

GENERALIZED PRICE’S LAW ON FRACTIONAL-ORDER ASYMPTOTICALLY FLAT STATIONARY SPACETIMES

KATRINA MORGAN AND JARED WUNSCH

ABSTRACT. We obtain estimates on the rate of decay of a solution to the wave equation on a stationary spacetime that tends to Minkowski space at a rate $O(|x|^{-\kappa})$, $\kappa \in (1, \infty) \setminus \mathbb{N}$. Given suitably smooth and decaying initial data, we show a wave locally enjoys the decay rate $O(t^{-\kappa-2+\epsilon})$.

1. Introduction

The goal of this work is to study the relationship among pointwise decay rates of waves, low- and high-frequency resolvent estimates, and the large scale behavior of the background geometry. We study solutions to the wave equation on asymptotically flat stationary 4-dimensional Lorentzian spacetimes with signature (3,1). The flat Minkowski metric, which we denote \mathfrak{m} , is given by $\mathfrak{m} = dt^2 - dx_1^2 - dx_2^2 - dx_3^2$. The spacetimes considered here are of the form $\mathfrak{g} = \mathfrak{m} + \mathfrak{h}$ where \mathfrak{h} has metric coefficients which decay like $r^{-\kappa}$ for some $\kappa \in (1, \infty) \setminus \mathbb{N}$. We say such a metric tends toward flat at a rate of $r^{-\kappa}$. We find that, given sufficiently differentiable and decaying Cauchy data, waves decay locally at a rate of $t^{-\kappa-2+\epsilon}$. The main new input to this decay estimate is a certain resolvent estimate valid uniformly near zero frequency. In previous work [34], the first author studied the case when $\kappa \in \mathbb{N}$ and established $t^{-\kappa-2}$ decay rates when the background geometry exhibits spherical symmetry, and $t^{-\kappa-2+\epsilon}$ decay in the absence of spherical symmetry.

The study of pointwise decay rates on asymptotically flat spacetimes arises in general relativity. In [35], physicist Richard Price gave a heuristic argument anticipating a t^{-3} pointwise decay rate for waves on the Schwarzschild spacetime, which describes space in the presence of a single, non-rotating black hole. This conjecture is known as Price’s Law. There has been much mathematical interest in studying pointwise decay rates of waves on relativistic geometric backgrounds, including the Schwarzschild and Kerr spacetimes (the latter describes the geometry resulting from a rotating black hole). Both these geometries tend toward flat at a rate of r^{-1} .

Price’s Law was proved in [39] (see also [16] for the mode-by-mode estimate as well as [12] and [22] for earlier decay estimates). Pointwise decay rates for the Kerr spacetime were studied in [14] and [17]. The techniques in [16], [39], [34], and the current work involve taking the Fourier transform in time and therefore do not readily extend to non-stationary geometries. In [32] the authors proved Price’s Law for non-stationary asymptotically flat spacetimes and established the t^{-3} decay rate for a class of perturbations of the Kerr spacetime. Results similar to the current work but also considering non-stationary spacetimes were obtained concurrently by Looi in [21]. The author uses vector field techniques and an iterative argument to obtain

decay rates which depend on the rate at which the background geometries become flat. While the results of [21] are in fact stronger than those obtained here in the sense that they apply to nonstationary metrics (and do not lose an ϵ power), the results of this paper give direct and precise estimates in the frequency domain—in particular, near zero frequency—which we hope will find further applications. For instance, given a situation in which high-frequency resolvent estimates are known directly (and there are many such situations, often with estimates obtained directly in frequency domain by semiclassical propagation of singularities; see e.g. [41]), we can easily combine the high- and low-frequency estimates to obtain new decay estimates in the time domain.

In the present work as well as in [34] and [39] an integrated local energy decay estimate is assumed to hold. Such estimates were established on the Schwarzschild geometry in [6], [11], and [23]. For the Kerr spacetime with low angular momentum, local energy estimates were proved in [2], [11], and [15]; the major challenge here was the trapping of null geodesics within a compact set in the space variables. The assumptions in [39] therefore hold for Schwarzschild and Kerr with low angular momentum. We discuss the integrated local energy decay estimates in more detail later in the introduction.

Finer analysis of the asymptotics of solutions on Schwarzschild space was obtained in [3], including a characterization of when the t^{-3} decay rate is a *lower* bound for the decay.

A different approach to Price’s law was pioneered by Hintz [19], who showed that the estimate in Price’s Law on Kerr backgrounds is sharp, and obtained explicit leading order asymptotics. Here, rather than use vector field methods in physical space, the author employs resolvent estimates, after Fourier transforming the equation in the appropriate variables. No explicit local energy decay assumption is employed: the author instead considers metrics for which the associated spectral family (given by formal Fourier transform in time) satisfies appropriate hypotheses on its inverse (the resolvent). Integrated local energy decay estimates are in fact intimately related to resolvent estimates. For example, in [29] the authors establish a full spectral characterization of local energy decay in the context of nontrapping asymptotically flat spacetimes.

An analysis of the asymptotics of solutions that distinguishes the contributions due to low angular modes from the more decaying contributions from higher ones has recently been carried out on Reissner–Nordström backgrounds in [5] and on Kerr backgrounds in [4].

Our approach blends the hypotheses of [39] and associated works with an adaptation of the new approach of Hintz [19], which in turn harnesses powerful low frequency resolvent estimates recently developed by Vasy [40]. The high frequency estimates thus follow [39] and [34], with the energy decay assumption being the crucial hypothesis ensuring that in return for enough derivatives of regularity, this part of the solution in fact decays at any desired rate. At low frequency, by contrast, the local energy decay assumption yields absence of resonances in the upper half-space, which is certainly one necessary condition for decay estimates to hold (and this is an explicit spectral hypothesis, e.g. in [19]). We obtain additional information about the low-frequency asymptotics of asymptotically Euclidean resolvents, which is essentially

independent of the small-scale geometry, via estimates from [19] and [40]: we apply these iteratively to obtain conormal estimates for the resolvent at zero frequency.

A crucial step in our analysis is thus writing an expansion of the resolvent at zero frequency (see Lemma 15). Low frequency resolvent expansions of the Laplacian were first studied in [20]. The geometric context we consider here reduces to analyzing perturbations of the flat Laplacian where the perturbation depends on the spectral parameter. The presence of such terms arises from the fully Lorentzian nature of the perturbations considered here. Closely related results on low frequency spectral behavior and local wave decay were previously obtained in the setting of scattering manifolds in [18], where the authors use an extremely precise description of the spectral measure at low frequencies to establish decay rates dependent on the dimension of the spacetime.

Questions similar to those treated here were studied in [8] and [7] where the authors established local decay rates for waves on asymptotically flat, stationary spacetimes which tend toward flat at different rates. There are several key differences compared with the current work. First, as noted above, we handle full Lorentzian perturbations of flat Minkowski space rather than restricting to perturbations of the Laplacian. This leads to the metrics considered in this paper containing $dt dx_i$ terms, which results in mixed space-time differential operators in our wave operator. Second, we allow for the possibility of unstable trapping on our background. In [8] and [7], a nontrapping assumption is used in order to obtain decay for the high frequency part of a solution to the wave equation (it is not needed for the low frequency part). Third, our result improves upon the established decay rates. Finally we note that [7] considers $(1+n)$ dimensional geometries for $n \geq 2$ and [8] considers n odd with $n \geq 3$. The current work only studies $(1+3)$ dimensional spacetimes.

1.1. Wave Equation. On flat Minkowski space, the wave operator, denoted \square , is given by $\square = \partial_t^2 - \Delta = \partial_t^2 - \sum_{i=1}^3 \partial_{x_i}^2$. More generally, on a Lorentzian spacetime with metric $g = g_{\alpha\beta}$, the wave operator is denoted \square_g and is given by the d'Alembertian

$$\square_g = \frac{1}{\sqrt{|g|}} \partial_\alpha \sqrt{|g|} g^{\alpha\beta} \partial_\beta$$

where $g^{\alpha\beta}$ are the dual metric coordinates.

1.2. Local Energy Decay. On stationary spacetimes, as considered here, solutions to the wave equation may have constant energy. If the background geometry allows solutions to spread out, then energy may decay within compact sets. We assume an integrated local energy estimate where the local decay is fast enough to be integrable in time. Estimates on the local decay of energy have a long history dating back to the work of Morawetz [33]. The specific version of the estimates used here come from the work of Metcalfe and Tataru in [31]. In addition to local energy decay, we will assume uniform energy bounds on solutions of the wave equation. Such bounds necessarily hold on stationary spacetimes where the Killing vector field ∂_t is everywhere time-like. The uniform energy estimate also holds on Schwarzschild and Kerr, although in these geometries ∂_t is not everywhere time-like (see e.g. [11], [23], [13] among others).

We consider the Cauchy problem

$$(1) \quad (\square_g + V)u = f, \quad u(0, x) = u_0, \quad \partial_t u(0, x) = u_1$$

where V is a suitably decaying potential (see the statement of the main theorem for conditions on V). The Cauchy data at time t is denoted $u[t] = (u(t, \cdot), \partial_t u(t, \cdot))$.

Definition 1. We say the evolution (1) satisfies the uniform energy bounds if:

$$(2) \quad \|u[t]\|_{\dot{H}^{k,1} \times H^k} \leq c_k(\|u[0]\|_{\dot{H}^{k,1} \times H^k} + \|f\|_{L^1 H^k}), \quad t \geq 0, \quad k \geq 0.$$

Here H^k denotes the usual Sobolev space, and we say $\phi \in \dot{H}^{k,1}$ if $\nabla \phi \in H^k$, and write

$$\|\phi\|_{\dot{H}^{k,1}} = \|\nabla \phi\|_{H^k}.$$

In the following definitions we use ∂ to denote the space-time gradient while ∇ is reserved for the gradient in spatial variables only. We write $\langle r \rangle := \sqrt{1 + r^2}$ and define the dyadic region $A_m := \{x : 2^m \leq \langle r \rangle \leq 2^{m+1}\}$. The local energy norm we use is defined by

$$\|u\|_{LE} = \sup_m \|\langle r \rangle^{-\frac{1}{2}} u\|_{L^2(\mathbb{R}_+ \times A_m)}.$$

Its H^1 analogue is given by

$$\|u\|_{LE^1} = \|\partial u\|_{LE} + \|\langle r \rangle^{-1} u\|_{LE},$$

and the dual norm is given by

$$\|f\|_{LE^*} = \sum_m \|\langle r \rangle^{\frac{1}{2}} f\|_{L^2(\mathbb{R}_+ \times A_m)}.$$

For functions with higher regularity we define the following norms

$$\|u\|_{LE^{1,N}} = \sum_{j \leq N} \|\partial^j u\|_{LE^1}, \quad \|f\|_{LE^{*,N}} = \sum_{j \leq N} \|\partial^j f\|_{LE^*}.$$

The spatial counterparts of the LE and LE^* space-time norms are

$$\|v\|_{\mathcal{LE}} = \sup_m \|\langle r \rangle^{-\frac{1}{2}} v\|_{L^2(A_m)}; \quad \|g\|_{\mathcal{LE}^*} = \sum_m \|\langle r \rangle^{\frac{1}{2}} g\|_{L^2(A_m)}$$

with the higher regularity norms defined by

$$\|v\|_{\mathcal{LE}^N} = \sum_{j \leq N} \|\nabla^j v\|_{\mathcal{LE}}, \quad \|g\|_{\mathcal{LE}^{*,N}} = \sum_{j \leq N} \|\nabla^j g\|_{\mathcal{LE}^*}.$$

Definition 2. We say the evolution (1) satisfies the local energy decay estimate if:

$$(3) \quad \|u\|_{LE^{1,N}} \leq c_N(\|u[0]\|_{H^{N,1} \times H^N} + \|f\|_{LE^{*,N}}), \quad N \geq 0.$$

Local energy decay is known to hold in several nontrapping geometries. For sufficiently small perturbations of flat space without trapping, local energy decay was established in [1], [28], and [30]. The case of stationary product manifolds was considered in [10], [7], and [38]. The nontrapping case was studied more generally in [29]. If trapping occurs then the local energy decay estimate does not hold ([36], [37]). However, if the trapping is sufficiently unstable (i.e. perturbing a trapped geodesic typically results in geodesics which escape to infinity) then a weaker form of local energy decay may hold (see, e.g., [42] in the case of the normally hyperbolic trapping that occurs in Kerr black hole spacetimes). In the case of trapping, there is necessarily a loss of derivatives on the right hand side of the estimate (see e.g. [9]).

Definition 3. We say the evolution (1) satisfies the weak local energy decay estimate if:

$$(4) \quad \|u\|_{LE^{1,N}} \leq c_N (\|u[0]\|_{\dot{H}^{N+\varpi,1} \times H^{N+\varpi}} + \|f\|_{LE^{*,N+\varpi}}), \quad N \geq 0$$

for some $\varpi > 0$.

This weak local energy decay estimate is generally obtained by using a cutoff function to remove the trapped set. The precise derivative loss depends on the trapping. Here we allow for an arbitrary but fixed loss, ϖ .

1.3. b-Sobolev spaces. Our hypotheses on the Cauchy data and our key estimates in the iteration at low frequency will be stated in the language of *b-Sobolev spaces*. These spaces, described in detail in [24] (and discussed further in Section 2 below), can be easily described in the context at hand, at least for positive integer orders, as the spaces of functions enjoying Kohn–Nirenberg symbol estimates to finite order with respect to L^2 (rather than the usual L^∞):

$$u \in H_b^{m,l}(\overline{\mathbb{R}^3}) \iff \|\langle x \rangle^{l+|\alpha|} \partial_x^\alpha u\|_{L^2(\mathbb{R}^3)} \leq C_\alpha, \quad |\alpha| \leq m.$$

The fact that the space is nominally defined on $\overline{\mathbb{R}^3}$, the radial compactification of \mathbb{R}^3 , is a nod to the fact that these spaces are more generally defined on manifolds with boundary; details follow in Section 2 and Appendix A below.

We will also use in our hypotheses the usual spaces of Kohn–Nirenberg symbols, defined by the estimates

$$(5) \quad a \in S^l(\mathbb{R}^n) \iff a \in C^\infty(\mathbb{R}^n), \sup \langle x \rangle^{-l+|\alpha|} |\partial_x^\alpha a| < C_\alpha \quad \text{for all } \alpha.$$

1.4. Statement of Main Theorem. Let \mathbf{g}, \mathbf{h} satisfy the hypotheses of Section 1.5 of [34]:

- (i) \mathbf{g} is stationary (i.e. the metric coefficients are time independent).
- (ii) The submanifolds $t = \text{constant}$ are space-like (i.e. the induced metric on the spatial submanifolds is positive definite).
- (iii) Let $\kappa \in (1, \infty) \setminus \mathbb{N}$. The metric \mathbf{g} is asymptotically flat in the sense that \mathbf{g} can be written as

$$\mathbf{g} = \mathbf{m} + \mathbf{h}$$

where

$$\mathbf{h} = \mathbf{h}_{00}(x)dt^2 + \mathbf{h}_{0i}(x)dt dx_i + \mathbf{h}_{ij}(x)dx_i dx_j$$

with $\mathbf{h}_{\alpha\beta} \in S^{-\kappa}(\mathbb{R}^3)$ for $\alpha, \beta \in \{0, 1, 2, 3\}$.

