

Precision mass measurements of $^{74-76}\text{Sr}$ using the multiple reflection time-of-flight mass spectrometer at TITAN

Z. Hockenberry,^{1,2} T. Murböck,^{1,3} B. Ashrafkhani,^{1,4} J. Bergmann,^{1,5} C. Brown,^{1,5} T. Brunner,^{1,6} J. Cardona,^{1,6} T. Dickel,^{1,7} E. Dunling,¹ J. D. Holt,^{1,2} C. Hornung,⁷ B. S. Hu,^{8,9} C. Izzo,¹⁰ A. Jacobs,^{1,11} A. Javaji,^{1,6} S. Kakkar,^{1,6} B. Koote,^{1,6} G. Kripko-Koncz,^{1,12} Ali Mollaebrahimi,^{1,7} D. Lascar,^{1,13} E. M. Lykiardopoulou,^{1,14} I. Mukul,¹ S. F. Paul,¹ W. R. Plaß,^{1,7} W. S. Porter,^{1,14,*} M. P. Reiter,^{1,5} J. Ringuette,^{1,15} H. Schatz,^{1,10,16,17} C. Scheidenberger,^{3,7,18} C. Walls,¹ Y. Wang,^{1,14} and A. A. Kwiatkowski,^{1,19}

¹*TRIUMF, 4004 Wesbrook Mall, Vancouver, British Columbia, Canada V6T 2A3*

²*Department of Physics, McGill University, 3600 Rue University, Montréal, Québec, Canada H3A 2T8*

³*II. Physikalisches Institut, Justus-Liebig-Universität Gießen, 35392 Gießen, Germany*

⁴*Department of Physics and Astronomy, University of Calgary, Calgary, Alberta, Canada T2N 1N4*

⁵*School of Physics and Astronomy, University of Edinburgh, Edinburgh EH9 3FD, United Kingdom*

⁶*Department of Physics and Astronomy, University of Manitoba, Winnipeg, Manitoba, Canada R3T 2N2*

⁷*GSI Helmholtzzentrum für Schwerionenforschung GmbH, Planckstraße 1, 64291 Darmstadt, Germany*

⁸*National Center for Computational Sciences, Oak Ridge National Laboratory, Oak Ridge, Tennessee 37831, USA*

⁹*Physics Division, Oak Ridge National Laboratory, Oak Ridge, Tennessee 37831, USA*

¹⁰*Facility for Rare Isotope Beams, East Lansing, Michigan 48859, USA*

¹¹*Argonne National Laboratory, Lemont, Illinois 60439, USA*

¹²*Helmholtz Research Academy Hesse for FAIR (HFHF), GSI Helmholtz Center for Heavy Ion Research, Campus Gießen, 35392 Gießen, Germany*

¹³*Center for Fundamental Physics, Northwestern University, Evanston, Illinois 60208, USA*

¹⁴*Department of Physics and Astronomy, University of British Columbia, Vancouver, British Columbia, Canada V6T 1Z1*

¹⁵*Department of Physics, Colorado School of Mines, Golden, Colorado 80401, USA*

¹⁶*Department of Physics and Astronomy, Michigan State University, East Lansing, Michigan 48824, USA*

¹⁷*Joint Institute for Nuclear Astrophysics Center for the Evolution of the Elements (JINA-CEE), East Lansing, Michigan 48854, USA*

¹⁸*Helmholtz Forschungssakademie Hessen für FAIR (HFHF), GSI Helmholtzzentrum für Schwerionenforschung, Campus Gießen, 35392 Gießen, Germany*

¹⁹*Department of Physics and Astronomy, University of Victoria, Victoria, British Columbia, Canada V8P 5C2*



(Received 10 August 2024; accepted 13 December 2024; published 23 January 2025)

We report precision mass measurements of $^{74-76}\text{Sr}$ performed with the TITAN multiple-reflection time-of-flight mass spectrometer. This marks a first-time mass measurement of ^{74}Sr and gives increased mass precision to both ^{75}Sr and ^{76}Sr , which were previously measured using storage ring and Penning trap methods, respectively. This completes the $A = 74$, $T = 1$ isospin triplet and gives increased precision to the $A = 75$, $T = 1/2$ isospin doublet, which are both the heaviest experimentally evaluated triplets and doublets to date. The new data allow us to evaluate coefficients of the isobaric multiplet mass equation for the first time at $A = 74$ and with increased precision at $A = 75$. With increased precision of ^{75}Sr , we confirm the recent measurement reported by CSRe that was used to remove a staggering anomaly in the doublets. New *ab initio* valence-space in-medium similarity renormalization group calculations of the $T = 1$ triplet are presented at $A = 74$. We also investigate the impact of the new mass data on the reaction flow of the rapid proton capture process in type I x-ray bursts using a single-zone model.

