nu+1} + g_{\kappa+1}^\nu)$ in [\(7.1\)](#) via an iterative argument, so we begin by writing the high frequency part of u as

$$u_{>1}(t, x) = \frac{1}{\sqrt{2\pi}} \int_{\mathbb{R}} \chi_{>1}(|\tau|) R_\tau(\tau f_0 + g_0) e^{i\tau t} d\tau \quad (7.2)$$

for $f_0 = g_\kappa^{\nu+1} \in Z^{\nu+1, \kappa}$ and $g_0 = g_{\kappa+1}^\nu \in Z^{\nu, \kappa+1}$.

We approximate $R_\tau(\tau f_0 + g_0)$ by $\tau^{-1} f_0$ and calculate the error. Direct calculation yields

$$P_\tau[R_\tau(\tau f_0 + g_0) - \tau^{-1} f_0] = (g_0 - iP^1 f_0) + \tau^{-1}(\Delta + P^2)(-f_0) =: f_1 + \tau^{-1} g_1.$$

Note $P^2 : Z^{p, q} \rightarrow Z^{p-2, q+\kappa}$, $P^1 : Z^{p, q} \rightarrow Z^{p-1, q+\kappa}$, and $\Delta : Z^{p, q} \rightarrow Z^{p-2, q+2}$. To see the latter, write Δ in spherical coordinates and viewing the derivatives as vector fields:

$$\Delta = \partial_r^2 + \frac{2}{r} \partial_r + r^{-2} \Delta_\omega = r^{-2}(S_r^2 - S_r) + 2r^{-2} S_r + r^{-2} \Omega^2.$$

Thus $f_1 \in Z^{\nu, \kappa+1}$ and $g_1 \in Z^{\nu-1, \kappa+2}$. Now we have

$$R_\tau(\tau f_0 + g_0) = \tau^{-1} f_0 + R_\tau(f_1 + \tau^{-1} g_1).$$

Next we reiterate the process and approximate $R_\tau(f_1 + \tau^{-1} g_1)$ by $\tau^{-2} f_1$:

$$\begin{aligned} R_\tau(f_1 + \tau^{-1} g_1) &= \tau^{-2} f_1 + R_\tau(\tau^{-1}(g_1 - iP^1 f_1) + \tau^{-2}(\Delta + P^2)(-f_1)) \\ &=: \tau^{-2} f_1 + R_\tau(\tau^{-1} f_2 + \tau^{-2} g_2) \end{aligned}$$

where $f_2 \in Z^{\nu-1, \kappa+2}$ and $g_2 \in Z^{\nu-2, \kappa+3}$. Further reiterating this process a total of J times we obtain the representation

$$R_\tau(\tau f_0 + g_0) = \underbrace{\sum_{j=0}^{J-1} \tau^{-j-1} f_j}_{=: \hat{u}_a} + \underbrace{\tau^{-J} R_\tau(\tau f_J + g_J)}_{=: \hat{u}_b} \quad (7.3)$$

where $f_j = g_{j-1} - iP^1 f_{j-1} \in Z^{\nu+1-j, \kappa+j}$ and $g_j = -(\Delta + P^2) f_{j-1} \in Z^{\nu-j, \kappa+1+j}$. We plug (7.3) into (7.2) and bound each term separately.

By Lemma 6.2 we see $|f_j| \lesssim \langle r \rangle^{-2-\kappa-j}$, and we calculate for any $N \geq 1$:

$$\begin{aligned} \left| \int_{\tau \in \mathbb{R}} \chi_{>1}(|\tau|) \hat{u}_a(\tau) e^{it\tau} d\tau \right| &\lesssim \sum_{j=0}^{J-1} \langle r \rangle^{-\kappa-2-j} \langle t \rangle^{-N} \left| \int \partial_\tau^N (\chi_{>1}(|\tau|) \tau^{-j-1}) e^{i\tau t} d\tau \right| \\ &\lesssim \langle t \rangle^{-N} \langle r \rangle^{-\kappa-2}. \end{aligned}$$