Theorem 4. *Let u solve the homogeneous Cauchy problem*

$$(\square_{\mathbf{g}} + V)u = 0, \quad u(0, x) = u_0, \quad \partial_t u(0, x) = u_1$$

for $V \in S^{-\kappa-2}(\mathbb{R}^3)$ and

$$u_0 \in H_b^{s+1, \kappa+7/2}, \quad u_1 \in H_b^{s, \kappa+7/2},$$

with

$$s > (\varpi + 1)(2\kappa + 9) + 2.$$

Assume the evolution (1) satisfies the uniform energy bounds (2) and the weak local energy decay estimate (4). Then $|u(t, x)| \leq C_\epsilon t^{-\kappa-2+\epsilon}$ for any $\epsilon > 0$, uniformly on compact sets in x .

1.5. Paper Overview. The operator $\square_g + V$ can be replaced by an operator of the form

$$P = \partial_t^2 - \Delta + \partial_t P^1 + P^2$$

where

$$P^1 \in S^{-\kappa} \partial_x + S^{-\kappa-1} \quad \text{and} \quad P^2 \in S^{-\kappa} \partial_x^2 + S^{-\kappa-1} \partial_x + S^{-\kappa-2}.$$

This is obtained by working in normalized coordinates and multiplying by $(g^{00})^{-1}$ so the coefficient on ∂_t^2 is 1. We refer the reader to [34] for details of the calculation. We work with this operator throughout. The resolvent, P_σ^{-1} , associated to P is given formally by inverse Fourier transforming the operator P and taking the inverse.

In section 3 we define the resolvent and extract relevant spectral information from the energy assumptions. In particular, we show the resolvent is well defined in the upper half plane and extends continuously to the real axis. The results in section 3 allow us to relate the inverse Fourier transform of a solution u to the initial data of the homogeneous Cauchy problem via the resolvent:

$$\check{u}(\sigma, x) = (2\pi)^{-\frac{1}{2}} P_\sigma^{-1} (-i\sigma u_0 + P^1 u_0 - u_1)$$

and to recover u by taking the Fourier transform (with integration along $\sigma \in \mathbb{R}$). The amount of decay we are able to obtain then depends on conormal regularity estimates (i.e. bounds on $(\sigma \partial_\sigma)^j P_\sigma^{-1} g$ for appropriately chosen g). We handle the high and low frequency cases separately. The high frequency part of u is sensitive to the trapping dynamics, which are controlled by our weak local energy decay assumption. Indeed, the spectral information derived from the energy assumptions is sufficient to handle the high frequency part of the solution to the wave equation, and we find that in exchange for enough derivatives on the Cauchy data, we could obtain any desired polynomial time-decay for this piece of the solution. The low frequency part of u , by contrast, is sensitive to the far away behavior of the background geometry, and it is this latter piece which ultimately dictates the final decay rate.

We establish conormal estimates for the high energy resolvent in section 4. We then turn to the low energy analysis in section 5. This includes deriving an expansion of the resolvent at zero energy, which is then used to find a helpful expression for the resolvent at low frequencies. The low frequency analysis utilizes conormal and b-Sobolev spaces, and we provide an overview of these function spaces in section 2. Finally, we prove the main theorem in section 6.

2. Background on Function Spaces

In dealing with low-frequency estimates, it will be convenient to compactify our asymptotically Euclidean space and to employ the language of conormal and b-Sobolev spaces on manifolds with boundary, as introduced by Melrose–Mendoza [25] based on Melrose’s b-calculus of pseudodifferential operators [26]. (See [24] for an extended exposition.) Some details of the local characterizations of these spaces near the boundary at infinity via Mellin transform have been relegated to Appendix A.

Let $\overline{\mathbb{R}^n} \simeq \overline{B^n}$ denote the radial compactification of Euclidean space to the unit ball, with $\text{RC} : \mathbb{R}^n \rightarrow (\overline{\mathbb{R}^n})^\circ$ given by the compactification map

$$\text{RC}(x) = \frac{x}{1 + \langle x \rangle}.$$

Note that $\text{RC}^*(\mathcal{C}^\infty(\overline{\mathbb{R}^n})) = S_{\text{cl}}^0(\mathbb{R}^n)$, the space of “classical” symbols on \mathbb{R}^n (those satisfying (5) with $l = 0$ and additionally admitting an asymptotic expansion in negative integer powers of $|x|$). The function

$$\rho \equiv |x \circ \text{RC}^{-1}|^{-1}$$

extends (except for a singularity at the origin) to a smooth function on $\overline{\mathbb{R}^3}$ that is a boundary defining function (it vanishes to exactly first order at the boundary). We will freely employ the abuse of notation

$$\rho = r^{-1} = |x|^{-1},$$

ignoring the pushforward/pullback by RC , and moreover will consider ρ to be extended to a globally smooth function on $\overline{\mathbb{R}^3}$, eliminating the singularity at $x = 0$.

More generally, we temporarily let X denote any manifold with boundary. Let $\mathcal{V}_b(X)$ denote the space of “b-vector fields,” i.e. those which are tangent to ∂X . If (ρ, y) are coordinates in a collar neighborhood of ∂X , with ρ a boundary defining function and y coordinates on ∂X , extended to the interior, then $\mathcal{V}_b(X)$ is locally spanned by $\rho \partial_\rho, \partial_y$ over $\mathcal{C}^\infty(X)$; in the special case $X = \overline{\mathbb{R}^n}$ these vector fields correspond to $-r \partial_r, \partial_\theta$; note in particular that their norm is $O(r)$ as $r \rightarrow \infty$.

The b-differential operators, $\text{Diff}_b^m(X)$, are defined as the \mathcal{C}^∞ -span of up to m -fold products of vector fields in $\mathcal{V}_b(X)$. Given a fixed volume form on X° (possibly singular at ∂X) we define $L^2(X)$ with respect to the volume form, and then set, for $k \in \mathbb{N}$,

$$H_b^k(X) = \{u \in L^2(X) : V_1 \dots V_j(u) \in L^2(X) \ \forall j \leq k, V_i \in \mathcal{V}_b(X)\}.$$

We can further define $H_b^s(X)$ for $s \in \mathbb{R}$ by interpolation and duality, or by use of the calculus of b-pseudodifferential operators microlocalizing the algebra $\text{Diff}_b(X)$ as in [24].

In this paper, we will always employ the standard metric volume form on \mathbb{R}^3 in defining L^2 and Sobolev spaces on $\overline{\mathbb{R}^3}$. This volume form is given, near the boundary $\rho = 0$, by

$$dV = \frac{d\rho d\theta}{\rho^4}$$

(with $d\theta$ shorthand for the volume form on S^2) hence appears quite singular on the compactification.¹ Note that Sobolev spaces with no subscripts will continue to denote “ordinary” Sobolev spaces on \mathbb{R}^3 . These can, if desired, be identified with the “scattering Sobolev spaces” on $\overline{\mathbb{R}^3}$ as introduced by Melrose [27], but we will not employ this terminology below.

We generalize the b-Sobolev spaces to *weighted* b-Sobolev spaces by simply setting

$$H_b^{m,\ell}(X) = \rho^\ell H_b^m(X).$$

¹We caution the reader that this choice of convention is not universal in the subject, with the “b-volume form” $d\rho d\theta/\rho$ also having a considerable popularity, since its use would eliminate the factors of $\rho^{3/2}$ that bedevil the accounting used here.

A related function space to the b-Sobolev space is that of *conormal functions* enjoying infinite-order iterated regularity under $\mathcal{V}_b(X)$, measured with respect to L^∞ rather than L^2 :

$$u \in \mathcal{A}^\gamma \iff V_1 \dots V_N(u) \in \rho^\gamma L^\infty(X) \text{ for all } N \in \mathbb{N}, V_j \in \mathcal{V}_b(X).$$

Such estimates are closely related to the standard Kohn–Nirenberg symbol estimates, since $\mathcal{V}_b(\overline{\mathbb{R}^3})$ is spanned over \mathcal{C}^∞ by $r\partial_r = -\rho\partial_\rho$, and ∂_θ ; thus radial compactification gives an isomorphism

$$\mathcal{A}^\gamma(\overline{\mathbb{R}^3}) \equiv S^{-\gamma}(\mathbb{R}^3),$$

with the symbol spaces S^\bullet defined by (5) above. The point is that the definition of the symbol spaces can be rephrased as iterated regularity under the vector fields $\langle x \rangle \partial_{x_j}$, which span \mathcal{V}_b .

The conormal spaces are also closely related to $H_b^{\infty, \ell}(X)$, but with a vexing shift in orders owing to the metric volume form:

$$(6) \quad H_b^{\infty, \ell}(X) \subset \mathcal{A}^{\ell+3/2}(X) \subset H_b^{\infty, \ell-}(X), \quad \ell \in \mathbb{R}.$$

3. Spectral Information from Energy Assumptions

We define a spectral family associated to P by

$$(7) \quad \begin{aligned} P_\sigma &\equiv e^{it\sigma} P e^{-it\sigma} \\ &= -\sigma^2 - \Delta - i\sigma P^1 + P^2 \end{aligned}$$

which is equivalent to formally inverse Fourier transforming the operator in t . In this section we use the energy assumptions to study the existence and boundedness of the operator P_σ^{-1} .

We will use the unitary normalization for the Fourier transform on \mathbb{R}^n :

$$(\mathcal{F}f)(\xi) = (2\pi)^{-n/2} \int f(x) e^{-i\xi x} dx,$$

with the inverse Fourier transform then given by its formal adjoint (and denoted $\mathcal{F}^{-1}f$ or \check{f}).

Proposition 5. *If $\text{Im } \sigma > 0$, the operator $P_\sigma : H^2 \rightarrow L^2$ is invertible. Furthermore, if u satisfies (1) then for $\text{Im } \sigma > 0$ we have*

$$(8) \quad (2\pi)^{1/2} \check{u}(\sigma, x) = P_\sigma^{-1}((2\pi)^{1/2} \check{f}(\sigma) - i\sigma u_0 + P^1 u_0 - u_1).$$

Proof. Let u solve $Pu = f$ for $t \geq 0$. By the uniform energy bound (2), $e^{it\sigma}u$ enjoys exponential energy decay at $t \rightarrow +\infty$ for $\text{Im } \sigma > 0$, hence, with $H(t)$ denoting the Heaviside function, $\mathcal{F}_{t \rightarrow \sigma}^{-1}(H(t)u)$ is an analytic function of σ in the upper half-space, taking values in the energy space.

Integrating by parts in t moreover gives, for $\text{Im } \sigma > 0$,

$$\begin{aligned} (2\pi)^{1/2} \mathcal{F}^{-1}(H(t)Pu)(\sigma, x) &= \int_0^\infty e^{+it\sigma} Pu dt \\ &= (2\pi)^{1/2} P_\sigma \mathcal{F}^{-1}(H(t)u)(\sigma) - \partial_t u(0) + i\sigma u(0) - P_1 u(0) \end{aligned}$$

Consequently,

$$\begin{aligned} (2\pi)^{1/2} P_\sigma \mathcal{F}^{-1}(H(t)u)(\sigma) \\ = -(2\pi)^{1/2} \mathcal{F}^{-1}(H(t)Pu)(\sigma, x) + \partial_t u(0) - i\sigma u(0) + P_1 u(0). \end{aligned}$$

Thus we can solve $P_\sigma v = g$, for $g \in L^2$, by solving the IVP (1) with $u_0 = 0$, $u_1 = g$ and setting

$$(9) \quad v(t, x) = \int_0^\infty u(t, x) e^{it\sigma} dt.$$

(A priori this construction would produce $v \in H^1$, but $v \in H^2$ then follows by ellipticity of P_σ .)

Now we turn to injectivity of P_σ . If $P_\sigma w = 0$ with $\text{Im } \sigma > 0$, we set

$$u(t, x) = e^{-it\sigma} w(x).$$

Then by (7) we obtain

$$Pu = e^{-it\sigma} P_\sigma w = 0.$$

Since $\text{Im } \sigma > 0$, $\|u\|$ is exponentially growing as $t \rightarrow +\infty$, contradicting the uniform energy bounds. Hence we obtain injectivity of P_σ . \square

The \mathcal{LE}_σ norm, in which we measure the resolvent $v = P_\sigma^{-1}g$, is defined by

$$(10) \quad \|v\|_{\mathcal{LE}_\sigma^N} = \|(|\sigma| + \langle r \rangle)^{-1} v\|_{\mathcal{LE}^N} + \|\nabla v\|_{\mathcal{LE}^N} + \|(|\sigma| + \langle r \rangle)^{-1} \nabla^2 v\|_{\mathcal{LE}^N}.$$

Next we transfer bounds in the local energy decay estimate to bounds on the resolvent measured in the \mathcal{LE}_σ norm which hold down to the real axis. Ultimately we wish to obtain decay rates for u by taking the Fourier transform of (8), but integrating in $\text{Im } \sigma > 0$ would lead to exponential blow-up in time. The extension of P_σ^{-1} to $\sigma \in \mathbb{R}$ allows us to integrate along $\sigma \in \mathbb{R}$.

The following proposition is analogous to results in [39] (see Proposition 9 and Corollary 12) where $\varpi = 3$ (see equation (4.2)). We provide an outline of the proof, focusing on how our assumed ϖ derivative loss in the weak local energy decay estimate (4) affects the derivative loss in the estimate (11).

Proposition 6. *If $\text{Im } \sigma \geq 0$ and $g \in \mathcal{LE}^{*,N+\varpi+1}$ for fixed $N \in \mathbb{N}$, then $v = P_\sigma^{-1}g$ satisfies*

$$(11) \quad \|v\|_{\mathcal{LE}_\sigma^N} \lesssim \|g\|_{\mathcal{LE}^{*,N+\varpi+1}}.$$

Proof. Recall that we may construct v for $\text{Im } \sigma > 0$ by solving the IVP with Cauchy data $(0, g)$ and evaluating the integral (9); this construction continues to make sense distributionally down to $\text{Im } \sigma = 0$, indeed with explicit weighted Sobolev estimates, as we will now see.

The weak local energy decay estimate allows us to establish the inequality

$$(12) \quad \|(\langle r \rangle^{-1} + |\sigma|)v\|_{\mathcal{LE}^N} + \|\nabla v\|_{\mathcal{LE}^N} \lesssim \sum_{j \leq N+\varpi} (1 + |\sigma|)^{N+\varpi-j} \|\nabla^j g\|_{\mathcal{LE}^*},$$

Note (12) follows formally from (4) by Plancherel after inverting the Fourier transform. From (12) we are able to bound the first two terms in the \mathcal{LE}_σ^N norm:

$$(13) \quad \|(\langle r \rangle^{-1} + |\sigma|)v\|_{\mathcal{LE}^N} + \|\nabla v\|_{\mathcal{LE}^N} \lesssim \|g\|_{\mathcal{LE}^{*,N+\varpi}}.$$

We refer the reader to [39] for details of the process (cf. equation (4.6)).