DOI: [10.1103/PhysRevC.111.014327](https://doi.org/10.1103/PhysRevC.111.014327)

I. INTRODUCTION

Isospin symmetry is based on the observation that protons (p) and neutrons (n) have the same spin and nearly identical masses [1], and they exhibit very similar pp , nn ,

and pn scattering lengths [2]. Within the isospin framework, protons and neutrons are described as particles distinguished by their isospin projection T_z , analogous to projections of angular momentum. Based on this symmetry, one would assume that isobaric nuclei that share the same nucleon number $A = N + Z$ but differ in their neutron (N) and proton (Z) numbers, exhibit similar structure. This is indeed true, as the study of nuclear isobars mirrored about the $N = Z$ line is an important avenue for understanding isospin symmetry in nu-

*Present address: University of Notre Dame, Notre Dame, Indiana 46556, USA.

clei [3]. Isospin is, however, broken by the difference in electric charge and mass between protons and neutrons and by a small contribution from isospin-violating interactions of nuclear origin [4,5]. Understanding exactly how these isospin-violating interactions manifest in the structure of nuclei near $N = Z$ has relevance to unitarity tests of the Cabibbo-Kobayashi-Maskawa (CKM) quark-mixing matrix [6,7], and the rapid proton capture process (rp-process) in type I x-ray bursts [8–10].

A method for studying the nature of isospin symmetry breaking near $N = Z$ is the isobaric multiplet mass equation (IMME) [11]. An isobaric multiplet is a set of nuclear states with the same isospin T and spin-parity J^π , but with differing isospin projections $T_z = (N - Z)/2$. As long as isospin-violating interactions can be expressed as two-body forces of a tensor of rank 2, the IMME states that the mass difference between these isobaric analog states follows a quadratic form:

$$ME(\alpha, T, T_z) = a(\alpha, T) + b(\alpha, T)T_z + c(\alpha, T)T_z^2, \quad (1)$$

where ME is the mass excess and the other quantum numbers of the system are encapsulated in α . The coefficients a , b , and c are empirically determined through fitting isobaric analog state data deduced from measurements of nuclear masses and level energies [11]. The most recent extensive evaluation of the IMME was performed by MacCormick and Audi who used experimental mass and nuclear level structure data in the $A = 10$ –60 region to evaluate multiplets with $T = 1/2$ up to 3 [11].

Attempts to theoretically reproduce the $T = 1/2$ doublet b coefficients have shown mixed results. These b coefficients are extracted from ground-state mass data of odd- A nuclei and their global trend shows a characteristic staggering pattern between the $A = 4n + 1$ and $A = 4n + 3$ nuclei (where n is a positive integer). This staggering pattern has been explained as being due to an increase in Coulomb repulsion between spin antialigned protons [12–14]. Evaluation of the $A = 69$ –75 doublets based on mass data from the 2020 Atomic Mass Evaluation (AME2020) [15] indicated a confounding departure from the regular staggering pattern. This staggering anomaly was first noted by Kaneko *et al.* [16], who were unable to reproduce it using a modern shell-model approach with inclusion of Coulomb and isospin nonconserving nuclear forces. Theoretical approaches such as density functional theory (DFT) [17] and valence-space in-medium similarity renormalization group (VS-IMSRG) [18] were also unsuccessful in reproducing this anomaly. Further discussion of the anomaly is given by Tu *et al.* [19]. It was recently shown by Li *et al.* [20] that the increased mass precision of ^{71}Kr and ^{75}Sr reported by Wang *et al.* [21] removes this staggering anomaly. Wang *et al.* [21] only reported four counts of ^{75}Sr during their experiment; therefore, independent confirmation of the mass would strengthen the result of the rectified staggering phase.

Isospin symmetry breaking in nuclei plays an important role in testing CKM matrix unitarity via precise determination of the up-down element V_{ud} [7]. Extractions of V_{ud} rely on global comparisons of superallowed β -decay $\mathcal{F}t$ values. These $\mathcal{F}t$ values are determined from experimentally deduced ft values that are corrected to be nucleus-independent using

theoretical corrections for radiative and isospin symmetry-breaking effects. Recent re-evaluations of the radiative corrections have led to an increased deviation in CKM matrix unitarity [22–24], thus leaving the isospin symmetry-breaking corrections as the leading contributor to uncertainty in V_{ud} [6]. Precise extraction of the isospin symmetry-breaking corrections is, therefore, a major task for this field to progress.

While isospin symmetry-breaking corrections for superallowed β decays have so far been primarily extracted via shell-model calculations [7], recent progress in *ab initio* theory has prompted an interest in using alternative methods such as VS-IMSRG theory [25,26] and the no-core shell model [27]. *Ab initio* methods hold the key promise of providing a more robust quantification of isospin symmetry-breaking corrections; however, careful benchmarking against experimental data is required to make sure the predicted symmetry-breaking correction factors meet accuracy demands for CKM unitarity tests.

Comparisons between theoretical and experimental IMME coefficients provide an ideal test bed to probe the reliability of theoretical nuclear forces and the isospin symmetry-breaking corrections determined from them. Recent VS-IMSRG calculations of the triplet c coefficients of all $A = 4n + 2$ nuclei in the $A = 10$ –74 range agreed with experimental coefficients within 250 keV, but lacked the precision to improve upon available isospin symmetry-breaking corrections [25]. To obtain improved correction terms, refinements of theory guided by further benchmarking against experimental data are needed.

This article reports on precision mass measurements of neutron-deficient 74 – ^{76}Sr . Our measurement of ^{75}Sr provides independent confirmation of the storage ring measurement by Wang *et al.* [21], supporting the removal of the staggering anomaly reported by Li *et al.* [20]. The new ^{74}Sr mass value allows us to extend the triplet coefficients into otherwise unexplored regions of the nuclear chart. Multishell VS-IMSRG calculations [28–30] at $A = 74$ were performed using two new χ EFT-derived interactions to explore the chiral nuclear force. We also investigated the impact of the new strontium masses on the rp-process using a single-zone type-I x-ray burst model [31].

