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Phoresis kernel theory for passive and active spheres with nonuniform phoretic mobility†

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By introducing geometry-based phoresis kernels, we establish a direct connection between the translational and rotational velocities of a phoretic sphere and the distributions of the driving fields or fluxes. The kernels quantify the local contribution of the field or flux to the particle dynamics. The field kernels for both passive and active particles share the same functional form, depending on the position-dependent surface phoretic mobility. For uniform phoretic mobility, the translational field kernel is proportional to the surface normal vector, while the rotational field kernel is zero; thus, a phoretic sphere with uniform phoretic mobility does not rotate. As case studies, we discuss examples of a self-phoretic axisymmetric particle influenced by a globally-driven field gradient, a general scenario for axisymmetric self-phoretic particle and two of its special cases, and a non-axisymmetric active particle.

1. Introduction

Phoresis encompasses a group of phenomena where particles move under the influence of a gradient in a physical property of their medium or a field within the fluid.^{1–4} Such phenomena include diffusiophoresis,^{5–9} prompted by the concentration gradient of a solute; electrophoresis,^{10–16} resulting from the spatial variation of electric potential; and thermophoresis,^{17–21} driven by temperature gradients. These phoretic movements play a vital role in various scientific and industrial processes, such as particle separation in mixtures, purification, pollution control, drug delivery system design, surface coatings, and the rapidly growing fields of microswimmers, active colloids and microrobotics.^{22–29}

Designing particles for specific purposes that undergo phoresis phenomena under a driving field or controlling their behavior requires a deep understanding of the particle–field interactions and the phoresis phenomena. In the linear regime, the majority of the interactions between particle and the surrounding field gradient occurs in the interaction layer, which is a region in the vicinity of the particle surface. In the regime where the length scale of the interaction layer is much smaller than the particle size, the effect of the phenomena inside the interaction layer is usually lumped into a quantity known as slip velocity. This is the fluid velocity at the outer

surface of the interaction layer. Models then use the slip velocity as a boundary condition to solve for the fluid flow to study the velocity of a single particle system^{1,10,11,13,30–34} or particle–particle^{35–40} and particle–boundary⁴¹ interactions.

In this paper, our focus is on the dynamics of an individual particle in an unbounded domain. The traditional approach to determining the velocity of a phoretic particle with a given surface composition involves calculating the driving field and obtaining the slip velocity over the particle surface, from which we calculate the translational and rotational velocities of the particle using the reciprocal theorem.^{4,42} To solve the inverse problem of finding a surface composition that leads to the desired dynamics, we typically engage in a series of trial-and-error designs, supported by intuition and experience, to determine the surface composition and corresponding slip velocities. From these, we derive the particle velocities, potentially converging to a surface composition that achieves the desired dynamics. With the calculation of slip velocity as an intermediate step, this trial-and-error process seems inevitable.

Since we typically have more direct control, albeit still not very precise, over flux than over slip velocity, to overcome this challenge or at least reduce the complications, we will bypass the need for calculating slip velocity. Our goal is to establish a direct connection between the field distribution or its flux with particle velocity through integral kernels. These kernels encompass the effects of geometry as well as the surface phoretic mobility, which is a measure of the interaction of the particle with the field, irrespective of the field's distribution. Thus, phoresis kernels provide deeper insight into how geometry and surface composition can influence phoresis phenomena. This knowledge can then be used to design phoretic particles for specific purposes or to optimize and fine-tune their properties.

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Particles undergoing phoresis are typically categorized into two classes: passive and active. Passive particles alter the distribution of the field around them but do not contribute to the creation of the field. For instance, a silica sphere in the presence of an external electric field undergoes electrophoresis; it influences the electric potential around it but does not create the field gradient that drives the phoresis phenomenon. Conversely, active particles interact with their environment in a way that leads to the creation of a local gradient of the driving field. An example is a platinum–gold bimetallic micro-rod⁴³ in a hydrogen peroxide solution, which interacts to generate a self-produced gradient of electric potential, resulting in self-electrophoresis.^{11–13}

Since active particles move under their self-generated field gradient, they are also referred to as self-phoretic or auto-phoretic particles. However, this does not preclude them from being driven by external fields, similar to passive particles. For example, a platinum–gold bimetallic micro-rod, while performing self-electrophoresis due to the local gradient of hydronium ions, may be subjected to a global concentration gradient of non-ionic molecules and undergo diffusiophoresis. Therefore, rather than focusing on the active or passive properties of the particle, we concentrate on the length scale of the field gradient.

We describe a phoresis process as globally-driven when the length scale of the field variation is of system size, much larger than the particle size. Conversely, we refer to a phoresis process as locally-driven when the length scale of the field gradient is on the order of the particle size. Within this framework, an active particle can be simultaneously locally-driven by its self-generated field gradient and globally-driven by another externally imposed field gradient. In general, a particle's phoretic motion may result from the superposition of multiple phoresis processes driven by distinct fields.

In both locally-driven and globally-driven phoresis processes, even though the particle velocity is a vectorial quantity for the entire particle, the driving field on the particle surface is distributed and varies locally across the particle surface. To bridge the dynamics and the field distribution within the phoresis kernel theory, we quantify the local contribution of the field to the particle velocity using integral kernels. These kernels depend on the geometry and surface properties of the particle and remain independent of the field's distribution. Thus, for both active and passive particles, the functional forms of the field kernels should be identical. However, the flux of the driving field in locally-driven processes is distributed over the particle's surface, whereas, for globally-driven processes, its far-field value is known. As a result, their respective flux kernels are different, and the surface integration of the normal component of the flux, weighted by these kernels, holds distinct interpretations for locally-driven and globally-driven phoresis phenomena.

In what follows, we begin by formulating the phoresis of a sphere in the linear regime of a thin interaction layer. We base our approach on Fair and Anderson's expressions for the rotational and translational velocities of a phoretic sphere, which rely on slip velocity.¹⁰ However, we replace the

integration involving slip velocity with the integration of the field, weighted by position-dependent field kernels, over the particle surface. Using expressions for fields over the particle surface for locally-driven and globally-driven phoresis processes, we derive the corresponding flux kernels. We then apply this formalism to study examples including the self-phoresis of axisymmetric particles with uniform phoretic mobility influenced by a globally-driven field gradient, the translational velocity of an axisymmetric self-phoretic swimmer with a distribution of phoretic mobility, and a non-axisymmetric active particle. The details of the derivations and calculations are provided in the ESI.†

2. Formulation of the problem

A. Driving field for phoresis

For a sphere situated within a fluid to undergo phoretic motion, the particle is subjected to at least one driving field Ψ corresponding to a phoresis process. In the limit of a thin interaction layer, we approximate the outer surface of the interaction layer, denoted by S^+ , as a sphere of radius a . For brevity, we refer to S^+ as the particle surface henceforth. The flux of the field is proportional to the gradient of the field,

$$\Gamma = -\mathcal{D}\nabla\Psi \quad (1)$$

where the proportionality constant $\mathcal{D}$ is a diffusion coefficient associated with the phoresis process.

The gradient of the field, and thus its flux, drives the phoresis process. The flux can be imposed globally or generated locally. Typically, a passive particle is driven by a global gradient, while an active particle can self-generate a local gradient even as it undergoes another phoresis process imposed by a global gradient. In this paper, we focus on phoresis processes which obey the following surface and far-field boundary conditions,

$$\begin{cases} \hat{n} \cdot \nabla\Psi|_{S^+} = -\mathcal{D}^{-1}\hat{n} \cdot \Gamma^{S^+}, \\ \nabla\Psi|_{\infty} = -\mathcal{D}^{-1}\Gamma^{\infty}, \end{cases} \quad (2)$$

where $\hat{n}$ is the unit normal to the particle surface. For the locally-generated flux Γ^{S^+} , the distribution of the normal component of the flux over S^+ may vary. Conversely, the far-field flux Γ^{∞} , which pertains to the globally-driven aspect of the phoresis process, is typically uniform. The boundary condition (2) includes both local and global gradients of the driving field. Common special cases include self-phoresis ($\Gamma^{S^+} \neq 0$) in the absence of global flux ($\Gamma^{\infty} = 0$),^{3,4} and globally-driven phoresis ($\Gamma^{\infty} \neq 0$) of a passive particle in the absence of surface flux ($\Gamma^{S^+} = 0$).¹⁰

B. Translational and rotational velocities

The interaction of the particle with the field Ψ near the particle surface, and the resulting flow within the interaction layer, can be lumped into a slip velocity,

$$v_{\text{slip}} = \mu\nabla_s\Psi, \quad (3)$$



tangent to the outer surface S^+ of the interaction layer. The phoretic mobility μ depends on the local thermodynamic conditions of the interface. Using the identity matrix $\mathcal{I}$, the expression $\nabla_S \Psi = (\mathcal{I} - \hat{n}\hat{n}) \cdot \nabla \Psi$ is tangent to S^+ , providing the surface gradient of the field over the particle surface.

In scenarios where more than one phoresis process occurs, the slip velocity over the particle surface is the sum of these phoresis-specific slip velocities,

$$\mathbf{v}_{\text{slip}} = \sum_{\alpha} \mathbf{v}_{\text{slip}}^{(\alpha)} = \sum_{\alpha} \mu^{(\alpha)} \nabla_S \Psi^{(\alpha)} \quad (4)$$

where α sums over different phoresis processes involved and the validity of this summation stems from the linearity of the governing equations.

The flow field around a particle moving with translational velocity $\mathbf{U}$ and rotational velocity $\mathbf{\Omega}$ in a Newtonian fluid of viscosity η follows the Stokes and continuity equations,

$$\eta \nabla^2 \mathbf{u} = \nabla P, \quad (5a)$$

$$\nabla \cdot \mathbf{u} = 0, \quad (5b)$$

respectively, subject to surface ($\mathbf{x}_S \in S^+$) and far-field boundary conditions,

$$\mathbf{u}(\mathbf{x}_S) = \mathbf{U} + \mathbf{\Omega} \times \mathbf{x}_S + \mathbf{v}_{\text{slip}}, \quad (6a)$$

$$\mathbf{u}(\mathbf{x} \rightarrow \infty) = 0. \quad (6b)$$

The translational and rotational velocities of the phoretic sphere are given by Fair and Anderson,¹⁰

$$\mathbf{U} = \frac{-1}{4\pi} \int_0^\pi \int_0^{2\pi} \mathbf{v}_{\text{slip}} \sin \theta d\theta d\phi, \quad (7a)$$

$$\mathbf{\Omega} = \frac{-3}{8\pi a} \int_0^\pi \int_0^{2\pi} \hat{n} \times \mathbf{v}_{\text{slip}} \sin \theta d\theta d\phi. \quad (7b)$$

Similar expressions were presented by Stone and Samuel⁴⁴ for the propulsion of microorganisms by surface distortions. The eqn (7) are surface integrations of integrands related to the slip velocity and, thus, to the surface gradient of the fields. In the next section, we discuss how we can eliminate the slip velocity from these equations, thereby relating the velocities $\mathbf{U}$ and $\mathbf{\Omega}$ directly to the driving field Ψ and fluxes Γ^{S^+} and Γ^∞ using phoresis kernels.

C. Phoresis kernel theory

Exploiting the linearity of the governing equations, the standard practice for obtaining the translational and rotational velocities for a phoresis process can be summarized in three steps:

$$\hat{n} \cdot \Gamma^{S^+} \xrightarrow[\text{eqn (17)}]{\text{step I}} \Psi \xrightarrow[\text{eqn (3)}]{\text{step II}} \mathbf{v}_{\text{slip}} \xrightarrow[\text{eqn (7)}]{\text{step III}} \begin{matrix} \mathbf{U} \\ \mathbf{\Omega} \end{matrix}. \quad (8)$$

First, we obtain the field Ψ from fluxes $\hat{n} \cdot \Gamma^{S^+}$ and Γ^∞ using eqn (17); second, calculate the surface gradient of the field, and obtain the slip velocity (3); and third, use the slip velocity in integrals (7) to yield the velocities.

In this paper, our goal is to provide a direct connection between the field Ψ and the fluxes $\hat{n} \cdot \Gamma^{S^+}$ and Γ^∞ to the translational $\mathbf{U}$ and rotational $\mathbf{\Omega}$ velocities using integral kernels. For this purpose, we need to circumvent the calculation of slip velocity, specifically step II, in the steps (8). In the next section, Background, we review earlier efforts to derive geometrical kernel in order to connect the surface flux to self-phoretic velocities for particles with uniform phoretic mobility. Subsequently, in the following section, we introduce the concepts of field-kernels and flux-kernels for spheres for more general scenarios where the particle can be passive or active and the phoretic mobility can be non-uniform.

1. Background. Nourhani and Lammert³ introduced a kernel-based integral equation for the translational self-phoretic velocity of an axisymmetric spheroidal swimmer, ranging from discotics to spheres to rod-like geometries, with uniform surface phoretic mobility, moving along its symmetry axis. The integral kernel weighs the local contribution of the surface flux to the particle's velocity. They showed that in rod-like geometries, the flux value near the particle's equator is almost zero, making the contribution of surface flux in this region to the velocity negligible. Therefore, in rod-like geometries, the major contribution of the surface flux to motion comes from the poles. From a design principle perspective, we can place the expensive active surface near the poles and a less expensive passive conductive region near the equator. As the aspect ratio decreases towards discotic geometries, the flux over a larger portion of the particle's surface contributes to the velocity. Schnitzer and Yariv³⁴ obtained an approximate expression for osmotic self-propulsion of slender particles with uniform phoretic mobility in terms of weighted integral of the surface flux. Lammert, Crespi, and Nourhani in a subsequent publication⁴ extended the flux-based geometrical kernel formalism to include translational and rotational velocities for particles of arbitrary geometry with uniform phoretic mobility. They recovered the earlier results^{3,34} as special cases.

The assumption of uniform phoretic mobility is the zeroth-order approximation for the surface properties of the swimmer. However, this approximation may not be suitable for scenarios where the surface properties of different materials on various parts of the swimmer's surface differ significantly, and we need to address the presence of a non-uniform distribution of phoretic mobility over the particle surface. The techniques and theorems introduced in ref. 3 and 4 are suitable for uniform phoretic mobility. To address the scenario of non-uniform phoretic mobility in this paper, we introduce two identities (12) for integration over a sphere's surface and employ them to derive a kernel-based formalism for the self-phoretic translational and rotational velocities of a sphere with a surface distribution of phoretic mobility.

2. Field and flux kernels for spheres. This paper focuses on the phoretic dynamics of a single sphere. The extension to other geometries and the interactions of the particle with other particles or boundaries will be the scope of future studies. We are interested in extending the scope of the velocity expressions in ref. 3 and 4, which are based on local flux Γ^{S^+} and uniform



phoretic mobility, to additionally include phoresis by the global flux Γ^∞ and non-uniform phoretic mobility. Moreover, beyond flux kernels, we will provide field kernels that connect the velocities to the values of the field over the particle surface S^+ .

For this purpose, using eqn (3) for slip velocity, we first rewrite the expressions for the translational and rotational velocities (7) in terms of surface integrals involving the surface gradients of the field,

$$\mathbf{U} = \frac{-1}{4\pi a^3} \int_{S^+} dS \mu (a \nabla_S \Psi), \quad (9a)$$

$$\boldsymbol{\Omega} = \frac{-3}{8\pi a^4} \int_{S^+} dS \mu \hat{\mathbf{n}} \times (a \nabla_S \Psi). \quad (9b)$$

and use them as the basis for obtaining the kernels. For a given phoresis phenomenon, we define translational $\mathbb{K}_t$ and rotational $\mathbb{K}_r$ field kernels that quantify the local contribution of the field over the particle surface to the translational and rotational velocities, respectively. For flux kernels we will have distinct local flux kernel $\mathcal{K}_i^{S^+}$ and global flux kernel $\mathcal{K}_i^\infty$ for translational ($i = t$) and rotational ($i = r$) velocities, using integrations equivalent to the integrations in eqn (9),

$$\int_{S^+} dS \mu (a \nabla_S \Psi) = \int_{S^+} dS \mathbb{K}_t \Psi \quad (10a)$$

$$= \frac{a}{D} \int_{S^+} dS \left(\mathcal{K}_t^{S^+} \hat{\mathbf{n}} \cdot \Gamma^{S^+} + \mathcal{K}_t^\infty \hat{\mathbf{n}} \cdot \Gamma^\infty \right), \quad (10b)$$

$$\int_{S^+} dS \mu \hat{\mathbf{n}} \times (a \nabla_S \Psi) = \int_{S^+} dS \mathbb{K}_r \Psi \quad (10c)$$

$$= \frac{a}{D} \int_{S^+} dS \left(\mathcal{K}_r^{S^+} \hat{\mathbf{n}} \cdot \Gamma^{S^+} + \mathcal{K}_r^\infty \hat{\mathbf{n}} \cdot \Gamma^\infty \right). \quad (10d)$$

Within these definitions, the kernels possess dimensions of phoretic mobility. As we will see shortly, in cases of non-uniform phoretic mobility, the kernels are functionals of the surface distribution of phoretic mobility.