Now consider the second order term in the \mathcal{LE}_σ^N norm. If $(\langle r \rangle^{-1} + |\sigma|)^{-1} \lesssim 1$ then

$$\|(\langle r \rangle^{-1} + |\sigma|)^{-1} \nabla^2 v\|_{\mathcal{LE}^N} \lesssim \|\nabla^2 v\|_{\mathcal{LE}^N} \lesssim \|\nabla v\|_{\mathcal{LE}^{N+1}} \lesssim \|g\|_{\mathcal{LE}^{*,N+\varpi+1}}.$$

This step is where the extra derivative loss in (11) versus (4) comes from. It is left to consider the case where $\langle r \rangle$ is large and $|\sigma| \lesssim 1$. Here we write

$$\|(\langle r \rangle^{-1} + |\sigma|)^{-1} (-\Delta + P^2)v\|_{\mathcal{LE}^N} \lesssim \| |\sigma| v \|_{\mathcal{LE}^N} + \|P^1 v\|_{\mathcal{LE}^N} + \|\langle r \rangle g\|_{\mathcal{LE}^N}.$$

The first two terms have already been shown to satisfy the desired bounds. For the third term, straightforward calculation yields $\|\langle r \rangle g\|_{\mathcal{LE}^N} \lesssim \|g\|_{\mathcal{LE}^{*,N}}$. The estimate is transferred to $\nabla^2 v$ using standard elliptic arguments. \square

4. High Energy Conormal Estimates

The goal of this section is to obtain pointwise bounds on $(\sigma \partial_\sigma)^M P_\sigma^{-1} g$ for $g \in \mathcal{LE}^*$. The results in this section are direct analogues to results in [39] and [34]. We provide sketches of the arguments, but detail how the regularity requirements depend on the loss, ϖ , in the weak local energy decay estimate (4).

We consider the vector fields $T \in \{\partial_{x_i} | i = 1, 2, 3\}$, $\Omega \in \{x_i \partial_{x_j} - x_j \partial_{x_i} | i, j = 1, 2, 3\}$, and $S = r \partial_r - \sigma \partial_\sigma$. Note if g is independent of σ then $Sg = r \partial_r g$. Since $r \partial_r = -\rho \partial_\rho$, this implies that

$$T^i \Omega^j S^k g \in L^2 \text{ for all } i + j + k \leq M \iff g \in H_b^{M,0}.$$

(Note that we include the T derivatives, which are powers of vector fields in $\rho \mathcal{V}_b$, merely to ensure differentiability at $r = 0$.) The assumptions here are stated in terms of the energy space \mathcal{LE}^* , and we note $\mathcal{LE}^* \subset r^{-\frac{1}{2}} L^2$.

In Lemmas 7 and 9 we take Q to represent an operator of the same form as $\sigma P^1 + P^2$ but let the exact coefficients change each time Q appears.

The following lemma is an analogue to [39] Proposition 10.

Lemma 7. *If $\text{Im } \sigma \geq 0$ and $g \in \mathcal{LE}^*$ satisfies*

$$(14) \quad \|T^i \Omega^j S^k g\|_{\mathcal{LE}^*} \lesssim 1, \quad i + (\varpi + 1)j + (\varpi + 1)k \leq M$$

for some positive integer M , then $v = P_\sigma^{-1} v$ satisfies

$$(15) \quad \|T^i \Omega^j S^k v\|_{\mathcal{LE}_\sigma} \lesssim 1, \quad i + (\varpi + 1)j + (\varpi + 1)k \leq M - \varpi - 1$$

Proof. Applying Proposition 6 to

$$P_\sigma T^i \Omega^j S^k v = T^i \Omega^j S^k g + [P_\sigma, T^i \Omega^j S^k] v$$

yields

$$\|T^i \Omega^j S^k v\|_{\mathcal{LE}_\sigma^N} \lesssim \|T^i \Omega^j S^k g\|_{\mathcal{LE}^{*,N+\varpi+1}} + \|[P_\sigma, T^i \Omega^j S^k] v\|_{\mathcal{LE}^{*,N+\varpi+1}}.$$

Since $\|T^i \Omega^j S^k v\|_{\mathcal{LE}_\sigma} = \|\Omega^j S^k v\|_{\mathcal{LE}_\sigma^i}$ we only require bounds on $\|\Omega^j S^k v\|_{\mathcal{LE}_\sigma^N}$. We illustrate the general method with concrete examples and highlight the role of ϖ in determining the requisite regularity.

Consider Ωv . Direct calculation shows $[P_\sigma, \Omega] = Q$ and

$$\|Q\phi\|_{\mathcal{LE}^{*,N}} \lesssim \|(\langle r \rangle^{-1} + |\sigma|)\phi\|_{\mathcal{LE}^N} + \|\nabla \phi\|_{\mathcal{LE}^N} =: \|\phi\|_{\mathcal{LE}_{\sigma,1}^N}$$

for ϕ with sufficient regularity and decay. Note that the $\mathcal{LE}_{\sigma,1}^N$ norm defined here is the first two terms of the \mathcal{LE}_{σ}^N norm (10), and by (13) we have $\|v\|_{\mathcal{LE}_{\sigma,1}^N} \lesssim \|g\|_{\mathcal{LE}^{*,N+\varpi}}$. Thus we find

$$\begin{aligned} \|\Omega v\|_{\mathcal{LE}_{\sigma}^N} &\lesssim \|\Omega g\|_{\mathcal{LE}^{*,N+\varpi+1}} + \|Qv\|_{\mathcal{LE}^{*,N+\varpi+1}} \\ &\lesssim 1 + \|v\|_{\mathcal{LE}_{\sigma,1}^{N+\varpi+2}} \\ &\lesssim 1 + \|g\|_{\mathcal{LE}^{*,N+2(\varpi+1)}}. \end{aligned}$$

We see there is one loss of $\varpi+1$ due to the estimate in Proposition 6 and a subsequent loss of $\varpi+1$ for each Ω . This justifies the requirement $i + (\varpi+1)j \leq M - \varpi - 1$.

Next we consider Sv . Note $[P_{\sigma}, S] = 2P_{\sigma} + Q$. We calculate

$$\begin{aligned} \|SP_{\sigma}^{-1}g\|_{\mathcal{LE}_{\sigma}^N} &\lesssim \|g\|_{\mathcal{LE}^{*,N+\varpi+1}} + \|Sg\|_{\mathcal{LE}^{*,N+\varpi+1}} + \|Qv\|_{\mathcal{LE}^{*,N+\varpi+1}} \\ &\lesssim 1 + \|v\|_{\mathcal{LE}_{\sigma,1}^{N+\varpi+2}} \\ &\lesssim 1 + \|g\|_{\mathcal{LE}^{*,N+2(\varpi+1)}}. \end{aligned}$$

As above, this demonstrates the requirement $i + (\varpi+1)j + (\varpi+1)k \leq M - \varpi - 1$. The general case follows by induction. Details are omitted here. \square

In Lemma 9 and Proposition 10 we will make use of the following result, which we quote from [39] Proposition 11:

Lemma 8. *If $\sigma \in \mathbb{R} \setminus \{0\}$ and $g \in \mathcal{LE}^{*,4}$, then $v = P_{\sigma}^{-1}g$ satisfies the outgoing radiation condition*

$$(16) \quad \lim_{m \rightarrow \infty} 2^{-\frac{m}{2}} \|(\partial_r - i\sigma)v\|_{L^2(A_m)} = 0.$$

A consequence of Lemma 8 and the proof of Lemma 7 is that for g satisfying (14), the radiation condition holds for $T^i \Omega^j S^k v$ with appropriate values of i, j , and k :

$$(17) \quad \lim_{m \rightarrow \infty} 2^{-\frac{m}{2}} \|(\partial_r - i\sigma)T^i \Omega^j S^k v\|_{L^2(A_m)} = 0, \quad i + (\varpi+1)j + (\varpi+1)k \leq M - 4.$$

As in the proof of Lemma 7, we see there is one loss of $\varpi+1$ due to Lemma 8 and a subsequent loss of $\varpi+1$ for each Ω and each S due to the commutator terms $[P_{\sigma}, \Omega]$ and $[P_{\sigma}, S]$.

Now we provide preliminary pointwise bounds on $T^i \Omega^j S^k P_{\sigma}^{-1}g$ using Sobolev embeddings and Proposition 7. The same result with different regularity assumptions can be found in [39] (cf. Proposition 16).

Let Γ denote the collection of all vector fields in Ω , T , and S . We write Γ^{α} to denote a product of these vector fields indexed by the multiindex α and let $\Gamma^{\leq n}$ denote a linear combination of Γ^{α} for $|\alpha| \leq n$: $\Gamma^{\leq n} := \sum_{|\alpha| \leq n} c_{\alpha} \Gamma^{\alpha}$. For the sake of notational simplicity, we write $v_{ijk} := T^i \Omega^j S^k v$ and $g_{ijk} := T^i \Omega^j S^k g$. Similarly we write $v_{\leq i \leq j \leq k} := T^{\leq i} \Omega^{\leq j} S^{\leq k} v$ and $g_{\leq i \leq j \leq k} := T^{\leq i} \Omega^{\leq j} S^{\leq k} g$.

Lemma 9. *Let $\text{Im } \sigma \geq 0$ with $|\sigma| \gtrsim 1$ and assume $g \in \mathcal{LE}^{*}$ satisfies (14).*

(i) *Then*

$$(18) \quad |T^i \Omega^j S^k P_{\sigma}^{-1}g| \lesssim (|\sigma| \langle r \rangle)^{-1}, \quad i + (\varpi+1)j + (\varpi+1)k \leq M - 5(\varpi+1)$$

(ii) *If, in addition, we have $\sigma \in \mathbb{R}$ and $M \geq 4$, then we have the outgoing radiation condition*

$$(19) \quad \lim_{r \rightarrow \infty} r(\partial_r - i\sigma)T^i\Omega^j S^k v(\sigma) = 0, \quad i + (\varpi + 1)j + (\varpi + 1)k \leq M - 5(\varpi + 1).$$

Proof. (i) To begin, we claim that establishing the estimate

$$(20) \quad \sum_m 2^{\frac{m}{2}} \|(\partial_r^2 + \sigma^2)(rv_{ijk})\|_{L_r^2 L_\omega^\infty(A_m)} \lesssim 1$$

will suffice to obtain (18). To see this, note that the fundamental solution to $(\partial_r^2 + \sigma^2)$ (for $\text{Im } \sigma > 0$) is given by $\phi_\sigma(s) = \sigma^{-1}e^{-i\sigma|s|}$. Thus (extending by continuity to $\text{Im } \sigma \geq 0$), $rv_{ijk} = (\partial_r^2 + \sigma^2)(rv_{ijk}) * \phi_\sigma$. Splitting

$$|rv_{ijk}| \leq \sum_m \int_{A_m} |(\partial_r^2 + \sigma^2)(rv_{ijk})(s)\phi_\sigma(r-s)| ds$$

we now use the fact that $|\phi_\sigma(s)| \lesssim |\sigma|^{-1}$ and apply the Cauchy–Schwarz inequality to each term in the sum to find

$$|\sigma||rv_{ijk}| \lesssim |\sigma| \cdot |\sigma|^{-1} \sum_m 2^{\frac{m}{2}} \|(\partial_r^2 + \sigma^2)(rv_{ijk})\|_{L_r^2 L_\omega^\infty(A_m)} \lesssim 1.$$

This yields (18) with $\langle r \rangle$ replaced by r , i.e.,

$$(21) \quad |rT^i\Omega^j S^k P_\sigma^{-1}g| \lesssim |\sigma|, \quad i + (\varpi + 1)j + (\varpi + 1)k \leq M - 5(\varpi + 1).$$

That the estimate in fact holds for $\langle r \rangle$ essentially follows from the fact that there is no preferred origin to our coordinate system, and our estimates are translation-invariant. In particular, letting \mathcal{T} denote the translation operation $\mathcal{T}f(x) = f(x-a)$ with a a fixed nonzero vector, we observe that \mathcal{T} commutes with translation vector fields T_i , while

$$\mathcal{T}\Omega - \Omega\mathcal{T}, \quad \mathcal{T}S - S\mathcal{T}$$

are both of the form $\sum c_j T_j \mathcal{T} = \sum c_j \mathcal{T} T_j$. Thus

$$T^i\Omega^j S^k \mathcal{T} - \mathcal{T}T^i\Omega^j S^k = \sum \Gamma^\alpha \mathcal{T} = \sum \mathcal{T}\tilde{\Gamma}^\alpha$$

where sums are over products of vector fields $\Gamma^\alpha, \tilde{\Gamma}^\alpha$ of the form $T^{i'}\Omega^{j'}S^{k'}$ with $i' + (\varpi + 1)j' + (\varpi + 1)k' \leq i + (\varpi + 1)j + (\varpi + 1)k$.

Note further that the hypotheses on P are translation-invariant, so that if $P_\sigma v = g$ then $P'_\sigma \mathcal{T}v = \mathcal{T}g$ with P'_σ an operator satisfying the same hypotheses. Owing to the commutation properties of \mathcal{T} with our rest operators, $\mathcal{T}g$ also satisfies (14), hence $\mathcal{T}v$ likewise satisfies (21). Now translating back and again using the commutation properties of the vector fields Γ^α shows that

$$(22) \quad |(x-a)T^i\Omega^j S^k v| \lesssim |\sigma|, \quad i + (\varpi + 1)j + (\varpi + 1)k \leq M - 5(\varpi + 1).$$

Adding (21) and (22) yields the desired estimate.

To establish (20) we write

$$(\partial_r^2 + \sigma^2) = -P_\sigma - (2r^{-1}\partial_r + r^{-2}\Delta_\theta + Q)$$

and commute P_σ with $T^i\Omega^j S^k$ to find

$$(23) \quad (\partial_r^2 + \sigma^2)(rv_{ijk}) = -r^{-1}\Delta_\theta v_{ijk} + rQv_{\leq i \leq j \leq k} - r g_{\leq i \leq j \leq k}$$

and bound each term on the right hand side as in (20).

To see where the vector field loss occurs, consider the term $r^{-1}\Delta_\theta v_{ijk}$ on the right hand side of (23). We wish to show this term satisfies (20). By the spherical Sobolev embedding, it suffices to show

$$(24) \quad \sum_m \|\langle r \rangle^{-\frac{3}{2}} \Delta_\theta v_{ijk}\|_{L^2(A_m)} + \|\langle r \rangle^{-\frac{3}{2}} \Omega^2(\Delta_\theta v_{ijk})\|_{L^2(A_m)} \lesssim 1.$$

We have

$$(25) \quad \|\langle r \rangle^{-\frac{1}{2}} v_{ijk}\|_{L^2(A_m)} \lesssim |\sigma| \|\langle r \rangle^{-\frac{1}{2}} v_{ijk}\|_{L^2(A_m)} \lesssim \|v_{ijk}\|_{\mathcal{LE}_\sigma} \lesssim 1$$

when $i + (\varpi + 1)j + (\varpi + 1)k < M - \varpi - 1$.

Replacing Δ_θ by $\sum_{|\alpha|=2} \Omega^\alpha$ and using (25) then yields

$$\begin{aligned} LHS \text{ of (24)} &= \sum_m \|\langle r \rangle^{-\frac{3}{2}} v_{i(j+2)k}\|_{L^2(A_m)} + \|\langle r \rangle^{-\frac{3}{2}} v_{i(j+4)k}\|_{L^2(A_m)} \\ &\lesssim 1 \end{aligned}$$

when $i + (\varpi + 1)j + (\varpi + 1)k \leq M - 5(\varpi + 1)$, as desired. The remaining terms on the right hand side of (23) are handled similarly.