II. EXPERIMENT

The experiment took place at TRIUMF’s Ion Trap for Atomic and Nuclear science (TITAN) [32], an experimental installation for precision mass measurements and in-trap nuclear-decay spectroscopy. The measurements were performed using TITAN’s multiple-reflection time-of-flight mass spectrometer (MR-ToF-MS) [33,34], which provided the means to overcome the low production cross sections and high in-beam contamination typical for this region of the nuclear chart. The primary radioactive ion beam (RIB) was produced using the isotope separation on-line method [35] at TRIUMF’s Isotope Separator and Accelerator (ISAC) facility [36]. A 50- μA , 480-MeV p^+ beam was impinged onto a niobium target to produce the neutron-deficient strontium isotopes of interest. To enhance the signal-to-background ratio, the strontium isotopes were selectively ionized by TRIUMF’s Resonant

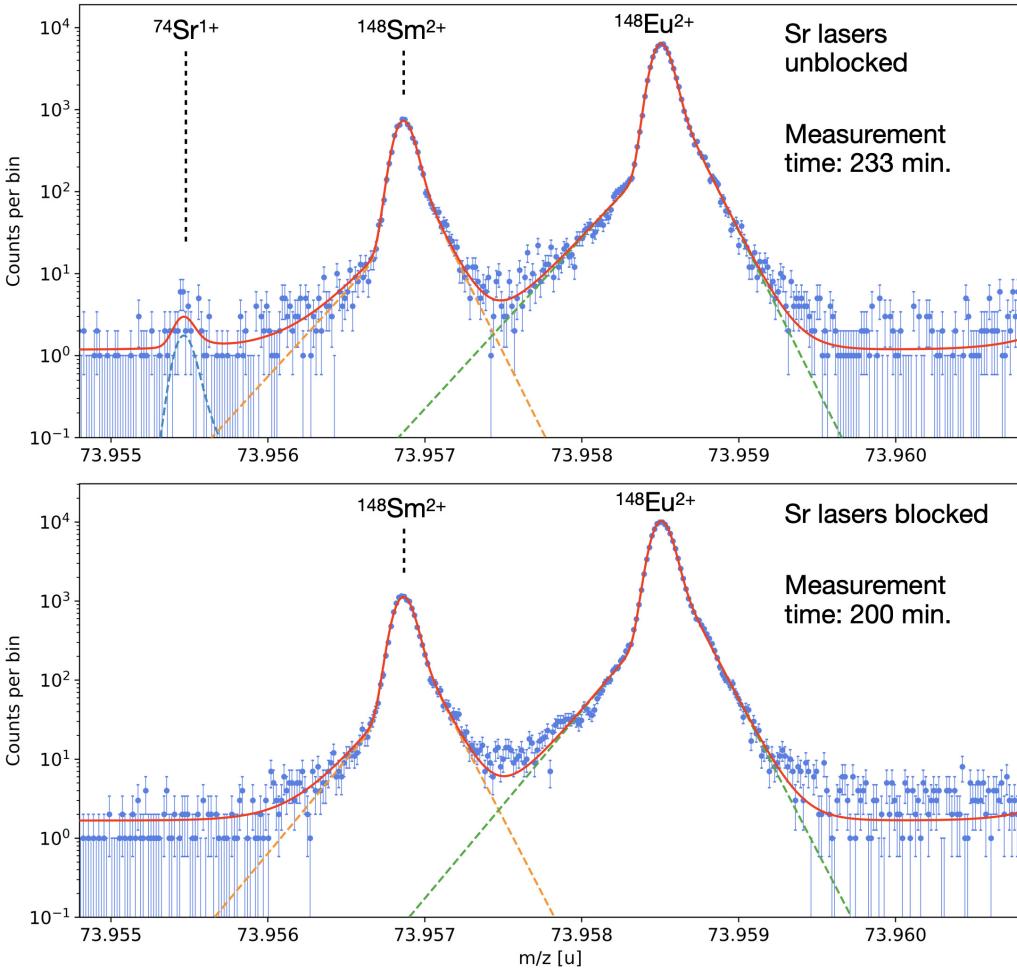


FIG. 1. Mass spectra obtained at $A = 74$. The spectra have been drift corrected using the time-resolved calibration procedure and fitted using hyper-EMG peak shapes. Error bars are $\pm 1\sigma$ assuming Poisson statistics in each bin. Multipeak fits to the data (red line) and the underlying single peak fits (dashed lines) are indicated. Unambiguous identification of ^{74}Sr is demonstrated with the TRILIS lasers, which were tuned to selectively ionize only strontium.

Ionization Laser Ion Source (TRILIS) using a two-step excitation scheme into an autoionizing state [37,38]. Following production, ISAC's magnetic mass separator system [36] filtered the RIB down to a range of a single mass unit using a mass separation power of $\frac{m}{\Delta m} \approx 2000$. The continuous RIB was guided at 20 keV into TITAN's helium-filled radio-frequency quadrupole (RFQ) trap [39] for beam cooling and bunching. The prepared RIB bunches were lowered to a transport energy of 1.3 keV using a pulsed drift tube and supplied at a repetition rate of 50 Hz to the buffer-gas-filled RFQ ion transport section of the MR-ToF-MS [40].