By (6.18), we have $|(\tau \partial_\tau)^\ell (\hat{u}_b(\tau) e^{i\tau \langle r \rangle})| \lesssim |\tau|^{\ell-J} \langle r \rangle^{-1}$ for $16\ell \leq \nu - J - 20$ and $\ell \leq \kappa + J$. We use this to calculate

$$\begin{aligned} \left| \int_{\tau \in \mathbb{R}} \chi_{>1}(\tau) \hat{u}_b(\tau) e^{it\tau} d\tau \right| &\approx \langle t - r \rangle^{-N} \left| \int \tau^{-N} \left[\sum_{\ell=0}^N (\tau \partial_\tau)^\ell (\chi_{>1}(|\tau|) \hat{u}_b e^{i\tau r}) \right] e^{i(t-r)\tau} d\tau \right| \\ &\lesssim \langle r \rangle^{-1} \langle t - r \rangle^{-N} \end{aligned}$$

for $J \geq 2$, $N \leq \kappa + J$, and $16N \leq \nu - J - 20$.

Combining the above results, we find

$$|u_{>1}(t, x)| \lesssim \langle t \rangle^{-N} \langle r \rangle^{-\kappa-1} + \langle r \rangle^{-1} \langle t - r \rangle^{-N}.$$

Theorem 1.5 then follows in the high frequency case if we take $J = 2$ and $N = \kappa + 2$ since the resulting requirement on ν is $\nu \geq 16\kappa + 54$, which is satisfied by our assumption $\nu \geq 31\kappa + 168$.

7.2. Low Frequency Case. ($|\tau| \lesssim 1$) Note Proposition 4.3 shows $|R_0 g_\lambda^\nu| \lesssim \langle r \rangle^{-1}$ for $1 \leq \lambda \leq \kappa + 1$. Therefore the terms of the form $R_0 g_\lambda^\nu$ in our expressions for $R_\tau g_\lambda^\nu$ in Proposition 5.6 can be included in the terms of the form $F_m(x)$. Thus we see

$$\begin{aligned} &R_\tau(\tau g_\kappa^{\nu+1} + g_{\kappa+1}^\nu) \\ &= R_\tau \left(\chi_{>|\tau|^{-1}}(r) (\tau g_\kappa^{\nu+1} + g_{\kappa+1}^\nu) \right) + \left(\sum_{m=0}^{\kappa} \tau^m F_m(x) e^{-i\tau \langle r \rangle} \right) + \tau^\kappa \varepsilon(r, \tau) e^{-i\tau \langle r \rangle} \quad (7.4) \\ &\quad + \tau^{\kappa+1} (R_\tau h_{\nu-3\kappa-3}) \end{aligned}$$

where

$$\varepsilon(r, \tau) = \langle r \rangle^{-1} \varepsilon_1(r \wedge |\tau|^{-1}) + \tau (\varepsilon_2(r \wedge |\tau|^{-1}) - \varepsilon_2(|\tau|^{-1})) \quad (7.5)$$

with $\varepsilon_1, \varepsilon_2 \in S_{rad}(\log r)$ and $r \wedge |\tau|^{-1} \approx \min(r, |\tau|^{-1})$ is smooth.

Using a decomposition as in (7.3), the first term on the right hand side of (7.4) becomes

$$R_\tau \left(\chi_{>|\tau|^{-1}}(r) (\tau g_\kappa^{\nu+1} + g_{\kappa+1}^\nu) \right) = \sum_{j=0}^{J-1} \tau^{-j-1} (\chi_{>|\tau|^{-1}} g_{\kappa+j}^{\nu+1-j}) + \sum_{M=J-1}^J \tau^{-M} R_\tau (\chi_{>|\tau|^{-1}} g_{\kappa+1+M}^{\nu-M}). \quad (7.6)$$