Both locally-driven and globally-driven field gradients, manifested in Γ^{S^+} and Γ^∞ respectively, as specified in eqn (2), contribute to the surface distribution of the field $\Psi(\mathbf{x}_S)$ around the particle. The field kernels $\mathbb{K}_t$ and $\mathbb{K}_r$, which deal directly with the field, are independent of the origin of the field gradient, whether locally Γ^{S^+} or globally Γ^∞ . The field kernels for both locally-driven and globally-driven aspects of phoresis processes are conceptually the same.

On the other hand, in the case of flux kernels, we deal with two conceptually distinct types of fluxes: one distributed over the particle surface Γ^{S^+} and the other a uniform far-field Γ^∞ . Therefore, the integrations of the normal component of these fluxes over the particle surface in eqn (10b) and (10d), weighted by the associated local flux kernels $\mathcal{K}_i^{S^+}$ and global flux kernels $\mathcal{K}_i^\infty$, have distinct interpretations, as we elaborate below. The integration involving local flux kernels $\mathcal{K}_i^{S^+}$ and the normal component of the locally-driven flux Γ^{S^+} has the same intuitive interpretation as the integration involving field kernels and the

field; the local flux kernels $\mathcal{K}_i^{S^+}$ define the local contribution of the surface flux Γ^{S^+} to the particle's translational and rotational velocities.

However, in the integrals involving the global flux kernels $\mathcal{K}_i^\infty$, although the integration is over the particle surface S^+ , the value of the flux is far-field. Therefore, the term $\hat{\mathbf{n}} \cdot \Gamma^\infty$ in the integration $\int_{S^+} dS \mathcal{K}_i^\infty \hat{\mathbf{n}} \cdot \Gamma^\infty$ in eqn (10b) and (10d), while integrated over the particle surface, does not represent the surface flux of the field over the particle surface. For example, in the case of globally driven phoresis of an impermeable, non-polarizable passive particle, the surface flux of the field $-\mathcal{D} \hat{\mathbf{n}} \cdot \nabla \Psi|_{\mathbf{x}_S}$ is zero and different from $\hat{\mathbf{n}} \cdot \Gamma^\infty$. To visualize the scenario and develop an intuition about the integral $\int_{S^+} dS \mathcal{K}_i^\infty \hat{\mathbf{n}} \cdot \Gamma^\infty$, imagine a uniform field gradient with a flux value Γ^∞ throughout the entire space. Then, we place an imaginary sphere in this field gradient without distorting the field gradient, and integrate $\hat{\mathbf{n}} \cdot \Gamma^\infty$ over the surface of the imaginary particle weighted by the kernel $\mathcal{K}_i^\infty$. Therefore, while the flux Γ^∞ is a far-field value, we use the same value for integration over the surface of the sphere in the velocity expressions (19) as if the presence of the sphere has not distorted the uniform distribution of the far-field flux over the particle surface.

Another interpretation is as follows. Writing the uniform far-field in the form $\Gamma^\infty = \Gamma^\infty \hat{\mathbf{e}}$, where $\hat{\mathbf{e}}$ is a unit vector defining the direction of the global field gradient, the integral $\Gamma^\infty \int_{S^+} dS (\mathcal{K}_i^\infty \hat{\mathbf{n}}) \cdot \hat{\mathbf{e}}$ is effectively proportional to the weighted integration of the projection of the direction of the far-field flux on the normal surface vector, that is, $\hat{\mathbf{n}} \cdot \hat{\mathbf{e}}$. This expression shows how the particle dynamics result from the coupling of the direction $\hat{\mathbf{e}}$ of the far-field flux to the surface characteristics and orientation of the particle, quantified by the second-rank tensor $\mathcal{K}_i^\infty \hat{\mathbf{n}}$.

3. Field kernels

For the field kernels $\mathbb{K}_t$ and $\mathbb{K}_r$, we combine eqn (9), (10a) and (10c) and reformulate the velocities in terms of weighted surface integrals of the fields such that

$$\mathbf{U} = \frac{-1}{4\pi a^3} \int_{S^+} dS \Psi \mathbb{K}_t[\mu], \quad (11a)$$

$$\boldsymbol{\Omega} = \frac{-3}{8\pi a^4} \int_{S^+} dS \Psi \mathbb{K}_r[\mu]. \quad (11b)$$

where the position-dependent field kernels $\mathbb{K}_i(\mathbf{x}_S)$ for $i = t, r$ are functionals of the surface distribution of the phoretic mobility $\mu(\mathbf{x}_S)$.

In case more than one phoresis process was involved similar to the slip velocity (4) the resulting velocities are the linear sum of the velocities (11) corresponding to the distinct phoresis processes. Keeping in mind the functional dependence of the kernels on phoretic mobility is critical, especially when more than one phoresis process is involved. For example, consider a particle that performs self-electrophoresis while also being



subjected to a global concentration gradient of non-electrolyte solutes, thus undergoing diffusiophoresis simultaneously. Each phenomenon has its own field, $\Psi^{(\text{self-elec})}$ and $\Psi^{(\text{diff})}$, and phoretic mobility, $\mu^{(\text{self-elec})}$ and $\mu^{(\text{diff})}$, respectively, along with their corresponding kernels and phoretic velocities. The particle's velocities will be the sum of the velocities resulting from each phoresis process. Fortunately, while the position dependence of the phoretic mobility for each phoresis process may differ depending on the distribution of the surface composition of the sphere, the functional dependence of the kernels is independent of the phoretic mobility distribution, as will be discussed soon and shown in eqn (13).

To obtain the field kernels $\mathbb{K}_t(\mathbf{x}_S)$ and $\mathbb{K}_r(\mathbf{x}_S)$ over the particle surface in the velocity expressions (11), we exploit the following two identities, the proofs of which are provided in Appendix A. For a function $\Phi(\mathbf{x}_S)$ over the surface of a sphere, we have

$$\int_{S^+} dS (a \nabla_S \Phi) = \int_{S^+} dS (2 \hat{\mathbf{n}} \Phi), \quad (12a)$$

$$\int_{S^+} dS \hat{\mathbf{n}} \times (a \nabla_S \Phi) = 0. \quad (12b)$$

By setting $\Phi = \mu \Psi$ in the above identities, using $\mu(a \nabla_S \Psi) = a \nabla_S(\mu \Psi) - \Psi(a \nabla_S \mu)$, and comparing the results with the integral expressions (10a) and (10c), we obtain the translational and rotational field kernels,

$$\mathbb{K}_t[\mu] = -a \nabla_S \mu + 2 \hat{\mathbf{n}} \mu, \quad (13a)$$

$$\mathbb{K}_r[\mu] = -\hat{\mathbf{n}} \times (a \nabla_S \mu) = \hat{\mathbf{n}} \times \mathbb{K}_t[\mu], \quad (13b)$$

respectively, as functionals of phoretic mobility distribution. For a constant field ($\Psi = \text{cte}$) over the surface of a particle, the slip velocity (3) is zero, and we expect no phoretic motion. This necessitates that the surface integral of the field kernels over the particle surface be zero,

$$\int_{S^+} dS \mathbb{K}_i = 0, \quad i = t, r. \quad (14)$$

which can be mathematically demonstrated by setting a constant field $\Psi(\mathbf{x}_S) = 1$ in the integral expressions (10a) and (10c).

A glance at the field kernels (13) immediately offers some physical insights. A particle with uniform phoretic mobility, for which $\nabla_S \mu = 0$, has a zero rotational field kernel, $\mathbb{K}_r[\mu = \text{cte}] = 0$, meaning the particle does not rotate. Therefore, a heterogeneous distribution of the field over a sphere's surface does not induce rotation for a sphere if the phoretic mobility is uniform. Furthermore, in the scenario of uniform phoretic mobility, the translational field kernel simplifies to a radial vector field over the sphere surface, $\mathbb{K}_t[\mu = \text{cte}] = 2\mu \hat{\mathbf{n}}$. Consequently, the calculation of the translational velocity is reduced to the surface integration of the field, weighted by the surface normal vector and phoretic mobility.

Therefore, for a spherical particle undergoing a phoresis process with uniform phoretic mobility, the particle does not

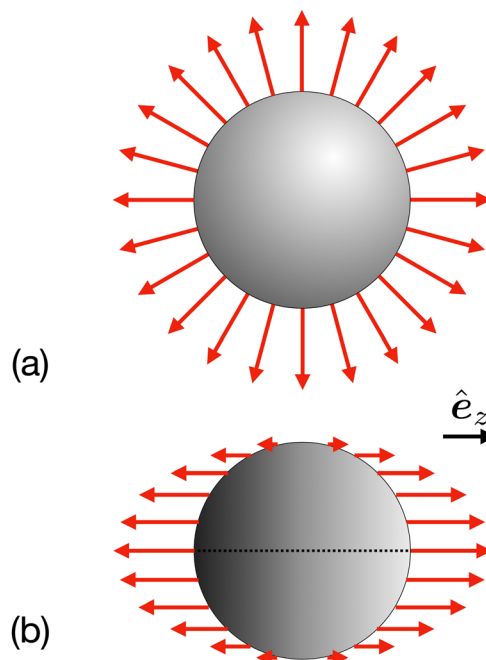


Fig. 1 For uniform phoretic mobility, the translational field kernel and flux kernels are (a) proportional to the normal $\hat{\mathbf{n}}$ of the surface, $\mathbb{K}_t = 2\mu \hat{\mathbf{n}}$, $\mathcal{K}_t^{S^+} = \mu \hat{\mathbf{n}}$, $\mathcal{K}_t^\infty = -3\mu \hat{\mathbf{n}}$, and (b) for the special case of an axisymmetric distribution of the field or flux, the $\hat{\mathbf{n}}$ term in kernels can be reduced to $\cos \theta \hat{\mathbf{e}}_z$, where $\hat{\mathbf{e}}_z$ is the symmetry axis. In this scenario, each kernel has a negligible value around the equator and maximum magnitude at the poles.

rotate, and its velocities are

$$\mu = \text{constant} \rightarrow \begin{cases} U = \frac{-\mu}{2\pi a^3} \int_{S^+} dS \Psi \hat{\mathbf{n}} \\ \Omega = 0 \end{cases} \quad (15)$$

In an axisymmetric scenario of constant phoretic mobility with the symmetry axis along $\hat{\mathbf{e}}_z$, the field is independent of the ϕ coordinate. Thus, the translational kernel $\mathbb{K}_t = 2\mu \hat{\mathbf{n}}$ effectively reduces to $\mathbb{K}_t = 2\mu \cos \theta \hat{\mathbf{e}}_z$, as shown in Fig. 1. Here, the field kernel provides us with another physical insight: in an axisymmetric scenario with uniform phoretic mobility, the field at the equator does not contribute to the motion, and the majority of the contribution comes from the poles.

The field kernels are solely functionals of phoretic mobilities and are independent of whether the distribution of the field over the particle surface is due to locally-driven or globally-driven field gradients. In the next section, we discuss how to quantify the local effect of geometry to the overall particle motion through flux kernels and their dependance on the origin of the flux.

4. Flux kernels

To obtain the flux kernels in terms of field kernels using eqn (10), we need to use the relationship between the fluxes Γ^{S^+} and Γ^∞ in the boundary conditions (2) and the driving field Ψ . In the limit of the thin interaction layer, the driving field



outside the interaction layer is often harmonic^{3,4,11} and satisfies the Laplace equation. Therefore, in what follows, we first provide the expression for the harmonic field and use it to find the flux kernels.

A. Harmonic driving field

For our analysis we use the spherical coordinate system, where the position vector $\mathbf{x} = (r, \theta, \phi)$ is identified by a radial coordinate r and two angular coordinates. Utilizing the orthogonality of spherical harmonics $Y_{\ell,m}(\theta, \phi)$, the solution to the Laplace equation,

$$\nabla^2 \Psi = 0, \quad (16)$$

subject to boundary conditions (2) yields the field,

$$\Psi(\mathbf{x}) = \mathcal{D}^{-1} \sum_{\ell=0}^{\infty} \sum_{m=-\ell}^{\ell} \frac{a}{\ell+1} \left(\frac{a}{r}\right)^{\ell+1} Y_{\ell,m} \left\{ \hat{\mathbf{n}} \cdot \mathbf{\Gamma}^{S^+} \right\}_{\ell,m} - \mathcal{D}^{-1} \left[1 + \frac{1}{2} \left(\frac{a}{r}\right)^3 \right] \mathbf{x} \cdot \mathbf{\Gamma}^{\infty}, \quad (17)$$

where we have employed the banana-curly-bracket notation $\{\cdot\}_{\ell,m}$ for the expansion coefficients in terms of spherical harmonics. That is, for a function $\Phi(\theta, \phi)$ defined over the surface of a sphere, we have

$$\Phi(\theta, \phi) = \sum_{\ell=0}^{\infty} \sum_{m=-\ell}^{\ell} \{\Phi\}_{\ell,m} Y_{\ell,m}(\theta, \phi) \quad (18a)$$

$$\{\Phi\}_{\ell,m} = \int_0^{\pi} \int_0^{2\pi} \Phi(\theta, \phi) Y_{\ell,m}^*(\theta, \phi) \sin \theta d\theta d\phi \quad (18b)$$

Since the gradient of the fields drives motion, constant terms in the fields do not contribute to the translational and rotational velocities, and we have ignored them in the solution (17).

B. Local and global flux kernels

To formulate the velocity expressions in terms of the local flux kernels $\mathcal{K}_i^{S^+}$ and global flux kernels $\mathcal{K}_i^{\infty}$, we combine eqn (9), (10b) and (10d) and recast the velocities in the form,

$$\mathbf{U} = \frac{-1}{4\pi a^2 \mathcal{D}} \int_{S^+} dS \left(\mathcal{K}_t^{S^+} \hat{\mathbf{n}} \cdot \mathbf{\Gamma}^{S^+} + \mathcal{K}_t^{\infty} \hat{\mathbf{n}} \cdot \mathbf{\Gamma}^{\infty} \right) \quad (19a)$$

$$\mathbf{\Omega} = \frac{-3}{8\pi a^3 \mathcal{D}} \int_{S^+} dS \left(\mathcal{K}_r^{S^+} \hat{\mathbf{n}} \cdot \mathbf{\Gamma}^{S^+} + \mathcal{K}_r^{\infty} \hat{\mathbf{n}} \cdot \mathbf{\Gamma}^{\infty} \right) \quad (19b)$$

in terms of weighted surface integrals of the normal components of the surface flux $\hat{\mathbf{n}} \cdot \mathbf{\Gamma}^{S^+}$ and far-field global flux $\hat{\mathbf{n}} \cdot \mathbf{\Gamma}^{\infty}$. As discussed following eqn (10), the integrands in these expressions have distinct interpretations. To obtain the local flux kernel, we consider a scenario where $\mathbf{\Gamma}^{S^+} \neq 0$ and $\mathbf{\Gamma}^{\infty} = 0$. For the global flux kernel, we study a situation where $\mathbf{\Gamma}^{S^+} = 0$ and $\mathbf{\Gamma}^{\infty} \neq 0$. The combination of these two scenarios leads to the boundary conditions (2).

In the first scenario, where phoresis is driven solely by a locally generated field gradient, by plugging the field

expression (17) over the particle surface,

$$\Psi(\mathbf{x}) = \frac{a}{\mathcal{D}} \sum_{\ell=0}^{\infty} \sum_{m=-\ell}^{\ell} \frac{1}{\ell+1} \left\{ \hat{\mathbf{n}} \cdot \mathbf{\Gamma}^{S^+} \right\}_{\ell,m} Y_{\ell,m} \quad (20)$$

into the kernel-based integral expressions (10) and performing some rearrangements, we obtain the flux kernels in terms of the field kernels for translational ($i = t$) and rotational ($i = r$) velocities,

$$\mathcal{K}_i^{S^+} = \sum_{\ell=0}^{\infty} \sum_{m=-\ell}^{\ell} \frac{1}{\ell+1} \{\mathbb{K}_i\}_{\ell,m} Y_{\ell,m} \quad (21)$$

which differs from the spherical harmonic expansion (18) of the field kernel $\mathbb{K}_i$ by the factor $1/(\ell+1)$ for each term.