The TITAN MR-ToF-MS consists of a buffer-gas-filled RFQ ion transport section followed by a linear Paul trap and the ToF mass analyzer. The linear Paul trap prepares and injects cold ion bunches into the mass analyzer, a symmetric assembly of two opposing ion mirrors. At the end of the ToF analysis, the ions are released through the back mirror onto a MagneTOF detector [41] which records the ion ToF. During this experiment, the MR-ToF-MS was operated in a mass-selective retrapping mode [42,43], allowing the device

to operate as an isobaric mass purifier and mass spectrometer. To this end, the ions were injected at two separate stages into the mass analyzer. The mass-selective retrapping was done after 20–40 isochronous turns in the mass analyzer. The ions were then injected back into the mass analyzer for the mass measurement stage, which involved 756–842 isochronous laps, resulting in mass-resolving powers of $R = \frac{m}{\Delta m} = 400\,000\text{--}530\,000$. The mass-selective retrapping mode strongly suppressed isobaric contamination, allowing a factor of 10 higher beam intensity during the measurement of ^{74}Sr . At each mass unit, the resonant laser ionization scheme was temporarily blocked, resulting in a substantial reduction in Sr yield (see Fig. 1), thus verifying the correct identification of isotope in the mass spectrum. More details of the TITAN MR-ToF-MS can be found in Ref. [40].

During the experiment, each ToF was converted to a mass-to-charge ratio ($\frac{m}{z}$) using the relation [44]

$$\frac{m}{z} = \frac{c(t - t_0)^2}{(1 + bN_{\text{it}})^2}, \quad (2)$$

TABLE I. Summary of mass measurements performed in this work. All masses were measured as singly charged ions. The third column gives N_{counts} , which is the number of counts determined by the area of the fitted mass peak.

Nuclide	$T_{1/2}$ (ms)	N_{counts}	Mass calibrant	Mass excess (keV)		
				Literature values	TITAN	TITAN - Literature
^{74}Sr	27.6(2.7)	36(12)	$^{148}\text{Sm}^{2+}$	-40 830#(100#) ^a	-40 970(31)	-140(105)
^{75}Sr	88(3)	424(22)	$^{75}\text{Rb}^+$	-46 620(220) ^a -46 200(150) ^b	-46 273(10)	347(220)
^{76}Sr	7 890(70)	2 108(46)	$^{76}\text{Rb}^+$	-54 250(30) ^a	-54 257(9)	-7(31)

^aValues from the Atomic Mass Evaluation (AME) 2020 [48] (# indicates extrapolations on the mass surface).

^bValue determined at the CSRe [21].

where t_0 , b , and c are calibration parameters, and N_{it} is the number of isochronous turns completed by the species. The calibration parameters c and t_0 were determined before the experiment with reference ions of well-known mass.

After the experiment, a time-resolved calibration was performed to correct for low-frequency (sub-Hz) fluctuations in the mirror electrode voltages and thermal expansion of the mass analyzer that cause a fluctuation in the ToF during the measurement. Time-resolved calibrations were performed for each spectrum according to the procedure outlined in Ref. [44]. To deduce peak positions and their respective mass values, each spectrum was fitted using the maximum likelihood approach implemented in the emgfit PYTHON package [45]. The fits were performed using hyper-exponentially modified Gaussian (hyper-EMG) peak shapes [46] whose tail order was determined using the procedure described in Ref. [47]. After fitting, a final mass calibration was performed using mass peaks in the spectrum with well-known literature values. Isobaric rubidium ions were used to calibrate $^{75,76}\text{Sr}$. Due to a lack of suitable isobaric reference species, the ^{74}Sr mass was determined using $^{148}\text{Sm}^{2+}$ as a mass calibrant. Statistical and systematic mass uncertainties were evaluated following the procedures and methods described in Refs. [44,45,47]. For $^{75,76}\text{Sr}$, a nonideal extraction of the ions from the mass analyzer dominated the mass uncertainty, while for ^{74}Sr the dominant contribution was statistical uncertainty.

III. RESULTS

The mass results are summarized in Table I with a comparison to literature values. Our measurement of ^{76}Sr is in excellent agreement with a Penning trap mass measurement at ISOLTRAP [49], providing independent confirmation of the value and improving the mass precision by a factor of 3. Recently, the experimental cooler storage ring (CSRe) in Lanzhou reported the first direct mass measurement of ^{75}Sr [21]. Our measurement is a variance-weighted mean determined from three independent measurements and improves upon the CSRe mass precision by a factor of 15. Our value is in good agreement with the CSRe value within uncertainty. Our final mass value of ^{74}Sr is a variance-weighted mean obtained from two independent measurements. This constitutes the first time that the ground-state mass of ^{74}Sr has been measured.

To visualize the staggering behavior of the doublet b coefficients, we consider the quantity

$$\Delta b = b(A) - b(A - 2), \quad (3)$$

which is plotted in Fig. 2. Data from the AME2020 updated with recent storage ring measurements [21] are shown with the black curve. As explained in Ref. [20], the mass measurements of ^{71}Kr and ^{75}Sr reported in Ref. [21] removed the staggering anomaly previously reported in Refs. [16,19]. Theoretical data from a modified homogeneously charged-sphere approximation [11] and density functional theory (DFT) [17] do well to reproduce the experimental data. However, the measurement of ^{75}Sr reported by Wang *et al.* [21] was limited to four counts and is therefore dominated by a large uncertainty of 150 keV. As seen from the red curve in Fig. 2, the inclusion of our new ^{75}Sr value is in good agreement with the previous value and also reduces the uncertainty by a factor of 15. This in turn supports the rectification of the doublet staggering phase that was shown in Ref. [20].