(ii) Note that we have

$$(\partial_r + i\sigma)(\partial_r - i\sigma)(rv_{ijk}) = (\partial_r^2 + \sigma^2)(rv_{ijk})$$

so that

$$(26) \quad \partial_r(e^{i\sigma r}(\partial_r - i\sigma)(rv_{ijk})) = (\partial_r^2 + \sigma^2)(rv_{ijk})e^{i\sigma r}.$$

By (20) we see that $(\partial_r^2 + \sigma^2)(rv_{ijk})$ is integrable in r , so that the limit $\lim_{|x| \rightarrow \infty} |(\partial_r - i\sigma)(rv_{ijk})|$ exists for each $\theta \in \mathbb{S}^2$ since

$$\lim_{|x| \rightarrow \infty} |(\partial_r - i\sigma)(rv_{ijk})| = \left| \int_0^\infty (\partial_r^2 + \sigma^2)(rv_{ijk})e^{i\sigma r} dr + v_{ijk}(0) \right|.$$

For fixed $\theta \in \mathbb{S}^2$, take

$$c_\theta = \lim_{|x| \rightarrow \infty} |(\partial_r - i\sigma)(rv_{ijk})| = \lim_{|x| \rightarrow \infty} |r(\partial_r - i\sigma)v_{ijk} + v_{ijk}|.$$

By part (i) of this proposition, $\lim_{r \rightarrow \infty} v_{ijk} = 0$ so that

$$\lim_{r \rightarrow \infty} |r(\partial_r - i\sigma)v_{ijk}| = c_\theta.$$

Thus we can write $|(\partial_r - i\sigma)v_{ijk}| = c_\theta r^{-1} + o(r^{-1})$.

On the other hand, since σ is real and $5(\varpi + 1) \geq 4$, by Lemma 8 v_{ijk} satisfies the radiation condition (17). Then Sobolev embedding yields

$$(27) \quad \lim_{m \rightarrow \infty} 2^{\frac{m}{2}} \|(\partial_r - i\sigma)v_{ijk}\|_{L_r^2 L_\theta^\infty(A_m)} = 0.$$

It follows that

$$0 \geq \lim_{m \rightarrow \infty} 2^{\frac{m}{2}} \|c_\theta r^{-1} + o(r^{-1})\|_{L_r^2(A_m)} = \lim_{m \rightarrow \infty} 2^{\frac{m}{2}} c_\theta 2^{-\frac{m}{2}} = c_\theta$$

Thus $c_\theta \equiv 0$ for $\theta \in \mathbb{S}^2$ so that $(\partial_r - i\sigma)v_{ijk} \in o(r^{-1})$, which concludes the proof of (19). □

Finally we provide the following result on the finite order conormal regularity of P_σ^{-1} for large σ (see [39] Proposition 17 and [34] Proposition 6.5). We give a sketch of the proof, with a focus on the numerology in the proposition.

Proposition 10. *Let $|\sigma| \gtrsim 1$ with σ real and take $g \in H_b^{s,\ell}$ with $\ell > \frac{1}{2}$. Then*

$$|(\sigma \partial_\sigma)^p (e^{-i\sigma r} P_\sigma^{-1} g)| \lesssim |\sigma|^{p-1} \langle r \rangle^{-1+p(1-\varepsilon)}, \quad p \leq \ell - \frac{1}{2}, \quad s \geq (2p+5)(\varpi+1)$$

for some $\varepsilon \in (0, 1]$.

Proof. Identifying the vector fields $T, \Omega, r\partial_r$ with b-vector fields and using the change of coordinates $\rho = r^{-1}$, we see $g \in H_b^{s,\ell}$ implies

$$\|r^{\ell-\frac{1}{2}} T^i \Omega^j S^k g\|_{\mathcal{LE}^*} \lesssim 1$$

for $i+j+k \leq s$.

Take $v = P_\sigma^{-1} g$ and write

$$\begin{aligned} (\sigma \partial_\sigma)^p (v e^{-i\sigma r}) &= ((-S + r(\partial_r - i\sigma))^p v) e^{-i\sigma r} \\ (28) \quad &= \left(\sum_{m=0}^p c_m [r(\partial_r - i\sigma)]^m (-S)^{p-m} v \right) e^{-i\sigma r} \\ &= \left(-S^p v + \sum_{m=1}^p \sum_{n=1}^m c_{mn} r^n (\partial_r - i\sigma)^n (-S)^{p-m} v \right) e^{-i\sigma r}. \end{aligned}$$

We claim $|(\partial_r - i\sigma)^p v_{ijk}| \lesssim |\sigma|^{p-1} r^{-1-p\varepsilon}$ with $\varepsilon \in (0, 1]$ for

$$i + (\varpi+1)j + (\varpi+1)k \leq s - 5(\varpi+1) - 2(\varpi+1)p.$$

The proposition follows from the claim. To see this, assume the claim holds. Then $|r^n (\partial_r - i\sigma)^n (-S)^{p-m} v| \lesssim |\sigma|^{n-1} r^{-1+n(1-\varepsilon)}$ when

$$(\varpi+1)p \leq s - 5(\varpi+1) + (\varpi+1)(m-2n)$$

for $0 \leq n \leq m \leq p$. The smallest value of $s - 5(\varpi+1) + (\varpi+1)(m-2n)$ in this range is $s - (5+p)(\varpi+1)$, which justifies the assumption $s \geq (2p+5)(\varpi+1)$.

It is left to prove the claim. When $p=0$, the claim follows from (18). For general p , applying $(\partial_r - i\sigma)^{p-1}$ to (23) we find

$$\begin{aligned} (29) \quad &(\partial_r + i\sigma)(\partial_r - i\sigma)^p (rv_{ijk}) \\ &= \sum_{m=0}^{p-1} \left((-1)^{p-m+1} c_m r^{-(p-m)} (\partial_r - i\sigma)^m \Delta_\theta v_{ijk} \right) + r(\partial_r - i\sigma)^{p-1} Q v_{\leq i \leq j \leq k} \\ &\quad + C(\partial_r - i\sigma)^{p-2} Q v_{\leq i \leq j \leq k} + (r(\partial_r - i\sigma)^{p-1} + C(\partial_r - i\sigma)^{p-2}) g_{\leq i \leq j \leq k}. \end{aligned}$$

Each $T^i \Omega^j S^k v$ term on the right hand side of (29) is bounded in magnitude by $|\sigma|^{p-1} \langle r \rangle^{-1-p\varepsilon}$ inductively using (18). We note the ε factor in the exponent shows up when considering, for example, the $\sigma S^{-\kappa} \partial_x$ term in Q which yields a $S^{-\kappa+1}(\partial_r -$

$i\sigma)^{p-1}v_{i+1,j,k}$ term on the right hand side of (29)². This requires

$$i + (\varpi + 1)j + (\varpi + 1)k \leq s - 5(\varpi + 1) - 2(\varpi + 1)p.$$

The loss of $2(\varpi + 1)p$ follows from the Δ_θ operator on the right hand side of (29), which we replace by Ω^2 .

To handle the $g_{\leq i \leq j \leq k}$ terms we note $g_{\leq i \leq j \leq k} \in H_b^{s-i-j-k, \ell}$ and by Sobolev embeddings we have $|\partial_r^p g_{\leq i \leq j \leq k}| \lesssim \langle r \rangle^{-\ell - \frac{3}{2} - p}$ if $s - i - j - k - p > 3$, which indeed holds from the above restrictions on i, j, k . Now we calculate

$$\begin{aligned} & |r(\partial_r - i\sigma)^{p-1}g_{\leq i \leq j \leq k}| + |(\partial_r - i\sigma)^{p-2}g_{\leq i \leq j \leq k}| \\ &= \left| r \sum_{m=0}^{p-1} c_m (-i\sigma)^m \partial_r^{p-1-m} g_{\leq i \leq j \leq k} \right| + \left| \sum_{m=0}^{p-2} (-i\sigma)^m c_m \partial_r^{p-2-m} g_{\leq i \leq j \leq k} \right| \\ &\lesssim |\sigma|^{p-1} \langle r \rangle^{-\ell - \frac{1}{2}} \end{aligned}$$

which yields the desired $|\sigma|^{p-1} \langle r \rangle^{-1-p}$ bound when $p \leq \ell - \frac{1}{2}$.

We write

$$\partial_r [(\partial_r - i\sigma)^p r v_{ijk}] e^{i\sigma r} = [(\partial_r + i\sigma)(\partial_r - i\sigma)^p (r v_{ijk})] e^{i\sigma r}.$$

When $p = 1$, Lemma 9 shows $\lim_{r \rightarrow \infty} (\partial_r - i\sigma)(r v_{ijk}) = 0$, so we can integrate from infinity to prove the claim. When $p > 1$ note that

$$(\partial_r - i\sigma)v_{ijk} = r^{-1}x \cdot \nabla v_{ijk} - i\sigma v$$

and thus

$$\begin{aligned} |(\partial_r - i\sigma)^2 v_{ijk}| &= |x r^{-1} (\partial_r - i\sigma) v_{(i+1)jk} - i\sigma (\partial_r - i\sigma) v| \\ &\lesssim |(\partial_r - i\sigma) v_{(i+1)jk}| + |\sigma| |(\partial_r - i\sigma) v|. \end{aligned}$$

Iterating, we find

$$|(\partial_r - i\sigma)^p v_{ijk}| \lesssim \sum_{m=0}^{p-1} |\sigma|^m |(\partial_r - i\sigma) v_{(i+p-1-m)jk}|$$

and thus $r(\partial_r - i\sigma)^p v_{ijk} \rightarrow 0$ as $r \rightarrow \infty$ by (19) for

$$i + (\varpi + 1)j + (\varpi + 1)k \leq s - 5(\varpi + 1) - (p - 1)$$

which is satisfied under the conditions of the claim since $-2(\varpi + 1)p \leq -p + 1$. Thus when $p > 1$ we can integrate from infinity just as in the $p = 1$ case to prove the claim. \square

²When $\kappa > 2$ we have $\varepsilon = 1$ since in this case $S^{-\kappa+1} \subseteq S^{-1}$. When $\kappa \in (1, 2)$, it should be possible to improve the estimate to $\varepsilon = 1$. We roughly illustrate the idea of the process by considering

$$(\sigma \partial_\sigma)(v e^{-i\sigma r}) = ([-S + r(\partial_r - i\sigma)]v) e^{-i\sigma r}$$

in which case we have

$$(\partial_r^2 + \sigma^2)(rv) = -r^{-1}\Delta_\theta v + rQv - rg$$

and the $S^{-\kappa}$ coefficient terms in Q do not satisfy the r^{-2} pointwise bounds. In this case we write

$$(\partial_r + i\sigma)(\partial_r - i\sigma)(rv) = S^{-\kappa}(\partial_r - i\sigma)(rv) + \mathcal{O}(r^{-2})$$

then solve for $(\partial_r - i\sigma)(rv)$ by integrating as above to get $(\partial_r - i\sigma)v \lesssim r^{-2}$, as needed.

5. Low Frequency Conormal Estimates

The purpose of this section is to provide information about the behavior of $(\sigma\partial_\sigma)^M P_\sigma^{-1}f$ for f in an appropriately chosen function space. We follow the approach introduced by Hintz in [19].

In order to obtain good asymptotic expansions as $\sigma \rightarrow 0$, it is useful to work with the conjugated operator

$$\tilde{P}(\sigma) \equiv e^{-ir\sigma} P_\sigma e^{ir\sigma}.$$

For $\sigma \in \mathbb{R}$, we define $\tilde{P}(\sigma)^{-1}$ to be $\lim_{\varepsilon \downarrow 0} \tilde{P}(\sigma + i\varepsilon)^{-1}$. Note that $\tilde{P}(0) = P_0$.

Working on the spatial manifold X , we find

$$\begin{aligned} \tilde{P}(\sigma) &= \rho^2 [-(\rho\partial_\rho)^2 + \rho\partial_\rho + \Delta_\theta] + 2i\rho\sigma(\rho\partial_\rho - 1) + \mathcal{A}^\kappa(X)\sigma^2 \\ &\quad + \mathcal{A}^{\kappa+1}(X)\sigma \text{Diff}_b^{-1}(X) + \mathcal{A}^{\kappa+2} \text{Diff}_b^{-2}(X), \end{aligned}$$

where the operator class

$$\mathcal{A}^{\kappa+2} \text{Diff}_b^{-2}(X)$$

refers to b-differential operators with conormal coefficients of the specified order.

Note that

$$(30) \quad \tilde{P}(\sigma) = \tilde{P}(0) - \sigma R$$

where

$$(31) \quad R = -2i\rho(\rho\partial_\rho - 1) + \mathcal{A}^{\kappa+1} \text{Diff}_b^{-1}(X) + \sigma\mathcal{A}^\kappa.$$

In Proposition 20 we study the low frequency behavior of $\tilde{P}(\sigma)^{-1}f$ by using a formal Neumann series argument to write

$$\tilde{P}(\sigma)^{-1} = \tilde{P}(0)^{-1} (\text{Id} + \cdots + (\sigma R \tilde{P}(0)^{-1})^N) + \tilde{P}(\sigma)^{-1} (\sigma R \tilde{P}(0)^{-1})^{N+1}.$$

The key feature of this expansion is that each iterative application of $R\tilde{P}(0)^{-1}$ results in the loss of one power of decay. After enough iterations, the output will be too large to apply $\tilde{P}(0)^{-1}$ again, which forces the iteration to stop with the final application of $\tilde{P}(\sigma)^{-1}$.

In subsection 5.1 we record a previously established preliminary estimate which will aid our subsequent calculations. We then prove the necessary mapping properties for $\tilde{P}(0)^{-1}$ and $\tilde{P}(\sigma)^{-1}$ in subsection 5.2. This includes an expansion for $\tilde{P}(0)^{-1}f$ in Lemma 15 (and in Lemma 16 for larger inputs). We note that in Lemma 15, we see that even if f is rapidly decaying, the expansion for $\tilde{P}(0)^{-1}$ is limited by the perturbative $\rho^\kappa L_1$ term in $\tilde{P}(0)$. It is this limitation which dictates when the Neumann series must end, which in turn dictates the final decay rate obtained in the proof of the main theorem. Finally, in subsection 5.3 we study the Neumann series and obtain the desired conormal estimates.

5.1. Preliminary results. The following result from [40] (cf. Theorem 1.1) will be used to analyze the low energy resolvent:

Theorem 11. *For $s, \ell, \nu \in \mathbb{R}$ with $\ell < -\frac{1}{2}$, $s + \ell > -\frac{1}{2}$, $\ell - \nu \in (-\frac{3}{2}, -\frac{1}{2})$ the bound*

$$\|(\rho + |\sigma|)^\nu u\|_{H_b^{s,\ell}(X)} \lesssim \|(\rho + |\sigma|)^{\nu-1} \tilde{P}(\sigma) u\|_{H_b^{s,\ell+1}(X)}$$

holds for bounded σ .

Note that this result is obtained from taking $r = s + l$ in Theorem 1.1 of [40] to obtain

$$\|(\rho + |\sigma|)^\nu u\|_{H_{sc,b,res}^{s,r,l}} \lesssim \|(\rho + |\sigma|)^{\nu-1} \tilde{P}(\sigma) u\|_{H_{sc,b,res}^{s-2,r+1,l+1}};$$

these triple-index Sobolev spaces measure a combined (“second microlocal”) b and scattering regularity. We then coarsen the estimate by estimating the right hand-side by

$$\|(\rho + |\sigma|)^{\nu-1} \tilde{P}(\sigma) u\|_{H_{sc,b,res}^{s,r+1,l+1}};$$

Since $H_{sc,b,res}^{s,s+l,l} = H_b^{s,l}$ the result then follows.

We remark here that giving up two derivatives as we do in Theorem 11 in return for weighted estimates does have consequences for the numerology of our decay hypotheses on initial data (mainly owing to the conjugating factor $e^{i\sigma/\rho}$ which means that non-Schwartz data has limited b -regularity after multiplication by this factor). It is possible that a finer accounting of regularity in our iteration, tracked in the second microlocal Sobolev spaces of [40], would yield more precise decay hypotheses.