The zero surface integral of the field kernel (14), $Y_{0,0}(\theta, \phi) = 1$, and the definition (18b) yield $\{\mathbb{K}_i\}_{0,0} = 0$. Additionally, the orthogonality of the spherical harmonics results in $\int_{S^+} dS Y_{\ell,m} = (4\pi)^{-1/2} a^2 \delta_{0\ell} \delta_{0m}$. Together with eqn (21), this leads to a zero surface integral of the flux kernels,

$$\int_{S^+} dS \mathcal{K}_i^{S^+} = 0, \quad i = t, r. \quad (22)$$

This confirms that a uniform distribution of activity does not lead to motility within the regime of our study.

In the second scenario, where the phoresis process is driven by a global gradient of the field in the absence of surface flux, $\mathbf{\Gamma}^{S^+} = 0$, the field (17) over the particle surface reduces to

$$\Psi(\mathbf{x}_S) = -\frac{3}{2} \frac{a}{\mathcal{D}} \hat{\mathbf{n}} \cdot \mathbf{\Gamma}^{\infty} \quad (23)$$

while the uniform flux $\mathbf{\Gamma}^{\infty}$ is far-field. By plugging this expression into the kernel-based integral expressions (10), we obtain the flux kernel to be proportional to the field kernel for a uniform global flux,

$$\mathcal{K}_i^{\infty} = -\frac{3}{2} \mathbb{K}_i, \quad i = t, r. \quad (24)$$

The zero surface integral of the field kernel (14) and the linearity of eqn (24) necessitate

$$\int_{S^+} dS \mathcal{K}_i^{\infty} = 0. \quad (25)$$

Although the flux $\mathbf{\Gamma}^{\infty}$ is uniform over the particle surface, what we integrate in eqn (19) is its normal component $\hat{\mathbf{n}} \cdot \mathbf{\Gamma}^{\infty}$, which is not uniform over the particle surface.

C. Discussion of the kernels

The field kernels (13) are functionals of phoretic mobility and are independent of the origin of the field gradient, provided the slip velocity follows the expression (3) for the involved phoresis processes. On the other hand, the flux kernels for locally-generated field gradients (21) and globally-generated field gradients (24) differ, and the integrated normal fluxes offer two distinct concepts.

In the uniform phoretic mobility scenario, $\nabla_S \mu = 0$, the rotational flux kernels $\mathcal{K}_r$, which are linearly proportional to the rotational field kernel $\mathbb{K}_r$, are also zero. Therefore, for an



active particle, heterogeneity in surface activity alone is not sufficient to induce rotation; heterogeneity in the surface distribution of phoretic mobility is necessary. Additionally, the translational flux kernels are proportional to the normal vector $\hat{n}$, thus,

$$\mu = \text{constant} \rightarrow \begin{cases} \mathcal{K}_t^{S+} = \mu \hat{n} \\ \mathcal{K}_t^{\infty} = -3\mu \hat{n} \\ \mathcal{K}_r^{S+} = \mathcal{K}_r^{\infty} = 0 \end{cases} \quad (26)$$

Therefore, for each phoresis process, the translational velocity is proportional to the surface integral of fluxes weighted by the unit normal. Similar to the scenario for the axisymmetric field, as shown in Fig. 1, in the case of axisymmetric normal flux, the $\hat{n}$ term in kernels (26) can be reduced to $\cos \theta \hat{e}_z$, where $\hat{e}_z$ is the symmetry axis. Again, we learn that in the axisymmetric scenario with constant phoretic mobility, the majority of the flux contribution to motion comes from the poles, and the flux around the equator has a negligible contribution.

In the next section, we study some examples to demonstrate the utility of the phoresis kernel theory. We start with a simple axisymmetric problem involving uniform phoretic mobility for a self-phoretic particle subjected to an additional globally-driven phoresis process, considering two scenarios: active-passive and sink-source surface configurations. Afterward, we tackle the general problem of an axisymmetric self-phoretic particle with multiple regions of distinct phoretic mobilities and also address two special cases. Finally, to present an example for the rotational kernel, we solve an analytically addressable problem for a non-axisymmetric scenario that contains both translational and rotational dynamics.

5. Examples with analytical solutions

A. Axisymmetric self-phoretic particle under global field gradient

In this example, we deal with axisymmetric Janus particles, each half covered with distinct materials and having uniform phoretic mobilities. The particles perform self-phoresis while simultaneously undergoing an additional globally-driven phoresis process. This latter process involves uniform phoretic mobility $\mu^{(\text{glb})}$ and is exposed to a globally-generated uniform flux in the $\hat{e}_z$ direction, that is, $\Gamma = \Gamma^{(\text{glb})} \hat{e}_z$ with field diffusivity $\mathcal{D}^{(\text{glb})}$.

The director $\hat{p}$ of the Janus particle points from one face to the other, defining the particle's orientation in space. Positioned along the symmetry axis, the direction of the director is arbitrary, yet it determines how we refer to the Janus particle based on its rear-front composition. For instance, in Fig. 2(a), the Janus particle is termed active-passive (a.p.), reflecting the orientation of its director: the normal component of the surface flux on the passive side is zero, while on the active side, it can be either negative or positive. Alternatively, in Fig. 2(b), both sides are active, and the Janus particle is described as sink-source (i.o.): the surface flux at the sink is negative, and at the source, it is positive.

The normal component of the surface flux in these two scenarios are represented by,

$$\hat{n} \cdot \Gamma^{(\text{a.p.})} = H\left(\theta - \frac{\pi}{2}\right) \Gamma_0^{(\text{a.p.})}, \quad \Gamma_0^{(\text{a.p.})} \neq 0, \quad (27a)$$

$$\hat{n} \cdot \Gamma^{(\text{i.o.})} = \left[1 - 2H\left(\theta - \frac{\pi}{2}\right)\right] \Gamma_0^{(\text{i.o.})}, \quad \Gamma_0^{(\text{i.o.})} > 0. \quad (27b)$$

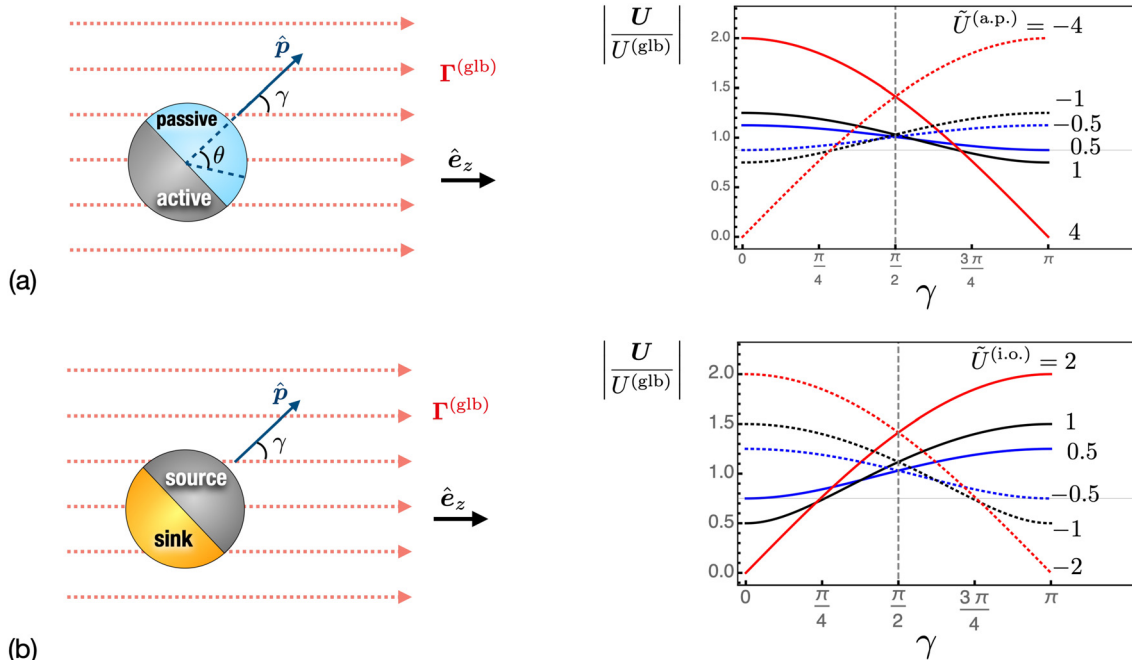


Fig. 2 Two self-phoretic Janus models of (a) active-passive (a.p.) and (b) sink-source (i.o.) composition are undergoing globally-driven phoresis by the far-field flux $\Gamma^{(\text{glb})}$. The angle γ defines the orientation of the Janus director $\hat{p}$ with respect to $\Gamma^{(\text{glb})}$. Solid lines represent positive values of $\tilde{U}^{(a)}$, while dashed lines represent negative values.



where $H(\theta)$ is the Heaviside step function. As shown in Fig. 2(a), the angle θ is the polar angle with respect to the Janus director $\hat{p}$. The globally-generated uniform flux $\Gamma^{(\text{glb})}$ forms an angle γ with the Janus director $\hat{p}$, such that $\hat{e}_z \cdot \hat{p} = \cos \gamma$.

For globally-driven phoresis, using the kernel for uniform phoretic mobility (26), the velocity expression (19a), and the identity $\int_{S^+} dS \hat{n} \hat{n} = \frac{4}{3} \pi a^2 \mathcal{I}$, we obtain the phoretic velocity,

$$\mathbf{U}^{(\text{glb})} = U^{(\text{glb})} \hat{e}_z = \frac{\mu^{(\text{glb})} \Gamma^{(\text{glb})}}{\mathcal{D}^{(\text{glb})}} \hat{e}_z. \quad (28)$$

For self-phoresis phenomena $\alpha = \text{a.p.}$, i.o., we define the ratio of velocity scales as

$$\tilde{U}^{(\alpha)} = \frac{\mu^{(\alpha)} \Gamma_0^{(\alpha)} / \mathcal{D}^{(\alpha)}}{U^{(\text{glb})}}. \quad (29)$$

Using the surface flux models (27), the translational velocity expression (19a), and the kernel for uniform phoretic mobility (26), we obtain the particle velocity for the two scenarios:

$$\mathbf{U} / U^{(\text{glb})} = \hat{e}_z + \frac{1}{4} \hat{p} \tilde{U}^{(\text{a.p.})} \quad (\text{active} - \text{passive}) \quad (30)$$

$$\mathbf{U} / U^{(\text{glb})} = \hat{e}_z - \frac{1}{2} \hat{p} \tilde{U}^{(\text{i.o.})} \quad (\text{sink} - \text{source}) \quad (31)$$

The particles have uniform phoretic mobilities, and thus their angular velocities are zero according to eqn (15).

Therefore, there is no deterministic change in orientation, γ , over time, and the globally-driven phoresis modulates the self-phoretic velocities, as shown in the curve of Fig. 2.

The scenario of an axisymmetric particle with uniform phoretic mobility does not take advantage of the terms in the translational and rotational kernels (13) that stem from the surface gradient of phoretic mobility. In the next two sections, we will take these terms into account to solve for problems with non-uniform phoretic mobility.

B. Axisymmetric self-phoretic particle with multiple surface regions

We start with a general scenario, shown in Fig. 3(a), where the surface of the sphere is partitioned into k axisymmetric regions such that the distribution of phoretic mobility over the sphere surface is described by

$$\mu(\theta) = \begin{cases} \mu^{(1)} & 0 \leq \theta < \theta_1 \\ \mu^{(2)} & \theta_1 \leq \theta < \theta_2 \\ \vdots & \vdots \\ \mu^{(k)} & \theta_{k-1} \leq \theta \leq \theta_k \equiv \pi \end{cases} \quad (32)$$

where θ_i 's defines the boundaries of the regions. The motion is along the symmetry axis $\hat{e}_z$, and using eqn (13a) and (21), the

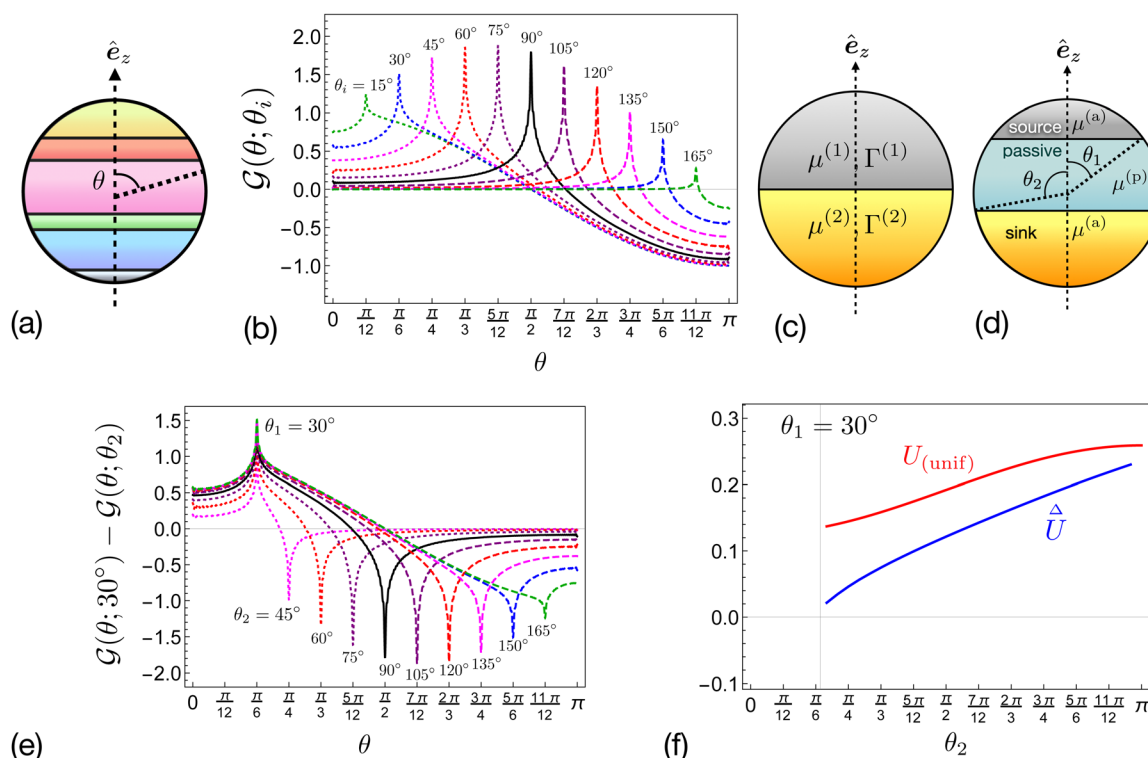


Fig. 3 (a) General axisymmetric scenario of a self-phoretic particle with multiple distinct regions. (b) The θ -dependent function $G(\theta - \theta_i)$ locally quantifying the disparity of the phoretic mobilities in the flux kernel (33b). (c) Self-phoretic particle with two active surfaces. (d) A source-sink self-phoretic particle with a passive region in the middle, and (e) the corresponding dimensionless component of the flux kernel due to the difference in the phoretic mobility of passive and active regions. (f) Dimensionless velocities for the model swimmer presented in Fig. 3(d).



corresponding field and flux kernels are

$$\hat{\mathbf{e}}_z \cdot \mathbb{K}_r = 2\mu^{(1)} \cos \theta + \sum_{j=1}^{k-1} [\mu^{(j+1)} - \mu^{(j)}] \mathbb{G}(\theta; \theta_j), \quad (33a)$$

$$\hat{\mathbf{e}}_z \cdot \mathcal{K}_t^{S+} = \mu^{(1)} \cos \theta + \sum_{j=1}^{k-1} [\mu^{(j+1)} - \mu^{(j)}] \mathcal{G}(\theta; \theta_j), \quad (33b)$$

respectively, where the dimensionless functions are

$$\mathbb{G}(\theta; \theta_i) = \sin \theta \delta(\theta - \theta_i) + 2 \cos \theta H(\theta - \theta_i), \quad (34a)$$

$$\begin{aligned} \mathcal{G}(\theta; \theta_i) = & \sum_{\ell=0}^{\infty} \frac{2\pi}{\ell+1} [Y_{\ell,0}(\theta_i, 0) \sin^2 \theta_i \\ & + \int_{\theta_i}^{\pi} Y_{\ell,0}(\theta, 0) \sin 2\theta d\theta] Y_{\ell,0}(\theta, 0), \end{aligned} \quad (34b)$$

and the Dirac delta function $\delta(\theta - \theta_i)$ is the derivative of the Heaviside function $H(\theta - \theta_i)$.

Fig. 3(b) shows the plot of $\mathcal{G}(\theta; \theta_i)$ for different values of θ_i . The function shows a peak at the boundary $\theta = \theta_i$. Therefore, we can expect $k - 1$ peaks in the flux kernel (33b) at the boundaries of the regions, and the strength of the peaks depends on the differences in phoretic mobilities of the neighboring regions forming the boundaries.