An interesting feature of the doublet b coefficients is a dampening of the staggering amplitude in the $A = 43\text{--}55$ $f_{7/2}$ subshell (visible in Fig. 2). Theoretical models have demonstrated some success in reproducing this feature, although

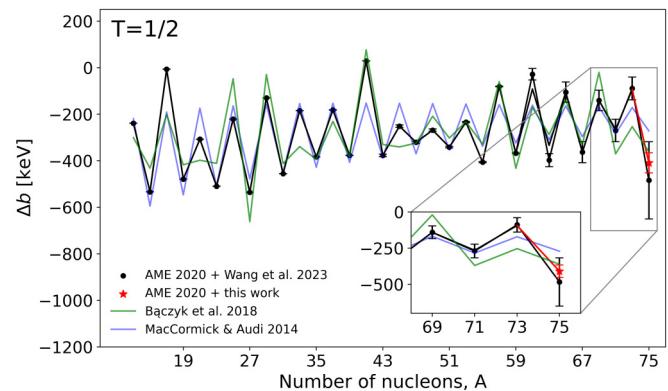


FIG. 2. The difference between the doublet b coefficients of adjacent isospin doublets. AME2020 values updated with recent storage ring measurements [21] (black) show the characteristic staggering pattern from $A = 10\text{--}75$. Our measurement of ^{75}Sr improves upon the precision of the value reported in Ref. [21] and confirms the continued staggering phase of Δb up to $A = 75$. For comparison, DFT calculations (green) [50] and a modified homogeneously charged-sphere model (blue) [11] show good agreement.

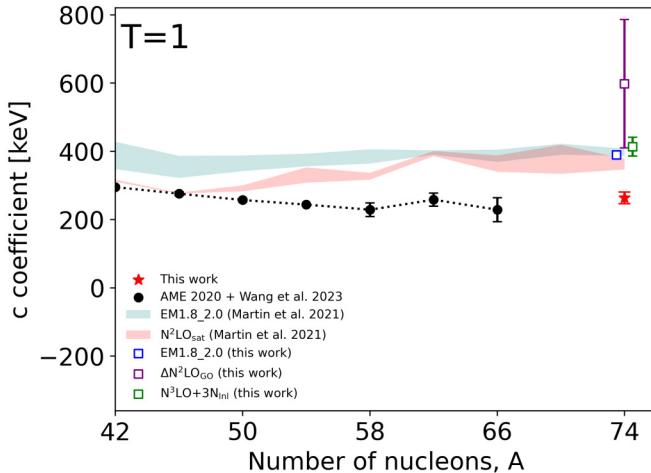


FIG. 3. The triplet c coefficients for $A = 4n + 2$ nuclei. AME2020 values updated with storage ring measurements of ^{58}Zr , ^{62}Ge , and ^{66}Se [21] (black) show the relatively constant trend of the c coefficient. Our new value of ^{74}Sr (red star) shows that this relatively constant trend continues. VS-IMSRG calculations (filled curves) from Ref. [25] are in good agreement, but consistently overcalculate the coefficient by 50–150 keV. VS-IMSRG calculations using two new χ EFT-derived interactions at $A = 74$ (empty squares) do not show a marked improvement.

they are in disagreement about the underlying mechanisms responsible for it. An extended Skyrme pn -mixed DFT approach [50] reproduced the $f_{7/2}$ dampening, but noted that the inclusion of isospin nonconserving forces increased the staggering amplitude. A nuclear shell-model approach with modern effective interactions [16] also reproduced the $f_{7/2}$ dampening, although the staggering amplitude was instead decreased with the inclusion of isospin nonconserving forces. In addition to the $f_{7/2}$ dampening, Bączyk *et al.* [17,50] predicted a dampening in the $g_{9/2}$ subshell at $A = 83$. Precision mass measurements at higher masses are needed to test these theoretical indications.

Experimentally evaluated triplet data have recently become available up to $A = 66$ through direct mass measurements of ^{58}Zn [21], ^{60}Ga [47], ^{62}Ge [21], and ^{66}Se [21,51]. Experimental evaluations at higher masses have so far been impossible due to a lack of data. Our measurement of ^{74}Sr completes the $A = 74$ triplet, which is now the heaviest experimentally evaluated triplet by 8 mass units. The experimental trend of the triplet c coefficients is shown for $A = 4n + 2$ nuclei in Fig. 3, including the new data point. This confirms that the relatively constant trend displayed by the $A = 4n + 2c$ coefficients continues to higher masses.

Results from a theoretical VS-IMSRG calculation by Martin *et al.* [25] predict a similarly constant trend, but consistently overcalculate by 50–150 keV and diverge near $A = 54$ and 58. In a follow-up to the calculations by Martin *et al.* [25], we perform new VS-IMSRG calculations for the $A = 74$ triplet. In the original work, two sets of χ EFT-derived NN + 3N interactions were used: 1.8/2.0 (EM) [52], which reproduce ground-state energies across the medium-mass region [53,54] and $N^2\text{LO}_{\text{sat}}$ [55]. The authors

used a storage truncation on 3N forces of $e_{3\text{max}} = 16$, but we have extended to $e_{3\text{max}} = 24$ using a novel storage scheme [56]. We then explored the chiral nuclear force with two new χ EFT-derived interactions: $\Delta N^2\text{LO}_{\text{GO}}$ [57] and $N^3\text{LO} + 3\text{N}_{\text{int}}$ [58,59]. Following the same procedure as Martin *et al.* we estimated the IMSRG truncation error using different choices for the reference nucleus: (i) self-referenced, (ii) $T_z = +1$, (iii) $T_z = 0$, and (iv) $T_z = -1$.