The term $\sum_{M=J-1}^J \tau^{-M} R_\tau (\chi_{>|\tau|^{-1}} g_{\kappa+1+M}^{\nu-M})$ can be written as $\tau^{\kappa+1} (R_\tau h_n)$ for $n = \min(\nu - J, J + \kappa)$ with h_n as in (5.17). This holds for any $J \geq 1$, and we will pick a suitable J once we have determined the necessary regularity for h_n in order for the theorem to hold. Indeed, direct calculation shows

$$|r^\ell (\partial_r + i\tau)^\ell T^i \Omega^j S^k \tau^{-N} \chi_{>1}(r\tau) g| \lesssim |r^N T^{\leq i} \Omega^{\leq j} S_r^{\leq k+\ell} g|.$$

It follows that

$$\|r^\ell (\partial_r + i\tau)^\ell T^i \Omega^j S^k \tau^{-\kappa-1-M} \chi_{>|\tau|^{-1}}(r) g_{\kappa+1+M}^{\nu-M}\|_{\mathcal{L}\mathcal{E}^*} \lesssim 1$$

for $\ell \leq M + \kappa + 1$ and $i + j + k + \ell \leq \nu - M$.

Motivated by the above calculations, we define

$$\begin{aligned} \hat{u}_a(\tau) &:= \sum_{j=0}^{J-1} \tau^{-j-1} (\chi_{>|\tau|^{-1}}(r) g_{\kappa+j}^{\nu+1-j}), & \hat{u}_b(\tau) &:= \sum_{m=0}^{\kappa} \tau^m F_m(x) e^{-i\tau \langle r \rangle}, \\ \hat{u}_c(\tau) &:= \tau^\kappa \varepsilon(r, \tau) e^{-i\tau \langle r \rangle}, & \text{and} & \quad \hat{u}_d(\tau) := \tau^{\kappa+1} (R_\tau h_{\nu-3\kappa-3} + R_\tau h_n) \end{aligned}$$

for fixed $J \geq 1$ and $n = \min(\nu - J, J + \kappa)$ and write the low frequency part of u as

$$u_{<1}(t, x) = \frac{1}{\sqrt{2\pi}} \int_{\mathbb{R}} \chi_{<1}(|\tau|) (\hat{u}_a(\tau) + \hat{u}_b(\tau) + \hat{u}_c(\tau) + \hat{u}_d(\tau)) e^{it\tau} d\tau. \quad (7.7)$$

To handle $\int \chi_{<1}(|\tau|) \hat{u}_a e^{it\tau} d\tau$ we calculate for any $N \geq 1$

$$\begin{aligned} \left| \int_{\mathbb{R}} \chi_{<1}(|\tau|) \hat{u}_a e^{it\tau} d\tau \right| &\lesssim \sum_{j=0}^{J-1} \langle r \rangle^{-\kappa-j-2} t^{-N} \left| \int_{\mathbb{R}} \partial_\tau^N \left(\chi_{<1}(|\tau|) \chi_{>1}(r|\tau) \tau^{-j-1} \right) e^{it\tau} d\tau \right| \\ &\lesssim \sum_{j=0}^{J-1} \sum_{\ell=0}^N \sum_{i=0}^{\ell} \langle r \rangle^{-\kappa-j-2} t^{-N} \left| \int_{\mathbb{R}} \chi_{<1}^{(N-\ell)}(|\tau|) \chi_{>1}^{(\ell-i)}(r|\tau) r^{\ell-i} \tau^{-j-i-1} e^{it\tau} d\tau \right| \\ &\lesssim \langle r \rangle^{-\kappa-2+N} \langle t \rangle^{-N}. \end{aligned} \quad (7.8)$$

If $r \geq \frac{t}{2}$, we take $N = 1$. If $r < \frac{t}{2}$, we take $N = \kappa + 2$. This yields

$$\left| \int_{\mathbb{R}} \chi_{<1}(|\tau|) \hat{u}_a e^{it\tau} d\tau \right| \lesssim \langle t \rangle^{-1} \langle t - r \rangle^{-\kappa-1}. \quad (7.9)$$

For the $\hat{u}_b$ term in (7.7) we find

$$\begin{aligned} \left| \int_{\mathbb{R}} \chi_{<1}(|\tau|) \hat{u}_b e^{it\tau} d\tau \right| &\lesssim \frac{1}{\langle t - r \rangle^N} \sum_{m=0}^{\kappa} \left| \int_{\mathbb{R}} \partial_\tau^N \left(\chi_{<1}(|\tau|) \tau^m F_m(x) \right) e^{i(t-\langle r \rangle)\tau} d\tau \right| \\ &\lesssim \langle r \rangle^{-1} \langle t - r \rangle^{-N} \end{aligned} \quad (7.10)$$

for any N .