To demonstrate the utility of the formalism, we revisit a model Janus particle presented by Golestanian *et al.*³¹ and shown in Fig. 3(c). The model is fairly general and is not limited to a source-sink scenario. For example, both sides can be sources consisting of distinct materials with different surface fluxes, thus leading to an asymmetry condition. The distribution of surface flux and phoretic mobility over the surface is given by

$$[\hat{\mathbf{n}} \cdot \mathbf{\Gamma}^{S+}, \mu] = \begin{cases} [\Gamma^{(1)}, \mu^{(1)}] & 0 \leq \theta < \frac{\pi}{2} \\ [\Gamma^{(2)}, \mu^{(2)}] & \frac{\pi}{2} \leq \theta \leq \pi \end{cases} \quad (35)$$

The corresponding flux kernel (33b) along the symmetry axis direction for this problem takes the form

$$\hat{\mathbf{e}}_z \cdot \mathcal{K}_t^{S+} = \mu^{(1)} \cos \theta + [\mu^{(2)} - \mu^{(1)}] \mathcal{G}\left(\theta; \frac{\pi}{2}\right). \quad (36)$$

Using the expression (19a) for translational velocity and $\int dS \cos \theta \hat{\mathbf{n}} \cdot \mathbf{\Gamma}^{S+} = 2 \int dS \mathcal{G}\left(\theta; \frac{\pi}{2}\right) \hat{\mathbf{n}} \cdot \mathbf{\Gamma}^{S+} = \pi a^2 [\Gamma^{(1)} - \Gamma^{(2)}]$ we reobtain eqn (9) in ref. 31,

$$\mathbf{U} = \frac{-1}{8\mathcal{D}} [\mu^{(1)} + \mu^{(2)}] [\Gamma^{(1)} - \Gamma^{(2)}] \hat{\mathbf{e}}_z. \quad (37)$$

An important outcome of this equation is that in a scenario where $\mu^{(1)} = -\mu^{(2)}$, despite asymmetry in the structure and composition resulting in distinct surface fluxes, the active particle does not move.

In another scenario, shown in Fig. 3(d), we address a source-sink problem with a conductive passive region in the middle. This configuration could represent a classic platinum-gold pair with a chemically passive region in the middle that supports

the electrochemical electron current from source to sink. Therefore, in the phoretic mobility model (32), we can approximate the source and sink to have the same phoretic mobility, $\mu^{(a)} := \mu^{(1)} = \mu^{(3)}$, which is distinct from the phoretic mobility of the passive part in the middle, $\mu^{(p)} := \mu^{(2)}$. Consequently, the translational flux kernel (33b) reduces to

$$\hat{\mathbf{e}}_z \cdot \mathcal{K}_t^{S+} = \mu^{(a)} \cos \theta + [\mu^{(p)} - \mu^{(a)}] [\mathcal{G}(\theta; \theta_1) - \mathcal{G}(\theta; \theta_2)]. \quad (38)$$

The source-sink surface flux model, satisfying the steady-state condition $\int_S dS \hat{\mathbf{n}} \cdot \mathbf{\Gamma}^{S+} = 0$, is

$$\hat{\mathbf{n}} \cdot \mathbf{\Gamma}^{S+} = \begin{cases} \Gamma_0 & 0 \leq \theta \leq \theta_1 \\ 0 & \theta_1 < \theta \leq \theta_2 \\ -\frac{1 - \cos \theta_1}{1 + \cos \theta_2} \Gamma_0 & \theta_2 < \theta \leq \pi \end{cases} \quad (39)$$

Using the flux-based velocity expression (19a) for this scenario gives the translational velocity along the symmetry axis,

$$\frac{\hat{\mathbf{e}}_z \cdot \mathbf{U}}{U_{(\text{ref})}} = U_{(\text{unif})} + \left(\frac{\mu^{(p)}}{\mu^{(a)}} - 1 \right) \frac{A}{U}, \quad (40)$$

where the dimensionless terms are

$$U_{(\text{unif})} = (2 + \cos \theta_1 - \cos \theta_2) \sin^2 \frac{\theta_1}{2} \quad (41a)$$

$$\begin{aligned} \frac{A}{U} = & \int_0^{\theta_1} d\theta \sin \theta [\mathcal{G}(\theta; \theta_1) - \mathcal{G}(\theta; \theta_2)] \\ & - \frac{1 - \cos \theta_1}{1 + \cos \theta_2} \int_{\theta_2}^{\pi} d\theta \sin \theta [\mathcal{G}(\theta; \theta_1) - \mathcal{G}(\theta; \theta_2)]. \end{aligned} \quad (41b)$$

The reference velocity scale in eqn (40),

$$U_{(\text{ref})} = \frac{-\mu^{(a)}}{2\mathcal{D}} \Gamma_0, \quad (42)$$

used for non-dimensionalization, corresponds to a scenario where the particle is half-coated by source and sink surfaces with no passive surface in the middle, that is, $\theta_1 = \theta_2 = \pi/2$.

As an example, for a source region identified by the angle $\theta_1 = 30^\circ$, Fig. 3(e) shows the dependence of $\mathcal{G}(\theta; \theta_1) - \mathcal{G}(\theta; \theta_2)$ in the flux kernel (38) for different sizes of the sink region, identified by θ_2 . The curves show two peaks at the boundaries, θ_1 and θ_2 , of the regions. Fig. 3(f) shows the corresponding dimensionless velocities (41), both of which increase monotonically with an increase in θ_2 .

C. Non-axisymmetric self-phoretic particle

In this section, we discuss the scenario of a non-axisymmetric self-phoretic particle shown in Fig. 4(a). A passive particle with surface phoretic mobility $\mu^{(p)}$ is covered by two active surfaces, each with approximately the same phoretic mobility $\mu^{(a)}$. Within the spherical coordinate system, the polar angle with respect to the z -axis is denoted by θ , and as shown in Fig. 4(b), the angle θ_0 defines the active cap region. Half of the cap (*i.e.*, $y > 0$) acts as the source with surface flux $\Gamma_0 > 0$, and the other half serves as the sink with surface flux $-\Gamma_0$. The surface flux Γ



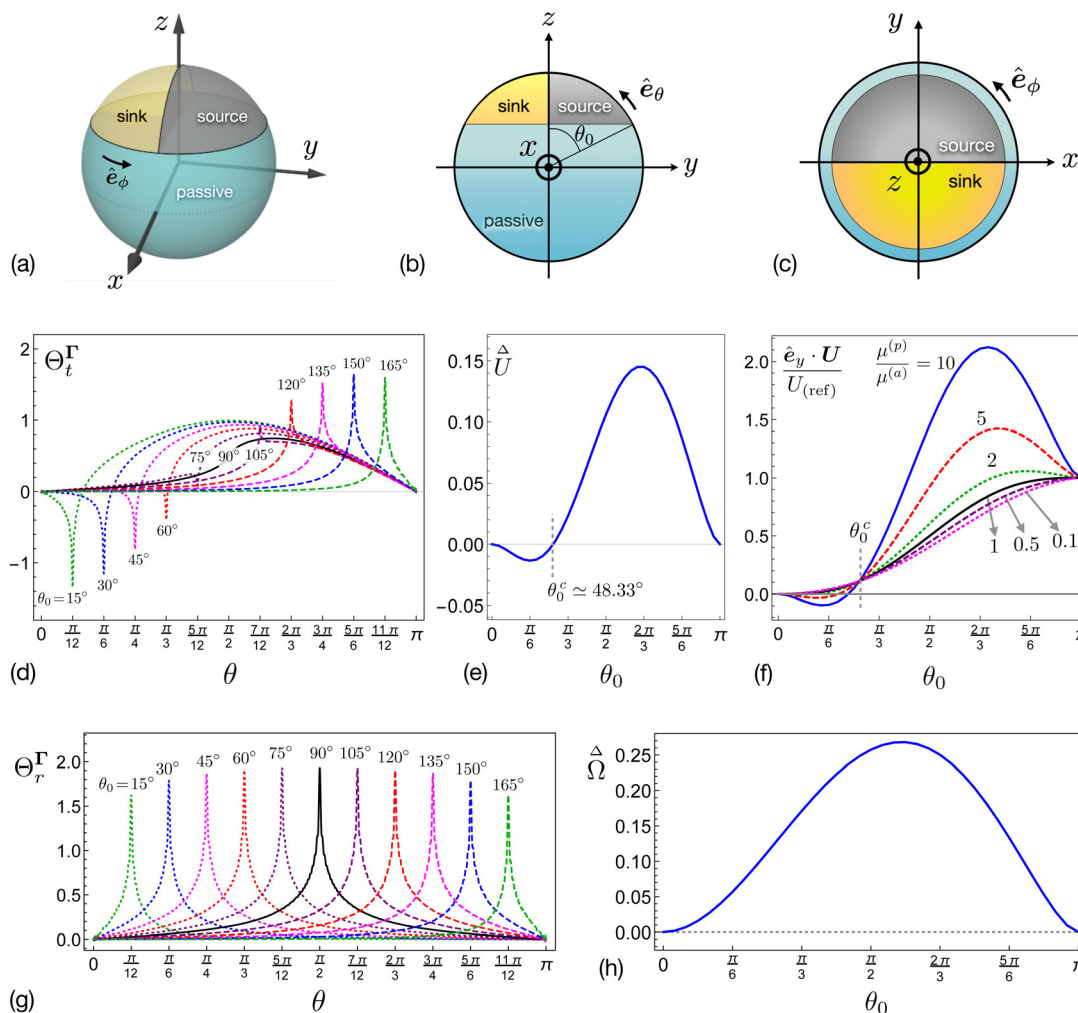


Fig. 4 Non-axisymmetric self-phoretic particle with rotational and translational dynamics. (a) The microswimmer is a passive particle covered by an active cap consisting of source and sink regions of equal area. (b) The size of the cap is quantified by θ_0 . (c) Top view. (d) The dimensionless function $\Theta_t^\Gamma(\theta; \theta_0)$ in the translational flux kernel (48) shows a peak at $\theta = \theta_0$. (e) The dimensionless function $\bar{U}$ in the translational velocity expression (49) determines the contribution of the difference in phoretic mobilities to motion. (f) The scaled velocity $\hat{\mathbf{e}}_y \cdot \mathbf{U}/U_{\text{ref}}$ for difference ratios of phoretic mobilities. (g) The θ -dependent function $\Theta_r^\Gamma(\theta; \theta_0)$ of the rotational flux-kernel (52b) shows peak at $\theta = \theta_0$. (h) The dimensionless function $\bar{\Omega}(\theta_0)$ shows the behavior of the magnitude of the angular velocity.

is symmetric about the yz plane, but antisymmetric about the xz plane.

Thus, the position-dependent phoretic mobility $\mu(\theta)$ is

$$\mu(\theta) = \mu^{(a)} + [\mu^{(p)} - \mu^{(a)}]H(\theta - \theta_0), \quad (43)$$

and the normal component of the flux $\hat{\mathbf{n}} \cdot \mathbf{\Gamma}^{S^+}$ over the surface of the sphere is given by

$$\hat{\mathbf{n}} \cdot \mathbf{\Gamma}^{S^+}(\theta, \phi) = [1 - H(\theta - \theta_0)][1 - 2H(\phi - \pi)]\Gamma_0. \quad (44)$$

We aim to obtain the translational $\mathbf{U}$ and rotational $\mathbf{\Omega}$ velocities of the self-phoretic particle.

1. Translational velocity. We start with the translational velocity, $\mathbf{U}$. As shown in Fig. 4(b), the yz -plane is the symmetry plane, and thus the particle has no preference to move in the $\hat{\mathbf{e}}_x$ or $-\hat{\mathbf{e}}_x$ direction. Therefore, there is no translation along the x direction, and

$$\hat{\mathbf{e}}_x \cdot \mathbf{U} = 0. \quad (45)$$

To show that there is no motion in the z direction, we first write the translational field kernel (13a) for our problem. Since $\hat{\mathbf{e}}_z \cdot \mathbb{K}_t = \sin \theta \partial_\theta \mu + 2\mu \cos \theta$ has no ϕ -dependence, the expansion coefficients $\{\hat{\mathbf{e}}_z \cdot \mathbb{K}_t\}_{\ell, m}$ for the flux kernel (21) are non-zero only when $m = 0$. Therefore, the z -component of the flux kernel, $\hat{\mathbf{e}}_z \cdot \mathbf{\mathcal{K}}_t^{S^+}$, has no ϕ -dependence. On the other hand, $\int_0^{2\pi} d\phi \hat{\mathbf{n}} \cdot \mathbf{\Gamma}^{S^+} = 0$. Hence,

$$\hat{\mathbf{e}}_z \cdot \mathbf{U} \propto \int_{S^+} dS \hat{\mathbf{e}}_z \cdot \mathbf{\mathcal{K}}_t^{S^+} \hat{\mathbf{n}} \cdot \mathbf{\Gamma}^{S^+} = 0. \quad (46)$$

Hence, the particle only moves along the y -direction, and using eqn (13a) the field kernel is given by

$$\hat{\mathbf{e}}_y \cdot \mathbb{K}_t = \sin \phi \left\{ \mu^{(a)} (2 \sin \theta) + [\mu^{(p)} - \mu^{(a)}] \Theta_t^\Psi \right\} \quad (47a)$$

$$\Theta_t^\Psi(\theta; \theta_0) = 2H(\theta - \theta_0) \sin \theta - \cos \theta \delta(\theta - \theta_0) \quad (47b)$$

The dependence of the field kernel (47a) on $\sin \phi$ necessitates that the flux kernel coefficients $\{\hat{\mathbf{e}}_y \cdot \mathbb{K}_t\}_{\ell, m}$ can be non-



zero only when $m = \pm 1$. Thus, using the expression (21) for the flux kernel of self-phoresis and plugging in the field kernel (47), we obtain the y -component of the flux kernel,

$$\hat{e}_y \cdot \mathcal{K}_t^{S+} = \sin \phi \left\{ \mu^{(a)} \sin \theta + \left[\mu^{(p)} - \mu^{(a)} \right] \Theta_t^\Gamma(\theta; \theta_0) \right\}, \quad (48a)$$

$$\Theta_t^\Gamma(\theta; \theta_0) = \sum_{\ell=1}^{\infty} \frac{2\pi}{\ell+1} \left\{ 2 \int_{\theta_0}^{\pi} Y_{\ell,1}(\theta', 0) (\sin \theta')^2 d\theta' - \sin \theta_0 \cos \theta_0 Y_{\ell,1}(\theta_0, 0) \right\} Y_{\ell,1}(\theta, 0). \quad (48b)$$

The θ -dependent part Θ_t^Γ of the translational flux kernel determines the significance of the difference between the phoretic mobilities of active and passive surfaces and is plotted for different values of θ_0 in Fig. 4(d). The function Θ_t^Γ has a peak at $\theta = \theta_0$.

As reference speed, similar to the axisymmetric scenario, we use that for $\theta_0 = \pi$ (a swimmer fully and symmetrically covered by two active surfaces), namely eqn (42). By plugging the flux kernel (48) and the model surface flux (44) into the velocity expression (19a), we obtain the scaled translational velocity along the y direction, similar to eqn (40),

$$\frac{\hat{e}_y \cdot \mathbf{U}}{U_{(\text{ref})}} = U_{(\text{unif})} + \left(\frac{\mu^{(p)}}{\mu^{(a)}} - 1 \right) \bar{U}, \quad (49)$$

with the dimensionless functions

$$U_{(\text{unif})} = \frac{\theta_0}{\pi} - \frac{\sin 2\theta_0}{2\pi}, \quad (50a)$$

$$\bar{U} = \frac{2}{\pi} \int_0^{\theta_0} d\theta \sin \theta \Theta_t^\Gamma(\theta; \theta_0). \quad (50b)$$

The function $\bar{U}$, plotted in Fig. 4(e), quantifies the effect of the disparity in phoretic mobilities between passive and active surfaces on the translational velocity.

We can see that the function $\bar{U}$ must go to zero at $\theta_0 = \pi$ by definition of $U_{(\text{ref})}$, and at $\theta_0 = 0$ because $\mu^{(p)}$ is meaningless there, so that the prefactor of $\bar{U}$ is effectively arbitrary. Apart from that, there is a critical value $\theta_0^c \approx 48.33^\circ$, such that the function is negative for smaller, and positive for larger values of θ_0 . Interestingly, at $\theta_0 = \theta_0^c$, the velocity is independent of $\mu^{(p)}$, even though the passive area is larger than the active area.

Fig. 4(f) shows the scaled velocity (49) of the microswimmer for different ratios of the phoretic mobility of the passive surface to that of the active surface. All curves intersect at the same point at the critical θ_0^c , where $\hat{e}_y \cdot \mathbf{U}/U_{(\text{ref})} \approx 0.11$. Furthermore, for some ratios of $\mu^{(p)}/\mu^{(a)}$, we observe maximum values for the scaled velocity at $\theta_0^{\text{max}} < \pi$, suggesting that a particle with smaller active surface coverage can move faster than a particle with a fully covered active surface.