It can be seen in Fig. 3 that increasing the model space for 3N forces in the 1.8/2.0 (EM) interaction does not significantly improve agreement with the experimental data. It can also be seen that neither of the two new interactions, $\Delta N^2\text{LO}_{\text{GO}}$ and $N^3\text{LO} + 3\text{N}_{\text{int}}$, show a marked improvement. However, there are still avenues that VS-IMSRG can take to improve agreement with experimental results. IMSRG(3) is a relatively new iteration of the VS-IMSRG method [60], which explicitly keeps three-body operators during the flow. While it is not yet computationally feasible to perform IMSRG(3) at such high masses, it is expected that better agreement with experiment would be demonstrated.

The new strontium mass values also affect the nuclear reaction flows of the rp-process in x-ray bursts. In particular, the new more precise ^{75}Sr mass removes mass uncertainties from determining the $^{74}\text{Rb}(p, \gamma)$ vs β -decay branching. This branching affects the rp-process production of the p nucleus ^{74}Se [61–63] and is thus important for assessing the possibility of the rp-process contributing to the origin of the light p nuclei. The improved ^{76}Sr mass provides an important constraint for the Q value of the $^{76}\text{Sr}(p, \gamma)^{77}\text{Y}$ reaction and thus on a possible proton-capture bypass of the ^{76}Sr waiting point in the rp-process [61]. The precise measurement of the ^{76}Sr mass is an important first step to reduce the uncertainty of this Q value.

To investigate the impact of the new strontium mass values on predictions of type-I x-ray bursts, we used a self-consistent single-zone model [31]. The single-zone model was first benchmarked using reaction rate Q values determined from current mass uncertainties in the AME2020. The proton-capture Q values were extremized by varying the masses of the respective initial and final nucleus by $\pm 3\sigma$. The Hauser-Feshbach reaction code TALYS 2.0 [64] was then used to calculate the astrophysical reaction rates $^{73}\text{Rb}(p, \gamma)^{74}\text{Sr}$, $^{74}\text{Rb}(p, \gamma)^{75}\text{Sr}$, $^{75}\text{Rb}(p, \gamma)^{76}\text{Sr}$, and $^{76}\text{Sr}(p, \gamma)^{77}\text{Y}$. The competing photodisintegration rates were calculated from detailed balance. Included in the calculations is an estimate of the unmeasured ^{77}Y mass, which was calculated from the well-known ^{77}Sr using the Coulomb displacement energy method [65]. To evaluate the impact of the new mass data on x-ray burst observables, the same procedure was then repeated using the new mass values.

The new masses better pin down the rp-process reaction flow between the ^{72}Kr and ^{76}Sr waiting points. Namely, we find a stronger ^{74}Rb β -decay branch and a reduction in mass flowing beyond the ^{76}Sr waiting point. The uncertainties induced by nuclear masses for the ^{76}Sr waiting point bypass $^{76}\text{Sr}(p, \gamma)$ have been reduced from 50% to 20%, with the remaining uncertainty arising from the unmeasured ^{77}Y mass that is needed together with the ^{76}Sr mass to obtain the important ^{76}Sr p -capture Q value. The new masses reduce the

^{76}Sr bypass flow further from 2% to just 0.7%, which leads to an increased production of stable $A = 74$ ashes. The previous mass-induced uncertainty for the $A = 74$ ashes was 16% and is now a negligible 0.7%. A more precise future measurement of the ^{77}Y mass would be desirable to confirm these results.

IV. CONCLUSION

In summary, precision mass measurements of neutron-deficient strontium ions were performed with TITAN's MR-ToF-MS, providing a first time measurement of ^{74}Sr and improved mass precision of both ^{75}Sr and ^{76}Sr . These measurements complete the heaviest experimentally evaluated isospin doublet measured to date and give improved precision to the heaviest isospin triplet, allowing us to study isospin symmetry breaking using the IMME. Our measurement of ^{75}Sr is in good agreement with a storage ring measurement [21] and improves upon the precision by a factor of 15. This value supports the rectification of the doublet staggering phase that was shown in Ref. [20]. The new data also extend the IMME coefficients for isospin triplets to $A = 74$, which is far higher than has been previously reached and confirms the constant trend of the c coefficients. New VS-IMSRG calculations at $A = 74$ were performed to explore the chiral nuclear force with two new interactions, but did not show a marked improvement over previous calculations by Martin *et al.* [25]. In a single zone x-ray burst model, the new masses altered

the reaction flow in the Sr region significantly, though the resulting impact on observables was found to be relatively small.