Next we consider the $\hat{u}_c$ term in (7.7). Recall $\hat{u}_c = \tau^\kappa \varepsilon(r, \tau) e^{-i\tau \langle r \rangle}$ with $\varepsilon(r, \tau)$ as in (7.5). Here we define for $j = 1, 2$

$$\varepsilon_j^m(r) := \int_0^r \beta_{\approx m}(\rho) \partial_\rho \varepsilon_j(\rho) d\rho$$

so that $\varepsilon_j(r) = \sum_{m \geq 0} \varepsilon_j^m(r)$. We note $|\varepsilon_j^m| \lesssim 1$ uniformly in m and for $N \geq 1$ we have

$$|\partial_r^N \varepsilon_j^m(r)| \lesssim 2^{-mN} \mathbf{1}_{\{r \approx 2^m\}}(r). \quad (7.11)$$

Furthermore, we see

$$\varepsilon_j^m(r) \equiv \begin{cases} 0, & r \ll 2^m \\ c_m, & r \gg 2^m \end{cases} \quad (7.12)$$

where c_m are constants. Now we define

$$\varphi^m(r, \tau) := \tau^\kappa \langle r \rangle^{-1} \varepsilon_1^m(r \wedge |\tau|^{-1}) + \tau^{\kappa+1} [\varepsilon_2^m(r \wedge |\tau|^{-1}) - \varepsilon_2^m(|\tau|^{-1})]$$

so that $\hat{u}_c = \sum_{m \geq 0} \varphi^m e^{-i\tau \langle r \rangle}$. We fix r and break the summation up into the cases $2^m \ll r$, $2^m \gg r$, and $2^m \approx r$ separately.

When $2^m \ll r$ we have $\varphi^m = \tau^\kappa \langle r \rangle^{-1} \varepsilon_1^m(|\tau|^{-1})$ since for $r \ll |\tau|^{-1}$, $\varepsilon_j^m(r) = c_m = \varepsilon_j^m(|\tau|^{-1})$. Thus we find

$$\begin{aligned} \sum_{m \ll \log r} \left| \int_{\mathbb{R}} \chi_{<1}(|\tau|) \varphi^m e^{i(t-\langle r \rangle)\tau} d\tau \right| &\lesssim \sum_{m \ll \log r} \langle r \rangle^{-1} 2^{m(N-(\kappa+1))} \langle t - \langle r \rangle \rangle^{-N} \\ &\lesssim \langle t + r \rangle^{-1} \langle t - r \rangle^{-\kappa-1}. \end{aligned} \quad (7.13)$$

To obtain the first inequality we change variables by $\tau \mapsto 2^{-m}\tau$ and use the fact that $\varepsilon_1^m(2^m|\tau|^{-1}) = 0$ for $2^{-m}|\tau| \gg 2^{-m}$ along with the observations $|\varepsilon_1^m(r)| \lesssim 1$ and (7.11). The second inequality uses different arguments depending on the relative sizes of r and t . If $t < 2r$, we break up the sum into $2^{-m} < |t - \langle r \rangle|$ (where we take $N = \kappa + 2$) and $|t - \langle r \rangle| \leq 2^m < r$ (where we take $N = 0$). If $t \geq 2r$, we use the fact that $|t - \langle r \rangle| \lesssim \langle t + r \rangle^{-1}$ and set $N = \kappa + 2$.