2. Rotational velocity. The analysis of the microswimmer's rotational dynamics requires less calculations compared to translational velocity. The constant term in the expression (43) for phoretic mobility does not contribute to the rotation of a phoretic sphere. As a result, the rotational velocity is essentially proportional to the difference in phoretic mobility between the passive and active surfaces. The following symmetry considerations indicate that the

angular velocity is purely along $\hat{e}_x$. Reflecting the particle about the antisymmetric xz plane effectively swaps the source and sink regions, leading to the transformation $\Gamma_0 \rightarrow -\Gamma_0$. This transformation, along with eqn (19b), results in the inversion of the z component of the angular velocity, $\hat{e}_z \cdot \boldsymbol{\Omega} \rightarrow -\hat{e}_z \cdot \boldsymbol{\Omega}$. However, this reflection is equivalent to rotating the sphere around $\hat{e}_z$ by an angle of π , which does not change the sign of $\hat{e}_z \cdot \boldsymbol{\Omega}$; consequently, $\hat{e}_z \cdot \boldsymbol{\Omega}$ must be zero. Furthermore, reflecting the particle about the symmetry yz plane preserves the direction of $\hat{e}_y \cdot \boldsymbol{\Omega}$, while it should reverse its direction. This requires $\hat{e}_y \cdot \boldsymbol{\Omega}$ to be zero as well.

Using the general expression (13b) for the rotational field kernel and the example phoretic mobility (43), we can derive the rotational field kernel $\mathbb{K}_r$. Subsequently, using the general expression for self-phoresis flux-kernel (21), we obtain the rotational flux kernel $\mathcal{K}_r^{S+}$ as follows

$$\mathbb{K}_r = -\hat{e}_\phi \left[\mu^{(p)} - \mu^{(a)} \right] \delta(\theta - \theta_0), \quad (51a)$$

$$\mathcal{K}_r^{S+} = -\hat{e}_\phi \left[\mu^{(p)} - \mu^{(a)} \right] \Theta_r^\Gamma(\theta; \theta_0). \quad (51b)$$

Expanding the θ -dependence in terms of spherical harmonics with zero azimuthal angle ($\phi = 0$),

$$\delta(\theta - \theta_0) = 2\pi \sum_{\ell=1}^{\infty} Y_{\ell,1}(\theta_0, 0) Y_{\ell,1}(\theta, 0), \quad (52a)$$

$$\Theta_r^\Gamma(\theta; \theta_0) = 2\pi \sum_{\ell=1}^{\infty} Y_{\ell,1}(\theta_0, 0) Y_{\ell,1}(\theta, 0) \left[\frac{\sin \theta_0}{\ell+1} \right]. \quad (52b)$$

shows that the differentiation between the field kernel and the flux kernel lies in the expansion coefficient within the square brackets for each ℓ . The dimensionless function $\Theta_r^\Gamma(\theta; \theta_0)$, plotted in Fig. 4(g) for different values of θ_0 , shows a peak at θ_0 .

Since a fully covered particle (half source/half sink) does not rotate, defining the characteristic angular velocity scale is not as straightforward as in the case of translational motion. We define the angular velocity scale as

$$\Omega_{(\text{ref})} = U_{(\text{ref})}/a, \quad (53)$$

representing the angular velocity of a sphere moving at a linear speed $U_{(\text{ref})}$, defined in eqn (42), along a circular trajectory of radius a . Thus, using the flux kernel (52b), the surface flux model (44), and the angular velocity expression (19b), we derive the scaled angular velocity as

$$\frac{\boldsymbol{\Omega}}{\Omega_{(\text{ref})}} = \hat{e}_x \left(\frac{\mu^{(p)}}{\mu^{(a)}} - 1 \right) \bar{\Omega}(\theta_0), \quad (54)$$

where the dimensionless function

$$\bar{\Omega}(\theta_0) = \frac{3}{\pi} \int_0^{\theta_0} d\theta \sin \theta \Theta_r^\Gamma(\theta; \theta_0) \quad (55)$$

is illustrated in Fig. 4(h) and is non-negative. Hence, the direction of rotation is primarily determined by the ratio of phoretic mobilities. As expected, the angular velocities are zero at the poles ($\theta_0 = 0$, no active surface) and ($\theta_0 = \pi$, full active surface) due to the lack of chiral asymmetry in the structure. The maximum angular



velocity occurs at a coverage of $\theta_0 = 103.32^\circ$, where the active area exceeds half the surface area of the sphere.

6. Conclusion

The kernel-based approach discussed in this work provides an analytical framework that simplifies the analysis of phoresis phenomena by breaking it down into elementary problems. This allows for individual solutions that can be compounded to address the main problem. Integral kernels enhance understanding of the impact of geometry and phoretic mobility variations on translational and rotational velocities. This necessitates a fundamental understanding of material composition and surface properties, especially when phoretic mobility is non-uniform across the particle surface. The connection between the driving field, its flux distribution, and particle velocities is established through integral kernels. These kernels quantify local contributions of the field or flux to velocities, amalgamating the effects of geometry and surface composition into a position-dependent weight function for integration. This framework offers insights into phoretic particle behavior and the fine-tuning of their properties based on geometrical considerations. Our formalism for studying the phoresis of particles, compliant with the boundary condition (2), paves the way for future studies on non-spherical particles.

Data availability

This article is a theory paper, has no data other than plotting the equations in the draft.

Conflicts of interest

There are no conflicts to declare.

Appendix

Appendix A: Proofs of identities

We establish the validity of identities (12). Given the geometric isotropy of a sphere, which lacks a preferred direction, we can demonstrate each identity along any arbitrarily chosen fixed direction, such as $\hat{e}_z$. Consider the expression:

$$\hat{e}_z \cdot (a \nabla_S \Phi) = \hat{e}_z \cdot (\mathcal{I} - \hat{n} \hat{n}) \cdot (a \nabla \Phi) = -\sin \theta \frac{\partial \Phi}{\partial \theta}, \quad (\text{A1})$$

we have

$$\begin{aligned} \hat{e}_z \cdot \int_{S^+} dS (a \nabla_S \Phi) &= a^2 \int_0^{2\pi} d\phi \int_0^\pi (-\sin^2 \theta) \frac{\partial \Phi}{\partial \theta} d\theta \\ &= a^2 \int_0^{2\pi} d\phi \left[\underbrace{(-\sin^2 \theta) \Phi}_{\theta=0}^{\theta=\pi} \right]_{\theta=0}^{\theta=\pi} \\ &\quad - a^2 \int_0^{2\pi} d\phi \int_0^\pi (-2 \sin \theta \cos \theta) \Phi d\theta \\ &= \hat{e}_z \cdot \int_{S^+} 2 \hat{n} \Phi dS \end{aligned} \quad (\text{A2})$$

where the first step utilizes integration by parts. In the last step, we apply $dS = a^2 \sin \theta d\theta d\phi$ and $\hat{e}_z \cdot \hat{n} = \cos \theta$. As $\hat{e}_z$ is an arbitrary constant unit vector, it can be eliminated from both sides, thus proving identity (12a).

The proof of the second identity is more straightforward. Given that $\hat{e}_z \cdot \hat{e}_\phi = 0$ and $\hat{e}_z \cdot \hat{e}_\theta = -\sin \theta$, we find:

$$\begin{aligned} \hat{e}_z \cdot \int_{S^+} dS \hat{n} \times (a \nabla_S \Phi) &= \int_{S^+} dS \frac{\partial \Phi}{\partial \theta} (\hat{e}_z \cdot \hat{e}_\phi) \\ &\quad + \int_{S^+} dS (-\hat{e}_z \cdot \hat{e}_\theta) \frac{1}{\sin \theta} \frac{\partial \Phi}{\partial \phi} \quad (\text{A3}) \\ &= \int_{S^+} dS \frac{\partial \Phi}{\partial \phi} = 0, \end{aligned}$$

where in the last step we used the periodic property $\Phi(\theta, 0) = \Phi(\theta, 2\pi)$.

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Supplementary Information for “Phoresis Kernel Theory for Passive and Active Spheres with Nonuniform Phoretic Mobility”

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This supplementary information provides detailed calculations for the problems presented in the main manuscript. It is intended for pedagogical purposes, primarily aimed at junior researchers, and thus includes a high level of detail that may be superfluous for more advanced researchers. The derivations are crafted with this educational intent in mind, although alternative approaches could yield shorter and more concise derivations.

1. HARMONIC DRIVING FIELD

For a sphere immersed in a fluid to exhibit phoretic motion, it may be exposed to a gradient of a harmonic driving field Ψ . In the limit of a thin interaction layer, outside the interaction layer, this field obeys the Laplace equation,

$$\nabla^2 \Psi = 0, \quad (1)$$

and is subjected to specific surface and far-field boundary conditions:

$$\begin{cases} \hat{\mathbf{n}} \cdot \nabla \Psi|_{S^+} = -\hat{\mathbf{n}} \cdot \mathbf{\Gamma}^{S^+}/\mathcal{D}, \\ \nabla \Psi|_{\infty} = -\mathbf{\Gamma}^{\infty}/\mathcal{D} \equiv -\mathbf{\Gamma}^{\infty} \hat{\mathbf{e}}_z/\mathcal{D}, \end{cases} \quad \text{or equivalently,} \quad \begin{cases} \partial_r \Psi|_{S^+} = -\hat{\mathbf{n}} \cdot \mathbf{\Gamma}^{S^+}/\mathcal{D}, \\ \partial_r \Psi|_{\infty} = -\mathbf{\Gamma}^{\infty} \cos \theta/\mathcal{D} = -r^{-1} \mathbf{\Gamma}^{\infty} \cdot \mathbf{x}/\mathcal{D}, \end{cases} \quad (2)$$

where S^+ represents the outer surface of the interaction layer ($r_{S^+} \approx a$), $\mathcal{D}$ is the diffusion coefficient associated with the phoretic process, and the relationship $\hat{\mathbf{e}}_r \cdot \hat{\mathbf{e}}_z = \cos \theta$ is used in the spherical coordinate system $\mathbf{x} = r\hat{\mathbf{e}}_r = (r, \theta, \phi)$.

For convenience, we divide the driving field $\Psi = \Psi^{(p)} + \Psi^{(a)}$ into the harmonic field $\Psi^{(p)}$, which accounts for the phoresis of a non-polarizable passive particle with the same geometry solely due to the far-field flux $\mathbf{\Gamma}^{\infty}$, and the harmonic field $\Psi^{(a)}$, which accounts for the self-phoresis of an active particle of the same geometry solely due to the surface flux $\mathbf{\Gamma}^{S^+}$. The boundary conditions for these two scenarios are

$$\begin{cases} \partial_r \Psi^{(p)}|_{S^+} = 0, \\ \partial_r \Psi^{(p)}|_{\infty} = -\mathbf{\Gamma}^{\infty} \cos \theta/\mathcal{D} = -r^{-1} \mathbf{\Gamma}^{\infty} \cdot \mathbf{x}/\mathcal{D}, \end{cases} \quad \begin{cases} \partial_r \Psi^{(a)}|_{S^+} = -\hat{\mathbf{n}} \cdot \mathbf{\Gamma}^{S^+}/\mathcal{D}, \\ \partial_r \Psi^{(a)}|_{\infty} = 0, \end{cases} \quad (3)$$

The general solution to the harmonic driving field $\Psi^{(\alpha)}$, expressed in terms of spherical harmonics

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$Y_{\ell m}(\theta, \phi)$, is

$$\Psi^{(\alpha)} = A_{00}^{(\alpha)} + \sum_{\ell=1}^{\infty} \sum_{-\ell \leq m \leq \ell} A_{\ell m}^{(\alpha)} r^{\ell} Y_{\ell m} + \sum_{\ell=0}^{\infty} \sum_{-\ell \leq m \leq \ell} B_{\ell m}^{(\alpha)} r^{-(\ell+1)} Y_{\ell m}, \quad (4)$$

with the radial derivative,

$$\frac{\partial \Psi^{(\alpha)}}{\partial r} = \sum_{\ell=1}^{\infty} \sum_{-\ell \leq m \leq \ell} \ell A_{\ell m}^{(\alpha)} r^{\ell-1} Y_{\ell m} + \sum_{\ell=0}^{\infty} \sum_{-\ell \leq m \leq \ell} [-(\ell+1)] B_{\ell m}^{(\alpha)} r^{-(\ell+2)} Y_{\ell m}. \quad (5)$$

We set $A_{00}^{(\alpha)} = 0$, or in other words, we are interested in the disturbance field. We will first obtain the solution for the passive phoresis $\Psi^{(p)}$ and then for the active self-phoresis $\Psi^{(a)}$, and add them up to obtain the harmonic driving field $\Psi = \Psi^{(p)} + \Psi^{(a)}$.

For the passive phoresis field, using $Y_{10} \propto \cos \theta$ and applying the far-field boundary condition yields $A_{10}^{(p)} Y_{10} = -\Gamma^{\infty} \cos \theta / \mathcal{D}$, $A_{1,1}^{(p)} = A_{1,-1}^{(p)} = 0$, and $A_{\ell m}^{(p)} = 0$ for $\ell \geq 2$. Then, applying the surface flux condition for the passive non-polarizable particle

$$\partial_r \Psi^{(p)}|_{r=a} = 0 = -\Gamma^{\infty} \cos \theta / \mathcal{D} + \sum_{\ell=0}^{\infty} \sum_{-\ell \leq m \leq \ell} [-(\ell+1)] B_{\ell m}^{(p)} a^{-(\ell+2)} Y_{\ell m} \quad (6)$$

yields $B_{0,0}^{(p)} = B_{1,1}^{(p)} = B_{1,-1}^{(p)} = 0$ and $B_{\ell m}^{(p)} = 0$ for $\ell \geq 2$. The equation $-\Gamma^{\infty} \cos \theta / \mathcal{D} - 2a^{-3} B_{10}^{(p)} Y_{10} = 0$ leads to $B_{10}^{(p)} Y_{10} = -\frac{1}{2} a^3 \Gamma^{\infty} \cos \theta / \mathcal{D}$. Thus, using $\hat{e}_z \cdot \mathbf{x} = r \cos \theta$ we obtain

$$\begin{aligned} \Psi^{(p)}(\mathbf{x}) &= A_{10}^{(p)} Y_{10} r + B_{10}^{(p)} r^{-2} Y_{10} \\ &= -\Gamma^{\infty} r \cos \theta / \mathcal{D} - \frac{a^3}{2r^2} \Gamma^{\infty} \cos \theta / \mathcal{D} \\ &= -\Gamma^{\infty} \hat{e}_z \cdot \mathbf{x} / \mathcal{D} - \frac{a^3}{2r^3} \Gamma^{\infty} \hat{e}_z \cdot \mathbf{x} / \mathcal{D} \\ &= -\Gamma^{\infty} \cdot \mathbf{x} / \mathcal{D} - \frac{a^3}{2r^3} \Gamma^{\infty} \cdot \mathbf{x} / \mathcal{D} \\ &= -\mathcal{D}^{-1} \left[1 + \frac{1}{2} \left(\frac{a}{r} \right)^3 \right] \Gamma^{\infty} \cdot \mathbf{x} \end{aligned} \quad (7)$$

Before proceeding with the active phoresis field, we define the banana-bracket notation $\{\!\!\{\cdot\}\!\!\}_{\ell m}$ for the expansion coefficients in terms of spherical harmonics. That is, for a function $\Phi(\theta, \phi)$ defined over the surface of a sphere, we have

$$\Phi(\theta, \phi) = \sum_{\ell=0}^{\infty} \sum_{m=-\ell}^{\ell} \{\!\!\{\Phi\}\!\!\}_{\ell m} Y_{\ell m}(\theta, \phi) \quad (8a)$$

$$\{\!\!\{\Phi\}\!\!\}_{\ell m} = \int_0^{\pi} \int_0^{2\pi} \Phi(\theta, \phi) Y_{\ell m}^*(\theta, \phi) \sin \theta d\theta d\phi \quad (8b)$$

Now, for the active phoresis field, the far-field zero radial gradient condition $\partial_r \Psi^{(a)}|_{\infty} = 0$ neces-

satisfies $A_{\ell m}^{(a)} = 0$. The surface flux condition yields

$$\partial_r \Psi^{(a)}|_{S^+} = \sum_{\ell=0}^{\infty} \sum_{-\ell \leq m \leq \ell} [-(\ell+1)] B_{\ell m}^{(a)} a^{-(\ell+2)} Y_{\ell m} = -\hat{\mathbf{n}} \cdot \mathbf{\Gamma}^{S^+} / \mathcal{D} = - \sum_{\ell=0}^{\infty} \sum_{m=-\ell}^{\ell} \mathcal{D}^{-1} \left\{ \hat{\mathbf{n}} \cdot \mathbf{\Gamma}^{S^+} \right\}_{\ell m} Y_{\ell, m} \quad (9)$$

where, in the last step, we have expanded the surface flux $\hat{\mathbf{n}} \cdot \mathbf{\Gamma}^{S^+}$ in terms of spherical harmonics. Exploiting the orthogonality of the spherical harmonics leads to $B_{\ell m}^{(a)} = a^{\ell+2}(\ell+1)^{-1} \mathcal{D}^{-1} \left\{ \hat{\mathbf{n}} \cdot \mathbf{\Gamma}^{S^+} \right\}_{\ell m}$ and we obtain the active self-phoresis field,

$$\Psi^{(a)}(\mathbf{x}) = \mathcal{D}^{-1} \sum_{\substack{\ell=0 \\ -\ell \leq m \leq \ell}}^{\infty} \frac{a}{\ell+1} \left(\frac{a}{r} \right)^{\ell+1} Y_{\ell, m} \left\{ \hat{\mathbf{n}} \cdot \mathbf{\Gamma}^{S^+} \right\}_{\ell m}. \quad (10)$$

Finally, we obtain the disturbance field by superposing the fields due to passive phoresis (7) and active self-phoresis (10),

$$\Psi(\mathbf{x}) = \Psi^{(a)}(\mathbf{x}) + \Psi^{(p)}(\mathbf{x}) = \mathcal{D}^{-1} \sum_{\substack{\ell=0 \\ -\ell \leq m \leq \ell}}^{\infty} \frac{a}{\ell+1} \left(\frac{a}{r} \right)^{\ell+1} Y_{\ell, m} \left\{ \hat{\mathbf{n}} \cdot \mathbf{\Gamma}^{S^+} \right\}_{\ell m} - \mathcal{D}^{-1} \left[1 + \frac{1}{2} \left(\frac{a}{r} \right)^3 \right] \mathbf{x} \cdot \mathbf{\Gamma}^{\infty}, \quad (11)$$

which is Eq (17) in the main manuscript.