We use this theorem to establish basic mapping properties of $\tilde{P}(0)^{-1}$ that will help us find a useful expression for $\tilde{P}(0)^{-1} f$.

Corollary 12. *Let $\ell \in (-\frac{3}{2}, -\frac{1}{2})$ and $s + \ell > -\frac{1}{2}$. Then*

$$\tilde{P}(0)^{-1} : H_b^{s,\ell+2} \rightarrow H_b^{s,\ell}.$$

5.2. Low frequency mapping properties. Following Hintz [19], we write

$$\tilde{P}(0) = \rho^2(L_0 + \rho^\kappa L_1)$$

with

$$\begin{aligned} L_0 &= -(\rho \partial_\rho)^2 + \rho \partial_\rho + \Delta_\theta, \\ L_1 &\in \mathcal{A}^0 \text{Diff}_b^2. \end{aligned}$$

The conclusion of the following lemma involves sums over finite-dimensional spaces of spherical harmonics, schematically denoted

$$\sum_{j=1}^M \rho^j Y_{j-1}$$

Since all Sobolev norms are equivalent on such spaces, we write

$$|Y_{j-1}|$$

to denote the supremum of each component, but will use tacitly the fact that this is equivalent to taking any desired b -Sobolev norm of these angular pieces, as well.

We begin by considering L_0^{-1} then argue perturbatively to analyze $\tilde{P}(0)^{-1}$.

Lemma 13. *Let $f \in H_b^{s,\gamma}$ with $\gamma > -\frac{1}{2}$ and $s > 0$. Assume $u \in H_b^{s,-\frac{1}{2}-}$ solves $L_0 u = f$. If $\gamma + \frac{3}{2} \notin \mathbb{N}$ then*

$$u = \sum_{j=1}^{\lfloor \gamma + \frac{3}{2} \rfloor} \rho^j Y_{j-1} + q$$

where Y_j is a linear combination of j^{th} order spherical harmonics, $q \in H_b^{s+2, \gamma-}$, and

$$\sum_{j=1}^{\lfloor \gamma + \frac{3}{2} \rfloor} |Y_{j-1}| + \|q\|_{H_b^{s+2, \gamma-}} \lesssim \|f\|_{H_b^{s, \gamma}}.$$

Proof. We are interested in the behavior of u as $\rho \rightarrow 0$. Take $\chi_{\partial}(\rho)$ to be a cutoff which is 1 on a neighborhood of ∂X and 0 for $\rho \geq 1$. Define $u_{\partial} := \chi_{\partial} u$. Then we have

$$(32) \quad L_0 u_{\partial} = f_{\partial} \in H_b^{s, \gamma}(X)$$

with $\text{supp}(f_{\partial}) \subset [0, 1]_{\rho}$.

The Mellin transform in ρ of a function g defined on X is given by $\mathcal{M}g(\xi) := \int_0^{\infty} \rho^{-i\xi} g(\rho) \frac{d\rho}{\rho}$. Taking the Mellin transform in ρ of Equation (32) yields

$$\hat{L}_0(\xi) \mathcal{M}u_{\partial}(\xi) = \mathcal{M}f_{\partial}(\xi), \quad \hat{L}_0(\xi) = -(i\xi)^2 + i\xi + \Delta_{\theta}.$$

The Mellin transform of f_{∂} is holomorphic in $\text{Im } \xi > -\gamma - \frac{3}{2}$ (see Proposition 25 in Appendix A). Since $u \in H_b^{s, -\frac{1}{2}-}$ by assumption, we can invert the Mellin transform of u_{∂} by integrating along a contour in $\text{Im } \xi > -1$. Thus we obtain

$$(33) \quad u_{\partial}(\rho) = \int_{\text{Im } \xi = -1+\varepsilon} \rho^{i\xi} \mathcal{M}u_{\partial}(\xi) d\xi.$$

For each $\xi \in \mathbb{C}$ we decompose $\mathcal{M}u_{\partial}(\xi)$ and $\mathcal{M}f_{\partial}(\xi)$ into spherical harmonics: $\mathcal{M}u_{\partial}(\xi) = \sum_{j=0}^{\infty} \sum_{m=-j}^j \mathcal{M}u_{mj}(\xi) y_{mj}$ where

$$\mathcal{M}u_{mj}(\xi) = \int_{\mathbb{S}^2} \mathcal{M}u(\xi) y_{mj} d\theta$$

and similarly for $\mathcal{M}f_{\partial}$. Define $\mathcal{Y}_j := \text{span}_m(y_{mj})$ and note $\hat{L}_0(\xi)|_{\mathcal{Y}_j} = -(i\xi)^2 + i\xi + j(j+1)$. Thus $\hat{L}_0(\xi)^{-1}|_{\mathcal{Y}_j} = \frac{1}{-(i\xi)^2 + i\xi + j(j+1)}$ has simple poles at $\xi = ij, -i(j+1)$. Using the spherical harmonic decomposition we have

$$\mathcal{M}u_{\partial}(\xi) = \sum_{j=0}^{\infty} \sum_{m=-j}^j \hat{L}_0(\xi)^{-1}|_{\mathcal{Y}_j} \mathcal{M}f_{mj}(\xi) y_{mj}$$

Since $f_{\partial} \in H_b^{s, \gamma-}$, we can push the contour of integration in (33) down to $\text{Im } \xi = -\gamma - \frac{3}{2} + \varepsilon$ and pick up residues at $\xi = -i, -2i, \dots, -\lfloor \gamma + \frac{3}{2} \rfloor i$. The residue of $\rho^{i\xi} \mathcal{M}u_{k-1}(\xi)$ at $\xi = -ik$ is

$$\rho^k \sum_{m=-k+1}^{k-1} i(2k-1)^{-1} \mathcal{M}f_{m(k-1)}(-ik) y_{m(k-1)}.$$

Thus

$$u_{\partial}(\rho, \theta) = \sum_{k=1}^{\lfloor \gamma + \frac{3}{2} \rfloor} \rho^k Y_{k-1} + \int_{\text{Im } \xi = -\gamma - \frac{3}{2} + \varepsilon} \rho^{i\xi} \mathcal{M}u_{\partial}(\xi) d\xi$$

where $Y_{k-1} = \sum_{m=-k+1}^{k-1} i(2k-1)^{-1} \mathcal{M}f_{m(k-1)}(-ik)y_{m(k-1)}$ is a linear combination of $(k-1)^{th}$ order spherical harmonics. By Cauchy-Schwarz,

$$|\mathcal{M}f_{\partial}(-ik)|^2 \lesssim \int_0^\infty \rho^{-2\gamma} |f_{\partial}|^2 \frac{d\rho}{\rho^4}$$

when $k < \gamma + \frac{3}{2}$. Thus

$$|\mathcal{M}f_{m(k-1)}(-ik)|^2 \leq \int_{\mathbb{S}^2} |\mathcal{M}f_{\partial}(-ik)y_{m(k-1)}| d\theta \lesssim \|f_{\partial}\|_{H_b^{s,\gamma}}$$

and $|Y_{k-1}| \lesssim \|f_{\partial}\|_{H_b^{s,\gamma}}$ as desired. Finally we have obtained

$$q = \int_{\text{Im } \xi = -\gamma - \frac{3}{2} + \epsilon} \rho^{i\xi} \hat{L}_0(\xi)^{-1} \mathcal{M}f_{\partial}(\xi) d\xi.$$

That $\mathcal{M}q$ is holomorphic in $\text{Im}(\bullet) > -\gamma - 3/2 + \epsilon$ follows since q differs from

$$\mathcal{M}^{-1}(\hat{L}_0(\xi)^{-1} \mathcal{M}f_{\partial}(\xi))$$

by subtraction of the poles of $\hat{L}_0(\xi)^{-1} \mathcal{M}f_{\partial}(\xi)$ in this region. Consequently, since $\hat{L}_0(\xi)|_{\mathcal{Y}_j}$ grows quadratically in both ξ and j , for any $\mu > -\gamma - 3/2 + \epsilon$ and for $\nu \in \mathbb{R}$,

$$|\mathcal{M}q_j(\nu + i\mu)| \leq C(1 + j^2 + \nu^2)^{-1} |\mathcal{M}f_j(\nu + i\mu)|,$$

hence

$$\|(\xi^2 + \Delta_{\theta})^{s+2} \mathcal{M}q(\nu + i\mu)\|_{L_{\nu}^2}^2 \leq C \left\| (1 + \xi^2 + \Delta_{\theta})^{s/2} \mathcal{M}f_{\partial}(\nu + i\mu) \right\|_{L_{\nu}^2}^2$$

and the estimate $\|q\|_{H_b^{s+2,\gamma-}} \lesssim \|f_{\partial}\|_{H_b^{s,\gamma}}$ follows, using Proposition 25. \square

The following lemma is analogous to Lemma 13 but with less decay assumed for f . It is stated separately because of minor technical changes in the numerology (note we now assume $u \in H_b^{s,\gamma-}$). The proof is the same and we provide an abbreviated argument.

Lemma 14. *Let $f \in H_b^{s,\gamma}$ with $\gamma \in (-\frac{3}{2}, -\frac{1}{2})$. Assume $u \in H_b^{s,\gamma-}$ solves $L_0 u = f$. Then*

$$\|u\|_{H_b^{s+2,\gamma-}} \lesssim \|f\|_{H_b^{s,\gamma}}.$$

Proof. The lemma is proved as in Lemma 13. Instead of (33) the assumptions on u allow us to integrate along $\text{Im } \xi = -\gamma - \frac{3}{2} + \epsilon$. Note $-\gamma - \frac{3}{2} + \epsilon \in (0, -1)$. Since f is assumed to have the same decay, we cannot push the contour of integration further into the lower half plane. The result follows immediately using Proposition 25 as before. \square

The preceding lemmas will now be used to establish asymptotic expansions for $\tilde{P}(0)^{-1}f$ which depend on the amount of decay assumed for f .

Lemma 15. *Let $f \in H_b^{s,\gamma}$ for $s > 0$ and $\frac{3}{2} < \gamma \leq \frac{3}{2} + \kappa$ with $\gamma + \frac{3}{2} \notin \mathbb{N}$. Then*

$$\tilde{P}(0)^{-1}f = \sum_{j=1}^{\lfloor \gamma - \frac{1}{2} \rfloor} \rho^j Y_{j-1} + q$$

with

$$q \in H_b^{s+2, \gamma-2-}$$

and

$$\sum_{j=1}^{\lfloor \gamma - \frac{1}{2} \rfloor} |Y_{j-1}| + \|q\|_{H_b^{s+2, \gamma-2-}} \lesssim \|f\|_{H_b^{s, \gamma-}}.$$

Proof. Define $u := \tilde{P}(0)^{-1}f$ and note that $u \in H_b^{s, -\frac{1}{2}-}$ by Corollary 12. Writing $L_0 = \rho^{-2}\tilde{P}(0) - \rho^\kappa L_1$ with $L_1 \in \mathcal{A}^0 \text{Diff}_b^2$, we find $L_0 u = \rho^{-2}f + \rho^\kappa L_1 u$. Then by Lemma 13 we have

$$u = \sum_{j=1}^{\lfloor \gamma - \frac{1}{2} \rfloor} \rho^j Y_{j-1} + q$$

with

$$\begin{aligned} \sum_{j=1}^{\lfloor \gamma - \frac{1}{2} \rfloor} |Y_{j-1}| + \|q\|_{H_b^{s+2, \gamma-2-}} &\lesssim \left\| \rho^{-2}f + \rho^\kappa L_1 \left(\sum_{j=1}^{\lfloor \gamma - \frac{1}{2} \rfloor} \rho^j Y_{j-1} + q \right) \right\|_{H_b^{s, \gamma-2-}} \\ &\lesssim \|f\|_{H_b^{s, \gamma-}} + \varepsilon \left(\sum_{j=1}^{\lfloor \gamma - \frac{1}{2} \rfloor} |Y_{j-1}| + \|q\|_{H_b^{s+2, \gamma-2-}} \right) \end{aligned}$$

for ρ small for some $0 < \varepsilon < 1$. (Recall that all norms are comparable to one another on the finitely many angular modes Y_j .) Bootstrapping the last terms on the right hand side above to the left hand side then yields the desired inequality. \square

Note even if f has faster decay as $\rho \rightarrow 0$ than assumed in Lemma 15, there is no improvement over the result for $f \in H_b^{s, \kappa + \frac{3}{2}}$ due to the perturbative $\rho^\kappa L_1$ term in $\tilde{P}(0)$.

As before, we prove a lemma analogous to Lemma 15 now assuming f has less decay. More regularity is also assumed for f due to the numerology in Corollary 12.

Lemma 16. *Let $f \in H_b^{s, \gamma}$ for $s > 1$ and $\gamma \in (\frac{1}{2}, \frac{3}{2})$ with $\gamma + \frac{3}{2} \notin \mathbb{N}$. Then $u = \tilde{P}(0)^{-1}f$ satisfies*

$$\|u\|_{H_b^{s+2, \gamma-2-}} \lesssim \|f\|_{H_b^{s, \gamma}}.$$

Proof. The proof is analogous to that of Lemma 15 except Corollary 12 now implies $u \in H_b^{s, \gamma-2-}$. Note $\gamma - 2 \in (-\frac{3}{2}, -\frac{1}{2})$. As before we find $L_0 u = \rho^{-2}f - \rho^\kappa L_1 u$. Now we use Lemma 14 to find

$$\|u\|_{H_b^{s+2, \gamma-2-}} \lesssim \|\rho^{-2}f\|_{H_b^{s, \gamma-2-}} + \|\rho^\kappa L_1 u\|_{H_b^{s, \gamma-2-}} \lesssim \|f\|_{H_b^{s, \gamma-}} + \epsilon \|u\|_{H_b^{s+2, \gamma-2-}}$$

for ρ small for some $0 < \epsilon < 1$. \square

Since $P(0)^{-1}f$ generates terms of the form $\rho^n Y_\ell$ and we will apply $P(0)^{-1}$ iteratively, we now consider the output of $P(0)^{-1}\rho^n Y_\ell$.

Lemma 17. *If $3 \leq n$ and $n \neq \ell + 3$, then*

$$\tilde{P}(0)^{-1}(\rho^n Y_\ell) = \rho^{n-2} Y_\ell + \sum_{j=1}^{\lfloor \kappa \rfloor + 1} \rho^j Y_{j-1} + \mathcal{A}^{\kappa+1-}$$

Proof. Note for general m we have $L_0 \rho^m Y_\ell = -(m+\ell)(m-\ell-1)\rho^m Y_\ell$ so $L_0^{-1}(\rho^m Y_\ell) = c\rho^m Y_\ell$ when $m \neq -\ell, \ell+1$.

Define $u := \tilde{P}(0)^{-1}(\rho^n Y_\ell)$. By Sobolev embeddings (6) we have $\rho^n Y_\ell \in \mathcal{A}^n \subseteq H_b^{\infty, n-\frac{3}{2}-}$. It follows from Corollary 12 that $u \in H_b^{\infty, -\frac{1}{2}-}$. As before we use $L_0 = \rho^{-2}\tilde{P}(0) - \rho^\kappa L_1$ to write $L_0 u = \rho^{n-2} Y_\ell + H_b^{\infty, \kappa-\frac{1}{2}-}$. The result then follows by Lemma 13 and using our assumptions on n to find $L_0^{-1}(\rho^{n-2} Y_\ell) = c\rho^{n-2} Y_\ell$. \square

We now consider the mapping properties of $P(\sigma)^{-1}$, which will be needed in the last term in the Neumann series. The following lemma is implicit in the proof of Lemma 2.16 in [19], and our proof follows that in [19].