When $2^m \gg r$, $\varepsilon_j^m(r \wedge |\tau|^{-1}) = 0$ so we are left only with the $\varepsilon_2^m(|\tau|^{-1})$ term in φ^m and we have $\varphi^m = \tau^{\kappa+1} \varepsilon_2^m(|\tau|^{-1})$. Thus we find

$$\begin{aligned} \sum_{m \gg \log r} \left| \int_{\mathbb{R}} \chi_{<1}(|\tau|) \varphi^m e^{i(t-\langle r \rangle)\tau} d\tau \right| &\lesssim \sum_{m \gg \log r} 2^{m(N-(\kappa+2))} |t - \langle r \rangle|^{-N} \\ &\lesssim \langle t + r \rangle^{-1} \langle t - r \rangle^{-\kappa-1}. \end{aligned} \quad (7.14)$$

The first inequality is obtained the same way we found the first inequality in (7.13). The second inequality again uses different arguments depending on the relative sizes of t and r . If $t \leq 2r$ we have $\langle r \rangle^{-1} \lesssim \langle t + r \rangle^{-1}$ and $\langle r \rangle^{-1} \lesssim \langle t - r \rangle^{-1}$ so the inequality follows taking $N = 0$. If $t > 2r$ we have $|t - \langle r \rangle|^{-1} \lesssim \langle t + r \rangle^{-1}$ and we break the sum into $2^m < |t - \langle r \rangle|$ (where we take $N = \kappa + 3$) and $2^m \geq |t - \langle r \rangle|$ (where we take $N = 0$).

When $2^m \approx r$, we have that φ^m is composed of terms as in the $2^m \ll r$ and $2^m \gg r$ cases. Thus we argue as above with the added benefit that the summation is finite in m to find

$$\sum_{m \approx \log r} \left| \int_{\mathbb{R}} \chi_{<1}(|\tau|) \varphi^m e^{i(t-\langle r \rangle)\tau} d\tau \right| \lesssim \langle t + r \rangle^{-1} \langle t - r \rangle^{-\kappa-1}. \quad (7.15)$$

Now (7.13), (7.13), and (7.15) yield

$$\left| \int_{\mathbb{R}} \chi_{<1}(|\tau|) \hat{u}_c(\tau) e^{it\tau} d\tau \right| \lesssim \langle t+r \rangle^{-1} \langle t-r \rangle^{-\kappa-1}. \quad (7.16)$$

Finally we handle the $\hat{u}_d$ term. We do our calculations for h_n then apply the results to $h_{\nu-3\kappa-3}$. We will use Proposition 6.8. Note $\tau^\ell \partial_\tau^\ell$ can be written as a linear combination of $(\tau \partial_\tau)^a$ with $1 \leq a \leq \ell$, so (6.29) and (6.30) hold for $(\tau \partial_\tau)^\ell$ replaced by $\tau^\ell \partial_\tau^\ell$.

We split up the $\langle r \rangle \lesssim |\tau|^{-1}$ and $\langle r \rangle \gtrsim |\tau|^{-1}$ cases using cutoff functions. When $\langle r \rangle \lesssim |\tau|^{-1}$, we first consider the case where $t \leq 2r$. Here we have

$$\left| \int_{\mathbb{R}} \chi_{<1}(\langle r \rangle |\tau|) \tau^{\kappa+1} (R_\tau h_n) e^{it\tau} d\tau \right| \lesssim \langle r \rangle^{-\kappa-2} \lesssim \langle t \rangle^{-\kappa-2}$$

by (6.29). When $2r < t$, we calculate

$$\begin{aligned} \left| \int_{\mathbb{R}} \chi_{<1}(\langle r \rangle |\tau|) \tau^{\kappa+1} (R_\tau h_n) e^{it\tau} d\tau \right| &\leq \left| \int_0^{\frac{1}{t}} |\tau|^{\kappa+1} d\tau \right| + \left| \int_{\frac{1}{t}}^\infty \chi_{<1}(r|\tau|) \tau^{\kappa+1} R_\tau h e^{it\tau} d\tau \right| \\ &\lesssim t^{-\kappa-2} \end{aligned}$$

for $16(\kappa+4) < n-20$. The inequality for the $\int_{\frac{1}{t}}^\infty \cdot d\tau$ term follows from integration by parts $\kappa+4$ times.