2. AXISYMMETRIC SELF-PHORETIC PARTICLE WITH MULTIPLE SURFACE REGIONS

For the general axisymmetric scenario with k axisymmetric regions, the distribution of phoretic mobility over the sphere surface is described by

$$\mu(\theta) = \begin{cases} \mu^{(1)} & 0 \leq \theta < \theta_1 \\ \mu^{(2)} & \theta_1 \leq \theta < \theta_2 \\ \vdots & \vdots \\ \mu^{(k)} & \theta_{k-1} \leq \theta \leq \theta_k \equiv \pi \end{cases} \quad (12)$$

where θ_i 's define the boundaries of the regions. We can rewrite the expression using the Heaviside step function:

$$H(\theta - \theta') = \begin{cases} 0 & \theta < \theta' \\ 1 & \theta \geq \theta' \end{cases} \quad (13)$$

as

$$\begin{aligned} \mu(\theta) &= \mu^{(1)} [1 - H(\theta - \theta_1)] + \mu^{(2)} [H(\theta - \theta_1) - H(\theta - \theta_2)] + \dots + \mu^{(k-1)} [H(\theta - \theta_{k-2}) - H(\theta - \theta_{k-1})] + \mu^{(k)} H(\theta - \theta_{k-1}) \\ &= \mu^{(1)} + [\mu^{(2)} - \mu^{(1)}] H(\theta - \theta_1) + [\mu^{(3)} - \mu^{(2)}] H(\theta - \theta_2) + \dots + [\mu^{(k)} - \mu^{(k-1)}] H(\theta - \theta_{k-1}) \\ &= \mu^{(1)} + \sum_{j=1}^{k-1} [\mu^{(j+1)} - \mu^{(j)}] H(\theta - \theta_j) \end{aligned} \quad (14)$$

In the last term of the first line, we have omitted $-\mu^{(k)}H(\theta-\theta_k)$ to include $\theta_k = \pi$ in the calculations. Additionally, this term is non-zero in the domain $\theta > \theta_k = \pi$, which does not enter the calculations.

The motion is along the symmetry axis $\hat{e}_z$. Using $\hat{e}_z \cdot \hat{e}_\theta = -\sin \theta$, we have $a \hat{e}_z \cdot \nabla_S = -\sin \theta \partial_\theta$. Combining this expression with $\hat{e}_z \cdot \hat{n} = \hat{e}_z \cdot \hat{e}_r = \cos \theta$ and $\partial_\theta H(\theta - \theta') = \delta(\theta - \theta')$, where $\delta(\theta)$ is the Dirac delta function, the translational field kernel is given by

$$\begin{aligned}
\hat{e}_z \cdot \mathbb{K}_t &= -a \hat{e}_z \cdot \nabla_S \mu + 2 \hat{e}_z \cdot \hat{n} \mu \\
&= \sin \theta \partial_\theta \mu + 2 \mu \cos \theta \\
&= \sin \theta \sum_{j=1}^{k-1} [\mu^{(j)} - \mu^{(j-1)}] \delta(\theta - \theta_{j-1}) + 2 \mu^{(1)} \cos \theta + 2 \cos \theta \sum_{j=1}^{k-1} [\mu^{(j)} - \mu^{(j-1)}] H(\theta - \theta_{j-1}) \\
&= 2 \mu^{(1)} \cos \theta + \sum_{j=1}^{k-1} [\mu^{(j)} - \mu^{(j-1)}] [\sin \theta \delta(\theta - \theta_{j-1}) + 2 \cos \theta H(\theta - \theta_{j-1})]
\end{aligned} \tag{15}$$

Defining the dimensionless function

$$\mathbb{G}(\theta; \theta_j) = \sin \theta \delta(\theta - \theta_j) + 2 \cos \theta H(\theta - \theta_j), \tag{16}$$

which is Eq. (34a) in the main manuscript, we can write the field kernel for translational velocity along the symmetry axis in the form

$$\hat{e}_z \cdot \mathbb{K}_t = 2 \mu^{(1)} \cos \theta + \sum_{j=1}^{k-1} [\mu^{(j)} - \mu^{(j-1)}] \mathbb{G}(\theta; \theta_j) \tag{17}$$

which is Eq. (33a) in the main manuscript.

To obtain the flux kernel for motion along the symmetry axis,

$$\hat{e}_z \cdot \boldsymbol{\kappa}_t^{S^+} = \sum_{\ell=0}^{\infty} \sum_{-\ell \leq m \leq \ell} \frac{1}{\ell+1} \{\hat{e}_z \cdot \mathbb{K}_t\}_{\ell m} Y_{\ell, m}, \tag{18}$$

we need to obtain $\{\hat{e}_z \cdot \mathbb{K}_t\}_{\ell m}$, and thus, $\{\cos \theta\}_{\ell m}$ and $\{\mathbb{G}(\theta; \theta_j)\}_{\ell m}$. Using $\cos \theta = A Y_{1,0}(\theta, \phi)$, where $A = 2\sqrt{\pi/3}$, we have

$$\{\cos \theta\}_{\ell m} = \int_0^\pi \int_0^{2\pi} A Y_{1,0} Y_{\ell, m}^*(\theta, \phi) \sin \theta d\theta d\phi = A \delta_{\ell,1} \delta_{m,0}, \tag{19}$$

leading to the first term of the flux kernel [\(18\)](#),

$$\sum_{\substack{\ell=0 \\ -\ell \leq m \leq \ell}}^{\infty} \frac{1}{\ell+1} \{2 \mu^{(1)} \cos \theta\}_{\ell m} Y_{\ell, m} = 2 \mu^{(1)} \sum_{\substack{\ell=0 \\ -\ell \leq m \leq \ell}}^{\infty} \frac{1}{\ell+1} A \delta_{\ell,1} \delta_{m,0} Y_{\ell, m} = 2 \mu^{(1)} \frac{1}{2} A Y_{1,0} = \mu^{(1)} \cos \theta. \tag{20}$$

Next, we calculate $\{\mathbb{G}(\theta; \theta_j)\}_{\ell m}$. Using

$$Y_{\ell, m}(\theta, \phi) = Y_{\ell, m}(\theta, 0) e^{im\phi}, \quad Y_{\ell, m}^*(\theta, \phi) = Y_{\ell, m}(\theta, 0) e^{-im\phi}, \tag{21}$$

for the first term in Eq. (16), we have

$$\begin{aligned}
\{\sin \theta \delta(\theta - \theta_j)\}_{\ell m} &= \int_0^\pi \int_0^{2\pi} \sin \theta \delta(\theta - \theta_j) Y_{\ell, m}^*(\theta, \phi) \sin \theta d\theta d\phi \\
&= \int_0^\pi \int_0^{2\pi} \sin \theta \delta(\theta - \theta_j) Y_{\ell, m}(\theta, 0) e^{-im\phi} \sin \theta d\theta d\phi \\
&= \int_0^{2\pi} e^{-im\phi} d\phi \int_0^\pi \sin^2 \theta \delta(\theta - \theta_j) Y_{\ell, m}(\theta, 0) d\theta \\
&= 2\pi \delta_{m, 0} \sin^2 \theta_j Y_{\ell, m}(\theta_j, 0) \\
&= 2\pi \sin^2 \theta_j Y_{\ell, 0}(\theta_j, 0) \delta_{m, 0}
\end{aligned} \tag{22}$$

For the second term in Eq. (16), we have

$$\begin{aligned}
\{2 \cos \theta H(\theta - \theta_j)\}_{\ell m} &= \int_0^\pi \int_0^{2\pi} 2 \cos \theta H(\theta - \theta_j) Y_{\ell, m}^*(\theta, \phi) \sin \theta d\theta d\phi \\
&= \int_0^\pi \int_0^{2\pi} 2 \cos \theta H(\theta - \theta_j) Y_{\ell, m}(\theta, 0) e^{-im\phi} \sin \theta d\theta d\phi \\
&= \int_0^{2\pi} e^{-im\phi} d\phi \int_0^\pi H(\theta - \theta_j) Y_{\ell, m}(\theta, 0) [2 \cos \theta \sin \theta] d\theta \\
&= 2\pi \delta_{m, 0} \int_0^\pi H(\theta - \theta_j) Y_{\ell, m}(\theta, 0) \sin 2\theta d\theta \\
&= 2\pi \delta_{m, 0} \int_{\theta_j}^\pi Y_{\ell, 0}(\theta, 0) \sin 2\theta d\theta
\end{aligned} \tag{23}$$

Combining Eqns. (22) and (23) yields

$$\{\mathbb{G}(\theta; \theta_j)\}_{\ell m} = \{\mathbb{G}(\theta; \theta_j)\}_{\ell, 0} \delta_{m, 0}, \tag{24}$$

where

$$\{\mathbb{G}(\theta; \theta_j)\}_{\ell, 0} = 2\pi \sin^2 \theta_j Y_{\ell, 0}(\theta_j, 0) + 2\pi \int_{\theta_j}^\pi Y_{\ell, 0}(\theta, 0) \sin 2\theta d\theta. \tag{25}$$

Therefore, we define

$$\begin{aligned}
\mathcal{G}(\theta; \theta_j) &= \sum_{\ell=0}^{\infty} \sum_{-\ell \leq m \leq \ell} \frac{1}{\ell+1} \{\mathbb{G}(\theta; \theta_j)\}_{\ell m} Y_{\ell, m}(\theta, \phi) \\
&= \sum_{\ell=0}^{\infty} \sum_{-\ell \leq m \leq \ell} \frac{1}{\ell+1} \{\mathbb{G}(\theta; \theta_j)\}_{\ell, 0} \delta_{m, 0} Y_{\ell, m}(\theta, \phi) \\
&= \sum_{\ell=0}^{\infty} \frac{1}{\ell+1} \{\mathbb{G}(\theta; \theta_j)\}_{\ell, 0} Y_{\ell, 0}(\theta, 0) \\
&= \sum_{\ell=0}^{\infty} \frac{2\pi}{\ell+1} \left[Y_{\ell, 0}(\theta_j, 0) \sin^2 \theta_j + \int_{\theta_j}^\pi Y_{\ell, 0}(\theta, 0) \sin 2\theta d\theta \right] Y_{\ell, 0}(\theta, 0),
\end{aligned} \tag{26}$$

which is Eq. (34b) in the main manuscript. In the last two lines $Y_{\ell, 0}(\theta, 0) \equiv Y_{\ell, 0}(\theta, \phi)$ is written to

express that the axisymmetric scenario is independent of ϕ . The flux kernel become

$$\hat{\mathbf{e}}_z \cdot \boldsymbol{\kappa}_t^{S^+} = \mu^{(1)} \cos \theta + \sum_{j=1}^{k-1} [\mu^{(j+1)} - \mu^{(j)}] \mathcal{G}(\theta; \theta_j), \quad (27)$$

which is Eq. (33b) in the main manuscript.

3. AXISYMMETRIC JANUS SWIMMER

The distribution of surface flux and phoretic mobility over the surface of the Janus swimmer is given by

$$[\hat{\mathbf{n}} \cdot \boldsymbol{\Gamma}^{S^+}, \mu] = \begin{cases} [\Gamma^{(1)}, \mu^{(1)}] & 0 \leq \theta < \frac{\pi}{2} \\ [\Gamma^{(2)}, \mu^{(2)}] & \frac{\pi}{2} \leq \theta \leq \pi \end{cases} \quad (28)$$

The corresponding flux kernel (27) along the symmetry axis for this problem takes the form

$$\hat{\mathbf{e}}_z \cdot \boldsymbol{\kappa}_t^{S^+} = \mu^{(1)} \cos \theta + [\mu^{(2)} - \mu^{(1)}] \mathcal{G}(\theta; \frac{\pi}{2}). \quad (29)$$

To calculate the translational velocity, we start with the first term

$$\begin{aligned} \int_S \cos \theta \hat{\mathbf{n}} \cdot \boldsymbol{\Gamma}^{S^+} dS &= 2\pi a^2 \Gamma^{(1)} \int_0^{\pi/2} \cos \theta \sin \theta d\theta + 2\pi a^2 \Gamma^{(2)} \int_{\pi/2}^{\pi} \cos \theta \sin \theta d\theta \\ &= 2\pi a^2 \Gamma^{(1)} \times \left(\frac{1}{2}\right) + 2\pi a^2 \Gamma^{(2)} \times \left(\frac{-1}{2}\right) \\ &= \pi a^2 [\Gamma^{(1)} - \Gamma^{(2)}] \end{aligned} \quad (30)$$

For the second term, we can write

$$\mathcal{G}(\theta; \frac{\pi}{2}) = \sum_{\ell=0}^{\infty} \mathcal{G}_{\ell}(\frac{\pi}{2}) Y_{\ell,0}(\theta, 0), \quad (31)$$

where

$$\mathcal{G}_{\ell}(\pi/2) = \frac{2\pi}{\ell+1} \left[Y_{\ell,0}(\frac{\pi}{2}, 0) + \int_{\pi/2}^{\pi} Y_{\ell,0}(\theta, 0) \sin 2\theta d\theta \right]. \quad (32)$$

Therefore, the surface integration of flux weighted by $\mathcal{G}(\theta; \frac{\pi}{2})$ is

$$\begin{aligned} \int_S \mathcal{G}(\theta; \frac{\pi}{2}) \hat{\mathbf{n}} \cdot \boldsymbol{\Gamma}^{S^+} dS &= \sum_{\ell=0}^{\infty} \mathcal{G}_{\ell}(\frac{\pi}{2}) \int_S Y_{\ell,0}(\theta, 0) \hat{\mathbf{n}} \cdot \boldsymbol{\Gamma}^{S^+} dS \\ &= 2\pi a^2 \sum_{\ell=0}^{\infty} \left[\Gamma^{(1)} \mathcal{G}_{\ell}(\frac{\pi}{2}) \int_0^{\pi/2} Y_{\ell,0}(\theta, 0) \sin \theta d\theta + \Gamma^{(2)} \mathcal{G}_{\ell}(\frac{\pi}{2}) \int_{\pi/2}^{\pi} Y_{\ell,0}(\theta, 0) \sin \theta d\theta \right] \\ &= 2\pi a^2 \Gamma^{(1)} \times \left(\frac{1}{4}\right) + 2\pi a^2 \Gamma^{(2)} \times \left(\frac{-1}{4}\right) = \frac{1}{2} \pi a^2 [\Gamma^{(1)} - \Gamma^{(2)}] \end{aligned} \quad (33)$$