Lemma 18. *Let $\alpha \in (0, 1)$ and $s > 1$. Then there exists σ_0 such that $|\sigma| \leq \sigma_0$ implies that for all $\delta > 0$ sufficiently small,*

$$\tilde{P}(\sigma)^{-1} : H_b^{s, 1/2-\alpha} \rightarrow |\sigma|^{-\alpha-\delta} H_b^{s, -3/2-\delta}.$$

In [19] this is used to show (by taking $s \rightarrow \infty$) that

$$\tilde{P}(\sigma)^{-1} : \mathcal{A}^{2-\alpha} \rightarrow |\sigma|^{-\alpha-0} \mathcal{A}^{-0}$$

(and indeed, a more refined statement holds on the resolved space).

Proof. We take $l = -3/2 - \delta$ in Theorem 11. Then the constraint on s is $s > 1 + \delta$, and is satisfied if $\delta > 0$ is sufficiently small. The constraint on ν is $\nu \in (-1 - \delta, -\delta)$; we take $\nu = -2\delta$ to obtain

$$\|(\rho + |\sigma|)^{-2\delta} u\|_{s, -3/2-\delta} \lesssim \|(\rho + |\sigma|)^{-1-2\delta} \tilde{P}(\sigma) u\|_{s, -1/2-\delta}.$$

Estimating

$$(\rho + |\sigma|)^{-1-2\delta} \leq \rho^{-1-\delta+\alpha} |\sigma|^{-\delta-\alpha}$$

allows us to bound the RHS by

$$|\sigma|^{-\alpha-\delta} \|\tilde{P}(\sigma) u\|_{s, 1/2-\alpha}.$$

Meanwhile, the LHS is clearly larger than $\|u\|_{s, -3/2-\delta}$, and the result follows. \square

We will also require a slightly different special case of Theorem 11, which we record for later use as a separate lemma.

Lemma 19. *For all $\delta > 0$ and $s > 1 + \delta$,*

$$\tilde{P}(\sigma)^{-1} : H_b^{s, -1/2-\delta} \rightarrow |\sigma|^{-1-2\delta} H_b^{s, -3/2-\delta}$$

Proof. We take $l = -3/2 - \delta$ in Theorem 11, which entails $s > 1/2 + \delta$ and $\nu \in (-1 - \delta, -\delta)$. Taking $\nu = -2\delta$, and estimating $(\rho + |\sigma|)^{-1-2\delta} \leq |\sigma|^{-1-2\delta}$ on the RHS gives the desired estimate. \square

Recall that

$$(34) \quad \tilde{P}(\sigma) = \tilde{P}(0) - \sigma R$$

where

$$(35) \quad R = -2i\rho(\rho\partial_\rho - 1) + \mathcal{A}^{\kappa+1} \text{Diff}_b^1(X) + \sigma\mathcal{A}^\kappa.$$

In the following calculations we make use of the fact that $H_b^{s_1, \gamma_1} \subset H_b^{s_2, \gamma_2}$ for $s_2 \leq s_1$ and $\gamma_2 \leq \gamma_1$. From (35) we see

$$R(\rho^n Y_\ell) = C_n \rho^{n+1} Y_\ell + \mathcal{A}^{\kappa+n+1} + \sigma \mathcal{A}^{\kappa+n},$$

where we crucially note that $C_1 = 0$. Likewise,

$$(36) \quad R(H_b^{s, \gamma}) \subset H_b^{s-1, \gamma+1} + \sigma H_b^{s, \gamma+\kappa} \quad \text{and} \quad R(\mathcal{A}^\kappa) \subset \mathcal{A}^{\kappa+1} + \sigma \mathcal{A}^{2\kappa}.$$

In consequence, the foregoing lemmas imply

$$(37) \quad \begin{aligned} R\tilde{P}(0)^{-1} : H_b^{s, \gamma} &\rightarrow \sum_{j=3}^{\infty} \rho^j Y_{j-2} + H_b^{s+1, \gamma-1-} + \sigma H_b^{s+2, \kappa-\frac{1}{2}-}, \\ \gamma &\in \left(\frac{3}{2}, \kappa + \frac{3}{2} \right], \quad s > 0, \end{aligned}$$

$$(38) \quad \begin{aligned} R\tilde{P}(0)^{-1} : H_b^{s, \gamma} &\rightarrow H_b^{s+1, \gamma-1-} + \sigma H_b^{s+2, \gamma+\kappa-2-}, \\ \gamma &\in \left(\frac{1}{2}, \frac{3}{2} \right), \quad s > 1, \end{aligned}$$

$$(39) \quad \begin{aligned} R\tilde{P}(0)^{-1} : \rho^n Y_\ell &\rightarrow \rho^{n-1} Y_\ell + \sum_{j=3}^{\infty} \rho^j Y_{j-2} + \mathcal{A}^{\kappa+2-} + \sigma \mathcal{A}^{\kappa+2}, \\ n &\geq 4, \ell \geq n-2 \end{aligned}$$

$$(40) \quad R\tilde{P}(0)^{-1} : \rho^3 Y_\ell \rightarrow \sum_{j=3}^{\infty} \rho^j Y_{j-2} + \mathcal{A}^{\kappa+2-} + \sigma \mathcal{A}^{\kappa+2}, \quad \ell \geq 1.$$

Note that we have written sums of $\rho^j Y_{j-2}$ terms with sums going out to infinity for simplicity in bookkeeping, but all but a finite number of these terms are subsumed in the conormal errors that we also carry along. When $\gamma \in (\frac{1}{2}, \frac{3}{2})$ we use Corollary 12 to find $\tilde{P}(0)^{-1} : H_b^{s, \gamma} \rightarrow H_b^{s, \gamma-2-}$ for $s > \frac{3}{2} - \gamma$.

Additionally, Lemma 19 and (35) yield the following estimate when we replace $\tilde{P}(0)$ with the full $\tilde{P}(\sigma)$ (and a factor of σ thrown in for purposes of later iteration):

$$(41) \quad \tilde{P}(\sigma)^{-1}(\sigma R) : H_b^{s, -3/2-\delta} \rightarrow |\sigma|^{-2\delta} H_b^{s-1, -3/2-\delta}, \quad s > 2 + \delta.$$

Since $[\sigma \partial_\sigma, \tilde{P}(\sigma)]$ is almost but not exactly $-\sigma R$, this is not quite the estimate we will need to obtain mapping properties of $(\sigma \partial_\sigma)^J \tilde{P}(\sigma)^{-1}$ below; rather we need a slight variant to take into account iterated commutators of $\sigma \partial_\sigma$ and $\tilde{P}(\sigma)$. To this end, recall that

$$\tilde{P}(\sigma) = \tilde{P}(0) - \sigma R$$

where $R = R_0 + \sigma \mathcal{A}^\kappa$, with R_0 independent of σ . Thus

$$[\sigma \partial_\sigma, \tilde{P}(\sigma)] = -\sigma R + \sigma^2 \mathcal{A}^\kappa,$$

and moreover all iterated commutators of $\sigma \partial_\sigma$ with $\tilde{P}(\sigma)$ are of the form

$$(42) \quad [\sigma \partial_\sigma, -\sigma R + \sigma^2 \mathcal{A}^\kappa] = -\sigma R + \sigma^2 \mathcal{A}^\kappa.$$

We thus remark for later use that changing σR by a multiplication operator in $\sigma^2 \mathcal{A}^\kappa$ does not change its mapping properties, since the mapping properties (36) apply a

fortiori when R is replaced by a multiplication operator in $\sigma\mathcal{A}^\kappa$ (such a term figured as part of R itself in the first place). Thus, more generally,

$$(43) \quad \tilde{P}(\sigma)^{-1}(\sigma R + \sigma^2\mathcal{A}^\kappa) : H_b^{s, -3/2-\delta} \rightarrow |\sigma|^{-2\delta} H_b^{s-1, -3/2-\delta}, \quad s > 2 + \delta.$$

5.3. Conormal estimates. We begin by stating the low-frequency estimates on the twisted resolvent that are essential to our energy decay results. We will employ the notation $\{\kappa\} \in [0, 1)$ for the fractional part of κ and $\lfloor \kappa \rfloor \in \mathbb{Z}$ for the floor function, so that

$$\kappa = \{\kappa\} + \lfloor \kappa \rfloor.$$

Proposition 20. *Let $s > 0$. The twisted resolvent enjoys the following low-frequency asymptotics: If $f \in H_b^{s, \kappa+3/2}$, then*

$$(\sigma\partial_\sigma)^M \tilde{P}(\sigma)^{-1} f \in |\sigma|^{1+\kappa-} L^\infty((-1, 1)_\sigma; H_{loc}^2) + \mathcal{C}^\infty((-1, 1)_\sigma; H_{loc}^2)$$

for

$$M \leq s + \lfloor \kappa \rfloor.$$

Proof. Suppose $\tilde{P}(\sigma)u = f$. We approximate the solution by applying a formal Neumann series argument: using the decomposition (34), we have for all N ,

$$(44) \quad \begin{aligned} (\tilde{P}(0) - \sigma R)^{-1} &= ((\text{Id} - \sigma R \tilde{P}(0)^{-1}) \tilde{P}(0))^{-1} \\ &= \tilde{P}(0)^{-1} (\text{Id} + \cdots + (\sigma \tilde{P}(0)^{-1} R)^N) \\ &\quad + \tilde{P}(0)^{-1} (\text{Id} - \sigma R \tilde{P}(0)^{-1})^{-1} (\sigma R \tilde{P}(0)^{-1})^{N+1} \\ &= \tilde{P}(0)^{-1} (\text{Id} + \cdots + (\sigma R \tilde{P}(0)^{-1})^N) \\ &\quad + (\tilde{P}(0) - \sigma R)^{-1} (\sigma R \tilde{P}(0)^{-1})^{N+1}. \end{aligned}$$

Define $f_n := (R \tilde{P}(0)^{-1})^n f$ so that

$$u = \tilde{P}(0)^{-1} f_0 + \cdots + \sigma^N \tilde{P}(0)^{-1} f_N + \sigma^{N+1} \tilde{P}(\sigma)^{-1} f_{N+1}.$$

By (37),

$$f_1 = R \tilde{P}(0)^{-1} f \in \sum_{j=3}^{\infty} \rho^j Y_{j-2} + H_b^{s+1, \kappa+\frac{1}{2}-} + \sigma H_b^{s+2, \kappa-\frac{1}{2}-}.$$

(Recall that we will use Y_j to denote finite linear combination of spherical harmonics of the given weight, without changing notation for each occurrence.)

Then (38), (39), (40) yield³

$$\begin{aligned} f_2 &= R \tilde{P}(0)^{-1} f_1 \\ &= \sum_3^{\infty} \rho^j (Y_{j-2} + \sigma Y_{j-2} + Y_{j-1}) + H_b^{s+2, \kappa-\frac{1}{2}-} + \sigma H_b^{s+3, \kappa-\frac{3}{2}-} + \sigma^2 H_b^{s+4, \kappa-\frac{1}{2}-}. \end{aligned}$$

We will frequently be faced with terms that have polynomial dependence on σ , and will not be especially interested in the degrees of the resulting polynomials (which

³We let the upper index of the sum of ρ^j terms equal infinity for brevity; in fact of course we could rewrite this as a finite sum, with all terms beyond $j = \lfloor \kappa \rfloor + 2$ being absorbed in the $H_b^{s+2, \kappa-\frac{1}{2}-}$ term, with the caveat that polynomial dependence on σ must then be allowed in that term.

could be bounded in terms of κ in the iteration below if desired). To streamline the resulting notation, we therefore write

$$\mathbb{C}[\sigma]Y_j, \mathbb{C}[\sigma]H_b^{s,l}$$

to indicate respectively polynomials in σ with coefficients in Y_j or $H_b^{s,l}$. In this notation, then, we continue applying $R\tilde{P}(0)^{-1}$ to establish inductively that

$$f_n \in \sum_{j=3}^{\infty} \rho^j \left(\sum_{\ell=j-2}^{j+n-3} \mathbb{C}[\sigma]Y_\ell \right) + H_b^{s+n, \kappa+\frac{3}{2}-n-} + \sigma \mathbb{C}[\sigma]H_b^{s+1+n, \kappa+1/2-n-}$$

for $1 \leq n \leq \lfloor \kappa \rfloor$. Hence, absorbing all terms in the first sum beyond $j = 3$ in the following term,

$$f_{\lfloor \kappa \rfloor} \in \rho^3 \left(\sum_{\ell=1}^{\lfloor \kappa \rfloor} \mathbb{C}[\sigma]Y_\ell \right) + \mathbb{C}[\sigma]H_b^{s+\lfloor \kappa \rfloor, \{\kappa\}+\frac{3}{2}-} + \sigma \mathbb{C}[\sigma]H_b^{s+\lfloor \kappa \rfloor+1, \{\kappa\}+1/2-}.$$

We can continue one more step with the iteration: setting

$$J = \lfloor \kappa \rfloor + 1,$$

we obtain

$$f_J \in \mathbb{C}[\sigma]H_b^{s+\lfloor \kappa \rfloor+1, \{\kappa\}+\frac{1}{2}-} + \sigma \mathbb{C}[\sigma]H_b^{s+\lfloor \kappa \rfloor+2, \{\kappa\}-\frac{1}{2}-} + \sigma^2 \mathbb{C}[\sigma]H_b^{s+3+\lfloor \kappa \rfloor, \{\kappa\}+\frac{1}{2}-}$$

where the terms containing the spherical harmonics Y_ℓ have been absorbed into the first term in f_J . We split this term into pieces

$$f_J = F_J + \sigma G_J$$

$$F_J \in \mathbb{C}[\sigma]H_b^{s+\lfloor \kappa \rfloor+1, \{\kappa\}+\frac{1}{2}-} + \sigma^2 \mathbb{C}[\sigma]H_b^{s+3+\lfloor \kappa \rfloor, \{\kappa\}+\frac{1}{2}-},$$

$$G_J \in \mathbb{C}[\sigma]H_b^{s+\lfloor \kappa \rfloor+2, \{\kappa\}-\frac{1}{2}-}.$$

Then

$$\begin{aligned} \tilde{P}(\sigma)^{-1}f &= \tilde{P}(0)^{-1}(f_0 + \cdots + \sigma^{J-1}f_{J-1}) + \sigma^J \tilde{P}(\sigma)^{-1}(F_J + \sigma G_J) \\ &= \tilde{P}(0)^{-1}(f_0 + \cdots + \sigma^{J-1}f_{J-1}) + \sigma^J \tilde{P}(0)^{-1}F_J \\ (45) \quad &+ \sigma^{J+1} \tilde{P}(\sigma)^{-1}(R\tilde{P}(0)^{-1}F_J) + \sigma^J \tilde{P}(\sigma)^{-1}(\sigma G_J) \\ &= \tilde{P}(0)^{-1}(f_0 + \cdots + \sigma^{J-1}f_{J-1}) + \sigma^J \tilde{P}(0)^{-1}F_J \\ &+ \sigma^{J+1} \tilde{P}(\sigma)^{-1}(R\tilde{P}(0)^{-1}F_J + G_J) \end{aligned}$$

where we have applied (44) with $N = 0$ in the penultimate step. Note that the terms with polynomial dependence in σ will be favorable for obtaining regularity of the resolvent at $\sigma = 0$: it is the final term

$$\tilde{P}(\sigma)^{-1}(R\tilde{P}(0)^{-1}F_J + G_J) \equiv \tilde{P}(\sigma)^{-1}W$$

with

$$W = (R\tilde{P}(0)^{-1}F_J + G_J)$$

that will require finer analysis.