ACKNOWLEDGMENTS

TITAN is funded by the Natural Sciences and Engineering Research Council (NSERC) of Canada and through TRI-UMF by the National Research Council (NRC). This work was partially supported by the National Science Foundation under Awards No. PHY-1430152 (JINA Center for the Evolution of the Elements), No. OISE-1927130 (IReNA), and No. PHY-2209429; the U.S. Department of Energy, Office of Science, under SciDAC-5 (NUCLEI Collaboration); the German Federal Ministry for Education and Research (BMBF) under Contracts No. 05P19RGFN1 and No. 05P21RGFN1; the German Research Foundation (DFG) under Contract No. 422761894; HGS-HIRe; and Justus-Liebig-Universität Gießen and GSI under the JLU-GSI strategic Helmholtz partnership agreement. The IMSRG code used is ragnar_imsrsg [66]. Theoretical efforts were further supported by NSERC under Grants No. SAPIN-2018-00027 and No. RGPAS-2018-522453 and the Arthur B. McDonald Canadian Astroparticle Physics Research Institute. Computations were based on the performed with an allocation of computing resources on Cedar at WestGrid and Compute Canada.

[1] E. Tiesinga, P. J. Mohr, D. B. Newell, and B. N. Taylor, *J. Phys. Chem. Ref. Data* **50**, 033105 (2021).

[2] Y. H. Lam, N. A. Smirnova, and E. Caurier, *Phys. Rev. C* **87**, 054304 (2013).

[3] M. Bentley and S. Lenzi, *Prog. Part. Nucl. Phys.* **59**, 497 (2007).

[4] W. Ormand and B. Brown, *Nucl. Phys. A* **491**, 1 (1989).

[5] J. Nolen, Jr., and J. Schiffer, *Annu. Rev. Nucl. Sci.* **19**, 471 (1969).

[6] J. C. Hardy and I. S. Towner, *Phys. Rev. C* **102**, 045501 (2020).

[7] I. S. Towner and J. C. Hardy, *Phys. Rev. C* **77**, 025501 (2008).

[8] W.-J. Ong *et al.*, *Phys. Rev. C* **95**, 055806 (2017).

[9] M. Saxena *et al.*, *Phys. Lett. B* **829**, 137059 (2022).

[10] O. Klochko and N. A. Smirnova, *Phys. Rev. C* **103**, 024316 (2021).

[11] M. MacCormick and G. Audi, *Nucl. Phys. A* **925**, 61 (2014).

[12] E. Feenberg and G. Goertzel, *Phys. Rev.* **70**, 597 (1946).

[13] J. Jänecke, *Phys. Rev.* **147**, 735 (1966).

[14] J. Jänecke, Systematics of Coulomb energies and excitation energies of isobaric analogue states, in *Proceedings of the Third International Conference on Atomic Masses, 28 August–1 September, 1967 at University of Manitoba, Winnipeg, Canada*, edited by R. C. Barber (University of Manitoba Press, Winnipeg, 1967), p. 583.

[15] W. Huang, M. Wang, F. Kondev, G. Audi, and S. Naimi, *Chin. Phys. C* **45**, 030002 (2021).

[16] K. Kaneko, Y. Sun, T. Mizusaki, and S. Tazaki, *Phys. Rev. Lett.* **110**, 172505 (2013).

[17] P. Bączyk, W. Satuła, J. Dobaczewski, and M. Konieczka, *J. Phys. G: Nucl. Part. Phys.* **46**, 03LT01 (2019).

[18] H. H. Li *et al.*, *Phys. Rev. C* **107**, 014302 (2023).

[19] X. Tu, Y. Sun, Y. Zhang, H. Xu, K. Kaneko, Y. A. Litvinov, and M. Wang, *J. Phys. G: Nucl. Part. Phys.* **41**, 025104 (2014).

[20] H. F. Li, X. Xu, Y. Sun, K. Kaneko, X. Zhou, M. Zhang, W. J. Huang, X. H. Zhou, Y. A. Litvinov, M. Wang, and Y. H. Zhang, *Phys. Rev. C* **110**, L021301 (2024).

[21] M. Wang *et al.*, *Phys. Rev. Lett.* **130**, 192501 (2023).

[22] C.-Y. Seng, M. Gorchtein, H. H. Patel, and M. J. Ramsey-Musolf, *Phys. Rev. Lett.* **121**, 241804 (2018).

[23] C.-Y. Seng, M. Gorchtein, and M. J. Ramsey-Musolf, *Phys. Rev. D* **100**, 013001 (2019).

[24] A. Czarnecki, W. J. Marciano, and A. Sirlin, *Phys. Rev. D* **100**, 073008 (2019).

[25] M. S. Martin, S. R. Stroberg, J. D. Holt, and K. G. Leach, *Phys. Rev. C* **104**, 014324 (2021).

[26] R. Stroberg, Isospin breaking in the IMSRG, PAINT2023 Workshop on Progress in *Ab Initio* Nuclear Theory (2023), <https://indico.triumf.ca/event/349/contributions/4257/>

[27] M. Gennari, Standard model corrections to Fermi decays in NCSM, PAINT2023 Workshop on Progress in *Ab Initio* Nuclear Theory (2023), <https://indico.triumf.ca/event/349/contributions/4294/>.

[28] H. Hergert, S. K. Bogner, T. D. Morris, A. Schwenk, and K. Tsukiyama, *Phys. Rep.* **621**, 165 (2016).

[29] S. R. Stroberg, A. Calci, H. Hergert, J. D. Holt, S. K. Bogner, R. Roth, and A. Schwenk, *Phys. Rev. Lett.* **118**, 032502 (2017).

[30] T. Miyagi, S. R. Stroberg, J. D. Holt, and N. Shimizu, *Phys. Rev. C* **102**, 034320 (2020).