ℓ	0	1	2	3	4	5	6	7	8	9	10
$\mathcal{G}_\ell(\pi/2)$	0	$\sqrt{\frac{\pi}{3}}$	$-\frac{\sqrt{5\pi}}{4}$	0	$\frac{\sqrt{\pi}}{4}$	0	$-\frac{3\sqrt{13\pi}}{64}$	0	$\frac{\sqrt{17\pi}}{32}$	0	$-\frac{35\sqrt{\frac{7\pi}{3}}}{512}$
$\int_0^{\pi/2} Y_{\ell,0}(\theta, 0) \sin \theta d\theta$	$\frac{1}{2\sqrt{\pi}}$	$\frac{\sqrt{3/\pi}}{4}$	0	$-\frac{\sqrt{7/\pi}}{16}$	0	$\frac{\sqrt{11/\pi}}{32}$	0	$-\frac{5\sqrt{15/\pi}}{256}$	0	$\frac{7\sqrt{19/\pi}}{512}$	0
$\int_{\pi/2}^\pi Y_{\ell,0}(\theta, 0) \sin \theta d\theta$	$\frac{1}{2\sqrt{\pi}}$	$-\frac{\sqrt{3/\pi}}{4}$	0	$\frac{\sqrt{7/\pi}}{16}$	0	$-\frac{\sqrt{11/\pi}}{32}$	0	$\frac{5\sqrt{15/\pi}}{256}$	0	$-\frac{7\sqrt{19/\pi}}{512}$	0

TABLE I. Terms for the integral involving second term of the flux kernel (29)

where Table II shows that only $\ell = 1$ contributes. Therefore,

$$\begin{aligned}
\hat{\mathbf{e}}_z \cdot \mathbf{U} &= \frac{-1}{4\pi a^2 \mathcal{D}} \int_{S^+} dS \hat{\mathbf{e}}_z \cdot \boldsymbol{\kappa}_t^{S^+} \hat{\mathbf{n}} \cdot \boldsymbol{\Gamma}^{S^+} \\
&= \frac{-1}{4\pi a^2 \mathcal{D}} \int_{S^+} dS \{ \mu^{(1)} \cos \theta + [\mu^{(2)} - \mu^{(1)}] \mathcal{G}(\theta; \frac{\pi}{2}) \} \hat{\mathbf{n}} \cdot \boldsymbol{\Gamma}^{S^+} \\
&= \frac{-1}{4\pi a^2 \mathcal{D}} \mu^{(1)} \int_{S^+} dS \cos \theta \hat{\mathbf{n}} \cdot \boldsymbol{\Gamma}^{S^+} + \frac{-1}{4\pi a^2 \mathcal{D}} [\mu^{(2)} - \mu^{(1)}] \int_{S^+} dS \mathcal{G}(\theta; \frac{\pi}{2}) \hat{\mathbf{n}} \cdot \boldsymbol{\Gamma}^{S^+} \\
&= \frac{-1}{4\pi a^2 \mathcal{D}} \mu^{(1)} \pi a^2 [\Gamma^{(1)} - \Gamma^{(2)}] + \frac{-1}{4\pi a^2 \mathcal{D}} [\mu^{(2)} - \mu^{(1)}] \frac{\pi}{2} a^2 [\Gamma^{(1)} - \Gamma^{(2)}] \\
&= \frac{-1}{4\mathcal{D}} [\Gamma^{(1)} - \Gamma^{(2)}] \left\{ \mu^{(1)} + \frac{1}{2} [\mu^{(2)} - \mu^{(1)}] \right\} \\
&= \frac{-1}{8\mathcal{D}} [\Gamma^{(1)} - \Gamma^{(2)}] [\mu^{(2)} + \mu^{(1)}]
\end{aligned} \tag{34}$$

which is equivalent to Eq. (37) in the main manuscript.

4. A SOURCE-SINK PROBLEM WITH A CONDUCTIVE PASSIVE REGION IN THE MIDDLE

The microswimmer consists of three regions with phoretic mobility and flux distributions,

$$\mu(\theta) = \begin{cases} \mu^{(1)} = \mu^{(a)} & 0 \leq \theta < \theta_1 \\ \mu^{(2)} = \mu^{(p)} & \theta_1 \leq \theta < \theta_2, \\ \mu^{(3)} = \mu^{(a)} & \theta_2 \leq \theta \leq \pi \end{cases}, \quad \hat{\mathbf{n}} \cdot \boldsymbol{\Gamma}^{S^+} = \begin{cases} \Gamma^{(1)} = \Gamma_0 & 0 \leq \theta < \theta_1 \\ \Gamma^{(2)} = 0 & \theta_1 \leq \theta < \theta_2, \\ \Gamma^{(3)} = f(\Gamma_0, \theta_1, \theta_2) & \theta_2 \leq \theta \leq \pi \end{cases}, \tag{35}$$

respectively. The relation $\Gamma^{(3)} = f(\Gamma_0, \theta_1, \theta_2)$ is obtained via the steady-state condition for the source-sink scenario,

$$\begin{aligned}
0 &= \int_S dS \hat{\mathbf{n}} \cdot \boldsymbol{\Gamma}^{S^+} = 2\pi a^2 \int_0^\pi \hat{\mathbf{n}} \cdot \boldsymbol{\Gamma}^{S^+} \sin \theta d\theta \\
&= 2\pi a^2 \Gamma_0 \int_0^{\theta_1} \sin \theta d\theta + 2\pi a^2 \Gamma^{(3)} \int_{\theta_2}^\pi \sin \theta d\theta \\
&= 2\pi a^2 \Gamma_0 (1 - \cos \theta_1) + 2\pi a^2 \Gamma^{(3)} (1 + \cos \theta_2)
\end{aligned} \tag{36}$$

which leads to

$$\Gamma^{(3)} = -\frac{1 - \cos \theta_1}{1 + \cos \theta_2} \Gamma_0. \quad (37)$$

Hence, the translational flux kernel reduces to

$$\begin{aligned} \hat{\mathbf{e}}_z \cdot \boldsymbol{\kappa}_t^{S+} &= \mu^{(1)} \cos \theta + [\mu^{(2)} - \mu^{(1)}] \mathcal{G}(\theta; \theta_1) + [\mu^{(3)} - \mu^{(2)}] \mathcal{G}(\theta; \theta_2) \\ &= \mu^{(a)} \cos \theta + [\mu^{(p)} - \mu^{(a)}] \mathcal{G}(\theta; \theta_1) + [\mu^{(a)} - \mu^{(b)}] \mathcal{G}(\theta; \theta_2) \\ &= \mu^{(a)} \cos \theta + [\mu^{(p)} - \mu^{(a)}] [\mathcal{G}(\theta; \theta_1) - \mathcal{G}(\theta; \theta_2)] \end{aligned} \quad (38)$$

which is Eq. (38) in the main manuscript.

If there was no passive region in the middle and the surface areas of the source and sink were equal, that is, $\theta_1 = \theta_2 = \frac{\pi}{2}$, we would have a Janus sphere, and we consider the microswimmer speed corresponding to this scenario as the reference velocity scale,

$$\begin{aligned} U_{(\text{ref})} &= \frac{-1}{4\pi a^2 \mathcal{D}} \int_S dS \hat{\mathbf{e}}_z \cdot \boldsymbol{\kappa}_t^{S+} \hat{\mathbf{n}} \cdot \boldsymbol{\Gamma}^{S+} \\ &= \frac{-1}{4\pi a^2 \mathcal{D}} (2\pi a^2 \mu^{(a)}) \left[\Gamma_0 \int_0^{\pi/2} \cos \theta \sin \theta d\theta + (-\Gamma_0) \int_{\pi/2}^{\pi} \cos \theta \sin \theta d\theta \right] \\ &= \frac{-\mu^{(a)}}{2\mathcal{D}} \Gamma_0, \end{aligned} \quad (39)$$

which is Eq. (42) in the main manuscript.

Now, for the three-region microswimmer, we start with the uniform phoretic mobility term in the flux kernel (38),

$$\begin{aligned} \frac{-1}{4\pi a^2 \mathcal{D}} \int_S dS \mu^{(a)} \cos \theta \hat{\mathbf{n}} \cdot \boldsymbol{\Gamma}^{S+} &= \frac{-2\pi a^2 \mu^{(a)} \Gamma_0}{4\pi a^2 \mathcal{D}} \left[\int_0^{\theta_1} \cos \theta \sin \theta d\theta - \frac{1 - \cos \theta_1}{1 + \cos \theta_2} \int_{\theta_2}^{\pi} \cos \theta \sin \theta d\theta \right] \\ &= U_{(\text{ref})} \left[\frac{1}{2} \sin^2 \theta_1 - \frac{1 - \cos \theta_1}{1 + \cos \theta_2} \left(-\frac{1}{2} \sin^2 \theta_2 \right) \right] \\ &= U_{(\text{ref})} (2 + \cos \theta_1 - \cos \theta_2) \sin^2 \frac{\theta_1}{2}. \end{aligned} \quad (40)$$

For the term with phoretic mobility difference we have

$$\begin{aligned} \frac{-1}{4\pi a^2 \mathcal{D}} \int_S dS [\mu^{(p)} - \mu^{(a)}] [\mathcal{G}(\theta; \theta_1) - \mathcal{G}(\theta; \theta_2)] \hat{\mathbf{n}} \cdot \boldsymbol{\Gamma}^{S+} \\ &= \frac{-2\pi a^2 \Gamma_0}{4\pi a^2 \mathcal{D}} [\mu^{(p)} - \mu^{(a)}] \int_0^{\theta_1} [\mathcal{G}(\theta; \theta_1) - \mathcal{G}(\theta; \theta_2)] \sin \theta d\theta \\ &\quad - \frac{1 - \cos \theta_1}{1 + \cos \theta_2} \frac{-2\pi a^2 \Gamma_0}{4\pi a^2 \mathcal{D}} [\mu^{(p)} - \mu^{(a)}] \int_{\theta_2}^{\pi} [\mathcal{G}(\theta; \theta_1) - \mathcal{G}(\theta; \theta_2)] \sin \theta d\theta \end{aligned} \quad (41)$$

Using the expression (39) for the reference speed we obtain

$$\begin{aligned}
& \frac{-1}{4\pi a^2 \mathcal{D}} \int_S dS [\mu^{(p)} - \mu^{(a)}] [\mathcal{G}(\theta; \theta_1) - \mathcal{G}(\theta; \theta_2)] \hat{\mathbf{n}} \cdot \mathbf{\Gamma}^{S^+} \\
&= \left(\frac{\mu^{(p)}}{\mu^{(a)}} - 1 \right) U_{(\text{ref})} \int_0^{\theta_1} [\mathcal{G}(\theta; \theta_1) - \mathcal{G}(\theta; \theta_2)] \sin \theta d\theta \\
&\quad - \frac{1 - \cos \theta_1}{1 + \cos \theta_2} \left(\frac{\mu^{(p)}}{\mu^{(a)}} - 1 \right) U_{(\text{ref})} \int_{\theta_2}^{\pi} [\mathcal{G}(\theta; \theta_1) - \mathcal{G}(\theta; \theta_2)] \sin \theta d\theta \\
&= \left(\frac{\mu^{(p)}}{\mu^{(a)}} - 1 \right) U_{(\text{ref})} \left\{ \int_0^{\theta_1} [\mathcal{G}(\theta; \theta_1) - \mathcal{G}(\theta; \theta_2)] \sin \theta d\theta - \frac{1 - \cos \theta_1}{1 + \cos \theta_2} \int_{\theta_2}^{\pi} [\mathcal{G}(\theta; \theta_1) - \mathcal{G}(\theta; \theta_2)] \sin \theta d\theta \right\}
\end{aligned} \tag{42}$$

Therefore, defining the dimensionless velocities,

$$U_{(\text{unif})} = (2 + \cos \theta_1 - \cos \theta_2) \sin^2 \frac{\theta_1}{2} \tag{43a}$$

$$\hat{U} = \int_0^{\theta_1} d\theta \sin \theta [\mathcal{G}(\theta; \theta_1) - \mathcal{G}(\theta; \theta_2)] - \frac{1 - \cos \theta_1}{1 + \cos \theta_2} \int_{\theta_2}^{\pi} d\theta \sin \theta [\mathcal{G}(\theta; \theta_1) - \mathcal{G}(\theta; \theta_2)], \tag{43b}$$

which are Eq. (41a) and Eq. (41b) in the main manuscript, respectively, the translational velocity along the symmetry axis becomes

$$\frac{\hat{\mathbf{e}}_z \cdot \mathbf{U}}{U_{(\text{ref})}} = U_{(\text{unif})} + \left(\frac{\mu^{(p)}}{\mu^{(a)}} - 1 \right) \hat{U}, \tag{44}$$

which is Eq. (40) in the main manuscript.

5. NON-AXISYMMETRIC SELF-PHORETIC PARTICLE

In this scenario, the position-dependent phoretic mobility $\mu(\theta)$ is given by

$$\mu(\theta) = \mu^{(a)} + [\mu^{(p)} - \mu^{(a)}] H(\theta - \theta_0), \tag{45}$$

and the normal component of the flux $\hat{\mathbf{n}} \cdot \mathbf{\Gamma}^{S^+}$ over the surface of the sphere is given by

$$\hat{\mathbf{n}} \cdot \mathbf{\Gamma}^{S^+}(\theta, \phi) = [1 - H(\theta - \theta_0)][1 - 2H(\phi - \pi)] \Gamma_0. \tag{46}$$

We aim to obtain the translational $\mathbf{U}$ and rotational $\mathbf{\Omega}$ velocities of the self-phoretic particle.

A. Translational Velocity

The particle's translational velocity is along the y -axis, with the field kernel

$$\hat{\mathbf{e}}_y \cdot \mathbb{K}_t = -a \hat{\mathbf{e}}_y \cdot \nabla_S \mu + 2 \hat{\mathbf{e}}_y \cdot \hat{\mathbf{n}} \mu = -\hat{\mathbf{e}}_y \cdot \hat{\mathbf{e}}_\theta \partial_\theta \mu + 2 \hat{\mathbf{e}}_y \cdot \hat{\mathbf{e}}_r \mu = \sin \phi (-\cos \theta \partial_\theta \mu + 2 \mu \sin \theta) \tag{47}$$

where we have used $\hat{\mathbf{e}}_y \cdot \hat{\mathbf{e}}_\theta = \cos \theta \sin \phi$ and $\hat{\mathbf{e}}_y \cdot \hat{\mathbf{e}}_r = \sin \theta \sin \phi$. Using the expression for the phoretic mobility (45) gives

$$\begin{aligned}\hat{\mathbf{e}}_y \cdot \mathbb{K}_t &= \sin \phi \left\{ -\cos \theta \left[\mu^{(p)} - \mu^{(a)} \right] \delta(\theta - \theta_0) + 2 \left(\mu^{(a)} + \left[\mu^{(p)} - \mu^{(a)} \right] H(\theta - \theta_0) \right) \sin \theta \right\} \\ &= \sin \phi \left\{ 2\mu^{(a)} \sin \theta + \left[\mu^{(p)} - \mu^{(a)} \right] [2H(\theta - \theta_0) \sin \theta - \delta(\theta - \theta_0) \cos \theta] \right\}\end{aligned}\quad (48)$$

and thus, we can write the field kernel for translation along the y -axis as

$$\hat{\mathbf{e}}_y \cdot \mathbb{K}_t = \sin \phi \left\{ \mu^{(a)} (2 \sin \theta) + \left[\mu^{(p)} - \mu^{(a)} \right] \Theta_t^\Psi \right\} = f^\Psi(\theta) \sin \phi \quad (49)$$

$$\Theta_t^\Psi(\theta; \theta_0) = 2H(\theta - \theta_0) \sin \theta - \delta(\theta - \theta_0) \cos \theta \quad (50)$$

which are Eq. (47a) and Eq. (47b) in the main manuscript, respectively.