We now analyze the regularity of the terms above. For $n = 1, \dots, J-1 = \lfloor \kappa \rfloor$, we have

$$\tilde{P}(0)^{-1} f_n = \sum_{j=1}^{\infty} \rho^j \sum_{\ell=j-1}^{j+n-1} \mathbb{C}[\sigma] Y_{\ell} + H_b^{s+n+2, \kappa-\frac{1}{2}-n-} + \sigma \mathbb{C}[\sigma] H_b^{s+1+n, \kappa-n-\frac{3}{2}-},$$

using Lemmas 15, 16, and 17. Likewise, by Lemmas 15 and 16,

$$\tilde{P}(0)^{-1} F_J \in \mathbb{C}[\sigma] H_b^{s+\lfloor \kappa \rfloor+3, \{\kappa\}-3/2-},$$

while using (38) gives

$$R\tilde{P}(0)^{-1} F_J \in \mathbb{C}[\sigma] H_b^{s+\lfloor \kappa \rfloor+2, \{\kappa\}-1/2-}.$$

In particular, we have now established

$$W \in \mathbb{C}[\sigma] H_b^{s+\lfloor \kappa \rfloor+2, \{\kappa\}-1/2-}.$$

hence Lemma 18 yields

$$(46) \quad \tilde{P}(\sigma)^{-1} W \in |\sigma|^{-1+\{\kappa\}-} L_{\sigma}^{\infty} H_b^{s+\lfloor \kappa \rfloor+2, -\frac{3}{2}-}.$$

where $L_{\sigma}^{\infty} H_b^{\cdot, \cdot}$ denotes a bounded function of $\sigma \in (-1, 1)$ with values in the given Sobolev space. We now sharpen this to obtain a partial conormality in σ at $\sigma = 0$: in particular, we claim

$$(47) \quad (\sigma \partial_{\sigma})^M \tilde{P}(\sigma)^{-1} W \in |\sigma|^{-1+\{\kappa\}-} L_{\sigma}^{\infty} H_b^{s+\lfloor \kappa \rfloor+2-M, -\frac{3}{2}-}$$

for all integers

$$M < s + \lfloor \kappa \rfloor + 1.$$

To accomplish this, we will need a small result about commutators of $\sigma \partial_{\sigma}$ with $\tilde{P}(\sigma)^{-1}$. In what follows, we use the letter Q to denote an operator of the form $(\text{constant}) \cdot (-\sigma R + \sigma^2 \mathcal{A}^k)$ but with the specific operator allowed to change in each occurrence.

Lemma 21. *For all $M \in \mathbb{N}$, there exist constants $C_{\ell m}$ such that*

$$(\sigma \partial_{\sigma})^M \tilde{P}(\sigma)^{-1} = \sum_{\ell+m \leq M} C_{\ell m} (\tilde{P}(\sigma)^{-1} Q)^{\ell} \tilde{P}(\sigma)^{-1} (\sigma \partial_{\sigma})^m.$$

(Note that some of the $C_{\ell m}$ are in fact zero.)

Proof. By induction on M , using the crucial fact (42) that iterated commutators of $\sigma \partial_{\sigma}$ with $\tilde{P}(\sigma)$ all have the form of Q . \square

Now for any m , Lemma 18 yields

$$\tilde{P}(\sigma)^{-1} (\sigma \partial_{\sigma})^m W \in |\sigma|^{-1+\{\kappa\}-} L_{\sigma}^{\infty} H_b^{s+\lfloor \kappa \rfloor+2, -\frac{3}{2}-}$$

(since $\sigma \partial_{\sigma}$ passes harmlessly through a $\mathbb{C}[\sigma]$ factor). Hence repeated use of (43) yields for $\ell \leq M$,

$$(\tilde{P}(\sigma)^{-1} Q)^{\ell} \tilde{P}(\sigma)^{-1} (\sigma \partial_{\sigma})^m W \in |\sigma|^{-1+\{\kappa\}-} L_{\sigma}^{\infty} H_b^{s+\lfloor \kappa \rfloor+2-\ell, -\frac{3}{2}-};$$

the constraint on M follows is the requirement that the b-regularity index, which is $s + \lfloor \kappa \rfloor + 2 - (M - 1)$ after the application of $(\tilde{P}(\sigma)^{-1} Q)^{M-1}$ still remain greater

than 2 as required for the M 'th and final application of (43). Thus by Lemma 21 we obtain (47).

Since the terms other than $\tilde{P}(\sigma)^{-1}W$ in (45) are smooth (indeed, polynomial) in σ , we finally arrive at the estimate

$$(48) \quad f \in H_b^{s, \kappa+3/2-} \implies (\sigma \partial_\sigma)^M \tilde{P}(\sigma)^{-1} f \in |\sigma|^{1+\kappa-} L^\infty((-1, 1)) + \mathcal{C}^\infty((-1, 1))$$

(implicitly with values in the space $H_b^{s+\lfloor \kappa \rfloor+2-M, -3/2-}$) for

$$(49) \quad M \leq s + \lfloor \kappa \rfloor.$$

Since our estimates are with respect to the spaces $H_b^{s+\lfloor \kappa \rfloor+2-M, -3/2-}$, the constraint (49) ensures that these lie in H_{loc}^2 . \square

The following lemma allows us to estimate b regularity of data with the oscillatory factor $e^{-i\sigma/\rho}$ inserted.

Lemma 22. *Let $s \geq 0$. If $f \in H_b^{s, \gamma+s}$ then $e^{-i\sigma/\rho} f \in H_b^{s, \gamma}$, uniformly for σ in a compact set.*

Proof. We compute

$$\left\| (\rho \partial_\rho)^\alpha \partial_\theta^\beta e^{-i\sigma/\rho} f \right\| = \left\| (\rho \partial_\rho - i\sigma/\rho)^\alpha \partial_\theta^\beta f \right\|.$$

As long as $\alpha + |\beta| \leq m$, then, the differential operator in question is in $\rho^{-m} \text{Diff}_b^m$ (with coefficients uniformly bounded for bounded σ) hence the RHS is bounded by the $H_b^{m, \gamma+m}$ norm. Thus we have obtained the result for integer s . The general case follows by interpolation. \square

We may thus translate our estimates back to the setting of the ordinary (unconjugated) resolvent:

Corollary 23. *Let $s > 0$. The resolvent enjoys the following low-frequency asymptotics: If $f \in H_b^{s, \kappa+3/2+s}$ then*

$$(\sigma \partial_\sigma)^M P_\sigma^{-1} f \in |\sigma|^{1+\kappa-} L^\infty((-1, 1)_\sigma; H_{loc}^2) + \mathcal{C}^\infty((-1, 1)_\sigma; H_{loc}^2)$$

for

$$M \leq s + \lfloor \kappa \rfloor.$$

Proof. Note that

$$P_\sigma^{-1} f = e^{i\sigma/\rho} \tilde{P}(\sigma)^{-1} e^{-i\sigma/\rho} f.$$

The leading factor $e^{i\sigma/\rho}$ is smooth in $\sigma \in (-1, 1)$ uniformly on compact sets in X° , hence it suffices to verify that $e^{-i\sigma/\rho} f$ satisfies the hypotheses of Proposition 20. This in turn follows from Lemma 22. \square

6. Proof of Main Theorem

Recall by Proposition 5 that if u solves the initial value problem, then

$$(2\pi)^{1/2}\check{u}(\sigma, \bullet) = P_\sigma^{-1}(-i\sigma u_0 + P_1 u_0 - u_1) \equiv P_\sigma^{-1}(\sigma f_0 + g_0)$$

with

$$(50) \quad f_0 = -iu_0, \quad g_0 = P_1 u_0 - u_1.$$

We take the Fourier transform to recover u and split the solution into low and high frequency parts (denoted u_L and u_H , respectively). Let $\chi_{<1}(|\sigma|)$ be a cutoff function equal to 1 on $(-1/2, 1/2)$ and supported in $(-1, 1)$ and take $\chi_{>1}(|\sigma|) = 1 - \chi_{<1}(|\sigma|)$. We write $u = u_L + u_H$ where

$$(51) \quad u_L = \int_{-\infty}^{\infty} \chi_{<1}(\sigma) e^{-i\sigma t} P_\sigma^{-1}(\sigma f_0 + g_0) d\sigma$$

and

$$(52) \quad u_H = \int_{-\infty}^{\infty} \chi_{>1}(\sigma) e^{-i\sigma t} P_\sigma^{-1}(\sigma f_0 + g_0) d\sigma.$$

It now suffices to treat the asymptotic behavior of the high and low-frequency contributions separately. We begin with u_L .

By Lemma 26 in Appendix B and Sobolev embedding, it will suffice in estimating u_L to show that

$$(53) \quad \check{u}_L(\sigma, x) \in |\sigma|^{\kappa+1-} I^M L_{\text{loc}}^\infty + \mathcal{C}^\infty$$

for some $M \geq \kappa + 1$ (where all spaces in σ are valued in H_{loc}^2).

To this end, we begin by noting (with a view to potential future applications) the sharp hypotheses that are necessary to obtain the estimate (53) for the low-frequency part of the solution: what we will in fact use is

$$(54) \quad \begin{aligned} u_0 &\in H_b^{3, \kappa+7/2}, \\ u_1 &\in H_b^{2, \kappa+7/2}. \end{aligned}$$

Now apply Corollary 23 with $s = 2$ to

$$g_0 = P_1 u_0 - u_1 \in H_b^{2, \kappa+7/2}$$

to obtain

$$(\sigma \partial_\sigma)^M P_\sigma^{-1} g_0 \in |\sigma|^{1+\kappa-} L^\infty((-1, 1)_\sigma; H_{\text{loc}}^2) + \mathcal{C}^\infty((-1, 1)_\sigma; H_{\text{loc}}^2), \quad M \leq 2 + \lfloor \kappa \rfloor.$$

Apply Corollary 23 with κ replaced by $\kappa' = \kappa - 1$ (hence the hypotheses on the perturbation are a fortiori satisfied for κ') and with $s = 3$ to

$$\sigma f_0 = -i\sigma u_0 \in \sigma H_b^{3, \kappa'+9/2}$$

to obtain for $M \leq 3 + \lfloor \kappa' \rfloor$

$$(\sigma \partial_\sigma)^M P_\sigma^{-1} \sigma f_0 \in |\sigma|^{2+\kappa'-} L^\infty((-1, 1)_\sigma; H_{\text{loc}}^2) + \mathcal{C}^\infty((-1, 1)_\sigma; H_{\text{loc}}^2),$$

i.e. for $M \leq 2 + \lfloor \kappa \rfloor$

$$(\sigma \partial_\sigma)^M P_\sigma^{-1} \sigma f_0 \in |\sigma|^{1+\kappa-} L^\infty((-1, 1)_\sigma; H_{\text{loc}}^2) + \mathcal{C}^\infty((-1, 1)_\sigma; H_{\text{loc}}^2).$$

By Lemma 26, as noted above (and since $2 + \lfloor \kappa \rfloor > 1 + \kappa$), this concludes the proof that u_L has the desired decay, and we now turn our attention to u_H , the high-frequency component of the solution.

We will decompose the expression $P_\sigma^{-1}(\sigma f_0 + g_0)$ in (52) via an iterative argument. We approximate $P_\sigma^{-1}(\sigma f_0 + g_0) \approx \sigma^{-1}f_0$ and let v_1 denote the error. Recall $f_0 \in H_b^{s+1, \kappa + \frac{7}{2}}$ and $g_0 \in H_b^{s, \kappa + \frac{7}{2}}$. Direct calculation shows

$$P_\sigma v_1 = (g_0 - iP^1 f_0) + \sigma^{-1}(\Delta + P^2)(-f_0) =: f_1 + \sigma^{-1}g_1$$

where $f_1 \in H_b^{s, \kappa + \frac{7}{2}}$ and $g_1 \in H_b^{s-1, \kappa + \frac{11}{2}}$ (since $P^1 : H_b^{m, \ell} \rightarrow H_b^{m-1, \ell + \kappa + 1}$ and $(\Delta + P^2) : H_b^{m, \ell} \rightarrow H_b^{m-2, \ell + 2}$). Now we have

$$P_\sigma^{-1}(\sigma f_0 + g_0) = \sigma^{-1}f_0 + P_\sigma^{-1}(f_1 + \sigma^{-1}g_1).$$

Next we iterate the process and approximate $P_\sigma^{-1}(f_1 + \sigma^{-1}g_1) \approx \sigma^{-2}f_1$:

$$\begin{aligned} P_\sigma^{-1}(f_1 + \sigma^{-1}g_1) &= \sigma^{-2}f_1 + P_\sigma^{-1}(\sigma^{-1}(g_1 - iP^1 f_1) + \sigma^{-2}(\Delta + P^2)(-f_1)) \\ &=: \sigma^{-2}f_1 + P_\sigma^{-1}(\sigma^{-1}f_2 + \sigma^{-2}g_2) \end{aligned}$$

where $f_2 \in H_b^{s-1, \kappa + \frac{11}{2}}$ and $g_2 \in H_b^{s-2, \kappa + \frac{11}{2}}$. Thus we have

$$(55) \quad P_\sigma^{-1}(\sigma f_0 + g_0) = \underbrace{\sigma^{-1}f_0 + \sigma^{-2}f_1}_{=: \check{u}_a} + \underbrace{P_\sigma^{-1}(\sigma^{-1}f_2 + \sigma^{-2}g_2)}_{=: \check{u}_b}$$

We plug (55) into our expression for u_H (see equation (52)) and bound each term separately. Note $|f_0| + |f_1| \lesssim \langle r \rangle^{-5-\kappa}$ so we calculate for any $M \geq 1$:

$$\begin{aligned} &\left| \int_{\sigma \in \mathbb{R}} \chi_{>1}(|\sigma|) \check{u}_a(\sigma) e^{-it\sigma} d\sigma \right| \\ &\lesssim \langle r \rangle^{-\kappa-5} \langle t \rangle^{-M} \left| \int \partial_\sigma^M [\chi_{>1}(|\sigma|)(\sigma^{-1} + \sigma^{-2})] e^{-i\sigma t} d\sigma \right| \\ &\lesssim \langle t \rangle^{-M} \langle r \rangle^{-\kappa-5}. \end{aligned}$$

By Proposition 10, we have $|(\sigma \partial_\sigma)^M(\check{u}_b(\sigma) e^{-i\sigma r})| \lesssim |\sigma|^{M-2} \langle r \rangle^{-1+M(1-\varepsilon)}$ for $s-2 \geq (\varpi+1)(2M+5)$ and $M < \kappa+5$ with $\varepsilon \in (0, 1]$. We use this to calculate

$$\begin{aligned} &\left| \int_{\sigma \in \mathbb{R}} \chi_{>1}(\sigma) \check{u}_b(\sigma) e^{-it\sigma} d\sigma \right| \\ &\approx \langle t-r \rangle^{-M} \left| \int \sigma^{-M} \left[\sum_{\ell=0}^M (\sigma \partial_\sigma)^\ell (\chi_{>1}(|\sigma|) \check{u}_b e^{-i\sigma r}) \right] e^{i(r-t)\sigma} d\sigma \right| \\ &\lesssim \langle r \rangle^{-1+M(1-\varepsilon)} \langle t-r \rangle^{-M}. \end{aligned}$$

Combining the above results, we find

$$|u_H(t, x)| \lesssim \langle t \rangle^{-M} \langle r \rangle^{-\kappa-1} + \langle r \rangle^{-1+M(1-\varepsilon)} \langle t-r \rangle^{-M}.$$

The main theorem (and indeed a finer estimate near the light cone for this part of the solution) then follows in the high frequency case when we take $M = \kappa + 2 - \varepsilon$ (and thus the b-regularity requirement is $s > (\varpi+1)(2\kappa+9)+2$). \square

Appendix A. Mellin transforms and b-Sobolev spaces

In this appendix, we describe the Mellin transform characterization of the b-Sobolev spaces defined in Section 2 above.