[31] H. Schatz, A. Aprahamian, V. Barnard, L. Bildsten, A. Cumming, M. Ouellette, T. Rauscher, F.-K. Thielemann, and M. Wiescher, *Phys. Rev. Lett.* **86**, 3471 (2001).

[32] J. Dilling, P. Bricault, M. Smith, and H.-J. Kluge, *Nucl. Instrum. Methods Phys. Res., Sect. B* **204**, 492 (2003).

[33] C. Jesch *et al.*, *Hyperfine Interact.* **235**, 97 (2015).

[34] T. Dickel *et al.*, *Hyperfine Interact.* **240**, 62 (2019).

[35] Y. Blumenfeld, T. Nilsson, and P. V. Duppen, *Phys. Scr. T152*, 014023 (2013).

[36] J. Dilling, R. Krücke, and L. Merminga, Ariel overview, in *ISAC and ARIEL: The TRIUMF Radioactive Beam Facilities and the Scientific Program*, edited by J. Dilling, R. Krücke, and L. Merminga (Springer Netherlands, Dordrecht, 2014), pp. 253–262.

[37] M. Baig, M. Yaseen, R. Ali, A. Nadeem, and S. Bhatti, *Chem. Phys. Lett.* **296**, 403 (1998).

[38] J. Lassen, *Spectrochim. Acta, Part B* (unpublished).

[39] T. Brunner *et al.*, *Nucl. Instrum. Methods Phys. Res., Sect. A* **676**, 32 (2012).

[40] M. Reiter *et al.*, *Nucl. Instrum. Methods Phys. Res., Sect. A* **1018**, 165823 (2021).

[41] Magnetof: A new class of robust sub-nanosecond tof detectors with exceptional dynamic range, <https://www.etp-ms.com/file-repository/8>, accessed: 2023-08-30.

[42] T. Dickel *et al.*, *Nucl. Instrum. Methods Phys. Res., Sect. A* **777**, 172 (2015).

[43] T. Dickel, M. I. Yavor, J. Lang, W. R. Plaß, W. Lippert, H. Geissel, and C. Scheidenberger, *Int. J. Mass Spectrom.* **412**, 1 (2017).

[44] S. Ayet San Andrés *et al.*, *Phys. Rev. C* **99**, 064313 (2019).

[45] S. F. Paul, emgfit, Fitting of time-of-flight mass spectra with Hyper-EMG models (2022), <https://pypi.org/project/emgfit/>.

[46] S. Purushothaman *et al.*, *Int. J. Mass Spectrom.* **421**, 245 (2017).

[47] S. F. Paul *et al.*, *Phys. Rev. C* **104**, 065803 (2021).

[48] M. Wang, W. Huang, F. Kondev, G. Audi, and S. Naimi, *Chin. Phys. C* **45**, 030003 (2021).

[49] G. Sikler *et al.*, *Nucl. Phys. A* **763**, 45 (2005).

[50] P. Bączyk, J. Dobaczewski, M. Konieczka, W. Satuła, T. Nakatsukasa, and K. Sato, *Phys. Lett. B* **778**, 178 (2018).

[51] X. Zhou *et al.*, *Nat. Phys.* **19**, 1091 (2023).

[52] J. Simonis, S. R. Stroberg, K. Hebeler, J. D. Holt, and A. Schwenk, *Phys. Rev. C* **96**, 014303 (2017).

[53] T. D. Morris *et al.*, *Phys. Rev. Lett.* **120**, 152503 (2018).

[54] S. R. Stroberg, J. D. Holt, A. Schwenk, and J. Simonis, *Phys. Rev. Lett.* **126**, 022501 (2021).

[55] A. Ekström, G. R. Jansen, K. A. Wendt, G. Hagen, T. Papenbrock, B. D. Carlsson, C. Forssén, M. Hjorth-Jensen, P. Navrátil, and W. Nazarewicz, *Phys. Rev. C* **91**, 051301(R) (2015).

[56] T. Miyagi, S. R. Stroberg, P. Navrátil, K. Hebeler, and J. D. Holt, *Phys. Rev. C* **105**, 014302 (2022).

[57] W. G. Jiang, A. Ekström, C. Forssén, G. Hagen, G. R. Jansen, and T. Papenbrock, *Phys. Rev. C* **102**, 054301 (2020).

[58] E. Leistenschneider *et al.*, *Phys. Rev. Lett.* **120**, 062503 (2018).

[59] V. Somà, P. Navrátil, F. Raimondi, C. Barbieri, and T. Duguet, *Phys. Rev. C* **101**, 014318 (2020).

[60] M. Heinz, A. Tichai, J. Hoppe, K. Hebeler, and A. Schwenk, *Phys. Rev. C* **103**, 044318 (2021).

[61] H. Schatz *et al.*, *Phys. Rep.* **294**, 167 (1998).

[62] N. N. Weinberg, L. Bildsten, and H. Schatz, *Astrophys. J.* **639**, 1018 (2006).

[63] Y. Herrera, G. Sala, and J. José, *Astron. Astrophys.* **678**, A156 (2023).

[64] A. Koning, S. Hilaire, and S. Goriely, *Eur. Phys. J. A* **59**, 131 (2023).

[65] H. Schatz and W.-J. Ong, *Astrophys. J.* **844**, 139 (2017).

[66] S. R. Stroberg, <https://github.com/ragnarstroberg/imsrg>.