Now we consider when $|\tau|^{-1} \lesssim \langle r \rangle$. Take $\phi(x, \tau) := \chi_{<1}(|\tau|) \chi_{>1}(\langle r \rangle |\tau|) \tau^{\kappa+1} (R_\tau h_n) e^{it\tau}$ and use (6.30) to calculate

$$|\partial_\tau^N \phi| \lesssim |\tau|^{\kappa-N} \langle r \rangle^{-1} + \sum_{i=0}^{N-1} \chi_{>1}^{(N-i)}(\langle r \rangle |\tau|) \langle r \rangle^{N-\kappa-1} \quad (7.17)$$

for $16N \leq n-20$, since $\langle r \rangle |\tau| \approx 1$ on the support of $\chi_{>1}^{(j)}(\langle r \rangle |\tau|)$ for $j \geq 1$. If $2r \leq t$, then $|t - \langle r \rangle|^{-1} \lesssim \langle t+r \rangle^{-1}$. We use (7.17) to find for $N \geq \kappa+2$ and $16N \leq n-20$

$$\begin{aligned} &\left| \int_{\mathbb{R}} \phi e^{i\langle r \rangle \tau} e^{i(t-\langle r \rangle)\tau} d\tau \right| \\ &= t^{-1} \langle t-r \rangle^{-N} \left| \int \partial_\tau \left[(\partial_\tau^N \phi) e^{-i\langle r \rangle \tau} \right] e^{it\tau} d\tau \right| \\ &\lesssim t^{-1} \langle t-r \rangle^{-N} \left(\int_{r^{-1} < |\tau| < 1} |\partial_\tau^{N+1} \phi| d\tau + \langle r \rangle \int_{\mathbb{R}} |\partial_\tau^N \phi| d\tau \right) \\ &\lesssim t^{-1} \langle t-r \rangle^{-N} \left(\langle r \rangle^{-1} \int_{\langle r \rangle^{-1}}^\infty |\tau|^{\kappa-N-1} d\tau + \int_{|\tau| \approx \langle r \rangle^{-1}} \langle r \rangle^{N-\kappa} d\tau + \int_{\langle r \rangle^{-1}}^\infty |\tau|^{\kappa-N} d\tau \right) \\ &\lesssim \frac{1}{t \langle t-r \rangle^{\kappa+1}} \frac{\langle r \rangle^{N-\kappa-1}}{\langle t+r \rangle^{N-\kappa-1}}. \end{aligned} \quad (7.18)$$

If $t < 2r$ then $1 \lesssim \frac{\langle r \rangle}{\langle t+r \rangle}$, $\langle r \rangle^{-1} \lesssim t^{-1}$, and $\langle r \rangle^{-1} < |t - \langle r \rangle|^{-1} < \infty$. We write

$$\begin{aligned} \left| \int_{\mathbb{R}} \phi e^{i\langle r \rangle \tau} e^{i(t-\langle r \rangle)\tau} d\tau \right| &\leq \int_0^{|t-\langle r \rangle|^{-1}} \tau^\kappa \langle r \rangle^{-1} d\tau + \left| \int_{|t-\langle r \rangle|^{-1}}^\infty \phi(\tau) e^{i(t-\langle r \rangle)\tau} d\tau \right| \\ &\lesssim t^{-1} |t - \langle r \rangle|^{-\kappa-1} + \langle r \rangle^{-1} |t - \langle r \rangle|^{-\kappa-1} \\ &\lesssim \frac{1}{t\langle t-r \rangle^{\kappa+1}} \frac{\langle r \rangle^{N-\kappa-1}}{\langle t+r \rangle^{N-\kappa-1}} \end{aligned} \quad (7.19)$$

for $2r > t$ and $16(\kappa + 4) \leq n - 20$.