If u is a function of ρ, θ ($\rho \in [0, \infty)$) set

$$\mathcal{M}u(\xi, \theta) = \int_0^\infty \rho^{-i\xi} u(\rho, \theta) \frac{d\rho}{\rho}.$$

Let $\mathcal{H}(\Omega)$ denote the space of holomorphic functions on the domain $\Omega \subset \mathbb{C}$ with values in $L^2(S^2)$.

Let L_{bb}^2 denote the space of u that are L^2 with respect to the “b”-density $d\rho/\rho$. Let $L_{bb,A}^2$ denote the subspace of functions supported in $\rho \in [0, A]$.

Lemma 24. *The range of \mathcal{M} on $L_{bb,A}^2$ is*

$$\{w(\xi, \theta) : w \text{ holomorphic in } \operatorname{Im} \xi > 0, \sup_{\mu > 0} A^{2\mu} \int_{\operatorname{Im} \xi = \mu} \|w(\xi, \theta)\|_{L^2(S^2)}^2 d\xi < \infty\}.$$

Proof. Set $x = \log \rho$. Then

$$\mathcal{M}u(\xi) = \int_{-\infty}^\infty e^{-i\xi x} u(e^x, \theta) dx = (2\pi)^{1/2} \mathcal{F}(u \circ \exp)(\xi),$$

with $u \circ \exp \in L^2(\mathbb{R}_x \times S_\theta^2)$ and supported in $x \in (-\infty, \log A]$. Let us take $A = 1$ for now, hence the support of $u \circ \exp$ is $(-\infty, 0]$. The Paley–Wiener theorem tells us that the Fourier transforms of such functions are precisely the space

$$\{w(\xi, \theta) : w \in \mathcal{H}(\operatorname{Im} \xi > 0), \sup_{\mu > 0} \int_{\operatorname{Im} \xi = \mu} \|w(\xi, \theta)\|_{L^2(S^2)}^2 d\xi < \infty\}.$$

This establishes the result for $A = 1$. More generally, $u \in L_{bb,A}^2$ iff $u(A^{-1}\rho, \theta) \in L_{bb,1}^2$, so the result follows from the fact that $\mathcal{M}[u(A^{-1}\rho, \theta)](\xi) = A^{i\xi} \mathcal{M}u$. \square

Now we observe that adding a weight simply shifts the domain of holomorphy, since $\mathcal{M}\rho^\alpha u(\xi) = \mathcal{M}u(\xi + i\alpha)$ hence

$$\begin{aligned} & \mathcal{M}\rho^\alpha L_{bb,A}^2 \\ & \rightarrow \{w(\xi, \theta) : w \in \mathcal{H}(\operatorname{Im} \xi > -\alpha), \sup_{\mu > -\alpha} A^{2\mu} \int_{\operatorname{Im} \xi = \mu} \|w(\xi, \theta)\|_{L^2(S^2)}^2 d\xi < \infty\}. \end{aligned}$$

Finally, we aim to keep track of b -regularity. Recall that H_b^m denotes the (un-weighted) b -Sobolev space of order m , measured with respect to the metric volume form, which is proportional to $d\rho/\rho^4 d\theta$; note in particular that this entails the numerology

$$H_b^0 = \rho^{3/2} L_{bb}^2.$$

More generally, recall that $H_b^{m,\ell} = \rho^\ell H_b^{m,0}$. We write $H_{b,A}^\bullet$ as before to denote the functions supported in $\rho \leq A$.

Proposition 25. *For $m \geq 0$ the Mellin transform is an isomorphism*

$$(56) \quad \mathcal{M} : H_{b,A}^{m,\ell} \rightarrow \{w(\xi, \theta) : w \in \mathcal{H}(\operatorname{Im} \xi > -\ell - 3/2), \sup_{\mu > -\ell - 3/2} A^{2\mu} \int_{\operatorname{Im} \xi = \mu} \|(\xi^2 + \Delta_\theta)^{m/2} w(\xi, \theta)\|_{L^2(S^2)}^2 d\xi < \infty\}.$$

Proof. For m an even integer, the result follows by the b-elliptic regularity of the differential operator $((rD_r)^2 + \Delta_\theta)^{m/2}$. For general m , it follows by interpolation and duality. \square

Appendix B. Fourier transforms of finite-regularity conormal distributions

We consider distributions in $|\sigma|^\alpha I^m L_c^\infty(\mathbb{R}; 0)$; this is defined as the space of distributions on \mathbb{R} lying in $|\sigma|^\alpha L_c^\infty$ (with c denoting compact support) and enjoying m -fold iterated regularity under vector fields tangent to the origin, which is to say, under powers up to the m th of $\sigma \partial_\sigma$. We will only deal with $m \in \mathbb{N}$ in order to keep the discussion simple. Thus $u \in |\sigma|^\alpha I^m L_c^\infty(\mathbb{R}; 0)$ if u is compactly supported and enjoys the estimate

$$|\partial_\sigma^j u| \leq C_j |\sigma|^{\alpha-j}, \quad j = 0, \dots, m.$$

Lemma 26. *Let $u \in |\sigma|^\alpha I^m L_c^\infty(\mathbb{R}; 0)$, and assume $\alpha > 0$ and $m \geq \alpha + 1$. Then*

$$\mathcal{F}(u) = O(\langle t \rangle^{-1-\alpha}).$$

Proof. (cf. Lemma 3.6 of [19].) We write

$$\mathcal{F}u = (2\pi)^{-1/2} \int u(\sigma) e^{i\sigma t} dt = w_0 + w_1$$

where

$$w_0 = (2\pi)^{-1/2} \int_{|\sigma| < t^{-1}} u(\sigma) e^{i\sigma t} dt, \quad w_1 = \mathcal{F}u - w_0.$$

Then

$$|w_0| \lesssim |t|^{-1-\alpha}$$

since $u \in |\sigma|^\alpha L^\infty$. On the other hand, integration by parts using the operator $(t^{-1}D_\sigma)^m$ in the integral expression for w_1 yields a bulk term bounded by

$$t^{-m} \int_{t^{-1} < |\sigma| < C} \sigma^{\alpha-m} d\sigma$$

(where C depends on the support of u), as well as boundary terms bounded by

$$t^{-m} \sigma^{-\alpha-m'} \Big|_{\sigma=t^{-1}}, \quad m' = 0, \dots, m-1.$$

Since $m \geq \alpha + 1$, these terms are all bounded by multiples of $|t|^{-1-\alpha}$ for $|t|$ large. \square

Acknowledgements

The authors are grateful to Peter Hintz for numerous helpful conversations, as well as to Jason Metcalfe, Mihai Tohaneanu, and Shi-Zhuo Looi. Two anonymous referees also supplied helpful comments and corrections.

The second author gratefully acknowledges support from Simons Foundation grant 631302. The first author gratefully acknowledges the support of NSF Postdoctoral Fellowship DMS-2002132.

References

- [1] S. Alinhac, *On the Morawetz–Keel–Smith–Sogge Inequality for the Wave Equation on a Curved Background*, Publications of The Research Institute for Mathematical Sciences **42** (2006) 705–720.
- [2] L. Andersson and P. Blue, *Hidden symmetries and decay for the wave equation on the Kerr spacetime*, Annals of Mathematics **182** (2015), no. 3, 787–853.
- [3] Y. Angelopoulos, S. Aretakis, and D. Gajic, *Late-time asymptotics for the wave equation on spherically symmetric, stationary spacetimes*, Adv. Math. **323** (2018) 529–621.
- [4] ———, *Late-time tails and mode coupling of linear waves on Kerr spacetimes*, arXiv preprint arXiv:2102.11884 (2021)
- [5] ———, *Price's law and precise late-time asymptotics for subextremal Reissner–Nordström black holes*, arXiv preprint arXiv:2102.11888 (2021)
- [6] P. Blue and A. Soffer, *Semilinear wave equations on the Schwarzschild manifold. I. Local decay estimates*, Adv. Differential Equations **8** (2003), no. 5, 595–614.
- [7] J.-F. Bony and D. Häfner, *Local energy decay for several evolution equations on asymptotically Euclidean manifolds*, in Annales scientifiques de l'École Normale Supérieure, Vol. 45, 311–335 (2012).
- [8] ———, *Improved local energy decay for the wave equation on asymptotically Euclidean odd dimensional manifolds in the short range case*, Journal of the Institute of Mathematics of Jussieu **12** (2013), no. 3, 635–650.
- [9] R. Booth, H. Christianson, J. Metcalfe, and J. Perry, *Localized energy for wave equations with degenerate trapping*, arXiv preprint arXiv:1712.05853 (2017)
- [10] N. Burq, *Décroissance de l'énergie locale de l'équation des ondes pour le problème extérieur et absence de résonance au voisinage du réel*, Acta Math. **180** (1998), no. 1, 1–29.
- [11] M. Dafermos and I. Rodnianski, *The red-shift effect and radiation decay on black hole spacetimes*, Communications on Pure and Applied Mathematics: A Journal Issued by the Courant Institute of Mathematical Sciences **62** (2009), no. 7, 859–919.
- [12] ———, *A new physical-space approach to decay for the wave equation with applications to black hole spacetimes*, in XVIth International Congress on Mathematical Physics, 421–432, World Sci. Publ., Hackensack, NJ (2010).
- [13] ———, *A proof of the uniform boundedness of solutions to the wave equation on slowly rotating Kerr backgrounds*, Inventiones mathematicae **185** (2011), no. 3, 467–559.
- [14] ———, *Lectures on black holes and linear waves*, Clay Math. Proc **17** (2013) 97–205.
- [15] M. Dafermos, I. Rodnianski, and Y. Shlapentokh-Rothman, *Decay for solutions of the wave equation on Kerr exterior spacetimes III: The full subextremal case $|a| < M$* , Annals of Mathematics **183** (2016), no. 3, 787–913.
- [16] R. Donninger, W. Schlag, and A. Soffer, *A proof of Price's law on Schwarzschild black hole manifolds for all angular momenta*, Adv. Math. **226** (2011), no. 1, 484–540.
- [17] F. Finster, N. Kamran, J. Smoller, and S.-T. Yau, *Decay of solutions of the wave equation in the Kerr geometry*, Communications in Mathematical Physics **264** (2006), no. 2, 465–503.
- [18] C. Guillarmou, A. Hassell, and A. Sikora, *Resolvent at low energy III: the spectral measure*, Transactions of the American Mathematical Society **365** (2013), no. 11, 6103–6148.
- [19] P. Hintz, *A sharp version of Price's Law for wave decay on asymptotically flat spacetimes*, arXiv preprint arXiv:2004.01664 (2020)
- [20] A. Jensen and T. Kato, *Spectral properties of Schrödinger operators and time-decay of the wave functions*, Duke mathematical journal **46** (1979), no. 3, 583–611.
- [21] S.-Z. Looi, *Pointwise decay for the wave equation on nonstationary spacetimes*.
- [22] J. Luk, *A vector field method approach to improved decay for solutions to the wave equation on a slowly rotating Kerr black hole*, Anal. PDE **5** (2012), no. 3, 553–625.
- [23] J. Marzuola, J. Metcalfe, D. Tataru, and M. Tohaneanu, *Strichartz estimates on Schwarzschild black hole backgrounds*, Communications in Mathematical Physics **293** (2010), no. 1, 37.
- [24] R. Melrose, *The Atiyah–Patodi–Singer Index Theorem*, A K Peters/CRC Press (1993).
- [25] R. Melrose and G. Mendoza, *Elliptic operators of totally characteristic type*, Mathematical Sciences Research Institute (1983).
- [26] R. B. Melrose, *Transformation of boundary problems*, Acta Math. **147** (1981), no. 3–4, 149–236.

- [27] ———, *Spectral and scattering theory for the Laplacian on asymptotically Euclidian spaces*, in Spectral and scattering theory (Sanda, 1992), Vol. 161 of *Lecture Notes in Pure and Appl. Math.*, 85–130, Dekker, New York (1994).
- [28] J. Metcalfe and C. D. Sogge, *Long-time existence of quasilinear wave equations exterior to star-shaped obstacles via energy methods*, SIAM journal on mathematical analysis **38** (2006), no. 1, 188–209.
- [29] J. Metcalfe, J. Sterbenz, and D. Tataru, *Local energy decay for scalar fields on time dependent non-trapping backgrounds*, American Journal of Mathematics **142** (2020), no. 3, 821–883.
- [30] J. Metcalfe and D. Tataru, *Decay estimates for variable coefficient wave equations in exterior domains*, in Advances in phase space analysis of partial differential equations, 201–216, Springer (2009).
- [31] ———, *Global parametrices and dispersive estimates for variable coefficient wave equations*, Mathematische Annalen **353** (2012), no. 4, 1183–1237.
- [32] J. Metcalfe, D. Tataru, and M. Tohaneanu, *Price’s law on nonstationary space-times*, Advances in Mathematics **230** (2012), no. 3, 995–1028.
- [33] C. S. Morawetz, *The decay of solutions of the exterior initial-boundary value problem for the wave equation*, Communications on Pure and Applied Mathematics **14** (1961), no. 3, 561–568.
- [34] K. Morgan, *The effect of metric behavior at spatial infinity on pointwise wave decay in the asymptotically flat stationary setting*, arXiv preprint arXiv:2006.11324 (2020)
- [35] R. H. Price, *Nonspherical perturbations of relativistic gravitational collapse. I. Scalar and gravitational perturbations*, Physical Review D **5** (1972), no. 10, 2419.
- [36] J. V. Ralston, *Solutions of the wave equation with localized energy*, Communications on Pure and Applied Mathematics **22** (1969), no. 6, 807–823.
- [37] J. Sbierski, *Characterisation of the energy of Gaussian beams on Lorentzian manifolds: with applications to black hole spacetimes*, Analysis & PDE **8** (2015), no. 6, 1379–1420.
- [38] C. D. Sogge and C. Wang, *Concerning the wave equation on asymptotically Euclidean manifolds*, Journal d’Analyse Mathématique **112** (2010), no. 1, 1–32.
- [39] D. Tataru, *Local decay of waves on asymptotically flat stationary space-times*, American Journal of Mathematics **135** (2013), no. 2, 361–401.
- [40] A. Vasy, *Resolvent near zero energy on Riemannian scattering (asymptotically conic) spaces, a Lagrangian approach*, Communications in Partial Differential Equations (2020) 1–47.
- [41] J. Wunsch, *Resolvent estimates with mild trapping*, Journées équations aux dérivées partielles (2012) 1–15.
- [42] J. Wunsch and M. Zworski, *Resolvent estimates for normally hyperbolic trapped sets*, Ann. Henri Poincaré **12** (2011), no. 7, 1349–1385.

DEPARTMENT OF MATHEMATICS, NORTHWESTERN UNIVERSITY, 2033 SHERIDAN ROAD, EVANSTON, IL 60208

E-mail address: `katrina.morgan@northwestern.edu`, `jwunsch@math.northwestern.edu`