To obtain the translational flux kernel we need to calculate $\{\hat{\mathbf{e}}_y \cdot \mathbb{K}_t\}_{\ell m}$. The dependence of the field kernel (49) on $\sin \phi$ necessitates that the flux kernel coefficients $\{\hat{\mathbf{e}}_y \cdot \mathbb{K}_t\}_{\ell m}$ can be non-zero only when $m = \pm 1$, and thus $\ell \neq 0$. Using

$$\{\hat{\mathbf{e}}_y \cdot \mathbb{K}_t\}_{\ell, 1}^* = \int_0^{2\pi} \int_0^\pi \hat{\mathbf{e}}_y \cdot \mathbb{K}_t Y_{\ell, 1} \sin \theta d\theta d\phi = \int_0^{2\pi} \int_0^\pi \hat{\mathbf{e}}_y \cdot \mathbb{K}_t (-Y_{\ell, -1}^*) \sin \theta d\theta d\phi = -\{\hat{\mathbf{e}}_y \cdot \mathbb{K}_t\}_{\ell, -1} \quad (51)$$

and $Y_{\ell, -1}(\theta, \phi) = -Y_{\ell, 1}^*(\theta, \phi) = -Y_{\ell, 1}(\theta, 0)e^{-i\phi}$ we can write

$$\begin{aligned}\hat{\mathbf{e}}_y \cdot \mathcal{K}_t^{S+} &= \sum_{\ell=0}^{\infty} \sum_{-\ell \leq m \leq \ell} \frac{1}{\ell+1} \{\hat{\mathbf{e}}_y \cdot \mathbb{K}_t\}_{\ell m} Y_{\ell, m}(\theta, \phi) \\ &= \sum_{\ell=1}^{\infty} \frac{1}{\ell+1} \left[\{\hat{\mathbf{e}}_y \cdot \mathbb{K}_t\}_{\ell, -1} Y_{\ell, -1}(\theta, \phi) + \{\hat{\mathbf{e}}_y \cdot \mathbb{K}_t\}_{\ell, 1} Y_{\ell, 1}(\theta, \phi) \right] \\ &= \sum_{\ell=1}^{\infty} \frac{1}{\ell+1} \left[\{\hat{\mathbf{e}}_y \cdot \mathbb{K}_t\}_{\ell, 1} e^{i\phi} - \{\hat{\mathbf{e}}_y \cdot \mathbb{K}_t\}_{\ell, -1} e^{-i\phi} \right] Y_{\ell, 1}(\theta, 0) \\ &= \sum_{\ell=1}^{\infty} \frac{1}{\ell+1} \left[\{\hat{\mathbf{e}}_y \cdot \mathbb{K}_t\}_{\ell, 1} e^{i\phi} + \{\hat{\mathbf{e}}_y \cdot \mathbb{K}_t\}_{\ell, 1}^* e^{-i\phi} \right] Y_{\ell, 1}(\theta, 0) \\ &= \sum_{\ell=1}^{\infty} \frac{1}{\ell+1} 2\text{Re} \left[\{\hat{\mathbf{e}}_y \cdot \mathbb{K}_t\}_{\ell, 1} e^{i\phi} \right] Y_{\ell, 1}(\theta, 0)\end{aligned}\quad (52)$$

where $\text{Re}[\cdot]$ returns the real part of the complex number. Evaluating

$$\begin{aligned}e^{i\phi} \{\hat{\mathbf{e}}_y \cdot \mathbb{K}_t\}_{\ell, 1} &= e^{i\phi} \{f^\Psi(\theta) \sin \phi\}_{\ell, 1} \\ &= e^{i\phi} \int_0^{2\pi} \int_0^\pi f^\Psi(\theta') \sin \phi' Y_{\ell, 1}^*(\theta', \phi') \sin \theta' d\theta' d\phi' \\ &= e^{i\phi} \int_0^{2\pi} e^{-i\phi'} \sin \phi' d\phi' \int_0^\pi f^\Psi(\theta') Y_{\ell, 1}(\theta', 0) \sin \theta' d\theta' \\ &= (\cos \phi + i \sin \phi)(-i\pi) \int_0^\pi f^\Psi(\theta') Y_{\ell, 1}(\theta', 0) \sin \theta' d\theta' \\ &= \pi(-i \cos \phi + \sin \phi) \int_0^\pi f^\Psi(\theta') Y_{\ell, 1}(\theta', 0) \sin \theta' d\theta'\end{aligned}\quad (53)$$

and taking its real part, combined with Eq. (52), leads to

$$\hat{\mathbf{e}}_y \cdot \boldsymbol{\kappa}_t^{S^+} = \sin \phi \sum_{\ell=1}^{\infty} \frac{2\pi}{\ell+1} Y_{\ell,1}(\theta, 0) \int_0^{\pi} f^{\Psi}(\theta') Y_{\ell,1}(\theta', 0) \sin \theta' d\theta' \quad (54)$$

and we are left with the calculation of the integral. For the uniform phoretic mobility term $\mu^{(a)}(2 \sin \theta)$, using $\sin \theta = BY_{1,1}(\theta, 0)$, where $B = -2\sqrt{2\pi/3}$, we have

$$\begin{aligned} Y_{\ell,1}(\theta, 0) \int_0^{\pi} [\mu^{(a)}(2 \sin \theta')] Y_{\ell,1}(\theta', 0) \sin \theta' d\theta' &= 2\mu^{(a)} Y_{\ell,1}(\theta, 0) \int_0^{\pi} BY_{1,1}(\theta', 0) Y_{\ell,1}(\theta', 0) \sin \theta' d\theta' \\ &= 2\mu^{(a)} Y_{\ell,1}(\theta, 0) \left(\frac{B}{2\pi} \delta_{\ell,1} \right) \\ &= 2 \frac{1}{2\pi} \mu^{(a)} BY_{1,1}(\theta, 0) \delta_{\ell,1} \\ &= \frac{1}{\pi} \mu^{(a)} \sin \theta \delta_{\ell,1} \end{aligned} \quad (55)$$

and thus,

$$\sum_{\ell=1}^{\infty} \frac{2\pi}{\ell+1} Y_{\ell,1}(\theta, 0) \int_0^{\pi} [\mu^{(a)}(2 \sin \theta')] Y_{\ell,1}(\theta', 0) \sin \theta' d\theta' = \sum_{\ell=1}^{\infty} \frac{2\pi}{\ell+1} \frac{1}{\pi} \mu^{(a)} \sin \theta \delta_{\ell,1} = \mu^{(a)} \sin \theta \quad (56)$$

For the term corresponding to the difference of the phoretic mobilities we need to use direct integration. Therefore, defining

$$\Theta_t^{\Gamma}(\theta; \theta_0) = \sum_{l=1}^{\infty} \frac{2\pi}{l+1} \left\{ 2 \int_{\theta_0}^{\pi} Y_{l,1}(\theta', 0) (\sin \theta')^2 d\theta' - \sin \theta_0 \cos \theta_0 Y_{l,1}(\theta_0, 0) \right\} Y_{l,1}(\theta, 0) \quad (57)$$

which is Eq. (48b) in the main manuscript, we can write the y -component of the flux kernel in the form,

$$\hat{\mathbf{e}}_y \cdot \boldsymbol{\kappa}_t^{S^+} = \sin \phi \left\{ \mu^{(a)} \sin \theta + [\mu^{(p)} - \mu^{(a)}] \Theta_t^{\Gamma}(\theta; \theta_0) \right\} = f^{\Gamma}(\theta) \sin \phi \quad (58)$$

which is Eq. (48a) in the main manuscript.

Therefore, we can obtain the translational velocity,

$$\begin{aligned}
\hat{\mathbf{e}}_y \cdot \mathbf{U} &= \frac{-1}{4\pi a^2 \mathcal{D}} \int_S dS \hat{\mathbf{e}}_y \cdot \boldsymbol{\kappa}_t^{S^+} \hat{\mathbf{n}} \cdot \boldsymbol{\Gamma}^{S^+} \\
&= \frac{-1}{4\pi a^2 \mathcal{D}} a^2 \int_0^{2\pi} \int_0^\pi \sin \theta d\theta d\phi [f^\Gamma(\theta) \sin \phi] \{[1 - H(\theta - \theta_0)][1 - 2H(\phi - \pi)] \Gamma_0\} \\
&= \frac{-\Gamma_0}{4\pi \mathcal{D}} \int_0^{2\pi} [1 - 2H(\phi - \pi)] \sin \phi d\phi \int_0^\pi \sin \theta d\theta f^\Gamma(\theta) [1 - H(\theta - \theta_0)] \\
&= \frac{-\Gamma_0}{4\pi \mathcal{D}} \times 4 \times \int_0^{\theta_0} \sin \theta d\theta f^\Gamma(\theta) \\
&= \frac{U_{(\text{ref})}}{\mu^{(a)}} \frac{2}{\pi} \int_0^{\theta_0} \sin \theta d\theta f^\Gamma(\theta) \\
&= \frac{U_{(\text{ref})}}{\mu^{(a)}} \frac{2}{\pi} \int_0^{\theta_0} \sin \theta d\theta \{ \mu^{(a)} \sin \theta + [\mu^{(p)} - \mu^{(a)}] \Theta_t^\Gamma(\theta; \theta_0) \} \\
&= U_{(\text{ref})} \frac{2}{\pi} \int_0^{\theta_0} \sin^2 \theta d\theta + U_{(\text{ref})} \left[\frac{\mu^{(p)}}{\mu^{(a)}} - 1 \right] \frac{2}{\pi} \int_0^{\theta_0} \sin \theta d\theta \Theta_t^\Gamma(\theta; \theta_0) \\
&= U_{(\text{ref})} \frac{2}{\pi} \frac{1}{2} \left(\theta_0 - \frac{1}{2} \sin 2\theta \right) + U_{(\text{ref})} \left[\frac{\mu^{(p)}}{\mu^{(a)}} - 1 \right] \frac{2}{\pi} \int_0^{\theta_0} \sin \theta d\theta \Theta_t^\Gamma(\theta; \theta_0) \\
&= U_{(\text{ref})} \left(\frac{\theta_0}{\pi} - \frac{\sin 2\theta}{2\pi} \right) + U_{(\text{ref})} \left[\frac{\mu^{(p)}}{\mu^{(a)}} - 1 \right] \frac{2}{\pi} \int_0^{\theta_0} \sin \theta d\theta \Theta_t^\Gamma(\theta; \theta_0) \tag{59}
\end{aligned}$$

Therefore, defining the dimensionless velocities,

$$U_{(\text{unif})} = \frac{\theta_0}{\pi} - \frac{\sin 2\theta_0}{2\pi}, \tag{60}$$

$$\hat{U} = \frac{2}{\pi} \int_0^{\theta_0} d\theta \sin \theta \Theta_t^\Gamma(\theta; \theta_0), \tag{61}$$

which are Eq. (50a) and Eq. (50b) in the main manuscript, respectively, we have the scaled translational velocity along the y direction,

$$\frac{\hat{\mathbf{e}}_y \cdot \mathbf{U}}{U_{(\text{ref})}} = U_{(\text{unif})} + \left(\frac{\mu^{(p)}}{\mu^{(a)}} - 1 \right) \hat{U}, \tag{62}$$

which is Eq. (49) in the main manuscript.

B. Rotational Velocity

For the rotational field kernel, we have

$$\mathbb{K}_r = -\hat{\mathbf{n}} \times (a \nabla_S \mu) = -\hat{\mathbf{e}}_r \times \hat{\mathbf{e}}_\theta \partial_\theta \mu = -\hat{\mathbf{e}}_\phi \partial_\theta \mu = -\hat{\mathbf{e}}_\phi [\mu^{(p)} - \mu^{(a)}] \delta(\theta - \theta_0) \tag{63}$$

which is Eq. (51a) in the main manuscript. Since the only ϕ dependence is in $\hat{\mathbf{e}}_\phi$, only $\{\mathbb{K}_r\}_{\ell,\pm 1}$ terms can be non-zero, and we have

$$\mathcal{K}_r^{S+} = \sum_{\ell=0}^{\infty} \sum_{-\ell \leq m \leq \ell} \frac{1}{\ell+1} \{\mathbb{K}_r\}_{\ell m} Y_{\ell,m}(\theta, \phi) = \sum_{\ell=1}^{\infty} \frac{1}{\ell+1} 2\text{Re} \left[\{\mathbb{K}_r\}_{\ell,1} e^{i\phi} \right] Y_{\ell,1}(\theta, 0) \quad (64)$$

Evaluating

$$\begin{aligned} \{\hat{\mathbf{e}}_\phi \delta(\theta - \theta_0)\}_{\ell,1} &= \int_0^\pi \int_0^{2\pi} \hat{\mathbf{e}}_\phi \delta(\theta - \theta_0) Y_{\ell,1}^*(\theta, \phi) \sin \theta d\theta d\phi \\ &= \int_0^{2\pi} e^{-i\phi} \hat{\mathbf{e}}_\phi d\phi \int_0^\pi \delta(\theta - \theta_0) Y_{\ell,1}(\theta, 0) \sin \theta d\theta \\ &= \int_0^{2\pi} e^{-i\phi} (-\sin \phi \hat{\mathbf{e}}_x + \cos \phi \hat{\mathbf{e}}_y) d\phi \int_0^\pi \delta(\theta - \theta_0) Y_{\ell,1}(\theta, 0) \sin \theta d\theta \\ &= \pi (i\hat{\mathbf{e}}_x + \hat{\mathbf{e}}_y) Y_{\ell,1}(\theta_0, 0) \sin \theta_0 \end{aligned} \quad (65)$$

and taking its real part leads to

$$\begin{aligned} 2\text{Re} \left[\{\mathbb{K}_r\}_{\ell,1} e^{i\phi} \right] &= -2\pi [\mu^{(p)} - \mu^{(a)}] Y_{\ell,1}(\theta_0, 0) \sin \theta_0 \text{Re} \left[(i\hat{\mathbf{e}}_x + \hat{\mathbf{e}}_y) e^{i\phi} \right] \\ &= -2\pi [\mu^{(p)} - \mu^{(a)}] Y_{\ell,1}(\theta_0, 0) \sin \theta_0 (-\sin \phi \hat{\mathbf{e}}_x + \cos \phi \hat{\mathbf{e}}_y) \\ &= -2\pi [\mu^{(p)} - \mu^{(a)}] Y_{\ell,1}(\theta_0, 0) \sin \theta_0 \hat{\mathbf{e}}_\phi \end{aligned} \quad (66)$$

and thus,

$$\mathcal{K}_r^{S+} = -\hat{\mathbf{e}}_\phi [\mu^{(p)} - \mu^{(a)}] \sum_{\ell=1}^{\infty} \frac{2\pi}{\ell+1} Y_{\ell,1}(\theta_0, 0) \sin \theta_0 Y_{\ell,1}(\theta, 0) \quad (67)$$

Thus, defining

$$\Theta_r^\Gamma(\theta; \theta_0) = 2\pi \sum_{\ell=1}^{\infty} Y_{\ell,1}(\theta_0, 0) Y_{\ell,1}(\theta, 0) \left[\frac{\sin \theta_0}{\ell+1} \right] \quad (68)$$

which is Eq. (52b) in the main manuscript, we have

$$\mathcal{K}_r^{S+} = -\hat{\mathbf{e}}_\phi [\mu^{(p)} - \mu^{(a)}] \Theta_r^\Gamma(\theta; \theta_0). \quad (69)$$

which is Eq. (51b) in the main manuscript.

Calculating the angular velocity is straightforward,

$$\begin{aligned}
\mathbf{\Omega} &= \frac{-3}{8\pi a^3 \mathcal{D}} \int_{S^+} dS \mathbf{\kappa}_r^{S^+} \cdot \hat{\mathbf{n}} \cdot \mathbf{\Gamma}^{S^+} \\
&= \frac{-3}{8\pi a^3 \mathcal{D}} [\mu^{(p)} - \mu^{(a)}] a^2 \int_0^{2\pi} \int_0^\pi \sin \theta d\theta d\phi [-\hat{\mathbf{e}}_\phi \Theta_r^\Gamma(\theta; \theta_0)] \{[1 - H(\theta - \theta_0)][1 - 2H(\phi - \pi)] \Gamma_0\} \\
&= \frac{-3\Gamma_0}{8\pi a \mathcal{D}} [\mu^{(p)} - \mu^{(a)}] \int_0^{2\pi} [1 - 2H(\phi - \pi)] (-\hat{\mathbf{e}}_\phi) d\phi \int_0^\pi \sin \theta d\theta \Theta_r^\Gamma(\theta; \theta_0) [1 - H(\theta - \theta_0)] \\
&= \frac{-3\Gamma_0}{8\pi a \mathcal{D}} [\mu^{(p)} - \mu^{(a)}] \times 4\hat{\mathbf{e}}_x \times \int_0^{\theta_0} \sin \theta d\theta \Theta_r^\Gamma(\theta; \theta_0) \\
&= \hat{\mathbf{e}}_x \frac{U_{(\text{ref})}}{a} \left[\frac{\mu^{(p)}}{\mu^{(a)}} - 1 \right] \frac{3}{\pi} \int_0^{\theta_0} \sin \theta d\theta \Theta_r^\Gamma(\theta; \theta_0)
\end{aligned} \tag{70}$$

Therefore, defining the reference angular velocity scale as

$$\Omega_{(\text{ref})} = U_{(\text{ref})}/a, \tag{71}$$

the scaled angular velocity as

$$\frac{\mathbf{\Omega}}{\Omega_{(\text{ref})}} = \hat{\mathbf{e}}_x \left(\frac{\mu^{(p)}}{\mu^{(a)}} - 1 \right) \hat{\Omega}(\theta_0), \tag{72}$$

which is Eq. (54) in the main manuscript. The dimensionless function is

$$\hat{\Omega}(\theta_0) = \frac{3}{\pi} \int_0^{\theta_0} d\theta \sin \theta \Theta_r^\Gamma(\theta; \theta_0), \tag{73}$$

which is Eq. (55) in the main manuscript.