Combining (7.18), and (7.19) then yields

$$\left| \int_{\mathbb{R}} \chi_{<1}(|\tau|) \tau^{\kappa+1} (R_\tau h_n) e^{it\tau} d\tau \right| \lesssim \frac{1}{\langle t+r \rangle^{\kappa+2}} + \frac{1}{t\langle t-r \rangle^{\kappa+1}} \frac{\langle r \rangle^{N-\kappa-1}}{\langle t+r \rangle^{N-\kappa-1}} \quad (7.20)$$

for $16(\kappa + 4) \leq n - 20$. Thus $n = \min(\nu - J, \kappa + J) \geq 16\kappa + 84$ for fixed $J \geq 1$ so that the results hold for $\nu \geq 31\kappa + 168$.

The statement of the main theorem then follows for $|\tau| \lesssim 1$ by (7.7), (7.9), (7.10), (7.16), and (7.20).

REFERENCES

- [1] S. Alinhac. On the morawetz-keel-smith-sogge inequality for the wave equation on a curved background. *Publications of The Research Institute for Mathematical Sciences*, 42:705–720, 2006.
- [2] L. Andersson and P. Blue. Hidden symmetries and decay for the wave equation on the kerr spacetime. *Annals of Mathematics*, 182(3):787–853, 2015.
- [3] P. Blue and A. Soffer. Semilinear wave equations on the schwarzschild manifold. i. local decay estimates. *Adv. Differential Equations*, 8(5):595–614, 2003.
- [4] J.-F. Bony and D. Häfner. The semilinear wave equation on asymptotically euclidean manifolds. *Communications in Partial Differential Equations*, 35(1):23–67, 2009.
- [5] J.-F. Bony and D. Häfner. Local energy decay for several evolution equations on asymptotically euclidean manifolds. In *Annales scientifiques de l'École Normale Supérieure*, volume 45, pages 311–335, 2012.
- [6] J.-F. Bony and D. Häfner. Improved local energy decay for the wave equation on asymptotically euclidean odd dimensional manifolds in the short range case. *Journal of the Institute of Mathematics of Jussieu*, 12(3):635–650, 2013.
- [7] R. Booth, H. Christianson, J. Metcalfe, and J. Perry. Localized energy for wave equations with degenerate trapping. *arXiv preprint arXiv:1712.05853*, 2017.
- [8] N. Burq. Décroissance de l'énergie locale de l'équation des ondes pour le problème extérieur et absence de résonance au voisinage du réel. *Acta Math.*, 180(1):1–29, 1998.
- [9] M. Dafermos and I. Rodnianski. The red-shift effect and radiation decay on black hole spacetimes. *Communications on Pure and Applied Mathematics: A Journal Issued by the Courant Institute of Mathematical Sciences*, 62(7):859–919, 2009.
- [10] M. Dafermos and I. Rodnianski. A proof of the uniform boundedness of solutions to the wave equation on slowly rotating kerr backgrounds. *Inventiones Mathematicae*, 185(3):467–559, 2011.
- [11] M. Dafermos and I. Rodnianski. Lectures on black holes and linear waves. *Clay Math. Proc.*, 17:97–205, 2013.
- [12] M. Dafermos, I. Rodnianski, and Y. Shlapentokh-Rothman. Decay for solutions of the wave equation on kerr exterior spacetimes iii: The full subextremal case — a — i m. *Annals of Mathematics*, 183(3):787–913, 2016.
- [13] R. Donniger, W. Schlag, and A. Soffer. A proof of price's law on schwarzschild black hole manifolds for all angular momenta. *Advances in Mathematics*, 226(1):484–540, 2011.

- [14] F. Finster, N. Kamran, J. Smoller, and S.-T. Yau. Decay of solutions of the wave equation in the kerr geometry. *Communications in Mathematical Physics*, 264(2):465–503, 2006.
- [15] K. Hidano, J. Metcalfe, H. Smith, C. Sogge, and Y. Zhou. On abstract strichartz estimates and the strauss conjecture for nontrapping obstacles. *Transactions of the American Mathematical Society*, 362(5):2789–2809, 2010.
- [16] M. Keel, H. F. Smith, and C. D. Sogge. Almost global existence for some semilinear wave equations. *Journal d'Analyse Mathématique*, 87(1):265–279, 2002.
- [17] H. Lindblad and M. Tohaneanu. Global existence for quasilinear wave equations close to schwarzschild. *Communications in Partial Differential Equations*, 43(6):893–944, 2018.
- [18] J. Luk. The null condition and global existence for nonlinear wave equations on slowly rotating kerr spacetimes. *Journal of the European Mathematical Society*, 15(5):1629–1700, 2013.
- [19] J. Marzuola, J. Metcalfe, D. Tataru, and M. Tohaneanu. Strichartz estimates on schwarzschild black hole backgrounds. *Communications in Mathematical Physics*, 293(1):37, 2010.
- [20] J. Metcalfe, M. Nakamura, and C. D. Sogge. Global existence of solutions to multiple speed systems of quasilinear wave equations in exterior domains. In *Forum Mathematicum*, volume 17, pages 133–168. Walter de Gruyter, 2005.
- [21] J. Metcalfe and C. D. Sogge. Long-time existence of quasilinear wave equations exterior to star-shaped obstacles via energy methods. *SIAM journal on mathematical analysis*, 38(1):188–209, 2006.
- [22] J. Metcalfe, J. Sterbenz, and D. Tataru. Local energy decay for scalar fields on time dependent non-trapping backgrounds. *arXiv preprint arXiv:1703.08064*, 2017.
- [23] J. Metcalfe and D. Tataru. Decay estimates for variable coefficient wave equations in exterior domains. In *Advances in phase space analysis of partial differential equations*, pages 201–216. Springer, 2009.
- [24] J. Metcalfe and D. Tataru. Global parametrices and dispersive estimates for variable coefficient wave equations. *Mathematische Annalen*, 353(4):1183–1237, 2012.
- [25] J. Metcalfe, D. Tataru, and M. Tohaneanu. Price's law on nonstationary space-times. *Advances in Mathematics*, 230(3):995–1028, 2012.
- [26] C. S. Morawetz. The decay of solutions of the exterior initial-boundary value problem for the wave equation. *Communications on Pure and Applied Mathematics*, 14(3):561–568, 1961.
- [27] R. H. Price. Nonspherical perturbations of relativistic gravitational collapse. i. scalar and gravitational perturbations. *Physical Review D*, 5(10):2419, 1972.
- [28] J. V. Ralston. Solutions of the wave equation with localized energy. *Communications on Pure and Applied Mathematics*, 22(6):807–823, 1969.
- [29] J. Sbierski. Characterisation of the energy of gaussian beams on lorentzian manifolds: with applications to black hole spacetimes. *Analysis & PDE*, 8(6):1379–1420, 2015.
- [30] H. F. Smith and C. D. Sogge. Global strichartz estimates for nontrapping perturbations of the laplacian. *Comm. Partial Differential Equations*, 25(11-12):2171–2183, 2000.
- [31] C. D. Sogge and C. Wang. Concerning the wave equation on asymptotically euclidean manifolds. *Journal d'Analyse Mathématique*, 112(1):1–32, 2010.
- [32] J. Sterbenz. Angular regularity and strichartz estimates for the wave equation. *International Mathematics Research Notices*, 2005(4):187–231, 2005.
- [33] D. Tataru. Local decay of waves on asymptotically flat stationary space-times. *American Journal of Mathematics*, 135(2):361–401, 2013.
- [34] M. Tohaneanu. Strichartz estimates on kerr black hole backgrounds. *Transactions of the American Mathematical Society*, 364(2):689–702, 2012.
- [35] S. Yang. Global stability of solutions to nonlinear wave equations. *Selecta Mathematica*, 21(3):833–881, 2015.

NORTHWESTERN UNIVERSITY, DEPT. OF MATHEMATICS, 2033 SHERIDAN RD., EVANSTON, IL 60208

Email address: `katrina.morgan@northwestern.edu`